

Solar Coronal Jets and Jet-like Features: Some Recent Investigations

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Coronal Jets

Wed Jan 10 16:13:36 2007

SAO /NASA/JAXA/NAOJ

Cirtain et al. (2007)

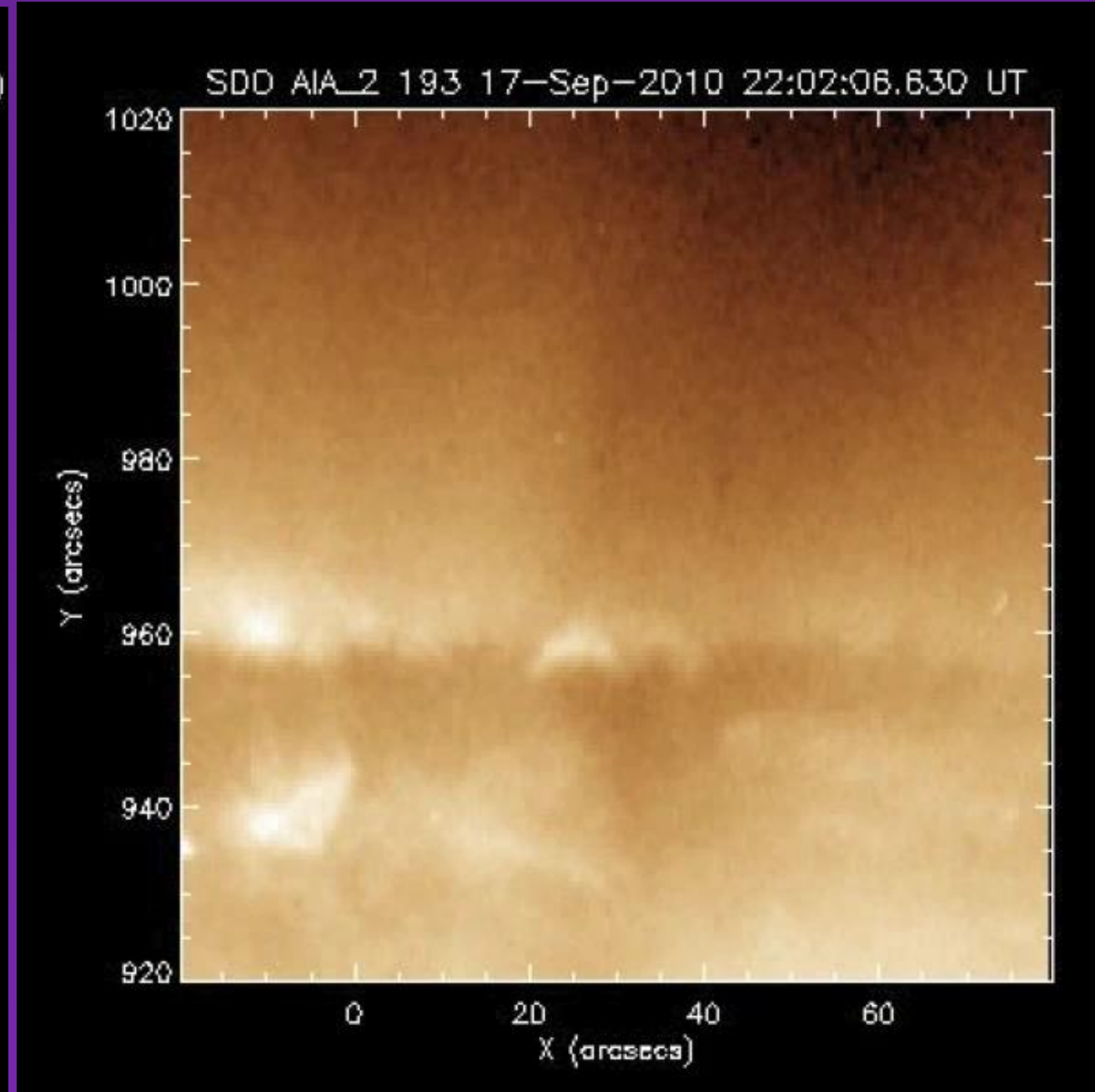
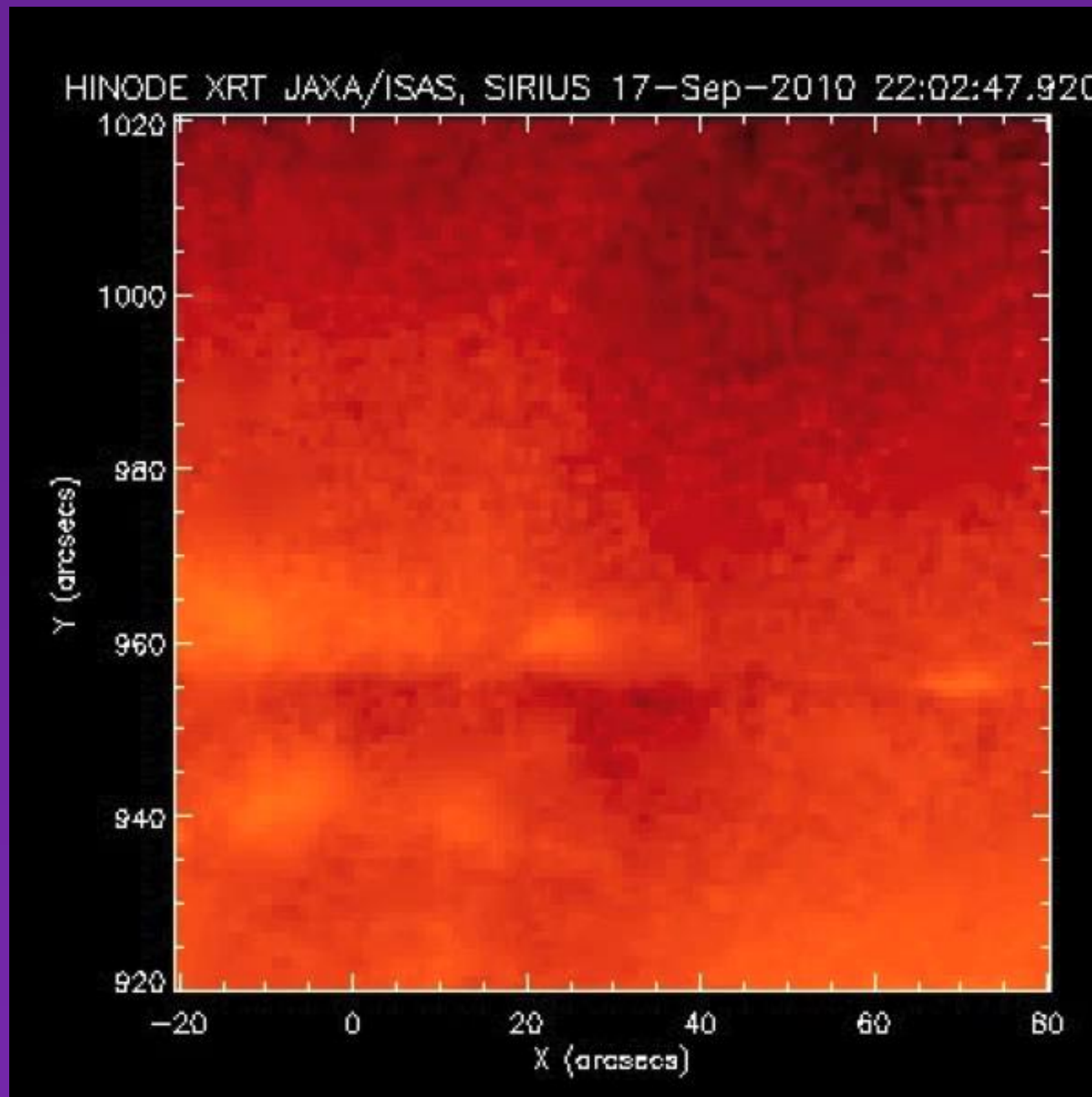
Introduction: Solar Coronal Jets

- Observed since the Yohkoh days (Shibata et al. 1992; also Shimojo et al. 1996, etc.) (Reviewed by Raouafi et al. 2016; Hinode Review (Sterling) 2019; Shen 2021.)
- Yohkoh (SXT) saw them mainly in active regions.
- Hinode/XRT found them to be plentiful in polar coronal holes (Cirtain et al. 2007; also Savcheva et al. 2007, etc.)
- Stereo EUVI+coronagraph (Nisticò et al. 2009, 2015).
- In polar coronal holes: size $\sim 50,000$ km x 8000 km; rate ~ 60 /day (Savcheva et al. 2007).
- Often have a “hot loop” at the jet’s base.
- Jets often display spinning (untwisting?) motions, both visually and spectroscopically (e.g., Pike & Mason 1998).

Coronal Hole Jets: “Minifilament eruptions”

XRT

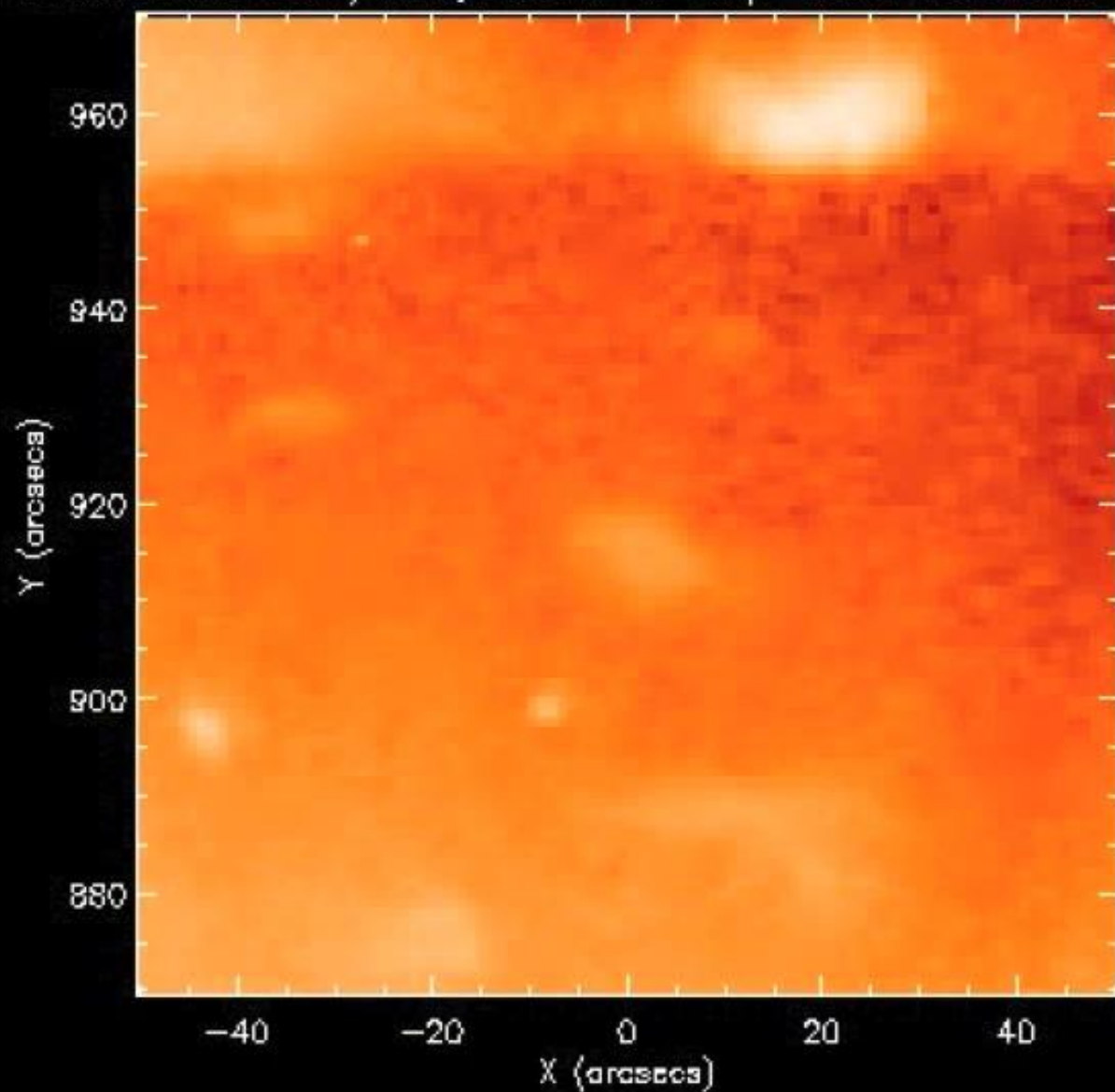
AIA 193



Sterling et al. (Nature, 2015): 20 Polar CH jets.

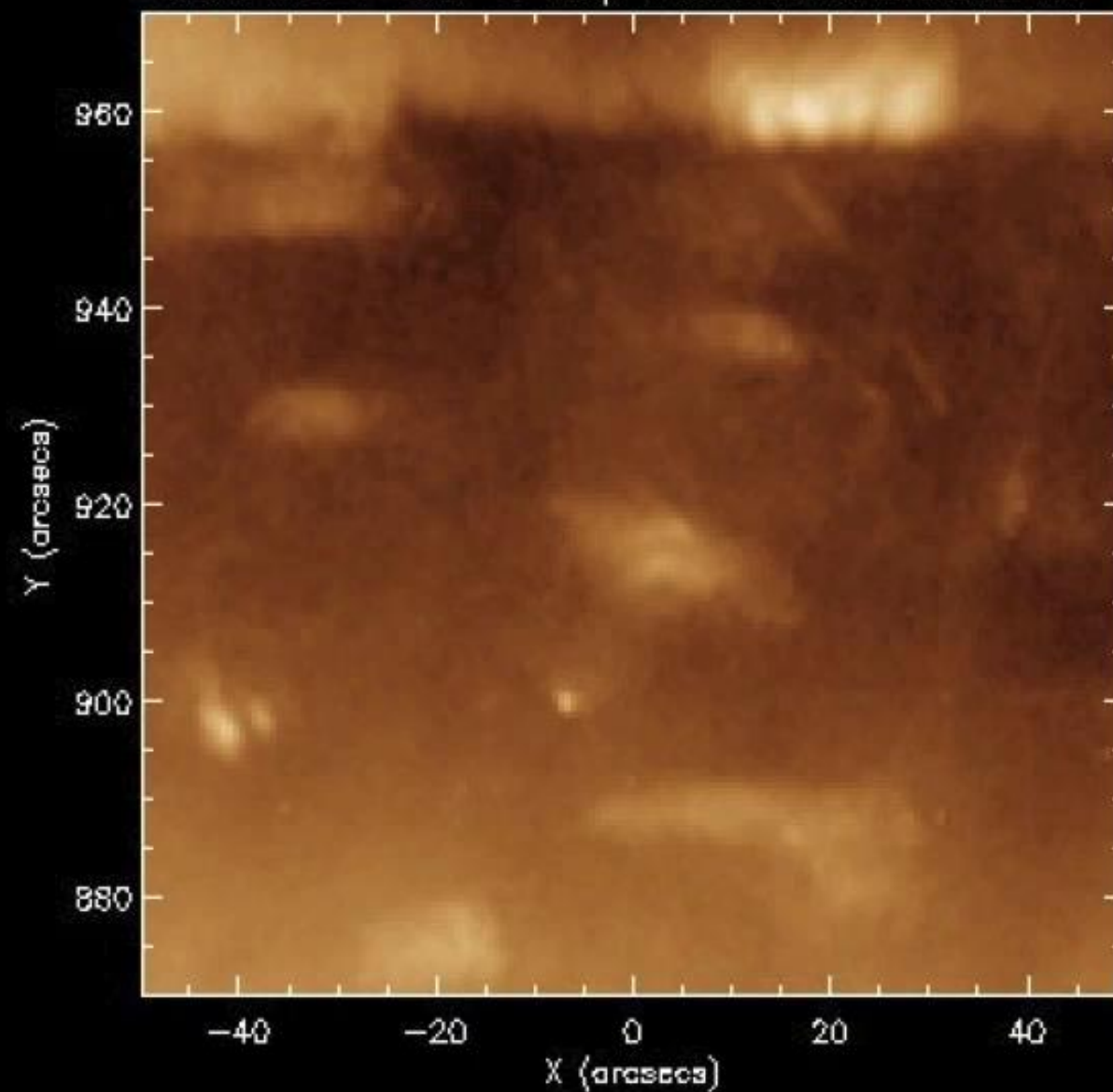
XRT

HINODE XRT JAXA/ISAS, SIRIUS 9-Sep-2010 21:50:23.471



AIA 193

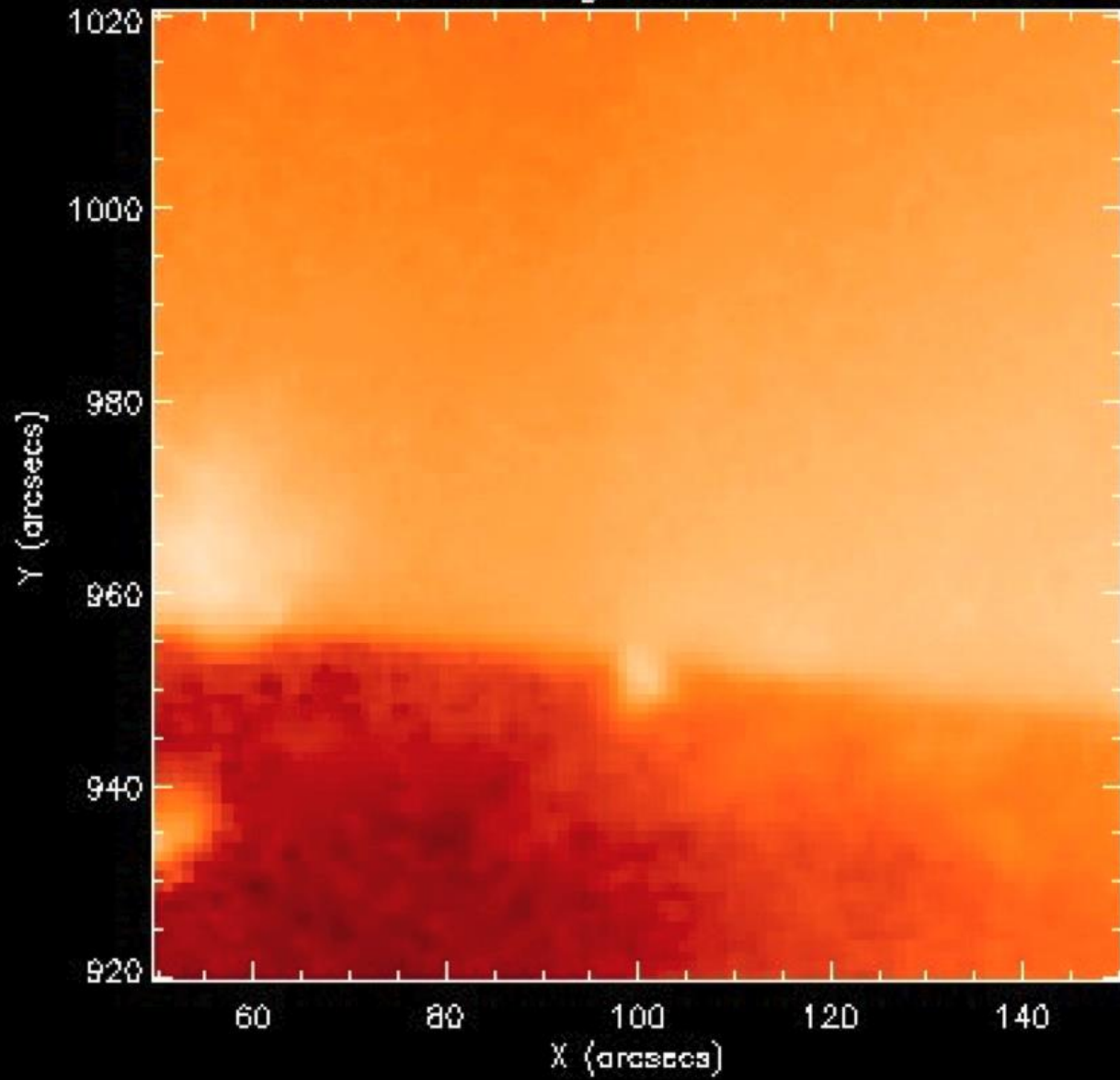
SDO AIA_2 193 9-Sep-2010 21:50:06.630 UT



Event 12

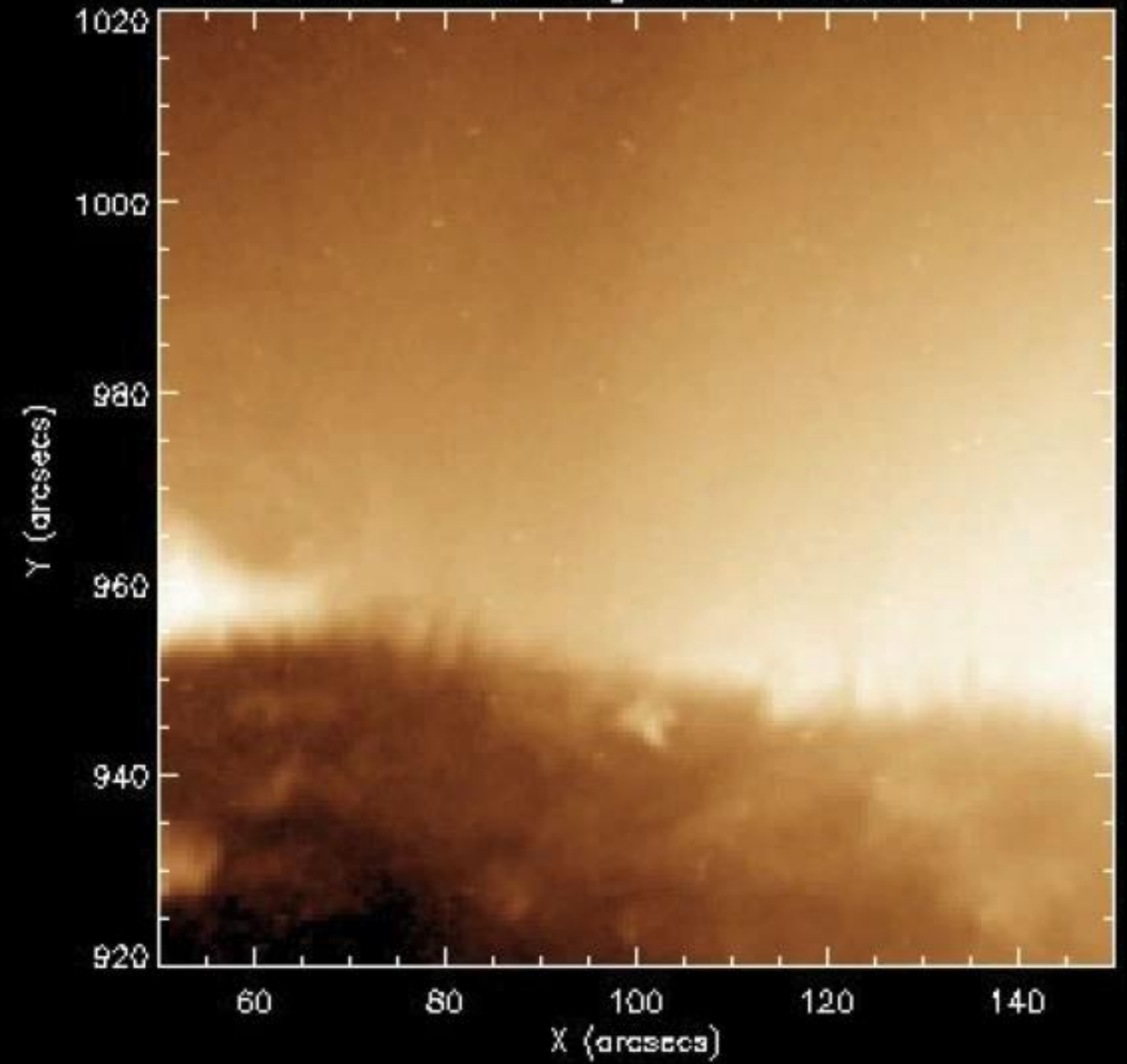
XRT

HINODE XRT 26-Aug-2010 14:07:21.527 UT



AIA 193

SDO AIA_2 193 26-Aug-2010 14:02:18.620 UT

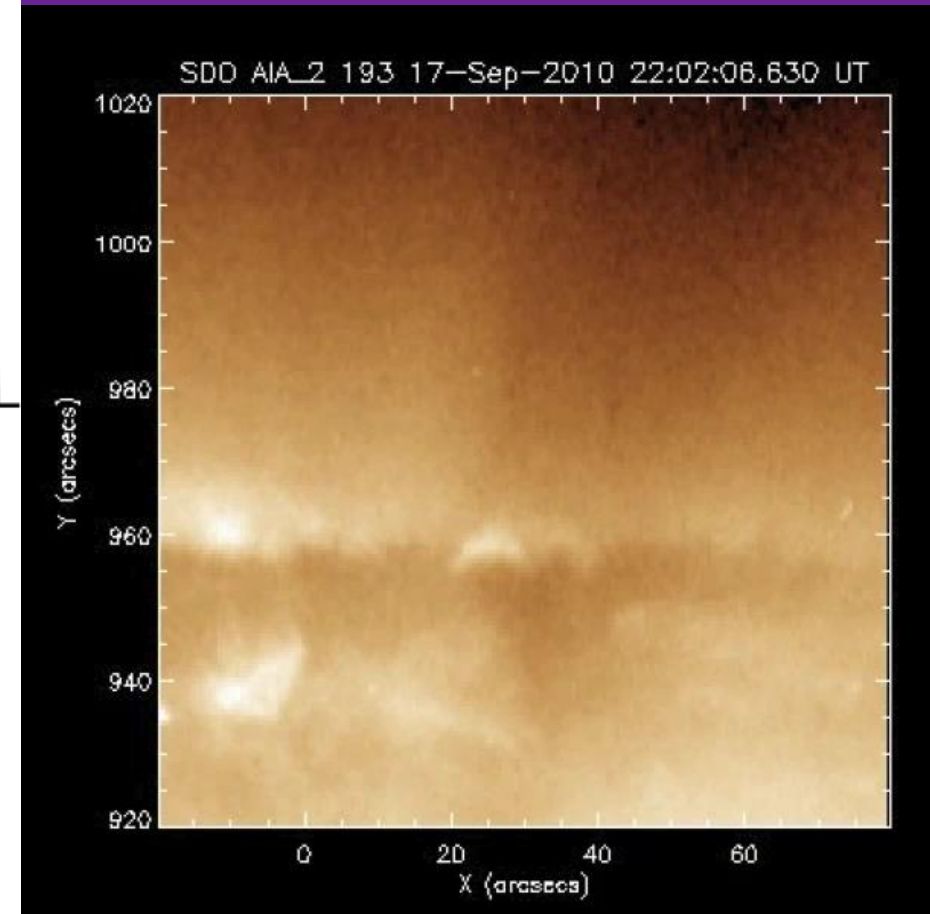
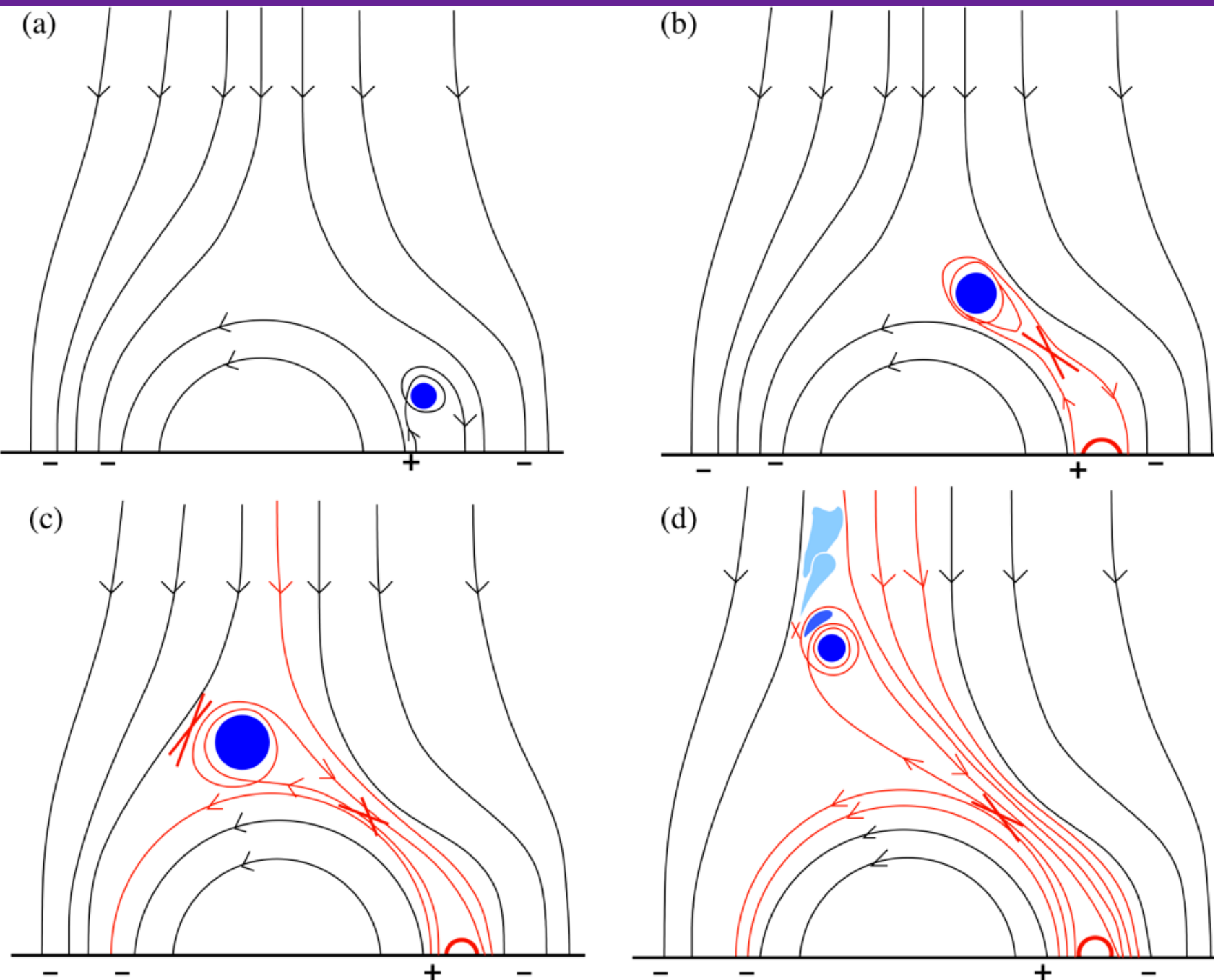


Event 3

“Normal” Filament Eruption (TRACE)



Minifilament-Eruption Model for (X-Ray) Jets



Sterling et al. (2015, 2016, 2017)

Quiet Sun jets work the same way (Panesar et al. 2016b)

Modeled by Wyper et al. 2017, 2018)

What Causes Miniature-Filament Eruptions?

- Did not look on-disk in this study, due to polar view. But....
- Adams et al. (2014) found no emerging flux in the jet region. Filament erupted from location where flux canceled. (Also, Hong et al. 2014.)
- Several other found cancelation leading to jets (e.g., Hong et al. 2011; Huang et al. 2012; Young & Muglach 2014a,b).
- Some others found jets from location of emerging flux+flux cancelation (e.g., Liu et al. 2011; Shen et al. 2012, 2017; Hong et al. 2012; Li et al. 2015).

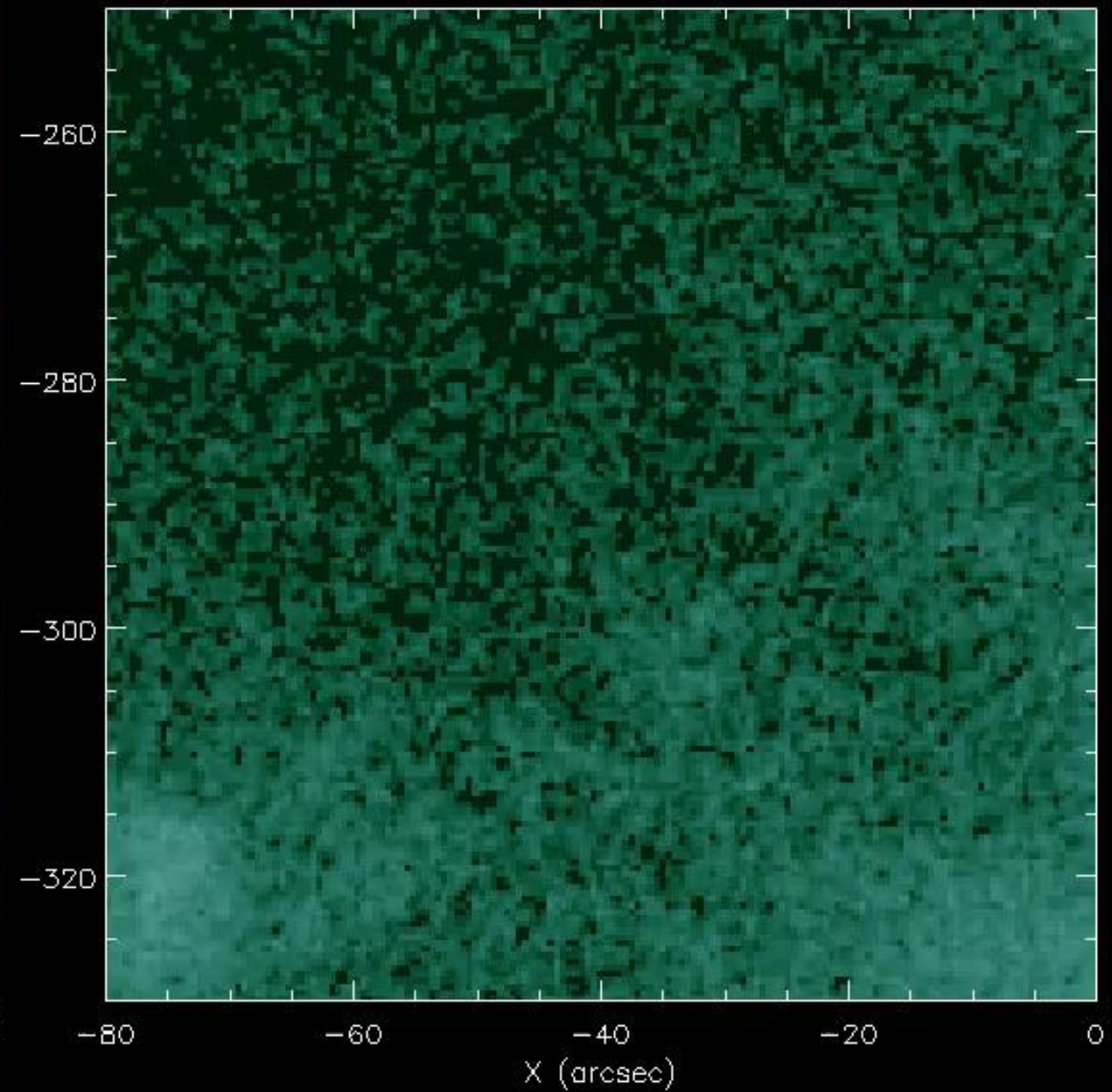
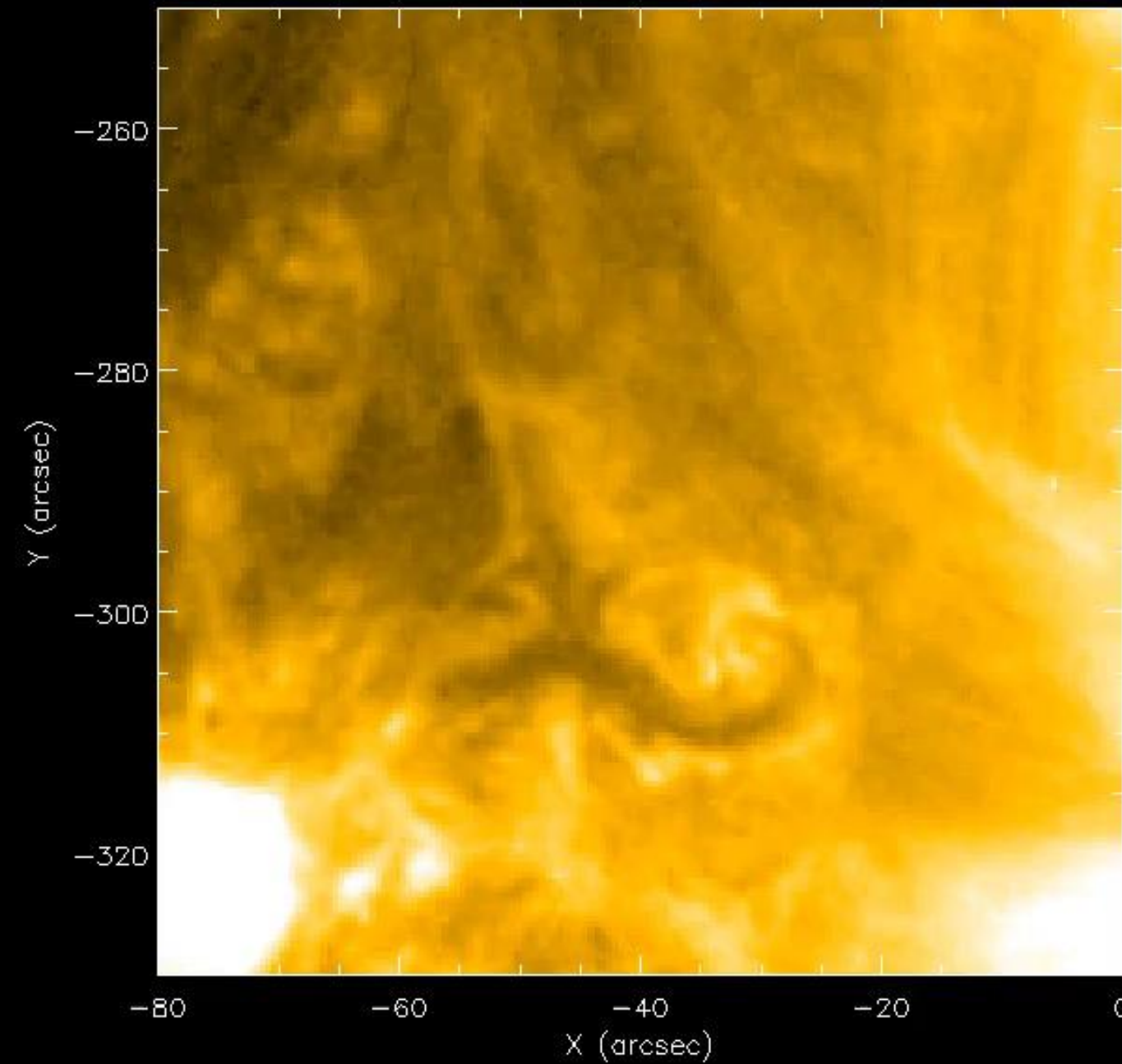
Quiet Sun Jets — Similar to PCH jets

AIA 171

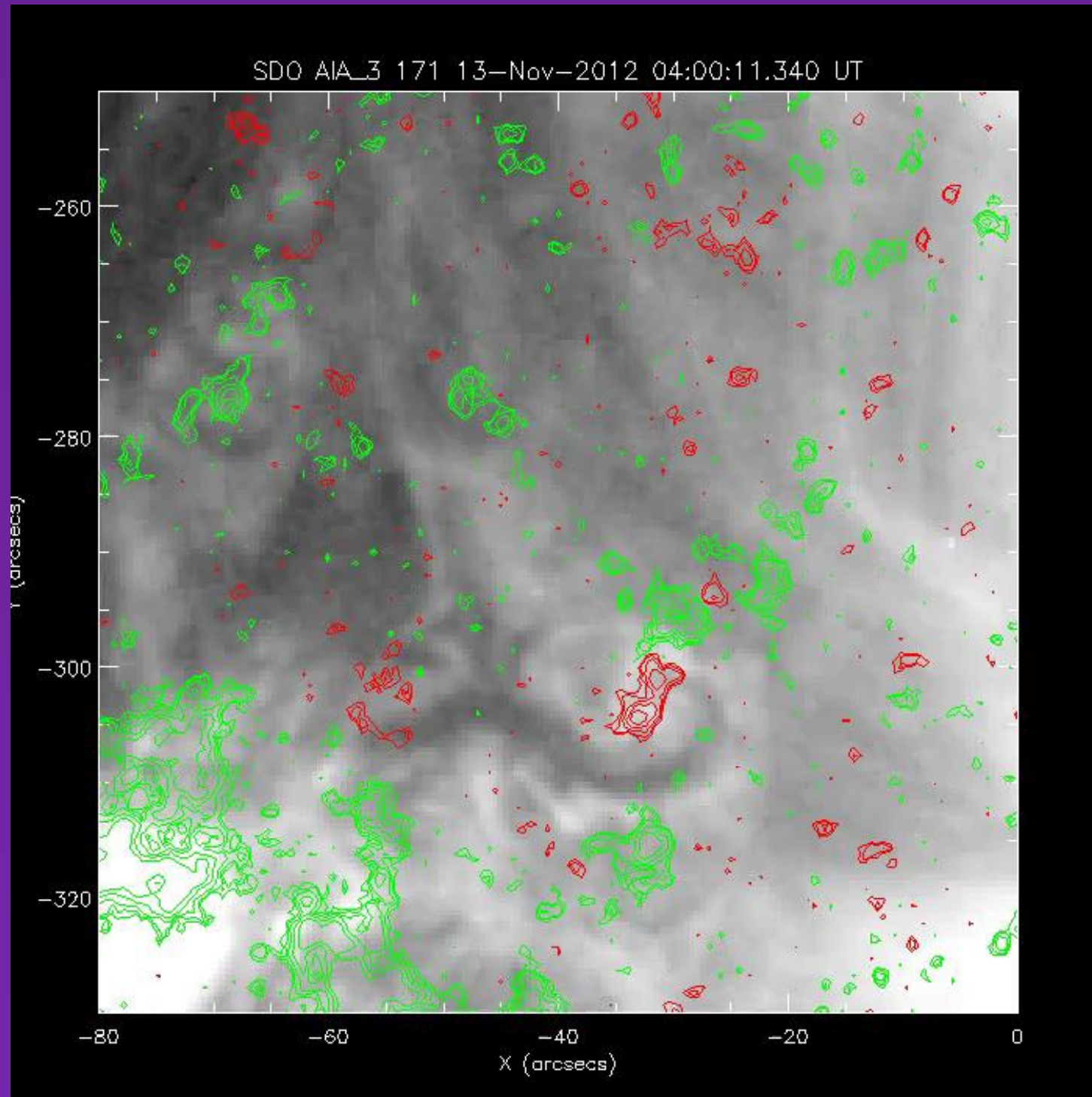
AIA 94

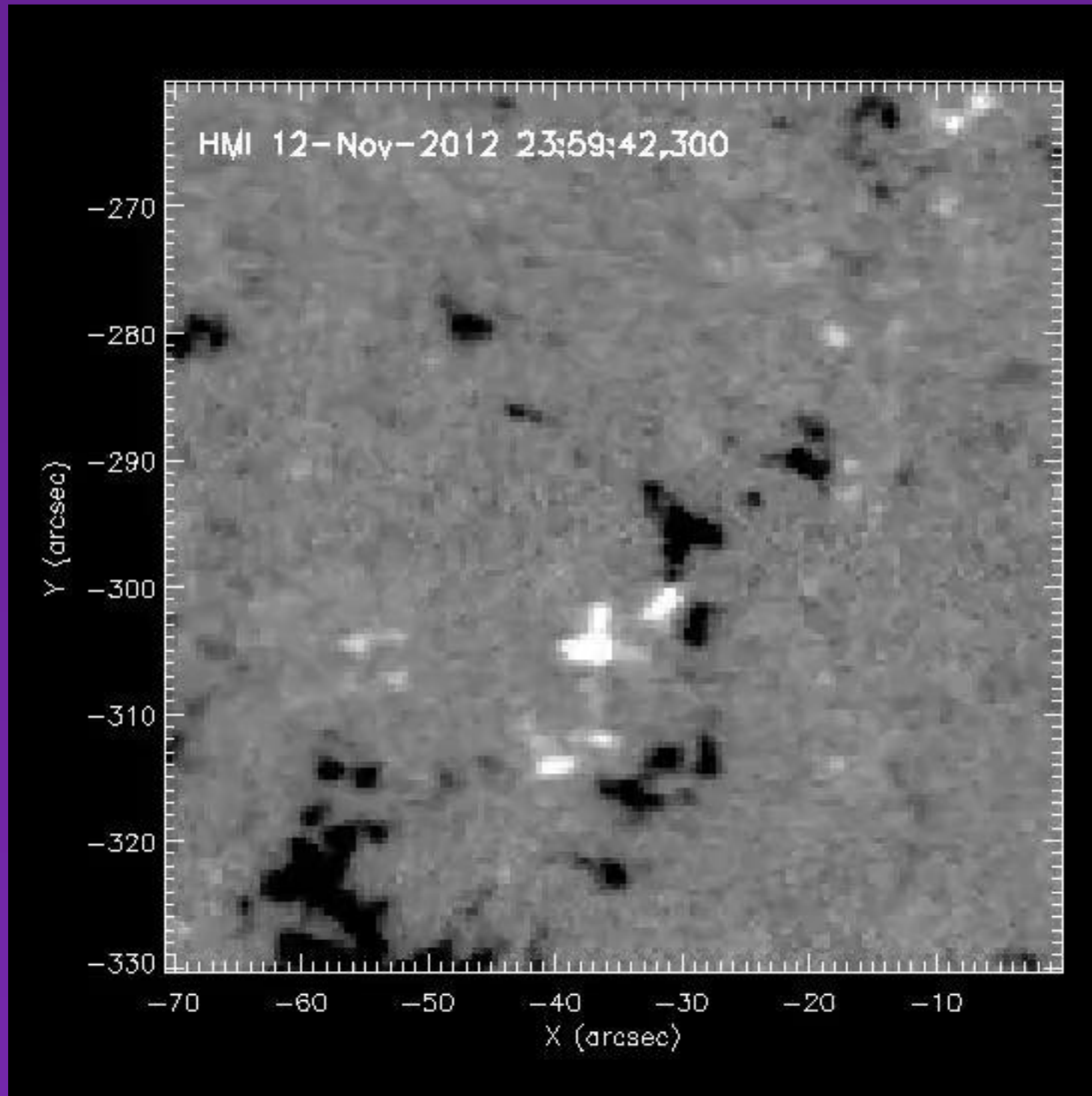
AIA 171: 13-Nov-2012 04:00:11 UT

AIA 94: 13-Nov-2012 04:00:01 UT



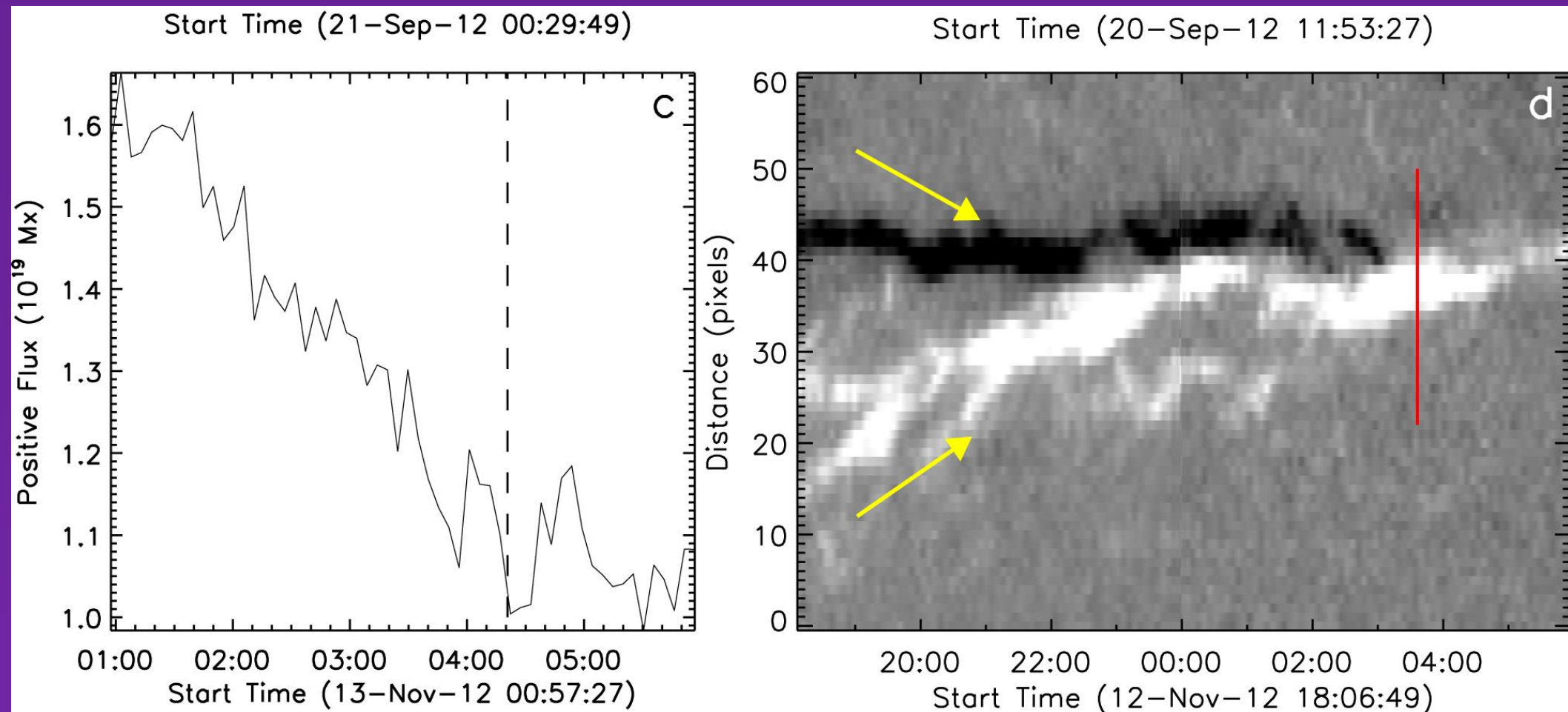
(Panesar et al. 2016b)





Panesar et al. (2016)

Same for QS jets: Occur at cancelation sites.



(Ave. Cancelation rate: $\sim 10^{18}$ Mx/hr.)

Panesar, Sterling, & Moore (2016b)

Large-ish Studies of Quiet Regions and Coronal Hole Jets:

- Sterling et al. (2015): 20 PCH jets - all have minifilament eruptions.
- Panesar, Sterling, & Moore (2016b): 10 QS jets
- Panesar, Sterling, & Moore (2018): 13 CH jets
 - All due to flux cancelation and minifilament eruptions (relatively large - easy-to-find jets).
- McGlasson et al. (2019) — 60 jets; 30 QS and 30 CHs. 85% consistent with flux cancelation and minifilament eruptions.
- Mulay et al. (2016). 20 AR jets - report 50% with flux cancelation.
- Kumar et al. (2019): 27 on-disk CH jets: 67% with minifilament eruptions, remainder filament-less eruptions; 22% with flux cancelation.

Active Region Coronal Jets

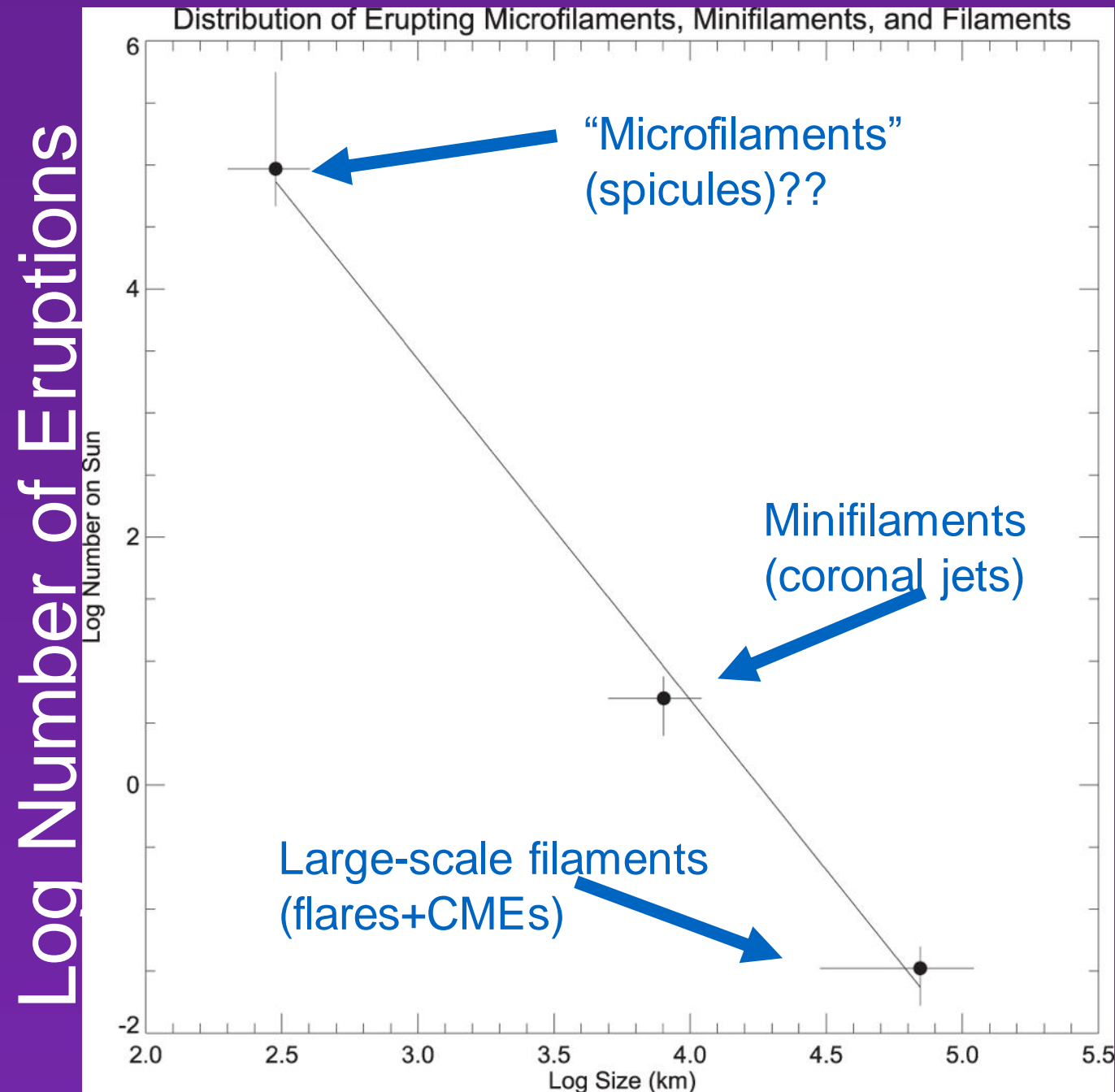
- Yohkoh studies (Shibata et al., Shimojo et al., many others).
- Raouafi et al. (2016).
- Mulay et al. (2016, 2017a, b, 2018) - AR-jet temps/emissions.
- Hong et al. (2017) — Minifil. eruption —> AR jet & Type III burst (also Shen et al. 2017, Moroccan et al. 2017).
- Panesar et al. (2016a); Sterling et al. (2016, 2017).

Current Views Regarding AR Jets

- Some can have different appearance from QS and CH jets.
- The ones we have studied follow the same basic picture: flux cancelation followed by a minifilament eruption that produces the jet.
- Different appearance likely due to complex magnetic environment and stronger fields.
- Some (many?) may have erupting minifilament flux rope, but little or no cool minifilament material.
- Still must study more!

Do Jets Exist on Smaller Size Scales?

Filament-Like Feature Eruptions on Smaller Scales??

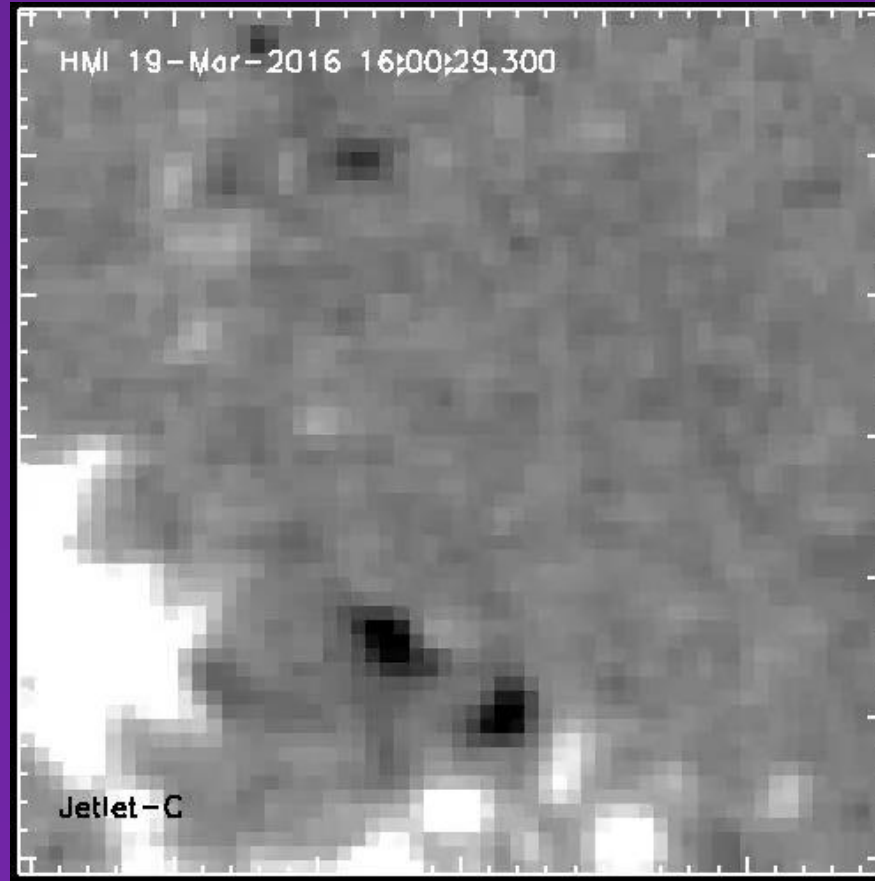
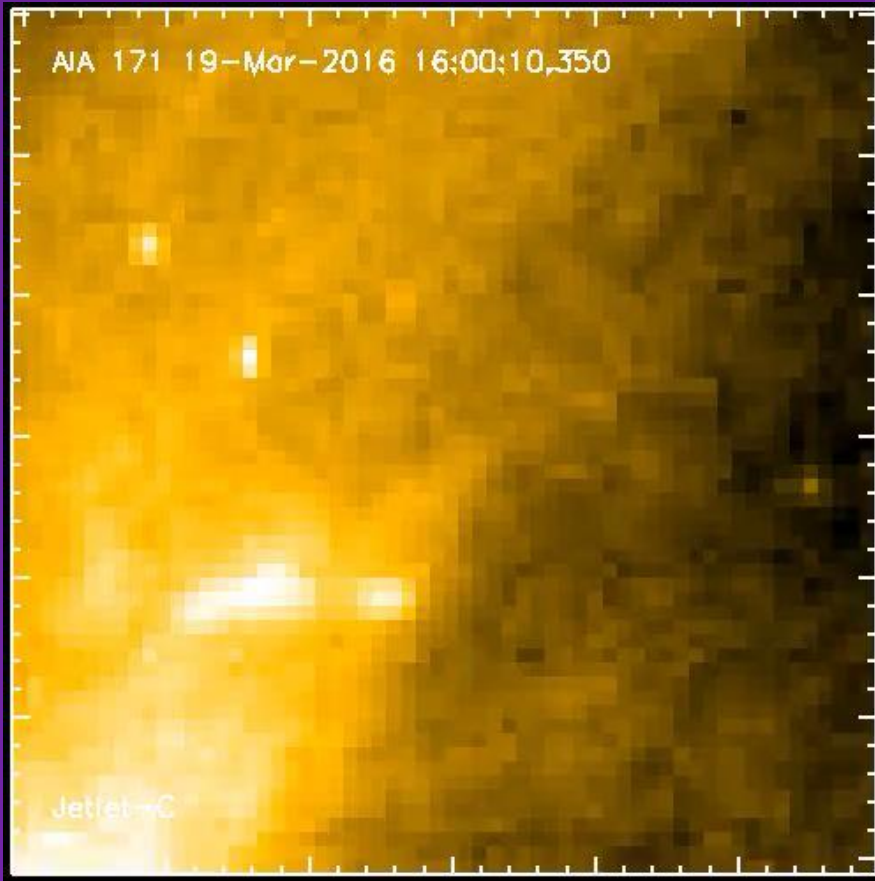
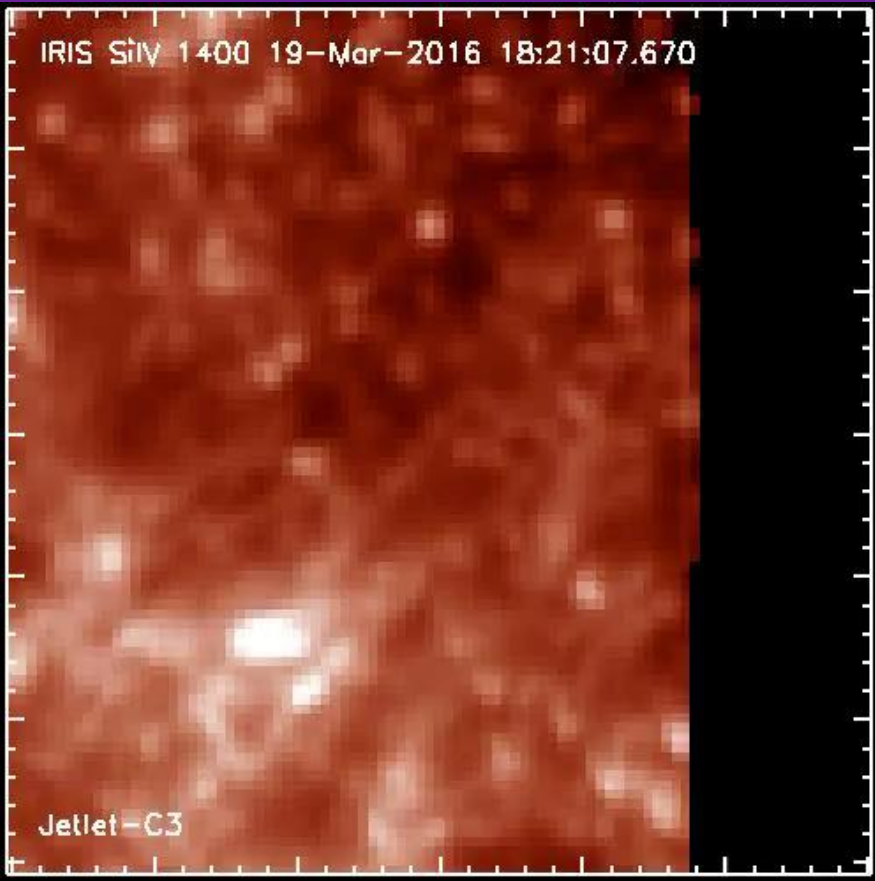


Log “Filament” Size

Sterling & Moore (2016)

Jetlets

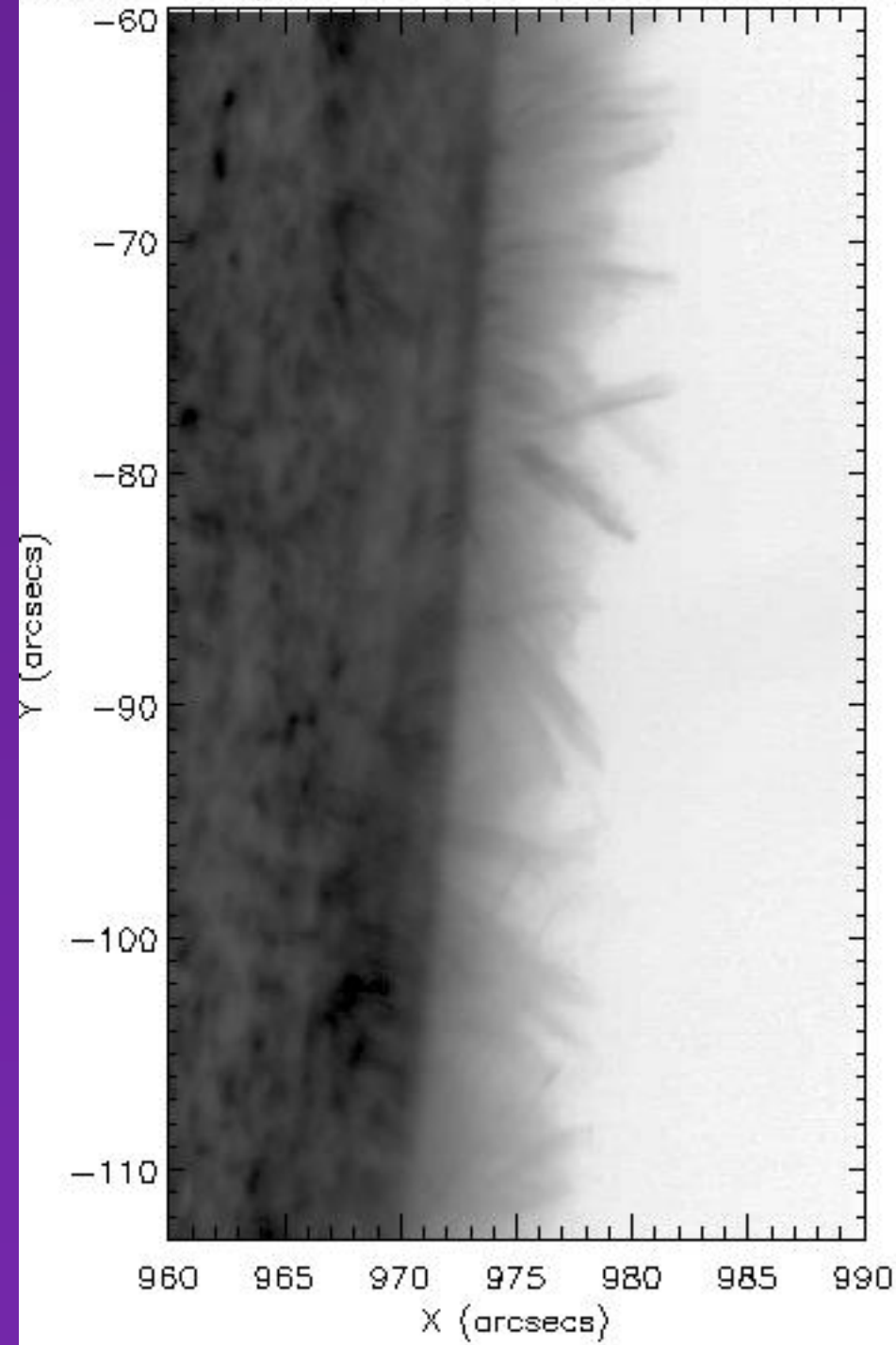
- Smaller than jets, larger than spicules.
- Seen in EUV, and in UV (IRIS).
- First identified at plume bases (Raouafi & Stenborg 2014).
- Also occur in more general network locations (Panesar et al. 2018, 2019).



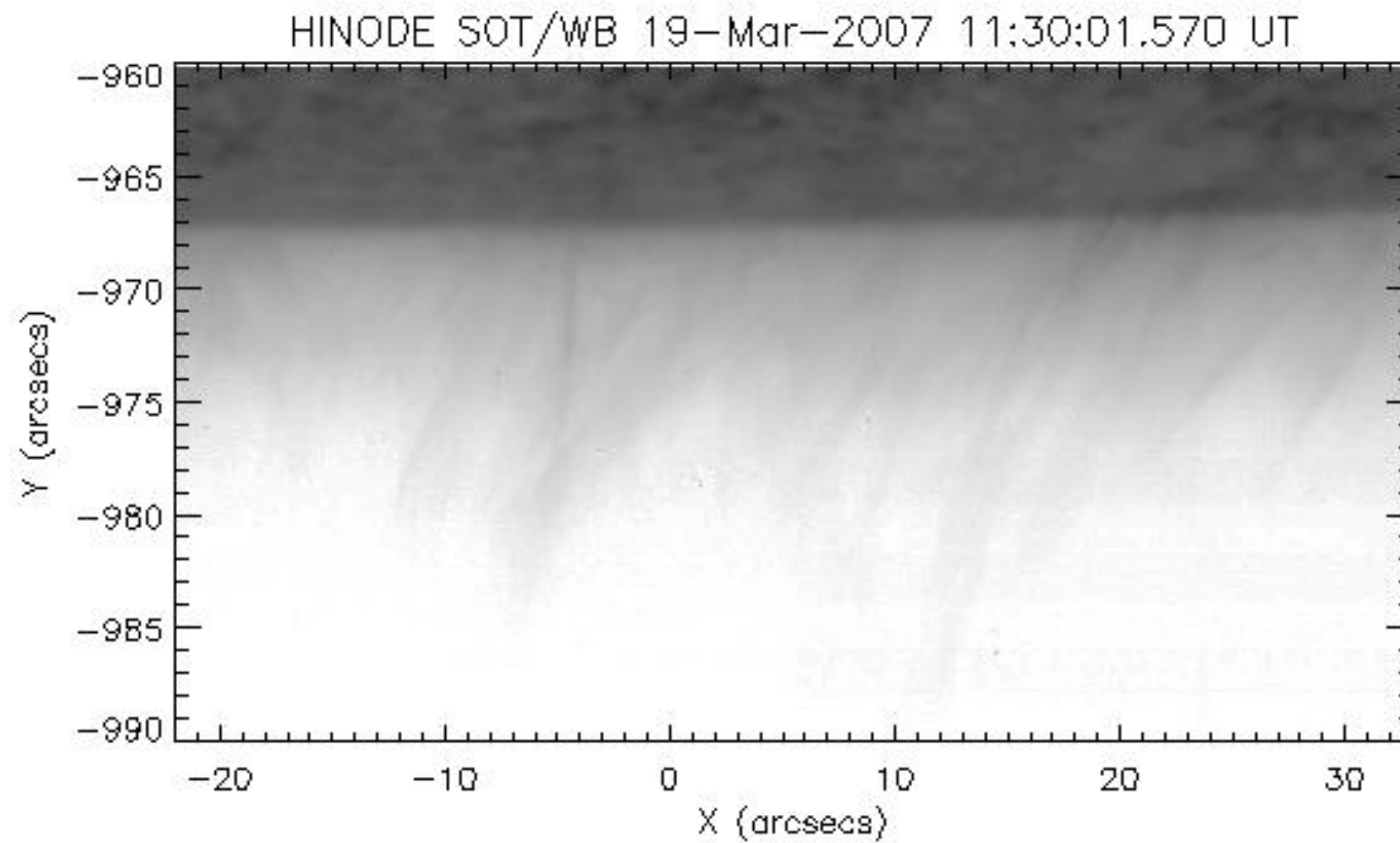
Panesar et al. (2018)

Spicules

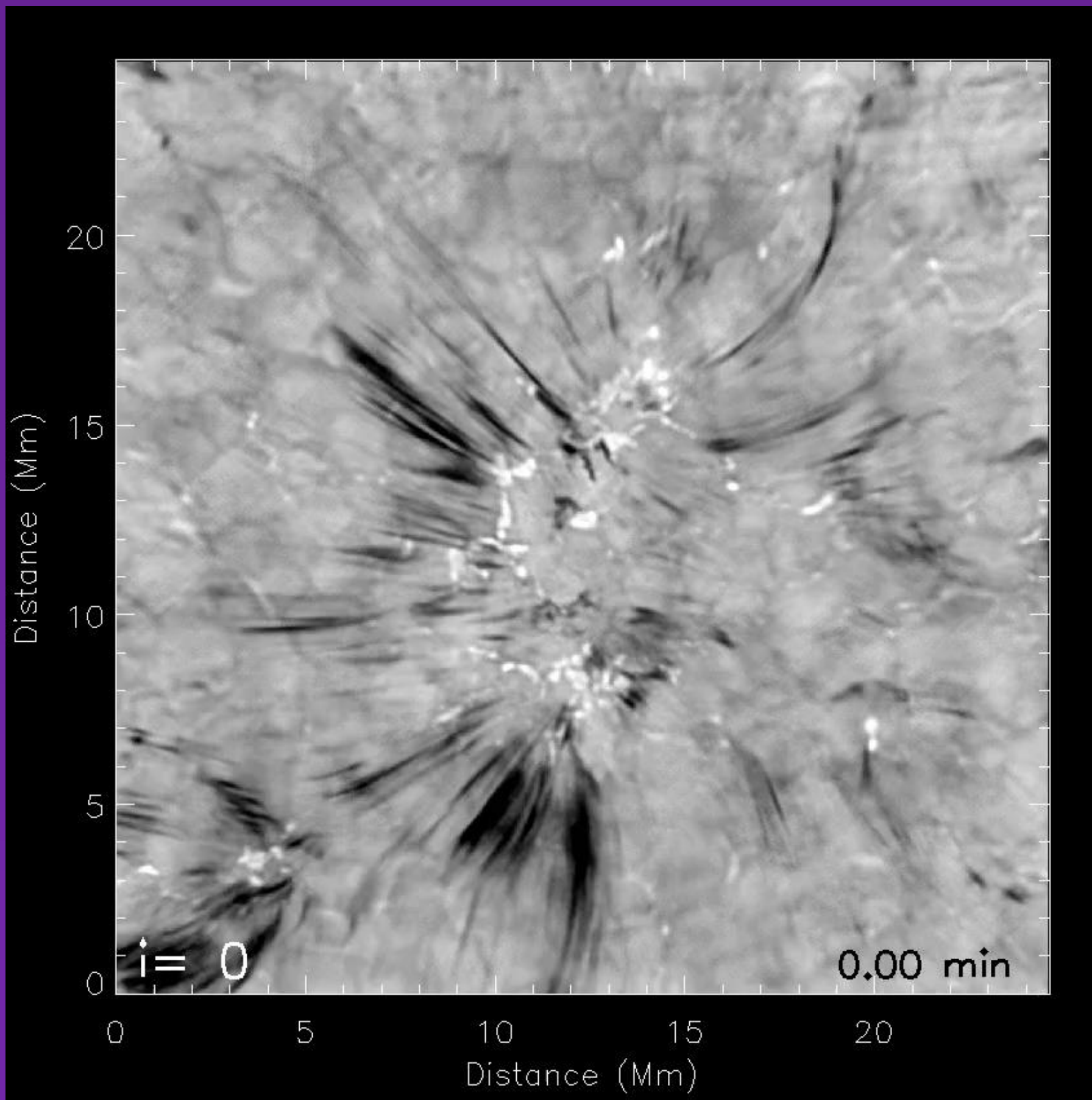
Hinode SOT/WB 21-Nov-2006 22:58:02.574 UT



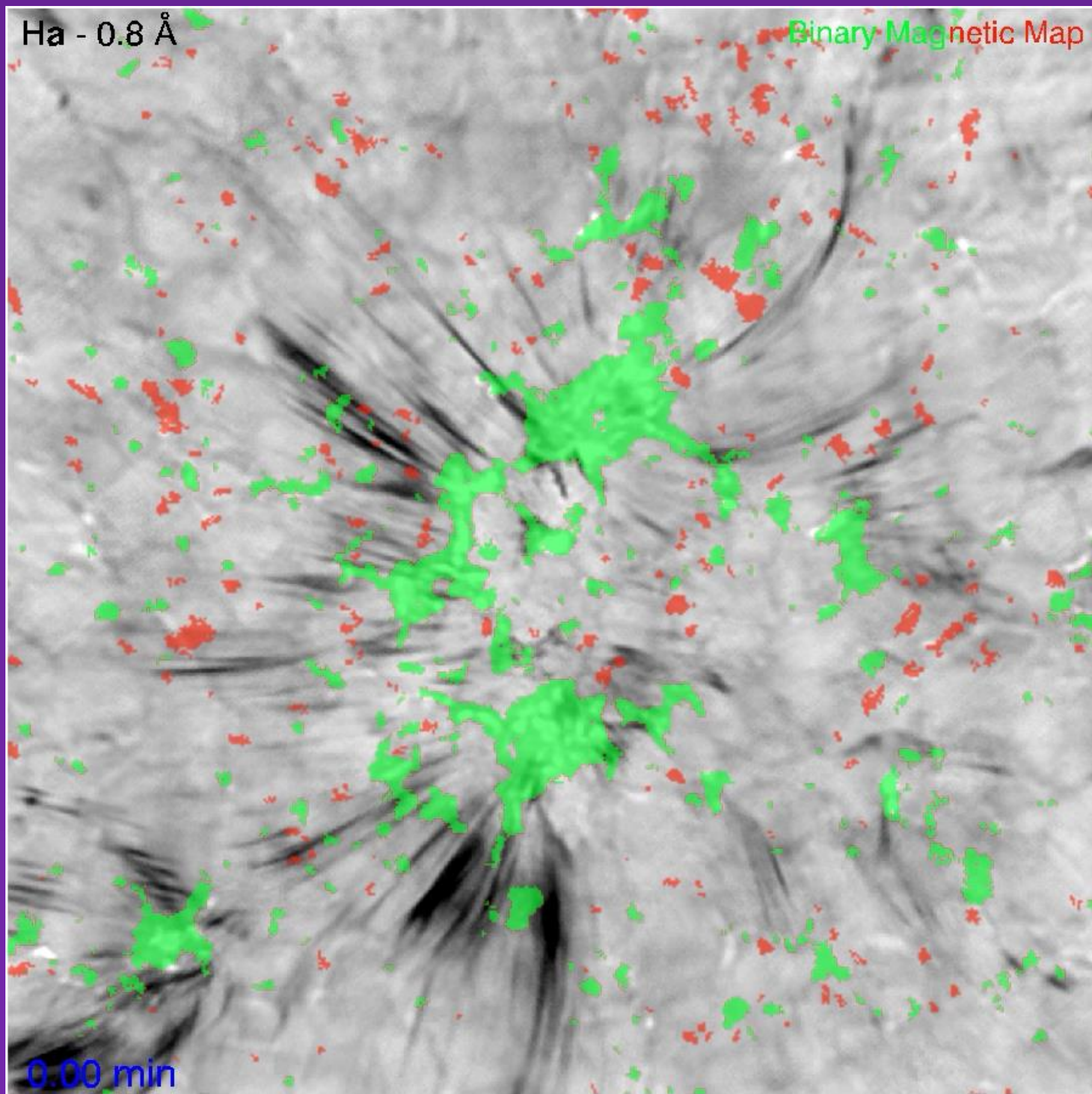
(Sterling 2021; Data set studied in Zhang et al. 2012²³)



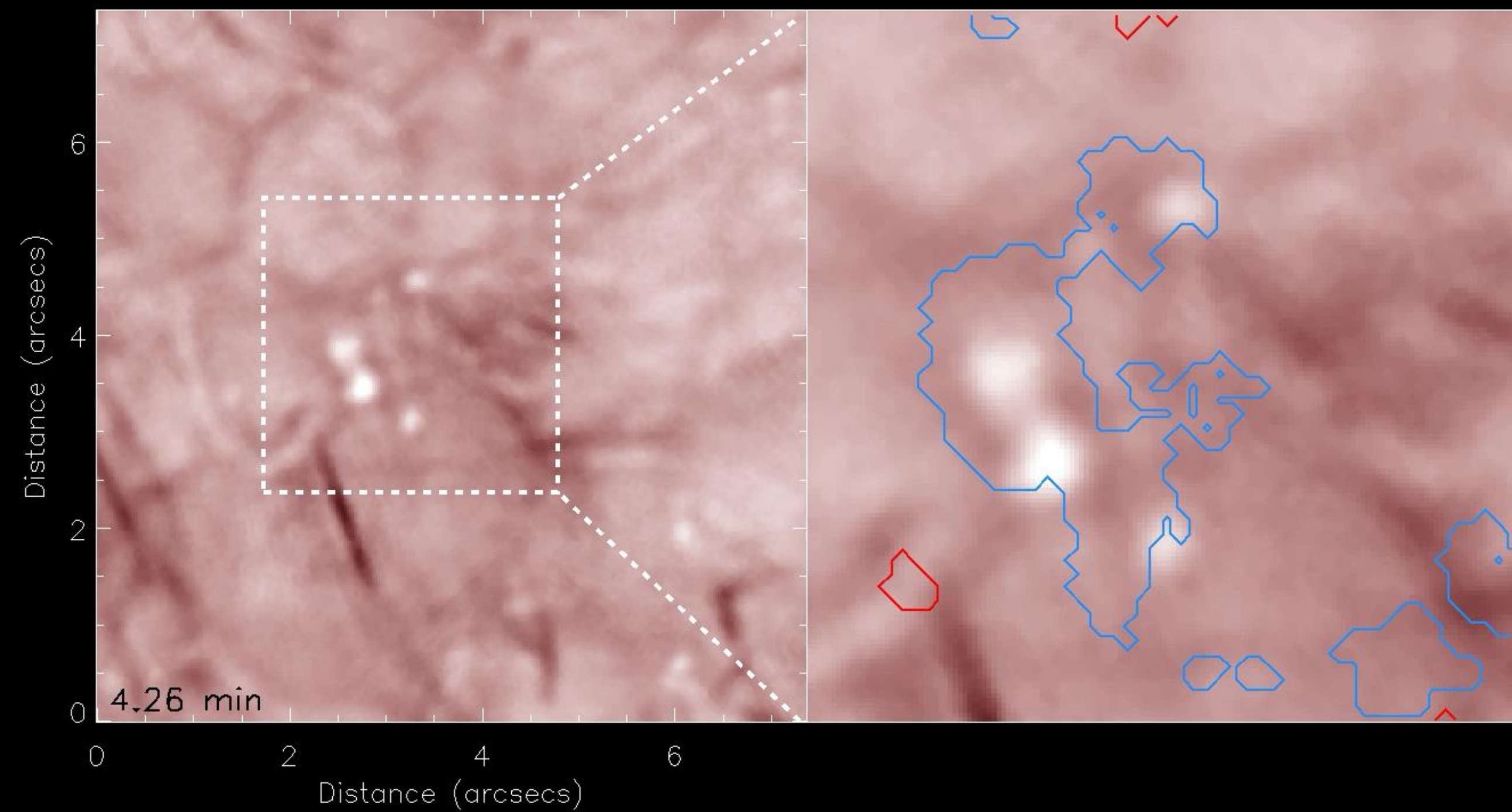
(Data set studied in Pereira et al. 2012)



(Sterling et al. 2020 [Samanta et al 2019])



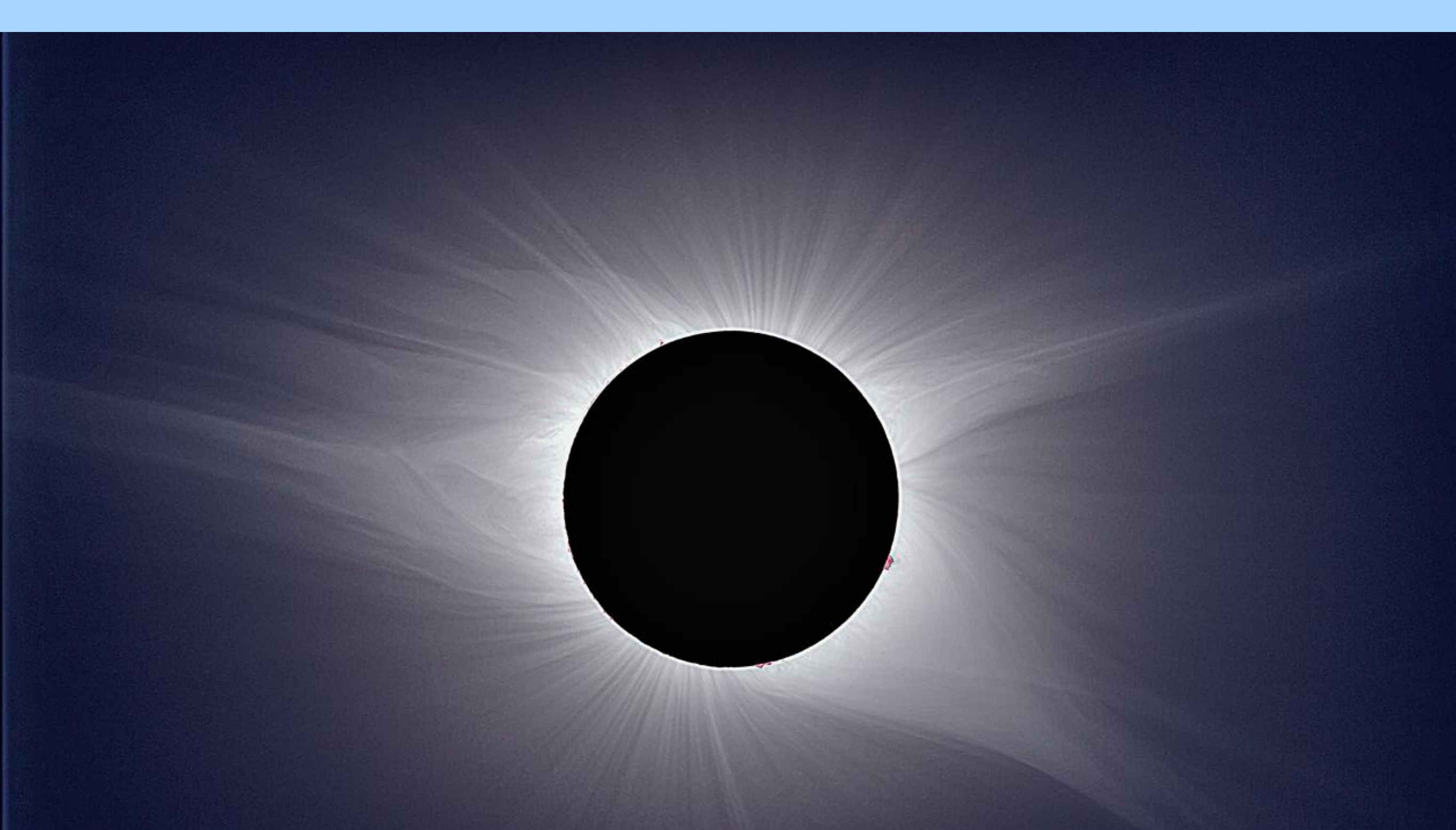
$H\alpha - 0.8 \text{ \AA}$



Some Outstanding Questions

- What causes jets? Strong evidence that it is flux cancelation in quiet Sun and CHs. Still must study more! (Shear only?? Kumar et al. 2018.)
- AR jets: Minifilaments sometimes less obvious (absent?). Also, “brightest” bright points sometimes in unexpected locations. (Result of complex field, multiple eruptions? Sterling et al. 2016, 2017.) Frequently see cancelation+emergence.
- How do jets scale to smaller structures (contribute to coronal heating? (Moore et al. 2015)). Some spicules? (Sterling et al. 2020.)
- If most jets result from flux cancelation, what about smaller jets, and spicules? (Samanta et al. 2019.)
- Role of *twist* in powering jets, and maybe spicules.

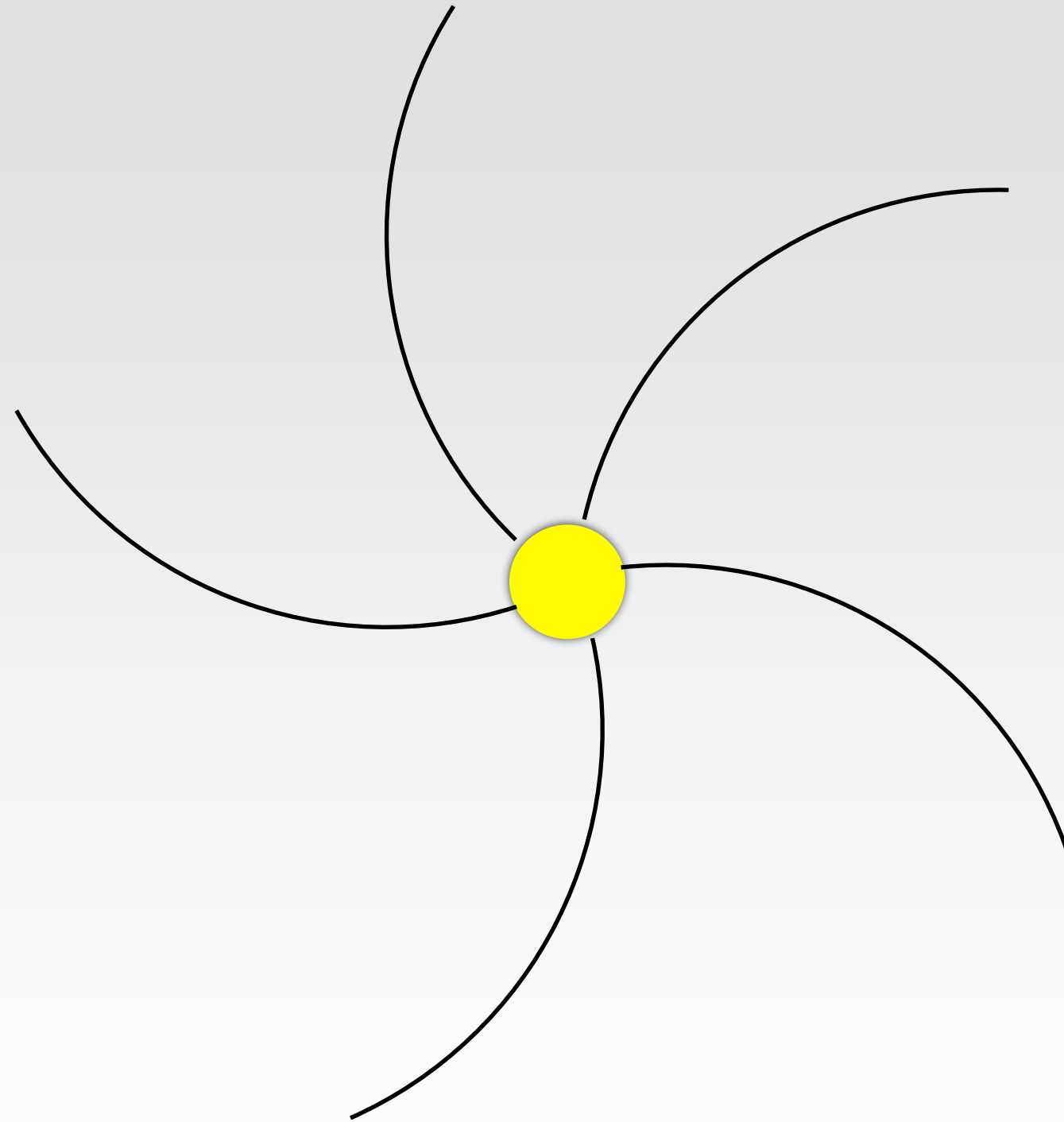
The Solar Wind, and “Switchbacks”

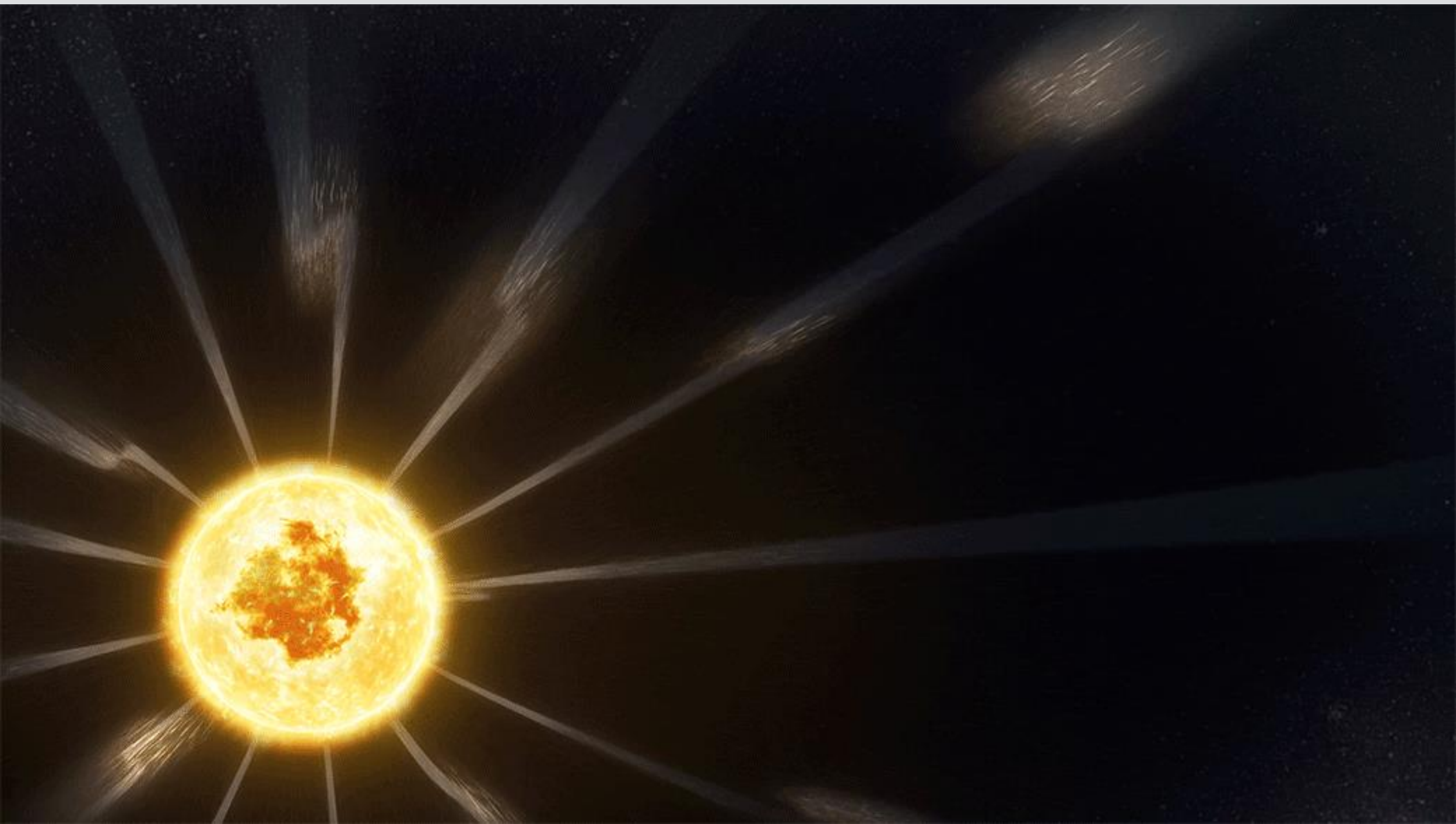


Alphonse Sterling, 21 August 2017
Lewisville, Idaho, USA
Takahashi FSQ106ed, f/5
Canon EOS700D, ISO 100; 1/1000, 1/500
1/250,1/125,1/60,1/30,1/15,1/8,1/4s

Corona – The Sun's outermost atmosphere

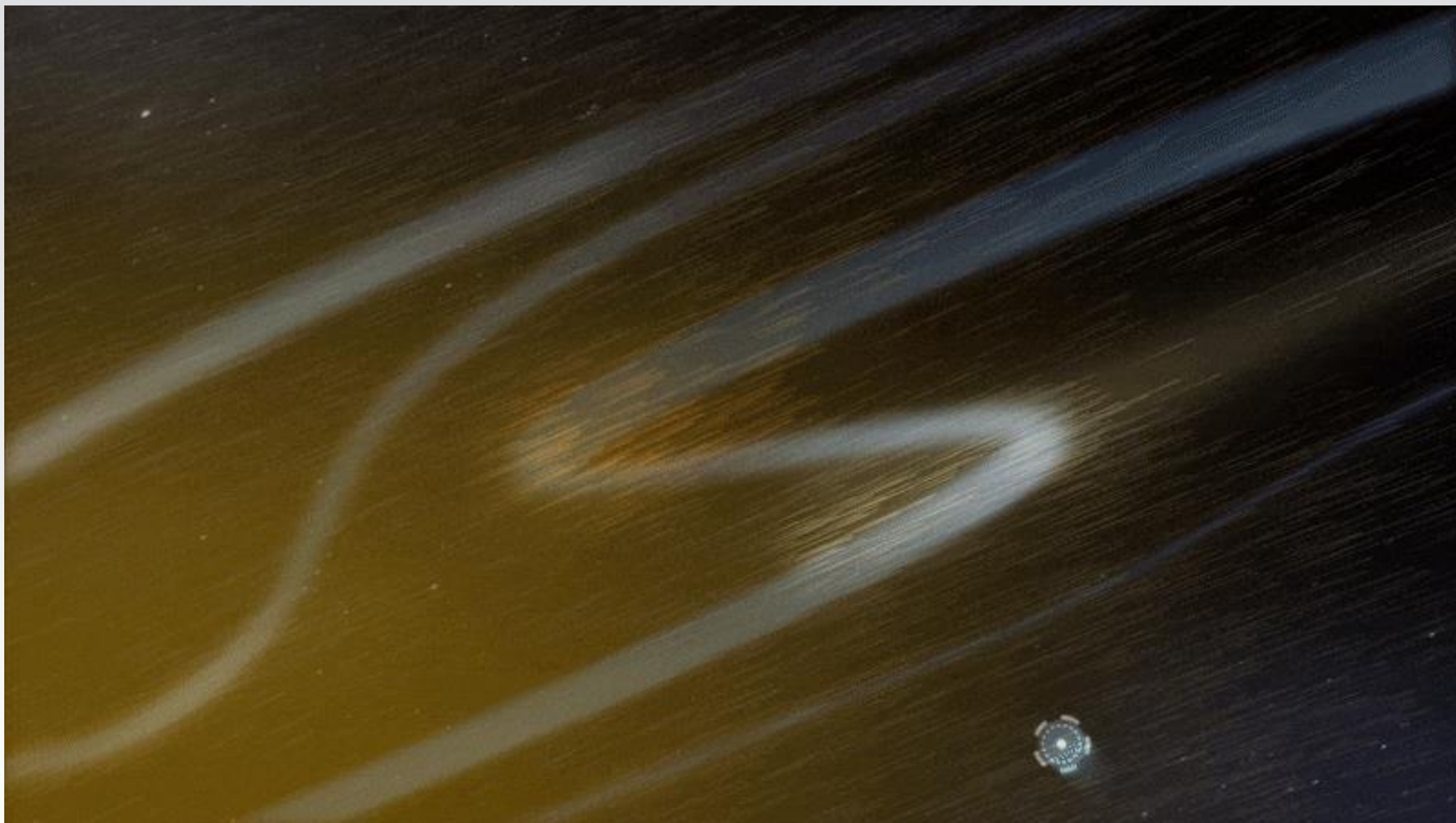
The Solar Wind, and "Switchbacks"





GSFC/Adriana Marique Gutierrez

<https://www.nasa.gov/feature/goddard/2021/switchbacks-science-explaining-parker-solar-probe-s-magnetic-puzzle>

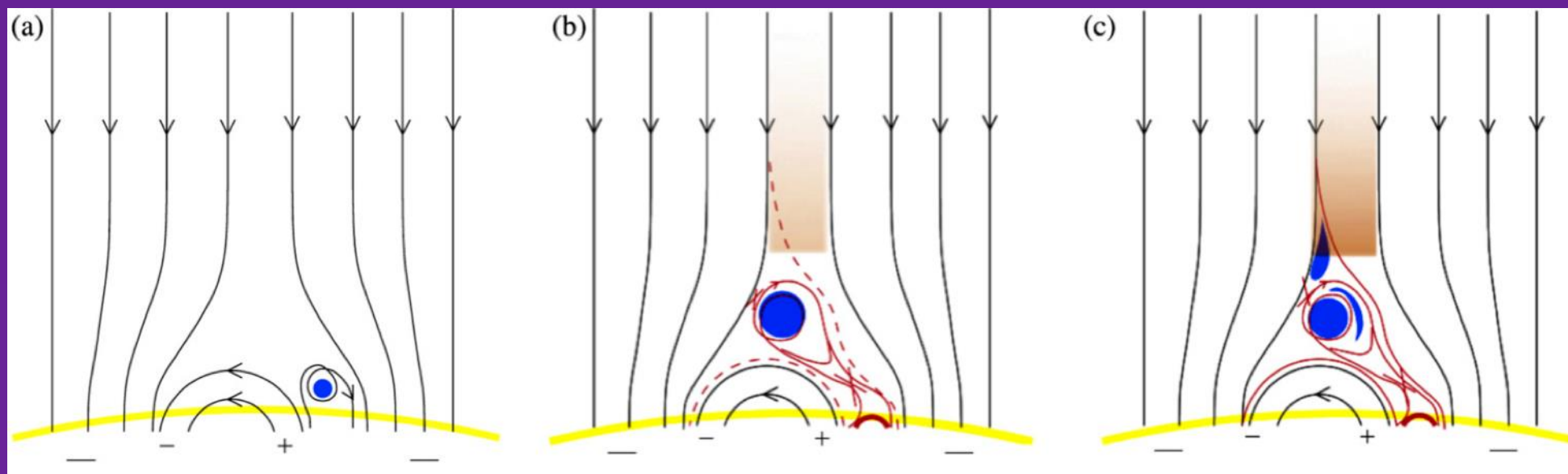


GSFC/Adriana Marique Gutierrez

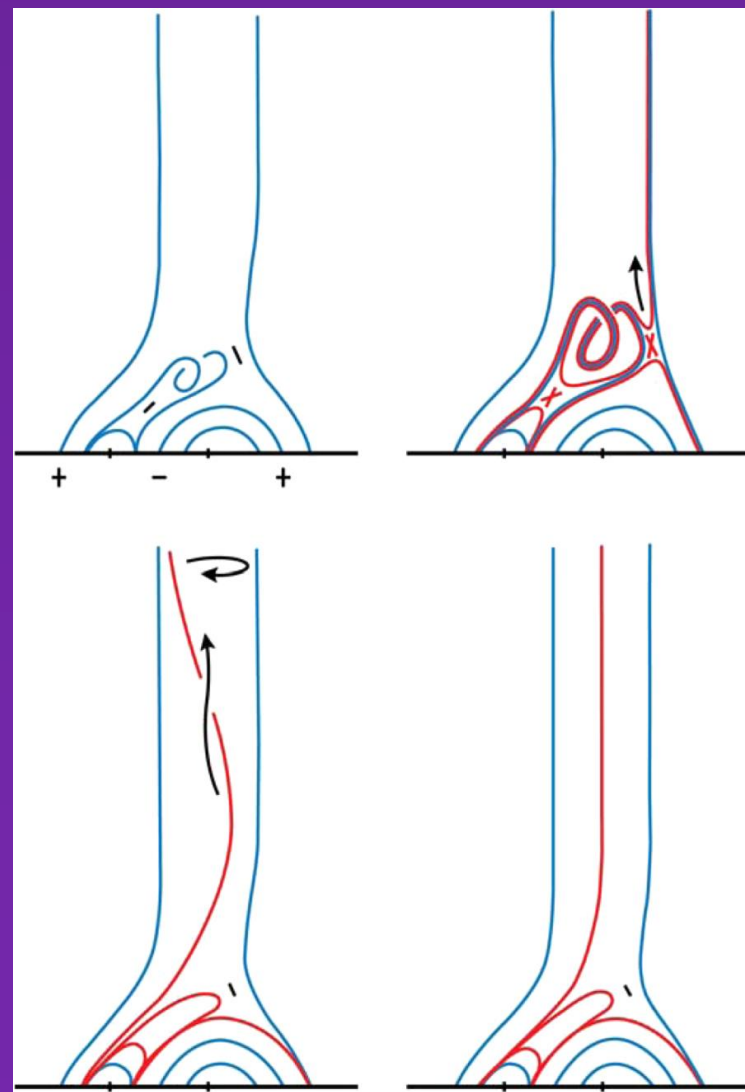
<https://www.nasa.gov/feature/goddard/2021/switchbacks-science-explaining-parker-solar-probe-s-magnetic-puzzle>

Parker Solar Probe (PSP) Switchbacks, and Coronal Jets

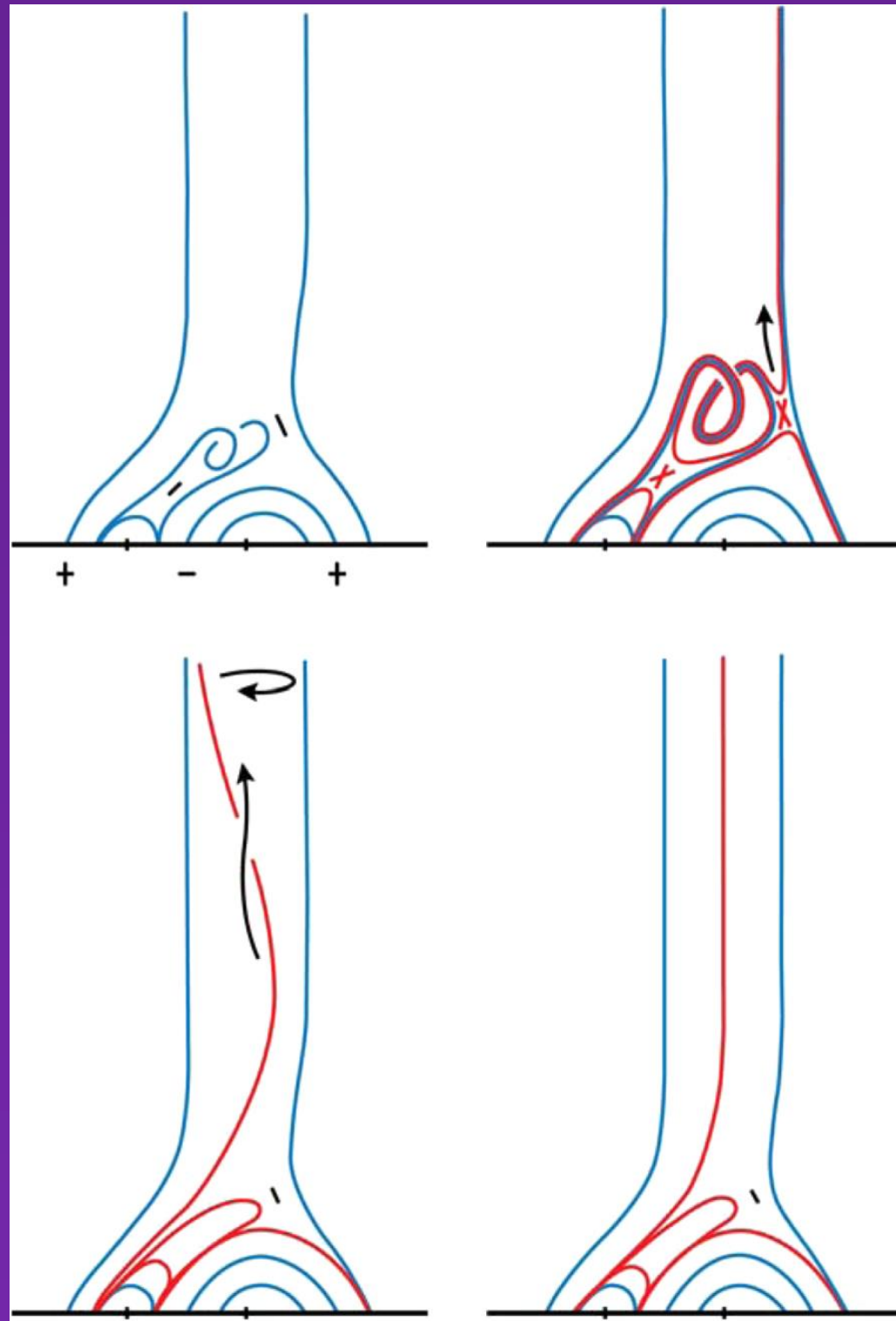
Sterling & Moore (2020) present an argument for how the PSP-observed switchbacks might result from coronal jets, based on the minifilament-eruption mechanism. (Also: Neugebauer & Sterling 2021.)



(Sterling et al. 2015)



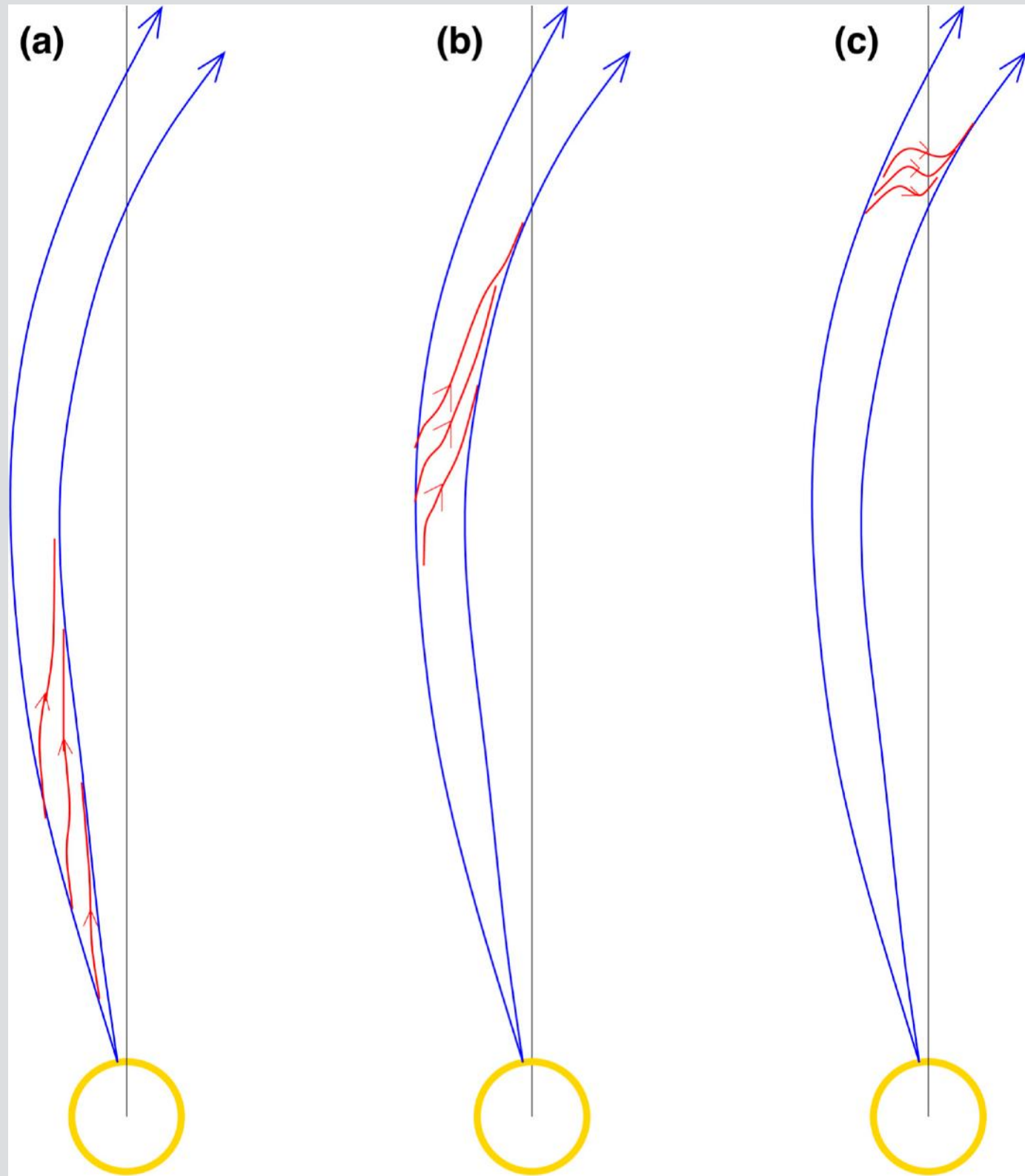
(Moore et al. 2015)



Moore et al. (2015). (Shibata & Uchida 1986-type mechanism; Patsouraos et al. 2008; Pike & Mason 1998; Yang et al. 2019.)

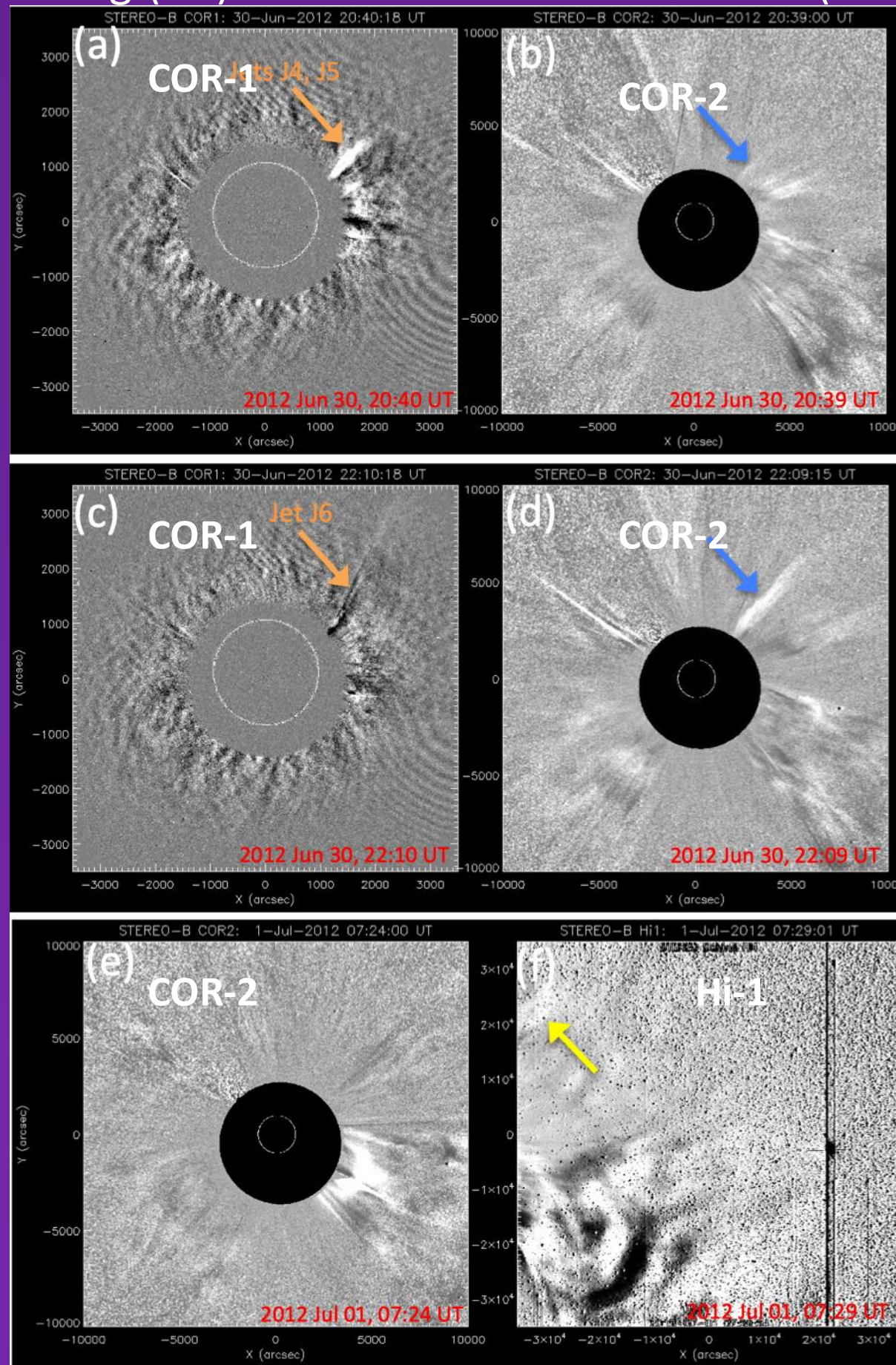
That packet would propagate out along the field at the Alfvén speed.

- That speed varies with distance from the Sun: ~ 1000 km/s in the corona, and ~ 100 km/s at $35 R_{\text{sun}}$.
- The wave packet, initially of size $\sim R_{\text{sun}}$ (Moore et al. 2015, Sterling & Moore 2020), will feel different Alfvén speeds at its front and near-Sun sides; this Alfvén-speed gradient will result in the wave packet becoming more compact with distance from the Sun.
- The result will be a feature that PSP could detect as a switchback.



Sterling & Moore (2020)

Far-Reaching (AR) Coronal Jets from STEREO-B (and AIA)



It is plausible that the jet mechanism occurs on smaller size scales, with increasing frequency, with “jetlets” (widths ~ 4000 km, Raouafi & Stenborg 2014; Panesar et al. 2018), and maybe even some spicules (widths $\sim \text{few} \times 10^2$ km; Sterling & Moore 2016).

If all can become switchbacks, the rates would be much larger than then the few $\times 100$ /day from canonical coronal jets.

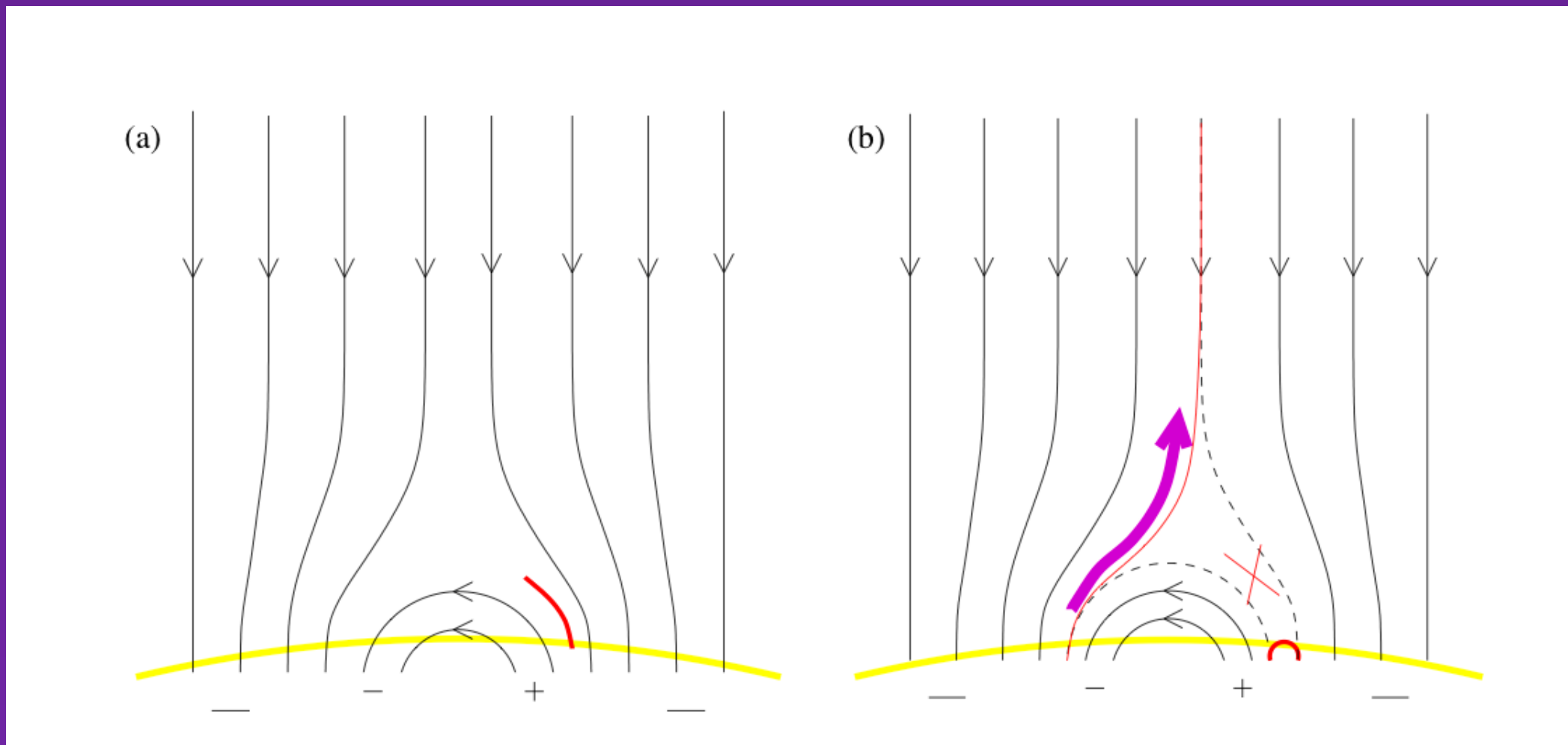
Summary

- Approaching a good understand of coronal jets, especially in QS and CHs: At least many jets are miniature filament eruptions triggered by flux cancelation. AR jets are likely similar, but altered due to the more-complex AR field.
- Needed: Jet simulations based on minifilament eruptions (e.g. Wyper et al), and *flux cancelation*!
- How does jet physics scale to different sizes? (Large eruptions? ``Jetlets''? Spicules??)
- Might jets (and/or jet-like features) be the source of the PSP-observed switchbacks?



Image:
Alphonse Sterling
21 August 2017,
Lewisville, Idaho

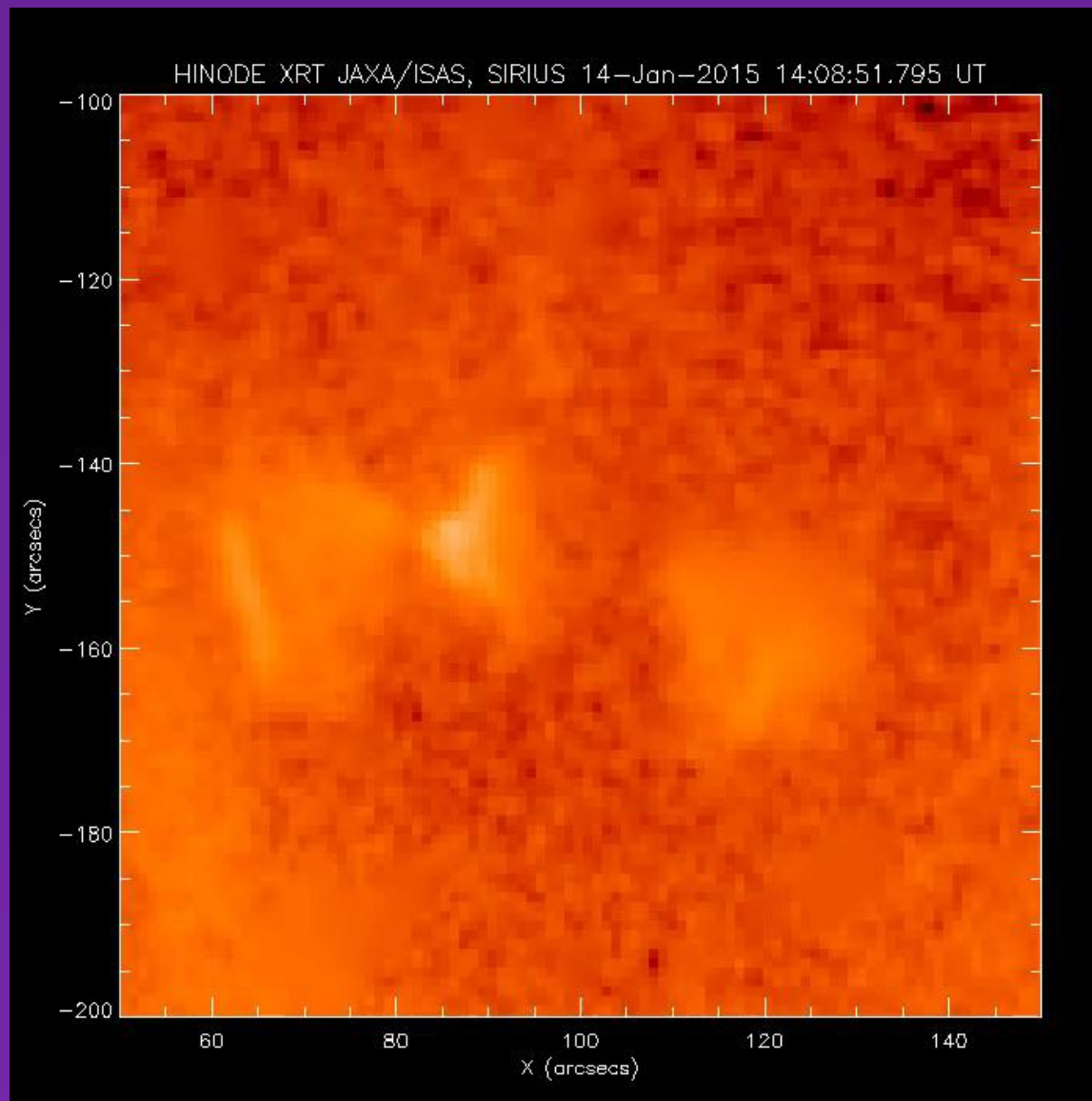
Emerging-Flux Model for (X-Ray) Jets



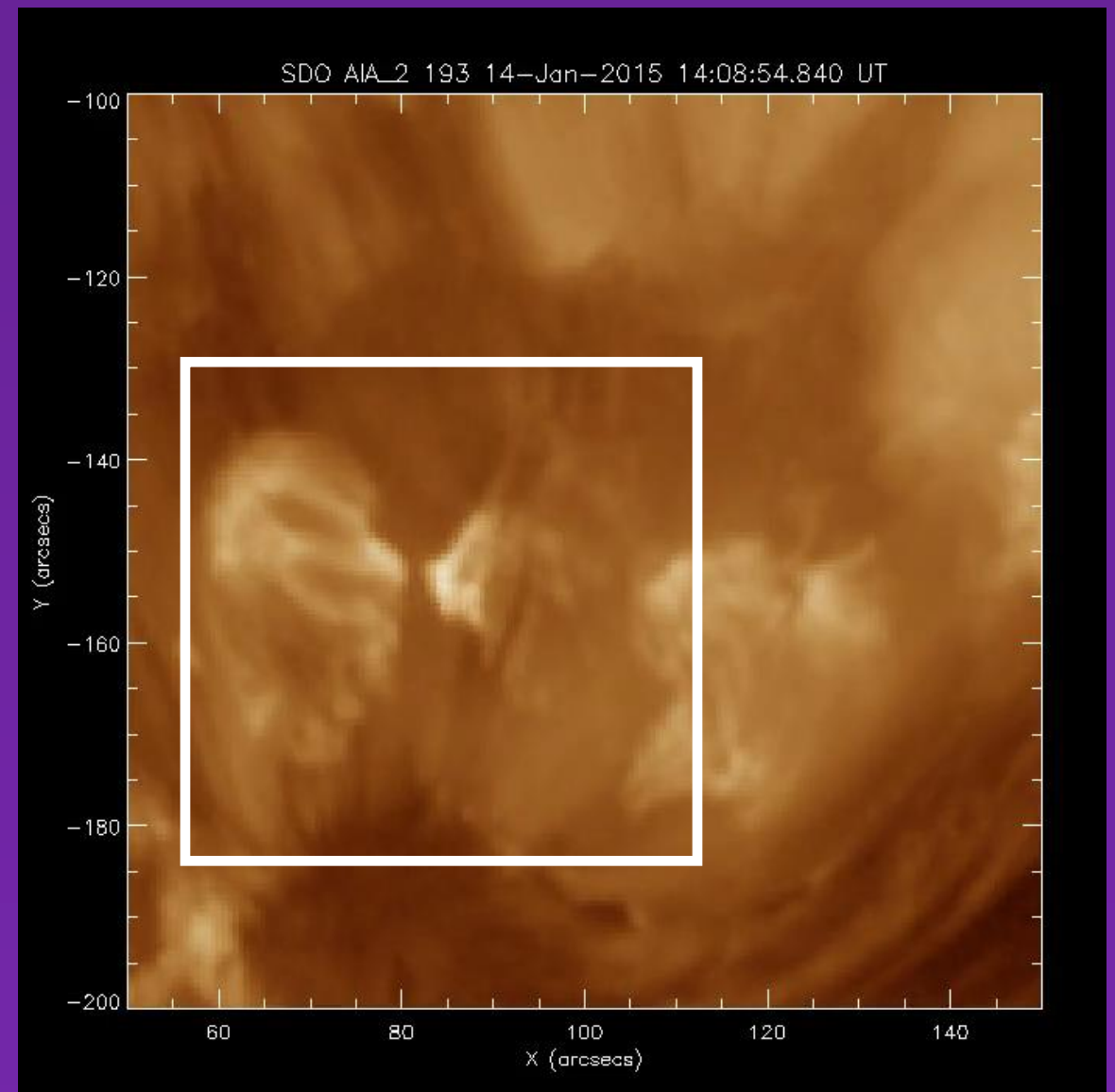
Supported by numerical simulations: Yokoyama & Shibata (1995), Nishizuka et al. (2008), Archontis et al. (2013), Moreno-Inertis et al. (2013), Fang et al. (2014), etc. (Cf. Heyvaerts, Priest, & Rust 1977.)

An Example: AR Jets

- 14 Jan 2015 (NOAA AR 12259), AIA, HMI, Hinode, IRIS.
- Sterling et al. (2017)



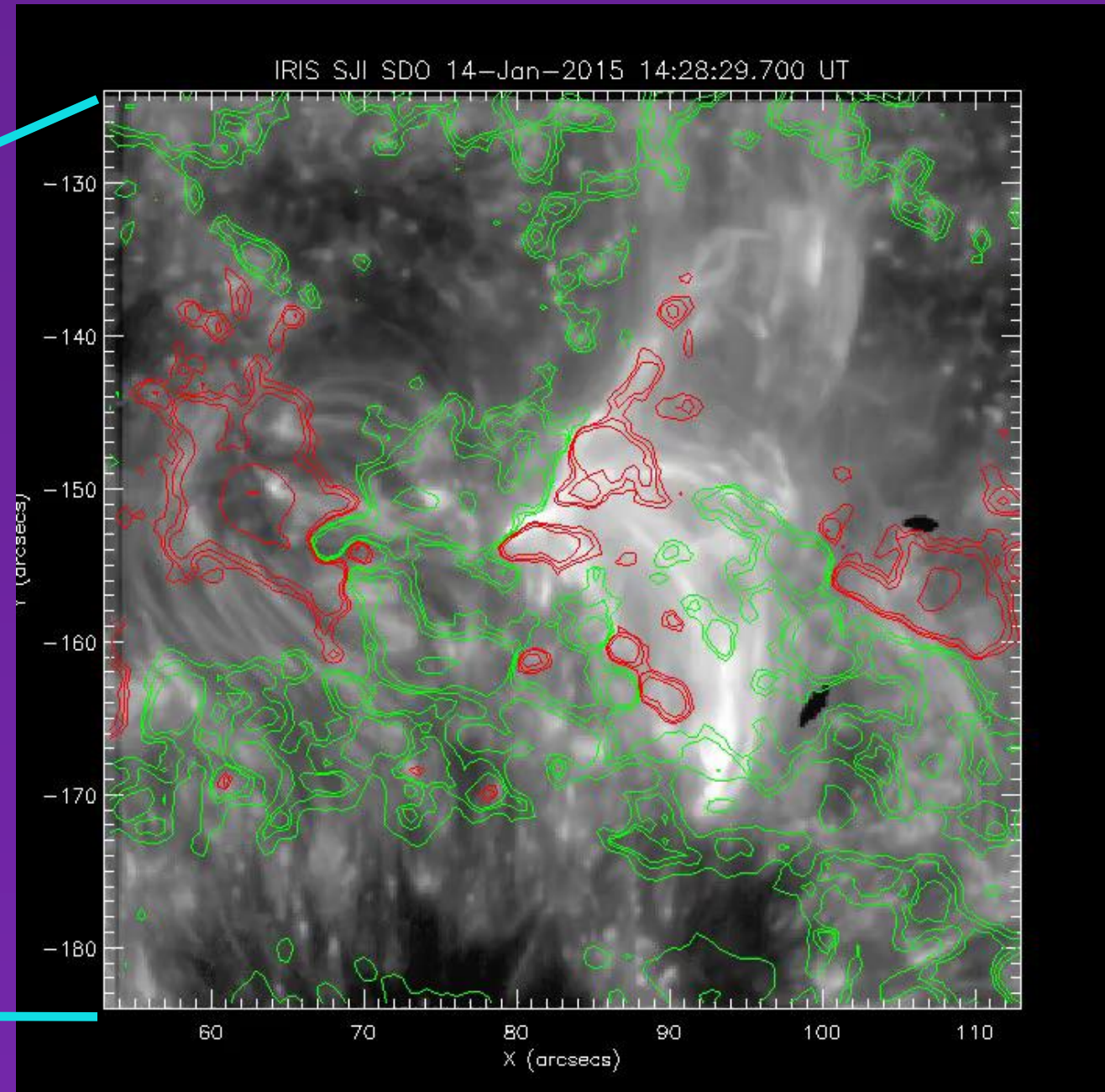
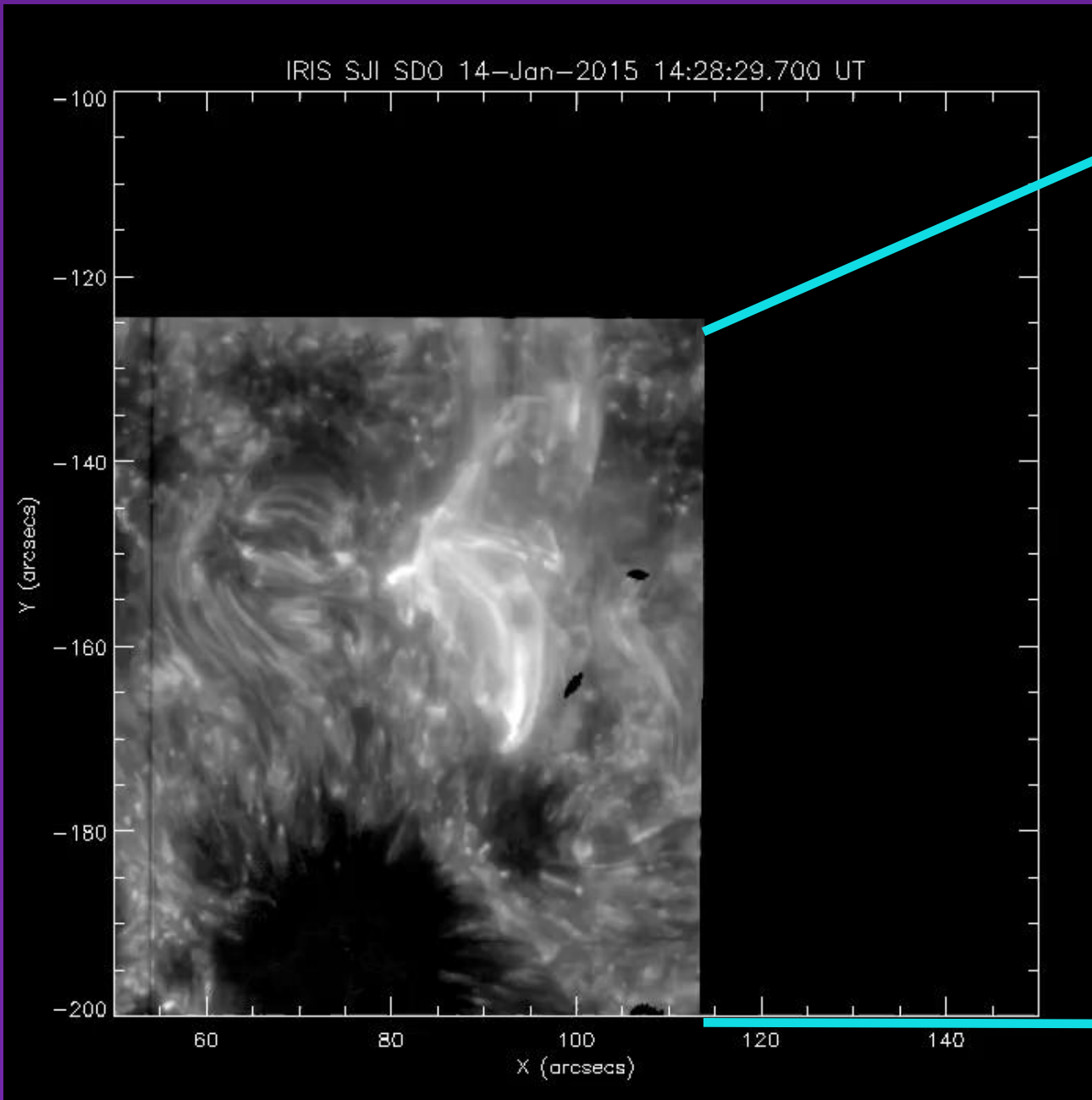
Hinode/XRT



AIA 193

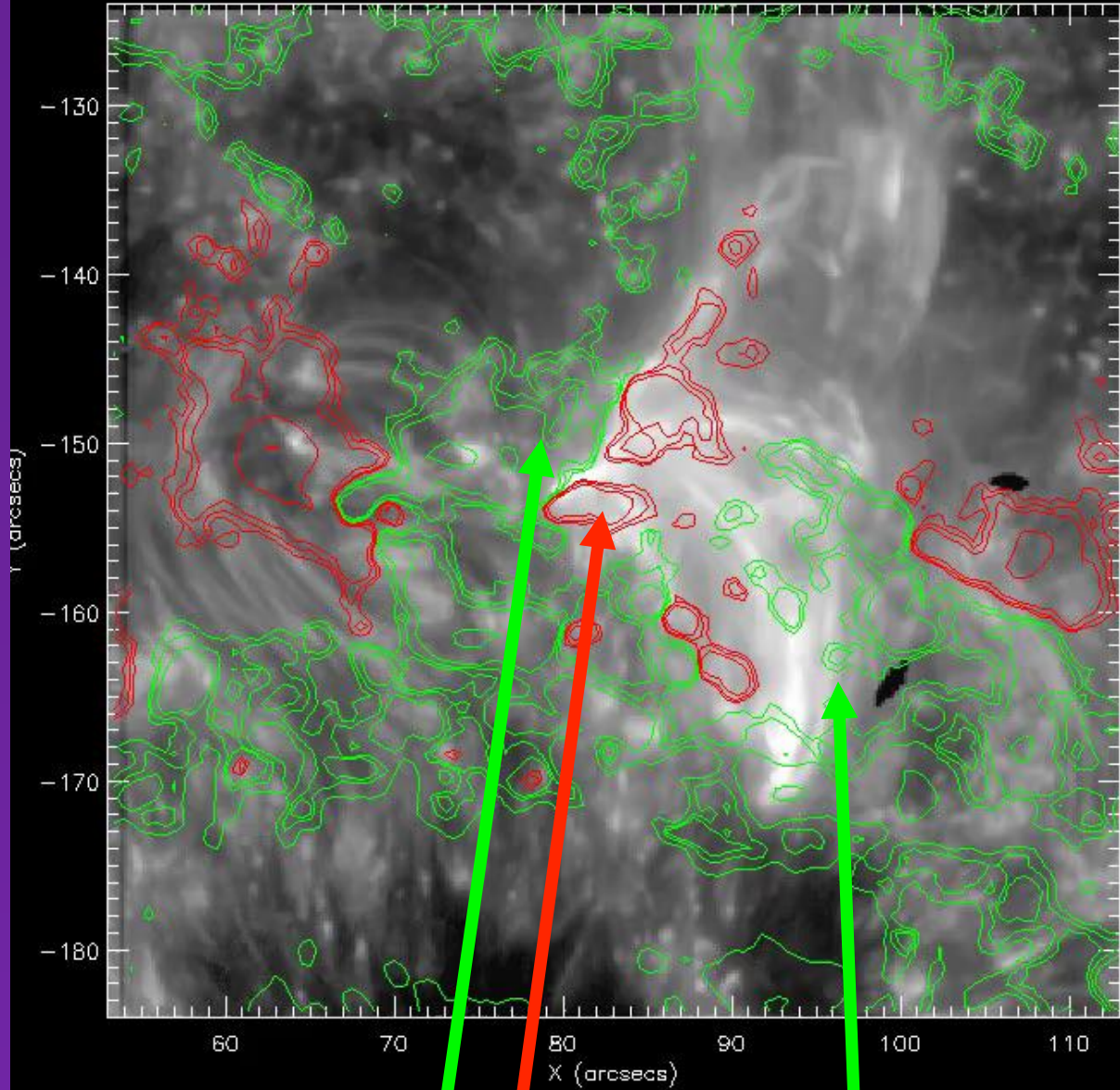
Minifilament hard to see (absent?). Work the same way??

Coronal Jets in Active Regions

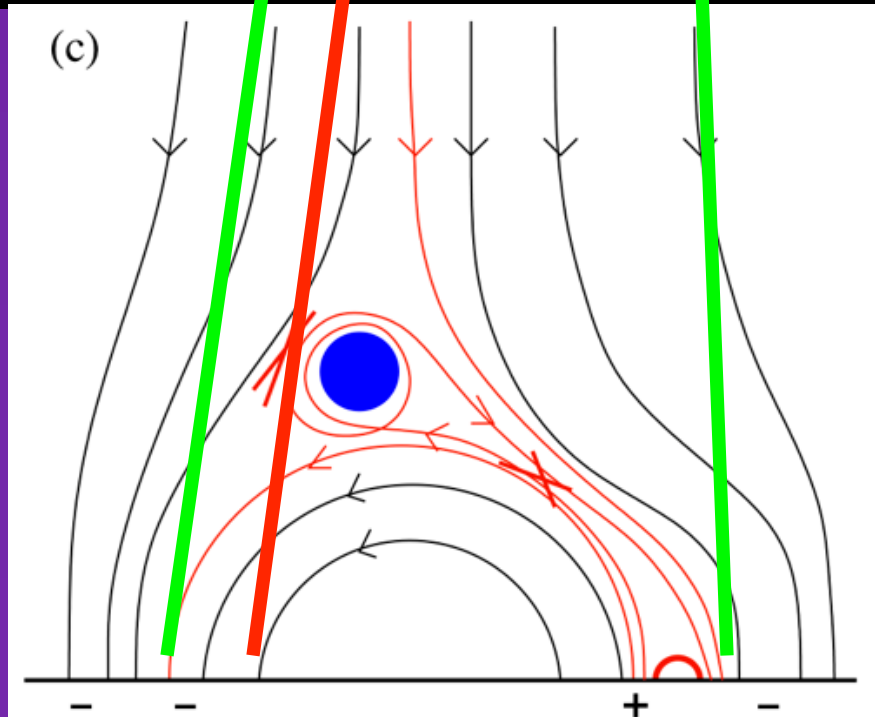
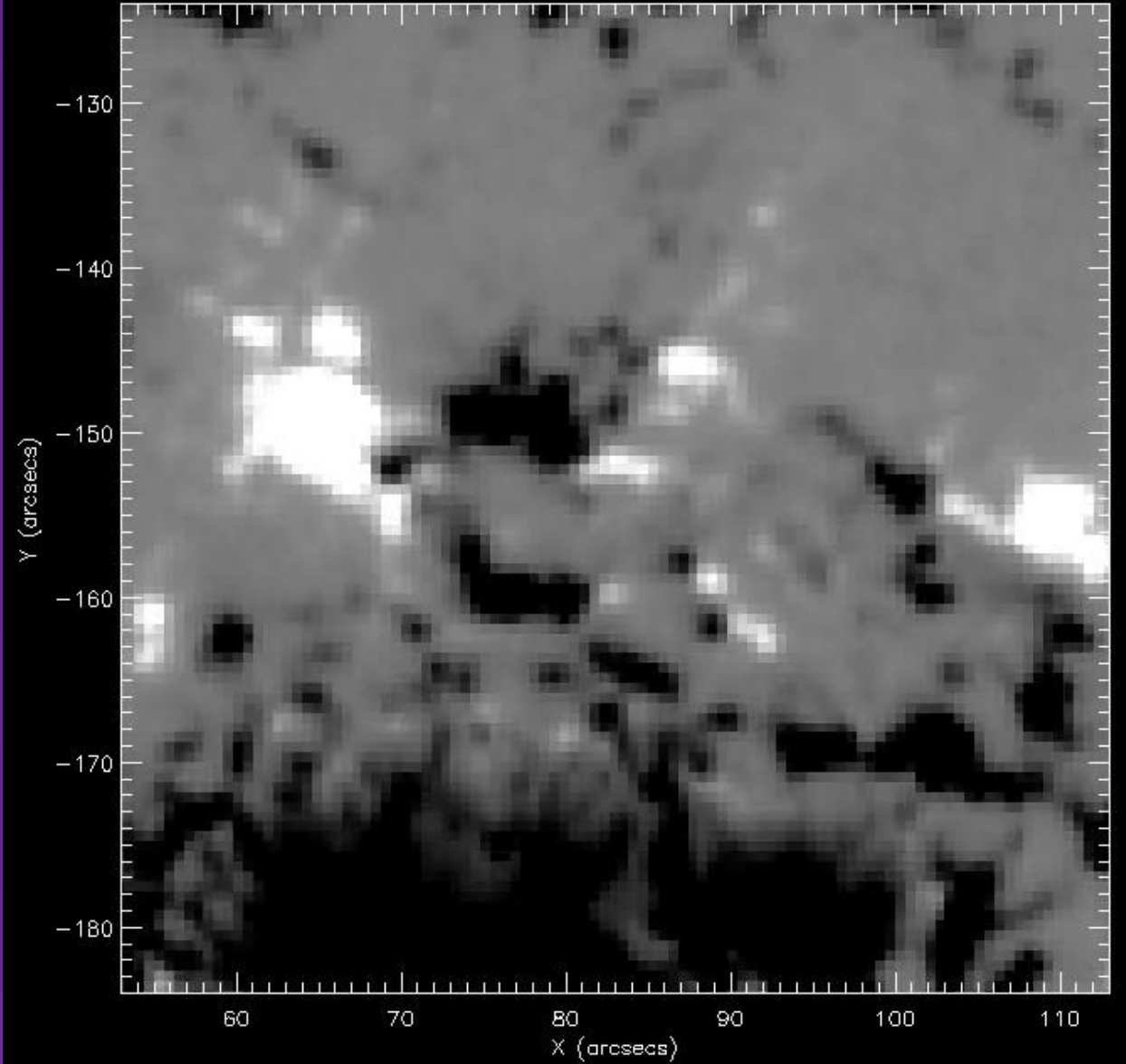


Sterling et al. (2017)

IRIS SJI SDO 14-Jan-2015 14:28:29.700 UT

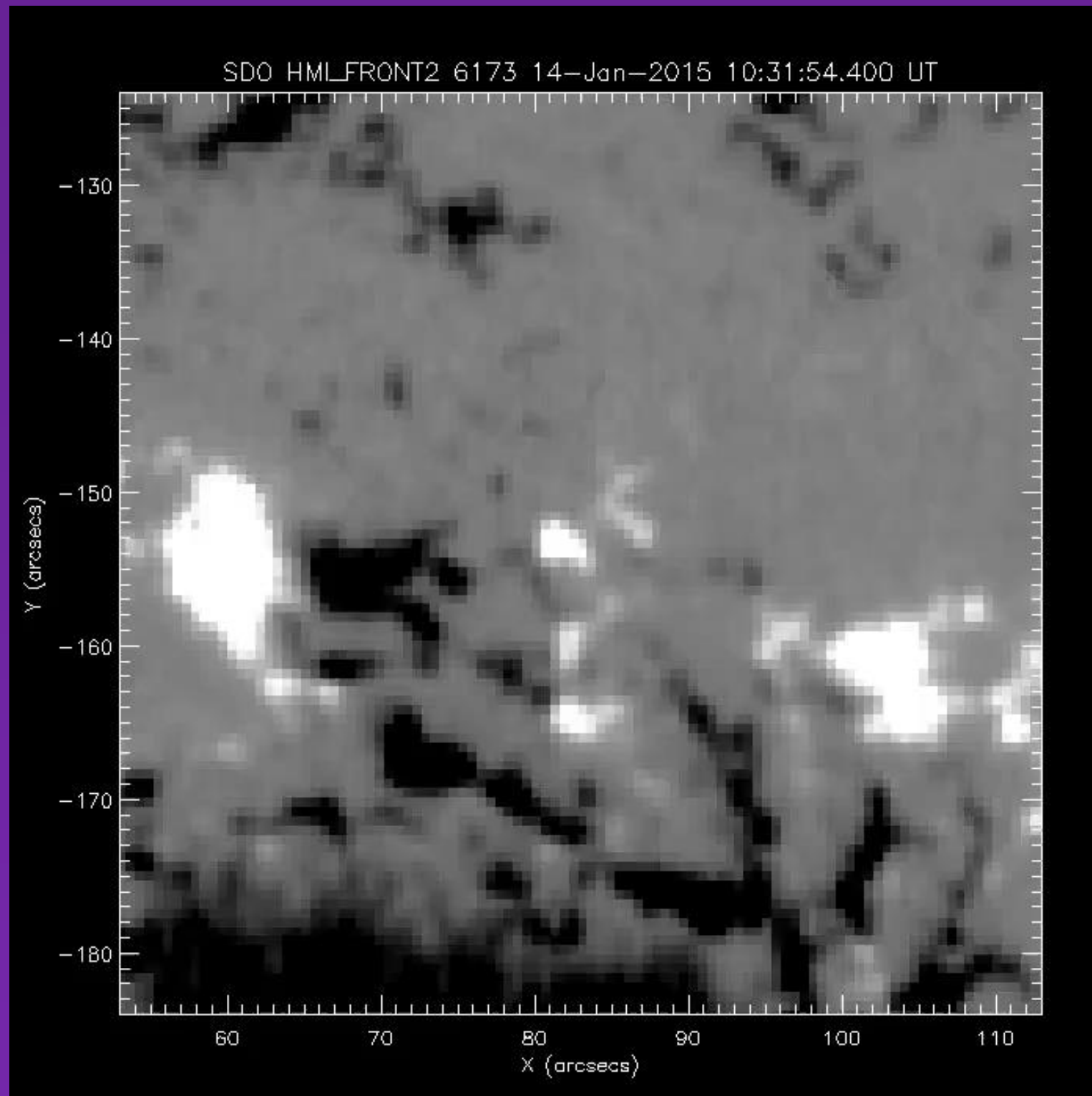


SDO HMLFRONT2 6173 14-Jan-2015 14:29:39.300 UT



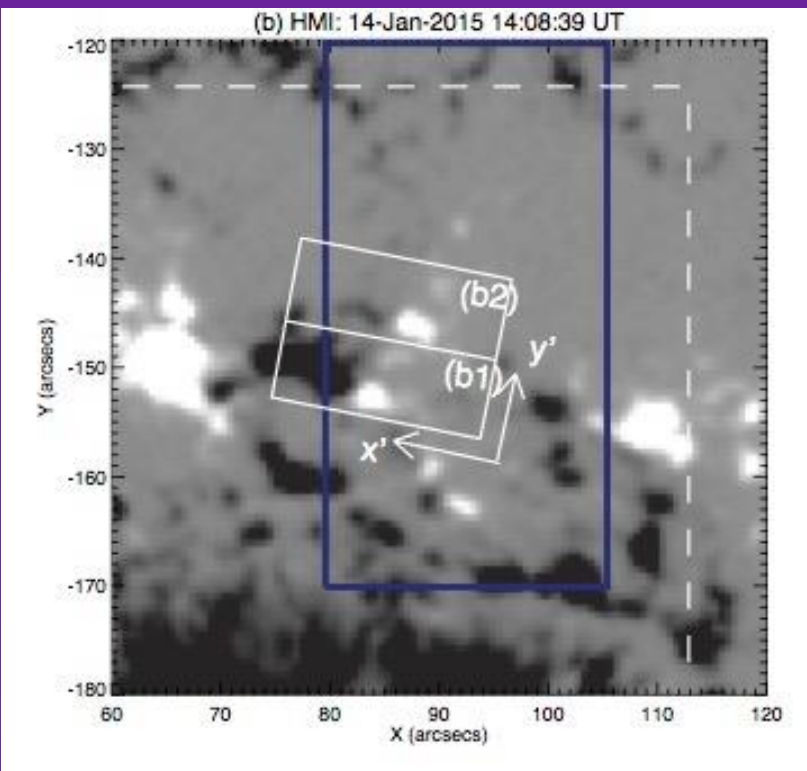
Sterling et al. (2017)

HMI of jetting region

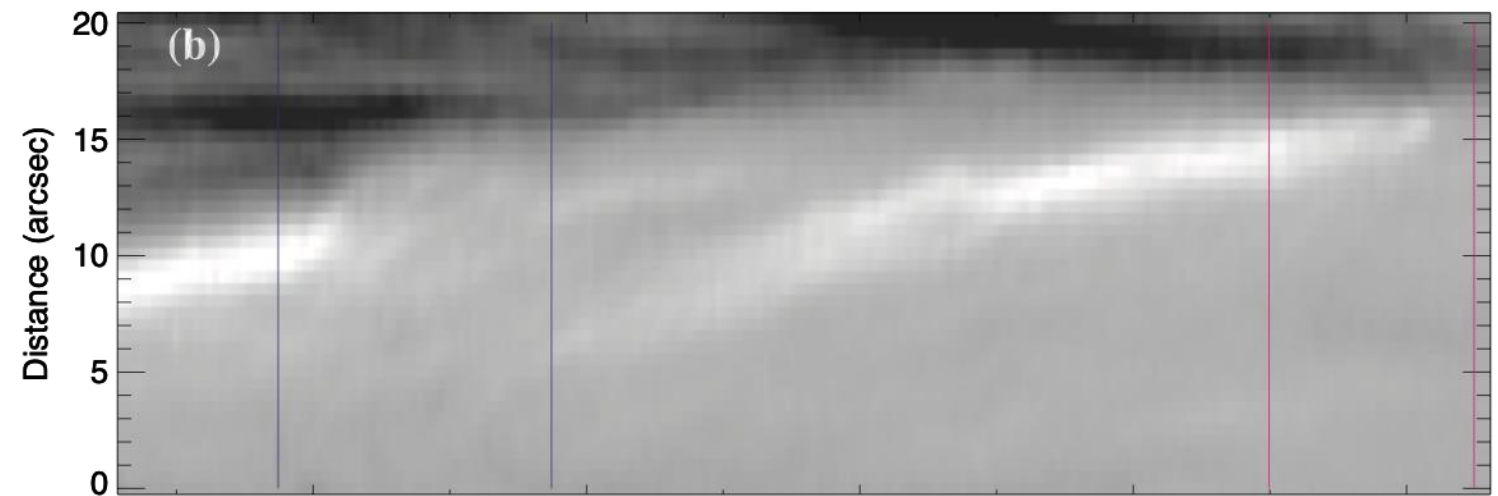
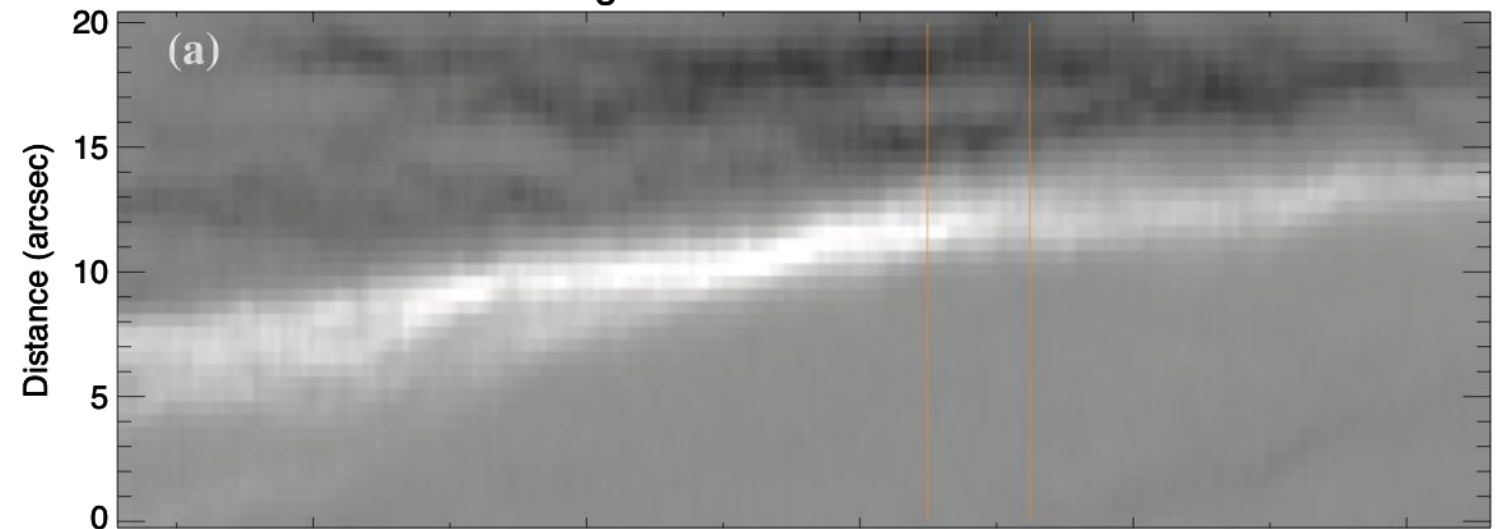


Jets occur at *flux cancelation* locations!

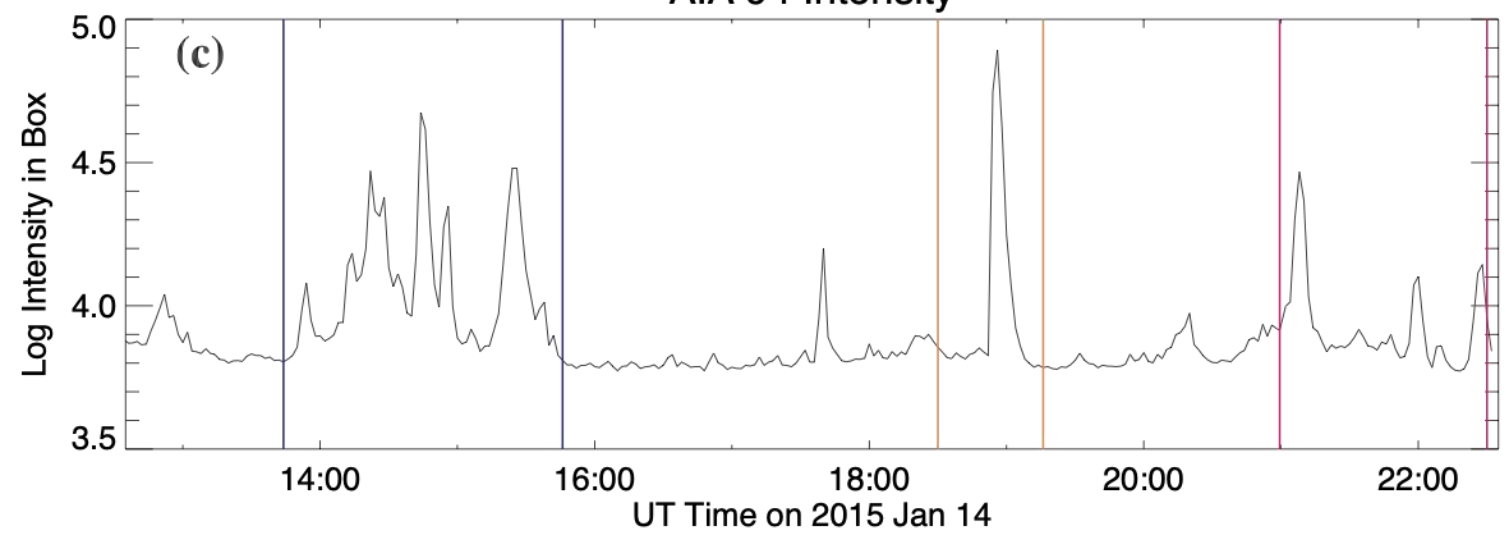
AR jets (Sterling et al. 2017)



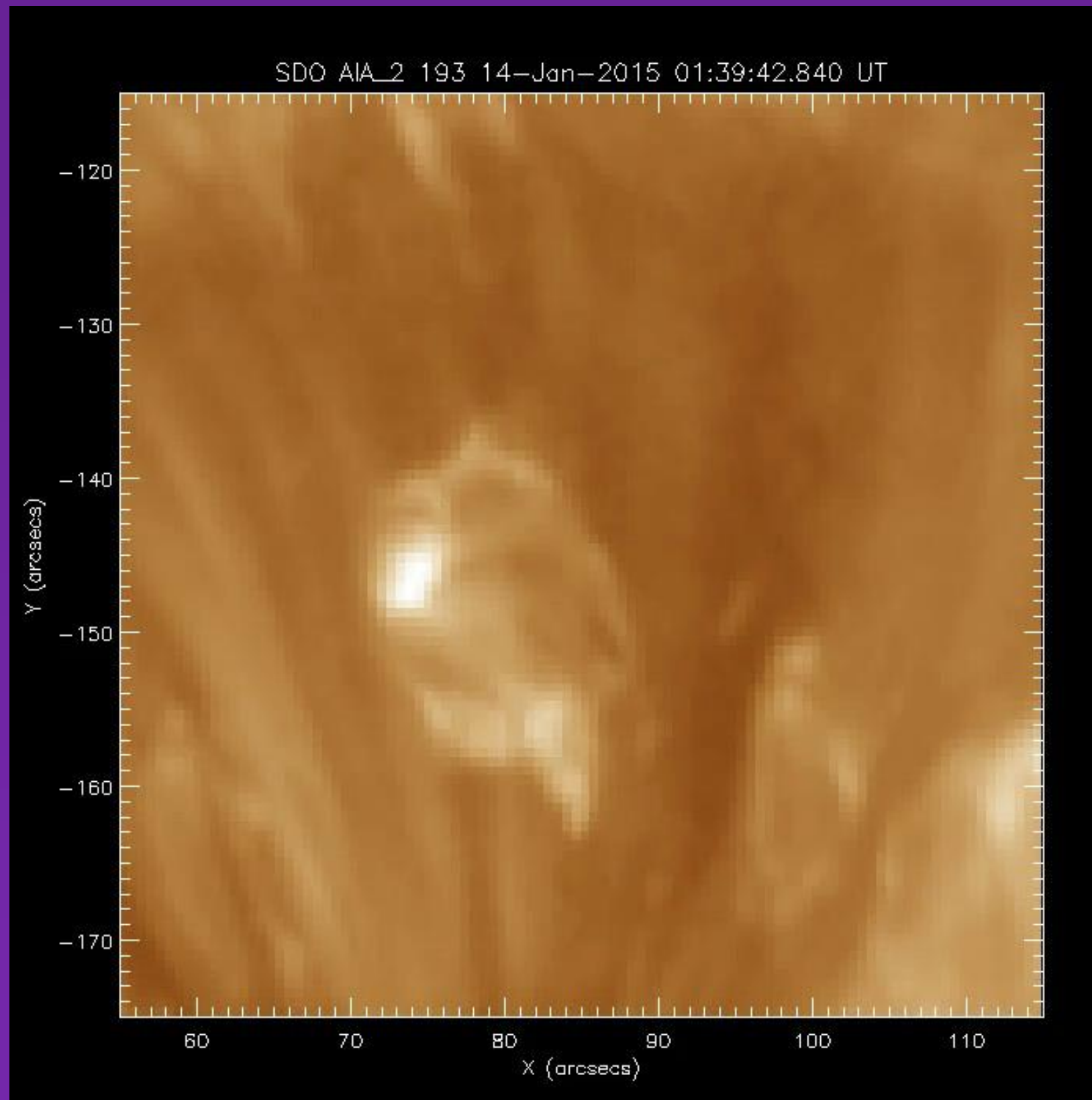
Magnetic Evolution with Time



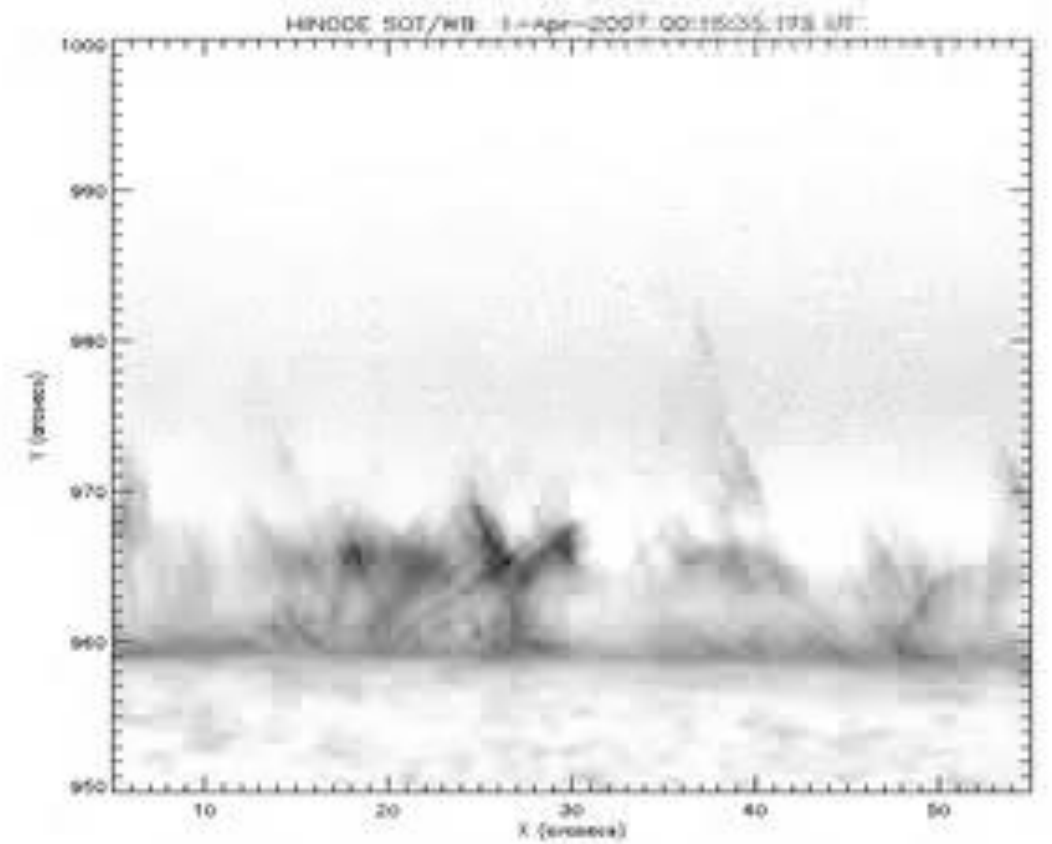
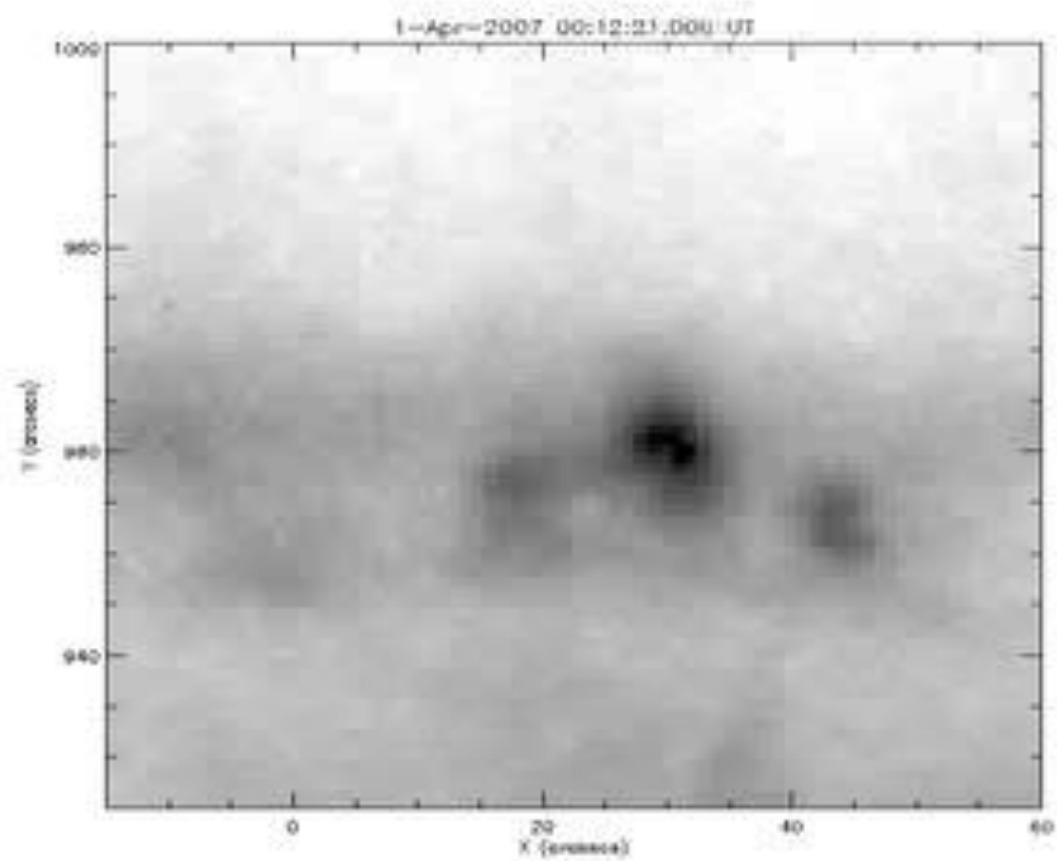
AIA 94 Intensity



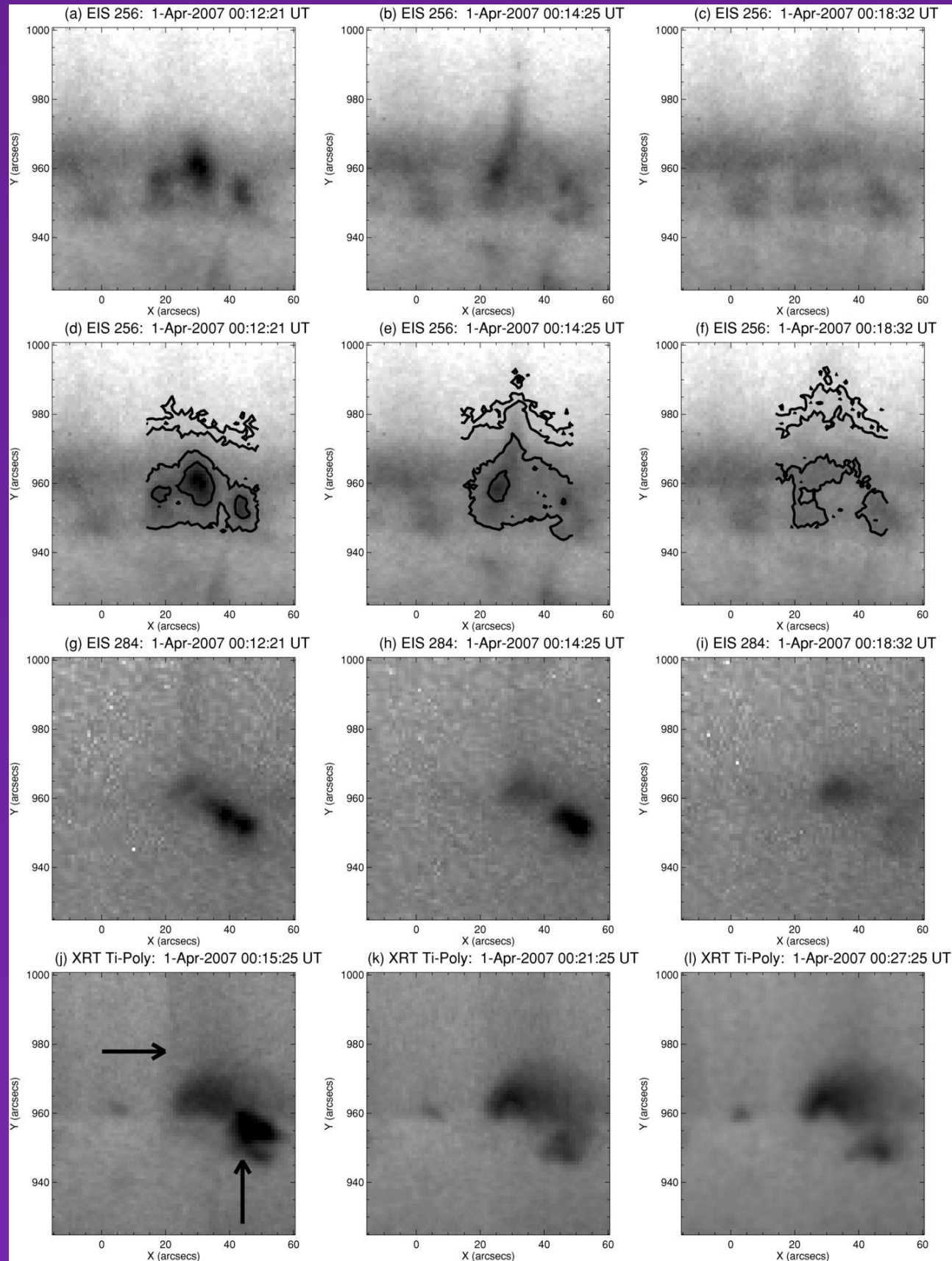
Minifilament “strand” visible from neighboring region, slightly different time.



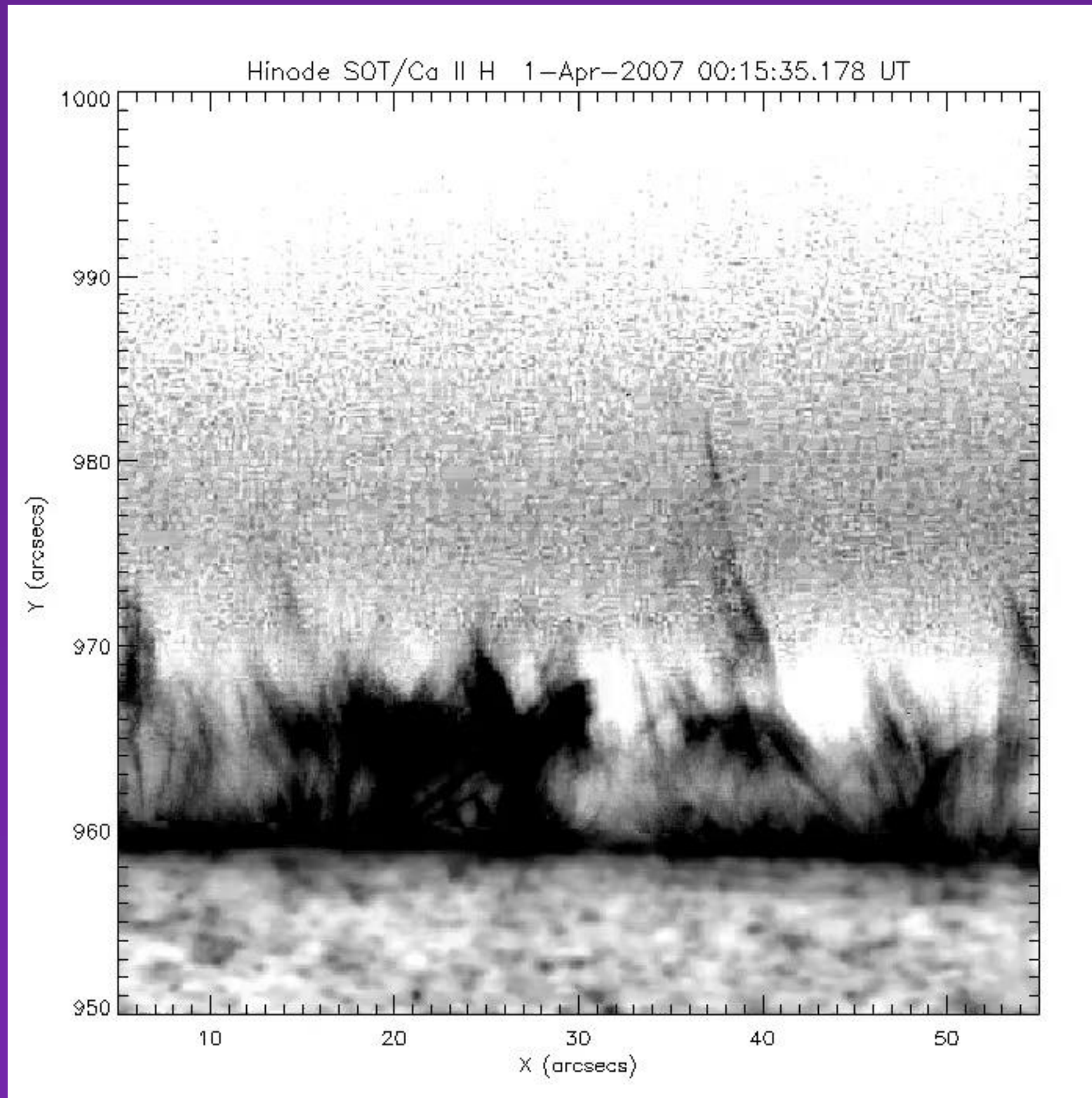
AIA 193



(Sterling, Harra, & Moore 2010)

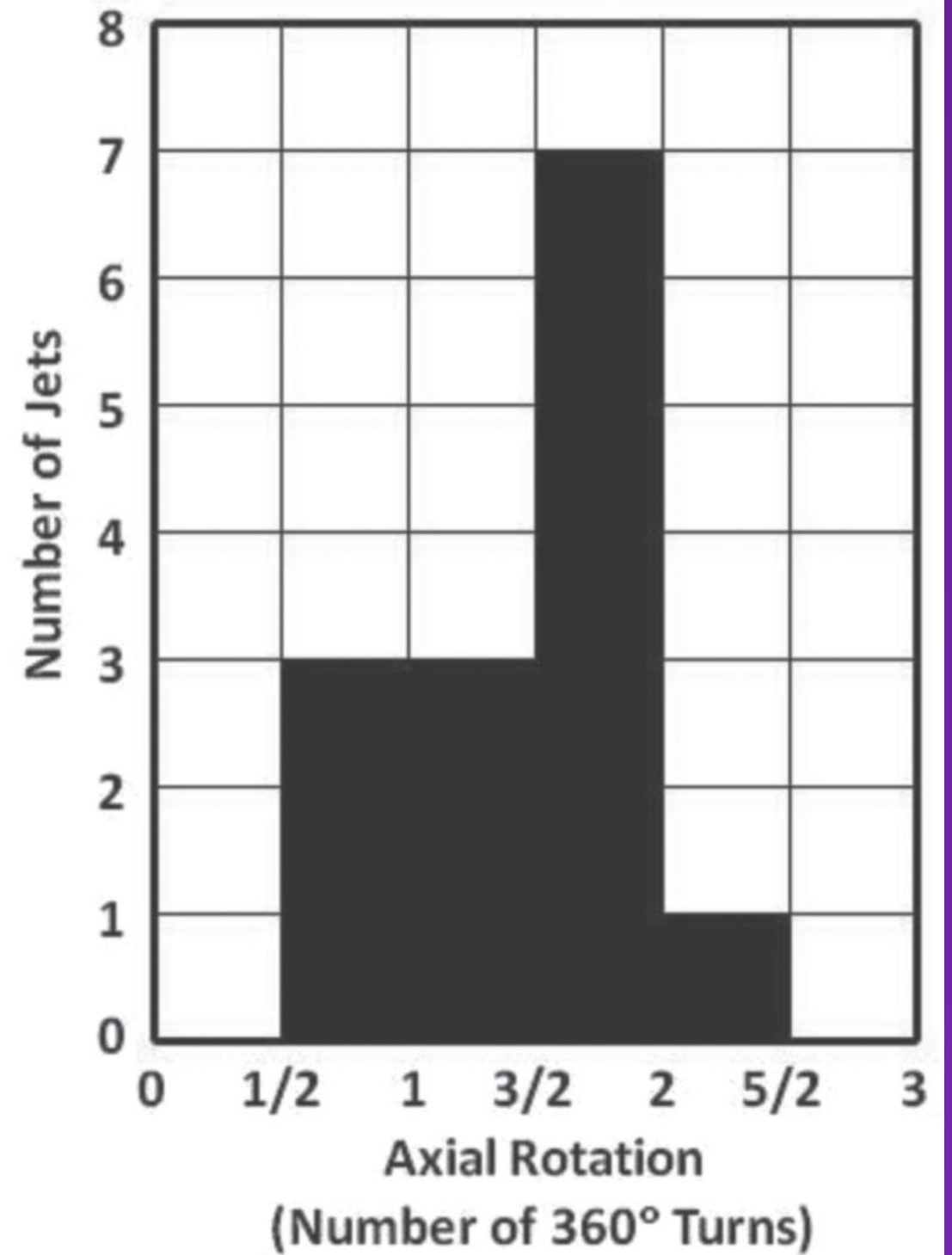
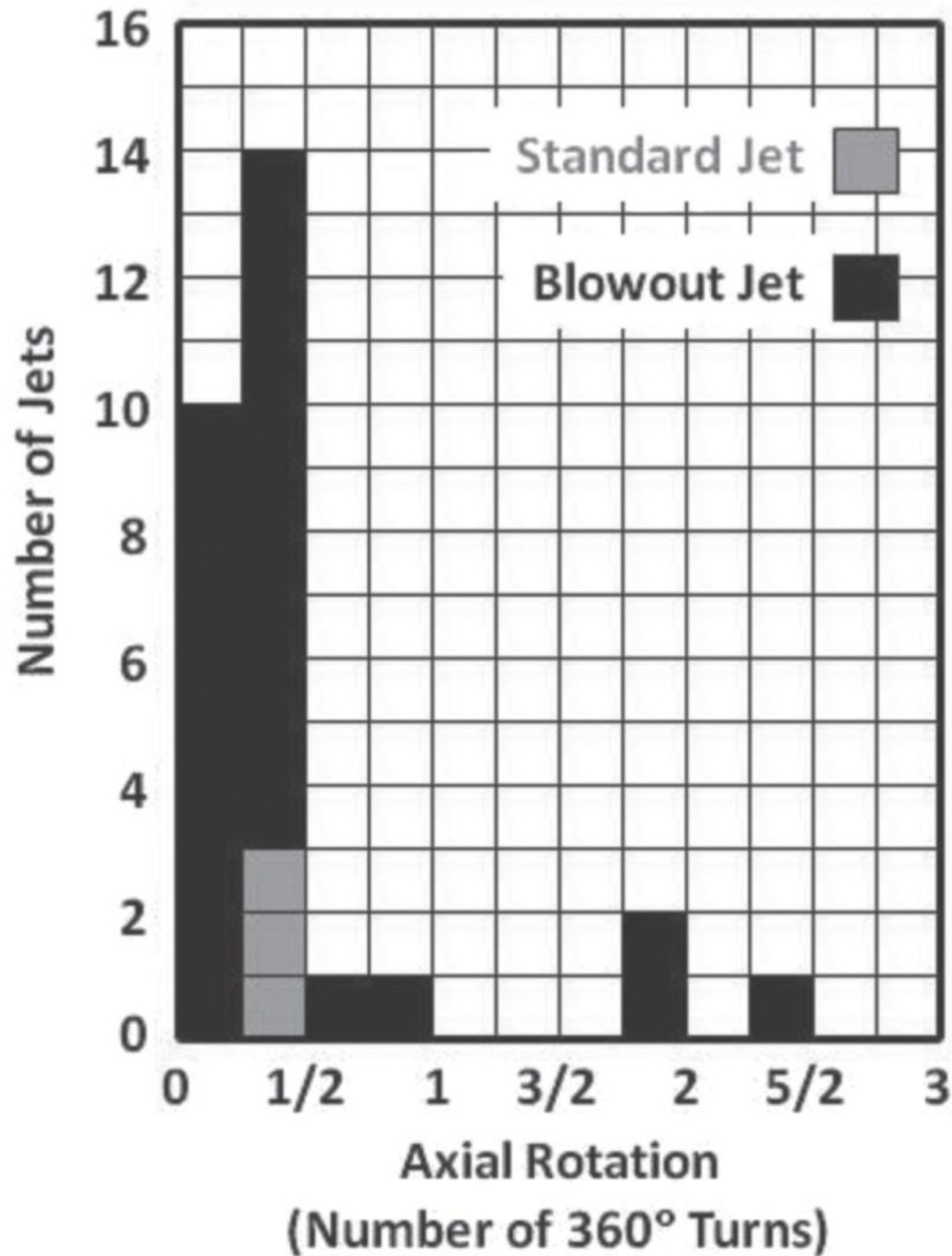


(Sterling, Harra, & Moore 2010)



(Sterling, Harra, & Moore 2010; Sterling et al. 2020)

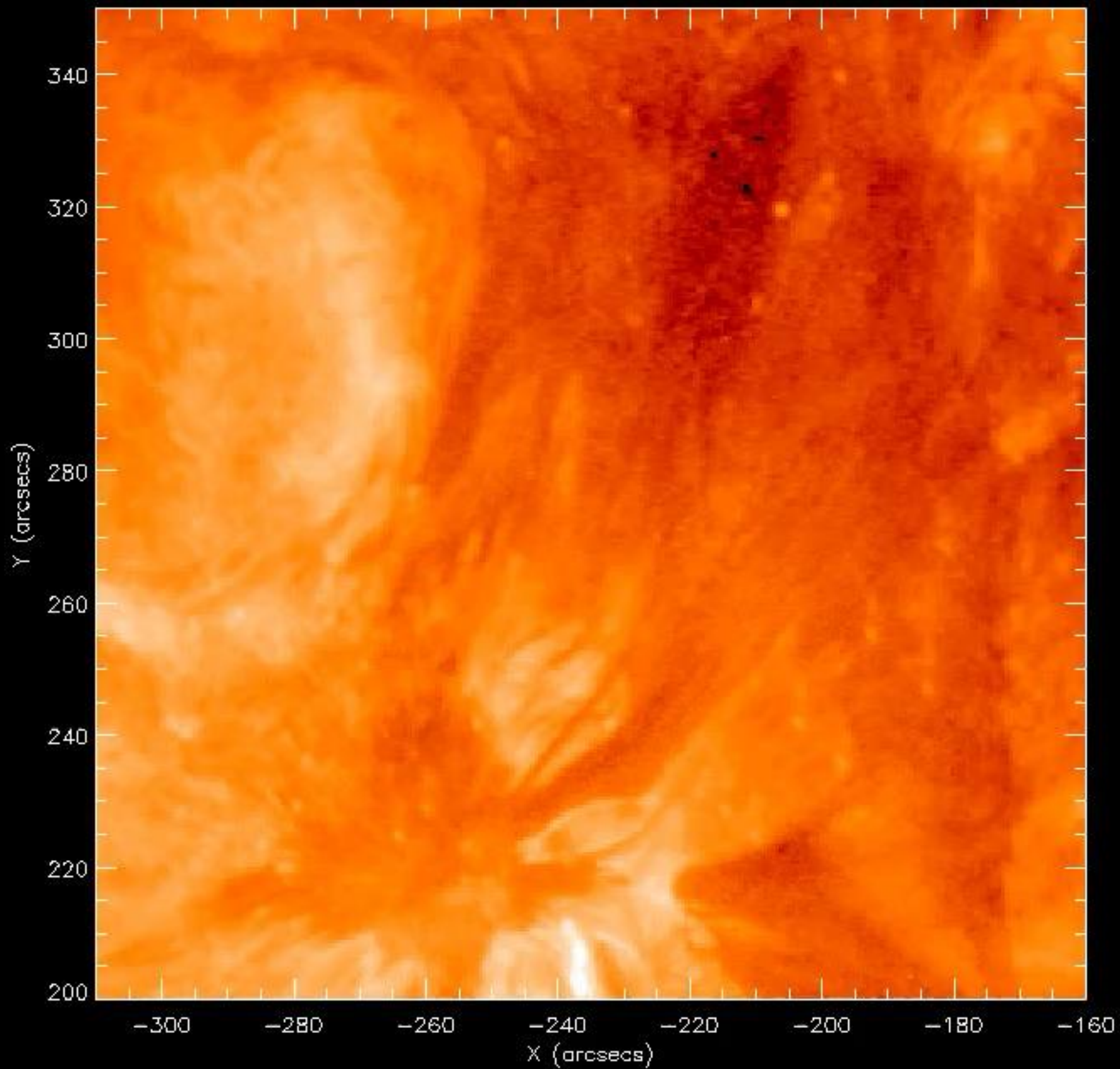
Twist in Jets



Random twisting PCH jets with cool component: Moore et al. (2013)

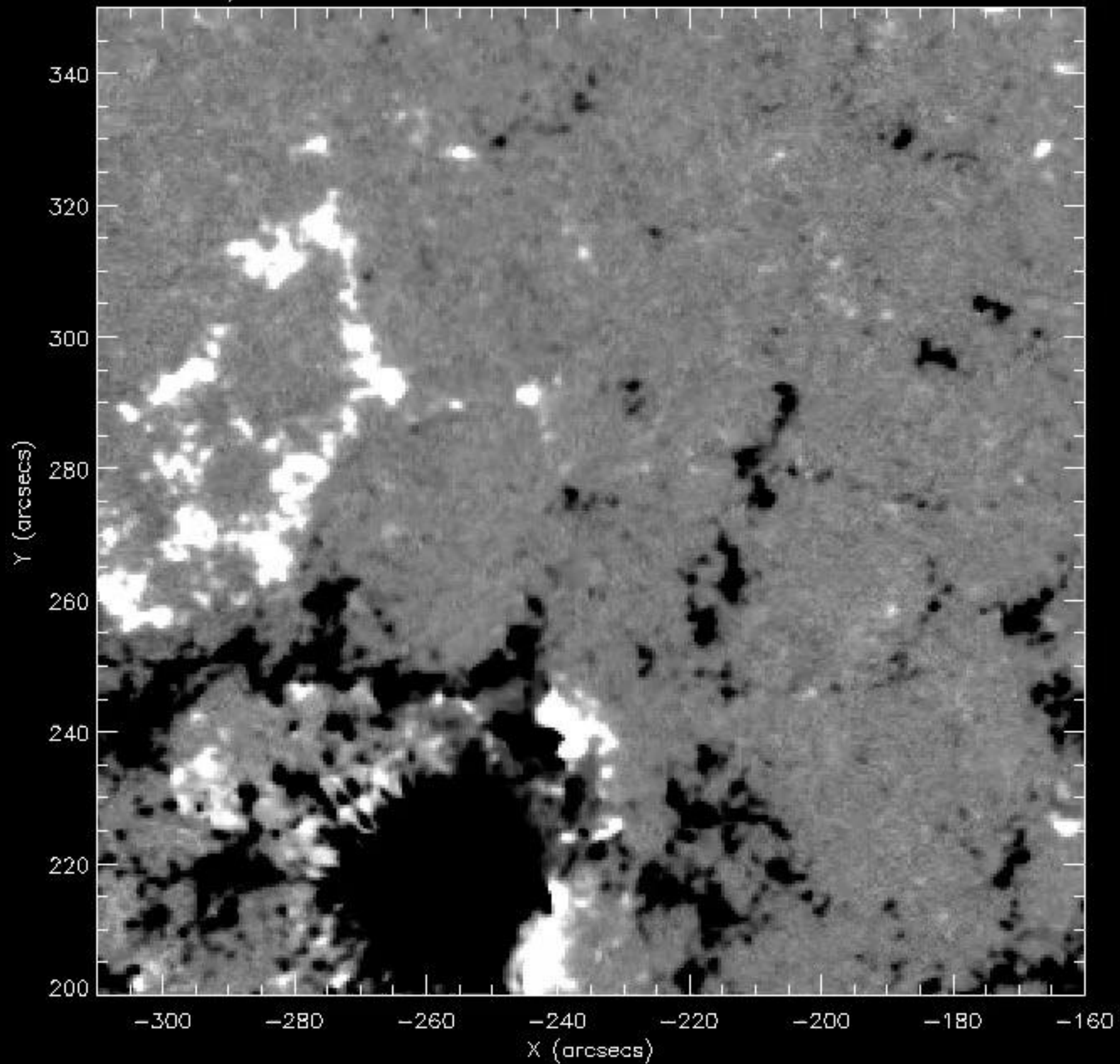
Narrow-CME-Producing Jets: (Moore et al. 2015)

SDO AIA_4 304 30-Jun-2012 19:00:44.130 UT



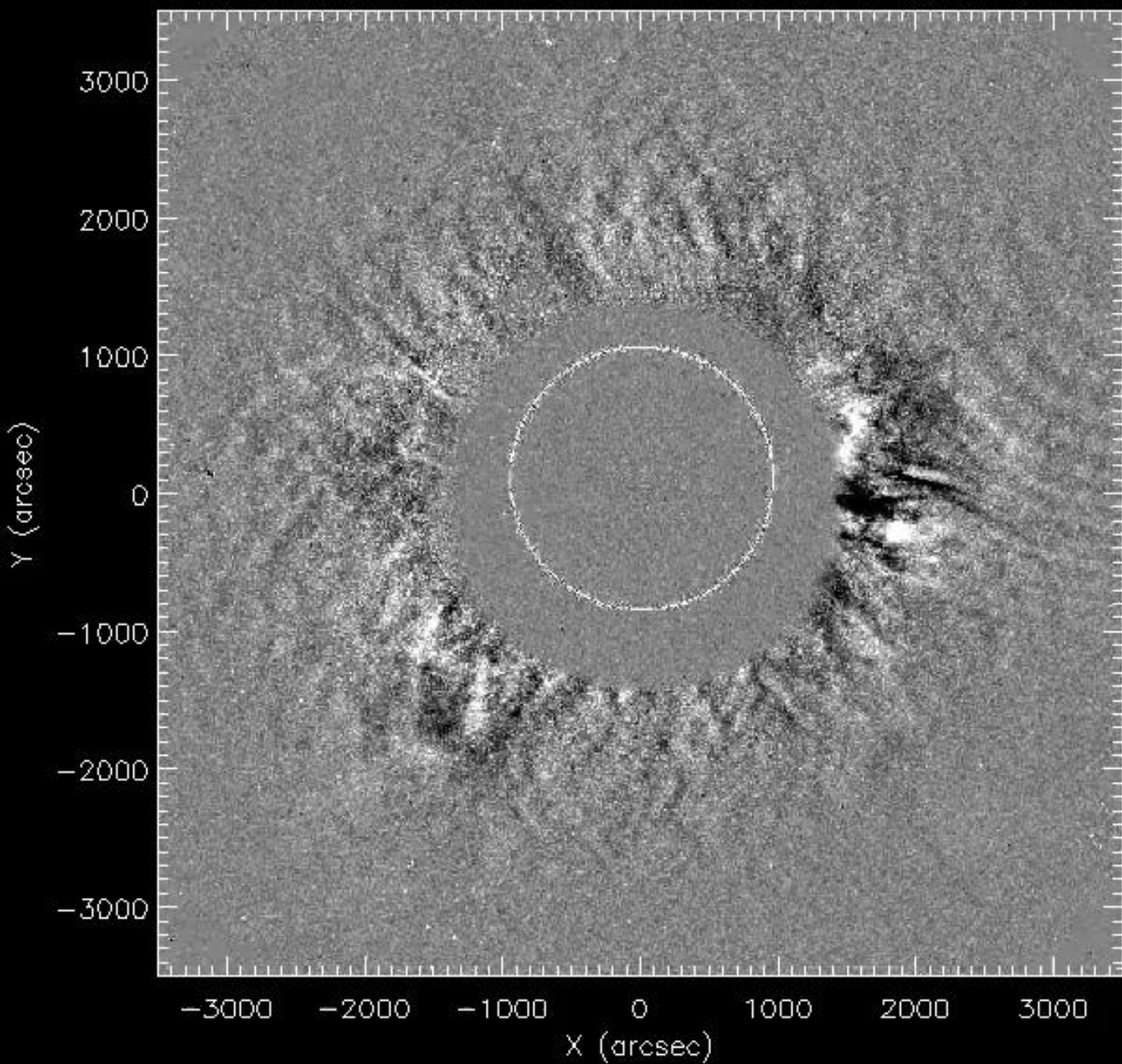
Sterling & Moore (2020)

SDO/HMI HMI_FRONT2 6173 30-Jun-2012 15:00:41.800 UT

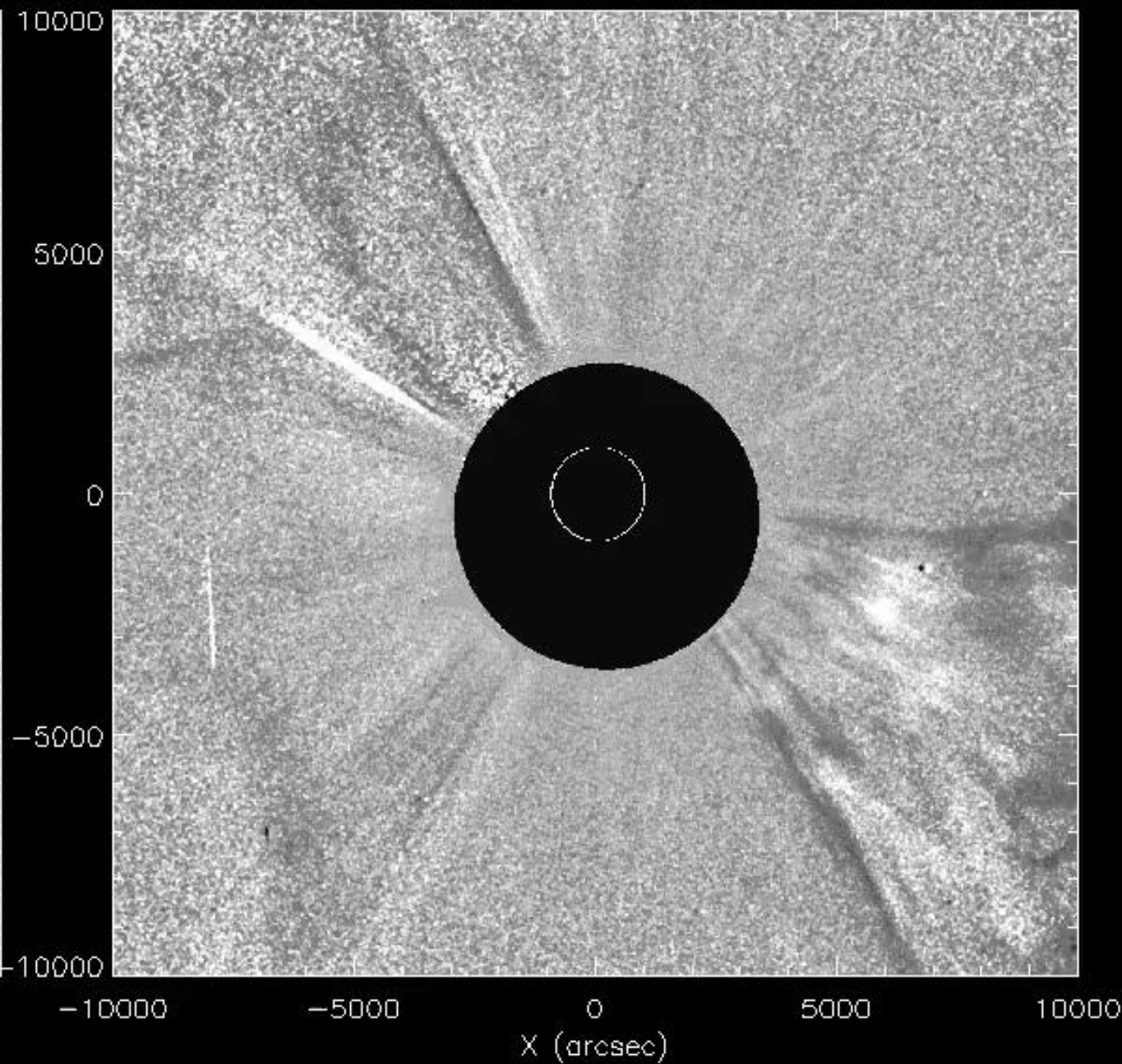


Sterling & Moore (2020)

STEREO-B COR1: 30-Jun-2012 17:25:18 UT

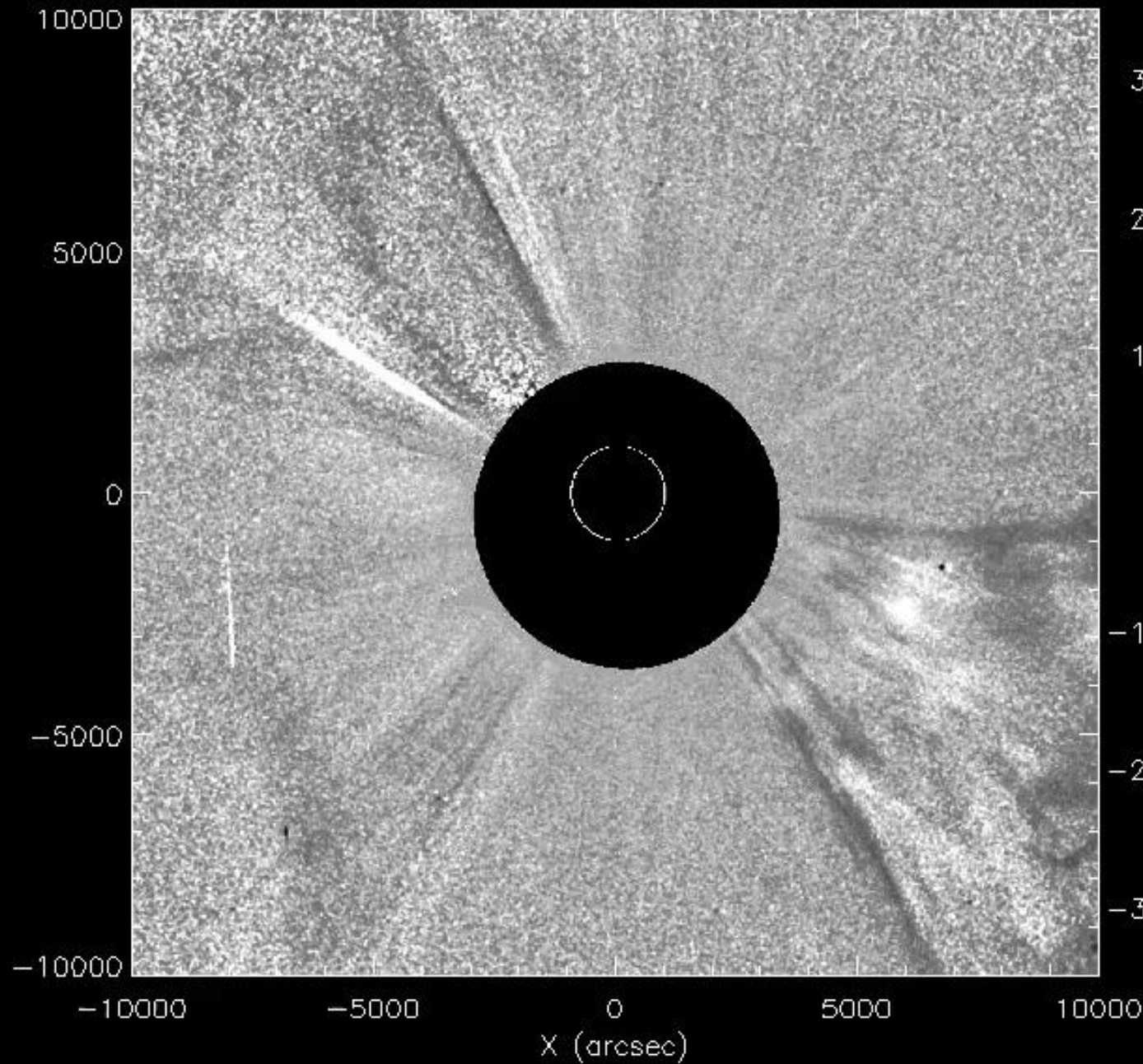


STEREO-B COR2: 30-Jun-2012 17:24:00 UT

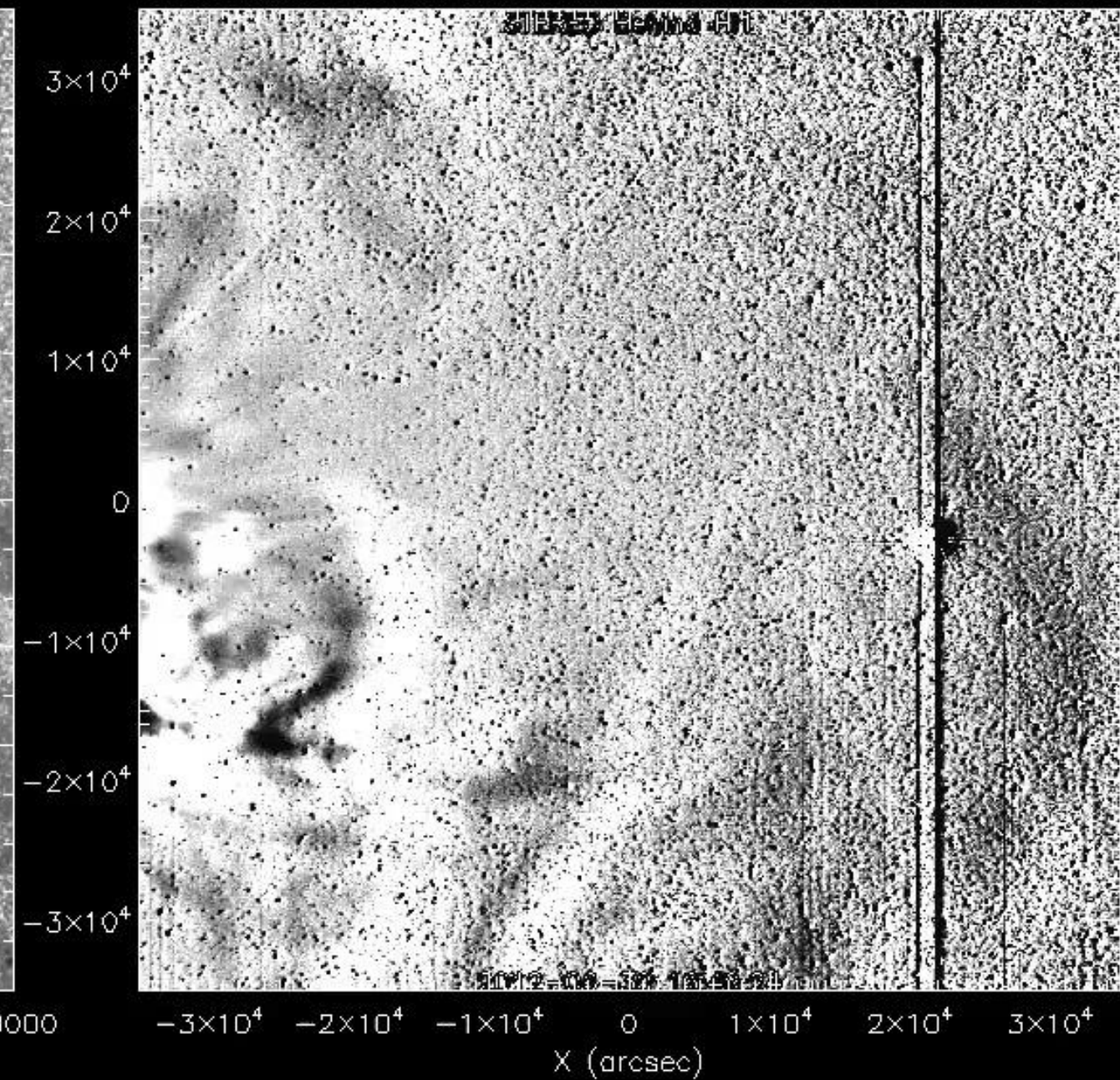


Sterling & Moore (2020)

STEREO-B COR2: 30-Jun-2012 17:24:00 UT



STEREO-B Hi1: 30-Jun-2012 16:49:01 UT

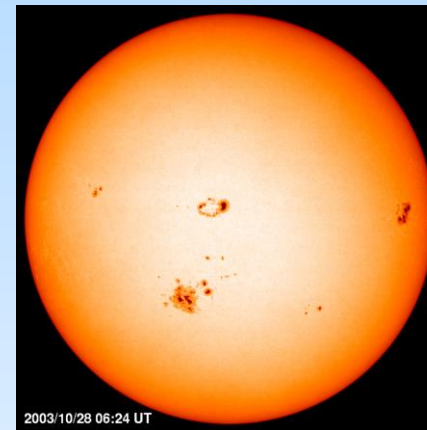


Sterling & Moore (2020)

Solar Interior and Atmosphere(s)

The Outer layers (Atmospheres) of the Sun:

- Photosphere

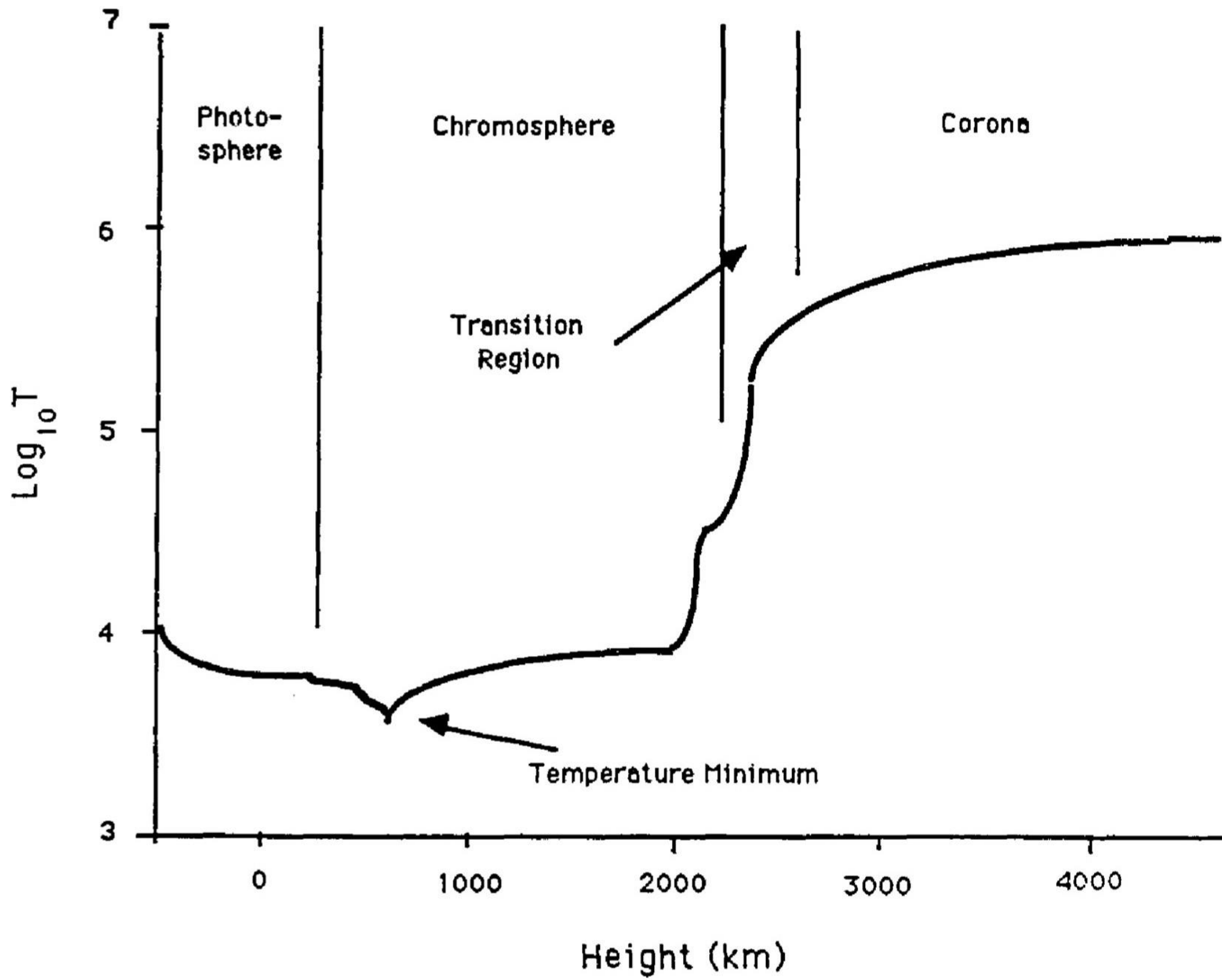


- Chromosphere

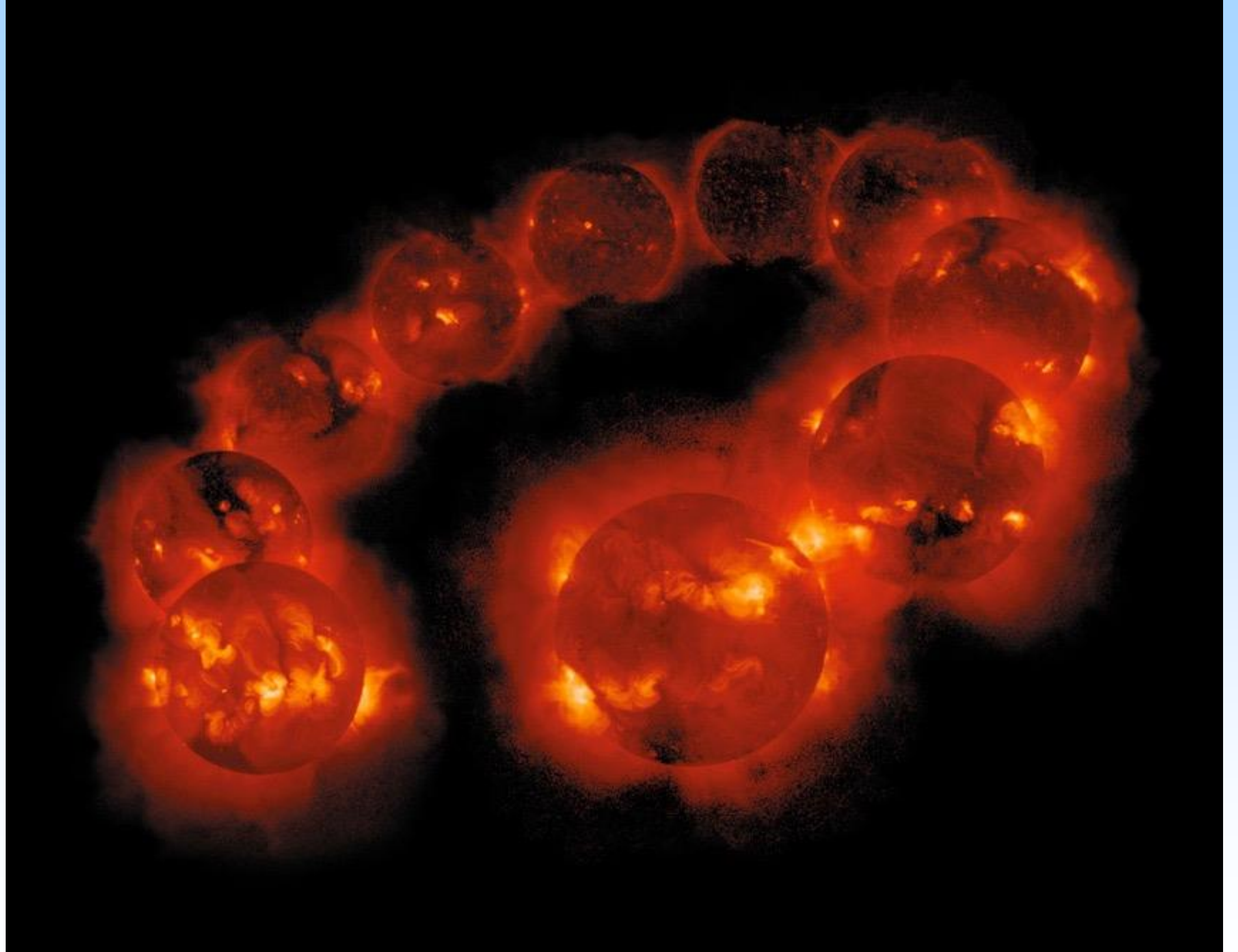


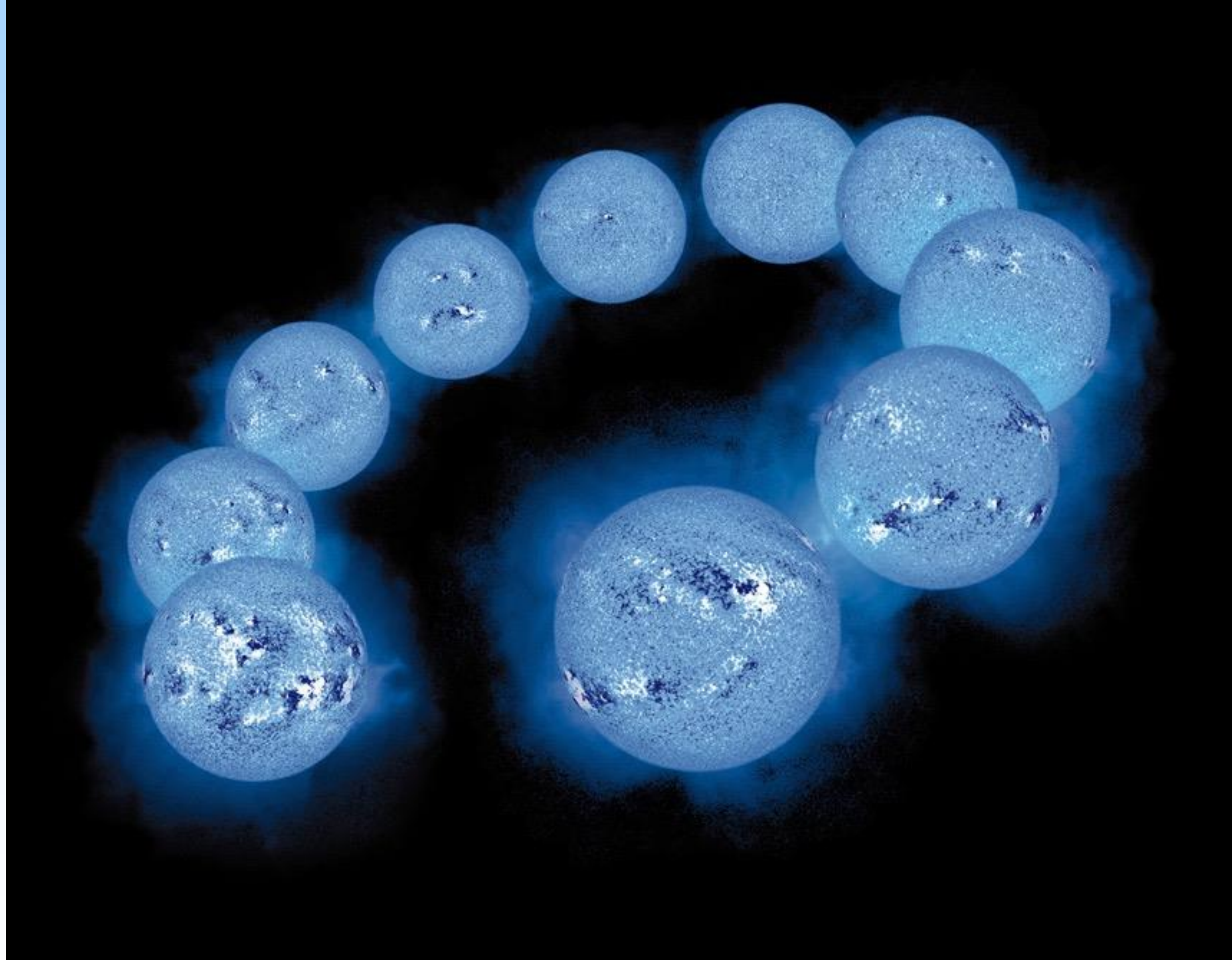
- Corona

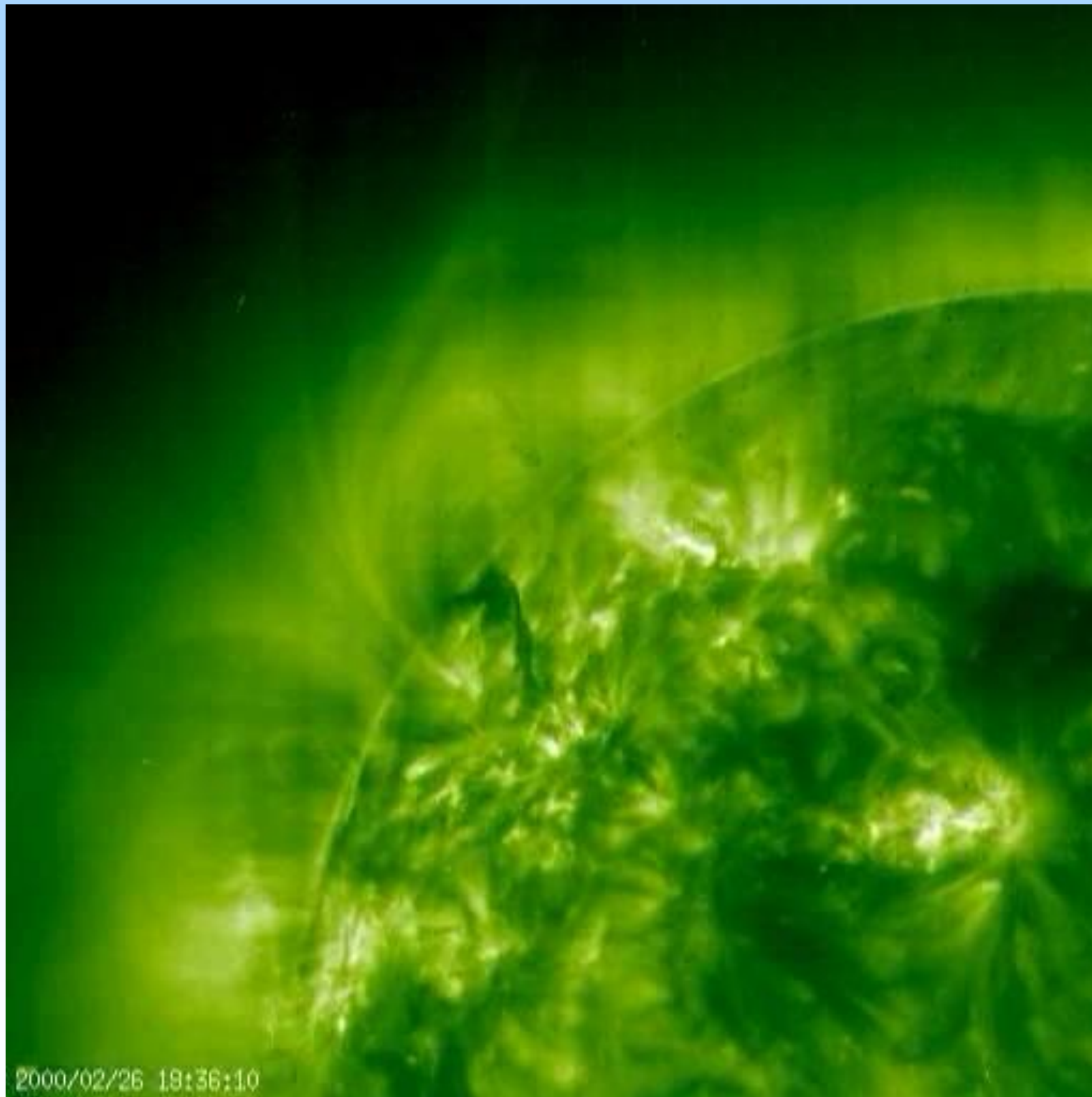




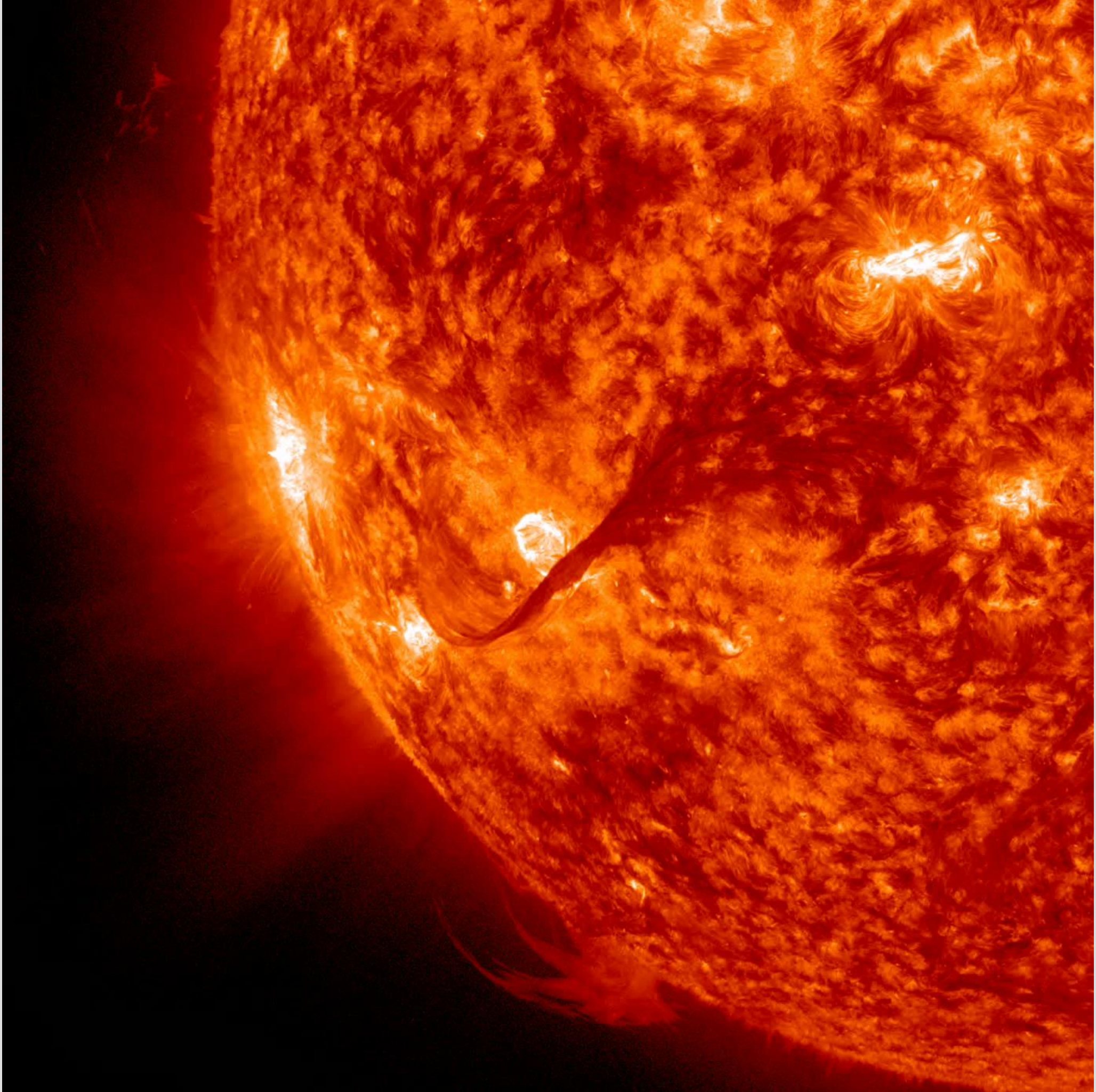
(Sterling, 1988)



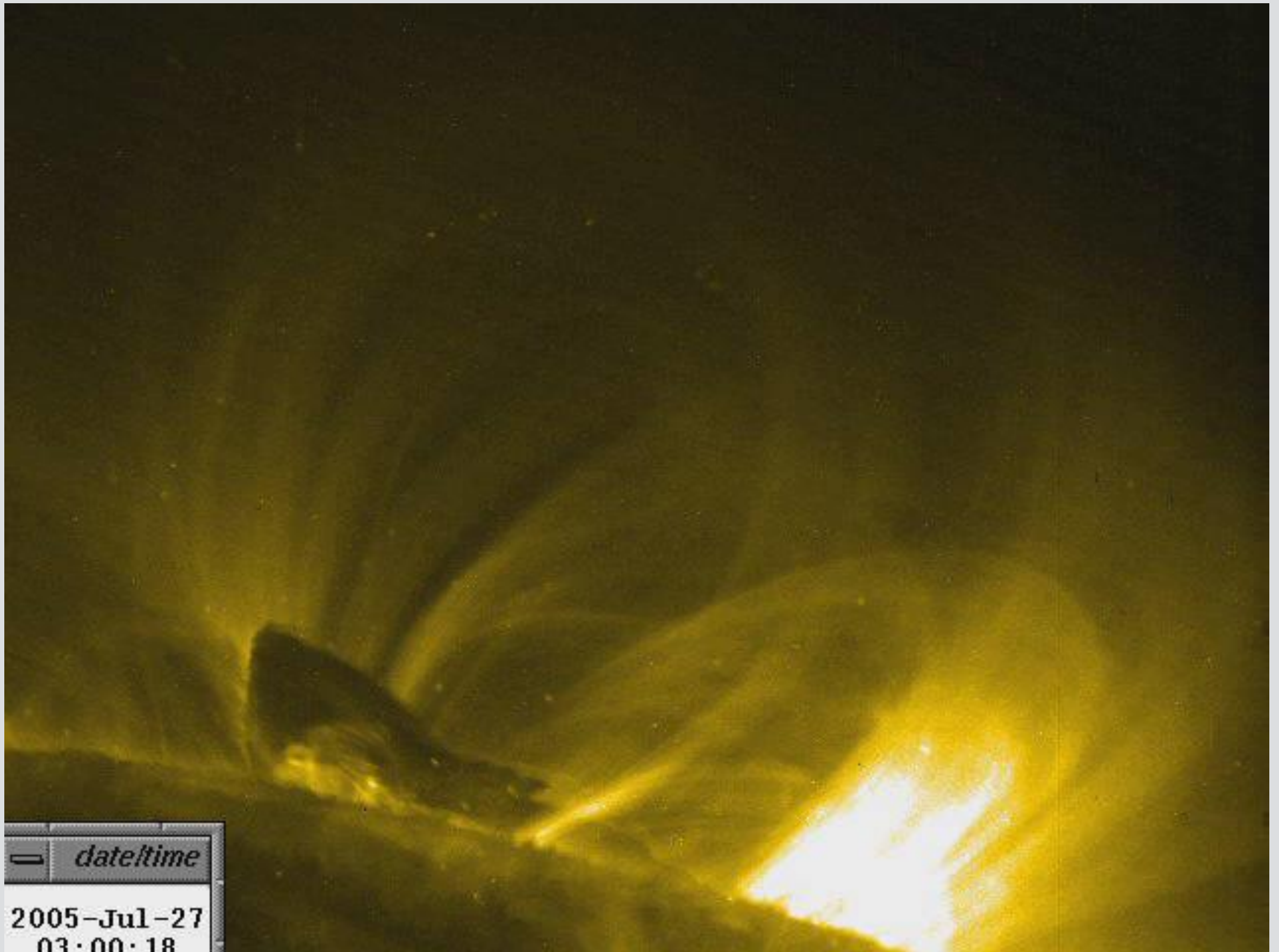




(Sterling & Moore 2004)



(TRACE 171)



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