

National Aeronautics and  
Space Administration



# EXPLORE MOON<sub>to</sub>MARS

## SCALPSS Project Overview

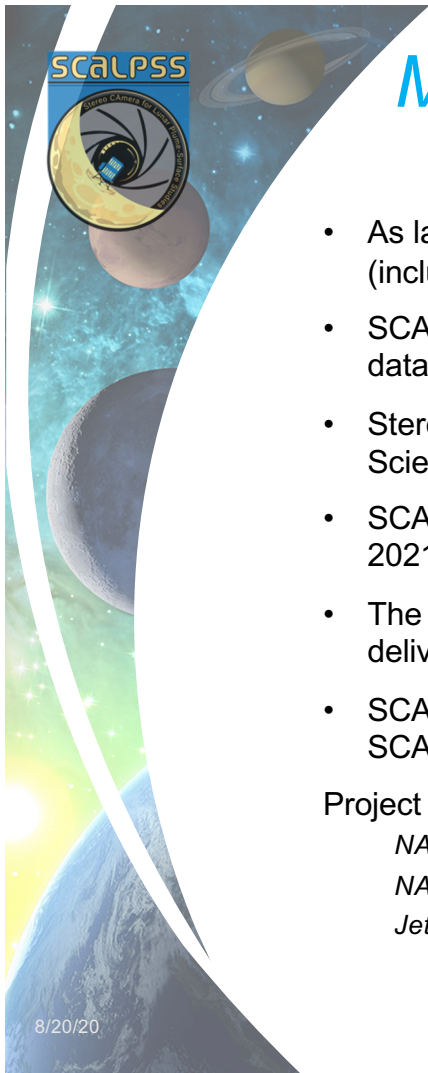
Lunar Surface Science Workshop

20 August 2020

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## Motivation and Project Overview



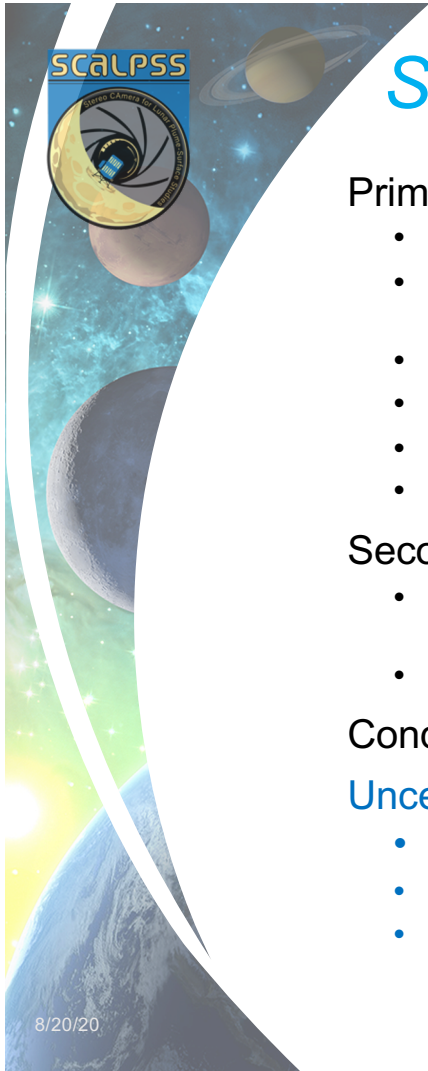
- As landers grow to human scale, understanding how these landers interact with the environment (including other landed assets), particularly during the landing event, becomes increasingly critical.
- SCALPSS is the first payload specifically designed to collect Plume Surface Interaction (PSI) flight data for verification of computational models. Complements NASA modeling/ground testing project.
- Stereo CAmera for Lunar Plume-Surface Studies (SCALPSS) was selected in February 2019 by the Science Mission Directorate (SMD) NASA Provided Lunar Payload (NPLP) Development Program
- SCALPSS is scheduled to fly on the Intuitive Machines NOVA-C lander to Vallis Schröteri in October 2021, as part of the initial Commercial Lunar Payload Services (CLPS) mission selections
- The SCALPSS payload is currently undergoing environmental testing in support of a payload delivery in October 2020
- SCALPSS 1.1, planned for CLPS flight 19D to Mare Crisium in early 2023, will be a “re-fly” of SCALPSS incorporating lessons learned/minor design changes to reduce risk and enhance science

### Project partners include:

*NASA Langley Research Center (Lead)*

*NASA Marshall Space Flight Center*

*Jet Propulsion Laboratory / California Institute of Technology*



# SCALPSS Science Objectives



## Primary Science Objectives

- Detect onset altitude of ejecta and erosion effects (*turn on @ 30 m altitude*)
- Record transient 3D crater morphology as the vehicle is descending to estimate transient crater formation volume rates (*15 fps through landing*)
- Capture post-landing 3D site-alteration morphology
- Measure crater depth, diameter and eroded volumes
- Estimate erosion rates
- Quantify uncertainty in photogrammetry measurements (*post-landing images*)

## Secondary Science Objectives

- Measure ejecta structure during onset and descent (may be able distinguish ejection angles, ejecta sheath morphology, etc.)
- Measure dust settling density as a function of time (*10 s post-landing @ 15 fps*)

Conduct comparisons between predictions and flight reconstructed data

Uncertainties will remain, due to regolith unknowns:

- Bulk density (compaction, porosity)
  - Particle size distribution
  - Particle shape and cohesion
- } At depth, down to [20 cm] (?)



# Stereo Photogrammetry

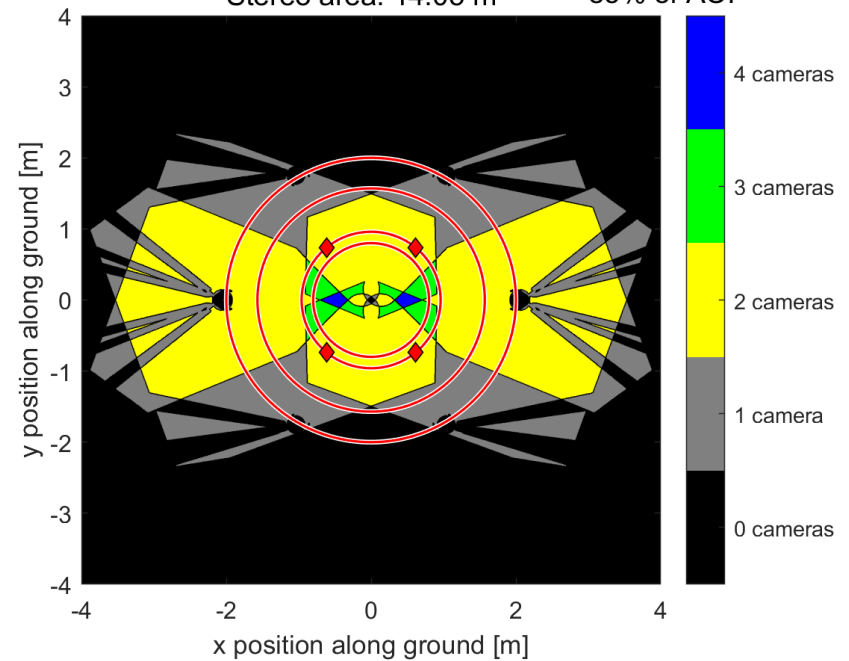
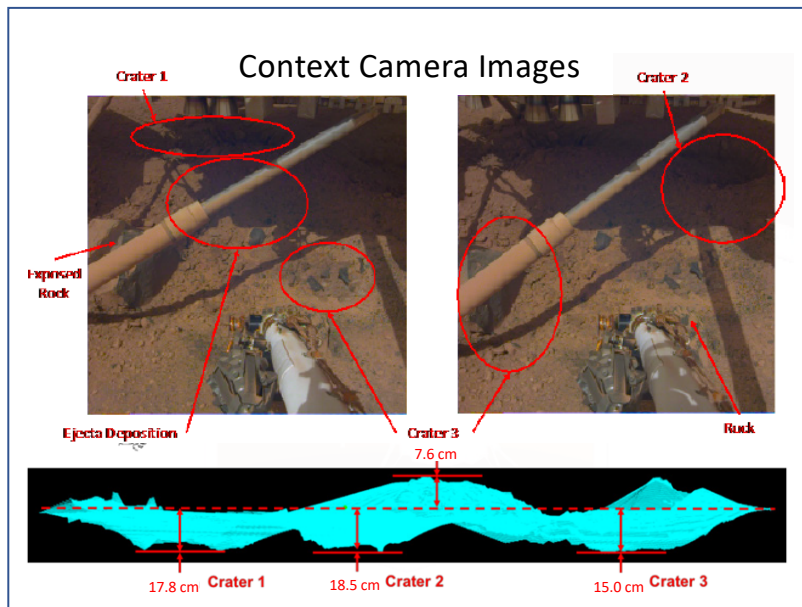


Mars InSight Topography Analysis

Stereo Coverage of SCALPSS Cameras

$R = 0.96 \text{ m}$ ,  $h = 0.95 \text{ m}$ ,  $\theta = 40^\circ$ ,  $\omega = 30^\circ$ ,  $\phi = 5^\circ$ ,  $\alpha = 0^\circ$ ,  $\beta = 90^\circ$

Stereo area:  $14.05 \text{ m}^2$  85% of AOI



Estimated Accuracy of Pseudo-Stereo Method:  $\pm 0.25 \text{ cm}$   
(Mehta & Liever, 2019)



# SCALPSS Payload Architecture



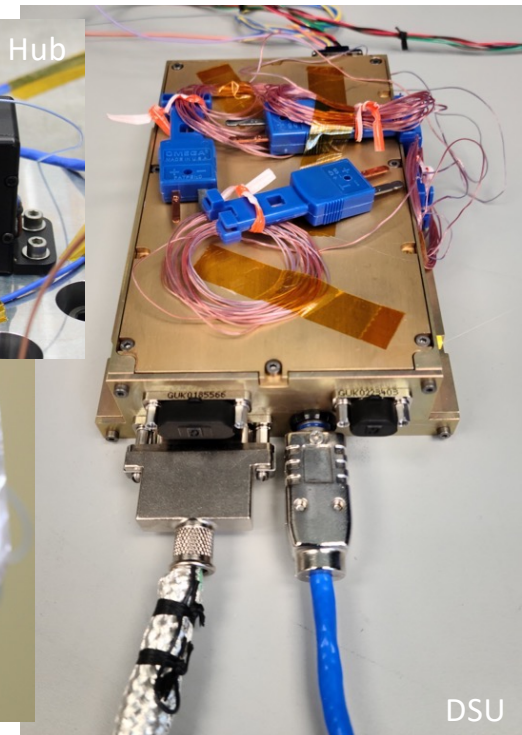
- SCALPSS is a distributed payload with 6 separate components
- Data Storage Unit (DSU)
  - build-to-print of M2020 EDLCam electronics
- USB Hub
  - hardened; same COTS part as used for M2020 EDLCam
- 2 pairs (4 ea.) short focal-length high resolution down-look cameras
  - COTS camera; same make as M2020 EDLCam cameras, but monochrome model
  - COTS lenses; short focal-length lenses used for SLS
  - optimized locations and pointing to maximize stereo coverage



USB Hub



Camera



DSU

A cosmic background image featuring a blue nebula in the upper right and a green nebula in the lower right, with numerous stars scattered throughout. A light blue horizontal band is centered across the image.

Questions?