

1 **Assessment of ICESat-2 sea ice surface classification with Sentinel-2 imagery:**
2 **implications for freeboard and new estimates of lead and floe geometry**

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14 **Key points**

- 15 • Specular leads identified by ICESat-2 show strong agreement with leads observed in
16 coincident Sentinel-2 imagery.
- 17 • The coincident images provide further evidence of the misclassification of dark leads due
18 to clouds, which are no longer used to derive freeboard in ICESat-2 data.
- 19 • We apply two metrics for estimating lead fraction and chord length across the entire
20 Arctic and Southern Ocean.

21 **Abstract**

22 NASA's Ice, Cloud, and Land Elevation Satellite-2 (ICESat-2) mission launched in September
23 2018 and is now providing high-resolution surface elevation profiling across the entire globe,
24 including the sea ice cover of the Arctic and Southern Oceans. For sea ice applications,
25 successfully discriminating returns between sea ice and open water is key for accurately
26 determining freeboard (the extension of sea ice above local sea level) and new information
27 regarding the geometry of sea ice floes and leads. We take advantage of near-coincident optical
28 imagery obtained from the European Space Agency (ESA) Sentinel-2 (S-2) satellite over the

29 Western Weddell Sea of the Southern Ocean in March 2019 and the Lincoln Sea of the Arctic
30 Ocean in May 2019 to evaluate the surface classification scheme in the ICESat-2 ATL07 and
31 ATL10 sea ice products. We find a high level of agreement between the ATL07 (specular) lead
32 classification and visible leads in the S-2 imagery in these two coincident images across all six
33 ICESat-2 beams, increasing our confidence in the freeboard products and deriving new estimates
34 of the sea ice state. The S-2 overlays provide additional, albeit limited, evidence of the
35 misclassification of dark leads due to clouds. Dark leads are no longer used to derive sea surface
36 and thus freeboard as of the third release (r003) of the ICESat-2 sea ice products. We show
37 estimates of lead fraction and more preliminary estimates of chord length (a proxy for floe size)
38 using two metrics for classifying sea surface (lead) segments across both the Arctic and Southern
39 Ocean for the first winter season of data collection.

40 **1. Introduction**

41 NASA's Ice, Cloud, and Land Elevation Satellite-2 (ICESat-2) is a new satellite mission
42 providing unprecedented (high resolution and accuracy) profiling of the Earth's surface,
43 especially over the high-latitude polar regions. ICESat-2 was launched successfully from
44 Vandenberg Air Force Base on September 15th, 2018 and started acquiring science quality data
45 on October 14th, 2018. A description of the ICESat-2 mission design and objectives can be
46 found in Markus et al., (2017) and Neumann et al., (2019).

47 ICESat-2 consists of a bespoke photon-counting Advanced Topographic Laser Altimeter
48 System (ATLAS) - a low pulse-energy laser split into a novel six-beam configuration of three
49 beam pairs (a strong and a weak beam). At the nominal orbital altitude, ATLAS generates
50 individual laser footprints of ~11 m (in diameter) on the Earth's surface (Magruder et al., 2020).
51 The high pulse repetition rate of 10 kHz means laser pulses are separated by only 70 cm along-
52 track; a much higher resolution and sampling than NASA's previous laser altimetry mission
53 ICESat (Zwally et al., 2002). ICESat-2 samples year-round and has only experienced minimal
54 downtime since launch.

55 Various data products are now being generated routinely from ICESat-2 which are
56 publicly disseminated through the National Snow and Ice Data Center (NSIDC,
57 <https://nsidc.org/data/icesat-2>). The primary science data product from ICESat-2 is the Level 2A

58 ATL03 geolocated photon heights (Neumann et al., 2019) from which most of the higher-level
59 science products are derived. For sea ice users, the primary products of interest are the Level 3A
60 along-track sea ice height and type (ATL07, <https://nsidc.org/data/ATL07>) and freeboard
61 (ATL10, <https://nsidc.org/data/ATL10>) products (Kwok et al., 2020a). The first winter-season of
62 Arctic sea ice heights and freeboards (October 2018 to March 2019) were presented in Kwok, et
63 al., (2019a). Sea ice thickness has been estimated from these freeboards using external snow
64 loading estimates from the NASA Eulerian Snow on Sea Ice Model (NESOSIM) and modified
65 versions of the Warren climatology (Petty et al., 2020) and through the combination of total (ice
66 plus snow) freeboard from ICESat-2 with estimates of ice freeboard from the CryoSat-2 radar
67 altimeter to derive snow depth and freeboard concurrently (Kwok et al., 2020b). The surface
68 heights and freeboards were assessed against coincident airborne elevation data collected by
69 NASA's Operation IceBridge (Kwok et al., 2019b). The along-track heights generated from the
70 Airborne Topographic Mapper (ATM, Studinger, 2014) flown on OIB showed strong agreement
71 with ATL07-derived heights, while the derived freeboard from ATM showed only small (0-4
72 cm) differences with ATL10 freeboards, depending on the method used to derive and compare
73 freeboard. Coincident imagery obtained during the OIB surveys by the Continuous Airborne
74 Mapping By Optical Translator (CAMBOT) alluded to erroneous classification of 'dark leads' in
75 the presence of clouds (Kwok et al., 2020c). A dark lead classification is associated with a dark
76 surface and relatively low photon rates unlike the high photon rates associated with specular
77 leads (described more in Section 2.1). The dark lead segment classification is still included in the
78 latest (r003) ATL07 sea ice height and surface classification product release, but these have now
79 been excluded from the derivation of freeboard in the latest (r003) ATL10 product release (the
80 ATL07 and ATL10 products are described more in the following section). This change remains
81 in-place for the upcoming r004 data release. The Kwok et al., (2020c) study highlighted the
82 significant utility of coincident high-resolution imagery for better understanding the performance
83 of the ICESat-2 sea ice surface classification algorithm. The assessment, however, was hindered
84 by the high consolidation of the ice pack in the region profiled during the spring 2019 OIB
85 campaign, limiting the number of leads/open water segments and the quality of the freeboard
86 estimates due to the lack of reliable sea surface tie-points.

87 The main objective of this study is to further assess the performance of the surface
88 classification scheme in ATL07 for discriminating between sea ice and open water. We take

89 advantage of near-coincident optical imagery from the Sentinel-2 satellite mission to assess this
90 specific aspect of the ATL07 algorithm performance. Successfully discriminating returns
91 between sea ice and open water has benefits beyond deriving freeboard. Additional metrics for
92 understanding the sea ice state include concentration, lead fraction (or density) and floe size. In
93 the marginal seas, where waves can break up and fracture the ice (Horvat et al., 2020 and
94 references therein), ice floes are thought to be smaller and less concentrated/consolidated (a
95 higher lead fraction) than within the pack ice. The geometry of floes and leads is an important
96 control on the strength of sea ice and its thermodynamic response to forcing (Feltham, 2005;
97 Horvat et al., 2016), particularly in marginal, coastal, and seasonally-ice-covered seas. The floe
98 size distribution (Rothrock & Thorndike, 1984) is increasingly being introduced into sea ice
99 components of climate models (Bateson et al., 2020; Roach et al., 2018) , but only limited basin-
100 scale observational constraints exist to-date, e.g. estimates derived from ESA's CryoSat-2 radar
101 altimeter for the Arctic Ocean only (Horvat et al., 2019). ICESat-2 offers the exciting potential to
102 provide new observational estimates of the floe size distribution, benefiting from the small
103 footprint, high precision and along-track sampling rate across the multiple beams. Satellite tracks
104 (e.g. those from ICESat-2) profile ice floes across random/unknown cross sections of the floe.
105 The along-track cross section of a floe measured by satellite is commonly referred to as a chord
106 length. A collection of floe chord lengths can provide statistics of floe geometry, e.g. moments of
107 the floe size distribution and the open water fraction, under certain assumptions about the
108 underlying floe geometry (Horvat et al., 2019).

109 The coincident S-2 scenes provide the ideal means for assessing new lead fraction and
110 chord length estimates from ICESat-2 data, complementing the freeboard/thickness estimates
111 already being generated. We choose to discuss lead fraction instead of sea ice concentration as
112 concentration/extent was discussed in Horvat et al., (2020) and an assessment/comparison with
113 the commonly used passive microwave-derived concentrations (taking into account the
114 significant sampling differences) is beyond the scope of this study.

115 **2. Data**

116 **2.1 ATL07 sea ice heights and surface classification**

117 We primarily use data from the ICESat-2 ATL07 sea ice height and type product. A full
118 description of the ATL07 product can be found in the ATL07/ATL10 Algorithm Theoretical
119 Basis Document (ATBD, Kwok et al., 2020a) and the recent changes to this algorithm for
120 Release 003 (r003) are discussed more in Kwok et al., (2020c), so here we provide only a basic
121 overview of the methodology relevant to the results/discussion presented in this study.

122 ATL07 is generated by aggregating 150 photons from the ATL03 geolocated photon
123 product (Neumann et al., 2019) independently along each of the six beams. The beams are
124 arranged in ‘strong’ and ‘weak’ beam pairs with each beam pair separated by ~3.3 kilometers in
125 the across-track direction and the strong/weak beams separated by ~90 m across-track and ~2.5
126 km along-track. A surface finding routine (ATL07/10 ATBD r003, section 4.2) first windows the
127 photon heights around an expected sea surface then extracts a best-guess Gaussian height
128 distribution (convolved with the expected system response) to the photon height histogram to
129 determine 1) a single segment height, 2) an associated quality flag based on the goodness of fit
130 and 3) associated metrics including photon rate. The heights are expressed relative to the WGS84
131 ellipsoid with the mean sea surface (MSS) and various time-variable geophysical corrections
132 removed: ocean tides, solid earth tides, ocean loading, solid earth pole tides, inverted barometer.
133 Surface-classified height segments are produced for each of the six beams (three strong and weak
134 pairs) independently. The photon rates of the strong beam are roughly 4 times higher than those
135 of the weak beam, which results in mean segment lengths of ~15 m for the strong beam and ~60
136 m for the weak beam (Kwok et al., 2019a). Adding the individual laser footprint size of ~11 m to
137 the segment length provides an estimate of the spatial resolution of the segments (i.e. a mean of
138 ~26 m x 11 m for the strong beam and 71 m x 11 m for the weak beam).

139 An empirically based decision-tree algorithm is used to discriminate the returns between
140 sea ice and open water (Kwok et al., 2016). The empirical thresholds were determined
141 principally by 1) data collected in campaigns prior to the launch of ICESat-2 by the Multiple
142 Altimeter Beam Experimental Lidar (MABEL) - a test-bed instrument for ICESat-2 (McGill et
143 al., 2013); and 2) post-launch evaluation of the ATLAS performance. The three inputs to this
144 decision tree (and the physical justifications) are as follows:

- 145 1. Photon rate (photon returns per laser pulse, the apparent reflectivity of the surface).
 146 Low photon rates are associated with water or very thin ice in open leads, which are
 147 generally darker than sea ice. However, specular and quasi-specular returns have
 148 been observed from smooth open water/thin ice surfaces in both ICESat and MABEL.
 149 2. Width of the Gaussian fit to the photon height distribution (the small-scale surface
 150 roughness). Used to partition the surface types into smooth/rough categories.
 151 3. Background rate (deviations compared to the photon rate indicate shadows, specular
 152 reflections and/or atmospheric contamination).

153 The result of the decision tree determines the radiometric surface type which includes the
 154 following classifications (denoted by the *height_segment_type* flag): cloud (0), snow covered,
 155 gray or rough ice (1), specular leads (2-5), smooth dark leads (6-7) and rough dark leads (8-9).
 156 This is considered the winter-time decision tree. As discussed in the introduction, in the latest
 157 ATL10 data release (r003), dark lead segments (*height_segment_type* = 6 to 9) have been
 158 dropped from the derivation of the local reference sea surface height (as detailed below) due to
 159 possible issues with cloud attenuation, so now only specular leads (*height_segment_type* = 2 to
 160 5) are used for deriving the reference sea surface height and freeboard (Kwok et al., 2020c).

161 Due to the expected uncertainty in this radiometric surface classification approach (Kwok
 162 et al., 2016), a further filtering is applied based on the local height distribution:

- 163 4. Local height filter, based on the distribution of local (10 km) smooth height segments
 164 (h_{sm}) – segments with a Gaussian width < 0.13 m. Specifically the lead segment
 165 height must be between the minimum of the smooth heights (h_{sm_min}) and the
 166 maximum of either the 2nd percentile of the smooth heights or $h_{sm_min} + 2 \sigma$ where $\sigma =$
 167 2-3 cm, the expected uncertainty in surface height over smooth surfaces.

168 The result of the radiometric decision tree and local height filter sets the sea surface
 169 height flag (*ssh_flag*, 0 = sea ice, 1 = sea surface). Sea surface segments are considered
 170 candidate lead segments for deriving the reference sea surface height and thus freeboard in
 171 ATL10. Note that the summertime (non-winter) decision tree simply extends the classification of
 172 ice, specular leads and dark leads as potential melt ponds (not a feature of this analysis).

173 Freeboards are calculated in ATL10 in 10 km along-track sections based on a reference
 174 sea surface derived from the available lead/sea surface segment heights (ATL07/10 ATBD r003,

175 section 5.1). Briefly, consecutive sea surface segments are grouped together into individual
176 ‘leads’ (to reduce noise in the individual sea surface estimates) before a single reference sea
177 surface estimate is produced for consecutive 10 km along-track sections that include at least one
178 sea surface segment, for each beam independently. A basic interpolation/end-filling procedure is
179 used to extend the reference sea surface estimates into regions where a sea surface segment does
180 not exist. A filtering is also applied to ensure that the 10 km sections are at least 25 km away
181 from the coast and have a concentration (from passive microwave data) $> 50\%$ - i.e. within the
182 pack ice and away from regions thought to be more affected by waves. The 10 km reference sea
183 surface heights must also lie within a set height window relative to the MSS (± 0.5 m for the
184 Arctic and ± 1 m for the Southern Ocean) and differences between consecutive sea surface
185 heights must be relatively small (see the ATL07/10 ATBD r003, section 5 for more details).
186 After the filtering, *candidate lead* segments (i.e. segments with *ssh_flag* = 1) used to generate a
187 reference sea surface that also pass the above filters are set as *lead* segments (*ssh_flag* = 2) to
188 make clear they have indeed been used to derive the reference sea surface height and thus
189 freeboard in ATL10. Freeboard segments are then derived by differencing the sea ice segment
190 heights from the local 10 km reference sea surface height. Negative freeboards are set to zero.
191 More detail about the methodology is given in Kwok et al., (2020a) and Kwok et al., (2020c).

192 **2.2 Sentinel-2 imagery**

193 Sentinel-2 is a constellation of two twin satellites, Sentinel-2A (S-2A) and Sentinel-2B
194 (S-2B), operated by the European Space Agency (ESA) and launched in June 2015 and March
195 2017, respectively. The satellites host the MultiSpectral Instrument (MSI), which provides 13
196 reflective-wavelength bands in the wavelength region between 443 nm and 2202 nm (visible,
197 near-infrared, short-wave infrared). Depending on the band, spatial resolution varies between 10,
198 20, and 60 m, with the images covering an area of ~ 110 km x 110 km. For a given area, the shift
199 by 180 degrees between the two sun-synchronous polar orbits and a 290 km wide swath
200 guarantee a revisit time of 5 days at the equator, which improves to one image per day at higher
201 latitudes due to overlapping swaths with different viewing angles. Systematic global coverage of
202 land surfaces and coastal waters (up to 20 km from the coast) is available between 84 °N and 56
203 °S. Additional imagery is collected near and over coastal regions of Antarctica, Baffin Bay and

204 the North Atlantic ([https://sentinel.esa.int/web/sentinel/user-guides/sentinel-2-msi/revisit-](https://sentinel.esa.int/web/sentinel/user-guides/sentinel-2-msi/revisit-coverage)
205 coverage).

206 We searched for coincident S-2 and ICESat-2 data over sea-ice covered portions of the
207 Arctic and Southern oceans. Our search algorithm matched the footprint of all S-2 images with a
208 nominal cloud coverage of less than 10% acquired during fall, winter, and spring months (fall
209 2018 to spring 2019) within the search regions (Arctic: Sep-May / Antarctic: Mar-Nov) with the
210 ICESat-2 reference ground track (RGT) times and locations (data from [https://icesat-](https://icesat-2.gsfc.nasa.gov/science/specs)
211 [2.gsfc.nasa.gov/science/specs](https://icesat-2.gsfc.nasa.gov/science/specs)). No images for winter months (e.g. December/January in the
212 Northern Hemisphere) are available due to the lack of sun illumination. We searched for
213 overlapping data pairs with an acquisition time difference of less than 2 hours to mitigate for ice
214 drift, resulting in 14 candidate S-2 images. This imagery subset was then checked for image
215 quality, presence/absence of sea ice, actual cloud cover, and for the extent of the overlap with
216 ICESat-2 data. Two optimal coincident images were found, 1) over the Lincoln Sea of the Arctic
217 Ocean in May 2019 (i.e. the end of winter, time difference of 94 minutes) and 2) over the
218 Western Weddell Sea of the Southern Ocean in March 2019 (i.e. the start of the austral winter,
219 time difference of only 7 minutes), as shown in Figure 1. This Arctic Ocean scene occurs late in
220 the winter/spring season, after typical dates of melt onset in the Arctic. However, the location is
221 in the coldest/thickest ice of the Arctic and no surface melt is visible in the imagery. Visual
222 inspection of these scenes showed clear coincidence between the imagery and the ICESat-2 data
223 with no obvious issues of drift misalignment (no correction was applied). Half of the Southern
224 Ocean image is contaminated by clouds, as is clearly visible in Figure 1. We thus split these into
225 ~50 km along-track ‘coincident profiles’ for visualization purposes (two Arctic profiles and one
226 Southern Ocean profile). 50 km was also a good balance between data coverage and ease of
227 interpretation/visualization.

228 Note that manual drift alignment/correction was attempted for some of the other
229 candidate images but the lack of notable features (e.g., leads and ridges) hindered our efforts.
230 Manual drift correction was successfully applied in generating coincident summer S-2/ICESat-2
231 scenes, as shown in Tilling et al., (2020).

232 3. Methods

233 3.1 Classification evaluation

234 To compare the IS-2 ATL07/10 data with the S-2 imagery we use a simple nearest
 235 neighbor interpolation scheme (the nearest geocoded pixel of the imagery to the given ATL07/10
 236 beam segment) to produce a coincident profile of S-2 surface reflectance. In particular, we use
 237 the red band, which is available at the highest 10 m spatial resolution and convert the Digital
 238 Number (DN) provided in the Level-2A data products to surface reflectance by dividing by the
 239 quantification value (DN/10,000). We compare these profiles qualitatively as the lack of perfect
 240 time coincidence with the imagery, together with the possible impact of ice drift and the
 241 contrasting resolutions of the data, make it challenging to carry out a more robust quantitative
 242 assessment.

243 3.2 Lead fraction

244 To estimate lead fraction, we take the segment-length-weighted ratio of sea surface
 245 segments to the total length of valid segments in 10 km along-track sections across the combined
 246 three strong beams (we combine data from the 3 beams before sectioning the data). Note that the
 247 IS-2 measurements include three beam pairs spread over ~6.6 km across-track, so for this to be
 248 treated as a two-dimensional lead fraction estimate one needs to assume that the underlying ice is
 249 isotropic and homogenous over this 10 km x 6.6 km window, something we plan to test more in
 250 future work. The along-track lead fraction is calculated as

$$251 \quad L_f = \frac{\sum_{N_{sl}} l_s^i}{\sum_{N_{all}} l_s^i}, \quad (1)$$

252 where l_s is the segment length, N_{all} is the total number of ice and lead segments in the given
 253 section (10 km along-track in this study), N_{sl} is the total number of lead segments (specular only
 254 in r003) and i is the height segment index.

255 The ATL07/10 sea surface classification approach is designed to be strict – i.e. possible
 256 leads are potentially thrown away by the height percentile filter (step 4 in Section 2.1) to reduce
 257 the likelihood of ice segments being erroneously used to derive a reference sea surface. This
 258 approach makes sense when one considers the high sensitivity of the freeboard estimate to errors

259 in the reference sea surface height calculation. However, for ice floe/lead geometry analyses,
260 such a strict filter could result in an underestimate of lead fraction (and an overestimate of chord
261 length, discussed in the following section). We therefore adopt two approaches for deriving lead
262 fraction: i) L_f^{v1} where we use the ATL07 sea surface flag for candidate leads ($ssh_flag = 1$) and
263 ii) L_f^{v2} where we use the ATL07 specular lead classification ($2 \leq height_segment_type \leq 5$)
264 together with a less strict 10 km local height percentile filter (20% instead of 2%, see step 4 in
265 Section 2.1) to determine the sea surface segments. In summary L_f^{v1} is derived using the
266 ssh_flag from the product, whereas L_f^{v2} is derived using a higher 20% local height filter. This is
267 a relatively crude way of exploring the sensitivity of these lead estimates to the underlying
268 classification algorithm which we will aim to expand on in future work.

269 **3.3 Chord length**

270 To calculate chord length, we devised a simple algorithm that splits the along-track
271 segment data into floe (or chord) groups based on the sea surface classification flag. The sea
272 surface segments are discarded. Similar to lead fraction, we use two approaches here; C_l^{v1}
273 utilizes the sea surface flag (ssh_flag) and C_l^{v2} uses the sea surface classification derived in this
274 study using the 20% local height filter. In both cases a ‘floe’ group needs to include at least 3
275 height segments and have a maximum spacing between consecutive segments less than 300 m.
276 300 m is the upper end of the tail of the distribution of segment lengths for the strong beams
277 (Kwok et al., 2020a) meaning a gap between segments greater than this value is very likely to be
278 driven by the presence of anomalously high segment lengths (which are generally more
279 uncertain) or a data gap. Data gaps between segments are caused primarily by atmospheric
280 scattering (e.g. by clouds), which could result in erroneously high chord lengths, especially as
281 clouds can form preferentially over leads (although this is still being debated, e.g. Li et al.,
282 2020). We also discard any floe groups that are smaller than 60 m, which we take to be the
283 minimum resolved chord length due to considerations of footprint size and segment length, and
284 groups that are longer than 50 km. We add 11 m, the approximate photon footprint resolution, to
285 all groups to finally derive estimates of chord length. We run this analysis for each of the three
286 strong beams independently.

287 Note that we do not provide estimates of lead length in this study as estimates were provided
288 in Kwok et al., (2019a) and this is thought to depend more on the resolution of the ATL07 data,
289 so investigations focused on the underlying ATL07 algorithm may be needed to produce more
290 reliable lead length estimates in future work.

291 For both the lead fraction and chord length estimates, we bin the along-track data to a 25 km
292 x 25 km polar stereographic grid (EPSG:3411 for the Arctic and EPSG:3412 for the Southern
293 Ocean) using data collected from the first Arctic (November 1, 2018 to April 30, 2019) and
294 austral (May 1, 2019 to September 30, 2019) winters.

295 **4. Results**

296 **4.1 Comparisons with Sentinel-2 imagery**

297 The S-2/ATL07 comparison for the ~50 km profile (P1) within the western Weddell Sea
298 on March 17th, 2019 for strong beam 1 (*gt11* in this backward spacecraft orientation) is shown in
299 Figure 2. This cropped S-2 image (top panel) depicts a scene of mixed ice surfaces - large
300 consolidated floes, small broken up floes, and the occasional lead opening. It also features very
301 flat/thin ice for the first ~20 km and thicker/rougher ice (relative heights ~2/2.5 m higher) for the
302 remaining 30 km along-track. The final ~5 km of the scene features clouds, as can be seen more
303 clearly in the zoomed-out Figure 1. The ATL07 radiometric surface classification scheme
304 detected 218 specular lead segments and 19 dark lead segments for this scene, with 196
305 candidate (and indeed utilized) sea surface segments. The ATL07/10 lead classification shows
306 strong agreement with the S-2 imagery. As most of the lead segments are specular, the photon
307 rate shows corresponding spikes in these same locations, along with drops in the background rate
308 (as expected from background photons scattering away from the detector over specular surfaces).
309 The scene highlights the ability of ICESat-2 to detect both narrow (10s of meters) and wide
310 (100s of meters) openings in the ice cover. Dark lead classified segments are produced at ~8 km
311 and especially at ~45-50 km along-track, with no leads visible in the imagery in both locations
312 (although in the latter it is harder to assess due to the clouds). The presence of clouds in the last
313 ~5 km corresponds with a clear attenuation in photon rate around this part of the scene, similar to
314 the Operation IceBridge comparisons given in Kwok et al., (2020c). The ~8 km along-track dark
315 lead is harder to diagnose, as this is associated with a sharp decline in photon rate, as one would

316 expect from a lead, but no obvious lead or cloud is visible in the imagery. These comparisons
317 benefit from the very small time difference (~ 7 minutes) between the ICESat-2 data and S-2
318 image acquisition. The derived lead fractions and chord lengths in this scene are also shown in
319 Figure 2 (panel 6). The lead fractions are estimated as 1.38% (v_1) and 1.54% (v_2), while the
320 mean chord lengths are estimated as 2.96 km (v_1) and 2.49 km (v_2). The remaining profiles using
321 the strong and weak beams are given in the Supplementary Information (Figures S1 to S5).
322 Summary statistics (lead classification counts, lead fraction and chord lengths) are provided in
323 Table 1. No obvious degradation of the classification performance is visible in the weak beam
324 profiles, although ICESat-2 generally detects fewer leads and lower lead fractions than with the
325 strong beams.

326 The S-2/ATL07 comparison for the northernmost ~ 50 km profile (P2) within the Lincoln
327 Sea, Arctic Ocean on May 25th, 2019 for the middle strong beam (*gt2l* in this backward
328 spacecraft orientation) is shown in Figure 3. The S-2 image depicts a scene of large consolidated
329 floes, small broken up floes/leads, but also two much larger (> 1 km) lead openings. No clouds
330 are visible in the imagery (or implied by attenuations in the photon rate). The heights extend
331 from just over 0 m to ~ 2 -2.5 m. The ATL07 radiometric surface classification scheme detected
332 416 specular leads and 2 dark leads for this scene, with 220 candidate (and again utilized) sea
333 surface segments. The ATL07/10 lead classification again shows strong agreement with the S-2
334 imagery. It is encouraging to note that the refrozen lead of low height/freeboard at ~ 10 km
335 along-track is not classified as a specular or dark lead. Only a slight decline photon rate is
336 observed here. At ~ 15 -20 km along-track, the S-2 scene shows some small but highly
337 consolidated ice floes, with some leads detected, highlighting the challenge of lead/floe detection
338 in the more consolidated ice regimes. The lead fractions for this scene are estimated as 2.55%
339 (v_1) and 6.37% (v_2), while the chord lengths are estimated as 1.82 km (v_1) and 1.99 km (v_2). The
340 difference between the v_1 and v_2 estimates are higher, and we can see this is mainly driven by the
341 inclusion (or absence) of the segments within the large lead openings. Including the extra sea
342 surface segments by using a higher (20%) height threshold in v_2 increases lead fractions, as
343 expected, however also results in the perhaps more counter intuitive result of increasing the
344 mean chord length as the ice floe groupings remain too small to be classified as a floe and are
345 simply discarded. The profiles of this image for the remaining strong and weak beams are given
346 in the Supplementary Information (Figures S6 to S10) and summarized in Table 1. We also

347 include in the Supplementary Information the profiles from P3 (Figures S11 to S16). We did not
348 include these in the main manuscript as they generally featured fewer leads compared to the P2
349 comparisons. These additional beam comparisons show again a remarkably high level of
350 agreement with the S-2 imagery in terms of the lead classification but some potentially missing
351 lead classifications within narrower cracks/leads suggested by the S-2 image. The summary
352 statistics are included in Table 1. Also noteworthy is that while the S-2 image in Figure 1 shows
353 clouds around the northwestern section of P3, no indication of clouds can be seen in the ICESat-
354 2 data (no photon rate attenuation), which is likely due to the higher time difference between the
355 ICESat-2 data and S-2 image acquisition here (~94 minutes).

356 In Figure 4 we zoom in further on the P2 profile to highlight the ICESat-2 performance
357 across this second large lead opening (at ~41 - 43 km along-track in Figure 3). The S-2 image,
358 despite being coarse when looking at this scale, suggests the presence of young gray ice within
359 this lead which is categorized as a dark lead in 2 of the ATL07 segments (note the drops in
360 photon rate compared to the increase in photon rate in the specular lead classifications). This
361 appears to be an accurate classification (they are of similar height to the surrounding specular
362 leads, so could be used to derive the reference sea surface) and highlights the potential utility of
363 the dark lead classification in ATL07/10. Another thing of note here is that due to the strict local
364 height filter in the ATL07 algorithm (Step 4 in Section 2.1) a significant fraction of these
365 segments are not classified as *candidate lead* segments (*ssh_flag* = 1). This results in some
366 erroneous ice segments and small ‘chords’ being included within the lead, biasing our statistics.
367 These erroneous chords are removed when we move to the higher height percentile filter (v_2),
368 resulting in an increase in chord length (0.74 km to 0.91 km) and a doubling in the lead fraction
369 (9.12% to 19.17%), albeit in this highly localized, lead profile.

370 4.2 Basin scale assessments

371 To better understand the sea ice classification and to provide context for the basin-scale
372 lead fraction/chord length estimates shown next, we first show basin-scale maps of key lead
373 classification metrics from ATL07 across the Arctic and Southern Ocean for the winter study
374 periods (Arctic: November 1, 2018 to April 30, 2019; Southern Ocean: May 1 2019 to
375 September 30, 2019). Figure 5 shows the radiometric lead fraction compared to all (segment
376 length weighted) segments, the ratio of radiometric specular lead segments to dark lead

377 segments, the fraction of specular lead segments that become sea surface segments ($ssh_flag \geq 1$)
 378 and the fraction of sea surface (lead) segments compared to all segments – i.e. the lead fraction.
 379 The radiometric lead and sea surface lead segments generally follow the spatial pattern expected
 380 from our past knowledge of the sea ice state, increases in leads (declines in concentration)
 381 towards the ice edge. In both hemispheres dark leads make up a significant (~50%) fraction of
 382 the total number of radiometric leads. In both hemispheres a significant (>70%) fraction of the
 383 specular leads pass the height filter and are assigned as sea surface segments. There is a clear
 384 increase in the fraction of discarded specular leads along the ice edge in both hemispheres, but
 385 also within the Canadian Arctic Archipelago. As ATL10 applies a stricter 50% concentration
 386 filter from passive microwave data to avoid wave contamination (ATL07 uses 15%), and a 25
 387 km coastal mask, many of these regions adjacent to the ice edge (in both hemispheres) and near
 388 coastlines will not be processed into freeboard. Sea ice concentrations from the monthly (final
 389 and near-real time) NSIDC Climate Data Record (CDR, Meier et al., 2017; Meier et al., 2017)
 390 averaged over this same time period across both hemispheres are given in Figure 6 for context.

391 Figure 7 shows the 25 km x 25 km gridded lead fractions calculated using the ssh_flag
 392 (L_f^{v1} , as in Figure 5d), and also using the 20% height filter (L_f^{v2}). L_f^{v2} consistently results in
 393 higher lead fractions as expected, especially around the ice edge. We also note the increased lead
 394 fractions in areas of known polynya formation - the North Water Polynya to the northwest of
 395 Greenland and, to a lesser degree, Terra Nova Bay Polynya in the northern Ross Sea. Neither are
 396 as visible in the passive microwave-derived concentrations (Figure 6). Mean Arctic Ocean lead
 397 fraction are given as 1.06 % (L_f^{v1}) and 1.80 % (L_f^{v2}). Mean Southern Ocean lead fractions are
 398 given as 0.66% (L_f^{v1}) and 1.24% (L_f^{v2}). The Southern Ocean lead fractions, while lower on
 399 average, show increased lead fractions within the pack ice (compare with sea ice concentration in
 400 Figure 6), especially in the eastern Southern Ocean. This is expected considering the open ocean
 401 boundaries and the general divergent (away from the coast) Antarctic ice drifts. The results also
 402 highlight the sensitivity of these estimates to this one aspect of the ATL07 lead finding
 403 algorithm. Lead fractions within the pack ice of the Central Arctic are very low (< 1%) which we
 404 discuss more in the discussion/summary section.

405 Figure 8 shows the 25 km gridded chord lengths calculated using the ssh_flag (C_l^{v1}) and
 406 the 20% height filter (C_l^{v2}). Mean Arctic Ocean chord lengths are given as 5.56 km (C_l^{v1}) and

407 5.23 km (C_l^{v1}). Mean Southern Ocean chord lengths are given as 5.43 km (C_l^{v1}) and 4.55 km
408 (C_l^{v2}). As in the lead fractions, there is high spatial variability not captured by these simple
409 hemispheric means. Lower chord lengths are produced within the eastern Arctic Ocean
410 (compared to the western Arctic Ocean) and also in the eastern Southern Ocean (compared to the
411 western Southern Ocean). The spatial variability is generally higher than the lead fraction as only
412 small changes in lead fraction can have a large impact on the derived chord lengths. The Arctic
413 shows a sharp increasing gradient within the more central Arctic region. Chord lengths are high
414 (consistently greater than 10 km) in the central Arctic, which we discuss more in the following
415 section.

416 Figure 9 shows the probability distribution of the individual chord lengths for the Arctic
417 and Southern Ocean using the C_l^{v1} and C_l^{v2} chord length estimates. The data are plotted on a
418 log-log scale to highlight the power law nature of the underlying distributions. The consistency
419 of the distributions across hemispheres and lead finding algorithm ($v1$ and $v2$) is encouraging.
420 The distributions become more variable at the tail of the distribution (> 10 km) due to the lower
421 sampling rate of these higher chord lengths. While log-log distributions can help visually
422 highlight power-law like behavior in empirical data, more robust statistical tests are needed to
423 truly test for whether data is well characterized by a power-law, especially due to the issues with
424 the less well-observed tails of the distribution (Clauset et al., 2009; Stern et al., 2018, Horvat et
425 al., 2019). We plan to explore this in future work as we compare against the chord length
426 estimates from other observed data, as discussed below.

427 **5. Discussion and summary**

428 The coincident Sentinel-2 (S-2)/ICESat-2 scenes provide crucial validation of the
429 ICESat-2 sea ice classification procedure, with the caveat that these represent only a small
430 fraction of the available sea ice data produced from ICESat-2 to-date. The specular lead
431 classification in the sea ice products shows strong agreement with the imagery across all beams,
432 while the reliability of the smaller quantity of dark leads found in these profiles was more
433 questionable. The Southern Ocean profile provided further evidence to that presented in Kwok et
434 al., (2020c) of their more uncertain reliability in the presence of clouds. Although the Arctic
435 Ocean profile showed 2 dark leads that appeared to be accurately classified (young, dark gray ice
436 in the imagery). More coincident S-2 scenes across different sea ice regimes (time and space)

437 would help provide further insight into the potential utility of the dark leads, especially as they
438 make up a significant fraction (~50-60%) of the total number of radiometric leads detected by the
439 current ATL07 algorithm. As discussed in Kwok et al., (2020c), the exclusion of dark leads from
440 the freeboard calculation has the downside of reducing coverage, especially in some of the more
441 consolidated sea ice regimes, providing motivation for continued study to ensure that this
442 exclusion is not overly restrictive.

443 It is interesting to note that the ATL07 segments over the relatively wide (> 1 km) leads
444 in our Arctic Ocean S-2 profile (Figure 3) are generally classified as specular, except for the 2
445 dark lead segments over regions of young, dark grey ice. Intuitively we might expect this all to
446 be considered a ‘dark lead’ as the larger opening increases fetch and the potential for surface
447 roughening. However, ICESat-2 segments are classified mainly as specular, until the photon rate
448 drops drastically (to less than that over the snow-covered ice) over the small region of young
449 gray ice. Similarly, the very thin cracks shown in the imagery (on the order of meters to tens of
450 meters, i.e. the segment resolution of the ATL07 data) were often classified as ice. It is unclear
451 from the S-2 imagery (10 m resolution) the extent to which these leads/cracks have refrozen and
452 thus how confident we can be in prescribing these segments as misclassified. Higher resolution
453 imagery (e.g. from Digital Globe’s Worldview imagery) could provide useful further insight here
454 and will be explored in future work.

455 The local height filter included in the lead (sea surface) finding algorithm further limits
456 the number of leads identified. This appears reasonable when the goal is accurately determining
457 a reference sea level for freeboard but less ideal for deriving sea ice state information, e.g. lead
458 fraction and chord length. This study presented a crude approach to relax this height filter
459 (increasing the percentile threshold from 2% to 20%), which is worthy of further investigation as
460 we seek to increase the utility of the ICESat-2 sea ice data.

461 Validation, or at least comparisons, of our basin-scale ICESat-2 estimates of lead fraction
462 and chord length with existing observational estimates are still needed. Basin-scale lead fraction
463 estimates have been produced from various satellite sensors, e.g. NASA’s Moderate Resolution
464 Imaging Spectroradiometer (MODIS) (Hoffman et al., 2019; Willmes & Heinemann, 2016) and
465 Advanced Microwave Scanning Radiometer (AMSR-E) (Röhrs & Kaleschke, 2012), and also
466 ESA’S CryoSat-2 and Envisat radar altimeters (Röhrs & Kaleschke, 2012; Tilling et al., 2019).

467 Chord length distributions have also been produced from a compilation of more sporadic
468 satellite/airborne imagery estimates (Stern et al., 2018). However, comparisons with these data
469 are hindered by the considerable differences in spatial/temporal sampling, and also the different
470 interpretations and definitions of leads across sensors and algorithms. As discussed earlier, the
471 lead classification algorithm in ATL07 appears to require clear sections of open water to trigger
472 a specular lead classification. Therefore, newly refrozen leads, which may be defined as leads in
473 other (generally coarser resolution) products, may be simply classified as ice. The radar
474 altimeter-derived estimates of lead fraction and chord length from CryoSat-2 (Tilling et al., 2019,
475 Horvat et al., 2019) provide arguably the most similar dataset for comparison/assessment. The
476 CryoSat-2 lead fractions and chord lengths shown in Tilling et al., (2019) appear consistently
477 higher and lower, respectively, than the ICESat-2 results presented here. For example, Tilling et
478 al., (2019, Figure 1) shows the percentage of CryoSat-2 waveforms classified as lead (a quantity
479 related to, but not the same as lead fraction) as $\sim 10\%$ in the Central Arctic but up to $\sim 60\%$ in
480 the more peripheral seas, while the chord lengths (Tilling et al., 2019, Figure 2) are ~ 3 km in the
481 Central Arctic and ~ 1 km in the peripheral seas. In contrast, the chord lengths presented in this
482 study are consistently > 10 km within the Central Arctic. There is better agreement on the
483 general spatial pattern of increasing lead fractions/decreasing chord lengths in the peripheral
484 seas, however. Radar altimeters such as CryoSat-2 are highly sensitive to the presence of
485 specular leads within the radar swath and have the added benefit of being largely unaffected by
486 clouds. However, the CryoSat-2 footprint is much larger (a pulse limited footprint of ~ 400 m
487 along track \times 1.65 km across track, however off-nadir leads can still dominate returns from
488 within the larger beam-limited across track footprint of 15 km) than ICESat-2 (mean of ~ 26 m \times
489 11 m for a given ATL07 strong beam segment). Future work will aim to better assess, and
490 hopefully reconcile, the lead fraction and chord length estimates from these two missions, taking
491 these significant sampling differences into account. We also hope to explore further statistical
492 testing of the power-law hypothesis for the floe length distribution, which was recently
493 challenged in Horvat et al., (2019) using CryoSat-2 chord length estimates. This reconciliation is
494 further motivated by the need to provide reliable observational constraints on the floe size
495 distributions being incorporated into sea ice components of Global Climate Models (Horvat, et
496 al., 2019). Airborne data from NASA's Operation IceBridge campaigns, which combine high-
497 resolution imagery and laser altimetry, should prove invaluable.

498 Our hope is that ICESat-2 sea ice data can be used to provide routine, and reliable, basin-
499 scale measurements of lead fraction, ice concentration and chord length estimates, in addition to
500 its primary mission requirement of delivering accurate estimates of sea ice height and freeboard.
501 Additional ICESat-2 sea ice algorithm testing and development is needed to further improve the
502 classification accuracy, which can be guided by additional comparisons with imagery (satellite
503 and airborne) and other airborne and satellite datasets. The ICESat-2 lead classification
504 algorithm utilizes fixed empirical thresholds (discussed in Section 2.1) which can easily be
505 tuned/calibrated as needed, as can other elements of the algorithm including the height percentile
506 threshold (which we simply explored in this study) and the photon aggregations. These
507 thresholds also need to be better explored in terms of appropriate strong/weak beam differences.
508 Efforts are also ongoing to provide further insight into the dark leads and their potential
509 reintroduction by utilizing a new cloud filter in the sea ice algorithm (Kwok et al., 2020c).
510 However, this assumes that the dark lead classifications unaffected by clouds are more reliable,
511 which is still unclear. The results and discussion presented here also raise the issue of how best
512 to classify newly refrozen leads or gray ice in the sea ice products, or even where exactly to draw
513 the line between sea ice and open water – another key consideration as we seek to increase the
514 utility of the ICESat-2 sea ice products.

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517 delivering and maintaining the high-quality sea ice data analyzed in this study. AP would also
518 like to thank Susan and Brian for the child-care support they provided in this challenging time,
519 enabling him to work on and complete this manuscript.

520 **Data availability**

521 Sentinel-2 imagery was derived from Copernicus data obtained at
522 <https://schihub.copernicus.eu>. The ICESat-2 ATL07 (sea ice height and type) and ATL10
523 (freeboard) data products used here are available through the NSIDC (<https://nsidc.org/data/atl07>
524 & <https://nsidc.org/data/atl10>). The gridded lead fraction and chord length estimates derived in
525 this study are available on Zenodo: <https://doi.org/10.5281/zenodo.4075390>.

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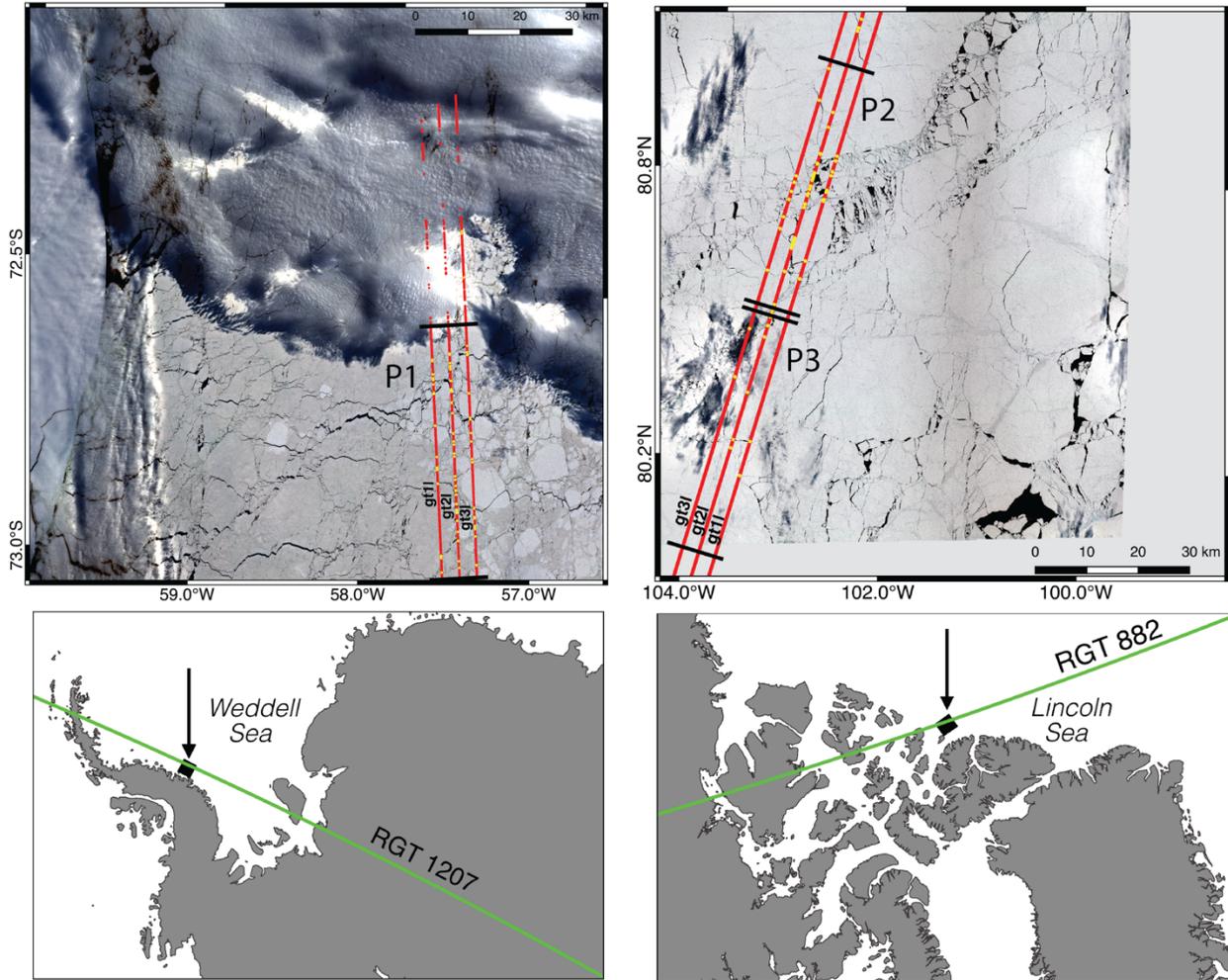
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Tables

	ground track	specular	dark lead	ssh	$L_f^{v_1}$ (%)	$L_f^{v_2}$ (%)	$C_l^{v_1}$ (km)	$C_l^{v_2}$ (km)
P1	gt1l	218	19	196	1.38	1.54	2.96	2.49
	gt1r	92	13	91	0.66	0.67	4.51	4.51
	gt2l	387	8	328	2.58	3.14	1.52	1.37
	gt2r	293	4	185	1.53	2.40	2.12	2.20
	gt3l	181	0	151	1.26	1.55	2.63	2.63
	gt3r	258	0	128	0.91	1.76	2.49	2.47
P2	gt1l	127	8	108	0.85	1.08	4.43	4.42
	gt1r	105	5	66	0.42	0.68	4.88	4.42
	gt2l	416	2	220	2.55	6.37	1.82	1.99
	gt2r	270	0	118	0.94	2.03	2.49	3.11
	gt3l	37	0	36	0.27	0.28	4.91	4.91
	gt3r	40	0	35	0.25	0.28	7.01	7.00
P3	gt1l	13	1	13	0.13	0.13	16.0	16.0
	gt1r	11	0	11	0.07	0.07	16.0	16.0
	gt2l	89	0	74	0.84	1.04	3.95	3.95
	gt2r	70	0	24	0.19	0.53	6.80	7.88
	gt3l	12	0	12	0.09	0.09	16.0	16.0
	gt3r	17	0	17	0.12	0.12	12.0	12.0

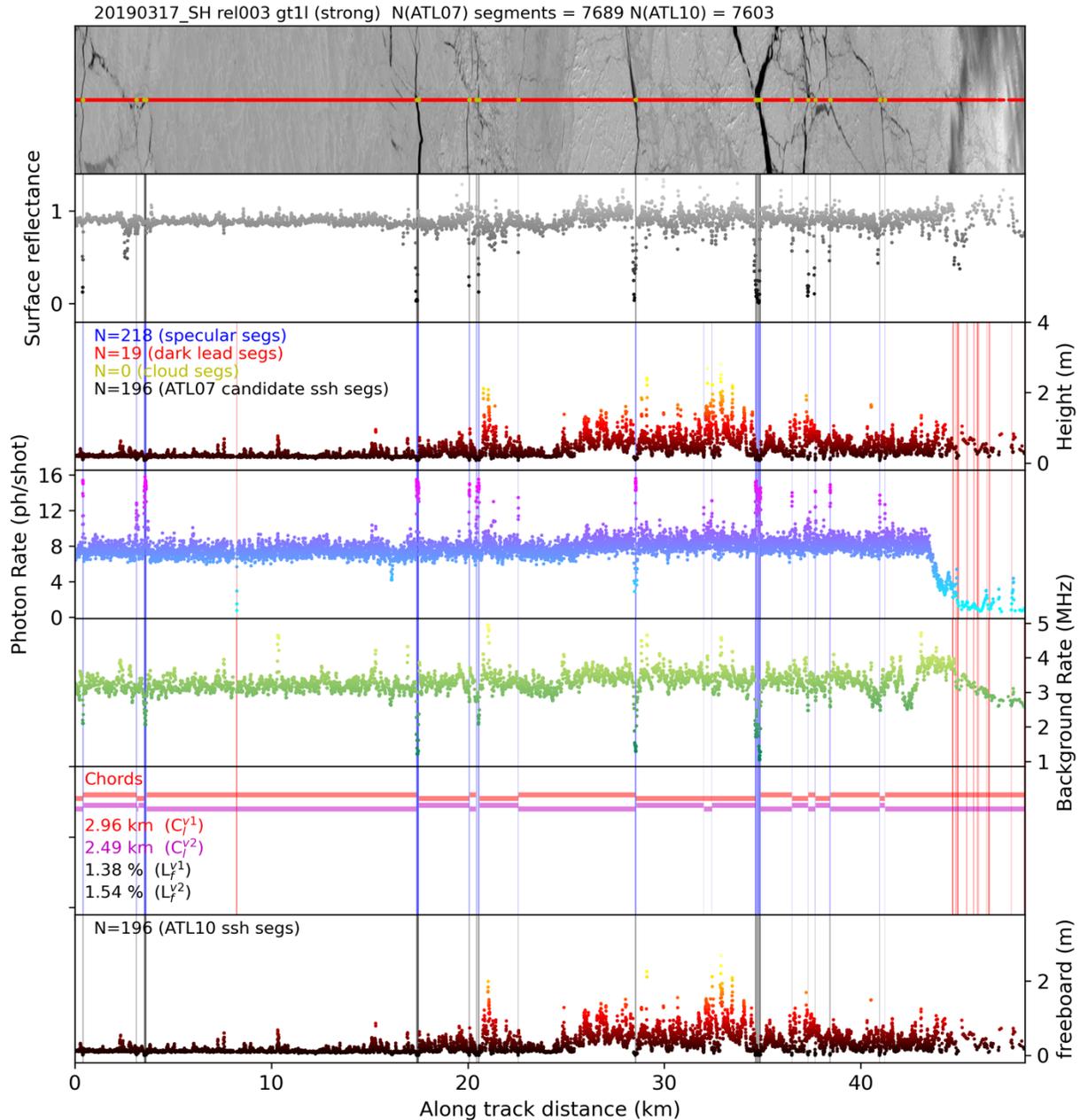
637 **Table 1:** Summary statistics of the ICESat-2 data in the three S-2 profiles (P1 - P3 in Figure 1)
638 across all six beams. Strong beams in this orientation are shaded gray. Columns 2-4: Number of
639 specular lead segments: *height_segment_type* = 2 - 4, dark leads segments: *height_segment_type*
640 = 5 - 9, sea surface segments: *ssh_flag* = 1 or 2 (all candidate lead segments become lead
641 segments in these profiles). Note that no cloud segment classifications were found in these
642 profiles, so they were not included in the Table. Columns 5-8: Lead fraction and chord length
643 estimates using our two different (v_1 and v_2) metrics.

644 **Figures**

645

646 **Figure 1:** ICESat-2 ATL07 strong beam data overlaid on near-coincident Sentinel-2 RGB image
 647 of (left) the Western Weddell Sea, Southern Ocean, on March 17th, 2019 and (right) the Lincoln
 648 Sea of the Arctic Ocean on May 25th, 2019. The time difference between the ICESat-2 and
 649 Sentinel-2 overpass for the Weddell Sea scene is ~ 7 minutes (ICESat-2 acquisition at 12:29
 650 UTC, Sentinel-2 acquisition at 12:22 UTC), time difference for the Lincoln Sea scene is ~ 94
 651 minutes (ICESat-2 acquisition at 00:36 UTC on 26 May 2019, Sentinel-2 acquisition at 23:01
 652 UTC on 25 May 2019). Red segments are classified as sea ice ($ssh_flag = 0$), yellow segments
 653 are classified as leads ($ssh_flag \geq 1$). The ICESat-2 footprint is not to scale. P1-P3 indicate the
 654 ~ 50 km profiles shown in Figures 2 and 3 and the Supplementary Information respectively.

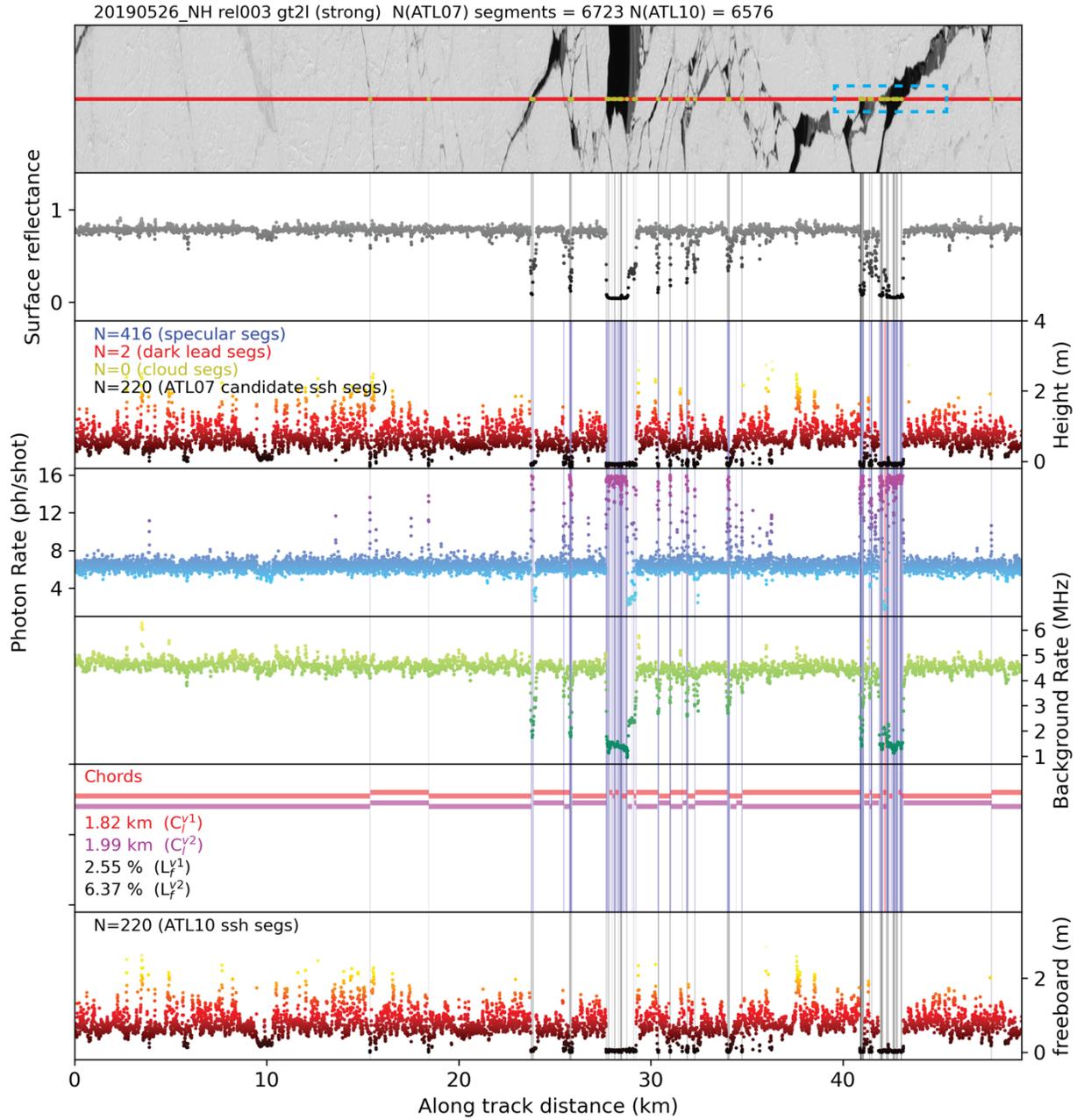
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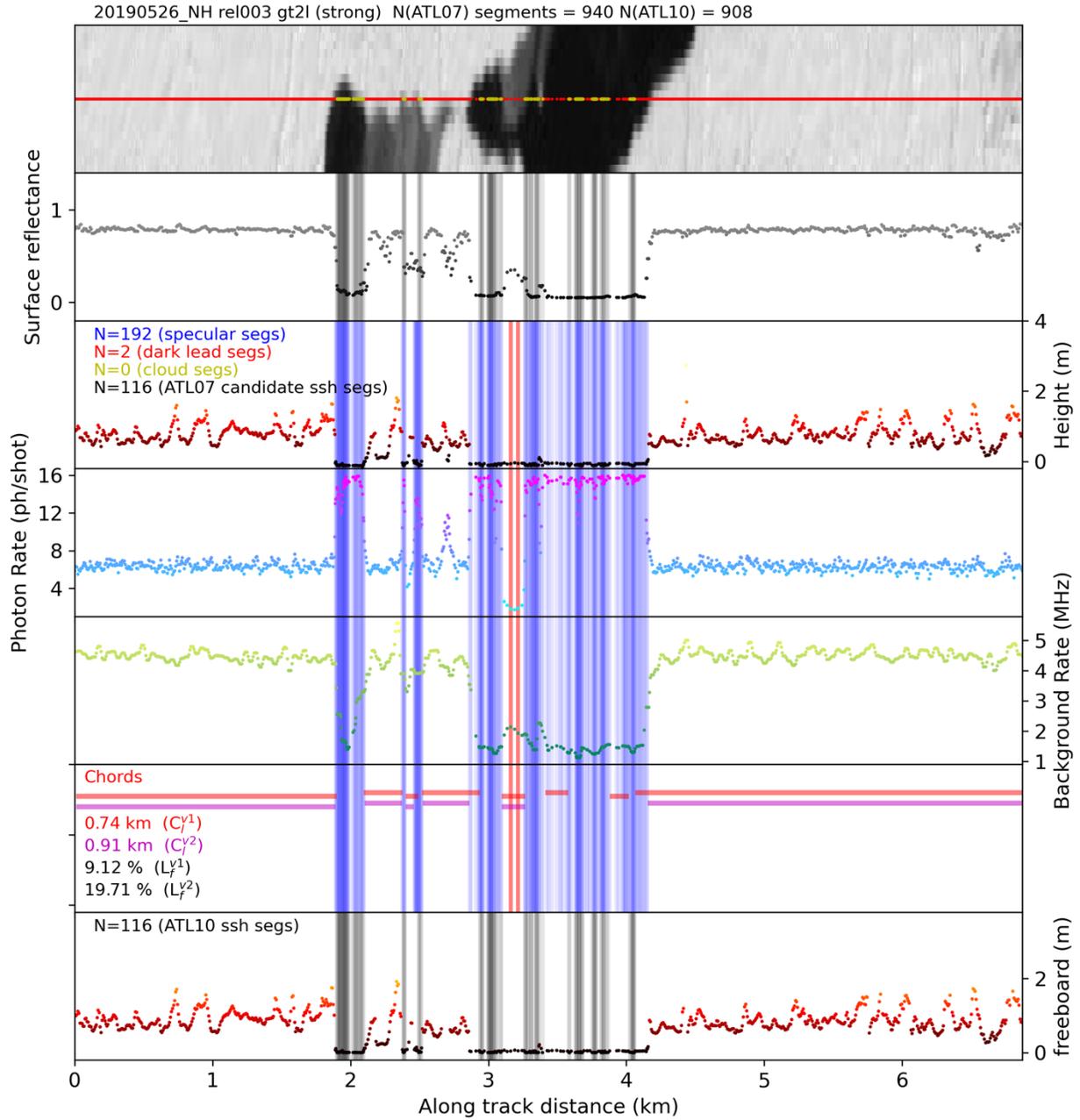
657 **Figure 2:** ICESat-2 ATL07 data from the *gt1l* ground-track (strong beam in this spacecraft
658 orientation) overlaid on Sentinel-2 RGB imagery for a ~ 50 km profile in the Western Weddell
659 Sea on March 17th, 2019 (P1, see Figure 1 for location). (top panel) Sentinel image with ICESat-
660 2 profile overlaid showing the sea surface flag (*ssh_flag*, red = ice, yellow = sea surface);
661 (second panel) surface reflectance calculated from the red band of the Sentinel-2 image from
662 nearest neighbor pixels to the ICESat-2 profile; (panels 3-5) segment height, photon rate and
663 background rate respectively from ATL07. Grey shading in panels 2-3 indicate *candidate* leads

664 (*ssh_flag* = 1) while in panels 4-5, blue shading indicates specular lead classifications
665 (*height_segment_type* = 2 to 5) and red shading indicates dark lead classifications
666 (*height_segment_type* = 6 to 9). Panel 7 shows the derived chord lengths as horizontal bars (red
667 = v_1 , magenta = v_2) with each chord grouping shifted vertically to indicate the groups, and
668 statistics of the mean chord lengths and lead fractions using these two approaches. Panel 8 shows
669 the freeboard in ATL10, with the shading indicating actual leads used to derive reference sea
670 surface (*ssh_flag* = 2). N indicates the number of segments in ATL07 and ATL10 as specified.
671



672

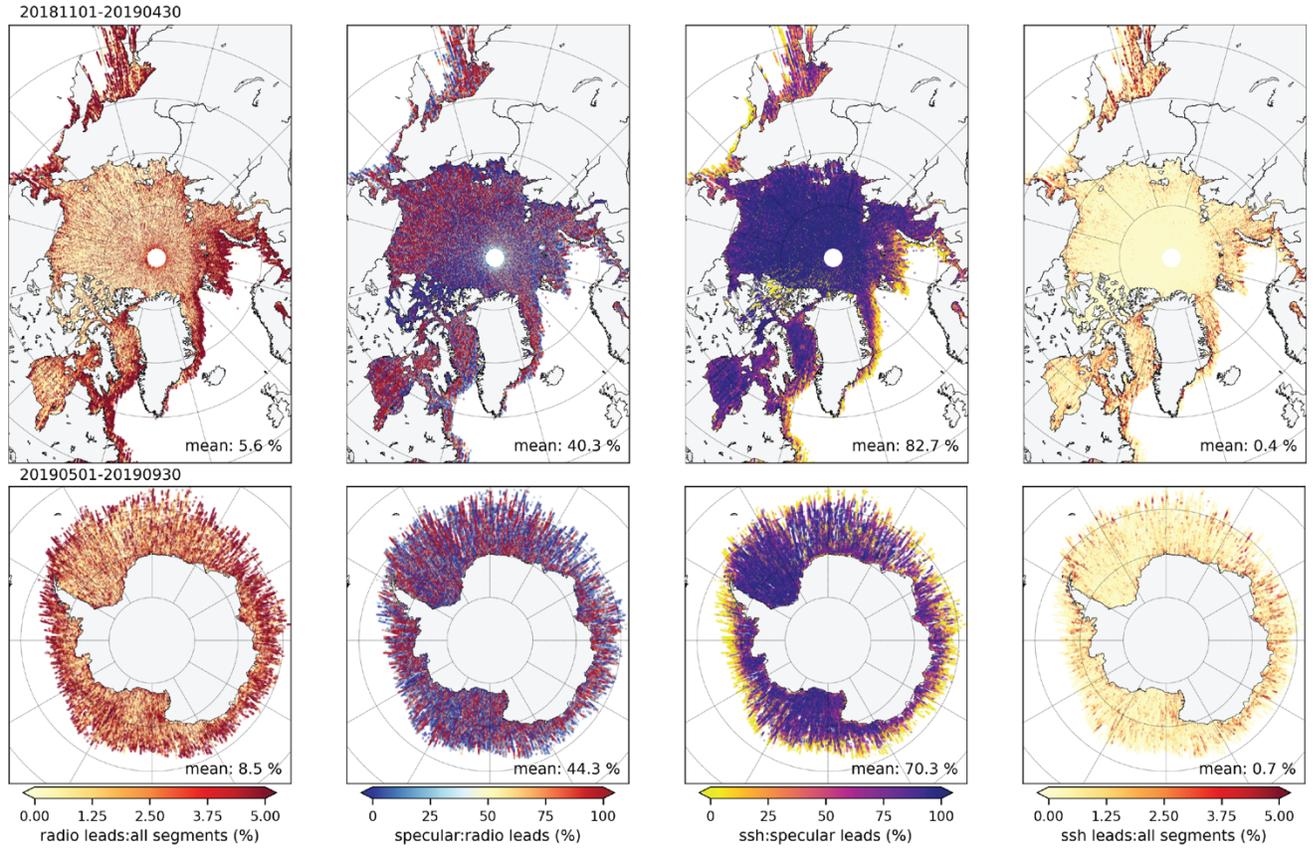
673 **Figure 3:** As in Figure 2 but for the P2 profile over the Lincoln Sea of the Arctic Ocean on May
 674 26th, 2019 for the *gt2l* ground-track (strong beam in this spacecraft orientation) (see Figure 1 for
 675 location). The cyan box in the top panel shows the location of the zoomed in profile in Figure 4.



676

677 **Figure 4:** As in Figure 3 (P2 profile) but zoomed in on the second large lead opening (cyan box
 678 in Figure 3).

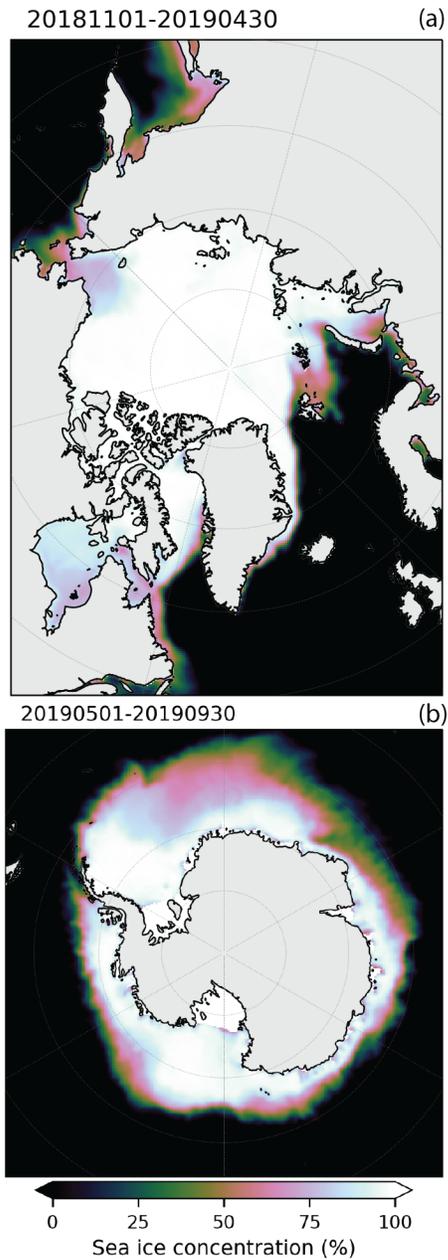
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680

681 **Figure 5:** (first column) percent of all ATL07 segments classified as radiometric leads
 682 (*height_segment_type* = 2 to 9), (second column) percent of radiometric leads classified as
 683 specular (*height_segment_type* = 2 to 5), (third column) percent of specular leads that become
 684 sea surface segments (*ssh_flag* ≥ 1), (fourth column) percent of all height segments that become
 685 sea surface segments. (Top row) Arctic Ocean, November 1 to April 30, 2019, (bottom row)
 686 Southern Ocean, May 1 to September 30, 2019. All these percentages are calculated in 10 km
 687 along-track sections across the three strong beams and are segment length weighted.

688



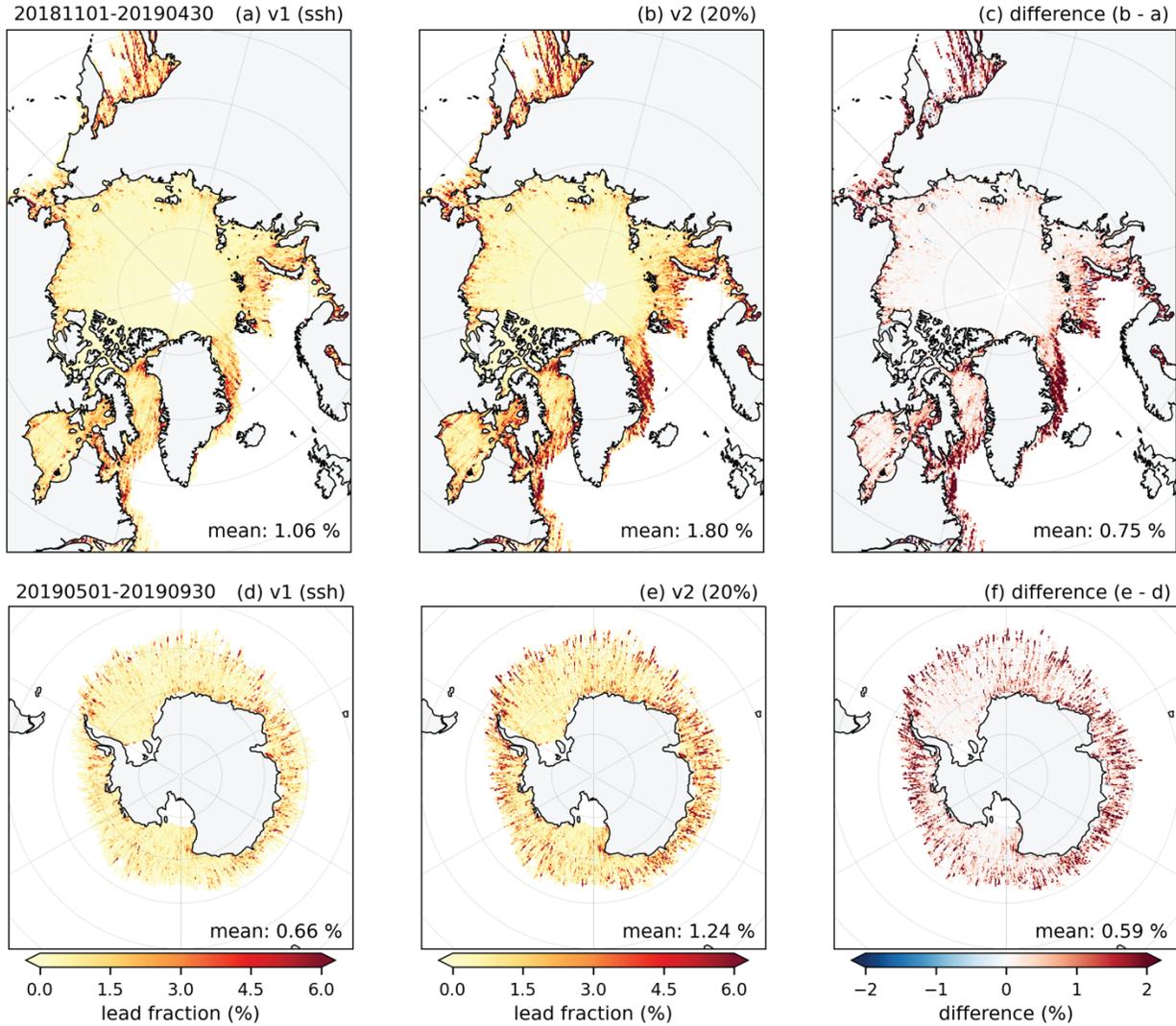
689

690 **Figure 6:** Mean winter sea ice concentration derived from passive microwave data (Climate

691 Data Record, CDR) for the (a) Arctic Ocean, November 1, 2018 to April 30, 2019 and (b)

692 Southern Ocean, May 1 to September 30, 2019.

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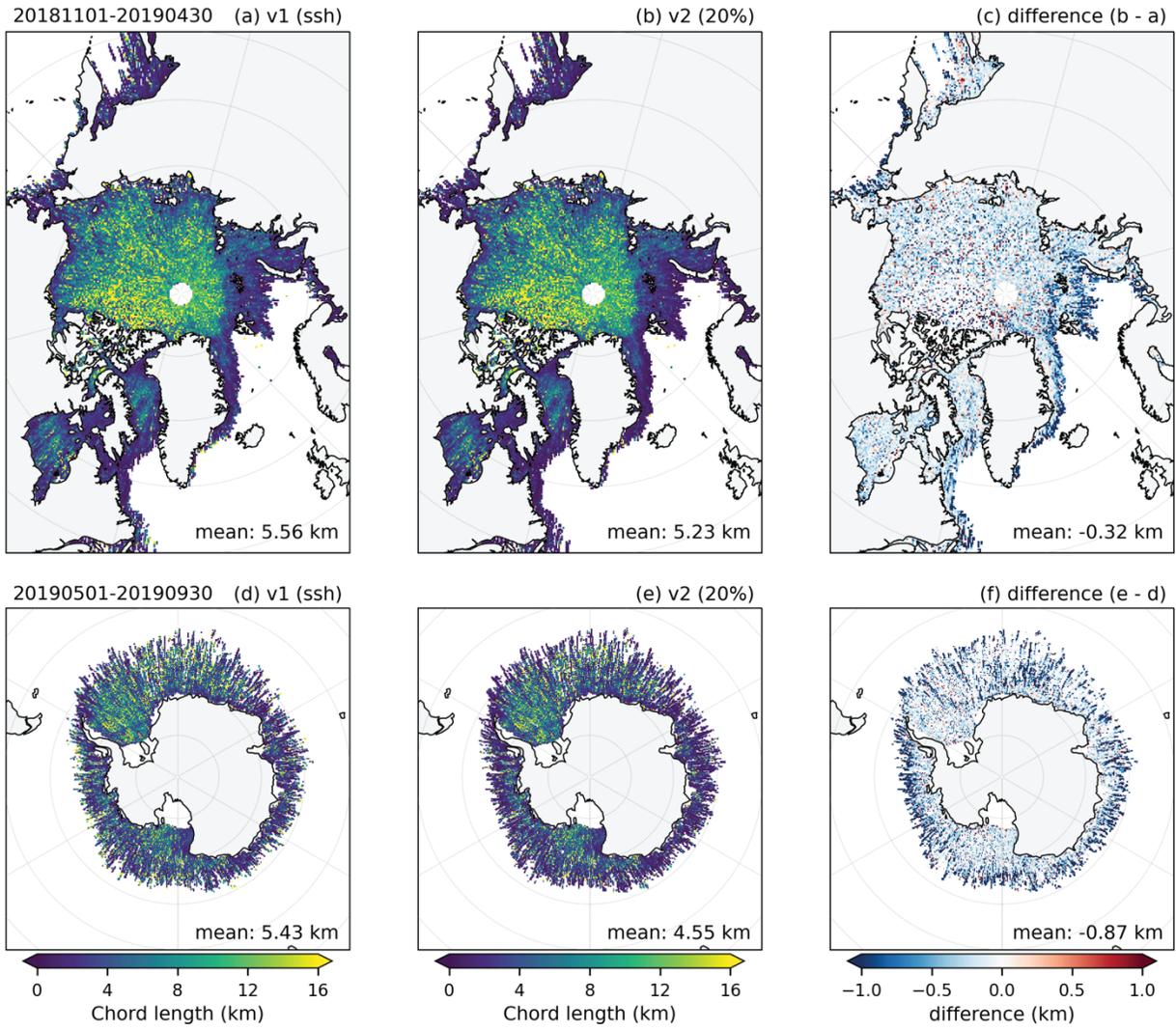
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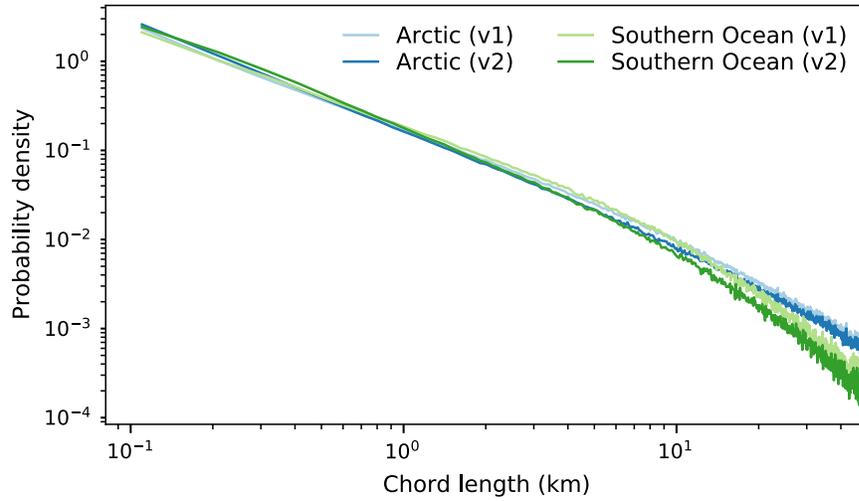
700

Figure 7: Lead fraction using (a and d) the *ssh_flag* (L_f^{v1}) and (b and e) the higher 20% height filter (L_f^{v2}) and the difference ($c = b$ minus a , $f = e$ minus d). Data are for the (top row) Arctic Ocean, November 1 to April 30, 2019, (bottom row) Southern Ocean, May 1 to September 30, 2019. All the fractions are calculated in 10 km along-track sections across the three strong beams and are segment length weighted. Data have binned to a 25 km x 25 km polar stereographic grid.



701

702 **Figure 8:** As in Figure 7 but for chord length (left) C_l^{v1} and (middle) C_l^{v2} and difference (right).



703

704 **Figure 9:** Probability distributions of the Arctic and Southern Ocean chord lengths (v1: C_l^{v1} and
 705 v2: C_l^{v2}) for the same winter-time periods as in Figure 8. Distributions are plotted on a log-log
 706 scale.

707