

Simple, compact, high-resolution monochromatic x-ray source for characterization of x-ray calorimeter arrays

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ABSTRACT

X-ray calorimeters routinely achieve very high spectral resolution, typically a few eV full width at half maximum (FWHM). Measurements of calorimeter line shapes are usually dominated by the natural linewidth of most laboratory calibration sources. This compounds the data acquisition time necessary to statistically sample the instrumental line broadening and can add systematic uncertainty if the intrinsic line shape of the source is not well known. To address these issues, we have built a simple, compact monochromatic x-ray source using channel cut crystals. A commercial x-ray tube illuminates a pair of channel cut crystals that are aligned in a dispersive configuration to select the $K\alpha_1$ line of the x-ray tube anode material. The entire device, including the x-ray tube, can be easily hand-carried by one person and may be positioned manually or using a mechanical translation stage. The output monochromatic beam provides a collimated image of the anode spot with magnification of unity in the dispersion direction (typically $100\ \mu\text{m}$ – $200\ \mu\text{m}$ for the x-ray tubes used here) and is unfocused in the cross-dispersion direction so that the source image in the detector plane appears as a line. We measured output count rates as high as 10 count/s/pixel for the Hitomi soft x-ray spectrometer, which had $819\ \mu\text{m}$ square pixels. We implemented different monochromator designs for energies of 5.4 keV (one design) and 8.0 keV (two designs), which have effective theoretical FWHM energy resolution of 0.125 eV, 0.197 eV, and 0.086 eV, respectively; these are well-suited for optimal calibration measurements of state-of-the-art x-ray calorimeters. We measured an upper limit for the energy resolution of our Cr $K\alpha_1$ monochromator of 0.7 eV FWHM at 5.4 keV, consistent with the theoretical prediction of 0.125 eV.

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I. INTRODUCTION

X-ray calorimeters have been developed over the last 30 years with the aim of providing high-resolution, high-efficiency imaging spectroscopy for astrophysics and other applications.¹ The Hitomi Soft X-ray Spectrometer (SXS), a production space flight spectrometer with a 36 pixel array of Si thermistor sensors with HgTe x-ray absorbers, achieved better than 5 eV full width at half maximum (FWHM) resolution at 6 keV in ground testing² and 5 eV on orbit,³

and detectors based on Transition Edge Sensors (TES) as well as Metallic Magnetic Calorimeters (MMCs) have achieved better than 2 eV resolution at 6 keV and better than 1 eV below 2 keV.^{4–9} Typical calorimeter pixels range in size from of order $0.01\ \text{mm}^2$ – $1\ \text{mm}^2$, depending on the application.

The spectral response of x-ray calorimeter detectors is dominated by a Gaussian broadening that is often characterized using Mn $K\alpha$ x rays from a radioactive ⁵⁵Fe source, which has a large advantage in terms of ease of use. The measured spectrum is a convolution

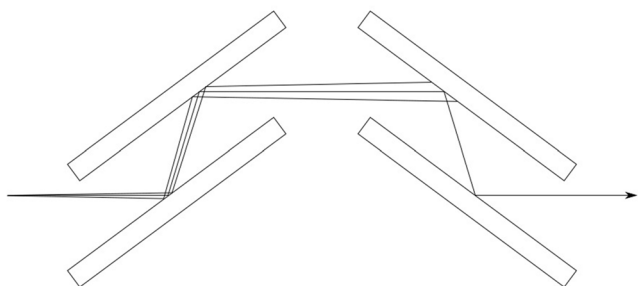


FIG. 1. Schematic diagram of a DuMond-Hart-Bartels monochromator illuminated by a divergent broadband point source. Rays of different energies satisfy the Bragg condition at different points on the first crystal pair, but only one ray satisfies the Bragg condition at the second crystal pair. (Actually, the angular width of the reflected bundle has a small but finite width. See Sec. IV.)

of the intrinsic line shape of Mn $K\alpha$ with the Gaussian core response of the detector. The intrinsic line shape of Mn $K\alpha$ is highly complex and is typically modeled using an eight Lorentzian empirical deconvolution.^{10,11} Since the Lorentzian components have linewidths of ~ 2 eV, the intrinsic width of the complex is a significant fraction of the measured broadening for state-of-the-art calorimeters. This adds considerably to the integration time required to measure the intrinsic linewidth to the desired statistical precision, given a fixed incident flux. This loss of experimental efficiency could be recovered with a sufficiently monochromatic x-ray source, which would return the maximum statistical precision achievable for the number of counts collected. Furthermore, characterization of non-Gaussian components of the detector response (from fluorescent escape photons, electron loss, and effects of incomplete thermalization)^{2,12} also

requires a monochromatic x-ray source to avoid confusion between the intrinsic source spectrum and non-Gaussian components.

Thus, there is a clear need for monochromatic sources for characterization of x-ray calorimeters. High performance monochromators for experiments at synchrotron facilities are well-developed, but travel to such facilities is prohibitive for the day-to-day activities of a cryogenic physics laboratory. Thus, we have built a series of small, portable monochromators using commercial table-top x-ray generators combined with channel cut crystals, and we have used these monochromators to characterize a number of x-ray calorimeter devices.

II. DESIGN

The scheme used in our monochromators is due to DuMond, who first proposed a four-reflection monochromator in the (+1, -1, -1, +1) configuration,¹³ as illustrated in Fig. 1, although it was not implemented at the time due to insufficient flux produced by x-ray generators. The main advantage of such a multiple reflection scheme is the suppression of reflectivity in the wings of the crystal rocking curve. A secondary advantage is the preservation of the original beam propagation direction. This design also has the interesting property that the output beam produces a collimated image of the x-ray generator spot in the dispersion direction with magnification of unity.

Subsequently, Beaumont and Hart and Bartels designed four-reflection monochromators for synchrotron beams.^{14,15} Hart simplified the alignment problem by using channels cut in monolithic crystals to provide pairs of aligned reflecting surfaces. The most important remaining degree of freedom in the alignment is the angle between the two channel cut crystals in the dispersion plane, which determines the energy passed by the monochromator.

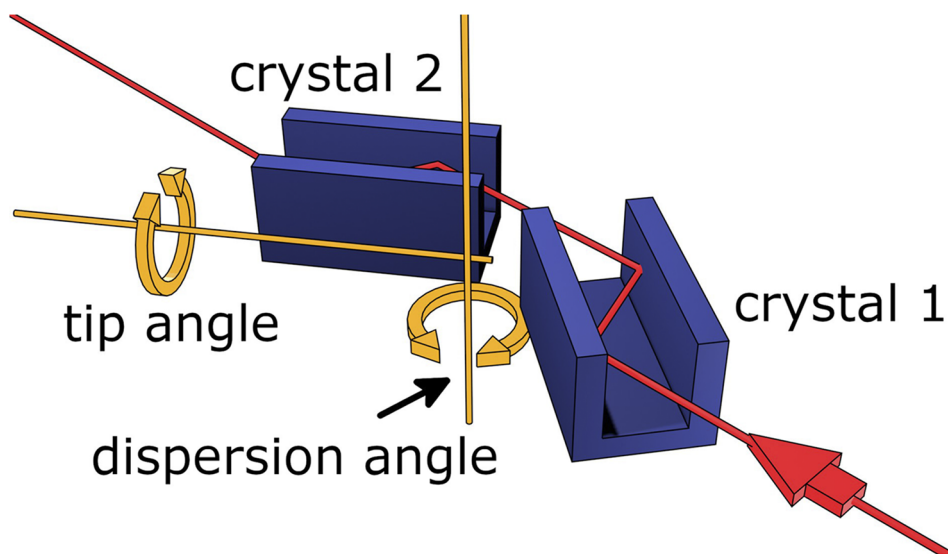


FIG. 2. Visualization of our implementation for a 2 + 2 reflection DHB monochromator using two channel cut crystals (blue). The beam path is indicated (red), as are the axes for adjusting the dispersion angle and tip angle of the second crystal (yellow). Note that adjusting the dispersion angle of the second crystal by θ is equivalent to an adjustment of both crystals by $\theta/2$ with respect to an imaginary plane of symmetry; this plane also rotates by $\theta/2$ with respect to the lab frame.

Such designs have been commonly implemented to take advantage of the bright x rays of synchrotron beamlines; however, because x-ray calorimeter arrays are typically designed to handle only relatively low count rates (e.g., ~ 1 count/s/pixel for the Hitomi SXS), we can achieve acceptable output flux using a commercial air-cooled 50 W x-ray generator.

To unambiguously characterize the performance of x-ray calorimeters, it is desirable to have a very monochromatic source with no flux in the line wings. Even a four-reflection monochromator design achieves this, although marginal decreases in FWHM can be achieved with more reflections, at the cost of some throughput. The number of reflections in a given crystal design can be chosen by changing the ratio of channel length to channel width for a given Bragg angle such that

$$l = \frac{2Nw}{\tan \theta_b}, \quad (1)$$

where l is the channel length, w is the channel width, θ_b is the Bragg angle, and $2N$ is the number of reflections per crystal in the design, with N constrained to be an integer. In this article, we use the notation $2N + 2N$ to denote the total number of reflections while calling attention to the symmetry in the optical design. All of our design implementations discussed in this article are for $4 + 4$ or $6 + 6$ reflection systems.

We designed the monochromators to be simple and compact, using commercially available parts to the extent possible. All alignment stages are operated by hand, and the number of degrees of freedom has been minimized to five: cross-dispersion translation and dispersion angle for each crystal, and the relative tip angle of the two crystals. In Fig. 2, we show an illustration of a $2 + 2$ reflection

channel cut crystal monochromator concept, indicating the path of the beam, as well as the axes for adjustment of the dispersion angle and relative tip angle of the second (downstream) crystal.

III. IMPLEMENTATION

In Fig. 3, we show one of our DuMond–Hart–Bartels (DHB) monochromators. A 0.5 in. thick aluminum enclosure provides stable mechanical support for the two channel cut crystals, which are mounted on commercial mirror mount stages, which are in turn mounted on commercial one-dimensional linear positioning stages. An Oxford¹⁶ Jupiter 5000 series x-ray tube is mounted at the entrance aperture of the enclosure. Three shafts allow external manipulation of the dispersion angle of each crystal, as well as the relative tip angle of the second crystal. The one-dimensional linear positioning stages are needed only for initial coarse adjustment and can be accessed by removing the walls of the enclosure from the base plate. The internal plate and back plate absorb any scattered x rays, allowing photons to travel only along the intended beam path.

In Fig. 4, we show one of the crystals used in our monochromators after it has been bonded to its substrate. The strain relief channel is visible closest to the bond. The crystals are supplied by Crystal Scientific,¹⁷ and their manufacturing specifications typically guarantee surface orientation tolerances to within 0.02° – 0.05° , spatial tolerances to within 0.05 mm–0.1 mm, and rocking curve FWHM tolerances to within 0.1 arc sec of the theoretical value.

In Table I, we give the dimensions of the crystals used in our monochromators. The ratio of channel width to length is chosen to allow potential use of the full channel width for the design energy

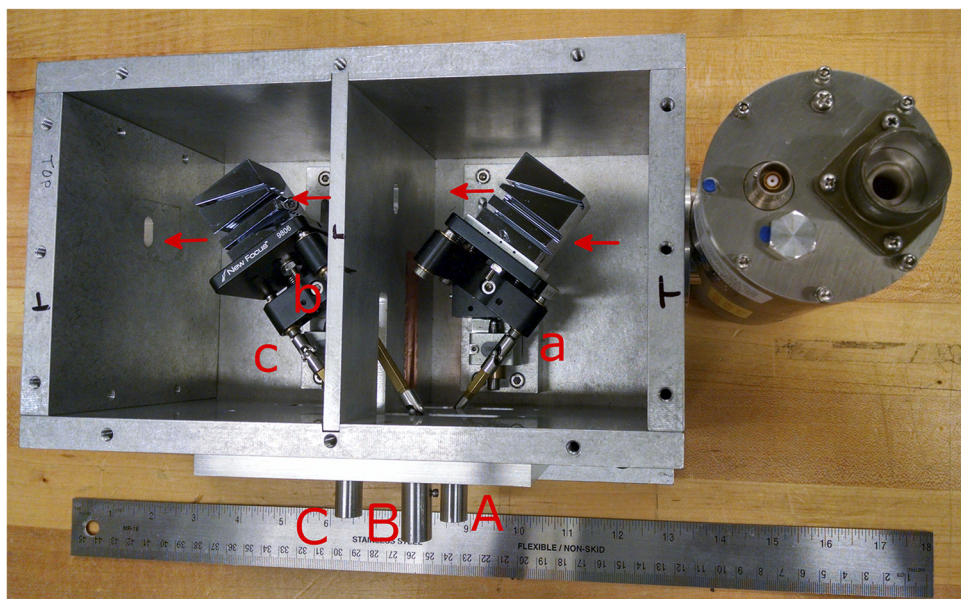


FIG. 3. Implementation of the $\text{Cr } K\alpha_1$ Si (220) $6 + 6$ reflection monochromator. The top plate has been removed to view the interior. The x-ray tube is the cylinder on the right, and the beam propagates from right to left. The arrows roughly indicate the propagation of the beam outside of the crystal channels. Uppercase labels denote external actuators, while lowercase labels refer to the corresponding adjusters on the optical mounting stage. A: First crystal dispersion angle; B: second crystal tip angle; C: second crystal dispersion angle.

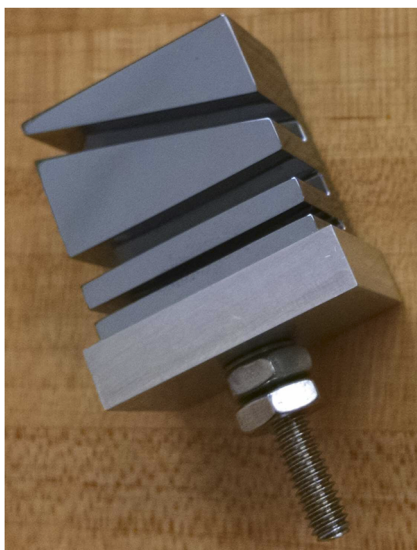


FIG. 4. Example of the Si (220) channel cut crystal attached to the substrate with mounting hardware. The crystal is bonded to the substrate with beeswax. The channel closest to the substrate is designed to prevent propagation of strain from the bond to the reflecting channels. The middle channel is for symmetric Bragg reflection, while the top channel is cut at an angle to the Si (220) plane and can be used for asymmetric Bragg reflection to achieve higher throughput at the price of lower resolution. We only used symmetric reflection in our monochromator.

using Eq. (1), with the exception of the Cr $K\alpha_1$ 6 + 6 reflection monochromator, for which we repurposed crystals designed for a 4 + 4 reflection Cu $K\alpha_1$ monochromator, taking advantage of the fact that the width-to-length ratio is similar for those two designs, so that almost the full width of the channel is usable. The widths were chosen to allow illumination of a significant part of a calorimeter array with the collimated image (in the dispersion direction) of an x-ray generator filling the full width of the channel; although the x-ray generators used here had much smaller spot sizes (see Sec. V B 1), in the future, we plan to retrofit x-ray generators with larger spot sizes.

A. Alignment

The crystals are aligned using a straightforward procedure. The first crystal (closest to the x-ray tube) is installed and coarsely aligned using a ruler and protractor, while the second crystal is not yet installed. A photon-counting x-ray detector with comparatively large area such as a silicon drift detector or proportional counter is positioned at the expected location of the output beam, coarsely

measured using a ruler. The x-ray generator is turned on, and the first crystal reflection angle is varied until the maximum intensity is achieved. Then, the generator is turned off, and the second crystal is installed and coarsely aligned. The x-ray detector is placed at the beam exit of the second crystal. The x-ray generator is turned on again, and the output intensity is maximized by varying the second crystal reflection angle, corresponding to the bright $K\alpha_1$ peak of the x-ray generator anode material. This step is very sensitive to the exact angle of the second crystal. While measuring and maximizing the intensity, it should be possible to find a second, smaller (local) maximum corresponding to $K\alpha_2$ at a slightly steeper reflection angle. Finally, the relative tip angle of the second crystal is adjusted, again by maximizing the intensity. The count rate is much less sensitive to this last degree of freedom, but optimizing the relative tip of the two crystals is important to maximize the monochromaticity of the exit beam along the cross-dispersion direction.

Alignment of a newly assembled monochromator typically takes several hours, while realigning a monochromator that is close to the correct alignment takes about an hour. We have verified the short-term alignment stability, for example, before and after positioning the monochromator housing for measurement campaigns. We have not explored the long term alignment stability rigorously, but we believe that realignment is rarely needed, if ever. The maintenance of alignment is verified by remeasuring the count rate of the monochromator for the same x-ray tube settings and checking that it is consistent with the value obtained during the previous alignment.

B. Mounting

As will be discussed in Sec. V B 1, the output beam of the monochromator in the dispersion direction is a collimated image of the x-ray source spot, which is smaller than many calorimeter pixels for typical commercial microfocus x-ray generators, while the beam diverges in the cross-dispersion direction so that the image on a calorimeter array is a thin line, and only a single column (or row) of such a detector array can be illuminated at one time. To align the monochromator beam to the detector array, and to allow illumination of all pixels, we mounted the entire apparatus (monochromator housing with the x-ray tube) on a motorized linear stage. The beam position can then be periodically cycled, allowing illumination of the full array over the course of an experiment.

IV. THEORETICAL PERFORMANCE

We used the web tool GID_SL^{18,19} to calculate single-reflection rocking curves $R_1(\Delta\theta)$ for relevant x-ray energies and crystal planes. Here, $\Delta\theta$ is the offset of the incident ray with respect to the nominal

TABLE I. Dimensions of crystals used in monochromators. l gives the length of the channel, and w gives the width.

Line	Crystal	Number of reflections	l (mm)	w (mm)	Bragg angle (deg)
Cr $K\alpha_1$	Si (220)	6 + 6	31.8	3.2	36.60 128
Cu $K\alpha_1$	Si (220)	4 + 4	31.8	3.2	23.65 103
Cu $K\alpha_1$	Si (400)	4 + 4	52.3	9.0	34.56 447

Bragg angle. Rocking curves were calculated for both σ and π polarization, which had to be considered separately; the results we derive for each polarization are averaged to obtain results for unpolarized light, where appropriate. We assumed symmetric Bragg diffraction from perfect crystals.

For each polarization, the rocking curves for $2N$ reflections (with N a positive whole number) from a single channel cut crystal were obtained by multiplying the single crystal rocking curves to the $2N$ power,

$$R_{2N}(\Delta\theta) = R_1(\Delta\theta)^{2N}. \quad (2)$$

The rocking curves are shown for both polarizations in Fig. 5 for one and four reflections from Cu $K\alpha_1$ Si (220).

The rocking curves give the reflectivity as a function of reflection angle with respect to the nominal Bragg angle at the specified energy; note that the actual peak reflectivity is slightly offset from the nominal Bragg angle as expected from theory. This rocking curve is really a one-dimensional version of what can more generally be described as a two-dimensional function of both energy and angle, $R(\Delta E, \Delta\theta)$. In the case of a rocking curve as presented in Fig. 5, the incident energy is fixed and no other energies are considered. To incorporate the effect of other nearby incident energies, we displace the rocking curve using a relation derived by dividing the Bragg equation by its derivative,

$$\frac{\tan \theta}{d\theta} = \frac{\lambda}{d\lambda} = -\frac{E}{dE}, \quad (3)$$

so that

$$d\theta = -\frac{dE}{E} \tan \theta. \quad (4)$$

We can thus write the reflectivity as a function of energy and angle in terms of the reflectivity as a function of angle at the nominal energy,

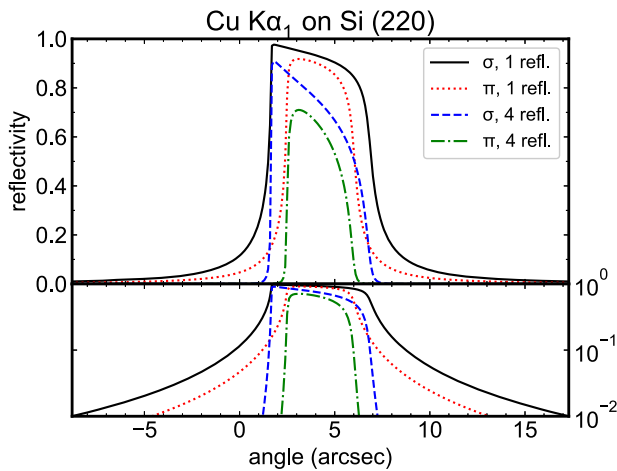


FIG. 5. Theoretical reflectivity as a function of angle for the Cu $K\alpha_1$ peak x-ray energy reflected from Si (220), in both σ and π polarizations. The top and bottom panels show the same rocking curves with a linear and logarithmic y axis, respectively. The four-reflection reflectivities are obtained by taking the single-reflection reflectivity to the fourth power. Note that the offset of the rocking curve center of mass on the x axis is with respect to the nominal Bragg angle for this x-ray energy and is expected from theory.

$$R(\Delta E, \Delta\theta) = R(\Delta\theta'), \quad (5)$$

where

$$\Delta\theta' = \Delta\theta - \frac{\Delta E}{E} \tan \theta \quad (6)$$

is the offset angle with respect to the nominal Bragg angle for a given offset energy ΔE when evaluated against the geometric reference of the nominal Bragg angle of $\Delta E = 0$. The reflectivity of the DHB 4N-reflection monochromator is obtained by multiplying the reflectivity for a single channel cut crystal with $2N$ reflections by itself but with the second reflectivity having the opposite signs in the displacement of angle as a function of energy displacement,

$$R_{\text{DHB,4N}}(\Delta E, \Delta\theta) = R_{2N}(\Delta E, \Delta\theta)R_{2N}(\Delta E, -\Delta\theta). \quad (7)$$

To get the transmitted line shape, we integrate the product of the reflectivity and the incident x-ray source flux over incident angles,

$$F_{E,\text{obs}}(\Delta E) = \int R_{\text{DHB,4N}}(\Delta E, \Delta\theta)F_{E,\theta}(\Delta E, \Delta\theta)d\Delta\theta, \quad (8)$$

where $F_{E,\text{obs}}$ is the observed flux per unit energy emerging from the monochromator and $F_{E,\theta}$ is the incident flux per unit energy per unit angle. If we assume that the source emits uniformly over the incident angle, as for an x-ray tube, we have

$$F_{E,\text{obs}} = F_{E,\theta}W_\theta(\Delta E), \quad (9)$$

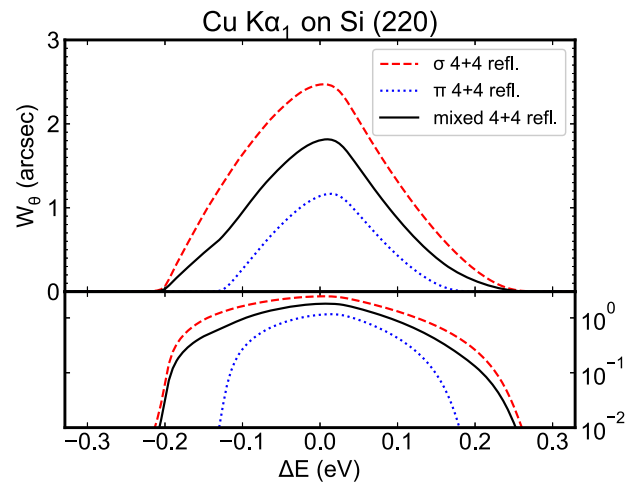


FIG. 6. Theoretical angular acceptance as a function of energy for both polarizations of Cu $K\alpha_1$ in a 4 + 4 reflection DHB monochromator using Si (220) crystals, as well as an average for unpolarized light (“mixed”). The top and bottom panels show the same acceptances with a linear and logarithmic y axis, respectively. These curves were evaluated at the nominal Bragg angle corresponding to the peak energy of Cu $K\alpha_1$, but because the center of mass of the rocking curve is slightly offset from the nominal Bragg angle (Fig. 5), the resulting angular acceptance peak was offset from the nominal energy by 0.35 eV. In practice, the true reflection angle is tuned to maximize the flux, and therefore, the x axis has been shifted to put the center of mass of the profile at zero. The magnitude of this shift can be used to calculate the expected dispersion angle for optimal alignment of the monochromator; however, in practice, it is so small that knowledge of this shift is not useful.

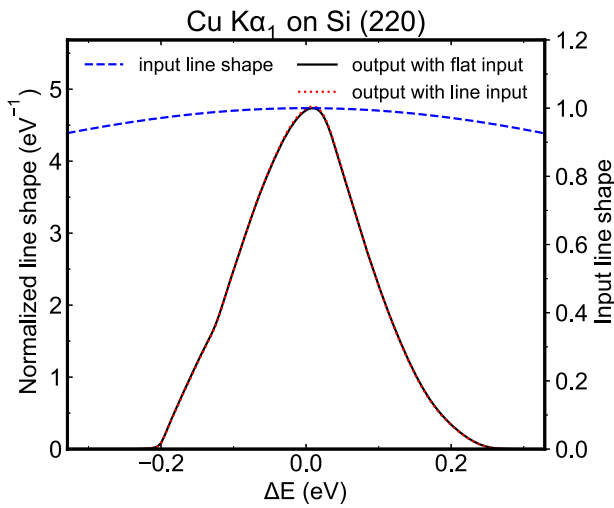


FIG. 7. Output line shape for a Cu $K\alpha_1$ 4 + 4 reflection DHB monochromator using Si (220) crystals. The black curve shows the line shape for a flat spectral input and has the same shape as the angular acceptance for mixed polarization. The blue dashed curve shows the input line shape from Cu $K\alpha_1$. The red dotted curve shows their product, which is the predicted true monochromator line shape. The similarity between the red and black curves shows that it is a good approximation to treat the input as flat since the monochromator resolution is much higher than the input line width.

where

$$W_\theta(\Delta E) = \int R_{\text{DHB},4N}(\Delta E, \Delta\theta) d\Delta\theta \quad (10)$$

is the *angular acceptance*. In the approximation that the incident spectral shape is flat (which is valid either for a continuum or for a monochromator tuned to the peak of a spectral line that is much broader than the rocking curve), the effective line shape of the monochromator is given by W_θ . Note that $W_\theta(E)$ is effectively a cross correlation of the rocking curve with its mirror image in θ , with the lag set by the dispersion from the Bragg equation.

In Fig. 6, we show calculations of W_θ for a 4 + 4 reflection Si (220) monochromator for the peak energy of Cu $K\alpha_1$. We average W_θ for σ and π polarizations to obtain W_θ for an unpolarized source. In Fig. 7, we show the effect of the Cu $K\alpha_1$ line shape on the output line shape of the monochromator. Since the input fluorescence line is much broader than the monochromator bandpass,

$W_\theta(\Delta E)$ is a good approximation for the true output line shape of the monochromator.

We can also compare the relative efficiency of a DHB monochromator by integrating over both energy and angle to obtain the *energy-angular acceptance* or simply acceptance,

$$W_{E,\theta} = \int R_{\text{DHB},4N}(\Delta E, \Delta\theta) d\Delta\theta d\Delta E. \quad (11)$$

Physically, the acceptance accounts for the throughput efficiency of a given energy resolution monochromator due to the rejection of energies outside the bandpass as well as incident angles outside the rocking curve for the accepted energies. The efficiency of a DHB monochromator thus scales with the square of the energy resolution (inversely with the square of the resolving power).

In Table II, we give the results of our calculations for the FWHM energy resolution and acceptance for the 4 + 4 Cu $K\alpha_1$ Si (220) monochromator, as well as for two other configurations we implemented a 6 + 6 reflection Cr $K\alpha_1$ Si (220) monochromator and a 4 + 4 reflection Cu $K\alpha_1$ Si (400) monochromator. We evaluated the formal FWHM energy resolution of the line shape as well as the effective FWHM energy resolution when the line shape is convolved with a 1 eV FWHM Gaussian ΔE_G ,

$$\Delta E_{\text{eff}} = \sqrt{\Delta E_{\text{conv}}^2 - \Delta E_G^2}. \quad (12)$$

Here, ΔE_{conv} is the FWHM of the convolution of the monochromator output with ΔE_G . The formal FWHM energy resolution is evaluated according to the formal definition of a full width at half maximum: the values of ΔE where the profile has half the strength of the peak value are found and then differenced to find the full width. The effective FWHM energy resolution is the value that has to be subtracted in quadrature from an uncorrected calorimeter energy resolution measured with a monochromator in order to obtain the true calorimeter energy resolution. The effective FWHM energy resolution is a figure of merit that more accurately describes the performance of the monochromator in a realistic scenario, such as a measurement of the FWHM energy resolution of calorimeter pixels. The effective FWHM is typically slightly smaller than the formal FWHM. The reason for this difference arises from the fact that convolution is more sensitive to the standard deviation, for which FWHM is not an ideal proxy for a non-Gaussian line shape. We also evaluated the effective FWHM with a range of Gaussian FWHMs and found that it does not change much as long as the Gaussian FWHM is significantly larger than the monochromator FWHM.

TABLE II. Predicted energy resolution and acceptance of DHB monochromators. The two values for FWHM energy resolution refer to the formal FWHM of the line shape (form.) and the effective FWHM manifested as an excess broadening when convolved with a 1 eV FWHM Gaussian (eff.).

Line	Crystal	No. of reflection	ΔE (FWHM, eV)		$W_{E,\theta}$ (eV arc sec)		
			Form.	Eff.	σ	π	Mixed
Cr $K\alpha_1$	Si (220)	6 + 6	0.143	0.125	0.208	7.5×10^{-5}	0.104
Cu $K\alpha_1$	Si (220)	4 + 4	0.201	0.197	0.584	0.182	0.383
Cu $K\alpha_1$	Si (400)	4 + 4	0.093	0.086	0.140	3.4×10^{-3}	0.072

V. MEASURED PERFORMANCE

A. Single crystal reflectivity

To verify the performance of individual crystals, we measured the reflectivity using a collimated monochromatic x-ray source. The source consisted of a water-cooled rotating anode x-ray generator with a Cu target illuminating a single channel cut crystal similar to those used in our DHB monochromators (hereafter referred to as the source crystal to avoid confusion with the sample crystals being characterized). The x-ray generator was operated with -30 kV cathode bias voltage and 100 mA cathode emission current. We estimate that the x-ray source spot size is ~ 300 μm . The source crystal was placed at a distance of 2.6 m from the source, with the channel oriented vertically so that the dispersion direction is horizontal. An exit slit with width 100 μm and height 1 mm was placed at a distance of 0.4 m from the source crystal and 3 m from the source. The source crystal was configured at an angle of 23.65° relative to the beam path to provide four reflections of Cu $K\alpha_1$ at 8047.78 eV. The beam divergence is set by the source spot size and exit slit width in comparison with their 3 m separation. Adding the source spot size and exit slit width in quadrature, we estimate an effective slit size of 316 μm and thus a divergence in the dispersion direction of 0.105 milliradian. Using Eq. (3), we find that this corresponds to an energy width of 1.93 eV. This is comparable to the FWHM of Cu $K\alpha_1$, so we expect both the natural line shape and the slit width to contribute to the source line shape.

The sample crystal was mounted on a goniometer allowing rotation of the crystal along the dispersion direction, and the reflected x rays were measured with a Xe-filled proportional counter. The distance from the source exit slit to the goniometer was 25 cm, and the distance from the goniometer to the proportional counter was 25 cm. The raw flux of Cu $K\alpha_1$ passing through the slit was measured by moving the sample out of the beam path and moving the proportional counter to the location of the unreflected beam. The experimental setup is the same as described in Fig. 2 of Mori *et al.*,²⁰ with the exception that the Si-PIN diode x-ray detector was replaced with the proportional counter, and the x-ray detector was not positioned at twice the goniometer angle, but rather at the correct position to intercept the reflected beam from the sample crystal, which is displaced parallel to the incident beam path.

We measured the reflected flux as a function of goniometer angle and displacement of the sample crystal along the channel direction. To understand how to interpret these measurements, we need to further develop the theory discussed in Sec. IV. The source and sample crystal act effectively as a DHB monochromator; however, we must now account for the fact that the source and sample crystals may be different. We still need to satisfy the Bragg equation for both crystals, so following Eq. (3), we have

$$\frac{-\Delta E}{E} = \frac{\Delta\lambda}{\lambda} = \frac{\Delta\theta_A}{\tan\theta_A} = \frac{\Delta\theta_B}{\tan\theta_B}. \quad (13)$$

Here, we use subscripts A and B to refer to the source and sample crystals, respectively. The change in the goniometer angle is the sum of the change in the two crystal reflection angles: $\Delta\theta_g = \Delta\theta_A + \Delta\theta_B$. Thus,

$$\Delta\theta_g = \frac{\Delta E}{E} (\tan\theta_A + \tan\theta_B). \quad (14)$$

If we denote the input flux from the x-ray generator as a function of energy and angle as $F_{E,\theta}$, and if we again assume that the input linewidth is much broader than the linewidth passed by the two crystals, then the flux observed as a function of goniometer angle is

$$F_{\text{obs}}(E)^{\text{AB}} = F_{E,\theta}(E) W_{E,\theta}^{\text{AB}}, \quad (15)$$

where the superscript AB indicates that the acceptance is evaluated for the reflectometer two-crystal system of the source and sample crystals and where the energy E is that allowed by the choice of goniometer angle. Here, $W_{E,\theta}^{\text{AB}}$ is calculated as in Eq. (11) but accounting for the heterogeneity of the source and sample crystal,

$$W_{E,\theta}^{\text{AB}} = \int R^A(\Delta E, \Delta\theta_A) R^B(\Delta E, \Delta\theta_B) d\Delta\theta_g d\Delta E. \quad (16)$$

Here, we use the abbreviated notation R^i , with $i = A, B$, to refer to R_{2N} for the source and sample crystals, respectively. Note that near the line center, the slit setting is irrelevant as the bandpass of the source-sample crystal system is much narrower than the slit width. However, as the goniometer is scanned, the slits do set the bounds of the scan angle where significant flux can be detected.

The flux collected by the x-ray detector with no sample crystal present is given by

$$F_{\text{obs}} = \int_{\Delta\theta_-}^{\Delta\theta_+} \int_0^\infty F_{E,\theta} R^A(\Delta E, \Delta\theta_A) dE d\Delta\theta_A. \quad (17)$$

The angular integration limits are set by the effective slit width. In analogy with Eq. (10), we define the *non-dispersive* angular acceptance

$$W_{2N,\theta} \equiv \int R_{2N}(E, \Delta\theta) d\Delta\theta. \quad (18)$$

Note that, in principle, $W_{2N,\theta}$ is still a function of energy, but since the rocking curve shape does not change much for a small change in energy, it is effectively a constant. Again, using A to refer to the source crystal, we then have

$$F_{\text{obs}}^A = W_{2N,\theta}^A \int_{\Delta E_-}^{\Delta E_+} F_{E,\theta} dE, \quad (19)$$

where we have converted the angular integration limits into equivalent energy integration limits. We have again assumed a uniform angular distribution for emission from the source.

Using Eqs. (15) and (19), the ratio of the flux observed with the sample crystal in and out is

$$\mathcal{R} = \frac{W_{E,\theta}^{\text{AB}}}{W_{2N,\theta}^A} \frac{F_{E,\theta}}{\int F_{E,\theta} dE}. \quad (20)$$

This is still not useful without knowledge of the line shape from the x-ray generator. Since we can scan the goniometer angle with the sample crystal installed, which is equivalent to scanning the energy passed by the system, we can thus effectively integrate the observed ratio over energy,

$$\mathcal{R}_E = \int \mathcal{R} dE = \frac{W_{E,\theta}^{\text{AB}}}{W_{2N,\theta}^A}. \quad (21)$$

Because the same integration limits imposed by the slits apply to both integrals over the source spectrum, the integrated spectrum

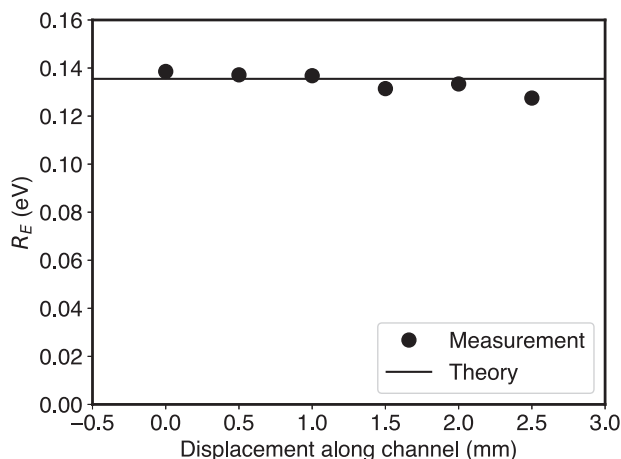


FIG. 8. Measured energy-integrated reflectivity ratio \mathcal{R}_E as a function of displacement along the channel width for a Si (220) crystal reflecting monochromatic Cu $K\alpha_1$ photons. The predicted R_E is shown as a solid horizontal line. This crystal was bonded to the substrate with epoxy.

terms cancel. Finally, note that for both $W_{E,\theta}^{AB}$ and $W_{2N,\theta}^A$, we must again evaluate them for both polarizations and average them to get the unpolarized values. We use the rocking curves discussed in Sec. IV to calculate the theoretical values of \mathcal{R}_E for comparison to our measurement data.

Results representative of our measurements are shown in Figs. 8 and 9. The first pair of crystals we used were Si (220) crystals with a ~ 3 mm wide channel. These were bonded to Al substrates with Henkel²¹ Loctite Hysol 9309 epoxy. We found that the reflectivity matches the prediction (Fig. 8). We repeated the procedure for

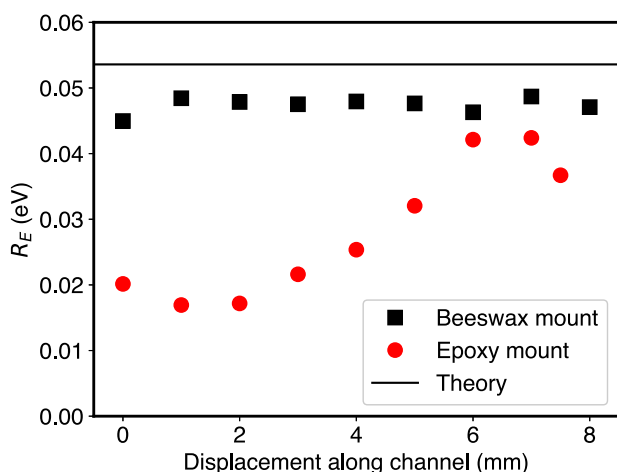


FIG. 9. Measured energy-integrated reflectivity ratio \mathcal{R}_E as a function of displacement along the channel width for a Si (400) crystal reflecting monochromatic Cu $K\alpha_1$ photons. This crystal was first bonded to the substrate with epoxy (red circles). Because the reflectivity was poor, the epoxy was removed and the crystal was rebonded with beeswax (black squares), leading to better reflectivity more closely matching the theoretical prediction.

a pair of Si (400) crystals with an ~ 9 mm wide channel. We found that the reflectivity was significantly less than the prediction. We suspected that strain due to the epoxy bond was responsible, so we removed the epoxy and rebonded the crystals to the substrates using beeswax. We remeasured the reflectivity and found that it was only slightly less than predicted (Fig. 9). It is not clear why the Si (400) crystals showed strain when bonded with epoxy, while the Si (220) crystals did not. This difference may be related to the relative width of the channels or to details of the epoxy bonding procedure for each crystal pair; however, beeswax gives good results and is much easier to unbond and rebond, so we recommend its use for the current application.

We note that the reduced reflectivity of the strained crystals would significantly reduce the output flux from the full monochromator. It is thus essential to achieve a sufficiently strain-free bond between the crystals and their substrates and prudent to validate the reflectivity of individual crystals before assembling the full monochromator.

B. Performance of full monochromator

We tested three different monochromator configurations and measured the count rate in each configuration incident on an Amptek²² XR-100CR Si-PIN detector with 13 mm^2 area positioned at the exit aperture of the monochromator box. We list the count rates in Table III. We also give count rates per angle in the cross-dispersion direction, using an estimated cross-dispersion capture angle of 16.0 mrad for our experiment configuration. The beam is collimated in the dispersion direction.

We note that the ratio of theoretical $W_{E,\theta}$ for the two Cu $K\alpha_1$ monochromators is 5.3, while the ratio of count rates observed is 12.0 (after correcting for the different emission current in the two experiments). It is not clear where this factor-of-two discrepancy in efficiency originates; the reflectivity reduction below theory shown in Fig. 9 is about 10% so that a 20% reduction in efficiency with respect to theory would be expected for the full monochromator.

1. Beam width

The optical design of the monochromator produces a collimated image of the anode spot in the dispersion direction and a radially diverging beam in the cross-dispersion direction. To characterize the beam width in the dispersion direction, we imaged the Cr $K\alpha_1$ Si (220) monochromator with a Princeton PI-SX:512 CCD. The x-ray tube was operated with an anode bias voltage $V = 10 \text{ kV}$ and emission current $I_{em} = 0.5 \text{ mA}$. The CCD has $24 \times 24 \mu\text{m}^2$ pixels, allowing measurement of the beam size with good resolution. It was operated at 1 MHz readout rate and with an exposure time of 1 s so that individual x-ray events could be counted, resulting in a frame time of 1.26 s and an out-of-time event fraction of 21%. The frame transfer direction was oriented parallel to the cross-dispersion axis of the monochromator so that the frame streak from out-of-time events runs in the same direction as the source image. Cosmic ray events were identified and removed based on both pulse height (PH) and charge distribution pattern. The image is shown in Fig. 10, and a histogram of the image in the dispersion direction is shown in Fig. 11. We fit the histogram with a Voigt function, obtaining a FWHM of 6.8 pixels or $162 \mu\text{m}$.

TABLE III. Measured count rates in test configuration. V is the x-ray tube anode bias voltage, and I_{em} is the cathode emission current.

Line	Configuration		Rate (c/s)	Rate (c/s/mrad)	V (kV)	I_{em} (mA)
	Crystal	Reflections				
Cr $K\alpha_1$	Si (220)	6 + 6	179	11.2	15	1
Cu $K\alpha_1$	Si (220)	4 + 4	921	57.6	20	0.7
Cu $K\alpha_1$	Si (400)	4 + 4	110	6.9	20	1

The manufacturer reported a measurement for the anode spot size of the Cr anode x-ray generator (serial number 59695) of $114.3 \pm 1.2 \mu\text{m}$ FWHM in the direction we aligned with the dispersion axis of the monochromator. The manufacturer's measurement was conducted at operating conditions of $V = 50 \text{ kV}$ and $I_{em} = 1 \text{ mA}$. The anode spot size is expected to change somewhat as a function of x-ray tube operating parameters, with spot size increasing with increased emission current and decreasing as anode voltage increases. The FWHM of the DHB rocking curve for Cr $K\alpha_1$ on Si (220) is 8 arc sec, which corresponds to $24 \mu\text{m}$ at the measurement distance of 60 cm between the crystal and the CCD. The correction due to angular divergence is thus almost negligible when added in quadrature with the anode spot size. We conclude that the measured beam width is consistent with the expectation of a collimated image of the anode spot, given the systematic uncertainty associated with the different tube operating conditions.

A beam width of $160 \mu\text{m}$ is smaller than the pixels of many calorimeter arrays, although arrays with smaller pixels have been fabricated. One possible application of a narrow beam monochromator is in tests for an undesired position dependent response. One example of this is shown in the work of Eckart *et al.*,²³ who used the narrow beam of the monochromator to probe position dependent response in a TES array developed for the Micro-X

sounding rocket mission.²⁴ The results of this test are summarized in Fig. 12.

2. Energy resolution

It is difficult to directly measure the energy resolution of a DHB monochromator of the design discussed in this article since the output x-ray flux is low by design, and furthermore, the resolving power of a spectrometer used to measure the energy resolution of the monochromator must be comparable to the monochromator itself.

Porst *et al.*⁵ measured the energy resolution of an MMC using an ^{55}Fe source, as well as our portable DHB monochromators (reproduced in Fig. 13). We can use this to estimate experimental constraints on the resolution of the Cr $K\alpha_1$ Si (220) monochromator since the resolution at Mn $K\alpha$ and Cr $K\alpha_1$ should be similar. The measured detector energy resolution from Mn $K\alpha$ after deconvolving the natural line shape¹⁰ is $1.71 \pm 0.12 \text{ eV}$, while the Gaussian FWHM when illuminated by the Cr $K\alpha_1$ monochromator is $1.82 \pm 0.01 \text{ eV}$, with no deconvolution. If we correct to this by subtracting the theoretical 0.125 eV FWHM of the monochromator in quadrature, we obtain $1.81 \pm 0.01 \text{ eV}$. In either case, the

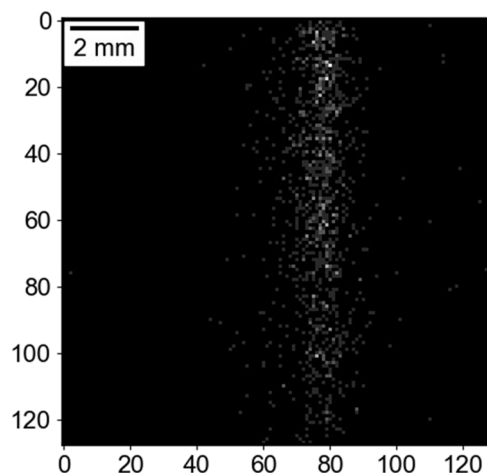


FIG. 10. CCD image of the monochromator beam. The native pixel size of the CCD is $24 \mu\text{m}$, and the pixels have been rebinned by a factor of four on both axes in this image.

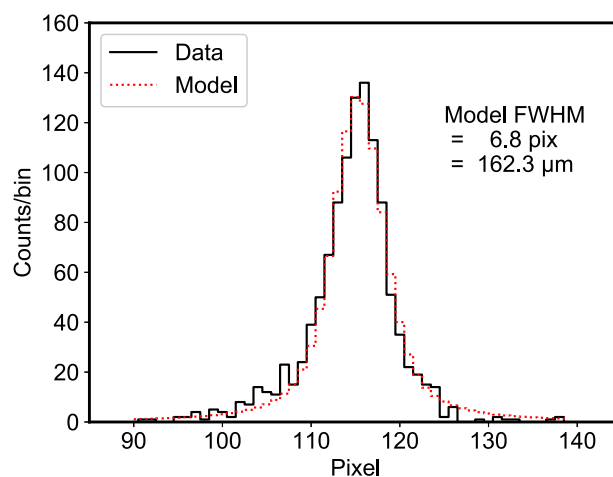


FIG. 11. Histogram in the dispersion direction of the CCD image of the monochromator beam. The pixel size is $24 \mu\text{m}$, and there has been no rebinning. The image was rotated before histogram accumulation to correct for a slight misalignment of the beam direction to the native pixel grid. The best fit Voigt function has a FWHM of 6.8 pixels, corresponding to $162.3 \mu\text{m}$.

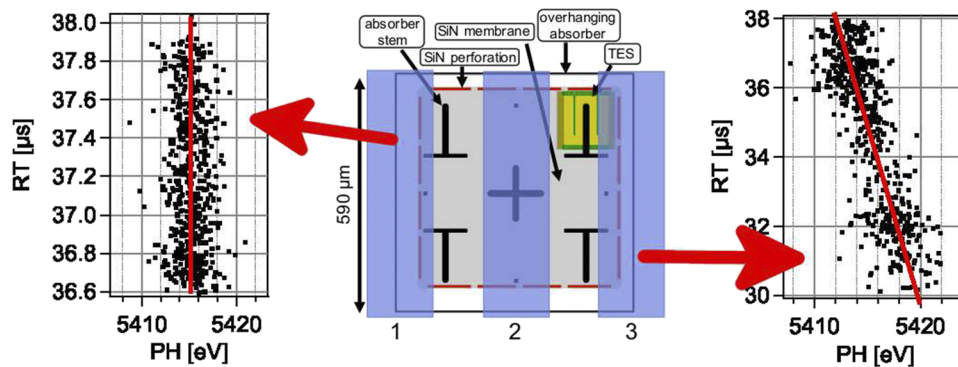


FIG. 12. Test of position dependent response in the TES detector.²³ The central diagram shows three beam positions where pulses were recorded. The left and right hand scatter plots show pulse rise time (RT) vs pulse height (PH) for events recorded at beam positions 1 and 3, respectively. The events in position 3 show a clear RT–PH correlation, which is due to a dependence of RT and PH on the distance from the TES sensor to where the photon was absorbed. To achieve the required energy resolution for this application, absorption of an x ray with a given energy should result in a similar pulse shape independent of where it arrives on the area of the absorber. The observed RT–PH correlation and its variation with the photon beam position on the absorber indicate problems with the thermalization of incident photons in the absorber. The results of this experiment were used to identify the origin of a measured distortion in the spectral response of these detectors that could then be addressed by changes to the pixel design and fabrication processes. Part of this figure is adapted with permission from Eckart *et al.*, “Development of TES microcalorimeter arrays for the micro-X sounding rocket experiment,” *IEEE Trans. Appl. Supercond.* **23**, 2101705 (2013). Copyright 2013 IEEE.

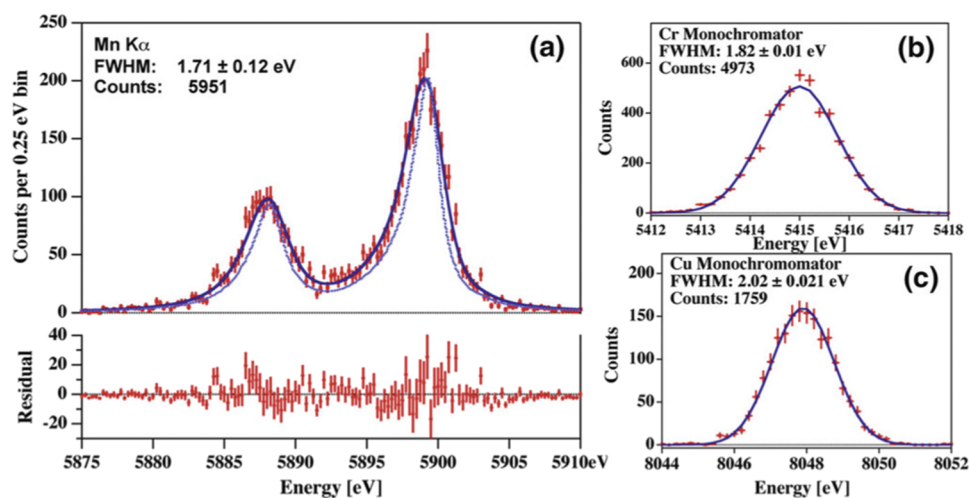


FIG. 13. Measurements of the energy resolution of a metallic magnetic x-ray calorimeter from Porst *et al.*⁵ Panel (a) shows a measurement of the Mn $K\alpha$ line from a radioactive ^{55}Fe source; the quoted energy resolution is the modeled intrinsic detector resolution after accounting for the known line shape of Mn $K\alpha$. Panels (b) and (c) show measurements using the 6 + 6 reflection Si (220) Cr $K\alpha$ monochromator and 4 + 4 reflection Si (400) Cu $K\alpha$ monochromator, respectively; the quoted energy resolution for these models does *not* account for the broadening from the monochromator. Reproduced with permission from Porst *et al.*, “Characterization and performance of magnetic calorimeters for applications in x-ray spectroscopy,” *J. Low Temp. Phys.* **176**, 617 (2014). Copyright 2013 Springer Nature.

resolution is consistent within 1σ with that measured using the ^{55}Fe source. Formally, the measured energy resolution of the monochromator obtained by differencing the ^{55}Fe and Cr $K\alpha_1$ measurements in quadrature and propagating the measurement uncertainty is $\Delta E_{\text{FWHM}} = 0.6 \pm 0.3$ eV. However, this is of questionable validity, since the propagated uncertainty on the energy resolution depends inversely on the best fit energy resolution itself, which is clearly very uncertain. If we instead neglect the relatively small uncertainty on the measurement with the monochromator and use the 1σ upper

and lower bounds of the Mn $K\alpha$ measurement to derive the bounds on the monochromator resolution, we obtain an upper limit of 0.9 eV FWHM. This is consistent with the theoretical resolution of 0.125 eV. We cannot use the MMC measurements presented in Fig. 13(c) using the Cu $K\alpha_1$ Si (400) monochromator in a similar way, both because the energy resolution at Cu $K\alpha_1$ is expected to be different from that at Mn $K\alpha$ and because the detector operating conditions had changed for this exposure due to a difference in the laboratory noise environment.

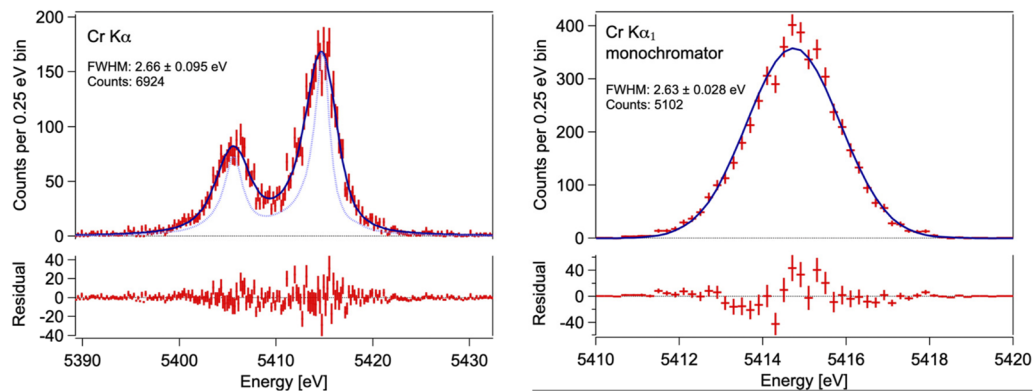


FIG. 14. X-ray spectra measured with a TES microcalorimeter. Illumination with Cr $K\alpha$ fluorescence (left) and the Cr $K\alpha_1$ monochromatic photons (right). The labeled Gaussian FWHM energy resolution for the Cr $K\alpha$ complex represents the intrinsic detector resolution after deconvolving the natural line shape,¹⁰ whereas the FWHM measured with the monochromator is reported with no deconvolution.

We also used a TES microcalorimeter to measure the energy resolution of the monochromator. The detector was from a prototype large pixel array²⁵ for the Athena X-ray Integral Field Unit (X-IFU).^{26,27} We obtained x-ray spectra from a Cr $K\alpha$ fluorescent source and the Cr $K\alpha_1$ monochromator, providing a comparison at the same incident energy. The measurements were performed consecutively on the same day with identical detector and cryogenic operating condition setpoints. The noise environment was stable. The results are presented in Fig. 14. Just as in our measurements using the MMC, we find that these two measurements are consistent at the 1σ level, and the inferred upper limit to the FWHM energy resolution of the monochromator is 0.7 eV, consistent with the theoretical resolution of 0.125 eV.

In both of the experiments described in this section, the reported error estimates are derived only from propagation of statistical uncertainties in fits to the data. The measurements are fully consistent with the theoretical prediction within the uncertainties. A more stringent measurement of the monochromator energy resolution is desirable, given ever-improving calorimeter energy resolutions, but will be challenging to achieve.

3. Polarization

The Cr $K\alpha_1$ 6 + 6 reflection Si (220) monochromator is theoretically predicted to be 99.96% σ -polarized, i.e., with the electric field in the dispersion direction. This makes it potentially of interest as a monochromatic calibration source for x-ray polarimeters.

We measured the polarization of this monochromator to be $99 \pm 1\%$ (statistical) using a micropattern Time Projection Chamber (TPC) polarimeter,²⁸ which obtains polarization information by analyzing the angle of the track created by the initial photoelectron. The polarimeter was calibrated using beamline X19-A at the National Synchrotron Light Source (NSLS), which has a calculated polarization of 98%. In Fig. 15, we show the number of events collected per phase angle bin in the TPC for both the synchrotron calibration and our measurement of the Cr $K\alpha_1$ 6 + 6 reflection Si (220) monochromator. The data are modeled with a sine plus a constant baseline. The ratio of the sine amplitude to the constant baseline is

known as the modulation factor and is proportional to the degree of polarization.

The polarization measurement of the monochromator may have some additional experimental errors that we estimate to be comparable to or smaller than the statistical uncertainty. The response of the polarimeter varies both with energy and, to a lesser extent, with position in the active area. We believe that the energy and position of the beams for the two measurements were sufficiently similar that these errors are small compared to the statistical uncertainties. There may also be an error in the assumed

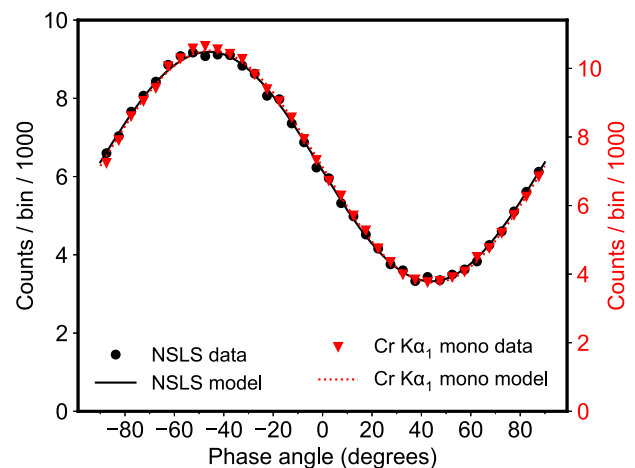


FIG. 15. Polarization measurements using the TPC polarimeter for the Cr $K\alpha_1$ Si (220) 6 + 6 reflection monochromator compared with calibration measurements from beamline X-19A at the NSLS synchrotron facility. The x axis denotes the reconstructed angle of the track of the initial photoelectron, while the y axis gives the number of counts accumulated per 5° angle bin. The left (black labels) and right (red labels) y axes are for the NSLS and Cr $K\alpha_1$ monochromator data and models, respectively. The solid lines show the best fit sine plus constant models. The modulation factor is derived by dividing the amplitude of the sine by the constant y-axis offset and is proportional to polarization.

synchrotron polarization of 98%, which is derived from a ray-tracing calculation²⁹ but has not been experimentally verified. Regardless of any possible systematic errors, our measurements are compatible with the theoretical prediction at the 1σ level.

VI. SUMMARY

We have reported on the design and construction of simple, portable x-ray monochromators intended for calibration of high-resolution x-ray calorimeters. The typical FWHM energy resolution of our designs is 0.1 eV–0.2 eV, and the output count rate may be as high as 500 counts/s/mm² at the monochromator exit aperture for typical operating conditions. The illumination pattern is a narrow strip with width corresponding to the x-ray tube spot size, ~100 μm –200 μm for the monochromators we have implemented.

Future work includes the development of further monochromators for different fixed energies, including Ti $K\alpha_1$, Fe $K\alpha_1$, and Au $L\alpha_1$ and β_1 ; implementation of computer-controlled, motorized positioning for the crystals; and use of bright, water-cooled x-ray generators with large anode spots, together with broad channels cut in the crystals. The latter is particularly important for the Athena X-IFU and other instruments featuring large arrays of calorimeter pixels so that much or all of the array may be simultaneously illuminated, allowing for more efficient use of limited calibration time. Many of these features have been implemented and are undergoing performance testing in our laboratory.

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DATA AVAILABILITY

The data that support the findings of this study are available from the corresponding author upon reasonable request.

REFERENCES

- ¹S. H. Moseley, J. C. Mather, and D. McCammon, "Thermal detectors as x-ray spectrometers," *J. Appl. Phys.* **56**, 1257–1262 (1984).
- ²M. E. Eckart, J. S. Adams, K. R. Boyce, G. V. Brown, M. P. Chiao, R. Fujimoto, D. Haas, J.-W. den Herder, A. Hoshino, Y. Ishisaki, C. A. Kilbourne, S. Kitamoto, M. A. Leutenegger, D. McCammon, K. Mitsuda, F. S. Porter, K. Sato, M. Sawada, H. Seta, G. A. Sneiderman, A. E. Szymkowiak, Y. Takei, M. S. Tashiro, M. Tsujimoto, C. P. de Vries, T. Watanabe, S. Yamada, and N. Y. Yamasaki, "Ground calibration of the ASTRO-H (Hitomi) soft x-ray spectrometer," *J. Astronomical Telesc. Instrum. Syst.* **4**, 021406 (2018).
- ³M. A. Leutenegger, M. Audard, K. R. Boyce, G. V. Brown, M. P. Chiao, M. E. Eckart, R. Fujimoto, A. Furuzawa, M. Guainazzi, D. Haas, T. Hayashi, J.-W. den Herder, R. Iizuka, M. Ishida, Y. Ishisaki, N. Kikuchi, C. A. Kilbourne, S. Koyama, S. Kurashima, Y. Maeda, M. Markevitch, D. McCammon, K. Mitsuda, H. Mori, N. Nakaniwa, T. Okajima, S. Paltani, R. Petre, F. S. Porter, K. Sato, T. Sato, M. Sawada, P. J. Serlemitsos, H. Seta, G. Sneiderman, Y. Soong, S. Sugita, A. E. Szymkowiak, Y. Takei, M. Tashiro, Y. Tawara, M. Tsujimoto, C. P. de Vries, T. Watanabe, S. Yamada, and N. Yamasaki, "In-flight verification of the calibration and performance of the ASTRO-H (Hitomi) soft x-ray spectrometer," *J. Astronomical Telesc. Instrum. Syst.* **4**, 021407 (2018).
- ⁴S. J. Smith, J. S. Adams, C. N. Bailey, S. R. Bandler, J. A. Chervenak, M. E. Eckart, F. M. Finkbeiner, R. L. Kelley, C. A. Kilbourne, F. S. Porter, and J. E. Sadleir, "Small pitch transition-edge sensors with broadband high spectral resolution for solar physics," *J. Low Temp. Phys.* **167**, 168–175 (2012).
- ⁵J.-P. Porst, S. R. Bandler, J. S. Adams, M. A. Balvin, S. E. Busch, M. E. Eckart, R. L. Kelley, C. A. Kilbourne, S. J. Lee, P. C. Nagler, F. S. Porter, J. E. Sadleir, G. M. Seidel, S. J. Smith, and T. R. Stevenson, "Characterization and performance of magnetic calorimeters for applications in x-ray spectroscopy," *J. Low Temp. Phys.* **176**, 617–623 (2014).
- ⁶S. J. Lee, J. S. Adams, S. R. Bandler, J. A. Chervenak, M. E. Eckart, F. M. Finkbeiner, R. L. Kelley, C. A. Kilbourne, F. S. Porter, J. E. Sadleir, S. J. Smith, and E. J. Wassell, "Fine pitch transition-edge sensor x-ray microcalorimeters with sub-eV energy resolution at 1.5 keV," *Appl. Phys. Lett.* **107**, 223503 (2015).
- ⁷W. B. Doriese, P. Abbamonte, B. K. Alpert, D. A. Bennett, E. V. Denison, Y. Fang, D. A. Fischer, C. P. Fitzgerald, J. W. Fowler, J. D. Gard, J. P. Hays-Wehle, G. C. Hilton, C. Jaye, J. L. McChesney, L. Miaja-Avila, K. M. Morgan, Y. I. Joe, G. C. O'Neil, C. D. Reintsema, F. Rodolakis, D. R. Schmidt, H. Tatsuno, J. Uhlig, L. R. Vale, J. N. Ullom, and D. S. Swetz, "A practical superconducting-microcalorimeter x-ray spectrometer for beamline and laboratory science," *Rev. Sci. Instrum.* **88**, 053108 (2017).
- ⁸S. Kempf, A. Fleischmann, L. Gastaldo, and C. Enss, "Physics and applications of metallic magnetic calorimeters," *J. Low Temp. Phys.* **193**, 365–379 (2018).
- ⁹M. Durkin, J. S. Adams, S. R. Bandler, J. A. Chervenak, S. Chaudhuri, C. S. Dawson, E. V. Denison, W. B. Doriese, S. M. Duff, F. M. Finkbeiner, C. T. FitzGerald, J. W. Fowler, J. D. Gard, G. C. Hilton, K. D. Irwin, Y. I. Joe, R. L. Kelley, C. A. Kilbourne, A. R. Miniussi, K. M. Morgan, G. C. O'Neil, C. G. Pappas, F. S. Porter, C. D. Reintsema, D. A. Rudman, K. Sakai, S. J. Smith, R. W. Stevens, D. S. Swetz, P. Szypryt, J. N. Ullom, L. R. Vale, N. A. Wakeham, J. C. Weber, and B. A. Young, "Demonstration of Athena X-IFU compatible 40-row time-division-multiplexed readout," *IEEE Trans. Appl. Supercond.* **29**, 2101005 (2019).
- ¹⁰G. Hölzer, M. Fritsch, M. Deutsch, J. Härtwig, and E. Förster, " $K\alpha_{1,2}$ and $K\beta_{1,3}$ x-ray emission lines of the 3d transition metals," *Phys. Rev. A* **56**, 4554–4568 (1997).
- ¹¹G. Hölzer, private communication to F. S. Porter (2004).
- ¹²M. E. Eckart, J. S. Adams, S. R. Bandler, S. Beaumont, J. A. Chervenak, A. M. Datesman, F. M. Finkbeiner, R. Hummatov, R. L. Kelley, C. A. Kilbourne, M. A. Leutenegger, A. R. Miniussi, S. J. Moseley, F. S. Porter, J. E. Sadleir, K. Sakai, S. J. Smith, N. A. Wakeham, and E. J. Wassell, "Extended line spread function of TES microcalorimeters with Au/Bi absorbers," *IEEE Trans. Appl. Supercond.* **29**, 2903420 (2019).
- ¹³J. W. M. DuMond, "Theory of the use of more than two successive x-ray crystal reflections to obtain increased resolving power," *Phys. Rev.* **52**, 872–883 (1937).
- ¹⁴J. H. Beaumont and M. Hart, "Multiple Bragg reflection monochromators for synchrotron x radiation," *J. Phys. E: Sci. Instrum.* **7**, 823–829 (1974).
- ¹⁵W. J. Bartels, "Characterization of thin layers on perfect crystals with a multi-purpose high resolution x-ray diffractometer," *J. Vac. Sci. Technol. B* **1**, 338–345 (1983).
- ¹⁶Oxford Instruments X-ray Technology, Scotts Valley, CA, USA.
- ¹⁷Crystal Scientific (UK) Ltd., Middle Barton, Whittingham, Alnwick, Northumberland, United Kingdom.
- ¹⁸S. A. Stepanov, E. A. Kondrashkina, R. Köhler, D. V. Novikov, G. Materlik, and S. M. Durbin, "Dynamical x-ray diffraction of multilayers and superlattices: Recursion matrix extension to grazing angles," *Phys. Rev. B* **57**, 4829–4841 (1998).
- ¹⁹See http://x-server.gmca.aps.anl.gov/GID_sl.html for access to GID_sl.
- ²⁰H. Mori, T. Okajima, W. W. Zhang, K.-W. Chan, R. Koenecke, J. R. Mazarella, A. Numata, L. G. Olsen, R. E. Riveros, and M. Yukita, "Reflective coatings for the future x-ray mirror substrates," in *Space Telescopes and Instrumentation 2018: Ultraviolet to Gamma Ray*, edited by J.-W. A. den Herder, S. Nikzad, and K. Nakazawa (SPIE-International Society for Optics and Photonics, 2018), Vol. 10699, pp. 1012–1019.

- ²¹Henkel Corporation Aerospace, 2850 Willow Pass Road, Bay Point, CA, USA).
- ²²Amptek Inc., 14 DeAngelo Drive, Bedford, MA, USA.
- ²³M. E. Eckart, J. S. Adams, S. R. Bandler, S. E. Busch, J. A. Chervenak, A. J. Ewin, F. M. Finkbeiner, R. L. Kelley, C. A. Kilbourne, F. S. Porter, J.-P. Porst, J. E. Sadleir, S. J. Smith, E. J. Wassell, and E. Figueroa-Feliciano, "Development of TES microcalorimeter arrays for the micro-X sounding rocket experiment," *IEEE Trans. Appl. Supercond.* **23**, 2101705 (2013).
- ²⁴J. S. Adams, R. Baker, S. R. Bandler, N. Bastidon, M. E. Danowski, W. B. Doriese, M. E. Eckart, E. Figueroa-Feliciano, D. C. Goldfinger, S. N. T. Heine, G. C. Hilton, A. J. F. Hubbard, R. L. Kelley, C. A. Kilbourne, R. E. Manzagol-Harwood, D. McCammon, T. Okajima, F. S. Porter, C. D. Reintsema, P. Serlemittos, S. J. Smith, J. N. Ullom, and P. Wikus, "First operation of TES microcalorimeters in space with the micro-X sounding rocket," *J. Low Temp. Phys.* **199**, 1062 (2020).
- ²⁵S. J. Smith, J. S. Adams, S. R. Bandler, G. L. Betancourt-Martinez, J. A. Chervenak, M. P. Chiao, M. E. Eckart, F. M. Finkbeiner, R. L. Kelley, C. A. Kilbourne, A. R. Miniussi, F. S. Porter, J. E. Sadleir, K. Sakai, N. A. Wakeham, E. J. Wassell, W. Yoon, D. A. Bennett, W. B. Doriese, J. W. Fowler, G. C. Hilton, K. M. Morgan, C. G. Pappas, C. N. Reintsema, D. S. Swetz, J. N. Ullom, K. D. Irwin, H. Akamatsu, L. Gottardi, R. den Hartog, B. D. Jackson, J. van der Kuur, D. Barret, and P. Peille, "Transition-edge sensor pixel parameter design of the microcalorimeter array for the x-ray integral field unit on Athena," in *Space Telescopes and Instrumentation 2016: Ultraviolet to Gamma Ray*, edited by J.-W. A. den Herder, T. Takahashi, and M. Bautz (SPIE-International Society for Optics and Photonics, 2016), Vol. 9905, pp. 772–790.
- ²⁶D. Barret, T. Lam Trong, J.-W. den Herder, L. Piro, M. Cappi, J. Houvelin, R. Kelley, J. M. Mas-Hesse, K. Mitsuda, S. Paltani, G. Rauw, A. Rozanska, J. Wilms, S. Bandler, M. Barbera, X. Barcons, E. Bozzo, M. T. Ceballos, I. Charles, E. Costantini, A. Decourchelle, R. den Hartog, L. Duband, J.-M. Duval, F. Fiore, F. Gatti, A. Goldwurm, B. Jackson, P. Jonker, C. Kilbourne, C. Macculi, M. Mendez, S. Molendi, P. Orleanski, F. Pajot, E. Pointecouteau, F. Porter, G. W. Pratt, D. Prêle, L. Ravera, K. Sato, J. Schaye, K. Shinozaki, T. Thibert, L. Valenziano, V. Valette, J. Vink, N. Webb, M. Wise, N. Yamasaki, F. Douchin, J.-M. Mesnager, B. Pontet, A. Pradines, G. Branduardi-Raymont, E. Bulbul, M. Dadina, S. Etori, A. Finoguenov, Y. Fukazawa, A. Janiuk, J. Kaastra, P. Mazzotta, J. Miller, G. Miniutti, Y. Naze, F. Nicastro, S. Scioritino, A. Simionescu, J. M. Torrejon, B. Frezouls, H. Geoffroy, P. Peille, C. Aicardi, J. André, C. Daniel, A. Clénet, C. Etcheverry, E. Gloaguen, G. Hervet, A. Jolly, A. Ledot, I. Paillet, R. Schmitter, B. Vella, J.-C. Damery, K. Boyce, M. Dipirro, S. Lotti, D. Schwander, S. Smith, B.-J. Van Leeuwen, H. van Weers, N. Clerc, B. Cobo, T. Dauser, C. Kirsch, E. Cucchetti, M. Eckart, P. Ferrando, and L. Natalucci, "The Athena x-ray integral field unit (X-IFU)," in *Proc. SPIE* **10699**, 106991G (2018).
- ²⁷F. Pajot, D. Barret, J. W. Den herder, M. Barbera, N. Webb, K. Mitsuda, S. Paltani, C. Kilbourne, M. Cappi, F. Douchin, J. Wilms, L. Piro, L. Valenziano, Gregor Rauw, R. Kelley, T. Lam-Trong, J. Huovelin, J. M. Mas-Hesse, A. Rozanska, H. Geoffroy, R. Den Hartog, C. Macculi, J.-M. Mesnager, and P. Peille, "High resolution imaging x-ray spectroscopy: The X-IFU instrument on Athena," in *42nd COSPAR Scientific Assembly (Cospar, 2018)*, Vol. 42, p. E1.17-35-18.
- ²⁸W. B. Iwakiri, J. K. Black, R. Cole, T. Enoto, A. Hayato, J. E. Hill, K. Jahoda, P. Kaaret, T. Kitaguchi, M. Kubota, H. Marlowe, R. McCurdy, Y. Takeuchi, and T. Tamagawa, "Performance of the PRAXyS x-ray polarimeter," *Nucl. Instrum. Methods Phys. Res., Sect. A* **838**, 89–95 (2016).
- ²⁹C. Y. Yang, J. E. Penner-Hahn, and P. M. Stefan, "NSLS X-19A beamline performance for x-ray absorption measurements," *Nucl. Instrum. Methods Phys. Res., Sect. A* **291**, 157–161 (1990).