

1 **Assessment of two stochastic cloud subcolumn generators using observed**
2 **fields of vertically resolved cloud extinction**

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ABSTRACT

18 We evaluate two stochastic subcolumn generators used in GCMs to emulate subgrid
19 cloud variability enabling comparisons with satellite observations and simulations of certain
20 physical processes. Our evaluation necessitated the creation of a reference observational
21 dataset that resolves horizontal and vertical cloud variability. The dataset combines two
22 CloudSat cloud products that resolve two-dimensional cloud optical depth variability of
23 liquid, ice, and mixed phase clouds when blended at ~ 200 m vertical and ~ 2 km horizontal
24 scales. Upon segmenting the dataset to individual “scenes”, mean profiles of the cloud fields
25 are passed as input to generators that produce scene-level cloud subgrid variability. The
26 assessment of generator performance at the scale of individual scenes and in a mean sense is
27 largely based on inferred joint histograms that partition cloud fraction within predetermined
28 combinations of cloud top pressure – cloud optical thickness ranges. Our main finding is that
29 both generators tend to underestimate optically thin clouds, while one of them also tends to
30 overestimate some cloud types of moderate and high optical thickness. Associated radiative
31 flux errors are also calculated by applying a simple transformation to the cloud fraction
32 histogram errors, and are found to approach values almost as high as 3 W m^{-2} for the cloud
33 radiative effect in the shortwave part of the spectrum.

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SIGNIFICANCE STATEMENT

36 The purpose of the paper is to assess the realism of relatively simple ways of producing
37 fine-scale cloud variability in global models from coarsely-resolved cloud properties. The
38 assessment is achieved via comparisons to observed cloud fields where the fine-scale
39 variability is known in both the horizontal and vertical directions. Our results show that while
40 the generators have considerable skill, they still suffer from consistent deficiencies that need
41 to be addressed with further development guided by appropriate observations.

42 **1. Introduction**

43 Satellite simulators are important tools of modern GCMs. Subcolumn cloud generators
44 are in turn critical components of satellite simulators striving to emulate subgrid-scale cloud
45 properties in order to bridge the coarse resolution of these models and the scales at which
46 satellite retrievals are performed. Emulation of subgrid variability is not only required for
47 mimicking satellite observations, but also contributes to improved calculations of quantities

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48 with non-linear dependences on cloud properties, such as radiative flux and precipitation rate
49 (Song et al. 2018). For some radiation schemes such as the Monte Carlo Independent Column
50 Approximation (McICA, Pincus et al. 2003), the coupling of the radiative transfer algorithm
51 with cloud fields resolved at subgrid scales is actually an inherent feature of their design.

52 Since simple profiles of averaged quantities at larger scales do not by themselves fully
53 constrain the distribution of profiles at smaller scales, there is considerable amount of
54 freedom (and uncertainty) in simulating the more detailed satellite views of clouds. The
55 emulated subgrid variability depends on how cloud microphysical properties are distributed
56 both horizontally and in height, and how cloud occurrence and the distributions of
57 microphysical properties overlap vertically. Within the satellite simulator framework, the tool
58 that handles subgrid variability is a cloud subcolumn generator. This tool produces stochastic
59 samples of subcolumn cloud profiles, which preserve the average profiles of the gridbox in
60 the limit of a large number of samples. An essential ingredient of subcolumn generators is a
61 set of rules on how the cloudy parts of a gridbox overlap vertically (Jakob and Klein, 1999).
62 The combination of cloud horizontal variability and how it correlates between various
63 atmospheric levels produces the subgrid variability of total (vertically integrated) cloud water
64 path and column cloud optical depth TAU. Note that subgrid variability of vertically
65 integrated (column) cloud properties can exist even if the properties themselves at individual
66 levels are distributed homogeneously; this is because the rules of cloud occurrence overlap
67 create variability in the number of layers that are cloudy in each subcolumn.

68 The subcolumn generator SCOPS (Subgrid Cloud Overlap Profile Sampler) of the
69 CFMIP (Cloud Feedback Modeling Intercomparison Project) Operational Satellite Simulator
70 Package (COSP, Bodas-Salcedo et al. 2011) assumes that vertical cloud occurrence follows a
71 combination of maximum and random overlap (Hillman et al. 2018) and that a cloudy layer's
72 condensate is homogeneous at the scale of model (GCM) gridboxes. While SCOPS
73 subcolumns are generally not passed outside of COSP, its underlying overlap assumptions
74 have been proven inadequate for the simulation of accurate cloud-sky radiative fluxes when
75 implemented in radiation schemes (Barker et al., 1999; Oreopoulos et al., 2012). A first step
76 then towards improvement of radiation from subcolumn generators would therefore be direct
77 statistical comparison of their cloud fields with observed subgrid cloud variability. Note that
78 horizontal coherence, i.e., the spatial arrangement of subcolumns does not need to be part of

79 the subgrid variability description and comparison since horizontal interactions within a
80 model gridbox are typically neglected.

81 Because subcolumn generators also emulate vertical variability, reference data
82 appropriate for generator validation must provide such information as well. An approach
83 embraced previously was to use simulated cloud fields more highly resolved than in GCMs as
84 proxy for observations, for example cloud fields from Multiscale Modeling Framework
85 (MMF) simulations (Hillman et al. 2018). In this treatment, model gridcolumns are treated as
86 observed subcolumns and serve as “truth” in comparisons with subcolumns produced by
87 SCOPS or other generators. The usefulness of such an evaluation is obviously limited by the
88 degree of realism of MMF clouds. Moving beyond previous work, here we make the case that
89 simulated cloud fields are not the only recourse, but rather that observed cloud fields exist as
90 a reference source for subgrid cloud variability for the purposes of subcolumn generator
91 validation and development. Such cloud fields come from combined observations of the
92 Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) lidar aboard the Cloud-Aerosol
93 Lidar and Infrared Pathfinder Satellite Observation (CALIPSO) satellite and its near-
94 synchronous CloudSat satellite carrying the Cloud Profiling Radar (CPR). The main feature
95 of the cloud fields retrieved from these two instruments is that the variability is resolved
96 horizontally not only for the atmospheric column as a whole, but also vertically within the
97 column.

98 Our paper thus uses such a reference dataset in an attempt to evaluate SCOPS and another
99 established subcolumn generator by Räisänen et al. (2004). It consists of two main parts: The
100 first part is dedicated to describing the construction of the observational reference dataset,
101 while the second part presents the approach and findings of our effort to assess the
102 performance of these two generators.

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104 **2. Quality assessment and improvement of reference dataset**

105 *a. Key Cloud Products*

106 Our construction of 2D cloud fields resolving horizontal (along the path of the two
107 observing satellites) and vertical (height) variability of visible cloud optical depth (COD) is
108 achieved by combining two CloudSat “release 5” (R05) products: 2B-CWC-RVOD
109 (Leinonen et al. 2016, hereafter 2BCWC) and the CALIPSO-enhanced 2C-ICE product

110 (Deng et al. 2015, hereafter 2CICE). To begin with, some quick clarifications on terminology
111 are in order before we proceed. Herein we use COD for “cell” optical depth and TAU for the
112 integrated/column optical depth, i.e., the vertical sum of CODs. In CloudSat terminology
113 CPR measurements are resolved vertically in 240 m “bins” and horizontally in 1.7 km (along
114 track) “rays”. A bin therefore represents a vertical layer, a ray essentially represents a
115 subcolumn, and a “cell” is a bin within a subcolumn. We refrain from using the term “bin”
116 for vertical layers because it is also used for histogram discretization.

117 1) LIQUID PHASE: 2B-CWC-RVOD (2BCWC)

118 This product focuses on clouds of liquid phase, with key inputs to the algorithm being the
119 CloudSat 2B-GEOPROF radar reflectivity profile (Marchand et al., 2008) and the column
120 cloud optical depth TAU from the Collection 6 level 2 Aqua Moderate Resolution Imaging
121 Spectroradiometer (MODIS) cloud product (MYD06) (Platnick et al., 2017), which has been
122 collocated with CloudSat CPR measurements and is available in CloudSat’s MOD06-1KM-
123 AUX product. The incorporation of MODIS TAU (TAU_{MODIS}) as a constraint in the retrievals
124 makes this a purely daytime product. The algorithm is based on the optimal estimation
125 framework (Rodgers 2000), with the measurement vector consisting of the logarithm of the
126 total column liquid optical depth TAU_{liq} and the profile of all valid CPR reflectivities. To
127 infer TAU_{liq} , the algorithm uses ice cloud optical depth provided by the 2CICE product
128 (Deng et al. 2015, see below) which is subtracted from MODIS total optical depth, i.e.,
129 $TAU_{liq} = TAU_{MODIS} - TAU_{2CICE}$. Ancillary temperature estimates from the European Centre
130 for Medium-Range Weather Forecasts (ECMWF) analysis provided as CloudSat’s ECMWF-
131 AUX product are used to delineate layers where ice, mixed, and liquid clouds are expected.
132 The state vector retrieved by the algorithm consists of the particle number concentration N_T ,
133 assumed constant throughout the column, and the profile of the logarithm of geometric mean
134 particle radius $r_g = \exp(\langle \ln r \rangle)$ (brackets denote the expected value). By assuming a
135 lognormal distribution of liquid particle size with geometric standard deviation σ_{log} , the
136 Liquid Water Content LWC profile can be derived one CloudSat subcolumn/ray at a time. A
137 preliminary version of the algorithm was previously outlined in Leinonen et al. (2016) and
138 has been revisited to better handle confounding factors such as precipitation, and overlying
139 ice and mixed-phase clouds. See [https://www.cloudsat.cira.colostate.edu/cloudsat-
140 static/info/dl/2b-cwc-rvod/2B-CWC-RVOD_PDICD.P1_R05.rev0_.pdf](https://www.cloudsat.cira.colostate.edu/cloudsat-static/info/dl/2b-cwc-rvod/2B-CWC-RVOD_PDICD.P1_R05.rev0_.pdf) for additional details.

141 2) ICE PHASE: 2C-ICE (2CICE)

142 The Cloudsat and CALIPSO Ice Cloud Property Product (2C-ICE, hereafter 2CICE)
 143 contains retrieved estimates of profiles of ice cloud water content, effective radius, and
 144 extinction coefficient for each active measurement ray that contains ice particles as indicated
 145 by CloudSat’s CPR and/or CALIPSO’s CALIOP. Specifically, the 2CICE cloud product uses
 146 combined inputs of measured radar reflectivity profiles from CloudSat’s 2B-GEOPROF
 147 product and measured attenuated backscattering coefficient profiles at 532 nm from CALIOP
 148 (Rogers et al. 2011) to constrain the ice cloud retrieval more tightly than the radar-only
 149 product, and to yield improved retrievals (Deng et al. 2015). As in 2BCWC, an optimal
 150 estimation framework is used with the measurement vector consisting of the CALIOP-
 151 measured backscattering coefficient profile and the CPR-measured reflectivity profile. The
 152 state vector is initialized with a priori estimates from extensive in situ measurements and/or
 153 literature-supported empirical relations and algorithms.

154 3) CLOUD MASK: 2B-CLDCLASS-LIDAR (2BCL)

155 The 2B-CLDCLASS-LIDAR product, hereafter 2BCL, combines CPR and CALIOP
 156 measurements for cloud phase determination and cloud scenario classification (Sassen and
 157 Wang 2008; Sassen and Wang 2012). Cloud classification is achieved by synthesizing
 158 information about the horizontal and vertical variability of cloud properties, the precipitating
 159 state of the cloud field, cloud temperature, and coincident MODIS radiances. For this work,
 160 we use 2BCL as a 2D (along track-height) cloud mask on the same grid as the previous two
 161 products. For each ray in the dataset, the height of the cloud top and base, as well as the cloud
 162 thermodynamic phase (liquid, ice, mixed) is extracted for vertically distinct cloud “objects”
 163 consisting of contiguous vertical layers. With each distinct cloud object assigned one of the
 164 three thermodynamic phases, all the vertical layers it encompasses have the same phase.

165 *b. Construction of base 2D cloud optical depth fields*

166 2CICE provides vertically (240m) and horizontally (~1.7 km) resolved (i.e., “cell”) ice
 167 water content (*IWC*) and particle effective radius (r_e) from which cell visible ice cloud optical
 168 depth COD_{ice} can be calculated according to (e.g., Stein et al. 2011):

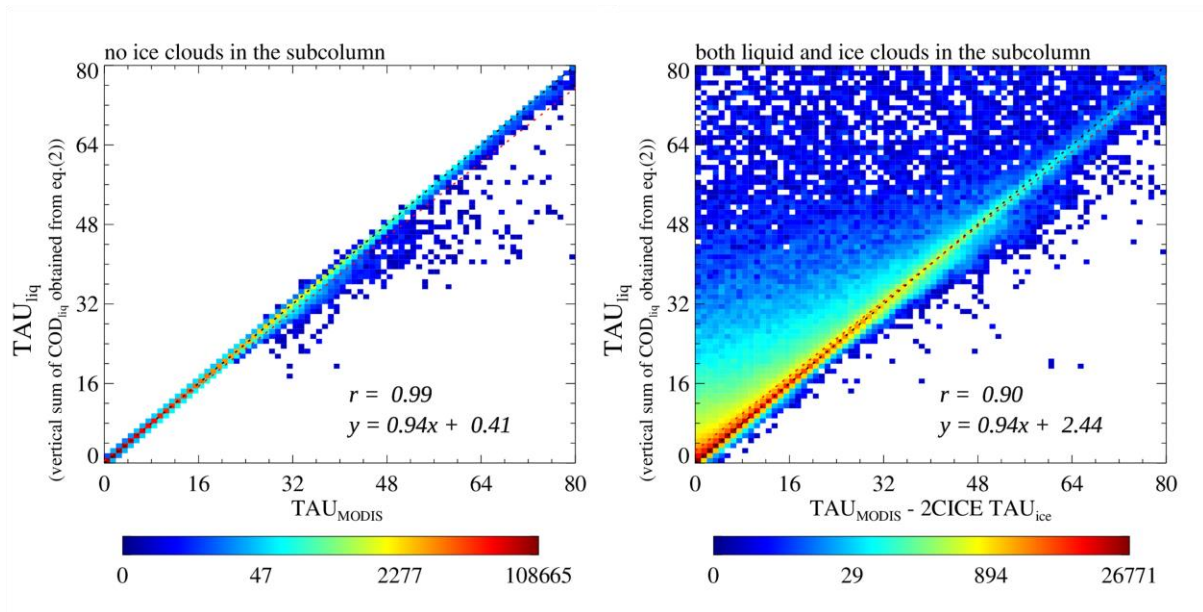
$$169 \quad COD_{ice} = \frac{3Q_{ext}IWC}{4\rho_{ice}r_e} \Delta z \quad (1)$$

170 where $\rho_{ice} = 0.92 \text{ g cm}^{-3}$ is the density of ice, $Q_{ext} = 2$ is the extinction efficiency at visible
 171 wavelengths, and $\Delta z = 240 \text{ m}$ is the cell’s physical thickness.

172 On the other hand, the 2D field of visible cloud optical depth for liquid phase clouds
 173 COD_{liq} is estimated from 2BCWC using the retrieved optical extinction coefficient. Per
 174 Leinonen et al. (2016), the cell's extinction coefficient is calculated by combining the cell's
 175 N_T , r_g , and the droplet distribution's geometric standard deviation σ_{log} :

$$176 \quad \sigma_{ext} = 2\pi N_T r_g^2 \exp(2\sigma_{log}^2) \quad (2)$$

177 Cell cloud optical depth COD_{liq} is then obtained by multiplying extinction with the cell's
 178 physical thickness Δz , i.e., $COD_{liq} = \sigma_{ext} \Delta z$.



179

180 Fig. 1. Comparison of either directly observed (subcolumns with liquid only clouds, left
 181 panel) or inferred (subcolumns with both liquid and ice clouds, right panel) $TAU_{MODIS,liq}$
 182 against TAU_{liq} from 2BCWC (eq. 2). One month of data (January 2007) with solar zenith
 183 angle below 45° were used, as in Leinonen et al. (2016).

184

185 We evaluate COD_{liq} obtained from eq. (2) by comparing its vertical integral TAU_{liq} with
 186 the collocated MODIS TAU_{liq} , inferred or directly observed. TAU_{MODIS} in MOD06-1KM-
 187 AUX corresponds to the TAU_{liq} used as constraint in the 2BCWC retrieval when there is no
 188 ice in the subcolumn; TAU_{liq} can also be inferred by subtracting 2CICE TAU_{ice} from total
 189 MODIS TAU_{MODIS} for subcolumns that have both liquid and ice clouds, as indicated above.
 190 A comparison for the first scenario is shown in the left panel of Fig. 1 for a sample month
 191 (January 2007), while a comparison for the second scenario is shown in the right panel. The
 192 general agreement seen in the figure serves primarily as a sanity check rather than an
 193 independent validation given that the 2BCWC retrievals are algorithmically constrained by

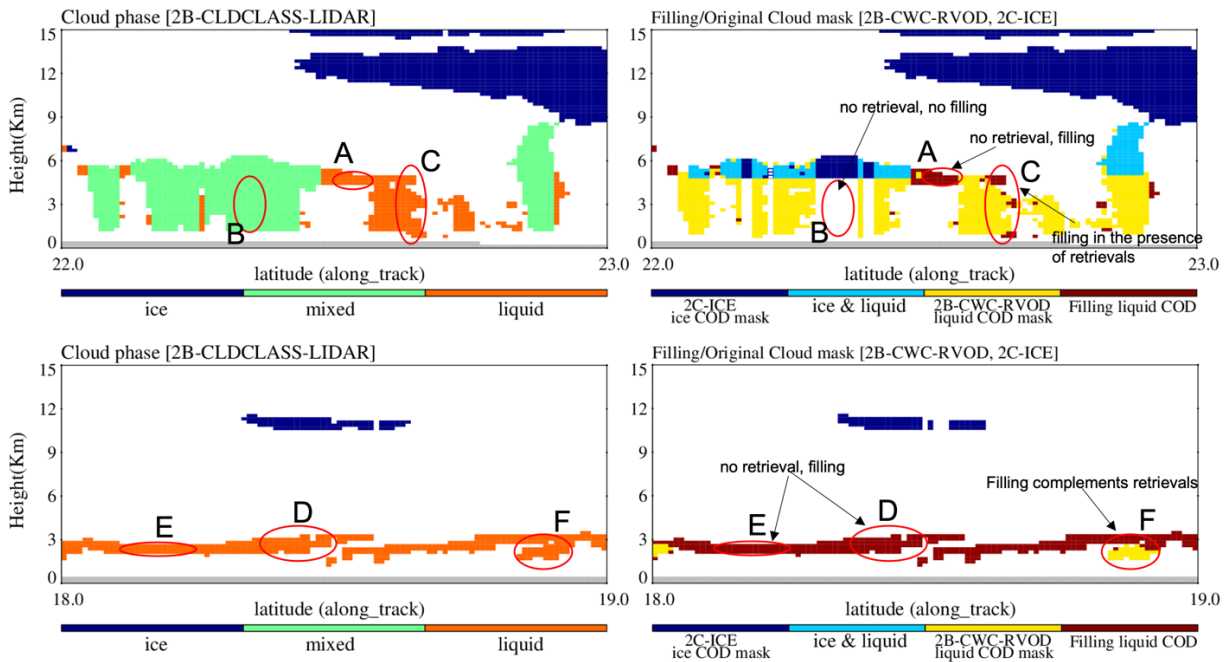
194 TAU_{MODIS}. Since MODIS is not sensitive to precipitating particles, the MODIS optical depth
195 constraint signifies that the combination of retrieved N_T , r_g in eq. (2) represents the extinction
196 of non-precipitating particles. However, occasional substantial deviations from the 1:1 to line
197 are indicative of imperfect filtering of rays containing precipitation (drizzle) in which case
198 the a priori constraints of the optimal estimation algorithm are inappropriate and may
199 introduce biases.

200 The blended 2D COD field from 2BCWC and 2CICE consists then of cells whose COD
201 comes from either 2CICE-alone (COD_{ice}), 2BCWC-alone (COD_{liq}), or combined CODs from
202 the two datasets (“mixed” phase cells with COD_{mixed} = COD_{liq}+COD_{ice}).

203 1) MISSING RETRIEVALS

204 Comparison of our 2D COD fields, constructed as described above, with CloudSat’s 2D
205 2BCL cloud mask product reveals that numerous cells identified as cloudy in that product do
206 not have COD > 0 in our combined product. This turns out to be a problem mostly confined
207 to liquid phase clouds whose CODs come from 2BCWC. It appears that most missing liquid
208 phase retrievals correspond to optically thin (and therefore of low CPR reflectivity) clouds
209 that are however still detected by the CALIOP lidar and therefore present in 2BCL. Since
210 upper-level clouds with substantial extinction (TAU ~ 5 and above) fully attenuate the lidar
211 beam, many of these low clouds seen by the lidar are either completely unobscured by clouds
212 above, or co-occur with upper level clouds still thin enough to allow the lidar beam to reach
213 the lower troposphere. Such clouds, present in 2BCL, but not in 2BCWC, are assigned COD
214 values according to a filling scheme described below.

215 Figure 2 shows two examples, each corresponding to a CALIOP-CloudSat “curtain”
216 within a 1° gridbox. In the first example (upper row) liquid cloud retrievals seem to be
217 missing (no yellow) in segment A. More missing liquid cloud retrievals apparently exist in
218 segment B, which is classified as being of mixed phase by virtue of being part of a distinct
219 cloud object in 2BCL assigned in its entirety to the mixed phase because of ice presence in its
220 upper part. The second example (bottom row) has far more missing liquid cloud retrievals:
221 indeed, very few yellow cells of successful 2BCWC retrievals exist compared to the large
222 liquid cloud cell population (orange cells) in 2BCL. Missing liquid retrievals can be found
223 both in parts of the scene with (segment D) and without (segment E) overlying high clouds.
224 Below we describe our simple approach to restore some of the missing retrievals for liquid
225 clouds that are apparently present (orange cells in the left panels of Fig. 2).



226

227 Fig. 2. Two examples 2BCL 2D cloud mask scenes (left panels) and reconstructed 2D COD
 228 scenes. Highlighted segments designated with letters are used as examples to explain major
 229 aspects of our filling scheme.

230

231 2) FILLING MISSING LIQUID PHASE RETRIEVALS

232 A simple scheme was devised to assign COD_{liq} values to cells with no retrievals in
 233 2BCWC but flagged as having liquid clouds in 2BCL. We follow the logic of the 2BCWC
 234 algorithm and calculate the subcolumn's TAU_{liq} by either simply assigning total TAU_{MODIS} to
 235 subcolumns with only liquid clouds, or assigning TAU_{MODIS} minus TAU_{ice} from 2CICE to
 236 mixed subcolumns. TAU_{liq} for the missing cells can then be calculated by subtracting the sum
 237 of existing 2BCWC COD_{liq} values in the subcolumn (if applicable, as in segment C of Fig.
 238 2). This adjusted TAU_{liq} is then partitioned into equal COD_{liq} values among all missing cells.
 239 The COD_{liq} profile of a subcolumn can therefore contain a mixture of COD_{liq} values from the
 240 original 2BCWC retrievals, and equal COD_{liq} values for all cells with missing retrievals from
 241 the partitioning of the adjusted TAU_{liq} . Assigning COD_{liq} values to cells with unavailable
 242 retrievals through a more sophisticated scheme would require information that is not
 243 currently at our disposal. Actually, application of even this procedure (constituting “Pass 1”
 244 of our filling scheme) is not always possible either because of unavailability of TAU_{MODIS} or
 245 because the adjusted TAU_{liq} after subtracting TAU_{ice} (and in some cases also the sum of
 246 2BCWC-provided COD_{liq}) from TAU_{MODIS} turns out to be negative. Missing COD_{liq} values

247 for such cases are filled with available neighboring COD_{liq} values (“Pass 2”), as described
248 below.

249 The possible Pass 1 scenarios to obtain profiles of COD_{liq} can then be summarized as
250 follows:

251 (1) Pure liquid subcolumn according to 2BCL, no 2BCWC retrievals in the subcolumn
252 (segment *E* in Fig. 2): Apportion equally $TAU_{MODIS} \equiv TAU_{MODIS,liq}$ to all cells with liquid
253 phase in 2BCL. The COD_{liq} profile consists of equal values.

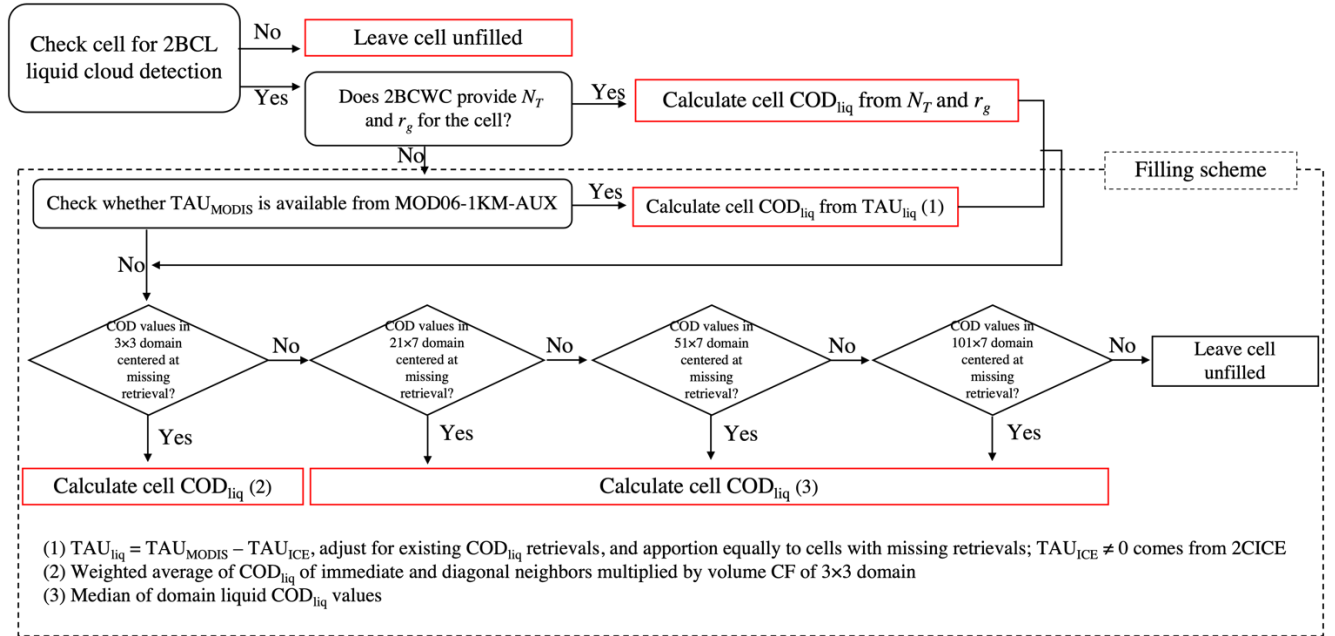
254 (2) Pure liquid subcolumn according to 2BCL with available 2BCWC retrievals (segment
255 *F* in Fig. 2): Subtract $TAU_{2BCWC,liq}$ from $TAU_{MODIS} \equiv TAU_{MODIS,liq}$ and if the result is
256 positive apportion equally to all missing cells with liquid phase in 2BCL. The COD_{liq} profile
257 is a mixture of 2BCWC retrievals and the equal values from the apportionment of this
258 adjusted TAU_{liq} corresponding to all cells of the subcolumn with missing retrievals.

259 (3) Overlying ice clouds and no retrievals in 2BCWC (segments *A* and *D* in Fig. 2):
260 Subtract TAU_{2CICE} from TAU_{MODIS} and if the result is positive apportion the resulting TAU_{liq}
261 equally to all missing cells with liquid phase in 2BCL. The COD_{liq} profile consists of equal
262 values.

263 (4) Both overlying ice clouds with 2CICE retrievals and liquid clouds with 2BCWC
264 retrievals (segment *C* in Fig. 2): Subtract the combined ice and liquid column optical depths
265 ($TAU_{2CICE} + TAU_{2BCWC}$) from TAU_{MODIS} and if the result is positive apportion the resulting
266 TAU_{liq} equally to all cells of liquid phase in 2BCL without 2BCWC retrievals. The COD_{liq}
267 profile is a mixture of 2BCWC retrievals and the equal values from the apportionment of
268 adjusted TAU_{liq} corresponding to all cells of the subcolumn with missing retrievals.

269 As previously mentioned, our scheme has also a second part (“Pass 2”) to fill 2BCL cells
270 with liquid clouds that remain unfilled after Pass 1 because TAU_{MODIS} is either unavailable or
271 inconsistent with available 2CICE and 2BCWC CODs in the subcolumn. Pass 2 applies a
272 “nearest-neighbor” (NN) scheme using cells with available COD_{liq} , from 2BCWC retrievals
273 or Pass 1. The scheme works as follows: We move from “left” to “right” (from start to end of
274 the data granule). We form a 3×3 domain centered around the missing value and calculate the
275 missing value as the weighted average of the available COD_{liq} values in the domain either
276 from 2BCWC, Pass 1, or Pass 2 on preceding subcolumns. We then multiply by the volume
277 liquid cloud fraction of the domain (fraction of cells in the domain with liquid clouds, i.e.,
278 number of cloudy cells divided by 9). Unity weights are used when averaging immediate

279 vertical and horizontal neighbors of the cell being filled, and $1/\sqrt{2}$ (pythagorean distance) for
 280 diagonal neighbors. If the 3×3 domain contains no COD_{liq} values, we expand the domain
 281 centered around the missing value to be a rectangle domain of size 21×7 (horizontal \times vertical)
 282 and calculate the missing value as the median of all available COD_{liq} values in the domain.
 283 Further expansion to 51×7 and 101×7 domains is built into the scheme, but is rarely invoked.



284
 285 Fig. 3. Flowchart summarizing the construction of our 2D COD_{liq} reference fields.

Provenance of COD_{liq} cells	Number of cells	Percentage of cells
Derived from r_g, N_T	289,197,600	80.7%
Filled using MODIS TAU_{liq}	45,438,111	12.7%
Filled using NN, 3×3 domain	22,188,733	6.2%
Filled using NN, 21×7 domain	1,270,571	0.4%
Filled using NN, 51×7 domain	184,610	0.1%
Filled using NN, 101×7 domain	82,989	0.0%
All COD_{liq} cells	358,362,614	100%

286 **Table 1.** Population information on the provenance of COD_{liq} cells in our 2D COD field
 287 constructed by combining 2BCWC and 2CICE retrievals.

288 The flowchart of Fig. 3 encapsulates the full scheme for constructing the 2D COD_{liq}
289 fields. Table 1 summarizes the provenance of COD_{liq} cells for the 2007 data we processed.
290 About 81% of COD_{liq} values come from 2BCWC; from the remaining 19%, about two-thirds
291 are filled via TAU_{MODIS} (explicit or adjusted MODIS TAU_{liq} , or inferred where appropriate
292 using TAU_{2CICE}), i.e., Pass 1, and the remaining one-third by NN filling (Pass 2), the vast
293 majority of which comes from very close neighbors (within the 3×3 domain).

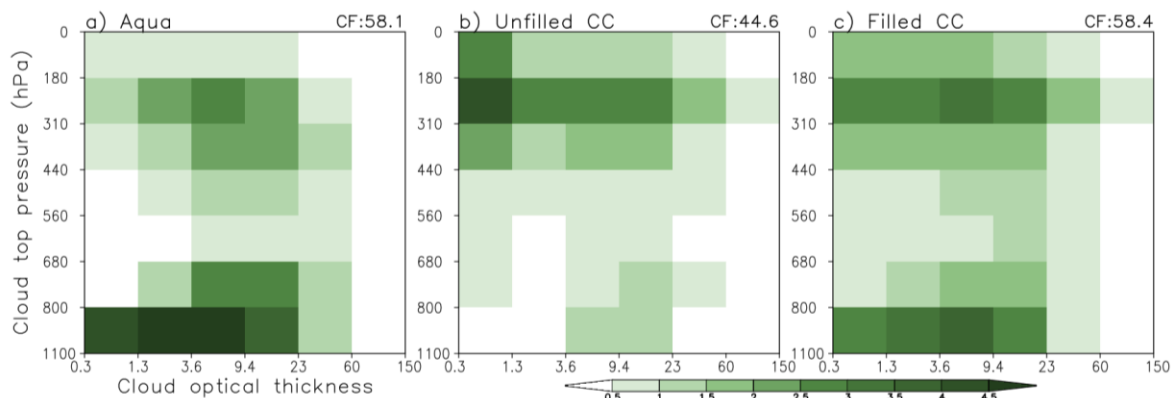
294 3) ASSESSMENT OF FILLING

295 A major application of the type of generators assessed in this work is to provide
296 subcolumns to COSP's International Satellite Cloud Climatology Project (ISCCP) and
297 MODIS simulators whose main diagnostic is joint cloud fraction (CF) histograms in cloud
298 top pressure (CTP) – TAU space. We therefore opt to use such histograms as the cornerstone
299 for evaluating not only the generators, but also the performance of the filling scheme.
300 Specifically, we examine whether average joint CTP-TAU histograms corresponding to the
301 modified COD fields are more alike to their counterparts from coincident Aqua cloud
302 observations than average joint histograms coming from the original (unfilled) COD fields.

303 The central role of CTP-TAU histograms in this paper merits some additional elaboration.
304 Given that we use in these histograms the subcolumn TAU and the CTP of its topmost cloudy
305 layer, one would initially think that the vertical distribution of COD does not matter, but
306 rather only its vertical integral, in accordance with the simplicity of our filling scheme's Pass
307 1. But the irrelevance of the COD profiles for the TAU of individual subcolumns casts doubt
308 at the same time on the appropriateness of CTP-TAU histograms as a rigorous evaluation
309 metric of the filling scheme. While it is true that using only TAU for individual observed
310 subcolumns considered in isolation makes the COD profile irrelevant, the details of the
311 vertical COD profile matter *for the ensemble of subcolumns* forming a scene. This is because
312 in addition to the observed scene CF profile (calculated as the fraction of the scene's
313 subcolumns with valid clouds in that layer), the generators also use the observed mean COD
314 profile as input. This layer mean COD across subcolumns *does depend* on how TAU is
315 vertically apportioned within individual subcolumns in the observations, making thus CTP-
316 TAU histograms coming from generators sensitive to the observed COD profile. Nonetheless,
317 the lack of better alternatives compels us to stay with the simple equal apportionment of
318 adjusted TAU_{liq} for Pass 1 of our filling scheme, and also CTP-TAU joint histograms as the

319 primary metric for evaluating both the filling scheme and the performance of the subcolumn
320 generators.

321 Aqua joint histograms were obtained from the equal area 3-hour histogram dataset used in
322 Cho et al. (2021) to derive MODIS Cloud Regimes (CRs) on the ISCCP grid. Because for the
323 year of our analysis (2007) Aqua and CloudSat-CALIPSO (CC) were part of the A-Train
324 constellation, temporal matching is already built into the dataset. We simply identify the
325 segment of the 2D COD field that falls within the 110 km gridbox on that day and create a
326 2D CTP-TAU histogram for that segment. To construct the 2D CTP-TAU histograms we
327 convert COD profiles from height to pressure coordinates using CloudSat’s ECMWF-AUX
328 product. We then eliminate all cloudy subcolumns with TAU < 0.3 (about 1% of all
329 subcolumns), by setting TAU = 0. This is done because MODIS detection and retrieval of
330 clouds with such low optical thickness is of low confidence, something accounted for in the
331 MODIS simulator (Pincus et al. 2012) used in this paper for generator evaluation. All
332 coincident joint histograms from the active CC and the passive MODIS observations are then
333 averaged. Figure 4 shows the comparison of global joint histograms resolving CF into 42-
334 bins, using ISCCP’s CTP-TAU bin discretization (Jakob and Tselioudis 2003), also used for
335 a CTP-TAU joint histogram version found in the MODIS cloud products.



337 Fig. 4. Globally-averaged joint CTP-TAU histograms from 2007 coincident observations of
338 (a) Aqua (left); (b) “Unfilled” COD fields from CC; (c) “Filled” COD fields from CC.

339 While MODIS clouds cannot be considered as “truth”, a certain degree of consistency
340 between passive and active retrievals is expected, with large discrepancies potentially being a
341 cause for concern. Our comparison is therefore highly instructive and clearly shows the
342 improvements brought by the filling scheme with respect to low clouds. The unfilled joint
343 histogram has a far smaller overall CF than the MODIS global histogram (44.6 vs 58.1, in %,
344 henceforth implied for all CF values) with the difference in low clouds, CTP > 680 hPa,

345 being even larger: low CF values are 28.5 for MODIS and 7.4 for unfilled CC 2D COD
346 fields. The reconstructed COD dataset compensates some of the low cloud difference with
347 more high clouds, especially in the bin of smallest TAU due to CALIOP's sensitivity to even
348 the optically thinnest clouds. But filling is essential in raising the low CF of CC to 19.7, still
349 not as high as MODIS (which may be overestimating low CF because MODIS pixels are
350 assumed overcast when cloud is detected), but good enough to bring down the overall CF
351 discrepancy to (absolute) 0.3%.

352 The evaluation of the filling scheme via such a CTP-TAU joint histogram comparison can
353 also be broken down by the MODIS CRs of Cho et al. (2021). These CRs represent the most
354 common cloud mixtures observed by MODIS at daily ~100 km scales as represented by the
355 mean of all CTP-TAU histograms deemed alike by a k-means clustering algorithm. The CTP-
356 TAU histograms of the CC COD segments collocated with MODIS-Aqua equal area
357 gridboxes are assigned to the MODIS CRs identified therein. For this comparison, histogram
358 averaging is therefore performed as before, but now separately for each MODIS CR. The
359 relative proximity of joint histograms coming from the unfilled and filled cloud fields to the
360 Aqua reference can be summarily captured by Euclidean Distances (EDs) between members
361 of histogram pairs. EDs derived for unfilled and filled CC COD fields are very similar for
362 CRs with relatively small populations of low clouds (CR1-CR6). But for the other CRs, CR7-
363 CR11, with plentiful low clouds, the lower EDs of the filled cloud fields indicate better
364 histogram resemblance, particularly for CR7-CR10. These results instill confidence in the
365 beneficial effects of our simple approach to fill cells with missing COD_{liq} values.

366

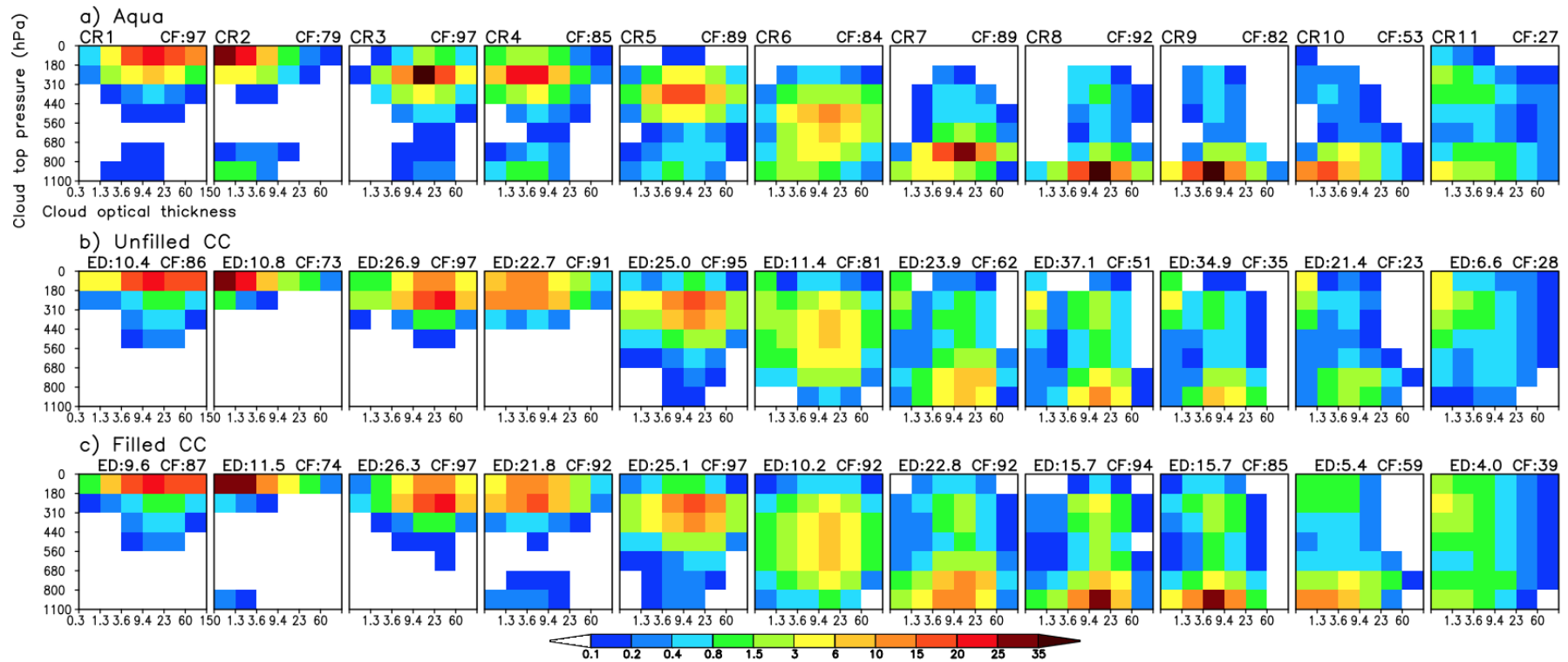
367 **3. Performance of subcolumn generators**

368 *a. Description of subcolumn generators and implementation specifics*

369 1) SCOPS GENERATOR

370 The SCOPS generator can produce subcolumns obeying random, maximum, or (our
371 choice for this paper) maximum-random overlap, and can treat separately convective and
372 stratiform clouds if such a distinction is known in the source dataset and of interest (neither
373 applies for this work). SCOPS takes as input the gridbox's or (in our case) the scene's mean
374 CF profile (the fraction of the scene at each layer with COD above the threshold that
375 indicates cloud presence) and then outputs a set of subcolumn cloud occurrence profiles

376 where the individual subcolumn cells are either overcast or cloud free (i.e., binary 0 or 100
377 cell CF). Condensate amounts are then assigned to the overcast cells according to the
378 standard implementation in COSP which assumes constant in-cloud condensate mixing ratio
379 across each layer equal to the gridbox layer mean provided by the host GCM. In our study
380 each SCOPS subcolumn is multiplied by the scene's mean COD rather than condensate
381 profile. The output is a 2D scene of COD that obeys the cloud occurrence overlap rules of
382 SCOPS and has horizontally uniform layer COD, but horizontally variable subcolumn TAU.
383 The scene COD field is then passed to the MODIS simulator to generate a CTP-TAU joint
384 histogram.



385

386 Fig. 5. Globally-averaged joint CTP-TAU histograms from coincident 2007 observations broken down by MODIS CR. (a) Aqua (top); (b)
 387 “Unfilled” COD fields from CC (middle); (c) “Filled” COD fields from CC (bottom). Above the CC panels, we also provide in addition to CF
 388 the CR-specific EDs from observations.

389 2) RAISANEN GENERATOR

390 The “Raisanen” generator also yields an ensemble of stochastically generated cloudy
391 subcolumns (Räisänen et al. 2004). Each layer within each subcolumn (cell) is assumed
392 homogeneous, with a CF of either 0 or 100, but condensate amount can vary across a vertical
393 layer consisting of the cells of the subcolumns at the same altitude, forming thus the layer’s
394 condensate probability density function (PDF) which can be described by an analytical
395 function such as gamma, beta or lognormal. In our case, we pass to the generator the CF,
396 mean COD, and COD variance profiles of the 2D scene, the latter coming from the CF profile
397 using the variance parameterization in Oreopoulos et al. (2012). The COD variance profile is
398 then used to create a profile of beta distribution PDFs.

399 The generator allows for a continuous range of cloud occurrence overlap rates between
400 maximum and random overlap according to the generalized overlap paradigm of Hogan and
401 Illingworth (2000). In this paradigm a weighting factor controls the relative contribution of
402 maximum (dominating for values of weighting factor close to one) and random overlap
403 (dominating for values of the weighting factor close to zero) to the combined CF of two
404 cloudy layers. The weighting factor is parameterized as an exponentially decaying function
405 with an e-folding distance or “decorrelation length” describing its rate of decrease as a
406 function of the separation distance between cloud layer pairs. Small values of decorrelation
407 length denote rapid decline of the weighting factor with separation distance (near-random
408 overlap) while large values denote a slow decline (near-maximum overlap). We use a
409 parameterization that captures the day-to-day latitude dependence of decorrelation length
410 with a Gaussian function fit to CloudSat observations of cloud occurrence overlap
411 (Oreopoulos et al. 2012). Similarly, vertical correlations of COD PDFs are captured by
412 correlations of COD ranks (i.e., Spearman rank correlations), also assumed to decay
413 exponentially with layer separation distance, according to a second decorrelation length. This
414 decorrelation length is also parameterized with a latitude and day-of-the-year-varying
415 Gaussian function which fits CloudSat reflectivity observations (Oreopoulos et al. 2012) and
416 represents a more rapid decay with vertical separation distance of COD rank than cloud
417 occurrence overlap.

418

419 *b. Results*

420 We use the one-year (2007) dataset of reconstructed CC COD fields to assess the skill of
421 the two subcolumn generators described above in simulating cloud subgrid variability of
422 scenes consisting of 100 subcolumns (~110 km given CC ray sampling). We only use oceanic
423 scenes (about 586 thousand scenes) because height in CC datasets is referenced relative to
424 local surface elevation, making the averaging of vertically resolved subcolumns ambiguous
425 when surface elevation varies.

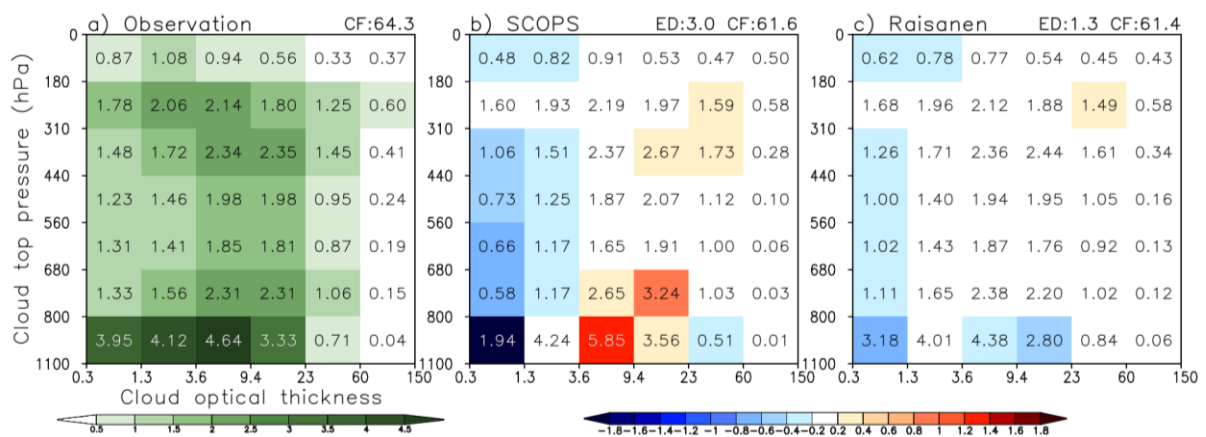
426 Our preferred (but not only) method of assessing skill is comparison of CTP-TAU joint
427 histograms coming from the MODIS simulator. These capture how clouds and their
428 condensate align vertically in real and simulated overlap scenarios to generate subcolumn
429 TAUs, and at the same time provide a rough cloud type discretization according to cloud top
430 height (pressure). Because in constructing these joint histograms the MODIS simulator
431 rejects subcolumns with $\text{TAU} < 0.3$, the reconstructed COD fields and the statistics derived
432 from them treat these subcolumns as cloud-free. All in all, the generators are evaluated using
433 comparisons between: (1) Mean CTP-TAU histograms obtained by extensive averaging of
434 individual scene histograms and their corresponding bin-resolved cloud radiative effects
435 (CREs); (2) Quantities derived from individual scene CTP-TAU histograms; (3) Profiles of
436 cumulative CF and (the closely related) CF exposed to space.

437 1) GRAND AVERAGE HISTOGRAMS

438 Figure 6 shows the one-year mean global ocean of 42-bin CTP-TAU joint histograms
439 from observations (left) and from the subcolumn generator reconstruction (middle and right
440 panels). The leftmost column of observed globally (ocean-only) averaged joint histograms
441 includes the numerical values of bin CF. The middle column shows the globally-averaged
442 histograms coming from SCOPS-generated subcolumns; the numbers stand for the bin CF
443 values while color displays differences from the reference observed histograms. The
444 rightmost column similarly conveys results from the Raisanen generator.

445 Once again, the overall resemblance of the reconstructed histogram to the true histogram
446 can be captured by the ED between the two histograms. According to that metric, the
447 Raisanen generator performs better overall (smaller ED) despite its slightly worse than
448 SCOPS underestimation of total CF. Both generators have reasonably good skill in
449 reproducing the total vertically projected CF. This is expected to some extent since they are
450 supplied observed CF profiles. Nevertheless, how the CFs of individual layers are overlapped
451 still matters: the underestimation by the generators suggest that they overlap clouds slightly

452 more maximally than in observations. A common deficiency of the two generators is the
 453 underestimation of optically thin ($\text{TAU} < 3.6$) clouds, which is less severe and mostly
 454 concentrated to $\text{TAU} < 1.3$ for the Raisanen generator. These underestimates are possibly
 455 related at least in part to the greater than observed tendency for maximum overlap and the
 456 resulting greater vertical cloud alignment that reduces the probability of optically thin clouds.
 457 SCOPS seems to compensate for the deficit of optically thin clouds with overestimates of
 458 clouds of moderate and large optical thickness, something that the Raisanen generator is
 459 much less prone to. Note that while absolute biases for the optically thickest TAU class (60-
 460 150) are mild for both generators, the small differences from observations correspond to
 461 small CFs to begin with (i.e., such clouds are rare). The CF of the second largest TAU class
 462 (23-60) is overestimated similarly by both generators (7.4 vs 6.7 in observations), but SCOPS
 463 suffers more error compensation. On the whole, the Raisanen generator produces fewer
 464 extreme biases (both overestimates and underestimates) than SCOPS, and most of the overall
 465 CF underestimate comes from optically thin clouds. SCOPS on the other hand would have
 466 suffered a much greater total CF underestimate due to thin clouds were it not for
 467 compensatory overestimates for clouds of moderate and high optical thickness.

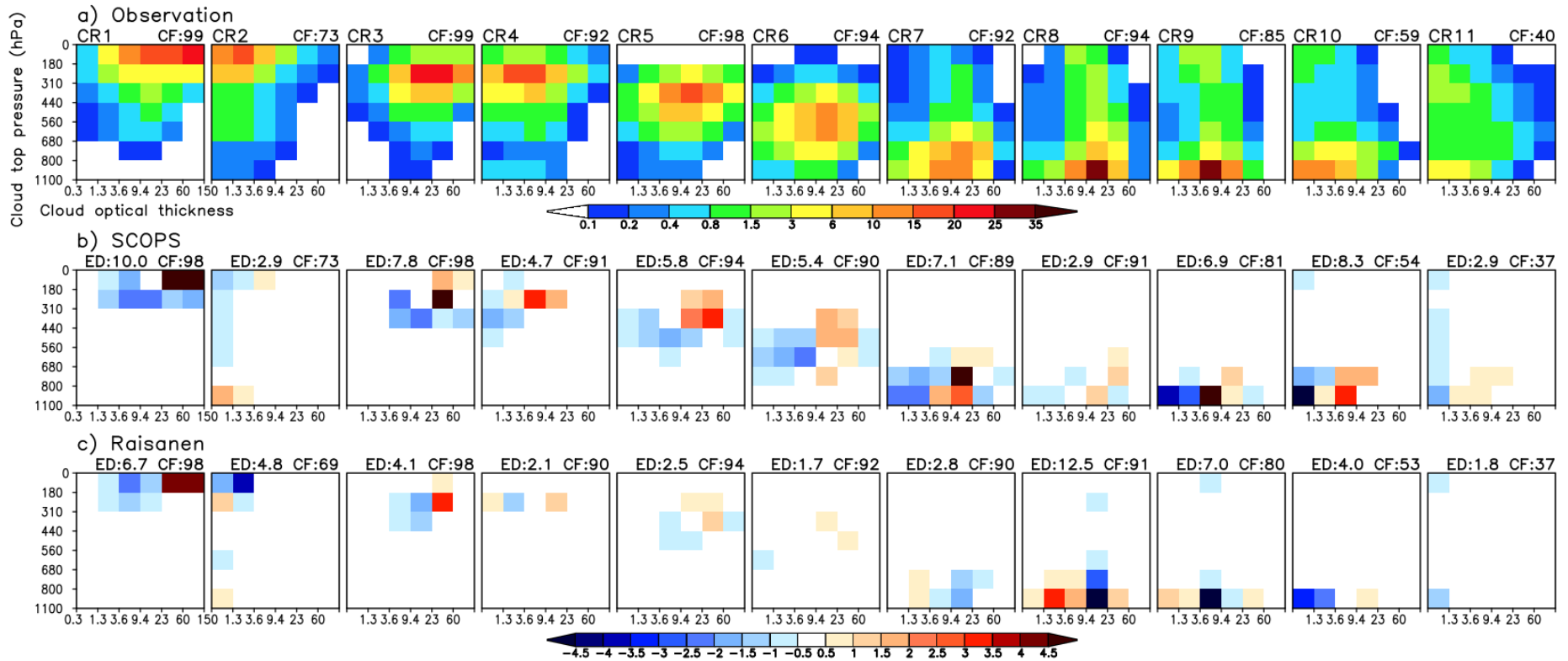


468

469 Fig. 6. One-year (2007) mean global ocean CTP-TAU joint histograms of observations (left
 470 panels) and from the subcolumn generator reconstruction (SCOPS-middle and Raisanen-right
 471 panels). For the middle and right panels the numbers represent the bin CF values themselves
 472 while the colors represent differences from the reference observed histograms.

473 The comparison of average histograms can be performed at a greater level of detail using
 474 again the Cho et al. (2021) MODIS CRs in the manner previously employed in assessing the
 475 quality of COD_{liq} filling. Mean joint CTP-TAU histograms by CR are compared between our
 476 reference COD fields and those produced by the two generators in Fig. 7. The top row shows
 477 the observed mean joint histograms by CR from CC COD fields for the year 2007 and for

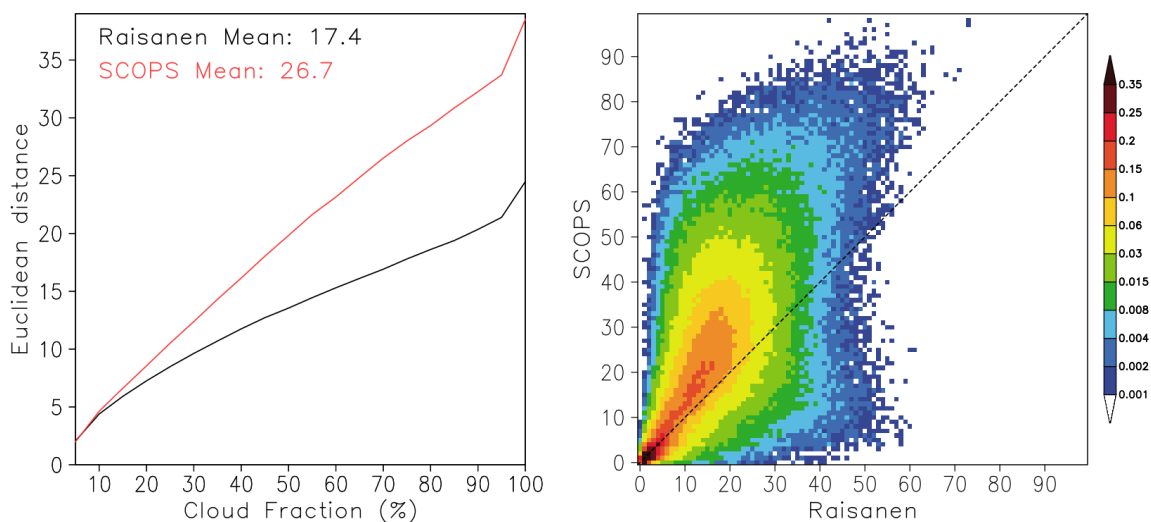
478 ocean only. The middle and bottom rows show joint histogram differences obtained by
479 subtracting the observed mean CR histogram from its reconstructed counterpart produced by
480 the two generators. Both generators are capable of closely reproducing the mean CF
481 corresponding to each CR, but again with a systematic underestimation. The Raisanen
482 generator performs overall better when performance is measured in terms of ED: for 8 out of
483 11 CRs, Raisanen EDs are smaller than SCOPS EDs. Raisanen is notably inferior for CR8
484 even though it reproduces the mean CF of this CR quite well. This CR along with CR9
485 appear to go against Raisanen's tendency of optical thin cloud underestimation; on the other
486 hand, SCOPS's underestimation of optically thin cloud is persistent across all CRs.



489 Fig. 7. Top row: Mean 42-bin joint CTP-TAU histograms from CC for 2007 aggregated by MODIS CR over oceans only. Middle row:
 490 difference between SCOPS-reconstructed and observed mean joint histograms by CR (negative values indicate underestimate by SCOPS).
 491 Bottom row: As middle row, but for the Raisanen generator.

492 2) INDIVIDUAL SCENES

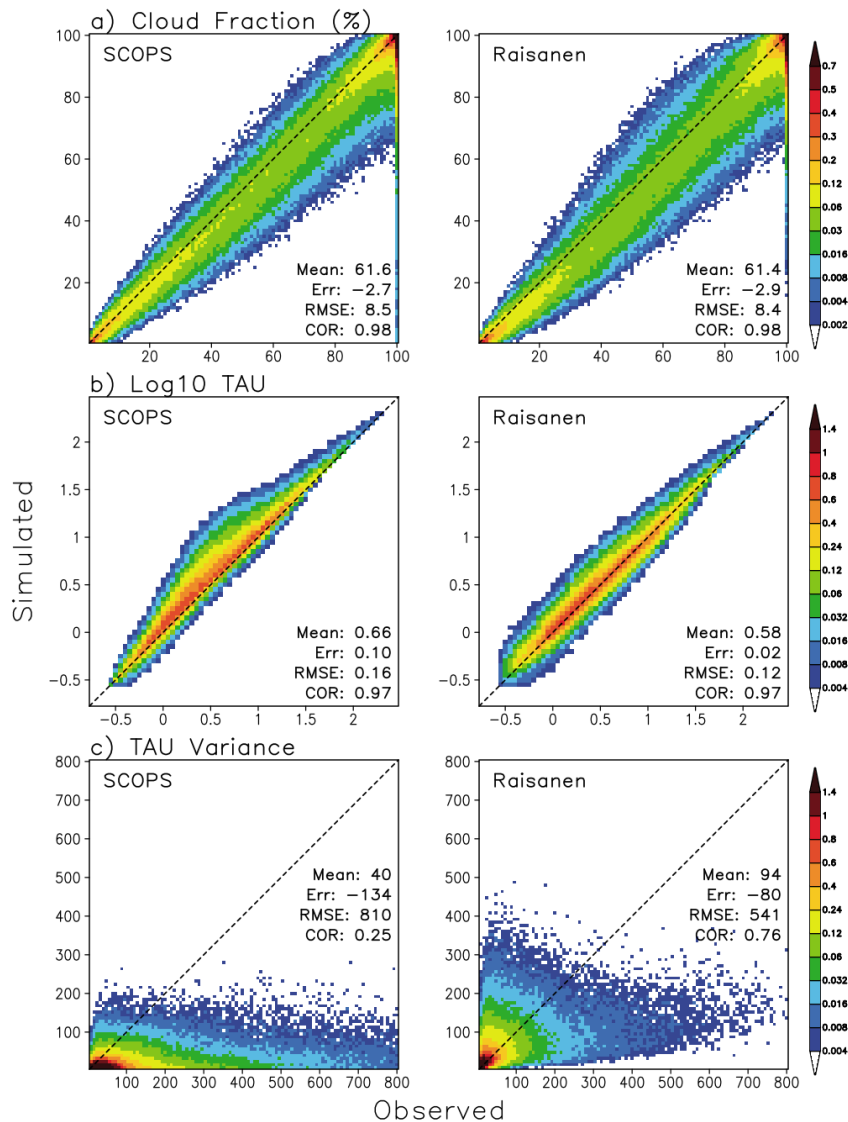
493 The performance of the two subcolumn generators can also be assessed at the scene level
494 by comparing the statistics of individual scene EDs and other metrics. Figure 8 depicts two
495 examples of how the EDs between simulated and observed 42-bin joint histograms for
496 individual 100-subcolumn scenes can be used to compare the performance of the two
497 generators. The left panel shows how the mean of scene EDs varies as a function of their CF.
498 The mean EDs of the two generators start to diverge at $CF \approx 20\%$, with the average EDs of
499 the Raisanen generator remaining consistently below those from the SCOPS generator,
500 indicating greater resemblance to observations on average. Two factors contribute to the
501 monotonic increases of ED with CF: larger CF values for joint histogram bins that are already
502 populated creating larger squared differences, and greater number of populated bins
503 contributing more terms to the sum of squared differences.



504

505 Fig. 8. Left panel: Average of individual 100-subcolumn scene EDs for the SCOPS and
506 Raisanen generators from scene 42-bin histograms, discretized by observed scene CF. Right
507 panel: Density plot of ED pairs from the two generators for individual scenes.

508 The right panel of Fig. 8 provides another glimpse of relative generator performance
509 using the same scene ED dataset. This time we create a density plot of ED pairs from the two
510 generators. The population of pairs above the diagonal containing scenes where SCOPS ED
511 exceeds Raisanen ED is much larger. There is a hint that SCOPS is doing better than
512 Raisanen at very small EDs (and thus likely small scene CFs), as indicated by the larger
513 density below the diagonal up to $ED \approx 5$, but the density asymmetry reverses quickly, with
514 far more scenes having greater ED for SCOPS than Raisanen above $ED \approx 10$.



515

516 Fig. 9. Top row: Density plot of pairs of observed and simulated (SCOPS: left column;
 517 Raisanen: right column) vertically projected scene CFs. Middle row: As top row, but for
 518 mean logarithmic TAU. Bottom row: As top row but for scene TAU variance.

519 Additional comparisons using density plots are shown in Fig. 9. This time we compare
 520 observed and simulated vertically projected scene-level CFs (top row), mean TAUs (middle
 521 row) and TAU variances (bottom row). The density plots for CFs indicate similar
 522 performance for the two generators and a preponderance of scene CF underestimates (i.e.,
 523 fewer points above the diagonal) over the full range of CFs, a result consistent with the
 524 overall underestimate of CF seen in Fig. 6. The two generators produce a broad overestimate
 525 of mean logarithmic TAU which is however less pronounced for Raisanen, consistent with
 526 the underestimate of optical thin and overestimate of optically moderate and thick cloud
 527 populations seen previously in averaged joint histogram results. One has to keep in mind

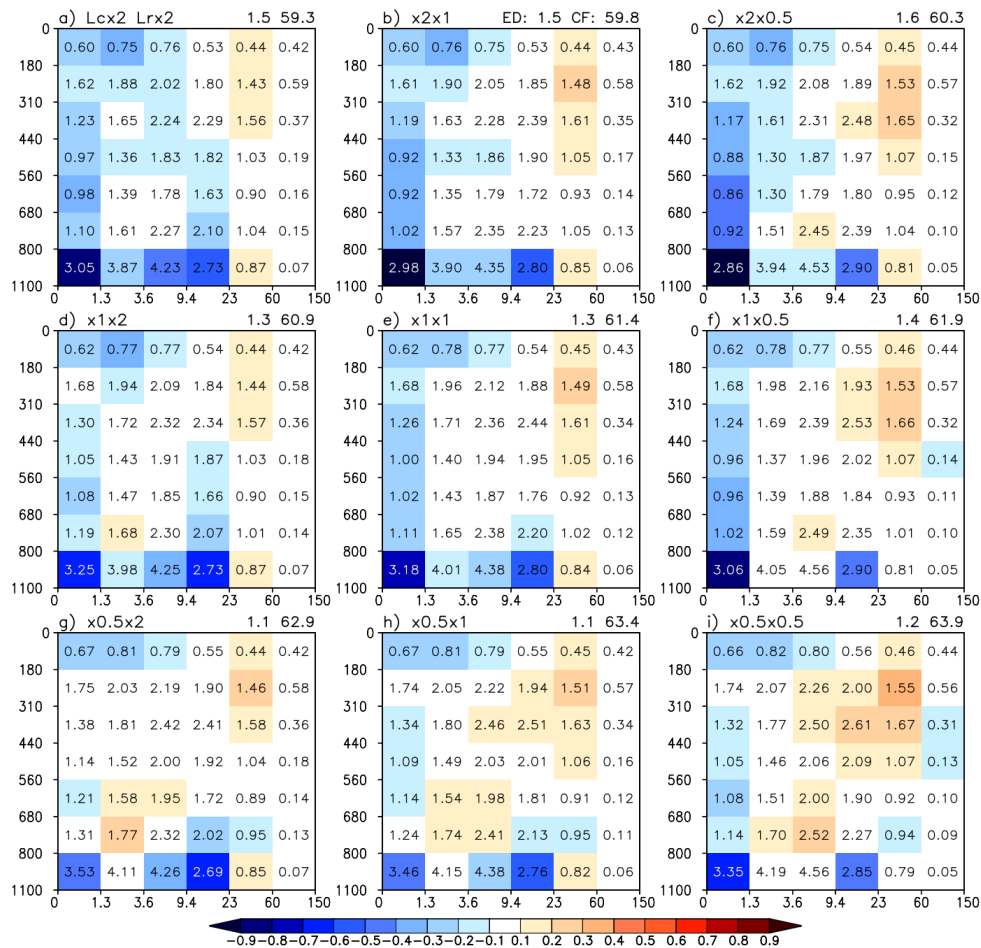
528 though when comparing to previous results that what is being assessed here is the mean
529 (logarithmic) TAU of individual scenes which corresponds to the TAU distribution of
530 individual scene joint histograms. The Raisanen generator appears to overestimate scene
531 mean TAU even when the scene is optically thin on average, something that does not occur
532 for SCOPS which is performing better for such scenes. But for intermediate TAUs Raisanen
533 is clearly better, ultimately yielding a smaller overall bias and RMSE. Where the two
534 generators diverge greatly is with respect to the variance of TAU: SCOPS's variance is far
535 below observations for the vast majority of scenes. This is hardly surprising given that
536 SCOPS distributes COD homogeneously across vertical layers and all variance of column
537 TAU comes from cloud occurrence overlap. Raisanen on the other hand tends to overestimate
538 variance of low and moderate magnitude. But once observed variance becomes very
539 pronounced (> 200), which is though quite rare, Raisanen typically underestimates it.

540 Even with this information, the radiative implications of generator performance at the
541 level of individual scenes are not easily predictable outside of an actual model
542 implementation. Yet, one can hypothesize on the potential impact of competing effects in the
543 shortwave part of the spectrum where cloud heterogeneity (subgrid variability) matters more:
544 reflected solar radiation would be underestimated when scene total vertically projected CF is
545 underestimated (both generators), but this would be compensated to some degree by
546 overestimates in mean TAU (both generators). Once variance of TAU is taken into account,
547 its smaller underestimate by Raisanen would contribute a smaller overestimate than SCOPS.
548 In other words, SCOPS can potentially provide greater compensation for its CF
549 underestimate through its mean TAU overestimate and TAU variance underestimate.
550 Radiative flux errors implied by generator deficiencies in producing correct subgrid
551 variability are discussed later, but only in the context of the grand-average joint histograms of
552 Fig. 6, and not for individual scenes where a more involved setup is required.

553 3) SENSITIVITY EXPERIMENTS WITH THE RAISANEN GENERATOR

554 As elaborated previously, the overlap of cloud occurrence and COD PDF in the Raisanen
555 generator is regulated by two decorrelation lengths which control the proportion with which
556 maximum and random overlap mix. In the results shown previously we used decorrelation
557 lengths values obtained by the parameterizations of Oreopoulos et al. (2012), specifically
558 their equations 10 and 11, which express decorrelation lengths as a function of latitude and
559 day of the year. This means that we have shown results from only a single realization of the

560 Raisanen generator. Yet, other decorrelation length combinations may give better agreement
 561 with observations. In this subsection, we therefore show results from a limited number of
 562 experiments applying simple scaling on the default values of the two decorrelation lengths.
 563 Specifically, we show results from eight experiments corresponding to all possible
 564 combinations of halving and doubling the default values of the two decorrelation lengths.



565

566 Fig. 10. Mean (2007 global average over ocean) CTP-TAU joint histograms for various
 567 experiments with the Raisanen generator using different decorrelation lengths, as indicated
 568 above each panel: “ $\times 0.5$ ” indicates halving, “ $\times 1$ ” same as default (Oreopoulos et al. 2012),
 569 and “ $\times 2$ ” doubling the decorrelation length; the scaling factor for cloud occurrence
 570 decorrelation length is given first and that for COD PDF overlap second. Also, above each
 571 panel the ED of the mean histogram from observations, and the histogram CF are provided.
 572 The center plot is the default experiment, i.e., Fig. 6c.

573 The center plot corresponds to the default experiment previously shown in Fig. 6c. Total
 574 CF and ED shown above each panel facilitate a quick assessment of performance. It can be
 575 seen that the observed CF (64.3) can be further approached by halving the cloud occurrence
 576 overlap decorrelation length (bottom row). This makes sense, because a smaller decorrelation

577 length makes cloud overlap more random which favors larger total CF. Doubling this
578 decorrelation length (top row) has the opposite than desired effect: overlap becomes more
579 maximum which decreases total CF compared to the already lower than observed CF of the
580 default experiment. Note that CF across rows can change even if the decorrelation length for
581 cloud occurrence overlap remains constant because of our previously discussed rejection of
582 $\text{TAU} < 0.3$ subcolumns which do not count towards cloudy skies; the number of such
583 subcolumns depends on the decorrelation length regulating COD PDF overlap.

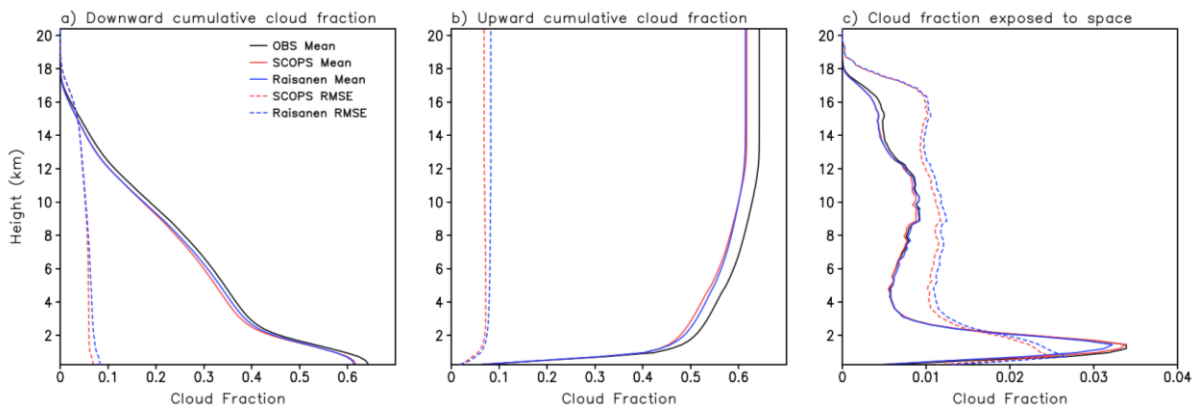
584 Matching total CF better does not guarantee superior (smaller) ED, as seen by the ED
585 value of the lower right panel which is (slightly) larger than that two panels of the same row
586 with worse CF. The best performing experiment is probably the one where decorrelation
587 length for COD PDF remains the same while decorrelation length for cloud occurrence is
588 halved (middle panel of bottom row). The fact that halving the default value for cloud
589 occurrence overlap improves results (compare middle and bottom row) is somewhat
590 surprising because previous results suggest that cloud occurrence overlap randomizes
591 substantially slower with cloud layer separation distance than COD PDF overlap (Räisänen et
592 al. 2004; Pincus et al. 2005; Oreopoulos and Norris 2011; Oreopoulos et al. 2012).

593 4) PROFILES OF OVERLAPPED CLOUD FRACTION

594 The generators by design should reproduce the observed mean CF profiles. However, the
595 cloud occurrence is imperfectly overlapped vertically and this can be captured by comparing
596 CFs of combinations of vertical layers between observations and simulations. Relevant
597 profiles conveying such information are those for cumulative CF and CF exposed to space as
598 in Barker (2008), the latter actually being the profile of differences between adjacent
599 cumulative CF values.

600 Figure 11 compares the profiles of mean downward and upward cumulative CF and CF
601 exposed to space between observations and simulations as well as the profiles of root mean
602 square differences in these quantities from the scene level data. Fig. 11 shows that in a mean
603 sense both generators handle cloud overlap quite well. The downward cumulative cloud
604 fraction profile (Fig. 11a) shows that the two generators handle the overlap of small CF in the
605 upper troposphere similarly and start to diverge only at a height of about 8 km, with the
606 cumulative CF of the Raisanen generator remaining closer to observations until abrupt CF
607 increases around 3 km (seen as a change in the slope of the curve) bring the two generator
608 curves closer together and with observations. This suggests errors in the CF overlap of the

609 generators at those heights which compensate for excessive overlap aloft causing
 610 underestimates of cumulative CF. At lower levels, the generators once again overlap too
 611 much, restoring the underestimates in cumulative CF. When CF is accumulated in the other
 612 direction (panel b) the excessive overlap of the generators starts appearing at around 1 km
 613 and once established continues unabated since getting closer to the observed curve would
 614 now require a severe underestimate of overlap. The RMSE curves from scene values suggests
 615 slightly more compensating error for Raisanen.



616

617 Figure 11. Downward (a) and upward (b) global (ocean) average cumulative CF for the
 618 observations and two generators. The rightmost panel (c) shows profiles of CF exposed to
 619 space. For all three panels, the profile of root mean square errors is also provided.

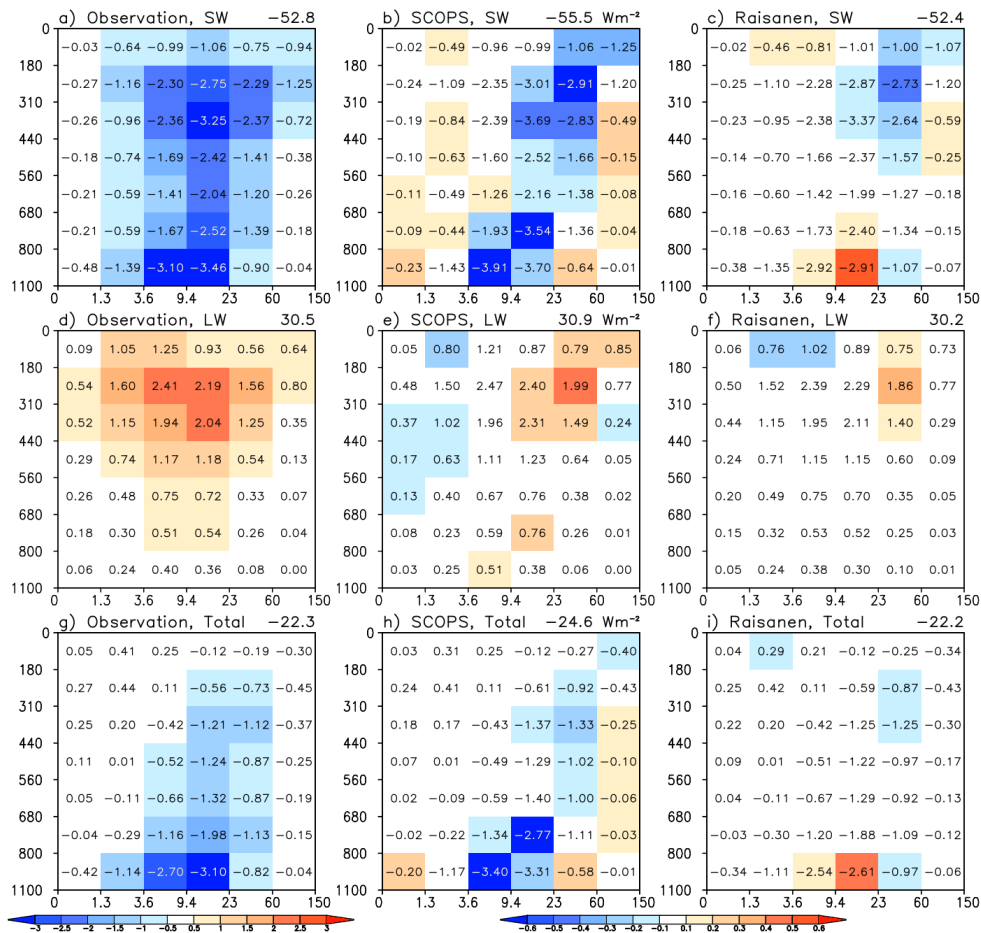
620 The rightmost panel shows the profile of CF exposed to space which peaks where the
 621 difference between two successive cumulative CF values is maximum. SCOPS outperforms
 622 Raisanen at that height, but the curves from the two generators are otherwise close and their
 623 deviations from observations are very small, until the highest levels of the troposphere where
 624 deviations re-emerge. Keep in mind that one can get good CF exposed to space even if the
 625 cumulative CF profile is biased because it is the shape of the profile that primarily matters.

626 5) FIRST ORDER RADIATIVE FLUX ERRORS

627 Here we present a simple method for translating the observed and simulated grand-
 628 averaged CF distribution resolved in CTP-TAU space shown in Fig. 6 to a cloud radiative
 629 effect distribution (CRE, i.e., the difference between all-sky and clear-sky fluxes) in the same
 630 phase space. To accomplish this, we use the concept of cloud radiative kernels (CRKs)
 631 introduced by Zelinka et al. (2012) in which the impact of cloud on the shortwave (SW),
 632 longwave (LW) and total (combined SW+LW) radiative flux is modeled for each of the 42
 633 bins of the CTP-TAU joint histogram. The CRKs give the change in radiative flux due to

634 clouds (i.e., compared to clear-sky conditions) per unit CF across all bins of the joint CTP-
635 TAU histogram. When this normalized difference between clear and overcast radiative
636 fluxes, i.e., the overcast CRE is multiplied with the distribution of CF shown in Fig. 6, a
637 distribution of CREs in the CTP-TAU phase space can be obtained. In our own
638 implementation of this CRE calculation, rather than using the model-derived Zelinka et al.
639 (2012) CRKs, we use roughly equivalent observational counterparts coming from the Clouds
640 and the Earth Radiant System (CERES) FluxByCldTyp product (Sun et al. 2022).
641 Specifically, from the monthly version of the product, we calculate globally-averaged ocean-
642 only pseudo-CRKs for the year 2007. These are shown in Fig. S1. When the pseudo-CRKs
643 are multiplied with the CF histograms of Fig. 6, the results of Fig. 12 are obtained. Above
644 each panel of Fig. 12, we provide the global (ocean only) CRE value.

645 As expected, LW CRE errors are less widespread than SW errors, based on the extent of
646 white space (indicating small errors) in the middle and right panels of the first and second
647 row. The Raisanen generator performs remarkably well on a global basis, but is aided in the
648 SW by non-negligible compensation of errors in individual bins; the bin errors are larger and
649 more extensive for SCOPS and result in a substantial 2.7 Wm^{-2} global SW CRE error.
650 Because SCOPS's LW CRE error is small, most of the SW CRE error propagates to total
651 CRE, which as a global value is actually perfect for Raisanen. The overestimates of SW CRE
652 by SCOPS are consistent with the previous discussion of Fig. 6, identifying underestimates of
653 optically thin clouds and overestimates of clouds with moderate and large optical thickness.
654 The distribution of SW CRE errors in Fig. 12 (colors) tracks the distribution of CF errors in
655 Fig. 6 (colors) to some extent, but the one-to-one mapping is imperfect because CF errors for
656 optically thin clouds are radiatively inconsequential. Binned LW CRE errors outside the ± 0.1
657 Wm^{-2} range are more frequent for SCOPS, but generally rare besides the more sensitive to
658 LW radiation high clouds. Errors in total CRE outside this range are limited to very few bins
659 for Raisanen, but quite more frequent for SCOPS when $\text{TAU} > 3.6$.



660

661 Fig. 12. Distributions of shortwave (SW), longwave (LW) and total = (SW+LW) CREs in the
 662 CTP-TAU phase space (numbers in all panels) and errors (colors in the middle and right
 663 panels) obtained by multiplying the “pseudo-CRKs” of Fig. S1 with the CF histograms of
 664 Fig. 6. Above each panel the global (ocean only) CRE value is provided.

665

666 4. Discussion and conclusions

667 We have created a dataset of 2D cloud optical depth fields that is extensive enough for
 668 statistical evaluation of cloud subcolumn generators used by satellite simulators in GCMs for
 669 emulating real-world subgrid variability. The dataset is based on active CloudSat-CALIPSO
 670 global observations for the year 2007. The appropriateness of such observations compared to
 671 those obtained from passive instruments stems mainly from their ability to resolve the
 672 vertical (height-dependent) variability of clouds; passive observations typically provide only
 673 column-integrated quantities. Our dataset was constructed by combining two CloudSat
 674 products. Upon doing this, it quickly became apparent that the portion of the cloud field
 675 contributed by liquid clouds required corrections as it was found to lack retrievals for a non-

676 negligible fraction of cells identified to be cloudy and of liquid phase in the CALIPSO-
677 enhanced CloudSat 2B-CLDCLASS-LIDAR product. We came up with a relatively simple
678 filling scheme for the missing clouds whose performance was assessed via comparisons with
679 MODIS-Aqua cloud retrievals. The comparisons were based on discretized cloud fraction
680 (CF) distributions in the form of joint histograms of cloud top pressure (CTP)-cloud optical
681 thickness, (TAU), and even encompassed a MODIS cloud regime (CR) segregation in order
682 to gauge possible dependences on cloud types.

683 The reconstructed 2D cloud optical depth fields including our improvements were
684 segmented into scenes comprising 100 subcolumns (~110 km) and whose mean CF and TAU
685 profiles were passed to the two generators to produce their own set of 100 subcolumns as a
686 simulated version of the scene. All three sets of subcolumns were passed to COSP's MODIS
687 simulator which transformed them to joint CTP-TAU scene histograms for potential
688 averaging across spatiotemporal scales. Note that there is no fundamental reason the
689 generators should be configured to also produce 100 subcolumns, since individual
690 subcolumns are never compared, but rather scene-level cloud field properties. When we
691 experimented with different numbers of subcolumns, we noticed a slow progressive
692 improvement in performance as the subcolumn number grew, but without much benefit
693 above 100 subcolumns. Actually, on the opposite side of fewer subcolumns, even as few as
694 20 would not have affected the results of this study substantiatively.

695 A sensible way of assessing the performance of the two generators in terms of joint
696 histograms is to compare the Euclidean distances of their grand averages against their
697 observational counterpart (which again can be made more detailed by applying a CR
698 breakdown), or statistics of EDs coming from individual scene histograms. Simpler
699 comparisons bypassing joint histograms altogether are of course also possible using
700 vertically-integrated quantities directly derived from the scene's subcolumns, such as
701 vertically-projected cloud fraction, (logarithmic) mean scene TAU, and variance of TAU.

702 Both types of comparisons described above reveal a clear superiority of the Raisanen
703 generator in our default implementation, i.e., using the parameterization of decorrelation
704 lengths for cloud occurrence and cloud optical depth PDFs proposed by Oreopoulos et al.
705 (2012). Sensitivity experiments allowing the decorrelation lengths to vary unveil that simple
706 modifications to this parameterization may yield even better results. Nevertheless, the main
707 deficiencies of both generators, namely an overestimation of overlap and an underestimation

708 of the occurrence of optically thin clouds remains a persistent theme. For SCOPS, this error
709 combines with overestimates of optically thicker clouds yielding an overestimate of
710 shortwave cloud radiative effect that is close to 3 Wm^{-2} on global scales, but for Raisanen
711 where optically thicker clouds are better simulated, the shortwave CRE error remains
712 contained.

713 We are fully aware that vertically integrated quantities and joint histograms describing
714 how integrated extinction and location of the highest cloud co-vary cannot be viewed as the
715 only way to evaluate subgrid variability simulated by generators. Even in a two-dimensional
716 world, the subgrid-scale profiles of cloud occurrence and cloud (liquid and ice) condensate
717 and particle size should be well-reproduced since they may matter for radiative heating rate
718 profiles and the physical parameterizations developed for GCMs. We have taken a first step
719 by evaluating subgrid cloud occurrence profiles via grand-average profiles of cumulative
720 cloud fraction and cloud fraction exposed to space, which revealed satisfactory skill of
721 similar parity for both generators. Nevertheless, one should keep in mind that knowing and
722 being able to simulate the subgrid variability of a wider range cloud properties is imperative
723 for an assessment of subgrid realism based on radiative flux and heating rate profiles. We
724 anticipate to be in a position to confront the simulators with such stricter tests in future
725 endeavors. Efforts of this kind would also benefit by a wider range of choices of empirical
726 and easy to use cloud subcolumn generators that operate on the (more readily available from
727 observations) condensed part of total water content, criteria that are unfortunately not met by
728 other existing generators (Norris et al. 2008; Larson and Schanen 2013).

729

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737

738 *Data Availability Statement.*

739 The CloudSat products used in this study are available at
740 <https://www.cloudsat.cira.colostate.edu>. The CERES FluxByCldType product is available at
741 <https://ceres.larc.nasa.gov/data/>. The COSP simulator which includes the MODIS simulator
742 can be downloaded from <https://github.com/CFMIP/COSPV2.0>.

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