



Lunar Surface Science Workshop 2021

Resource Mapping in Extreme Environments

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Robert W. Moses

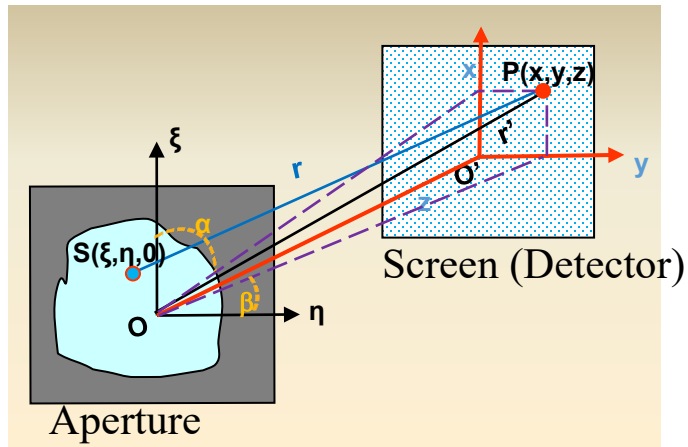
NASA Langley Research Center
Hampton, VA

August 18-19, 2021

Approximation of Diffraction Phenomena



$S(\eta, \xi, 0)$ at a coordinate $(\eta, \xi, 0)$ on an aperture and illuminates a point, $P(x, y, z)$ at a coordinate (x, y, z) on a screen.



$$u(P) = \frac{iA}{\lambda} \iint \frac{e^{-ikr}}{r} d\xi d\eta$$

where λ = wavelength, k = wavenumber of light.

r is a temporary variable for integration, r' is a fixed engineering parameter.

Approximation
with r'

$$u(P) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} S(\xi, \eta, 0) \exp[ik(l\xi + m\eta)] d\xi d\eta$$

where $r \cong r' - (l\xi + m\eta)$, $l = \frac{x}{r'} = \cos \alpha$; $m = \frac{y}{r'} = \cos \beta$

$$z > \frac{a^2}{\lambda}$$

Fraunhofer Terms

$$r = r' - \frac{x\xi + y\eta}{r'} + \left[-\frac{(x\xi + y\eta)^2}{2r'^3} + \frac{\xi^2 + \eta^2}{2r'} \right] + \dots$$

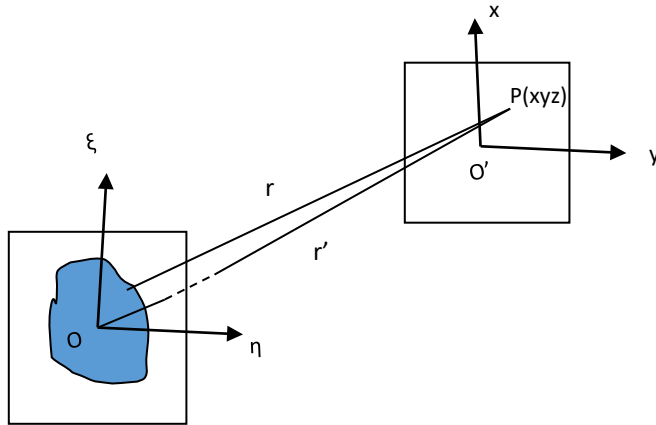
Fresnel Terms

$$u(P) = \frac{i}{\lambda z} \exp[-ikz] \iint S(\xi, \eta, 0) \exp\left[\frac{-ik}{2z} \left[(x - \xi)^2 + (y - \eta)^2\right]\right] d\xi d\eta$$

$$a < z < \frac{a^2}{\lambda}$$

Diffraction & Fresnel Number

$$u(P) = \frac{iA}{\lambda} \iint \frac{e^{-ikr}}{r} d\xi d\eta$$



$$F = \frac{a^2}{z\lambda}$$

a = radius of aperture
z = distance along OA
λ = wavelength

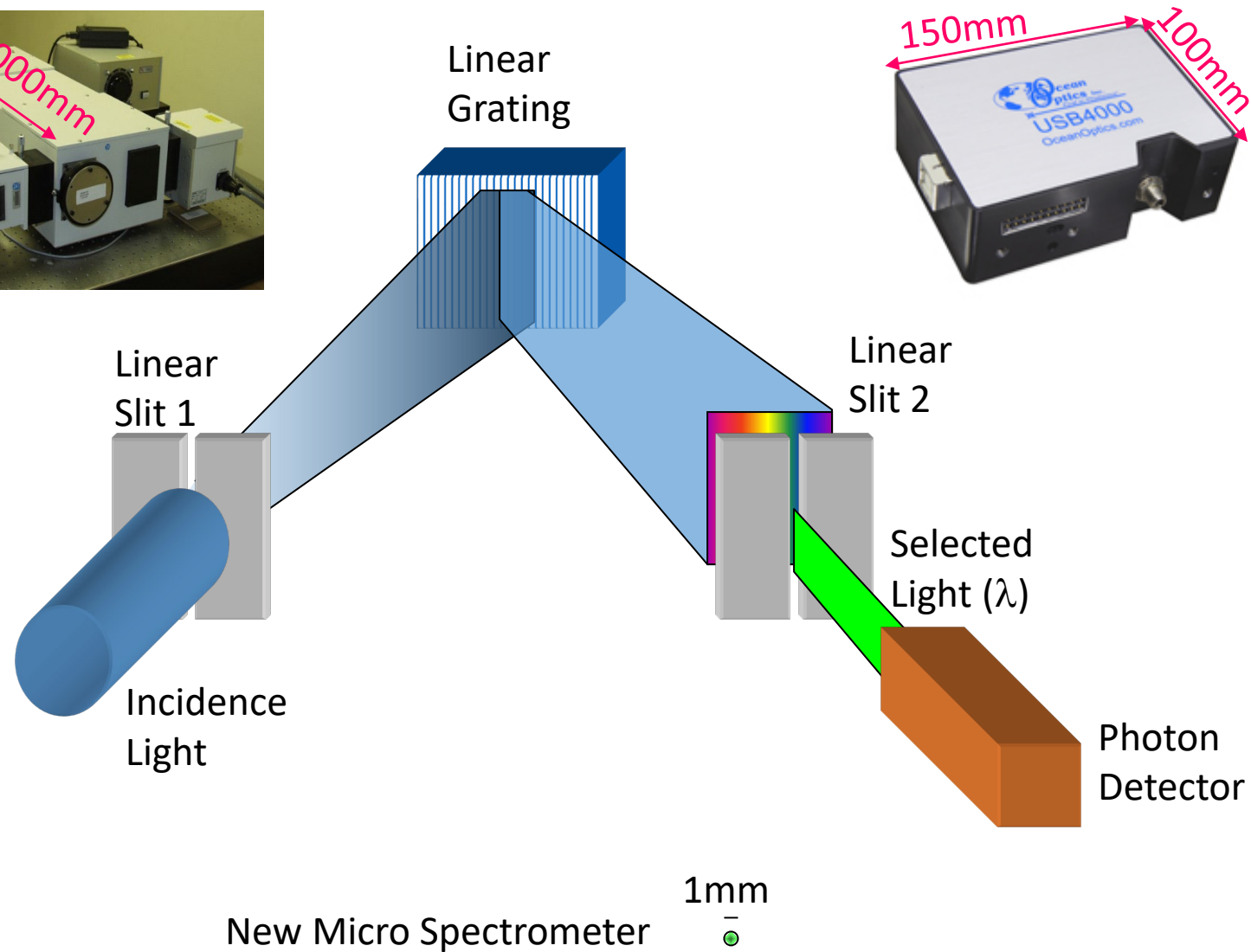
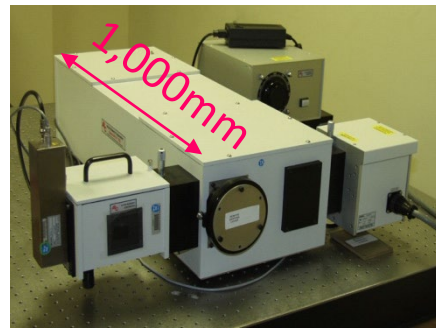
The value of F determines the valid diffraction approximation.

$F \ll 1$; Fraunhofer Diffraction

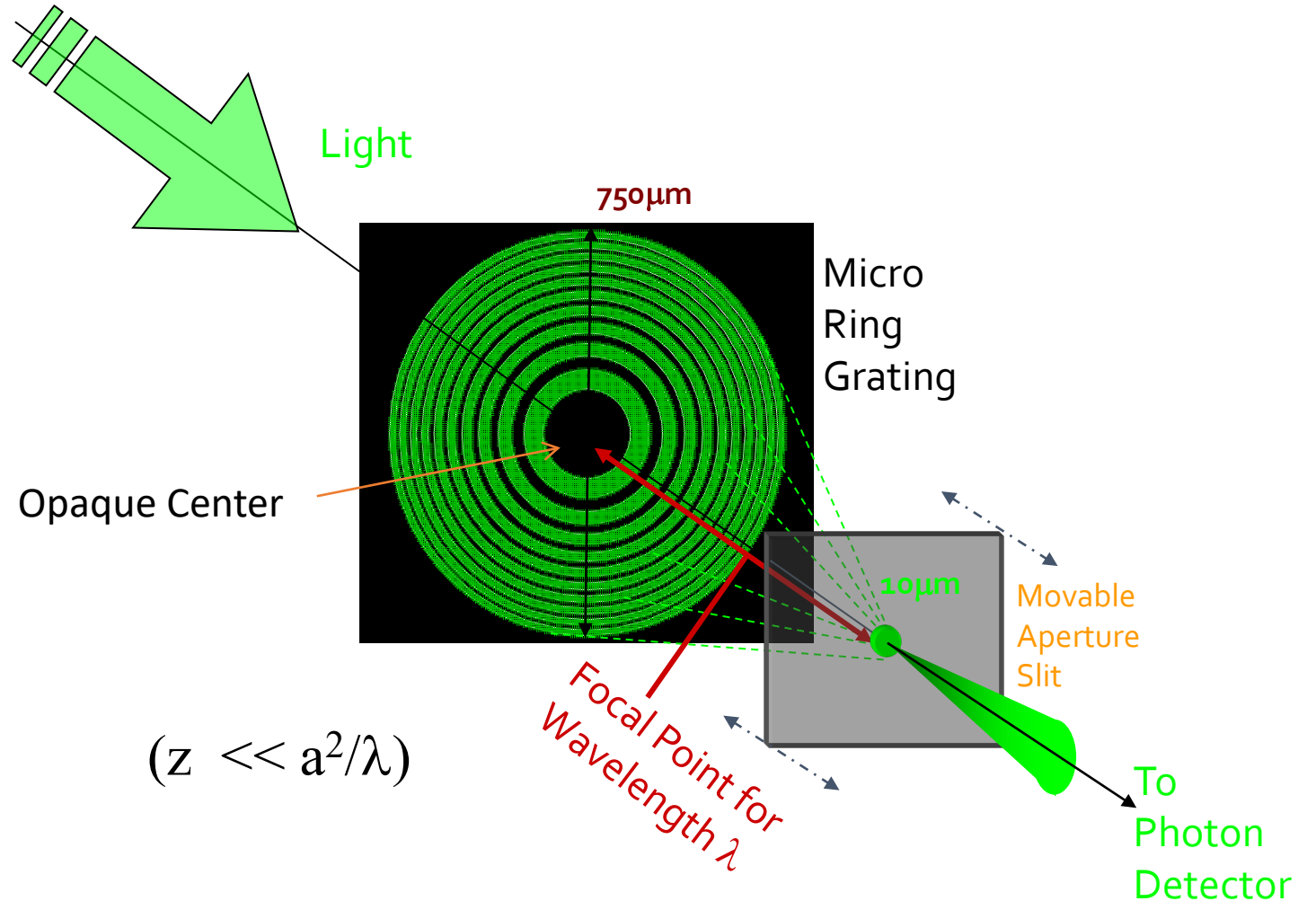
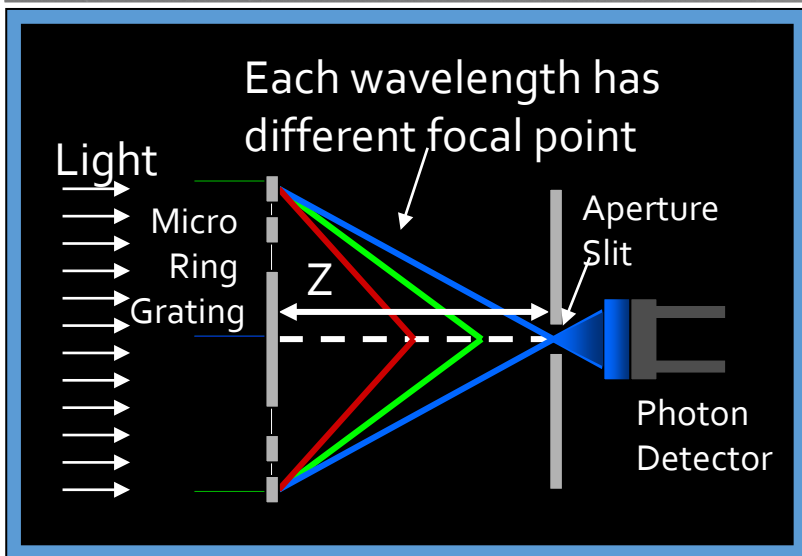
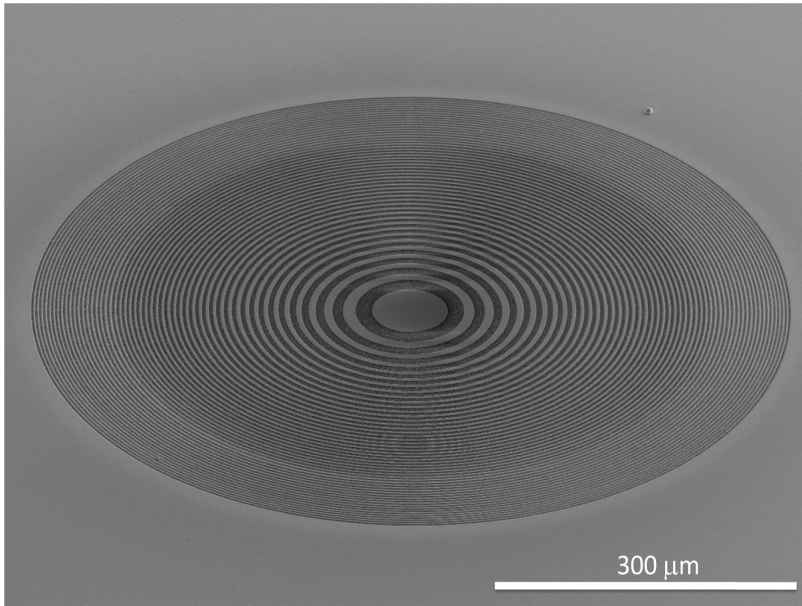
$F \gg 1$; Fresnel Diffraction

Most modern conventional spectrometers use the Fraunhofer diffraction because the resolution improves with a longer optical path.

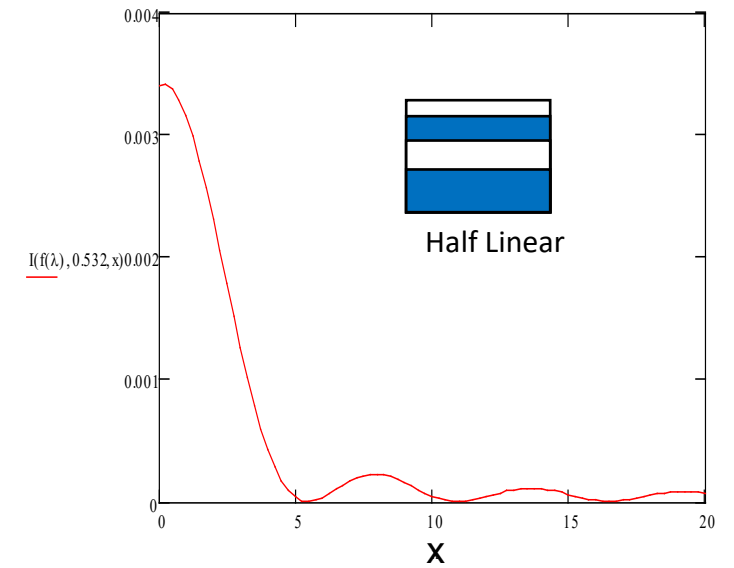
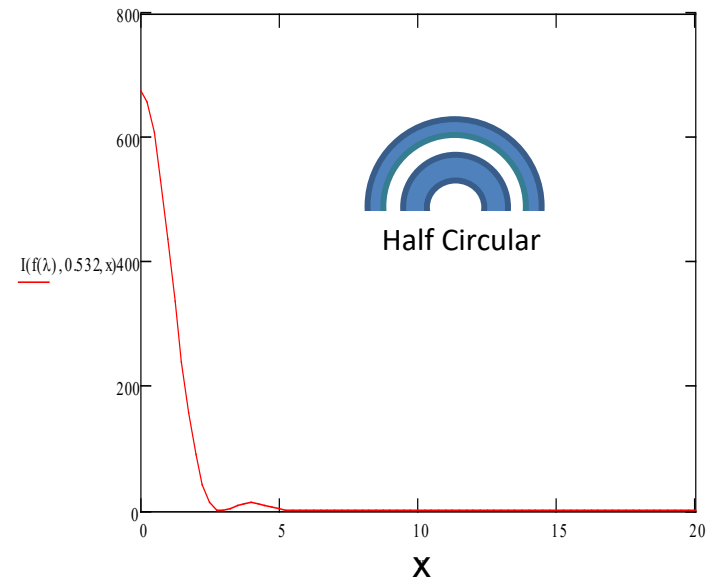
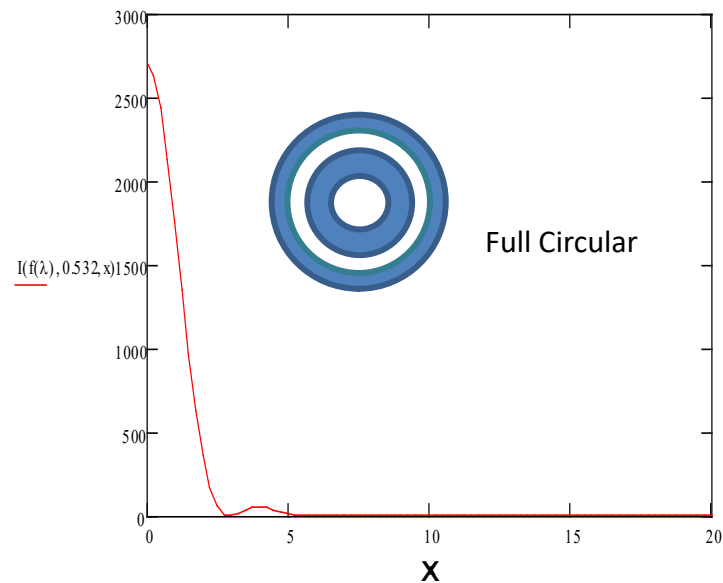
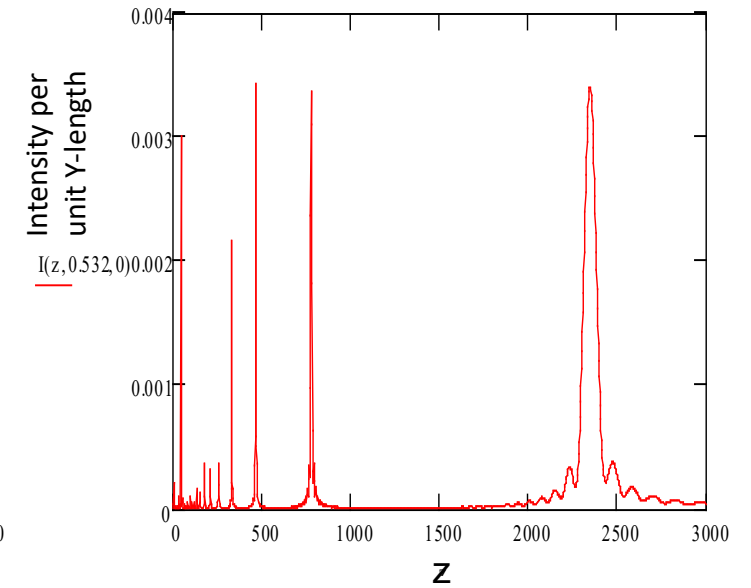
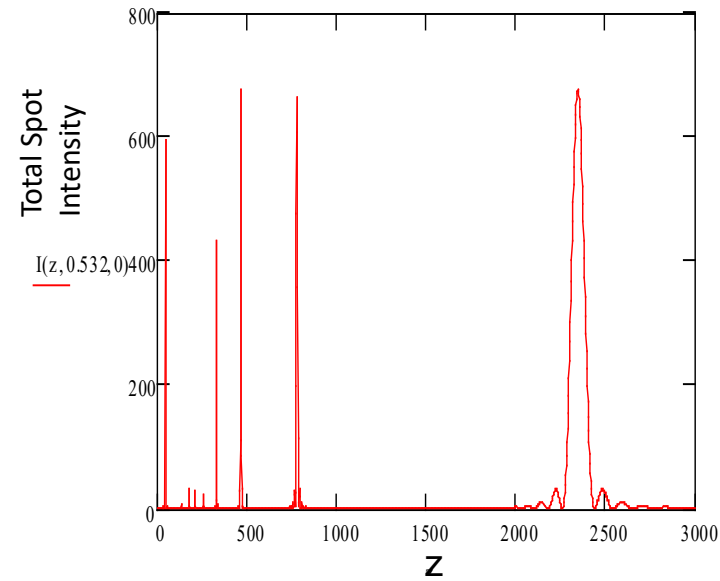
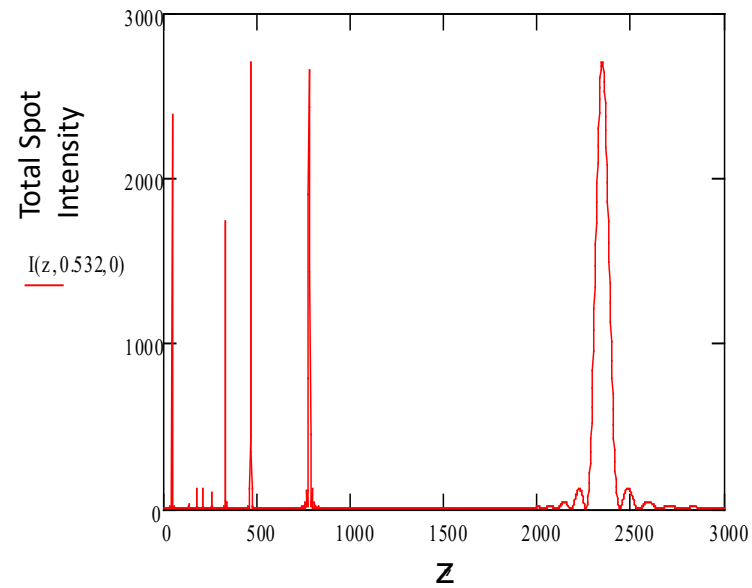
Conventional Fraunhofer Spectrometers



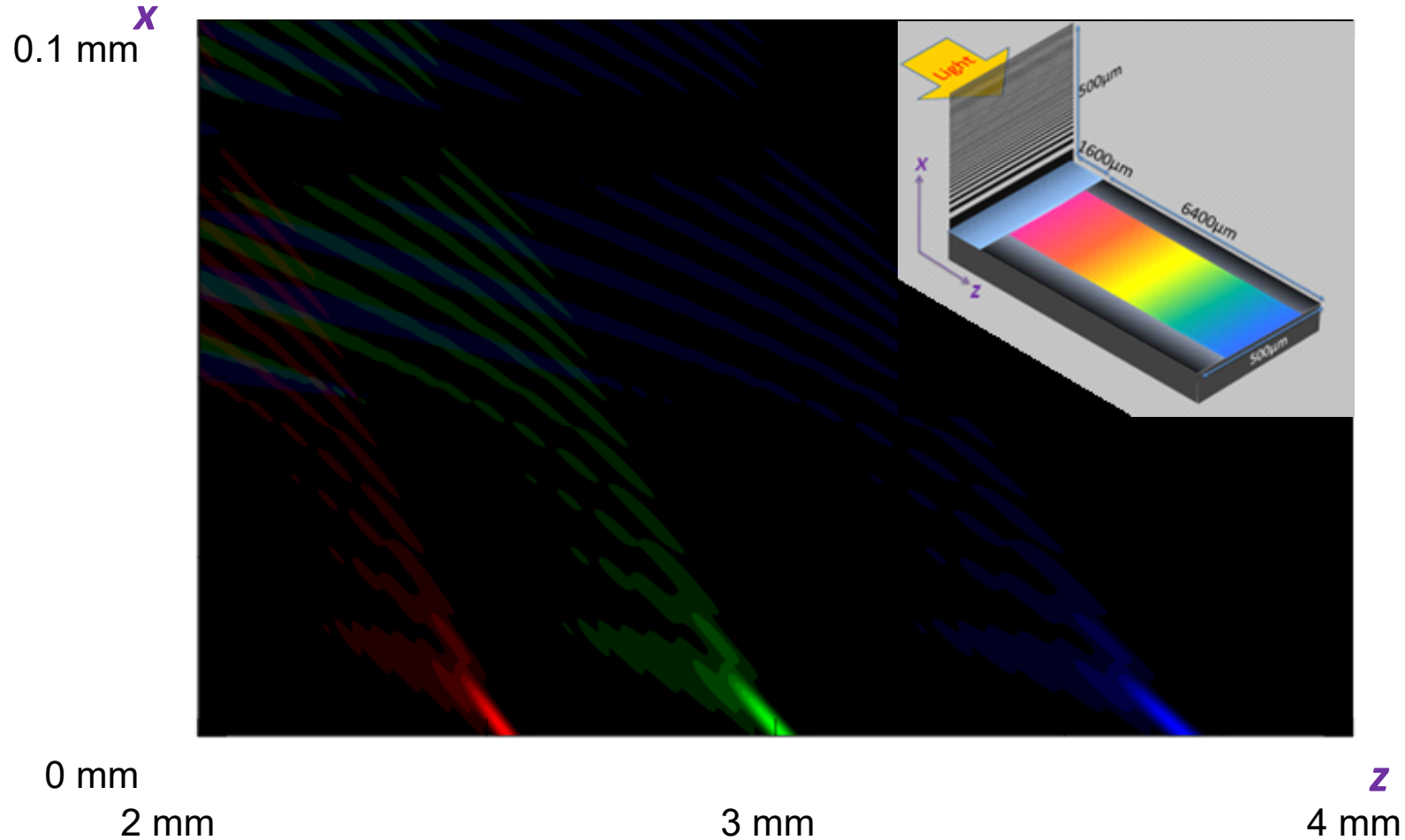
Working Principle of Fresnel Spectrometer with Negative Zone Plate



Optical Responses of Circular and Linear Gratings



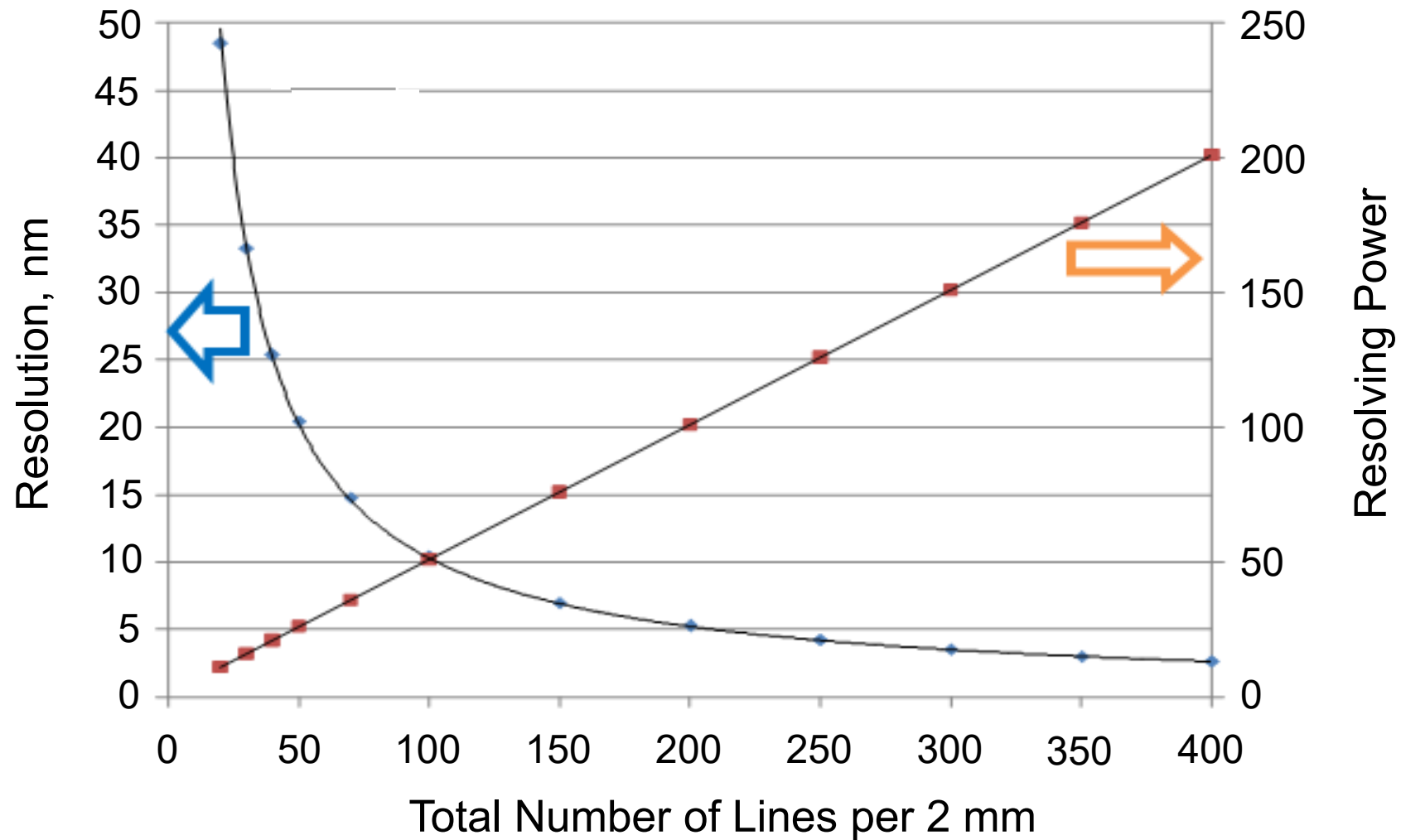
Fall of Lights onto Detector Surface



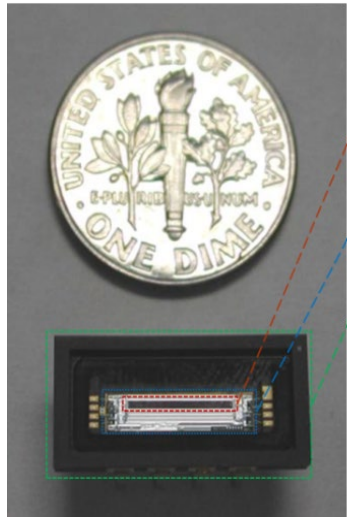
(632, 532, 432nm) **156 lines in 500 μ m**



Spectral Data & Resolution



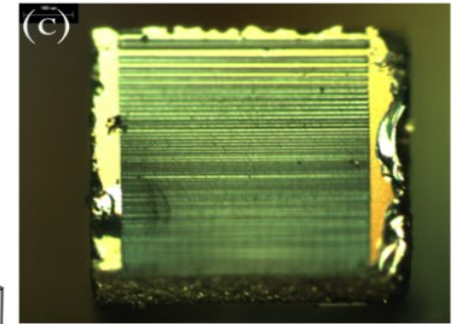
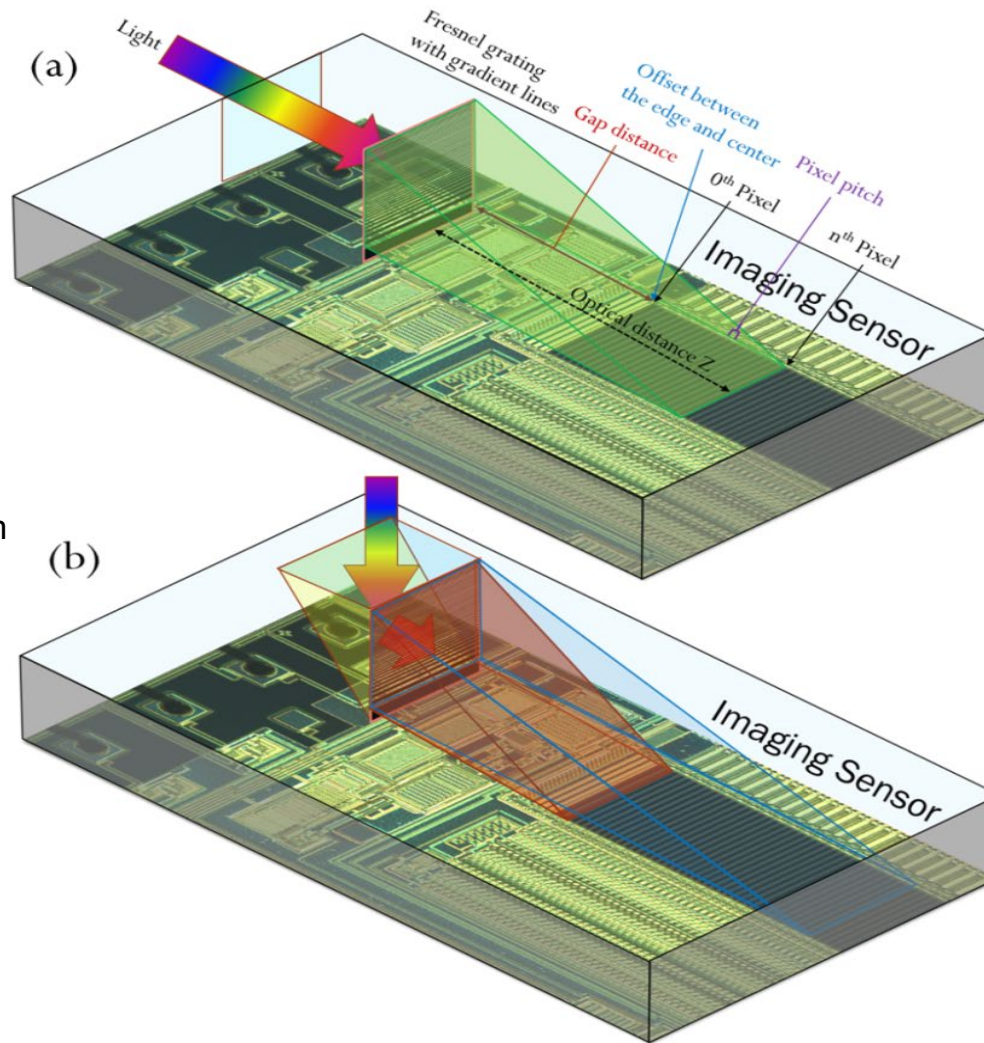
Spectrometer Chip



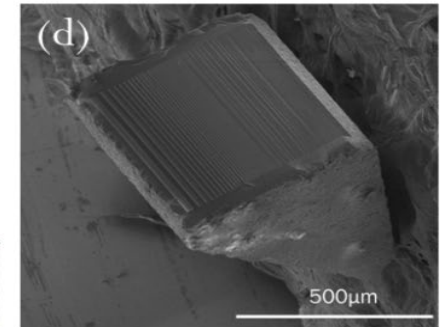
Active Area: 6.4 mm x 0.5 mm
(Spectrum Area)

Die Size: 9.5 mm x 2.5 mm

Packaging: 15.8 mm x 7.87 mm

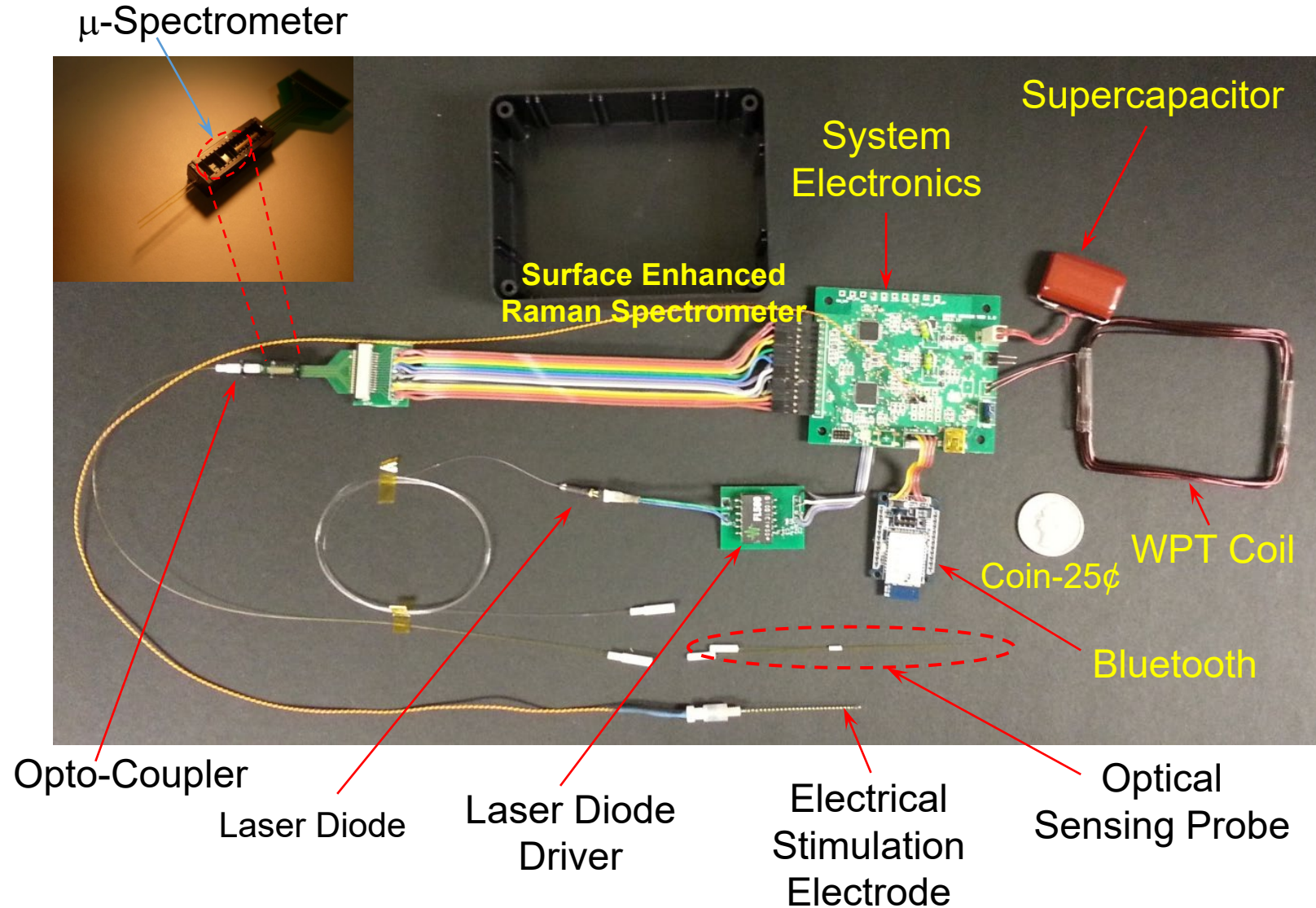


700 μm

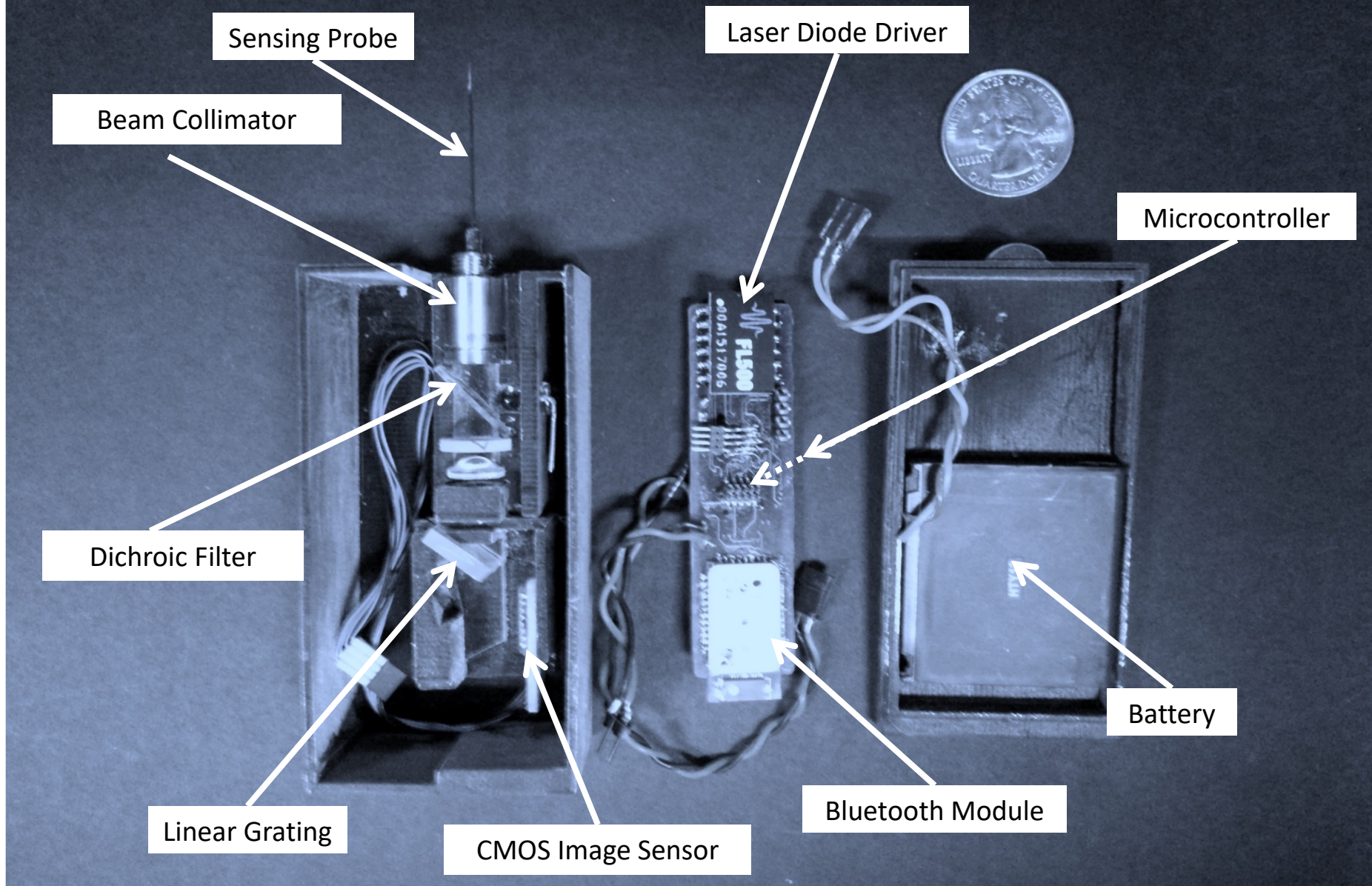


500 μm

1-D Scan Prototype System

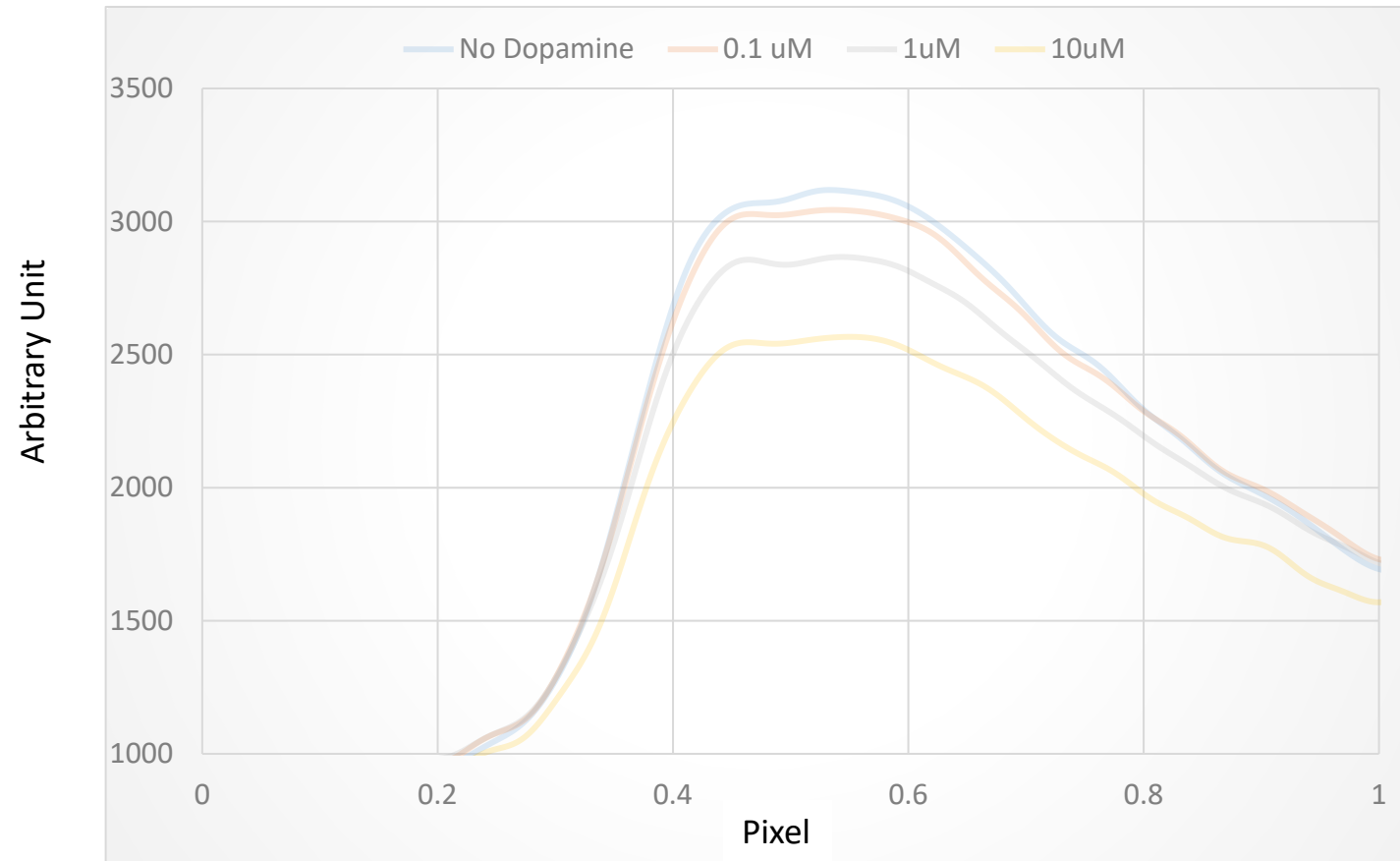


1-D Scan Prototype System

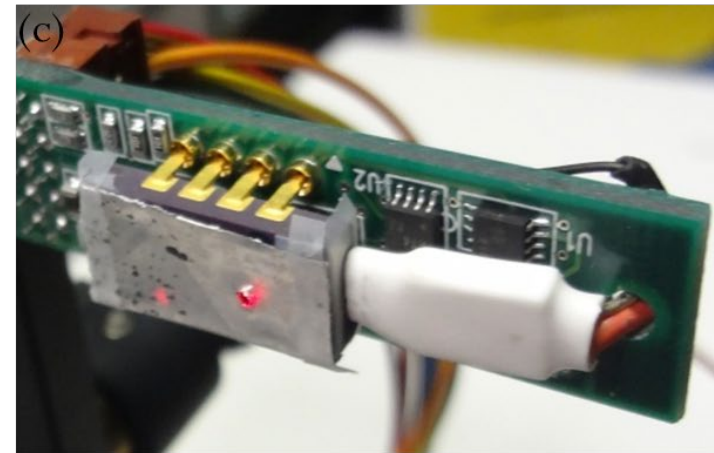
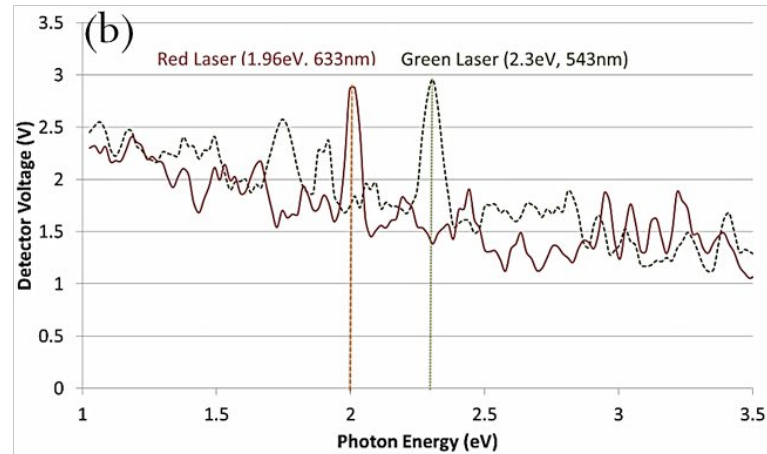
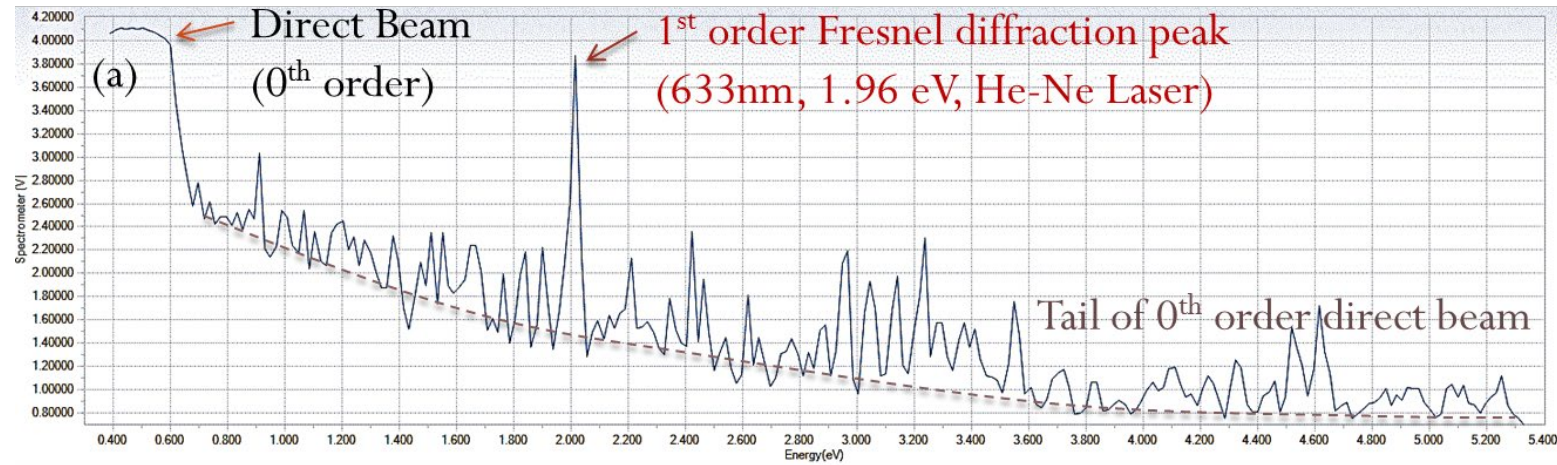




Dopamine Sensing Results



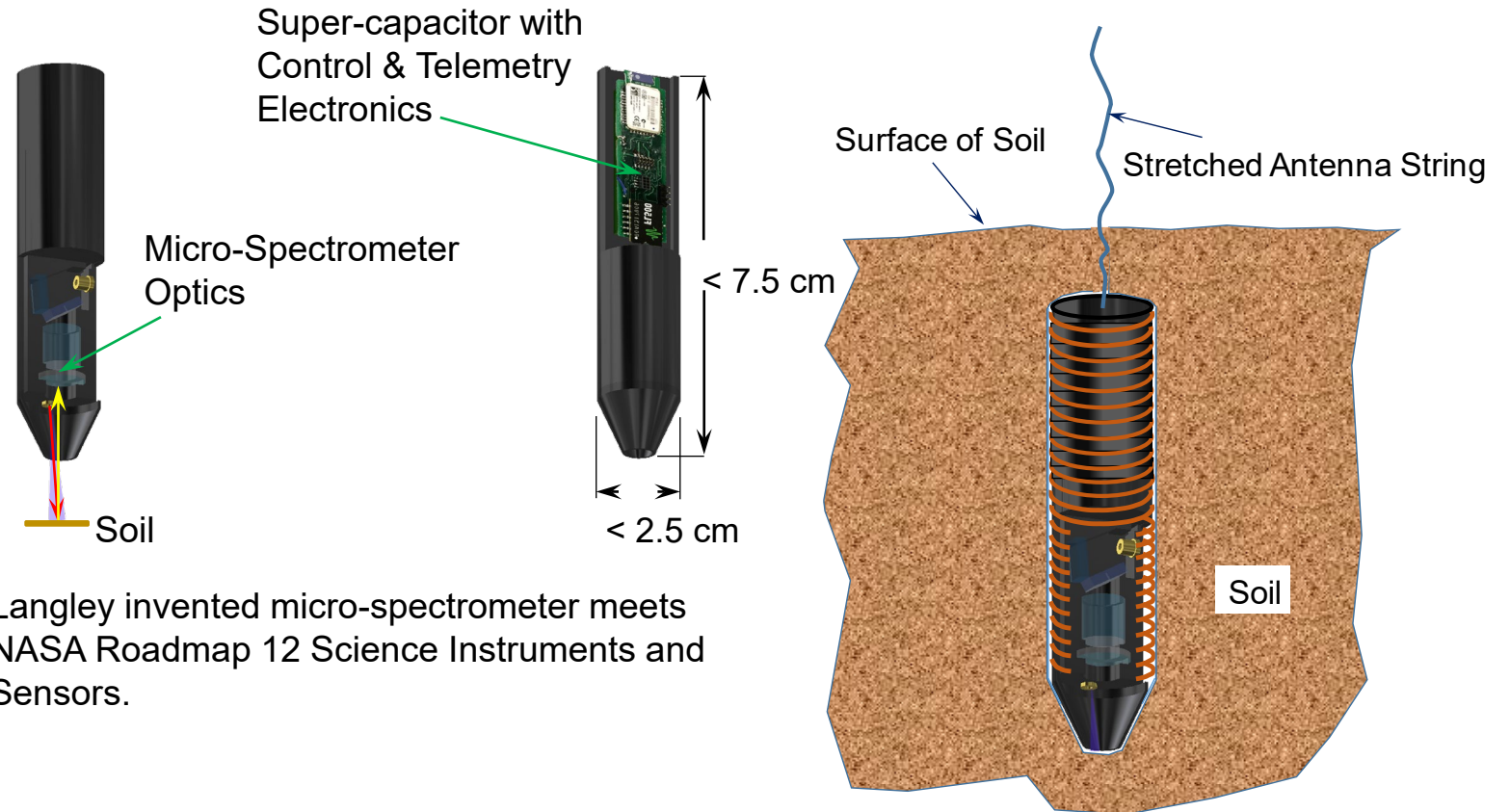
Optical Sensor Response



Bullet μ -Spectrometer

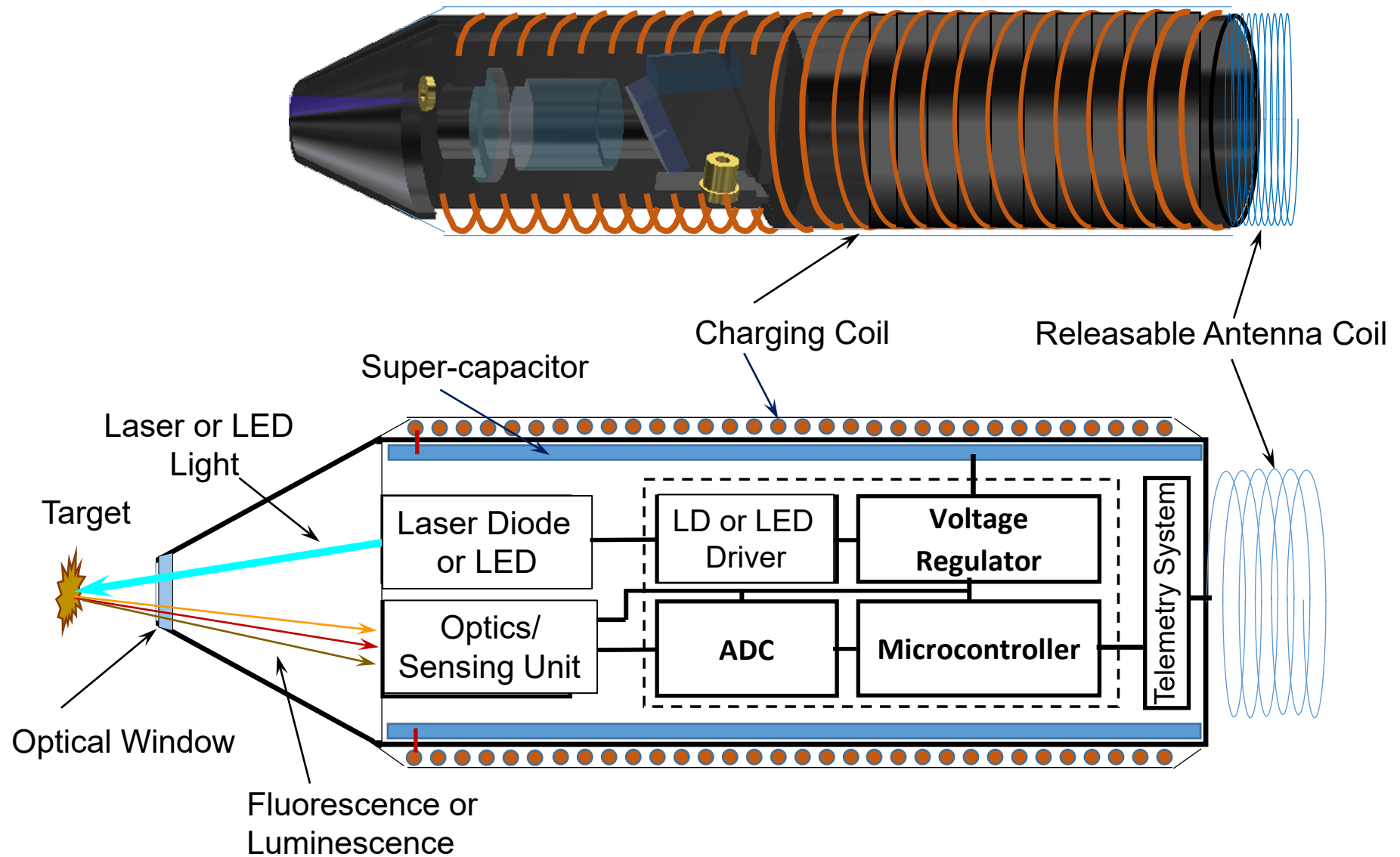
Micro-Spectrometer

- Principle of Fresnel diffraction
- Light source: Deep or vacuum UV from burst-mode LED
- Pulse mode power feed from super-capacitor
- Bullet-like consumable micro-spectrometers (< \$100)
- Shooting by compressed air gun

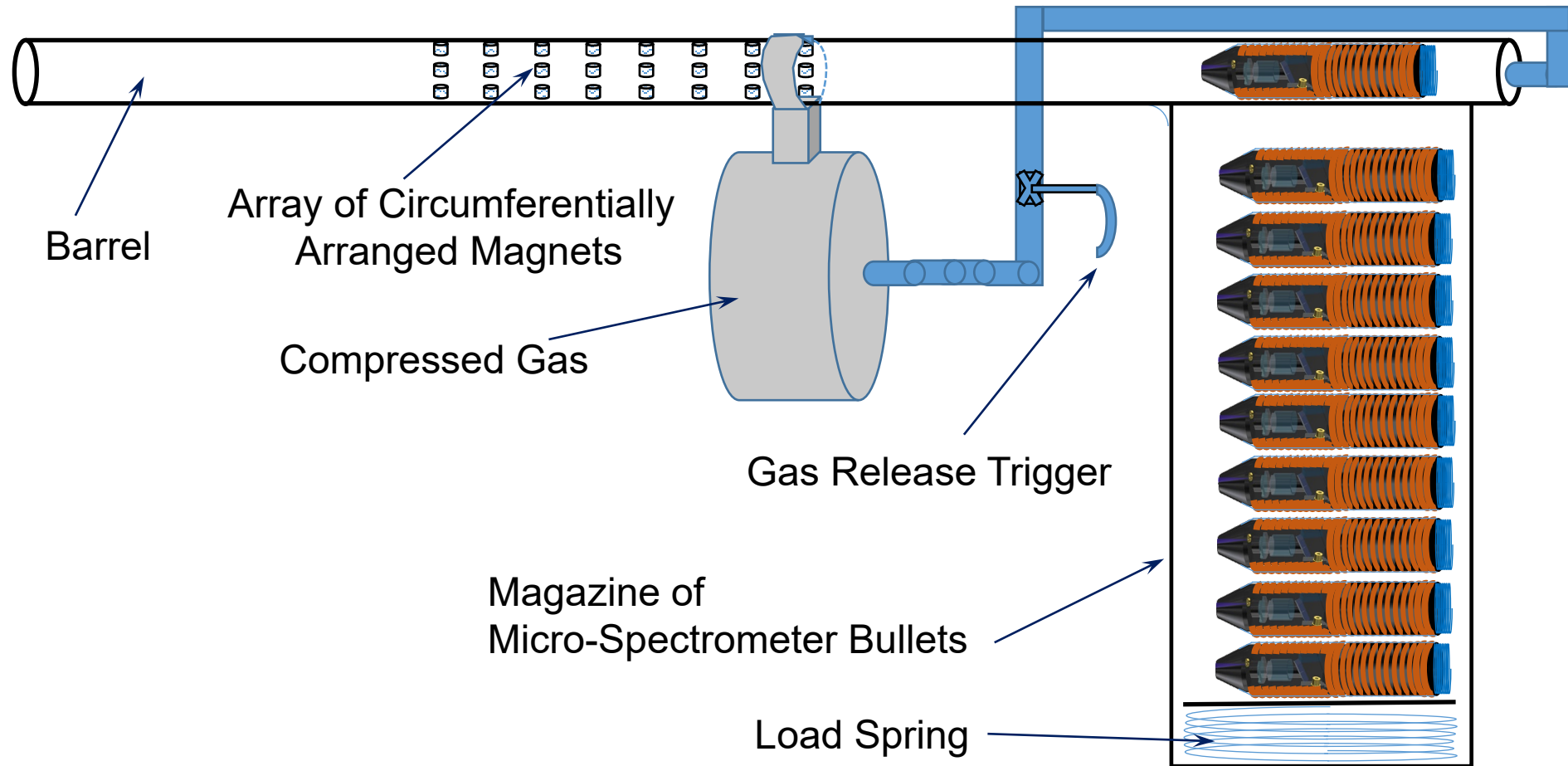


Langley invented micro-spectrometer meets NASA Roadmap 12 Science Instruments and Sensors.

Bullet μ -Spectrometer

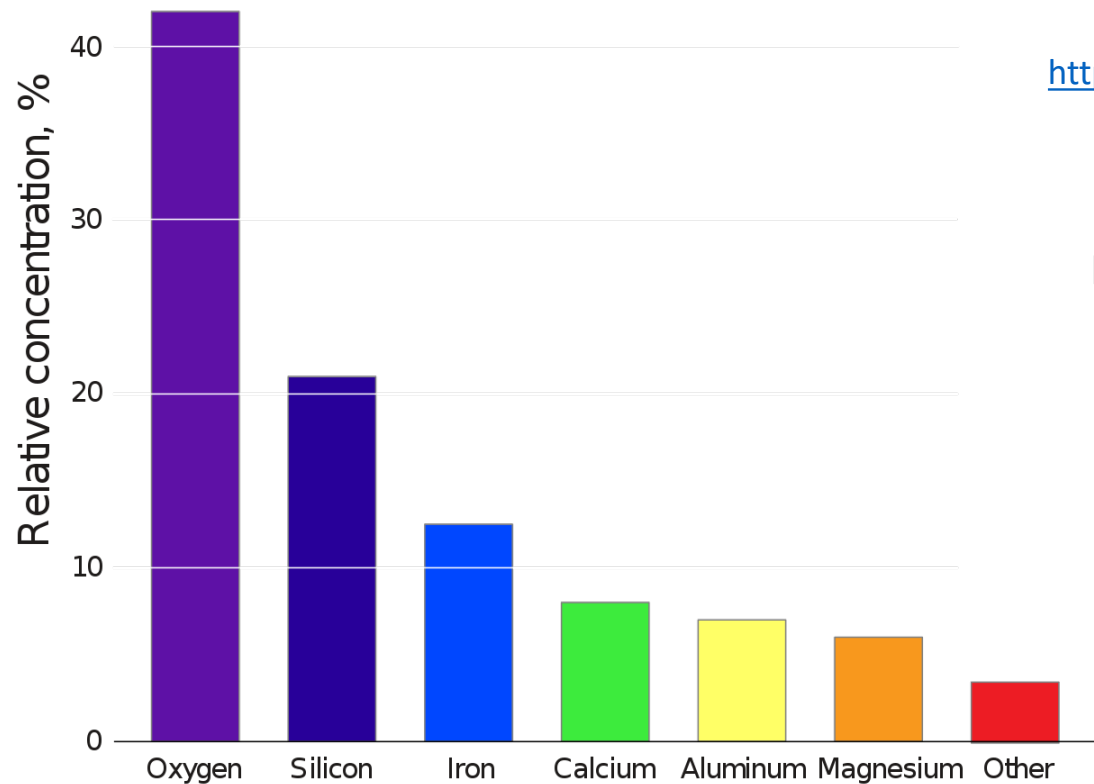


Gas Gun for Bullet μ -Spectrometer

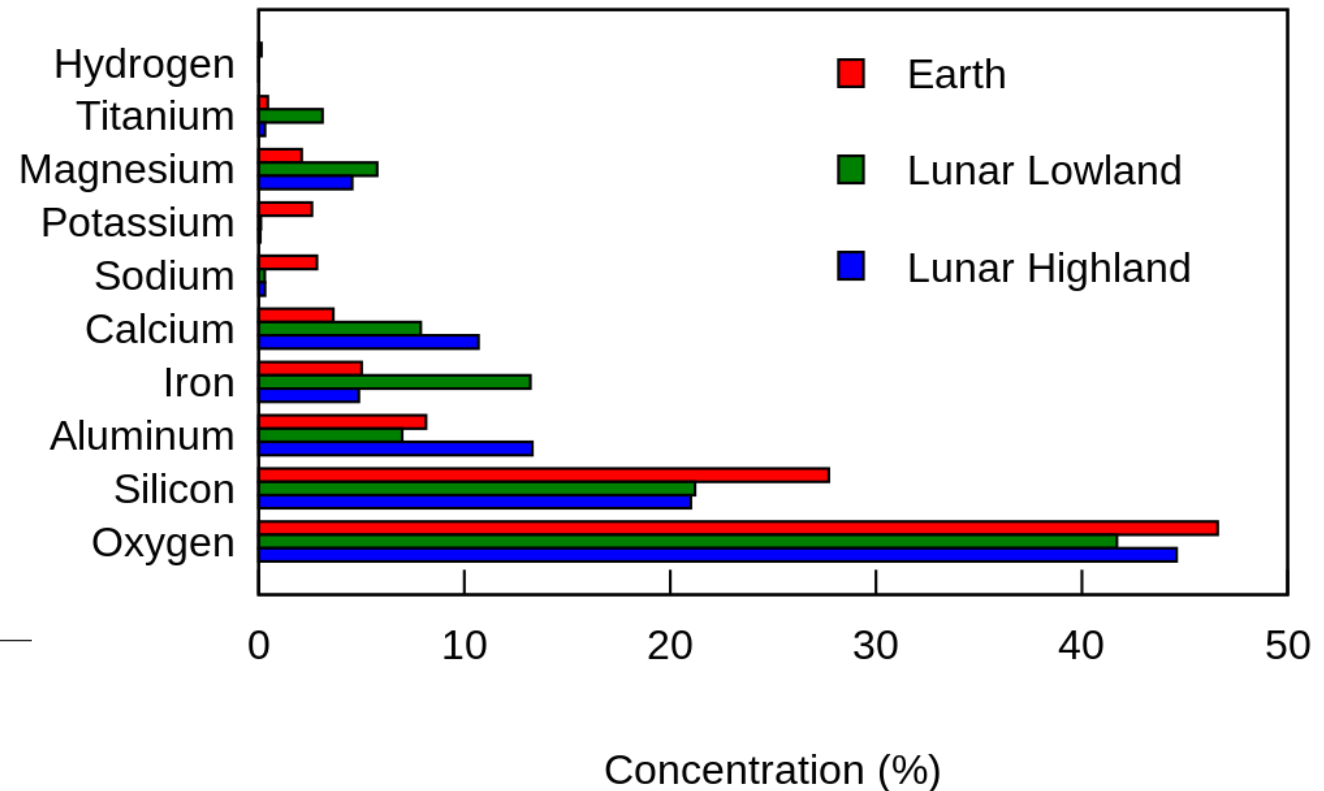




Composition of Lunar Regolith



https://en.wikipedia.org/wiki/Lunar_soil#/media/File:Moon_vs_earth_composition.svg

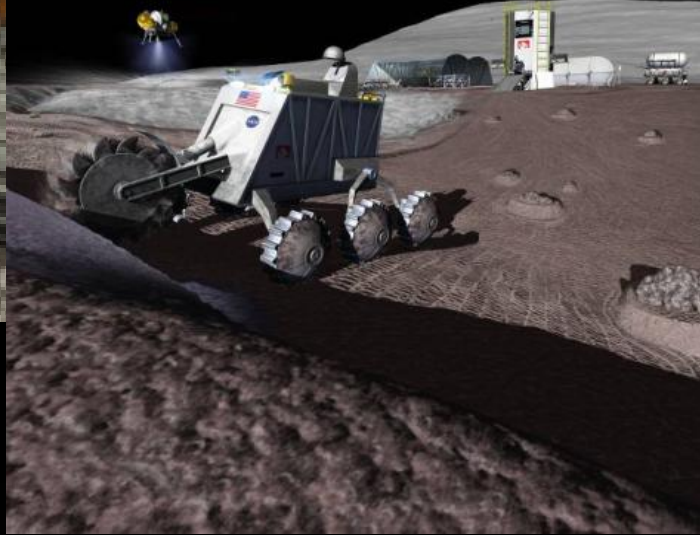


https://en.wikipedia.org/wiki/Lunar_soil#/media/File:Composition_of_lunar_soil.svg



**Resource Prospecting –
Looking for Resources**

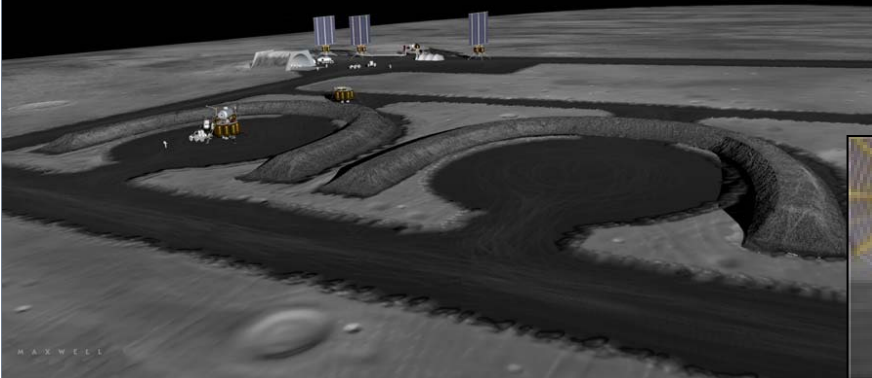
**Excavation & Regolith
Processing for H₂ & O₂
Production, Binders &
Aggregates**



Images credit: NASA

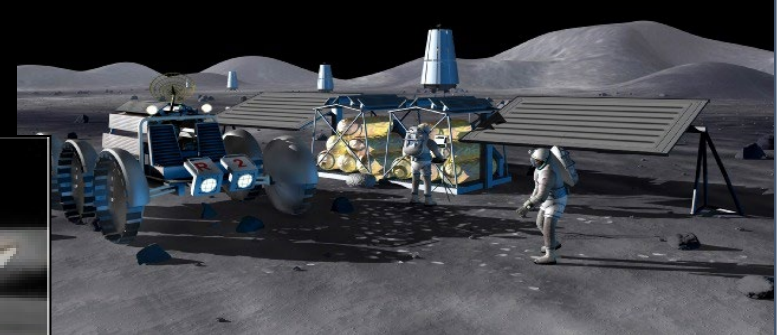
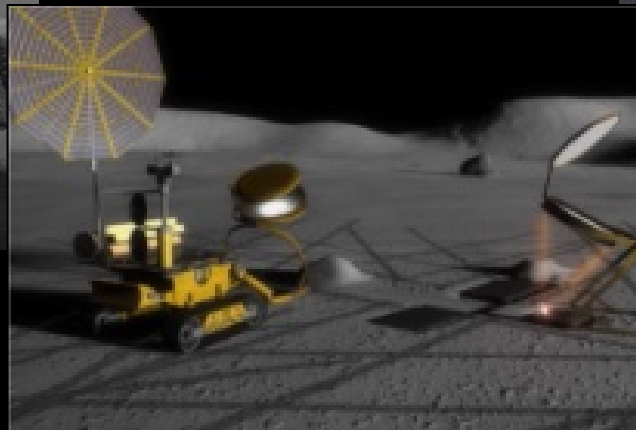


**Propellant Processing with
Lander & Pad Infrastructure**



**Habitat, Hangars, Dust Free
Zones, Landing Pads, Berm, and
Road Construction**

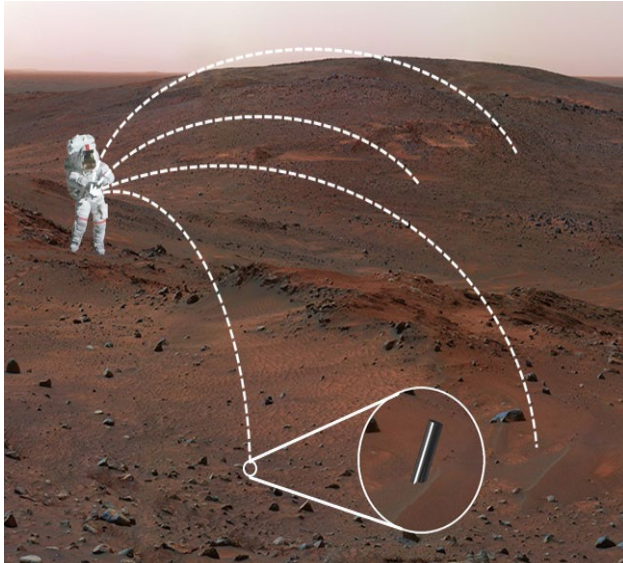
**Thermal Energy Storage
Construction**



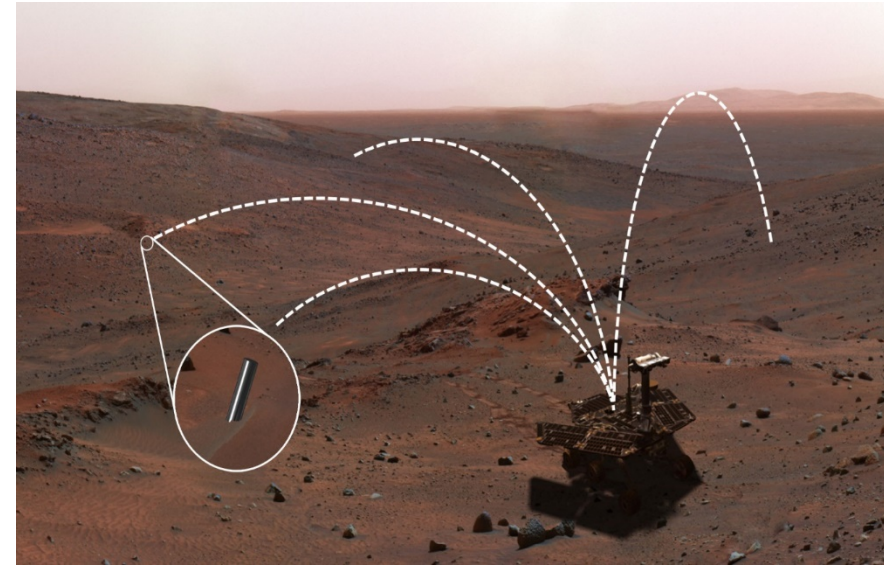
**Construction of Consumables
Depots for Crew & Power (O₂,
H₂)**

Mars Soil Assay by Bullet μ -Spectrometer

Astronaut with Micro-Spectrometer Gun



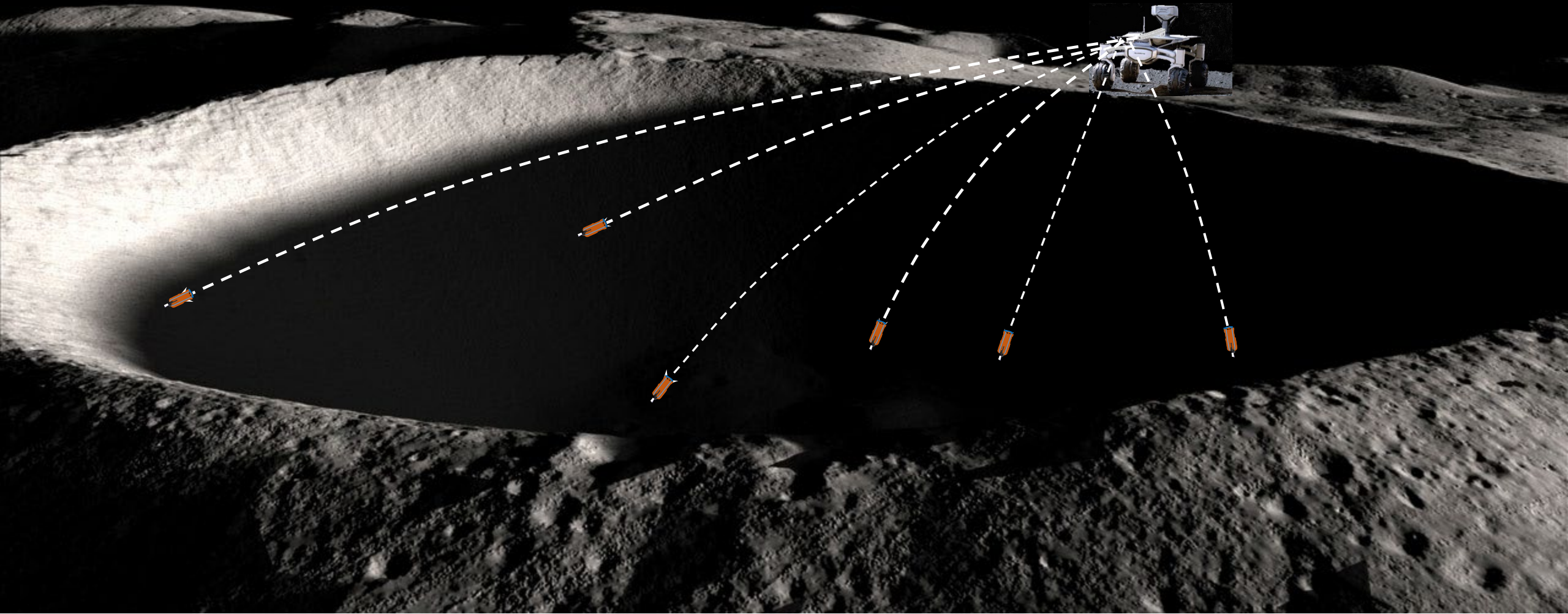
Rover with Micro-Spectrometer Gun



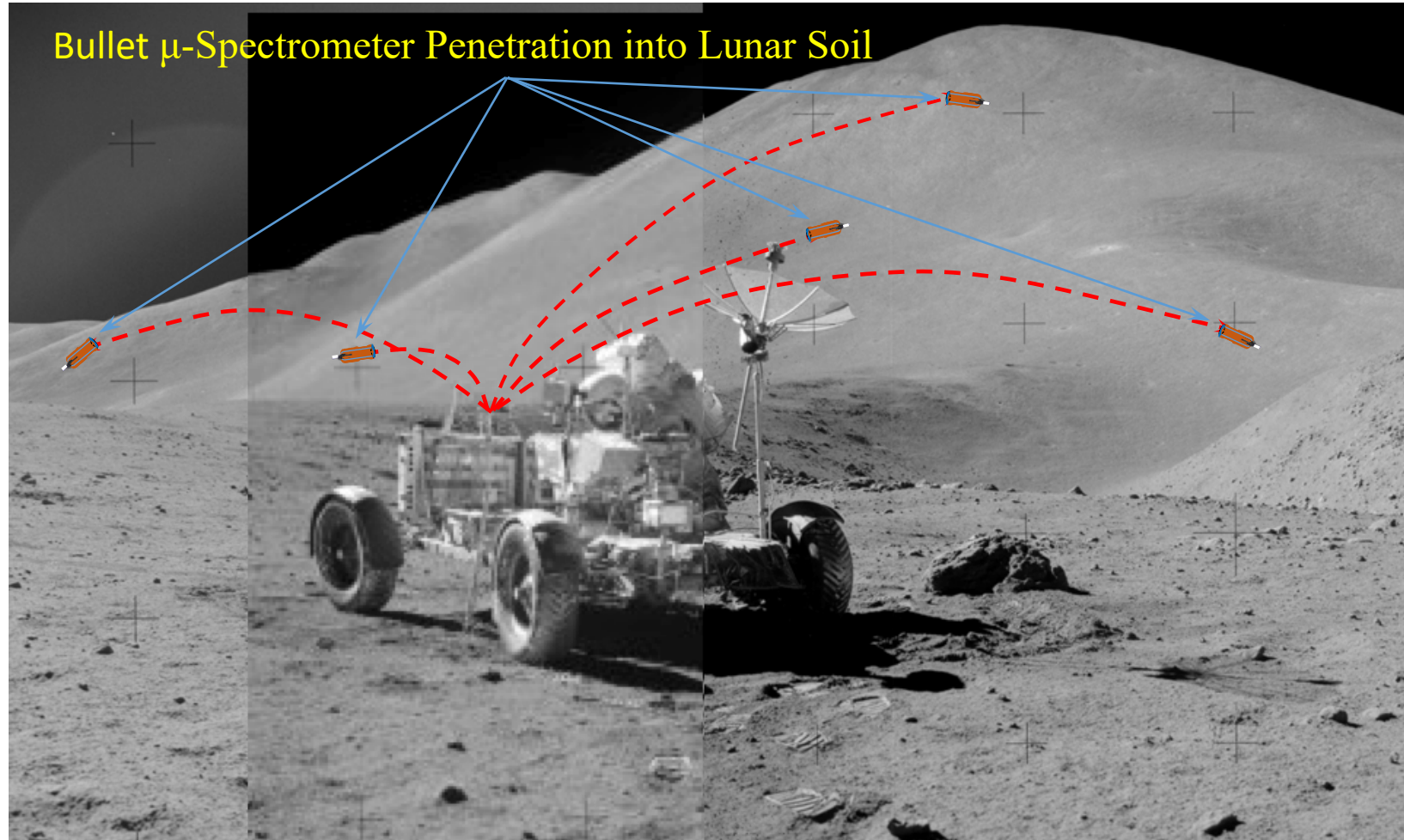
Bullet-like micro-spectrometers are shot out broadly to assay organic and inorganic components of Lunar/Mars/Asteroids soils.

Lunar Soil Assay by Bullet μ -Spectrometer

Penetration of Bullet μ -Spectrometer into Steep Hillside and Bottom of Crater

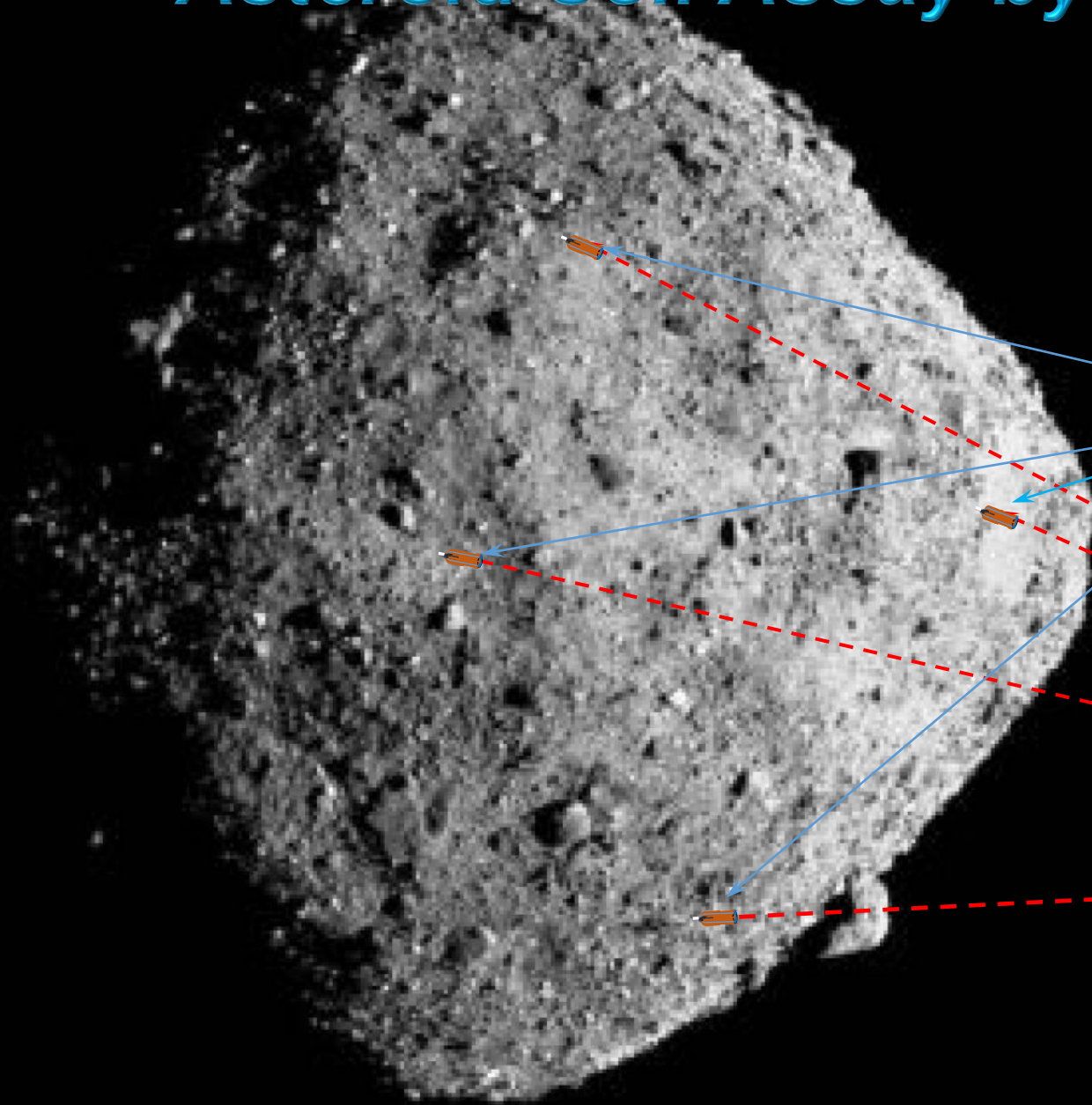


Lunar Soil Assay by Bullet μ -Spectrometer

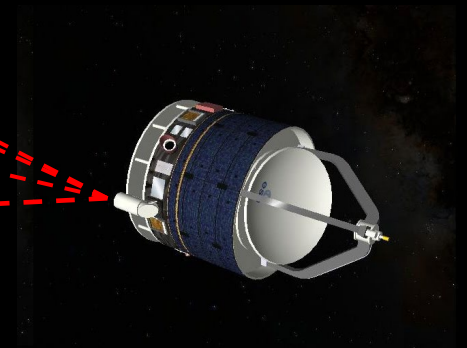


Images credit: NASA

Asteroid Soil Assay by Bullet μ -Spectrometer

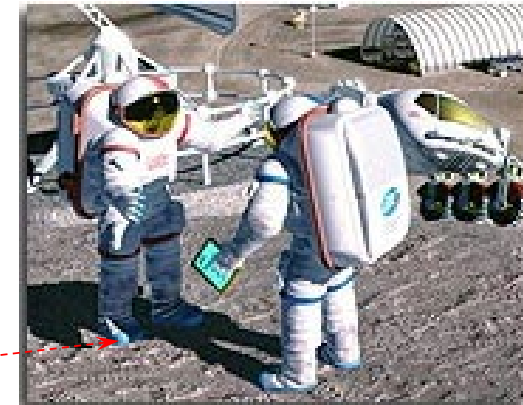
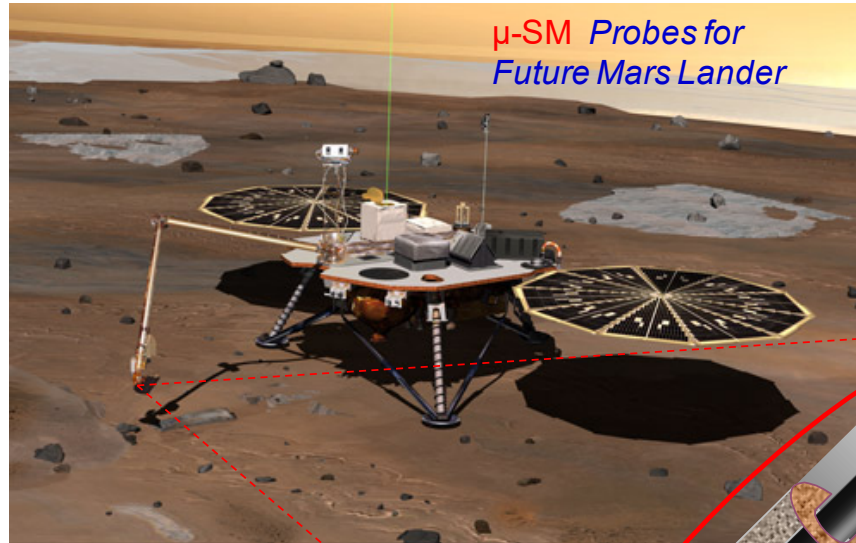


Penetration of Bullet μ -Spectrometer
into Asteroid

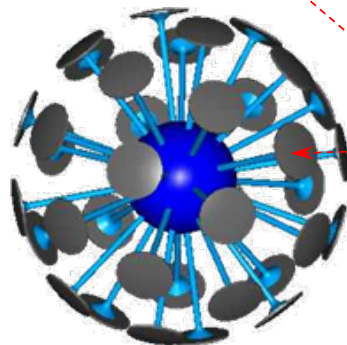


Images credit: NASA

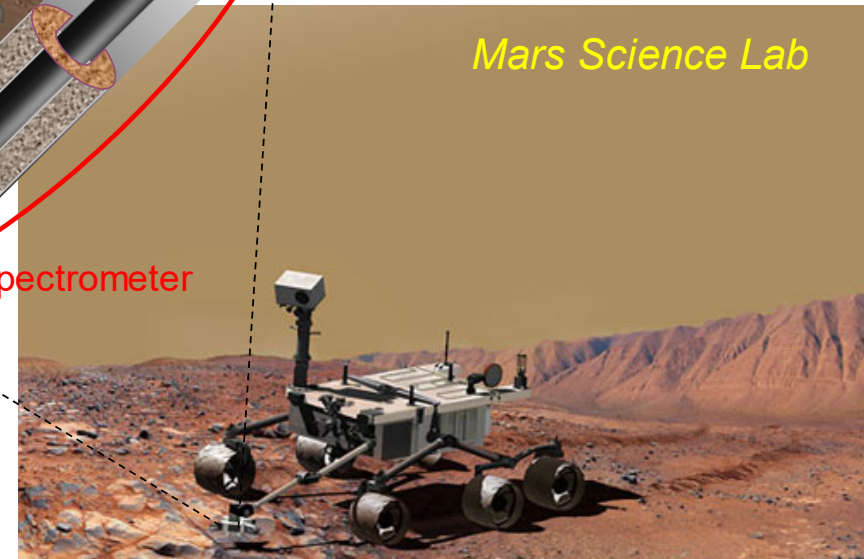
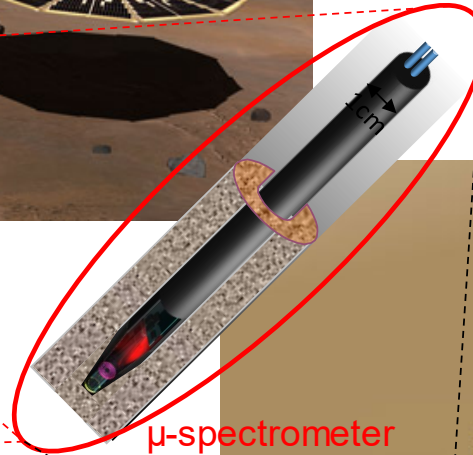
μ -Spectrometer for Lunar & Mars Exploration



μ -SM imbedded Astronaut Shoes
and Rover Tires



μ -SM imbedded Tumbleweed Rover



Concluding Remarks



- Applications of micro-spectrometer bullets include mineral mapping on the Moon and Mars, especially hillsides and deepened bottom floor of craters.
- Laboratory prototype was fully developed and tested for dopamine samples.
- Bullet type micro-spectrometer and deployment gun were designed for space applications.
- Based on the spectral data of the minerals of interest, line and pixel densities of the micro-spectrometer will be tailored for measurement accuracy and clarity.
- Device components, such as avalanche photo-diode, burst-mode LED, cylindrical super-capacitor, and telemetry device require further development for enhanced and versatile performance of bullet shape micro-spectrometers.