Saharan dust effects on North Atlantic sea surface skin temperatures

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Key Points:

1. Using independent shipboard measurements and MERRA-2 reanalysis in the tropical Atlantic Ocean.

2. The dust outbreaks can decrease the surface shortwave radiation up to 190 W/m² and increase the surface longwave radiation by up to $14W/m^2$.

3. Skin sea surface temperature response to the abnormal surface radiative fluxes can range from a net cooling to a tiny warming: -0.24K to 0.06K.

1 Abstract

2 Saharan dust outbreaks frequently propagate westward over the Atlantic Ocean; accurate 3 quantification of the dust aerosol scattering and absorption effect on the surface radiative fluxes 4 (SRF) is fundamental to understanding critical climate feedbacks. By exploiting large sets of 5 measurements from many ship campaigns in conjunction with reanalysis products, this study 6 characterizes the sensitivity of the SRF and skin Sea-Surface Temperature (SST_{skin}) to the 7 Saharan dust aerosols using models of the atmospheric radiative transfer and thermal skin effect. Saharan dust outbreaks can decrease the surface shortwave radiation up to 190 W/m², and an 8 9 analysis of the corresponding SST_{skin} changes using a thermal skin model suggests dust-induced 10 cooling effects as large as -0.24 K during daytime and a warming effect of up of 0.06 K during 11 daytime and nighttime respectively. Greater physical insight into the radiative transfer through 12 an aerosol-burdened atmosphere will substantially improve the predictive capabilities of weather 13 and climate studies on a regional basis.

1. Introduction and Background

15	Skin sea surface temperatures (SST_{skin}) provide a key indicator of the air-sea exchanges at				
16	the ocean surface as well as the acquisition, storage, transport and release of heat. The majority				
17	of the incoming solar radiation is absorbed in the upper ocean; most of the resulting heat gain is				
18	released to the lower atmosphere, helping to drive weather and climate. The Surface Radiative				
19	Fluxes (SRF) are influenced by the atmospheric state [Kato et al., 2020; Wild et al., 2013;				
20	Yamada and Hayasaka, 2016]. The Saharan Air Layer (SAL) is a particular atmospheric				
21	anomaly that frequently occurs at large scales in the tropical North Atlantic Ocean and has				
22	consequences on the surface radiation budget [Evan et al., 2009; Yu et al., 2006], on the				
23	development of severe storms [Zhang et al., 2007], and on the accuracy of satellite-derived				
24	SST _{skin} [Luo et al., 2019; Merchant et al., 2006]. The SAL is a typically warm and dry air layer				
25	that can reside up to 5km in altitude, and is often be accompanied by dust aerosols [Carlson and				
26	Prospero, 1972; Dunion and Velden, 2004; Nalli et al., 2011; Wu, 2007].				
27	The dust aerosol particles are found with diameters exceeding 10 μ m even after long-				
28	range transport (e.g., Kramer et al, 2020), allowing them to absorb, scatter and emit infrared				
29	radiation [Kok et al., 2017]. Indeed, many studies have made efforts to quantify dust direct				
30	radiative effects (e.g., Doherty and Evan [2014]; Hansell et al. [2010]; Song et al. [2018];				
31	Thorsen et al. [2020]; Zuidema et al. [2008]). The ocean gains most of its heat from the				
32	absorption in the upper-layer of shortwave radiation during the daytime and loses the heat				
33	through the ocean surface nearly all the time through infrared emission and turbulent fluxes.				
34	Thus, the anomalies in SRF due to aerosol radiative effects can impact the sea surface				
35	temperatures (SSTs) and upper ocean heat content, especially the diurnal warm layer, also				
36	referred to as the diurnal thermocline [Gentemann et al., 2009]. Numerous studies have				

37 attempted to analyze the impact of dust aerosols on SST: Foltz and McPhaden [2008] used 38 combinations of satellite-derived aerosol optical depth (AOD) and in-situ measurements obtained 39 from the PIRATA mooring data to investigate the relations between SST anomalies and AOD 40 anomalies, they found that the dust outflows were always associated with a reduction in solar 41 radiation, and about 35% of the SST variability was related to dust outbreaks; other SST cool 42 anomalies were due to wind stress. Evan et al. [2009] used 26 years of satellite-derived SST and 43 AOD retrievals to model the response of the temperature of the mixed layer to changes in dust 44 AOD, which were shown to decrease over the period studied, with aerosols accounting for 69% 45 of the upward trend in the mixed-layer temperature. Martínez Avellaneda et al. [2010] analyzed 46 the MODerate resolution Imaging Spectroradiometer (MODIS) derived AOD and TRMM 47 Microwave Imager (TMI) derived SST with a physical model to estimate the Saharan dust 48 impact on North Atlantic SST, the dust associated SST shortwave cooling was found to range 49 from 0.2 K to 0.4 K. About 30% of the climatological SST variance could be explained by the 50 dust cooling scheme at low wind speed conditions. Lau and Kim [2007] estimated the solar 51 radiation attenuation by dust aerosols which could explain 30% - 40% of the SST cooling 52 patterns, and that the cooling rate depends on the mixed layer depth. Evan et al. [2011] presented 53 the effects of the dust aerosols on air-sea interface variability with satellite datasets; they found 54 the dust-forced SST anomalies were related to the Atlantic Meridional Mode [Vimont and 55 Kossin, 2007].

However, all of the above studies were performed using satellite datasets and models.
Compared to the Top of Atmosphere (TOA) radiation being directly monitored by satellites,
estimates of the perturbations of the surface radiative budget caused by dust aerosols are difficult
to derive from satellite data due to the additional complexities in the retrieval process, and

60 imperfect understanding of the atmospheric radiative effects [Kato et al., 2013]. The 61 determination of aerosol radiative forcing at the sea surface can also be undermined by the 62 limited availability of direct measurements. Furthermore, estimating the correct dust effect on 63 SST in different geographical locations requires a sufficiently large in situ dataset and accurate 64 models of the thermal skin effect and diurnal heating. None of the aforementioned studies are 65 based on a long-term analysis of in situ data. The lack of in situ measurements at the sea surface 66 has hampered studies of the effects of dust aerosol SRF on SST over the tropical North Atlantic 67 Ocean.

68 This study uses measurements from a suite of AERosol and Ocean Science Expedition 69 (AEROSE) campaigns that have provided appropriate atmospheric and oceanic datasets [Nalli et 70 al., 2011], augmented with reanalysis dust aerosol profiles throughout the tropical North Atlantic 71 Ocean, with the objective of a better understanding of the aerosol radiative forcing, both infrared 72 and shortwave, on the SST_{skin} . Since the SRF variations are modulated by radiative feedback 73 processes, a critical tool is provided by radiative transfer models, which allow the isolation of different internal and external factors and an examination of their relationships with each other. 74 75 In this study, the shortwave and longwave radiative fluxes are calculated with the Rapid 76 Radiative Transfer Model for General Circulation Models Applications (RRTMG; Iacono et al. 77 [2008]). The downwelling radiative effects of dust also have an impact on the SST_{skin} . A widely-78 used skin layer and diurnal model that includes convection and insolation absorption effects was 79 developed by Fairall et al. [1996] and further refined by Zeng and Beljaars [2005], Takaya et al. 80 [2010], and Gentemann et al. [2009]. In our study, the dust radiative effects on SST_{skin} are 81 simulated by models of the thermal skin used in the NASA Global Modeling and Assimilation 82 Office (GMAO) Weather Analysis and Prediction System

83	(https://gmao.gsfc.nasa.gov/weather_prediction/), MERRA-2 (Modern-Era Retrospective
84	analysis for Research and Applications, version 2; Randles et al. [2017]); MERRA-2 reanalysis
85	used daily ¹ / ₄ deg Reynolds SST [Reynolds and Smith, 1994] as foundation SST before Apr 2006,
86	thereafter the daily OSTIA product [Donlon et al., 2012] was used, see [Akella et al., 2017] for
87	details and the aerosol information is also from the NASA GMAO MERRA-2 (Modern-Era
88	Retrospective analysis for Research and Applications, version 2; Randles et al. [2017])
89	atmospheric reanalysis, see Section 2.2 for further details. The extensive datasets from AEROSE
90	campaigns can be used as valuable inputs for the RRTMG model, as well as examining the dust
91	radiative effects on SST _{skin} . Greater physical insight into the downwelling radiative effect of
92	Saharan dust aerosols on SST_{skin} will lead to substantially improving the predictive capabilities of
93	weather and climate models in the Atlantic Ocean area.

2. Data and Study Fields

94 **2.1 AEROSE in situ data**

95 In situ measured data from research ships collected during a series of AEROSE 96 campaigns [Morris et al., 2006; Nalli et al., 2011] onboard the NOAA Ship Ronald H. Brown 97 and the R/V Alliance and remotely derived datasets are used to assess the Saharan dust effects on 98 SST_{skin}. AEROSE was a sequence of Atlantic field campaigns from 2004 to the present, aiming to 99 take accurate oceanic and atmospheric measurements of the tropical Atlantic Ocean under 100 Saharan dust outbreaks [*Nalli et al.*, 2011]; Figure 1 shows the AEROSE ship tracks of each 101 year, where the colors indicate day of the year. Table 1 summarizes the cruise starting and 102 ending dates as well as the number of radiosondes deployed in each of the AEROSE cruises that 103 were used in this study. A total of 751 radiosonde profiles over a span of 231 days were used.

The measurements made during these campaigns provided data that were required as inputs forradiative transfer models and models of the thermal skin and diurnal heating.

106 One of the key instruments is the Marine-Atmospheric Emitted Radiance Interferometer 107 (M-AERI), which is a Fourier-transform infrared spectro-radiometer which measures spectra in 108 the wavenumber range of 500-3000 cm⁻¹ (3.3 -20 μ m) [*Minnett et al.*, 2001]. An M-AERI was 109 mounted on the ships for each AEROSE cruise. Highly accurate SST_{skin} can be retrieved from M-110 AERI measurements. An error budget of the SST_{skin} derived from the M-AERI measurements 111 gives a root mean square accuracy of about 40 mK. The M-AERI infrared spectra can also be 112 used to retrieve the lower troposphere temperature and humidity profiles from the measurements 113 of CO₂ emission spectra [Szczodrak et al., 2007]. M-AERI SST_{skin} retrievals have been widely 114 used to validate satellite-derived SST_{skin} [Kearns et al., 2000; Kilpatrick et al., 2015; Luo et al., 115 2019; Luo et al., 2020a; Minnett et al., 2020a], and reanalysis model output SST_{skin}, such as those 116 from MERRA-2 [Luo et al., 2020b] and ERA-5 [Luo and Minnett, 2020]. 117 During the AEROSE cruises, two to four radiosondes were deployed each day to measure

the vertical air temperature and water vapor profiles. Figure 2, rows 1-2, show the relative humidity and rows 3-4 show the air temperature along each of the AEROSE tracks. These intensive observations provide us with the opportunity to quantify dust aerosol radiative forcing on the SRF and SST_{skin} ; the dense network of observations will benefit the radiative transfer model simulations.

123 **2.2 MERRA-2 data**

124 This study uses data from the NASA MERRA-2 [*Gelaro et al.*, 2017], which contains 125 geolocated dust aerosol mixing ratio at 72 standard pressure levels with one or three hours 126 temporal resolution which is extraordinarily useful for this study; aerosols in MERRA-2 are

127 constrained via data assimilation, see *Randles et al.* [2017] and *Buchard et al.* [2017] for further

details.. Figure 2 (rows 5-6) shows the MERRA-2 dust mass mixing ratio at the radiosonde

129 deployment location along with each AEROSE track; clearly, as intended, AEROSE campaigns

130 have encountered significant Saharan dust outflows.

131 **2.3 RRTMG**

132RRTMG [*Iacono et al.*, 2008] was developed by Atmospheric and Environmental133Research, Inc. and uses a correlated-k method to improve the computational efficiency of134radiative transfer calculations by dividing the longwave spectrum into 16 and shortwave135spectrum into 14 continuous bands. The atmospheric relative humidity and air temperatures are136from the radiosondes, the *SST_{skin}* and near surface air temperatures are those from M-AERI, the137dust aerosol inputs are taken from the MERRA-2. The surface shortwave and longwave radiation138are the outputs from RRTMG.

MERRA-2 provides the dust mixing ratio of each layer but RRTMG requires AOD as
input; the 550 nm AOD for each layer in RRTMG can be expressed as:

141
$$\tau_{550nm} = \sum_{i=1}^{i=5} x_i \times b_{ext,i}(RH, 550nm) \times \delta z \tag{1}$$

MERRA-2 includes five different sizes of dust particles, where x_i is the dust aerosol mass mixing ratio of each size interval i with the effective dust radii of 0.64, 1.34, 2.32, 4.2, and 7.75 µm; therefore dust is size-resolved into 5 size bins according to *Tegen et al.* [2004]. The dust extinction coefficient is $b_{ext,i}$ and δz is the atmospheric layer thickness, RH is the relative humidity of the dust layer. The dust extinction coefficient, single-scattering albedo (SSA), and asymmetry factor vary as a function of the dust size and composition, and these parameters were taken from the lookup table of *Randles et al.* [2017]. The dust aerosols were assumed to be well mixed within each layer of the modelled atmosphere; the overall extinction coefficients, SSAs,
and asymmetry factors were calculated following equations A1-A3 of *Thorsen et al.* [2020].

151 RRTMG longwave calculations require specifying the aerosol scattering properties at 152 each of the 16 spectral bands; the Ångström exponent formula [Ångström, 1929] given in 153 following Equation 2 was used to explicitly specify the spectral AOD, τ_{λ} , for a given wavelength 154 λ in the radiative transfer calculations:

155
$$\frac{\tau_{\lambda}}{\tau_{550nm}} = \left(\frac{\lambda}{\lambda_{550nm}}\right)^{-\alpha} \quad (2)$$

156 where α is the aerosol Ångström exponent determined from the MERRA-2 Dust Angstrom 157 parameter. MERRA-2 provides three-hourly averaged dust properties, so the dust profiles are 158 linearly interpolated to the radiosonde deployment times and locations. Other inputs, such as the 159 sea surface emissivity and surface albedo are also taken from MERRA-2.

160 **2.4** *SST*_{skin} model

161 The Group for High Resolution Sea Surface Temperature (GHRSST) has defined 162 different kinds of SSTs according to the depth and produced a suite of SST products [Donlon et 163 al., 2007]. Sea-Bird Thermosalinographs (TSG) were mounted near the ships' seawater intakes 164 to measure the sea temperature that can approximate the foundation SST (SST_{fnd}) where the 165 temperature is presumed to be free of diurnal variability.

166 There are many diurnal heating and cooling models with varying complexity and

- used in ERA-5 [Akella et al., 2017; ECMWF, 2016], and some others use ship measurements
- 169 [Gentemann et al., 2009; Minnett et al., 2011; Zhang et al., 2020]. The SST_{fnd} for this study is
- 170 taken from the ship-based thermosalinographs. The SST_{skin} variations can be expressed by the

¹⁶⁷ dependences on forcing parameters. Some models are driven by surface fluxes, such as those

171 cool skin layer and the warm layer schemes [Akella et al., 2017; ECMWF, 2016; Takaya et al.,

172 2010; Zeng and Beljaars, 2005].

173
$$SST_{skin} = SST_{fnd} - \Delta SST_d + \Delta SST_W(z)$$
(3)

where SST_{fnd} , ΔSST_d and ΔSST_W are TSG SST_{fnd} , the vertical temperature difference across the cool-skin layer, and the temperature difference between the sub-skin and foundation temperatures, i.e., the warm-skin layer. The ΔSST_d , can be expressed as:

$$\Delta SST_d = \frac{\delta}{\rho_w c_w k_w} (Q_c) \tag{4}$$

177 where ρ_w , k_w , c_w are the water density, thermal conductivity and volumetric heat capacity 178 respectively, δ is the thermal skin layer thickness. Q_c is the surface net heat flux in the cool-skin 179 layer, which is positive downward:

$$Q_c = H_s + H_L - LW_{net} - f_s SW_{net}$$
⁽⁵⁾

180 where H_s , H_L , LW_{net} and SW_{net} are the sensible and latent heat flux, net longwave and solar 181 radiation at the surface, but only a fraction, f_s , of the surface insolation is absorbed by the near 182 surface ocean, given by *Fairall et al.* [1996] and *Fairall et al.* [2003], f_s depends on the 183 thickness of the skin layer. The heat fluxes are positive downward. The sensible and latent heat 184 flux can be calculated by the updated version 3.6 Coupled Ocean–Atmosphere Response 185 Experiment (COARE) algorithm [*Fairall et al.*, 2003]. The longwave and shortwave heat fluxes 186 are from RRTMG output with and without aerosol.

187 The diurnal warming can be expressed as:

188
$$\frac{\partial \left(SST_{-\delta} - SST_{fnd}\right)}{\partial t} = \frac{Q_w + R_s - R(-d)}{d\rho_w c_w \nu/(\nu+1)} - \frac{(\nu+1)ku_{*w}f(La)}{d\phi_t (d/L)} \left(SST_{-\delta} - SST_{fnd}\right)$$
(6)

189 where $SST_{-\delta}$ is the temperature at the base of the cool skin layer, *d* is the diurnal warm layer 190 depth, ν is the profile shape, having a value of 0.2 [*Gentemann and Akella*, 2018], u_{*w} is the 191 water friction velocity, κ is the von Karman constant of 0.4, $\phi_t(d/L)$ is the stability function. *La* 192 is the Langmuir number, $La = \sqrt{u_{*w}/u_s}$, and $f(La) = La^{-2/3}$. The Stokes velocity, u_s , is 193 taken as 0.01m/s, as by *Akella et al.* [2017]. R(-d) is the intensity of solar radiation at depth -194 d.

195 The stability parameter, d/L, includes the Obukhov length, L, given by

$$L = \frac{\rho_w c_w u_{*w}^3}{\kappa g \alpha_w Q_w} \quad (7)$$

197 The stability function $\phi_t(d/L)$ is:

198
$$\phi_t(d/L) = \begin{cases} 1 + \frac{5d/L + 4(d/L)^2}{1 + 3d/L + 0.25(d/L)^2} & \text{when } d/L \ge 0\\ (1 - 16d/L)^{-0.5} & \text{when } d/L < 0 \end{cases}$$
(8)

199 The net heat flux, Q_w , available to heat the warm layer can be expressed as:

$$200 Q_w = SW_{net}^w + LW_{net} - H_s - H_L (9)$$

201 Note the SW_{net}^w is warm layer absorbed net shortwave radiation, $SW_{net}^w = SW_{net} - SW_{PEN}$, where SW_{PEN} is the penetrating short-wave radiation:

203
$$SW_{PEN} = SW_{net} \sum_{i=1}^{N=3} a_i \exp(-db_i)$$
 (10)

204 the coefficients a_i and b_i can be obtained from Zeng and Beljaars [2005].

Integrating equation (6) in time is part of the warm-skin layer scheme, the results of the warm layer and the cool layer schemes are added together (Equation 3) to derive the SST_{skin} . The net solar radiation at the surface SW_{net} and the net longwave radiation LW_{net} at the surface are taken from the RRTMG simulation outputs with and without aerosols (Section 3.1); so the SST_{skin} response to the dust can be expressed as the difference between the SST_{skin} schemes with and without aerosols.

3. Results and Discussion

3.1Net radiative effects of dust at the sea surface from RRTMG

The sea surface heat budget is related to the shortwave and longwave net radiative heat fluxes and the turbulent latent and sensible fluxes at the surface. To quantify dust aerosol effects on the sea surface heat budget and SST_{skin} , the downward shortwave and longwave radiation were estimated using radiative transfer calculations.

216 The downward irradiance is sensitive to the solar zenith angle, air temperature, water 217 vapor, cloud and aerosol properties [Kato et al., 2018]. Because directly measured SRF is 218 available in only a limited area, the current global SRF data [Kato et al., 2018; Loeb et al., 2020; 219 Loeb et al., 2019] are calculated using radiative transfer models with input from satellite 220 measurements and reanalysis fields. The AEROSE campaigns encountered significant Saharan 221 dust outflow events as shown in Figure 2 (rows 5 and 6); and the measurements taken during the 222 AEROSE campaigns provide basic inputs for atmospheric radiative transfer calculations in the 223 tropical North Atlantic area.

224 RRTMG calculations with and without aerosols indicate the dust radiative effects on the
225 SRF, which are defined here as the dust contaminated SRF minus the clear-sky SRF. Figure 3
226 shows the surface shortwave and longwave radiation changes due to the presence of dust
227 aerosols.

228	Since the dust aerosols scatter and absorb the solar radiation that would otherwise reach
229	the sea surface, there is a reduction in the surface shortwave radiation, as shown in Figure 3:
230	most of the negative anomalies are smaller than -90 W/m^2 , however, there are some extreme
231	values from -120 W/m ² to -190 W/m ² for AOD > 0.8. Because most of the radiosondes were
232	deployed shortly before AQUA satellite overpass times (around 13:30 ascending node time), the
233	stronger solar insolation increases the calculated shortwave anomalies at these times to be above
234	the diurnal-mean. The dust also emits longwave radiation which warms the sea surface; the dust
235	aerosols introduced positive longwave flux anomalies as large as 14 W/m^2 in our simulations.
236	Note that different (yearly) AEROSE campaigns would have encountered different vertical
237	distribution of dust aerosols (shown in Fig.2), which could explain the differences among
238	different years in Fig.3, for instance the much lower dust in 2013 (leg 2) leading to smaller
239	shortwave radiation reduction. Most of the large anomalies occurred during the AEROSE 2007
240	and 2019 cruises, because there were strong dust outbreaks during that time (Figure 2). Based on
241	the RRTMG calculations, the dust aerosols introduce negative shortwave anomalies of -130
242	$W/m^2/AOD$ and positive longwave anomalies of 12 $W/m^2/AOD$. Our calculated anomalies are
243	significantly higher than other studies, for example Yu et al. [2006] found -68 W/m ² /AOD and
244	Song et al. [2018] found -83 W/m ² /AOD direct radiative anomalies. This may be due to different
245	model inputs, as other studies used satellite data as input, the abnormal atmospheric cases can
246	introduce biases in the satellite-derived SRFs [Kato et al., 2020; Loeb et al., 2020]. Also, our
247	study is limited to outbreaks of Saharan dust aerosols in the North Atlantic area while other
248	studies were based on larger research areas where there may be aerosols with different
249	characteristics.

250 It is worth noting that for the RRTMG calculations, any inaccuracies in the MERRA-2 251 dust aerosols profiles, such as reported by Kramer et al. [2020], could introduce errors and 252 uncertainties. MERRA-2 assimilates the AOD derived from satellite and in-situ measurements 253 and determines the vertical distribution through the model physics. This study assumes that 254 MERRA-2 provides reasonably accurate dust profiles; MERRA-2 AODs have been validated 255 using independent dust aerosol observations from satellite, aircraft, and ground-based 256 observations [Bosilovich et al., 2015; Buchard et al., 2017; Gelaro et al., 2017], the results show 257 the AOD in MERRA-2 generally compared well.

3.2 SSTskin response to dust radiative effects

259 There have been several models [Gentemann et al., 2009; Minnett et al., 2011; Zhang and 260 Zhang, 2012] that discuss the SST_{skin} variations as a function of wind speed: the SST_{skin} variation 261 is weak under stronger wind and well-mixed conditions. As indicated by Foltz and McPhaden 262 [2008]; Lau and Kim [2007]; Martínez Avellaneda et al. [2010], found most of the cooling 263 anomalies at our study regions are driven by surface wind stress, which modulates the sensible 264 and latent heat loss and the horizontal current-introduced heat advection. The shortwave 265 radiation anomalies do not produce as large an effect on SST_{skin} variability as wind stress; thus, 266 only the situations with wind speed < 4m/s are selected here to reduce the wind stress forcing on 267 SST_{skin}.

Figure 4 shows the time series of the SST_{skin} changes due to dust aerosol (top), and the distributions of the SST_{skin} change at the times and places of radiosonde launches (bottom). Clearly, the significant SST_{skin} changes are within the dust outflow region; overall, the dust aerosols introduce cooling anomalies to SST_{skin} . When the wind speed is low, the absorbed solar radiation leads to a stable stratification in the upper ocean that results in an increase in

273 temperature located near the ocean surface, resulting in a diurnal warm layer. The diurnal 274 warming temperature increase can be several K, and it has been captured in field measurements 275 [*Donlon et al.*, 2002; *Gentemann and Minnett*, 2008; *Minnett et al.*, 2011] and in satellite data 276 [*Gentemann et al.*, 2003; *Marullo et al.*, 2010]. However, the dust aerosols reduce the downward 277 shortwave radiation, reducing the SST_{skin} by as much as -0.24 K. The overall cooling magnitudes 278 are consistent with those of previous studies [*Foltz and McPhaden*, 2008].

279 The SST_{skin} response to dust-induced radiation changes for wind speed < 4m/s are plotted 280 against AOD at 550 nm wavelength in Figure 5, separately by daytime (left) and nighttime 281 (right). From the daytime linear regression analyses shown in Figure 5 (left), the AOD of 1.4-1.6 282 can introduce cooling effects as large as -0.24 K; it can also introduce warming effects of 0.04 K 283 during the daytime, depending on the solar zenith angle and dust layer's temperature and 284 altitude. In general, when the solar zenith angle is small, the dust aerosols can block more 285 downward shortwave radiation that reaches the Earth's surface and further decrease the SST_{skin}. 286 However, when the solar zenith angle is $> 50^{\circ}$, the dust aerosol introduced surface shortwave 287 anomalies may not be as large as the longwave anomalies; the increased downward longwave 288 radiations can warm the sea surface. Thus, the dust aerosols can introduce both warm and cold 289 anomalies to SST_{skin} during daytime, depending on the solar zenith angle. The results of studying 290 different dust aerosol temperatures and altitudes will be reported in Section 4.

Since the SST is nearly everywhere higher than the air temperature, the upward longwave radiation is greater than downward. With the dust-introduced increased net longwave radiation $(LW_{net} \text{ in equation 5})$, the surface net heat flux Q_c will decrease, so the temperature drop across the thermal skin layer (ΔSST_d in equation 3) will be reduced; thus the dust aerosol longwave radiation can introduce a warm SST_{skin} anomaly during the nighttime, as shown in Figure 5

(right). However, the nighttime warm anomalies have only been calculated for low wind speeds; the averaged magnitudes of warm anomalies are usually < 0.03 K, and such a longwave heating anomaly by dust layer has usually been ignored in published studies. The dust-induced *SST*_{skin} anomalies are related to the AOD: there were strong dust outbreaks during 2007, 2009, and 2019 AEROSE cruises as shown in Figure 2 (rows 5-6); the corresponding *SST*_{skin} anomalies were larger during these years and near 15°N, 25°W, where the downward SRF had the greatest impact on *SST*_{skin}.

4. Cases of the Effects of Aerosol Variability

As discussed above, the downward SRF is modified by the dust aerosols. A comparison of individual radiosonde profiles with the MERRA-2 reanalyzed dust aerosol fields was made to assess the effects of dust layers encountered at different altitudes and with different vertical extents during the AEROSE cruises.

307 The AEROSE cruises sampled the Saharan dust outflow regions each year; the dust can 308 be lifted to 400 hPa as shown in Figure 2 (rows 5 and 6). Figure 6 shows four situations when the 309 dust aerosols reduce the SST_{skin} by -0.238 K to -0.151 K. The relative humidity and air 310 temperature profiles measured by radiosondes are plotted with blue and Magenta lines; the M-311 AERI derived SST_{skin} is shown as the green star; the shading indicates the MERRA-2 DU002 312 dust mixing ratio as shown on color bar at right. The MERRA-2 dust mixing ratio during these 313 four radiosonde deployments indicate large scale dust aerosol outflows; the elevated dust aerosol 314 layers are always associated with dry air layers as shown in Figure 6. On July 22, 2009, 14:44 315 UTC, the dust aerosol layer was lifted to 800 hPa, but the dust concentration was not as large as 316 in other situations shown; thus, the dust aerosols only reduced the SST_{skin} by -0.151K. On the

317 days of July 26, 2009 and August 2, 2011, there were intense dust outbreaks extending vertically 318 from 900 hPa to 500 hPa; the dust aerosols blocked large amounts of the downward shortwave 319 radiation reaching the sea surface and further decrease the SST_{skin} by -0.238K and -0.163K 320 respectively. The dust particles absorbed the shortwave radiation and warmed the lower 321 troposphere, as shown by the temperature inversions at the bottom of the dust layer; the dry 322 layers associated with the dust aerosol are also obvious. On February 4, 2013, the dust layer was 323 not lifted to such high altitudes as on other days. The radiosonde captured the air temperature 324 and relative humidity anomalies around 980 hPa associated with the aerosol dust layers. 325 However, the dust layer temperatures were always lower than the SST_{skin}, which introduced a 326 cold effect on the SST_{skin}.

327 Data from the AEROSE cruises provide a valuable way to test the dust effect on SST_{skin} 328 under a variety of meteorological conditions. Figure 7 shows another four situations when the 329 dust aerosols increase the SST_{skin} by 0.017 K to 0.081 K. The two radiosonde profiles on May 14, 330 2007 show a warm effect when dust aerosol is present, and although the dust aerosols occur at 331 high-altitude, to 650 hPa, the relatively high temperatures (299K-300K from 850hPa to 900hPa 332 for dust layers) compared to the SST_{skin} lead to comparable warm effects. On the days when the 333 ship Ronald H. Brown entered significant, large-scale warm dust outflow events, the warm 334 SST_{skin} effects were more pronounced during nighttime or for large solar zenith angles. Similar 335 situations can be seen on March 17, 2019, but the warm effects were not as large as during 336 AEROSE 2007 when the dust layer temperature was higher than the SST_{skin}, the aerosol dust 337 layers emitted radiation warmed the SST_{skin}.

338 The results from various radiosonde cases show the dust aerosol effect on SST_{skin} depends 339 on several factors, such as the temperature contrast between the dust layer and SST_{skin} , the 340 characteristics of the dust layer, concentration and altitude.

5. Summary

341 Previously published research has been directed at quantifying the direct dust effects on 342 the SRF and SST using satellite data (e.g., Evan et al. [2009]; Foltz and McPhaden [2008]; 343 Martínez Avellaneda et al. [2010]; Song et al. [2018]); however, the downward SRF derived 344 from satellite measurements over the ocean have large uncertainties over the tropical Atlantic 345 Ocean [Kato et al., 2018]. Given the inaccuracies in satellite retrievals, accurate independent 346 shipboard measurements in the tropical Atlantic Ocean area provide an independent 347 representation of the atmosphere and ocean that can be used to investigate the influence of the dust aerosols on SST_{skin} variability. The NASA MERRA-2 reanalysis fields augment the 348 349 radiosonde data to characterize the vertical dust aerosol profiles at the times and places where 350 radiosondes were launched, and to provide inputs for radiative transfer calculations. This study 351 includes the RRTMG-simulated surface shortwave and net longwave downwelling radiative 352 changes due to dust and calculates the corresponding thermal skin layer temperature changes. 353 The radiative transfer model was driven with oceanic and atmospheric variables taken from 354 AEROSE cruises and from MERRA-2 fields resulting in estimates of radiative forcing of the 355 SST_{skin}.

As our ship-based sampling covered weak to strong dust outbreaks, the dust aerosol effects on *SST_{skin}* also vary temporally and spatially. Based on the RRTMG model calculations under various dust distributions, we estimate the dust can introduce a reduction of up to 190

 W/m^2 in surface shortwave radiation at around 13:30 local time and an increase of 14 W/m^2 359 360 surface longwave radiation. As the SST variability is mainly responsive to wind-induced 361 turbulent latent and sensible heat loss at the surface [Foltz and McPhaden, 2008; Lau and Kim, 362 2007], we have simulated the SST_{skin} variations with models of the thermal skin layer for wind 363 speeds < 4 m/s. The dust aerosols can introduce warm and cold anomalies to SST_{skin} during 364 daytime, depending on the solar zenith angle, dust layer concentration, temperature and altitude; 365 the reduction in surface shortwave radiation can decrease the SST_{skin} by as much as -0.24 K. The 366 anomalous increase in the surface longwave radiation is associated with an increase in SST_{skin} of 367 up to 0.06 K, which is identifiable at daytime and nighttime. The cooling and warming anomalies 368 cover a broad region with dust outbreaks and are strongest in the central sub-tropical North 369 Atlantic Ocean, which are consistent with previous studies [Evan et al., 2011; Foltz and 370 *McPhaden*, 2008].

371 This study combines in situ measurements with reanalysis fields to improve our 372 knowledge of the dust forcing on SRF and *SST*_{skin}, which is key to understanding better the roles 373 of aerosols and their feedbacks in the climate system. Reanalyses such as MERRA-2, as they 374 evolve into more comprehensive earth system reanalyses, should include coupling feedbacks 375 between ocean (via SST_{skin}), meteorology and aerosols. For instance, the satellite brightness 376 temperatures that are assimilated in MERRA-2 and ERA-5 are mostly clear-sky radiances, that 377 is, measurements that are deemed to be in cloudy, aerosol, or rain conditions are not assimilated 378 in the infrared; ERA-5 includes "all-sky" microwave radiance data assimilation, see Hersbach et 379 al., 2020 [https://rmets.onlinelibrary.wiley.com/doi/full/10.1002/qj.3803] for details. Since 380 reanalyses (e.g., MERRA-2) capture detailed aerosol variability, how does that impact brightness 381 temperature simulation? And in-turn, does that impact meteorological assimilation? The study of

Kim et. al., 2018 [https://gmao.gsfc.nasa.gov/pubs/docs/Kim1018.pdf] showed that indeed more accurate brightness temperatures can be simulated by including aerosol information in radiative transfer modeling, however the computation cost was burdensome. A more recent work by Choi et al., 2020 [https://agupubs.onlinelibrary.wiley.com/doi/10.1029/2019MS001890] also provide encouraging results in this direction. In closing we recommend that more detailed determination of the dust effects on SRF requires better knowledge of dust radiative properties and vertical profiles derived, for example, from ship-based lidar.

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- 398 https://www.aoml.noaa.gov/phod/pne/cruises.php and NOAA NCEI
- 399 https://www.nodc.noaa.gov/archivesearch/rest/find/document?f=searchPage&searchText=ron+br
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- 587

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Figures:



Figure 1. Cruise tracks of each AEROSE campaign, the points indicate the radiosonde deployment locations. The colors indicate day of the year.



Figure 2. Rows 1-2: Relative humidity measured by radiosondes launched from the ships, the dust introduced dry layers are visible on some days. Rows 3-4: Air temperatures measured by radiosondes [Luo et al., 2020b]. Rows 5-6: MERRA-2 dust mixing ratio at radiosonde deployment location and times along each AEROSE track, the shading indicates the dust mixing ratio as shown on the right.



Figure 3. Left: RRTMG-calculated Saharan dust shortwave reduction at the sea surface with AOD at 550 nm at the radiosonde deployment stations. Right: RRTMG-calculated Saharan dust longwave increases. The colors indicate the year.



Figure 4. Top: Time series of the SST_{skin} changes due to dust. The x-axis is the radiosonde deployment number, and the y-axis shows the simulated SST_{skin} changes, which are calculated as the difference between the SST_{skin} with and without dust. The unit is K. The colors indicate the deployment year. Bottom: Geographic distributions of the calculated SST_{skin} changes due to dust. The colors indicate the SST_{skin} change due to dust, as shown on the right with the unit of K. Note there are many points which have almost zero SST_{skin} change. The latitude and longitude ranges are reduced compared to Figure 1.



Figure 5. Left: Calculated daytime SST_{skin} changes due to dust with AOD at 550 nm, the SST_{skin} changes are large with high aerosol concentrations. Right: Calculated nighttime SST_{skin} changes.



Figure 6. Atmospheric and SST_{skin} profiles when the dust aerosols reduce the SST_{skin} . Magenta and blue lines indicate relative humidity and air temperature profiles from radiosonde. Green star indicates the corresponding M-AERI measured SST_{skin} when radiosondes were deployed. The shading indicates the MERRA-2 DU002 dust mixing ratio as shown on the right with the unit of kg/kg.



Figure 7. As Figure 6, but the dust aerosols increase the SST_{skin}.

Tables

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CRUISES	NUMBER OF RADIOSONDES	START	END	DAYS OF DATA
2007	96	2007-05-07	2007-05-28	22
2008	74	2008-04-29	2008-05-19	21
2009	78	2009-07-11	2009-08-11	31
2011	102	2011-07-21	2011-08-20	31
2013 Leg 1	111	2013-01-09	2013-02-13	36
2013 Leg 2	97	2013-11-11	2013-12-08	28
2015	92	2015-11-17	2015-12-14	28
2019	101	2019-02-24	2019-03-29	34
Total	751	2007-05-07	2019-03-29	231

591 Table 1 Details of the AEROSE cruises.