

# Optimal Estimation Framework for Ocean Color Atmospheric Correction and Pixel-level Uncertainty Quantification

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**Abstract:** Ocean color remote sensing requires compensation for atmospheric scattering and absorption (aerosol, Rayleigh, and trace gases), referred to as atmospheric correction (AC). AC allows inference of parameters such as spectrally resolved remote sensing reflectance ( $R_{rs}(\lambda)$ ;  $\text{sr}^{-1}$ ) at the ocean surface from the top-of-atmosphere reflectance. Often, the uncertainty of this process is not fully explored. Bayesian inference techniques provide a simultaneous AC and uncertainty assessment via a full posterior distribution of the relevant variables, given the prior distribution of those variables and the radiative transfer (RT) likelihood function. Given uncertainties in the algorithm inputs, the Bayesian framework enables better constraints on the AC process by using the complete spectral information compared to traditional approaches that use only a subset of bands for AC. This paper investigates a Bayesian inference research method (Optimal Estimation, OE) for ocean color AC by simultaneously retrieving atmospheric and ocean properties using all visible and near-infrared spectral bands. The OE algorithm analytically approximates the posterior distribution of parameters based on normality assumptions and provides a potentially viable operational algorithm with a reduced computational expense. We developed a Neural Network (NN) RT forward model look-up-table-based emulator to increase algorithm efficiency further and thus speed up the likelihood computations. We then applied the OE algorithm to synthetic data and observations from the MODerate resolution Imaging Spectroradiometer (MODIS) on NASA's Aqua spacecraft. We compared the  $R_{rs}(\lambda)$  retrieval and its uncertainty estimates from the OE method with in-situ validation data from the SeaWiFS Bio-optical Archive and Storage System (SeaBASS) and Aerosol Robotic Network Ocean Color (AERONET-OC) datasets. The OE algorithm improved  $R_{rs}(\lambda)$  estimates relative to the NASA standard operational algorithm by improving all statistical metrics at 443, 555, and 667 nm. Unphysical negative  $R_{rs}(\lambda)$ , which often appear in complex water conditions, was reduced by a factor of 3. The OE-derived pixel-level  $R_{rs}(\lambda)$  uncertainty estimates were also assessed relative to in-situ data and were shown to have skill.

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## 1. Introduction

The atmospheric correction (AC) process in ocean color (OC) remote sensing involves separating and removing the atmospheric contributions (aerosol and gas scattering and absorption) and ocean surface signal from the spectral reflectances observed by a satellite radiometer at the top of the atmosphere (TOA) [1–4]. The science of OC aims to quantify and assess the biogeochemical properties of aquatic ecosystems by interpreting their visible water-leaving spectra. These spectral reflectance signals emerging from the water body primarily

47 depend on the inherent optical properties (IOPs; absorption and scattering properties) of the  
48 biogeochemical constituents dissolved or suspended within the water column, in combination  
49 with the IOPs of seawater itself. These constituents include organic and inorganic hydrosols  
50 suspended in seawater, colored dissolved organic matter (CDOM), and photosynthetic  
51 pigments within phytoplankton. The primary heritage OC data product is the near-surface  
52 concentration of the photosynthetic pigment chlorophyll-a (Chl-a; mg m<sup>-3</sup>), which provides a  
53 convenient and widely-used proxy for phytoplankton biomass [5]. Phytoplankton biomass is  
54 an essential component of the Earth's carbon cycle, and producing climate-quality OC data  
55 records is generally necessary for Earth climate studies [6–8].

56 Chlorophyll-a concentrations are typically derived through an empirical relationship based  
57 on coincident in-situ observations of Chl-a and the aforementioned water-leaving radiometric  
58 signal, namely spectral remote sensing reflectances ( $R_{rs}(\lambda)$ ; sr<sup>-1</sup>), which are the radiances  
59 exiting the water column normalized to downwelling surface irradiance. However, since the  
60 atmospheric radiance contribution to the TOA signal is typically between 85-90% of the total,  
61 a small uncertainty in the AC can lead to large uncertainties in the ocean radiances and derived  
62 OC products [3]. The Rayleigh scattering of the atmosphere is effectively known, based on  
63 assumed molecular properties [9], yet it can introduce additional uncertainties in the AC [10].  
64 However, the aerosol signal must be inferred from the satellite observations since the aerosol  
65 type and concentration vary spatially and temporally in the atmosphere [11].

66 Inferring useful information from satellite-derived radiometry is accomplished by solving  
67 the inverse problem, which is ill-posed and under-constrained for AC [12,13]. The TOA  
68 reflectance of multi-spectral, single viewing sensors such as the Sea-viewing Wide Field-of-  
69 view Sensor (SeaWiFS) [14], MODerate resolution Imaging Spectroradiometer  
70 (MODIS) [15], and Visible Infrared Imaging Radiometer Suite (VIIRS) [16], contain less  
71 information than what is required to find an unambiguous solution to a complex Atmosphere-  
72 Ocean (AO) model. To address this issue, future NASA missions will dedicate more advanced  
73 instruments to increase the observed information. For example, NASA's Plankton, Aerosol,  
74 Cloud, ocean Ecosystem (PACE) mission will host three instruments that will measure the AO  
75 system with unprecedented spectral and angular information [17]. The primary instrument is  
76 the Ocean Color Instrument (OCI), which is being developed at the Goddard Space Flight  
77 Center (GSFC) and is a hyperspectral scanning radiometer that measures the light from 320 to  
78 890 nm at 5-nm spectral resolution and 2.5-nm spectral sampling, and at seven discrete short-  
79 wave infrared (SWIR) channels: 940, 1,038, 1,250, 1,378, 1,615, 2,130, and 2,260 nm. The two  
80 other instruments, the Hyper-Angular Rainbow Polarimeter 2 (HARP2) and the Spectro-  
81 Polarimeter for Exploration (SPEXone), are aimed at studying aerosols and clouds and are  
82 multi-angular polarimeters (MAPs) developed and contributed by external partners. To employ  
83 OCI's unprecedented hyperspectral capabilities for ocean applications, an accurate AC process  
84 with capabilities beyond the current algorithm designed for multispectral sensors is necessary.  
85 One such OCI algorithm has been developed, which relies on the proven heritage AC  
86 capabilities of the NASA standard algorithm, extended to hyperspectral data and with added  
87 capabilities to seamlessly utilize the SWIR channels for AC in coastal and inland waters (e.g.,  
88 the multiband AC (MBAC) algorithm) [18]. In addition to this OCI-only AC, the MAPs will  
89 provide more complex aerosol information to constrain the AC for OCI [19]. Thus,  
90 establishing a probabilistic framework that can combine the information from two or three  
91 independent instruments with different spatial and spectral resolutions, information content,  
92 and measurement uncertainty characteristics using a Bayesian framework is a logical next step  
93 to advance the AC performance and the quality of OC retrievals from PACE.

94 Deterministic (that is, non-stochastic) AC methods have been and are currently being used  
95 as the standard processing algorithms for satellite remote sensing of OC [3,4,11,20,21]. These  
96 methods maximize the likelihood (i.e., match radiative transfer prediction models to the  
97 observations) of the AC parameters such as the aerosol, surface, and ocean optical properties  
98 and often do not directly provide an estimate of the uncertainty on these parameters or consider

99 the uncertainty in the algorithm inputs and parameters. The forward likelihood model is  
100 parametrized from radiative transfer simulations in a (pre-computed) look-up-table (LUT) for  
101 computational efficiency. These LUTs contain the modeled TOA reflectances for a pre-  
102 determined set of relevant parameters within a typical range.

103 NASA's current operational AC algorithm for OC sensors is based on Gordon and  
104 Wang [11], with the current implementation detailed in [20]. The algorithm determines and  
105 removes atmospheric (i.e., Rayleigh and aerosol) and surface (i.e., whitecaps, glint)  
106 reflectances through a LUT search of pre-computed reflectances as derived using vector  
107 radiative transfer (VRT) simulations. One LUT contains the spectral TOA Rayleigh reflectance  
108 for different geometries and surface wind speeds. The aerosol reflectance LUTs are  
109 parametrized for 80 different aerosol optical models representing the range of relative humidity  
110 (RH) and fine-mode volume fractions [22]. These models assume a complex refractive index  
111 and bimodal effective radius and variance for coarse and fine aerosol particles, determined from  
112 Aerosol Robotic Network (AERONET) observations [23,24]. The absorption coefficients of  
113 trace gases such as ozone, water vapor, oxygen, and methane are stored in LUTs and applied  
114 to compensate for atmospheric path absorption given the gas concentration and an assumed  
115 vertical profile. Ancillary information, including relative humidity, ozone and water vapor  
116 concentrations, and wind speed, are provided as auxiliary inputs to constrain the inversion.

117 The aforementioned models are explicitly parameterized to ensure that the inversion is not  
118 mathematically ill-posed and act as a constraint to reduce ambiguity and the potential for  
119 degenerate solutions. The aerosol optical models are assumed to be non- to weakly- absorbing  
120 and to have a fixed vertical profile. With these assumptions, only two pieces of information are  
121 needed for the AC: aerosol optical depth (AOD, i.e., loading) and spectral dependence (i.e.,  
122 from the optical model), both of which can be determined using a pair of near-infrared (NIR)  
123 or SWIR wavelengths (dependent on the sensor). However, the presence of strongly absorbing  
124 aerosol types confounds this process, and the AC typically produces either underestimated or  
125 non-physical negative ocean radiances in the blue part of the spectrum . This is because the  
126 algorithm relies on the extrapolation of the model information determined from the longer NIR  
127 or SWIR wavelengths (where the ocean is dark) to the visible (where it is not). The spectral  
128 information in the longer wavelengths is insufficient to discern absorbing from non-absorbing  
129 aerosols, as they differ primarily in the shorter wavelengths and do not have a discriminating  
130 signature in the NIR. Thus, the solution can be ambiguous and aerosol absorption cannot be  
131 reliably inferred unless the algorithm is constrained by additional external information.

132 Pixel-level Uncertainty Quantification (UQ) is critical in assessing the fidelity of  
133 geophysical retrievals within the Earth system. UQ also allows for identifying issues and  
134 limitations in retrieval algorithms due to inherent modeling assumptions, measurement  
135 uncertainties, and gaps in knowledge and sources of uncertainties. Traditionally, uncertainties  
136 in  $R_{rs}$  are based on the reported average discrepancy between the satellite-derived and in-situ  
137  $R_{rs}$  [27,28]. UQ has been attempted through various techniques such as Bayesian  
138 approaches [29,30], Monte Carlo simulations [31], or analytical error propagation of sensor  
139 random noise [32]. A new approach was developed to estimate pixel-level uncertainties for  
140 Sentinel-3 Ocean and Land Colour Imager (OLCI) based on an ensemble of neural network  
141 atmospheric correction models for coastal waters, showing an estimate of the  $R_{rs}$  uncertainty  
142 product that is feasible to apply operationally [33].

143 Because of the ill-posed nature of the problem, Bayesian approaches are well-suited for AC  
144 and indeed have been applied widely for aerosol [34,35], cloud [36,37], atmospheric trace gas  
145 profiling [38,39], and OC [29,40–42] retrievals. For a given model, the aerosol and ocean  
146 properties can be inferred, along with the associated uncertainties, in the form of a posterior  
147 distribution. Bayes theorem calculates conditional probabilities and updates a prior belief when  
148 new data (evidence) is introduced [43] such that  $P(x|y_{obs}) \propto P(y_{obs}|x) \times P(x)$ , where  $P$   
149  $(x|y_{obs})$  is the posterior distribution or probability of the variables needed for AC,  $x$ , given the  
150 observed data,  $y_{obs}$ . The posterior distribution is proportional to the likelihood function,  $P(y_{obs}$

151  $|x)$ , and the prior probability of the variables  $P(x)$ . The likelihood function describes the  
152 probability of the observed TOA reflectance,  $y_{obs}$ , given the variables  $x$ . Here, the likelihood  
153 function is the forward model based on RT and  $x$  are the variables that describe the state of the  
154 ocean and atmosphere, such as the aerosol and ocean optical properties or ancillary data. The  
155 AC algorithm requires some prior information,  $P(x)$ , such as the relative humidity, surface  
156 pressure, ozone, and water vapor that can, along with their uncertainties, be directly  
157 incorporated into the prior (in contrast to a non-Bayesian retrieval where these values are  
158 assumed to be true). In the deterministic sense, the likelihood is typically written as  $\mathbf{y}_{obs} = \mathbf{F}(\mathbf{x})$   
159  $+ \epsilon$ , where  $\mathbf{F}(\mathbf{x})$  is the forward operator (model), and  $\epsilon$  is the uncertainty associated with that  
160 model. In Bayesian terminology, the likelihood probability is modeled as a statistical  
161 distribution, assumed normal in this case, with mean and variance determined from the forward  
162 model.

163 There are various numerical techniques that approximate Bayes' theorem. The grid  
164 approximation is the most straightforward inference engine by approximating the continuous  
165 variables,  $x$ , on a finite parameters grid. The posterior is calculated by multiplying the  
166 likelihood probability and prior probability evaluated at each grid point: a non-iterative brute  
167 force approach. The Generalized Nonlinear Retrieval Analysis (GENRA) algorithm for cloud  
168 properties retrievals utilizes the grid approximation to retrieve, for example, the posterior of  
169 two independent parameters: cloud optical depth and effective radius [44]. Expanding the grid  
170 to higher dimensions, however, can be computationally challenging. But, when the dimension  
171 is low (e.g.,  $<5$ ), the method is tractable and yields inference results within a reasonable  
172 computational time [30]. This manuscript will focus on the normal or quadratic inference  
173 approximation, Optimal Estimation (OE), as used in Rodgers's (2000) formalism [45]. This is  
174 a widely used inverse algorithm within the atmospheric science community [34,35,37–39,46].

175 Due to the high computational demand of an OE inference algorithm that fully considers  
176 the correlation structure in the observations and model, a fast likelihood function (i.e., forward  
177 model) evaluation is necessary. There are several ways to approximate the forward model in  
178 the iterative inversion process. A rigorous RT computation is the most accurate; however, it is  
179 computationally slow for a complex AO system. The LUT parametrization of the RT, such as  
180 NASA's operational tables, are pre-computed and stored for a pre-determined grid of  
181 parameters, thus requiring multi-dimensional interpolation for each iteration in the retrieval.  
182 The LUT parametrization is accurate and sufficiently fast (for low-dimensional problems) for  
183 deterministic inversion using, for example, non-Bayesian methods or low-dimensional OE.  
184 However, in the high-dimensional inverse problem of the coupled AO system, the likelihood  
185 function based on the associated LUT interpolations becomes computationally costly. We  
186 developed a deep Neural Network (NN) with a simple multi-layer perceptron (MLP)  
187 architecture that efficiently and accurately emulates the forward RT LUT parameterization to  
188 speed up the forward model computations. In this case, the Forward RT NN is a non-linear  
189 function approximator of the radiative transfer equation.

190 Forward model emulators using MLP-NN have been used to speed up the RT computations  
191 in modeling solar radiation [47] and satellite sensor simulators [48]. A forward model  
192 emulator was also used in inverting geophysical properties using a Gaussian Process model for  
193 land surface parameter inference in spectroscopic remote sensing of land and ocean  
194 surfaces [50,51], and polarimetric remote sensing of aerosols [52,53]. Note that the NN is not  
195 required for a Bayesian retrieval; it is merely a tool adopted to increase the computational  
196 efficiency of the analysis. A forward RT NN emulator provides advantages over an inverse RT  
197 NN model that estimates geophysical parameters from observations. The forward NN model is  
198 easier to train as there is a 1-1 mapping between the geophysical inputs to the RT and the  
199 predicted TOA reflectances, avoiding the ill-posed, non-uniqueness, and overfitting problem  
200 common with inverse neural networks due to multicollinearity among variables [54]. The  
201 Jacobian matrix (see later) is necessary for the iterative inversion scheme. A NN forward model  
202 can efficiently provide the Jacobian using the backpropagation chain rule algorithm and, with

203 modern computer languages, Automatic Differentiation (AD) [55–57]. Additionally, the  
204 forward RT NN can be used in any iterative or stochastic inversion models that allow  
205 uncertainty propagation or estimation, which is more challenging for an inverse NN. However,  
206 several studies aimed at assessing the variability in NN weights and their relationship to  
207 geophysical parameter uncertainties showed promising results within their application  
208 domain [54,58,59].

209 This work aims to establish an inference framework for the AC that can be potentially  
210 applied to global datasets for a wide range of environmental conditions and provide pixel-level  
211 uncertainty. The algorithm relies on the simultaneous estimation of the atmospheric parameters  
212 (i.e., AOD and fine-mode fraction, and ancillary related parameters), as well as the ocean’s  
213 inherent optical properties established through the Generalized Inherent Optical Properties  
214 (GIOP) model [60] (i.e., absorption coefficients of seawater, phytoplankton and colored  
215 dissolved plus detrital matter and backscattering coefficients for seawater and particle matter).  
216 It can exploit the information content of all spectral bands available for an instrument. Our OE  
217 algorithm finds the optimal solution to the TOA reflectance state vector and estimates the pixel-  
218 level uncertainty (i.e., the error covariance matrix) of  $R_{rs}$ . The availability of a spectral error  
219 covariance matrix can be used as an input for estimating IOP and biogeochemical product  
220 uncertainties [61]. The model considers the uncertainty at the TOA due to instrument random  
221 noise, ancillary data uncertainty, and the systematic and forward model uncertainty estimated  
222 at the Marine Optical BuoY (MOBY) site.

223 The OE algorithm effectively recasts the standard NASA algorithm approach into a  
224 Bayesian framework. The goal of this framework, however, manifests in several ways:

- 225 • We aim to assess the performance of the algorithm’s retrievals of  $R_{rs}$  for a wide range of  
226 water conditions and provide validation metrics relative to in-situ data and compared to  
227 the NASA standard algorithm.
- 228 • We aim to assess the performance of the pixel-level uncertainty of  $R_{rs}$  relative to the  
229 error between the in-situ data and the satellite retrievals.
- 230 • We aim to assess the algorithm’s performance on an entire scene retrieval since the  
231 algorithm is computationally fast as it relies on the NN model to emulate the forward  
232 calculations and provide the Jacobian matrix necessary for the optimization and the  
233 uncertainty estimates.

234  
235 The structure of this paper is as follows. Section 2 provides details of the physical forward  
236 model based on radiative transfer computations for the atmospheric LUTs, with the analytical  
237 forward model of the GIOP algorithm described in [Supplement 1](#). We follow that with a  
238 discussion on the development of the NN model that serves as the likelihood function for the  
239 OE algorithm and the associated NN training process. Section 3 details the assumed uncertainty  
240 sources. Section 4 describes the OE algorithm architecture, selection of priors, uncertainty  
241 propagation, and derivation of  $R_{rs}$  through the AC process. Section 5 describes the validation  
242 datasets, including the in-situ SeaBASS and AERONET-OC datasets and satellite imagery  
243 from the Moderate Resolution Imaging Spectroradiometer (MODIS), along with matchup  
244 statistics and uncertainty validation metrics. In Section 6, we evaluate the performance of the  
245 NN model as well as the  $R_{rs}$  retrieval from the OE algorithm. The OE algorithm is also evaluated  
246 on a real validation dataset and compared with the operational algorithm. Finally, we discuss  
247 the results and provide a conclusion in Section 7.

## 249 **2. Methods**

### 250 *2.1 Forward model*

251 The TOA reflectance is based on a radiative coupling of various components of the atmosphere,  
252 ocean, and surface [20]. The forward model relates the retrievable geophysical parameters to

253 the TOA observations measured by the satellite sensor. For a clear (cloud-free) ocean pixel, the  
 254 TOA reflectance is calculated as follows:

$$\rho_t(\lambda;Geom) = \left( \rho_{path}(\lambda;Geom) + \rho'_w(\lambda;Geom) + \rho'_{surface}(\lambda;Geom) \right) \times T_g \quad (1)$$

$$(\lambda;Geom).$$

255 It is a function of *Geom* (i.e., solar zenith  $\theta_0$ , sensor zenith  $\theta$ , and relative azimuth  $\varphi$ ), and  
 256 wavelength,  $\lambda$ ;  $\rho_{path}(\lambda;Geom)$  is the path reflectance due to scattering and absorption by air  
 257 molecules (Rayleigh scattering) and aerosols bounded by the sea surface;  $\rho'_w(\lambda;Geom)$  is the  
 258 ocean body reflectance, and  $\rho'_{surface}(\lambda;Geom)$  is the reflectance contribution from surface  
 259 glint and whitecaps, where both  $\rho'_w(\lambda;Geom)$  and  $\rho'_{surface}(\lambda;Geom)$  are expressed at the TOA  
 260 after propagation through the atmosphere.  $T_g(\lambda;Geom)$  is the two-way absorbing gas  
 261 transmittance along the solar and sensor zenith. The path reflectance is a summation of two  
 262 terms, the Rayleigh reflectance and the aerosol reflectance (including the aerosol-Rayleigh  
 263 interaction):

$$\rho_{path}(\lambda;Geom) = \rho_r(\lambda;Geom) + \rho_a(\lambda;Geom). \quad (2)$$

264 The  $\rho_r(\lambda;Geom)$  term is calculated through the tabulation of VRT simulations. The  
 265 Rayleigh optical depth is calculated from [9]. Although the path reflectance term is shown in  
 266 Eqs. (1) and (2) as a function of only wavelengths and geometry, the Rayleigh reflectance is  
 267 also a function of surface pressure and wind speed. The former is needed to know the total  
 268 number of air molecules in the atmospheric column. The latter is to account for the interaction  
 269 of Rayleigh scattering with the wind-roughened sea surface. The surface roughness model is  
 270 from Cox and Munk (1954), and the effect of pressure variation is modeled by [62].

271 The second term in Eq. (2) is the aerosol reflectance, calculated through the VRT  
 272 simulations for each of 80 different bimodal aerosol models from [22], consisting of assumed  
 273 aerosol microphysical properties for a pre-determined set of 8 near-surface atmospheric RHs  
 274 and 10 fine-mode volume fractions. The aerosol vertical profile in the atmosphere is taken  
 275 from [63]. The aerosol reflectance calculations include the effects of multiple scattering and  
 276 molecule-aerosol interaction within the atmosphere. Note that these simulations also provide  
 277 the molecule-aerosol diffuse transmittance along the solar and sensor directions,  $t_{sol}(\lambda, Geom)$   
 278 and  $t_{sen}(\lambda, Geom)$ , respectively, used later to propagate the water and surface reflectance to the  
 279 TOA.

280  $\rho'_w(\lambda;Geom)$  is the ocean reflectance at TOA. The bottom of atmosphere (BOA) ocean  
 281 reflectance  $\rho_w(\lambda;Geom)$  is calculated through a forward model that provides the ocean  
 282 reflectance as a function of Chl-a, *Geom*, and spectral IOPs. The BOA reflectance contribution  
 283 is attenuated by the diffuse transmittance of the atmosphere, such that  $\rho'_w(\lambda;Geom) = t_{sen}$   
 284  $(\lambda, Geom) \times \rho_w(\lambda;Geom)$ . The BOA ocean reflectance are generated from an ocean  
 285 reflectance model (ORM) that derives the above-water remote sensing reflectance,  $R_{rs}(\lambda; sr^{-1})$ ,  
 286 which is converted from nadir geometry to the desired solar and sensor path geometries using  
 287 the bidirectional reflectance distribution function ( $f_{brdf}$ ) of [64], and then propagated to the  
 288 TOA as:

$$\rho'_w(\lambda;Geom) = \pi R_{rs} t_{sol} t_{sen} / f_{brdf}. \quad (3)$$

289  $R_{rs}(\lambda)$  is modeled using the quasi-single scattering approximation ORM [65] included  
 290 within the Generalized Inherent Optical Property algorithm framework (GIOP) [60]. Given the  
 291 IOP data as an input to the GIOP ORM forward model, we can simulate a realistic  $R_{rs}(\lambda)$   
 292 distribution for various conditions observed by ocean color sensors. Details of the GIOP  
 293 forward model are provided in [Supplement 1](#).

294 The surface reflectance,  $\rho'_{surface}(\lambda;Geom)$ , is the light scattered by the air-sea interface. It  
 295 has two terms: the direct sun glint reflectance and the whitecap reflectance, both of which are  
 296 driven by the ocean surface wind speed. It is important to remember that the sky glint reflection  
 297 was calculated through the VRT model of the Rayleigh signal. However, the direct glint signal  
 298 is calculated by the two-way attenuation of the direct solar beam that is modulated by the  
 299 surface glint reflectance,  $L_{GN}(\lambda)$ , which is modeled using Cox and Munk (1954) wave slope  
 300 statistics [66]. The TOA direct glint reflectance is then  $\pi L_{GN} T_{sol} T_{sen} / \mu_0$ , where  $\mu_0$  is the  
 301 cosine of the solar zenith angle, and spectral (and geometric for  $T$ ) dependency is implied. The  
 302 whitecap irradiance reflectance at the BOA,  $\rho_{wc}(\lambda)$ , is based on Koepke (1984) [67] combined  
 303 with the windspeed-dependent fractional coverage model of Stramksa and Petelski (2003) [68]  
 304 and the whitecap albedo spectral-dependence in the red and near-infrared from [69]. The BOA  
 305 irradiance reflectance is then propagated to TOA, similar to the ocean reflectance, as  $\rho_{wc} t_{sol}$   
 306  $t_{sen}$ , with spectral and geometric dependency implied.

307 We also account for the main absorbing gases in the atmosphere, including  $O_3$ ,  $H_2O$ , and  
 308  $O_2$ . The  $H_2O$  and  $O_2$  transmittance are based on the HITRAN 2016 line by line (LBL)  
 309 spectroscopic dataset [70]. Assuming the US standard atmospheric profile, we calculate the  
 310 LBL transmittance for different column water vapor (CWV) values. Then we apply the  
 311 instrument spectral response function (SRF) to the LBL transmittances and store them in a  
 312 LUT. Spectral  $H_2O$  transmittance at each  $Geom$ ,  $T_{wv}$ , is then interpolated from the LUT for a  
 313 given slant water vapor (WV) concentration along the path as  $cwv/\mu$ , where  $\mu$  is the cosine of  
 314 the path zenith angle. The  $O_2$  transmittance is calculated similarly for different path lengths of  
 315 the atmosphere given the observation geometry. The  $O_3$  transmittance is calculated from the  $O_3$   
 316 optical depth assuming the Beer-Lambert-Bouguer law, where the optical depth is determined  
 317 from the spectral  $O_3$  absorption coefficient [71] integrated with the sensor SRFs, and the  $O_3$   
 318 concentration.  $H_2O$  and  $O_3$  concentrations are taken from ancillary sources.

## 319 2.2 Neural network forward model

### 320 2.2.1 Data generation

321 In this work, the NN training dataset is derived from NASA's operational atmospheric LUTs.  
 322 Hence, the TOA reflectance can be represented as

$$323 \rho_t(\lambda) = \mathbf{F}(RH, O_3, Pr, WS, WV, fmf, \tau_a, a_{ph}, a_{dg}, b_{bp}, \gamma, Chl - a, \theta_0, \varphi, \theta_v), \quad (4)$$

324 where  $\mathbf{F}$  is the atmospheric LUT and ORM forward model operator,  $\lambda$  is sensor (here MODIS  
 325 Aqua) band center wavelengths within the solar spectrum,  $RH$  is the relative humidity in the  
 326 atmosphere,  $O_3$  is the column ozone concentration in Dobson units,  $Pr$  is the atmospheric  
 327 pressure in mbar,  $WS$  is the wind speed in m/s,  $WV$  is the column water vapor concentration in  
 328 cm,  $fmf$  is the aerosol volume fine-mode fraction,  $\tau_a$  is the AOD at 869 nm,  $a_{ph}$  is the  
 329 phytoplankton absorption coefficient at 443nm,  $a_{dg}$  is the colored dissolved and detrital matter  
 330 absorption coefficient at 443nm,  $b_{bp}$  is the particulate backscattering coefficient at 443nm, and  
 331  $\gamma$  is the slope of the backscattering coefficient.  
 332  
 333

Table 1. The range of all the parameters used in the NN training.

<i>Variable</i>	<i>Range</i>	<i>Distribution</i>	<i>Distribution Log<sub>10</sub> mean</i>	<i>Distribution Log<sub>10</sub> standard deviation</i>
$\lambda$ (nm)	412:869	-	-	-
<b>RH</b> (%)	30:95	Uniform	-	-
<b>O<sub>3</sub></b> (DU)	200:500	Uniform	-	-
<b>Pr</b> (mbar)	800:1100	Uniform	-	-
<b>WS</b> (m/s)	0.1:15	Uniform	-	-
<b>WV</b> (cm)	0.01:30	Lognormal	0.173	0.53
<b>f<sub>mf</sub></b> (unitless)	0:1	Uniform	-	-
$\tau_a$ (unitless)	0:0.4	Lognormal	-1.03	0.316
<b>a<sub>ph</sub></b> (m <sup>-1</sup> )	0.001:5	Lognormal	-1.5	0.45
<b>a<sub>dg</sub></b> (m <sup>-1</sup> )	0.001:5	Lognormal	-1.2	0.63
<b>b<sub>bp</sub></b> (m <sup>-1</sup> )	0.0001:0.1	Lognormal	-2.35	0.44
<b>Chl-a</b> (mg m <sup>-3</sup> )	0.05:50	Lognormal	-0.217	0.724
$\gamma$ (nm <sup>-1</sup> )	0:2	Uniform	-	-
$\theta_0$ (°)	5:77	MODISA geometry	-	-
$\varphi$ (°)	0:180	MODISA geometry	-	-
$\theta_v$ (°)	0:65	MODISA geometry	-	-

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The input parameters were generated for random uniform distribution with data ranges given in Table 1, with a few exceptions. Aerosol optical depth,  $\tau_a$  which was modeled with a log-normal distribution such that low optical depth cases have greater representation than higher optical depths [29,72,73]. Similarly, we assumed a log-normal distribution for the column water vapor,  $WV$ ,  $a_{ph}$ ,  $a_{dg}$ , and  $b_{bp}$  with the distribution mean and standard deviation reported in Table 1. The ocean IOPs are based on monthly mean (Level-3, L3) climatology products from MODIS Aqua, as distributed by the OB.DAAC, but the range was extended to include more extreme cases as observed in Level-2 (L2) data. The geometric parameters  $\theta_0$ ,  $\varphi$ , and  $\theta_v$  were all sampled from two MODIS Aqua orbits for a day in the summer and winter seasons, thus covering the entire solar geometry range of the sensor’s imaging duty cycle. As expected, the NN training is highly sensitive to the choice of geometries since radiant path geometry is a primary driver for signal variations at TOA. Sampling from observed orbit geometries ensures that the NN training considers only realistic solar and viewing geometry combinations, thus improving performance. However, we did not include covariance between the other parameters, which are all assumed independent.

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### 2.2.2 Training process

We generated spectral TOA reflectance,  $\rho_t$ , from the standard algorithm LUT for 16 million different data points. After excluding data points with the normalized sun glint radiance  $> 0.005$ , similar to the operational algorithm, we ended with  $\sim 9$  million data points for the training. The training was performed using the open-source machine learning platform Keras-TensorFlow (Keras.io). The NN input layer vector has 15 parameters (Table 1), and the output layer is the TOA reflectance,  $\rho_t$ , at 13 MODIS wavelengths from 412 to 869 nm. We found by trial and error that four hidden layers provide a good performance of the NN, with additional layers just adding forward model computational cost in the retrieval with negligible performance improvement. The Rectified Linear Unit (ReLU) activation function was used for the NN hidden layers [74]. We trained the NN with the Adam optimization algorithm for 10,000 epochs and with a batch size of 1,000 [75]. The dataset was split into the training set (85%) and a test set (15%). The mean squared error cost function between the training dataset (i.e.,  $\rho_t$ ) and the predicted values was minimized through the optimization process of the NN weights. We compared the NN performance on the training and test (independent) sets for all



365 the training epochs, showing a continuous decline in the cost function for both training and  
366 testing, indicating that the NN did not overfit the training data.

### 367 **3. Uncertainty sources**

368 It is important to properly account for an instrument's measurement uncertainty when  
369 validating the uncertainty estimates of the inferred variables. The measurement uncertainty  
370 includes both random and systematic components. The random component (noise) is calculated  
371 using the instrument's signal-to-noise ratio (SNR). In this work, we assume the sensor's  
372 uncorrelated random noise effects. We calculate the noise-equivalent radiance as follows:

$$NE\Delta L(\lambda) = [C_0(\lambda) + C_1(\lambda) \times L_t(\lambda)] \times S(\lambda), \quad (5)$$

373

374 where,  $C_0(\lambda)$  and  $C_1(\lambda)$  are linear fit coefficients of the noise model from [76], and  $S(\lambda)$   
375 is the spectrally-dependent spatial weight that brings all bands to a common 1 km spatial  
376 resolution [18]. The standard deviation of the signal is radiance-dependent and calculated as:

$$\sigma_n(\lambda) = \frac{NE\Delta L(\lambda)}{L_t(\lambda)}. \quad (6)$$

377

378 The standard deviation of the radiance is then converted to noise-equivalent reflectance after  
379 normalizing by the solar irradiance at a specific solar angle.

380 Systematic (calibration) errors in measurements are challenging to characterize post-launch  
381 due to the lack of an accurate absolute calibration apparatus on-orbit. Typically, the systematic  
382 uncertainty is correlated between bands. The Marine Optical BuoY (MOBY) site, off the coast  
383 of Lanai, Hawaii, is the system vicarious calibration site for all NASA-supported ocean color  
384 missions. NOAA has continuously operated MOBY since 1996 as the in-situ calibration source  
385 for vicarious calibration and a source of high-quality  $R_{rs}$  data [77,78]. There are 523 co-located,  
386 coincident MODIS Aqua-MOBY matchups, to date, of which a smaller fraction are used for  
387 the system vicarious calibration to derive the gain corrections at the TOA. Our approach relies  
388 on estimating the total uncertainty between the observed and predicted TOA reflectance in this  
389 work. Similar to [78], we calculate the predicted TOA reflectance by propagating the in-situ  
390 MOBY  $R_{rs}$  to the TOA while simultaneously solving for the aerosol properties.

391 The residual uncertainty between the observed and predicted TOA reflectance represents  
392 the total uncertainty at TOA defined below:

$$S_t = S_n + S_a + S_w + S_b, \quad (7)$$

393

394 where these terms represent error covariance matrices, with subscripts  $t$  for the total  
395 uncertainty,  $n$  for random noise,  $a$  is for ancillary data uncertainty,  $w$  is for the in-water  
396 component from MOBY, and  $b$  for the uncertainty due to instrumental systematic artifacts as  
397 well as the forward model uncertainty (e.g., RT simplifications). It is valid to sum these terms  
398 assuming that each is independent. The terms  $S_n$  and  $S_a$  are known given the SNR model and  
399 the ancillary data uncertainty. The uncertainty in MOBY  $R_{rs}$  observations is not well known for  
400 all conditions, but is expected to be a few percent [79]; thus, we assume it is negligible in this  
401 work as a first approximation. The term  $S_b$  can then be estimated and taken as a measure of  
402 systematic and forward model uncertainty used in the retrieval process.

### 403 **4. Optimal Estimation (OE)**

404 Optimal Estimation finds the most probable values of the unknown parameters in Table 1 by  
405 minimizing a cost function that incorporates the likelihood function, priors, and uncertainties.  
406 The likelihood and priors are assumed to be normal distributions, characterized by a one sigma  
407 width and correlations for all measurement pairs. The cost function near the solution is typically  
408 the weighted sum of squared differences between the forward model and the measurements,  
409 plus a similar weighted squared difference between the state and prior knowledge of the state.  
410 For non-linear problems such as the radiative transfer in the AO system, an iterative constrained  
411 optimization is used to minimize the cost function. Also, for simplicity, a conjugate Gaussian

412 distribution of the error covariance matrices is assumed, and, therefore, the computationally  
 413 intensive sampling of the distributions is unnecessary. Note strictly that these are “uncertainty”  
 414 rather than “error” covariance matrices, as in this case, the true value is not known (uncertainty  
 415 is a measure of dispersion, and error is a departure from the truth) [12]. However, we use the  
 416 common “error covariance” terminology for convenience (JCGM, 2008). OE involves  
 417 determining the maximum a posteriori (MAP) solution, which is a single point estimate of the  
 418 approximately normal distribution at the mode of the posterior, obtained by minimizing the  
 419 negative log posterior (known as the cost function,  $\chi^2$ ):

$$-2 \log_e P(\mathbf{x} | \mathbf{y}_{\text{obs}}, \mathbf{x}_a) = [\mathbf{y} - \mathbf{F}(\mathbf{x}, \mathbf{b})]^T \mathbf{S}_e^{-1} [\mathbf{y} - \mathbf{F}(\mathbf{x}, \mathbf{b})] + [\mathbf{x} - \mathbf{x}_a]^T \mathbf{S}_a^{-1} [\mathbf{x} - \mathbf{x}_a]. \quad (8)$$

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421 For this study, the forward model,  $\mathbf{F}(\mathbf{x}, \mathbf{b})$ , is the forward radiative transfer calculated for a  
 422 given state vector  $\mathbf{x}$  and, while  $\mathbf{b}$  represents the parameters that are used as an input to the  
 423 forward model, but not part of the state vector.  $\mathbf{y}_{\text{obs}}$  is a vector that contains the spectral  
 424 observed TOA reflectance, while  $\mathbf{x}_a$  is the prior state vector (knowledge of the state vector  $\mathbf{x}$   
 425 before measurements).  $\mathbf{S}_e$  is the measurement error covariance matrix, and  $\mathbf{S}_a$  is the prior error  
 426 covariance matrix. The diagonal elements of these matrices are the variances, while the off-  
 427 diagonal elements represent the correlated standard uncertainties in the state variables. Both  
 428 matrices need to be positive semi-definite (i.e., non-negative). The forward model parameters  
 429 (state vector) are:

$$\mathbf{x} = [RH, O_3, Pr, WS, WV, fmf, \tau_a, a_{ph}, a_{dg}, b_{bp}]. \quad (9)$$

430

431 The state vector  $\mathbf{x}$  in Eq. (9) includes the ancillary data as retrievable parameters, which is  
 432 different from many other approaches that either assume they are known perfectly or that they  
 433 are known imperfectly with some uncertainty (in which case this uncertainty is typically  
 434 propagated to TOA and included in  $\mathbf{S}_e$ ). Suppose the uncertainty of the ancillary data is known  
 435 or assumed. In that case, it is logical to have them as part of the state vector  $\mathbf{x}$  since the ancillary  
 436 data do influence the observations. Meanwhile, the non-retrievable parameters  $\mathbf{b}$  include Chl-  
 437 a and  $\gamma$ .

438

439 The iterative process to find a solution to the state vector,  $\mathbf{x}$ , follows the modified Gauss-  
 440 Newton optimization method by Levenberg-Marquardt (LM) [80,81]. We used the Python  
 441 library SciPy which implements the least-squares algorithm. The LM algorithm is very efficient  
 442 and provides a high convergence rate. Once a solution is found, we can estimate the error  
 443 covariance matrix at the estimated parameters. This is calculated using error propagation  
 through the Jacobian matrix,  $\hat{\mathbf{K}}$ , expressed as:

$$\hat{\mathbf{S}} = \left( \hat{\mathbf{K}}^T \mathbf{S}_e^{-1} \hat{\mathbf{K}} + \mathbf{S}_a^{-1} \right)^{-1}, \quad (10)$$

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where  $\hat{\mathbf{K}}$  is the partial first derivative of the forward function with respect to the state vector  
 (i.e.,  $\partial \mathbf{F} / \partial \mathbf{x}$ ).

This  $\hat{\mathbf{S}}$  term is the retrieval uncertainty of the state vector parameters and combines  
 uncertainty introduced by the measurements with the a priori constraints ( $\mathbf{S}_a$ , see later).

The OE technique described here is based on the normal distribution approximation of prior,  
 likelihood, and measurement uncertainty that may cause problems [12]. The LM algorithm may  
 converge to a local rather than the global minimum when the posterior is multi-modal. This can  
 occur for high-dimensional retrievals that are not properly constrained. The retrieval  
 uncertainty may also be over- or under-estimated if the forward model is highly nonlinear near  
 the solution, which is typically only an issue for poorly constrained parameters.

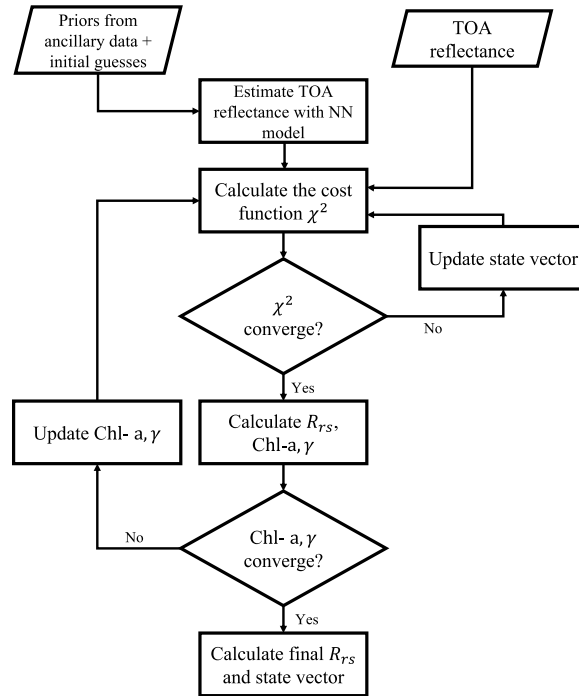


Fig. 1. Flow diagram of the OE algorithm

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Figure 1 shows the OE algorithm flow diagram. The required inputs include the TOA reflectance observed from MODIS, the ancillary data, and a priori information about the atmospheric and oceanic state. The prior distribution describes our current knowledge of the parameters of interest, and its mean is used as the first guess in the iterative inversion. Typically there are three types of priors: non-informative such as unbounded uniform distribution; weakly informative, such as bounded uniform or normal distribution with large variance; and informative such as normal with small variance. When non-informative priors are used, the prior does not affect the posterior, and the inference is then identical to the estimate of the likelihood. We used a normal prior distribution with no correlation between parameters in our analysis.  $RH, O_3, Pr, WS$ , and  $WV$  are obtained from ancillary data sources (National Centers for Environmental Prediction; NCEP). Assuming the mean is known, the uncertainty (standard deviation of the normal distribution) is assumed to be 1 mbar for  $Pr$ , 1 m/s for  $WS$ , 5% of the mean for  $RH$ , 1% of the mean for  $O_3$ , and 10% of the mean for  $WV$  [82,83]. The  $f_{mf}, \tau_a, a_{ph}, a_{dg}$ , and  $b_{bp}$  priors are assumed weakly informative normal with mean values obtained from the 4-km MODIS Aqua climatology obtained from the OB.DAAC and with a large standard deviation of 10. The standard deviation is much larger than the range of data, but the priors are bounded within their physical values in the inversion. The values for  $\gamma$  and Chl-a were used as the first guess and are obtained from climatology data as well. Given the latitude and longitude of each observation, we interpolate to the nearest neighbor of the global L3 image.

With the initial values of the input parameters, the TOA reflectance is calculated by evaluating the NN forward likelihood model, and a  $\chi^2$  value is calculated (Eq. 8). The algorithm iteratively updates the state vector until it converges. In the next step, we calculate the  $R_{rs}$  by performing the AC outlined in the following section. Neither Chl-a nor the backscattering slope is part of the state vector. Including them creates a highly ill-posed problem. The spectral backscattering requires simultaneously solving for its shape and magnitude. To avoid this problem, but provide a calculation of both, we utilize empirical relationships to estimate  $\gamma$  and Chl-a from  $R_{rs}$ . We use the OCx algorithm for Chl-a [84], and we use the Quasi Analytical

485 Algorithm (QAA) for  $\gamma$  [85]. The Chl-a and  $\gamma$  are iteratively adjusted and the OE AC correction  
 486 is repeated until they converge (i.e., they change by >2%), with a maximum of 10 iterations  
 487 (typically 2-3 are needed). It is important to note that we assume the uncertainty from Chl-a  
 488 and  $\gamma$  do not propagate into the  $R_{rs}$  uncertainty since Chl-a would only impact the uncertainty  
 489 in the BRDF correction, while  $\gamma$  does not play a role in the  $R_{rs}$  uncertainty estimate other than  
 490 constraining the AC.

#### 491 4.1 Remote sensing reflectance ( $R_{rs}$ ) retrieval

492 Our approach to the AC is a two-step one. First, the OE algorithm estimates the atmosphere-  
 493 related parameters in the previous section. Second, the inferred parameters are ingested into a  
 494 proper atmospheric correction similar to the operational algorithm. That involves removing the  
 495 atmospheric and glint signal from TOA observations and compensating for the atmospheric  
 496 diffuse and direct transmittance once these properties are inferred. We start by relating  $R_{rs}$   
 497 to the normalized water-leaving radiance (for simpler notation,  $\lambda$  is not included in the following  
 498 equations),

$$499 \quad R_{rs} = \frac{L_{wn}}{F_0}, \quad (sr^{-1}), \quad (11)$$

500 where  $L_{wn}$  is the normalized water-leaving radiance after the BRDF correction and  $F_0$  is the  
 501 extraterrestrial solar irradiance at 1 astronomical unit.  $L_{wn}$  is connected to TOA observations  
 502 by

$$503 \quad L_{wn} = \frac{f_{brdf} tL_w}{t_{sen} t_{sol} \mu_0 f_{sol}}, \quad (12)$$

504 where  $tL_w$  is the water-leaving radiance measured at TOA,  $t_{sen}$  and  $t_{sol}$  represent the diffuse  
 505 transmittance along the viewing and solar direction, respectively,  $\mu_0$  is the cosine of the solar  
 506 zenith angle,  $f_{sol}$  is the earth-sun distance correction factor, and  $f_{brdf}$  is the BRDF correction  
 507 factor:

$$508 \quad tL_w = \frac{F_0 \mu_0}{\pi} \times \left[ \frac{\rho_t}{T_{gsol} T_{gsen}} - \rho_{path+surf} \right]. \quad (13)$$

509  $\rho_t$  is the observed TOA reflectance.  $T_{gsol}$  and  $T_{gsen}$  represent the gas transmittance (ozone  
 510 and water vapor in this case) along the solar and viewing directions, respectively,  $\rho_{path+surf}$   
 511 is the TOA reflectance with a black ocean that includes only the reflectance from Rayleigh,  
 512 aerosols, glint, and white caps reflectance. The dark ocean TOA reflectance is calculated using  
 513 LUTs such that:

$$514 \quad \rho_{path+surf} = \mathbf{F}_a(Pr, WS, RH, fmf, \tau_a, \theta_0, \varphi, \theta_0). \quad (14)$$

515 The diffuse transmittance of the atmosphere needs to be calculated and is simply estimated  
 516 from the LUTs:

$$517 \quad t_{sol} = \mathbf{F}_{tsol}(Pr, RH, fmf, \tau_a, \theta_0), \quad (15)$$

518 and

$$519 \quad t_{sen} = \mathbf{F}_{tsen}(Pr, RH, fmf, \tau_a, \theta_v). \quad (16)$$

520 The above equations can therefore be used to solve for  $R_{rs}$

$$521 \quad R_{rs} = \frac{f_{brdf}}{\pi t_{sen} t_{sol}} \times \left[ \frac{\rho_t}{T_{gsol} T_{gsen}} - \rho_{path+surf} \right]. \quad (17)$$

522 To estimate the uncertainty in the  $R_{rs}$  estimate, we can easily propagate the uncertainties  
 523 from the inferred parameters through the above equations step-by-step. In the OE method, we

524 can calculate the Jacobian matrices of  $\mathbf{F}_a$ ,  $\mathbf{F}_{tsol}$ , and  $\mathbf{F}_{tse}$  which are denoted as  $\mathbf{K}_a$ ,  $\mathbf{K}_{tsol}$ , and  
 525  $\mathbf{K}_{tse}$ , respectively. We can then simplify the estimate of the  $R_{rs}$  as follows:

$$R_{rs}(\lambda) = \mathbf{F}_{AC}(\rho_t(\lambda), RH, O3, Pr, WS, WV, F_{MF}, \tau_a, \theta_0, \varphi, \theta_v), \quad (18)$$

526 where  $\mathbf{F}_{AC}$  is the Atmospheric Correction function. Using the chain rule, we can efficiently  
 527 calculate its Jacobian matrix,  $\mathbf{K}_{AC}$ , to estimate the error covariance matrix of the remote sensing  
 528 reflectance,  $\mathbf{S}_{Rrs}$ , as follows:  
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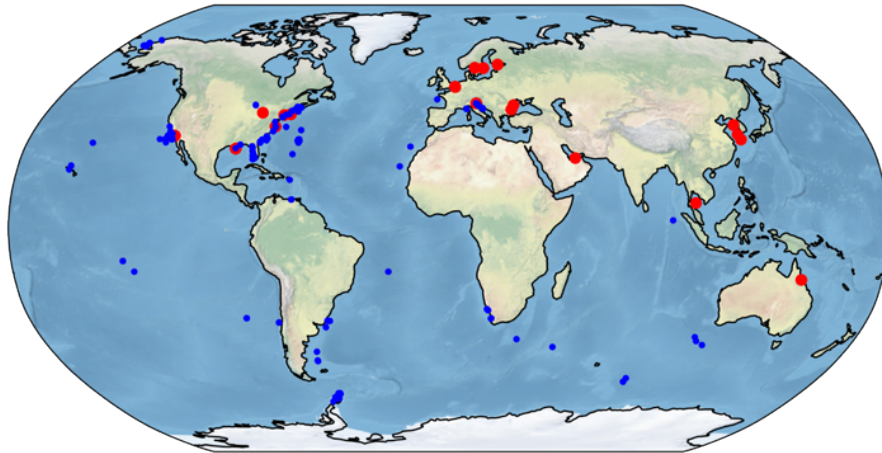
$$\mathbf{S}_{Rrs} = \mathbf{K}_{AC}^T \hat{\mathbf{S}} \mathbf{K}_{AC} + \mathbf{K}_{TOA}^T \mathbf{S}_e \mathbf{K}_{TOA}. \quad (19)$$

530  
 531 The second term in Eq. (19) accounts for propagating sensor noise to  $R_{rs}$  directly where  
 532  $\mathbf{K}_{TOA}$  is the Jacobian of  $R_{rs}$  with respect to the TOA reflectance. This method is a two-step  
 533 approach, where both terms on the right-hand side of Eq. (19) are assumed to be independent.

## 534 5. Validation data

### 535 5.1 In-situ radiometry

536 The in-situ  $R_{rs}$  data were obtained from the NASA SeaBASS database ([seabass.gsfc.nasa.gov](http://seabass.gsfc.nasa.gov))  
 537 includes above and in-water radiometry as well as retrievals from AERONET-OC (Version 2.0,  
 538 Level 2.0) sites ([aeronet.gsfc.nasa.gov](http://aeronet.gsfc.nasa.gov)) [86,87]. The AERONET-OC sites shown in Figure 2,  
 539 marked in red circles, are primarily located in coastal water near land. We used Version 2.0 for  
 540 consistency with the latest validation statistics used in the operational algorithm of the SeaWiFS  
 541 Data Analysis System (SeaDAS) and applied Level 2.0 quality filtering to ensure the highest  
 542 quality data. A complete list of the locations and characteristics of the AERONET-OC sites are  
 543 found on the AERONET-OC webpage and in [86]. The SeaBASS data points are marked in  
 544 blue circles shown in Figure 2, including samples in open ocean conditions. Accordingly, the  
 545 data exhibits a large dynamic range of  $R_{rs}$ . Full details on the  $R_{rs}$  dynamic range for all datasets  
 546 are available on the SeaBASS web page.



547  
 548 Fig. 2. Map of the SeaBASS (blue circles) and AERONET-OC (red circles) sites used in the validation.

### 549 5.2 MODIS Aqua

550 TOA reflectance data from MODIS onboard the Aqua satellite (MODIS-A) were used in this  
 551 study to validate  $R_{rs}$  matchups. MODIS-A level-1A (L1A) data were obtained from NASA's  
 552 OB.DAAC and processed to level-1B (L1B) after georeferencing. Satellite match-ups  
 553 coincident with the in situ validation dataset were identified following [88]. Satellite  
 554 measurements are derived from a box of pixels (i.e., 5 km  $\times$  5 km) centered on the location of  
 555 the in situ measurement. The satellite value is defined as the filtered mean of unflagged pixels  
 556 in the box, and the spatial homogeneity and other quality criteria at the validation point are  
 557 evaluated. Since in-situ data are rarely collected at the precise moment when a satellite views

558 its location, we allow a time window threshold of  $\pm 3$ -hours around the ground truth  
 559 observations. The length of that window is a compromise between being short enough to  
 560 minimize differences due to temporal variability in the ocean and being long enough to create  
 561 a sufficient volume of successful match-ups with satellite observations. The L1B file was then  
 562 processed to L2 using the SeaDAS standard algorithms to obtain geophysical products as well  
 563 as the TOA reflectance after applying the Ocean Biology Processing Group (OBPG)  
 564 calibrations of reprocessing R.2018 (e.g., polarization correction and vicarious  
 565 calibration) [78]. The standard L2 products were stored and used for the validation  
 566 comparisons. Since the vicarious calibration is an AC-specific procedure, we removed the  
 567 vicarious gains from the TOA reflectance by dividing the standard algorithm gains for the OE  
 568 L2 processing. We then use the modified TOA reflectance in the OE algorithm, as shown in  
 569 Figure 1.

## 570 5.2 Statistical metrics

571 When comparing satellite-derived  $R_{rs}$  with the in-situ value, we use several metrics, primarily  
 572 mean bias,  $\delta$ , and the mean absolute error (MAE or  $|\delta|$ ) both of which are routinely used to  
 573 assess model skill in SeaBASS [89]. We also calculated the root mean squared errors (RMSE)  
 574 ( $\Delta$ ) and the Pearson and Spearman squared,  $R^2$  correlation as well as the centered (bias-  
 575 corrected) MAE  $|\delta|_c$  and RMSE  $\Delta_c$  and the mean absolute relative error,  $|\psi|_m$ . We adopt the  
 576 IOCCG report 18 [28] notation assuming the satellite observations are denoted  $x_{i=1,N}$  and in-  
 577 situ denoted  $y_{i=1,N}$  and the following metrics are:

$$\delta = \frac{1}{N} \times \sum_{i=1}^N y_i - x_i, \quad (20)$$

$$|\delta| = \frac{1}{N} \times \sum_{i=1}^N |y_i - x_i|, \quad (21)$$

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$$\Delta = \sqrt{\frac{1}{N} \times \sum_{i=1}^N (y_i - x_i)^2}, \quad (22)$$

579

$$|\psi|_m = 100 \times \frac{1}{N} \times \sum_{i=1}^N \frac{|y_i - x_i|}{x_i}. \quad (23)$$

580

581 The centered statistics  $|\delta|_c$  and  $\Delta_c$  simply involve removing the average bias between  $y_i$   
 582 and  $x_i$ , thus showing the algorithm performance without any potential bias either in the  
 583 algorithm or the in-situ data.

## 584 5.2 Uncertainty validation

585 Our assumption to account for all sources of uncertainties at TOA relies on the MOBY  
 586 vicarious calibration to be representative of the global oceans. To validate this assumption, we  
 587 provide a closure analysis by comparing the satellite-derived  $R_{rs}$  and their associated  
 588 uncertainties to the in-situ measurements. Since we derive the pixel-level uncertainty, we can  
 589 use a statistical ensemble method to compare the derived uncertainty to the error between the  
 590 satellite-derived  $R_{rs}$  and the in-situ  $R_{rs}$ . The uncertainty estimated by OE is a normal distribution  
 591 with a standard deviation obtained through the analytical error propagation technique.  
 592 Meanwhile, the error defined as the difference between the retrieved and in-situ truth  $R_{rs}$  is an  
 593 instantaneous realization of that uncertainty distribution. Thus, a direct pixel-level comparison  
 594 between pixel-level uncertainty and retrieval errors is irrelevant. A more appropriate approach  
 595 to compare the two quantities is to calculate the normalized error distribution  $\Delta_N$  following the  
 596 approach of [90,91], where

$$\Delta_N = \frac{\Delta_s}{\sqrt{u_{sat}^2 + u_{ref}^2}}. \quad (24)$$

597  $\Delta_s$  is the error (i.e., difference) between the satellite-derived and in-situ  $R_{rs}$ .  $u_{sat}^2$  is the  
598 variance in the pixel-level uncertainty derived from the OE algorithm, while  $u_{ref}^2$  is the variance  
599 in the in-situ measurements. In an ideal scenario, where all sources of uncertainties are  
600 accounted for in the satellite and in-situ data, with perfect error propagation and with  
601 uncertainties following a normal distribution, the normalized error distribution should follow a  
602 normal distribution with a zero mean (i.e., no bias) and with 1 variance (or standard deviation).  
603 It is possible to examine the normality of the normalized error by plotting the cumulative  
604 distribution function (CDF). This provides an assessment of the average comparison between  
605 the total error and the OE-provided uncertainties. It is also possible to extend this analysis to  
606 assess the variability of the error to uncertainty relationship across the dynamic range of errors  
607 which would require stratifying the errors and comparing the 68<sup>th</sup> percentile of the error to the  
608 mean of the uncertainty within a bin (i.e., dividing the data by the expected error into equally  
609 populated bins) [91]. The choice of the number of bins depends on the available data volume  
610 in order to have a representative sample within each. In our analysis, we choose not to bin the  
611 data and provide a comparison between the 68<sup>th</sup> percentile of the error  $\Delta_s$ , and the mean of the  
612 uncertainty,  $\overline{u_{sat}}$ .

## 613 6. Results

### 614 6.1 Neural network performance

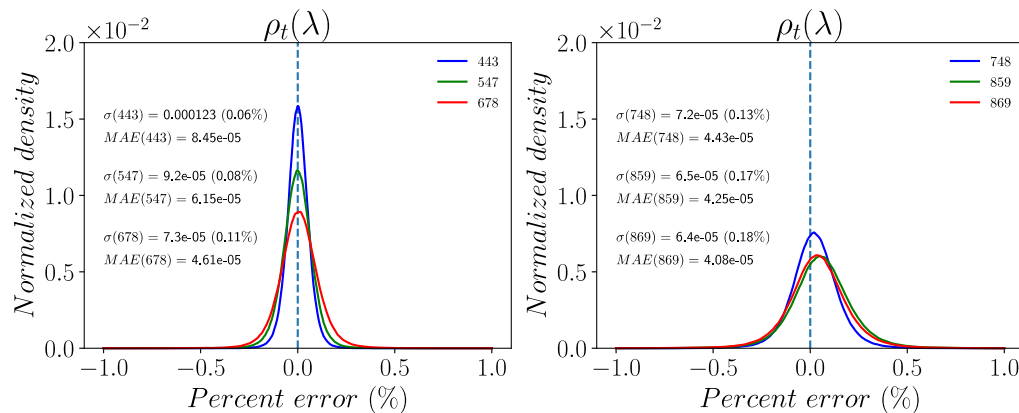
615 Our initial analysis of the NN prediction error on the testing dataset indicates that the error  
616 varies systematically with radiant path geometry. Figure 3 shows the percent prediction error  
617 histogram of the independent dataset (i.e., the 15% of the dataset reserved for testing) for three  
618 visible wavelengths (443, 547, and 678 nm) and three NIR wavelengths (748, 859, and 869  
619 nm).

620 The percent error is calculated as follows:

$$621 \quad \% \text{ error} = 100 \times (\rho_t^{NN}(\lambda) - \rho_t^{LUT}(\lambda)) / \rho_t^{LUT}(\lambda), \quad (25)$$

621

622 where  $\rho_t^{NN}(\lambda)$  is the TOA reflectance calculated by the forward NN model, and  $\rho_t^{LUT}(\lambda)$  is  
623 the TOA reflectance calculated from the atmosphere-ocean RT-based LUT model. We also  
624 calculate the mean absolute error,  $|\delta|$ , where  $y_i$  is retrieved data (i.e.,  $\rho_t^{NN}(\lambda)$ ),  $x_i$  is the truth  
625 (i.e.,  $\rho_t^{LUT}(\lambda)$ ), and N is the number of data points (approximately 1.3 million).



626

627 Fig. 3. Histograms of the percent error between the NN derived TOA reflectance and the LUT using an independent  
628 validation data set for 443, 547, and 678 nm (right panel), and 748, 859, and 869 nm (left panel). Errors are mostly  
629 smaller than 0.2% in reflectance.  $\sigma$  is the standard deviation of the absolute error, while the value in parenthesis is for  
630 the percent error. MAE is the mean absolute error.

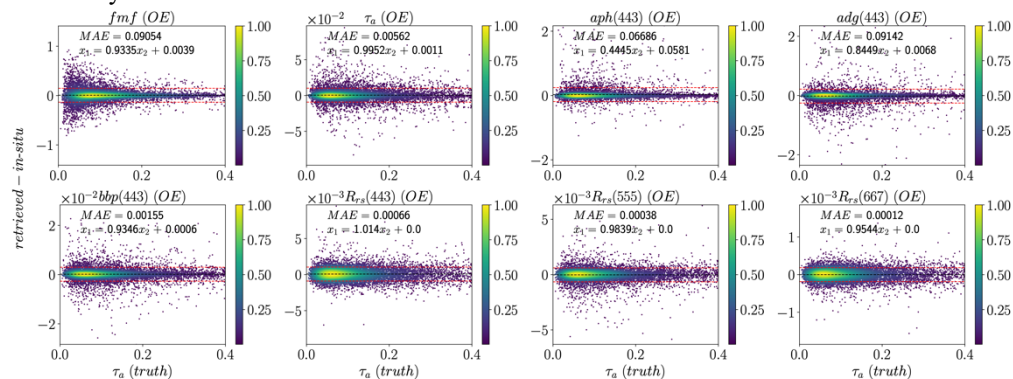
631

632 Figure 3 shows a larger percent error at longer wavelengths, with a slight bias at 859 and  
633 869 nm. Overall, the performance of the NN is excellent with an error < 0.2% for 82% of the

634 testing cases in the worst-case scenario and  $<0.06\%$  in the best case, similar to the instrument's  
 635 radiometric noise and within the bounds of the vicarious calibration uncertainty [78]. We  
 636 parametrized the NN model uncertainty,  $\sigma_{NN}$ , as a function of the geometry to account for the  
 637 forward model uncertainty needed in the inference process. The NN was trained with the AC  
 638 LUTs, which were calculated with a coarse grid that can cause interpolation errors. However,  
 639 in [18], we showed that the LUT interpolation error is the smallest fraction of the total  
 640 uncertainty. Therefore, this forward model uncertainty here is a fraction of the total forward  
 641 model uncertainty, which is unknown and likely systematic because of the simplification of the  
 642 physics (i.e., not accounting for absorbing aerosols and not including other unknown  
 643 unknowns) [12].

## 644 6.2 Synthetic data analysis

645 Out of the NN test dataset, we extracted 10,000 cases of TOA reflectances and the ‘truth’  
 646 geophysical parameters used in the OE algorithm. Before passing to the algorithm, we added  
 647 random and systematic radiometric uncertainty to the TOA reflectance derived in section 3 and  
 648 to the ancillary data input as a prior. The input data set spanned a wide range of environmental  
 649 conditions and geometries with statistical samples representing the NN training and testing  
 650 data. In Figure 4, rather than showing a scatter plot, we show the scatter density histogram plot  
 651 for each retrieved parameter of the OE algorithm. The color bar indicates the normalized  
 652 density of the data frequency. The plot shows the difference (error) between the retrieved data  
 653 and the truth. Thus a perfect retrieval would show a zero error on the y-axis. We choose the x-  
 654 axis that is relevant to the AC process. Since the AC and the TOA reflectance strongly depend  
 655 on the AOD, dependence between the AC parameters and the  $R_{rs}$  is expected. The black dashed  
 656 lines are the mean of the difference between the retrieved and the truth, while the red dashed  
 657 lines indicate the  $\pm$  standard deviation around the mean of the difference. A bias between the  
 658 retrieval and the truth would manifest in the black dashed line deviating away from zero. Above  
 659 zero, the retrieval is overestimated and vice versa for underestimated retrieval. A larger spread  
 660 between the retrieval and truth would lead to a more significant deviation of the red dashed  
 661 lines away from the mean black dashed lines.

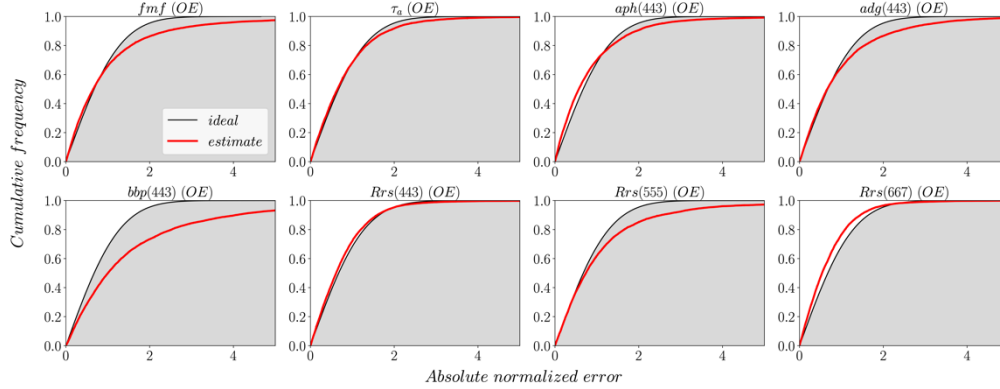


662 Fig. 4: Scatters density histogram of the synthetic data retrievals using the OE algorithm. The color bar indicates the  
 663 data normalized density ranging from 0 to 1. MAE is the mean absolute error between the retrieved and truth, while  
 664  $x_1 = ax_2 + b$  is the regression line between retrieved and truth with  $a$  being the slope and  $b$  is the bias.  
 665  
 666

667 Figure 4 shows the retrieval performance for two parameters related to the AC ( $fmf$ ,  $\tau_a$ ) and  
 668 three ocean-related parameters from the GIOP model  $a_{ph}$ ,  $a_{dg}$ ,  $b_{bp}$  all at 443 nm. The  $R_{rs}$  at  
 669 443, 555, and 665 nm were calculated after performing the AC by removing the atmospheric  
 670 signal contribution from the TOA. There is a negligible bias in the retrieval for all parameters,  
 671 particularly for  $R_{rs}$  with no dependence on the  $\tau_a$ . The  $fmf$  error shows a slight dependence  
 672 on  $\tau_a$  at low values, where the uncertainty is increased at low AOD. The MAE ( $|\delta|$ ) of  $R_{rs}$  is  
 673 0.00066, 0.00038, 0.00012 for 443, 555, and 665nm, respectively, showing that the absolute



674 magnitude of uncertainty is higher at shorter wavelengths consistent with what is observed  
 675 based on real data validation statistics [27].  
 676 To evaluate the retrieval uncertainty for each case, we calculate the CDF of the normalized  
 677 error distribution,  $\Delta_N$ , for each retrieval parameter. For a perfect retrieval and uncertainty  
 678 estimate, the calculated normalized error would agree with the ideal case across the normalized  
 679 error range. Figure 5 shows the CDF of  $\Delta_N$  in red compared to the ideal case of a standard  
 680 normal in black. When the red curve is within the grey shaded region, the uncertainty is  
 681 underestimated and overestimated when the curve is in the white region. Overall, there is a  
 682 good agreement for all parameters except  $b_{bp}$ , where the uncertainty is underestimated.



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 684 Fig. 5. CDF plot of the absolute normalized error,  $\Delta_N$ , for all retrieval parameters of the synthetic dataset. The  
 685 estimated CDF from the OE algorithm is shown in red, and the ideal CDF for a standard normal is shown in black.  
 686 The grey shaded region shows where the uncertainty is underestimated.  
 687

688 In Table 2, we compare the mean uncertainty estimate  $\overline{u_{sat}}$  of the retrieval as compared to  
 689 the 68<sup>th</sup> percentile of the retrieval error,  $\overline{\Delta_s}$ .  
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 692 Table 2.  $\overline{\Delta_s}$  is the 68<sup>th</sup> percentile of the error between the truth and the retrieval and  $\overline{u_{sat}}$  is the mean  
 693 uncertainty for each parameter.

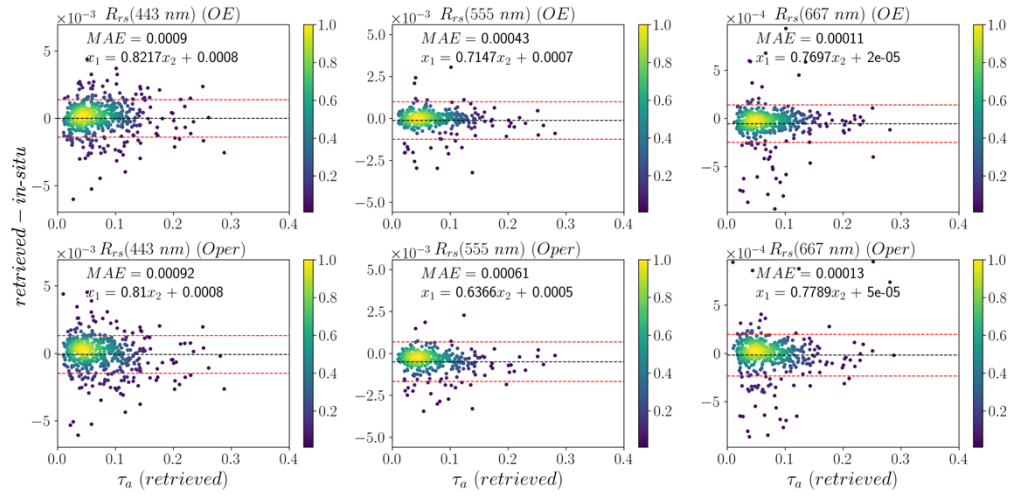
	$fmf$	$\tau_a$	$a_{ph}$	$a_{dg}$	$b_{bp}$	$R_{rs}(443)$	$R_{rs}(555)$	$R_{rs}(667)$
$\overline{\Delta_s}$	0.00865	0.0054	0.0373	0.0564	0.00140	0.00076	0.000449	0.00015
$\overline{u_{sat}}$	0.00633	0.0050	0.0372	0.0460	0.00063	0.00078	0.000363	0.00016

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 695 The results in Table 2 complement Figure 5, indicating a good agreement between the two  
 696 results, except for underestimation of  $b_{bp}$  uncertainty.

### 697 6.3 In-situ validation

#### 698 6.3.1 SeaBASS

699 The SeaBASS dataset provides an overall assessment of the OE algorithm in a wide range of  
 700 water conditions. Figure 6 shows the error between the  $R_{rs}$  retrieval and the in-situ truth for  
 701 three wavelengths at 443, 555, and 667nm. The first row is for the OE algorithm, while the  
 702 second row is for the operational retrieval using the SeaDAS/l2gen L2 processing software.  
 703 The matchup analysis shows a lower MAE for the three OE algorithm bands than the  
 704 operational one. There is no apparent correlation with  $\tau_a$ , as a primary source of AC errors in  
 705 all cases.  
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Fig. 6. Scatter density histograms of the SeaBASS data retrievals using the OE algorithm. The first row is for the OE algorithm, and the second row is for the operational algorithm. The color bar indicates the data normalized density ranging from 0 to 1. MAE is the mean absolute error between the retrieved and truth, while  $x_1 = ax_2 + b$  is the regression line between retrieved and truth with  $a$  being the slope and  $b$  is the bias.

For a quantitative assessment, we provide in Table 3 all metrics and validation statistics of the matchups for  $R_{rs}$ . In Table 3, the numbers highlighted in bold are for the OE algorithm, while the non-bolded numbers in parentheses are for NASA's current operational AC algorithm.

**Table 3. Matchups statistics for the SeaBASS dataset. OE statistics are in bold font-weight, while NASA's operational AC are in normal font and in parentheses. N is the number of negative  $R_{rs}$  retrievals.**

	N (N)	$\delta$	$ \psi _m$ (%)	$ \delta $	$ \delta _c$	$\Delta$	$\Delta_c$	$R^2$ (Pearson)	$R^2$ (Spearman)
$R_{rs}(443)$	589	$-1 \times 10^{-6}$	<b>22.0</b>	$9 \times 10^{-4}$	$9 \times 10^{-4}$	$1.39 \times 10^{-3}$	$1.39 \times 10^{-3}$	<b>0.69</b>	<b>0.73</b>
	(0, 1)	$(-8.1 \times 10^{-3})$	(23.7)	$(9.2 \times 10^{-4})$	$(9.2 \times 10^{-4})$	$(1.39 \times 10^{-3})$	$(1.40 \times 10^{-3})$	(0.69)	(0.71)
$R_{rs}(555)$	438	$-1.3 \times 10^{-4}$	<b>13.7</b>	$4.3 \times 10^{-4}$	$4.5 \times 10^{-4}$	$1.12 \times 10^{-3}$	$1.14 \times 10^{-3}$	<b>0.76</b>	<b>0.73</b>
	(0, 0)	$(-5 \times 10^{-4})$	(18.6)	$(6.1 \times 10^{-4})$	$(1 \times 10^{-3})$	$(1.28 \times 10^{-3})$	$(1.54 \times 10^{-3})$	(0.76)	(0.66)
$R_{rs}(667)$	490	$-5.5 \times 10^{-5}$	<b>43.4</b>	$1.1 \times 10^{-4}$	$1.4 \times 10^{-4}$	$2.02 \times 10^{-4}$	$2.23 \times 10^{-4}$	<b>0.8</b>	<b>0.3</b>
	(1, 14)	$(-2 \times 10^{-5})$	(63.1)	$(1.3 \times 10^{-4})$	$(1.3 \times 10^{-4})$	$(2.17 \times 10^{-4})$	$(2.19 \times 10^{-4})$	(0.75)	(0.23)

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The mean bias,  $\delta$ , between the in-situ and retrieved  $R_{rs}$  is smaller for the OE algorithm, while the  $|\psi|_m$  and  $|\delta|$  is reduced for all bands. The improvement at 443nm is marginal at 1.7%, however, the  $|\psi|_m$  is reduced by 4.9% and 19.7% for the OE algorithm for 555 and 667nm, respectively. We also calculate the centered statistics after removing the mean bias showing consistent results where the OE algorithm outperforms the operational algorithm. The Spearman correlation is improved for the OE algorithm. In contrast, the Pearson correlation shows no improvement except for 667 nm, possibly due to spurious outliers that the metric can be sensitive to.

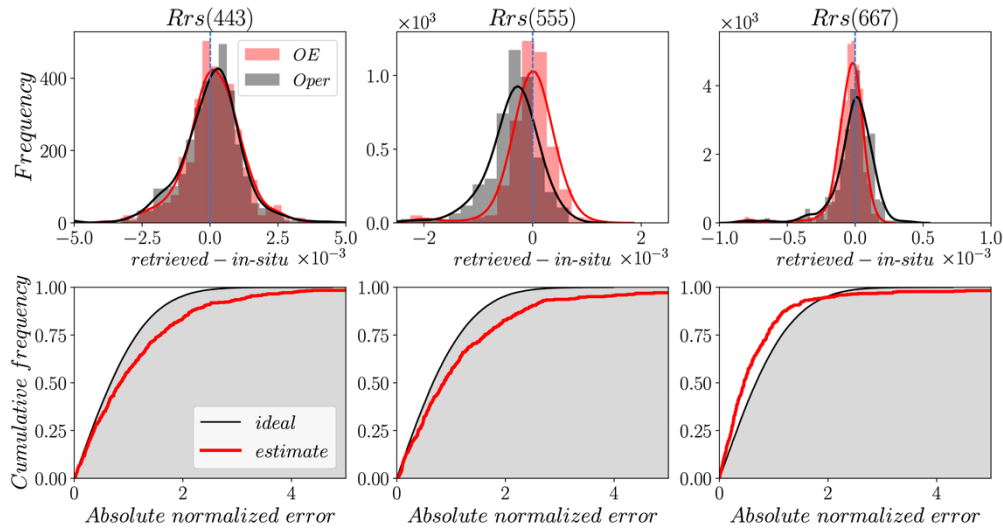


Fig. 7. Top row is a histogram of the difference between the retrieved and in-situ  $R_{rs}$  at 443, 555, and 667 nm, respectively for the OE algorithm in red, and the operational algorithm in black. The bottom row is the CDF of the absolute normalized error  $\Delta_N$  for  $R_{rs}$  at the same three bands, where the red curve is estimated from the OE algorithm, and the black curve is the ideal case for a standard normal.

Figure 7 shows the histogram of the error as the difference between  $R_{rs}$  in-situ and retrieved at 443, 555, and 667 nm, respectively in the top row. The error is mostly centered around zero, and the histogram follows a normal distribution indicating that a random process likely dominates the error. The mode of the  $R_{rs}$  histogram is closer to zero for the OE algorithm than for the operational algorithm, for the green and red bands, while the distributions look very similar for the blue band. The bottom row of Figure 7 is the CDF of the absolute normalized error  $\Delta_N$ . These results indicate that  $R_{rs}$  uncertainty is underestimated relative to the in-situ matchup errors for the blue and green bands with higher underestimation for larger errors. At 667 nm, the uncertainty was overestimated for lower errors and vice versa for higher errors. It is important to note that the matchup errors would implicitly include other sources of errors such as in-situ data uncertainties, adjacency effects, and temporal and spatial mismatch.

Table 4.  $\overline{\Delta_s}$  is the 68<sup>th</sup> percentile of the error between the truth and the retrieval and  $\overline{u_{sat}}$  is the mean uncertainty for SeaBASS  $R_{rs}$  matchups.

	$R_{rs}(443)$	$R_{rs}(555)$	$R_{rs}(667)$
$\overline{\Delta_s}$	0.00091	0.00030	0.00008
$\overline{u_{sat}}$	0.00071	0.00025	0.00012

We also compare the mean uncertainty estimate  $\overline{u_{sat}}$  of the retrieval as compared to the 68<sup>th</sup> percentile of the retrieval error,  $\overline{\Delta_s}$  for the SeaBASS matchups in Table 4. There is a good agreement between the two metrics, however, the uncertainty is underestimated slightly at 443 and 555 nm and overestimated for 667 nm.

### 6.3.2 AERONET-OC

We extended our matchup validation analysis to the AERONET-OC coastal water sites. The plot in Figure 8 shows the retrieval error for  $R_{rs}$  at 443, 555, 667 nm versus the retrieved  $\tau_a$ . The results for the OE algorithm show a smaller MAE but comparable spread to the operational algorithm. The plot shows little dependence of the error on the retrieved  $\tau_a$  indicating no aerosol-dependent bias in the AC.

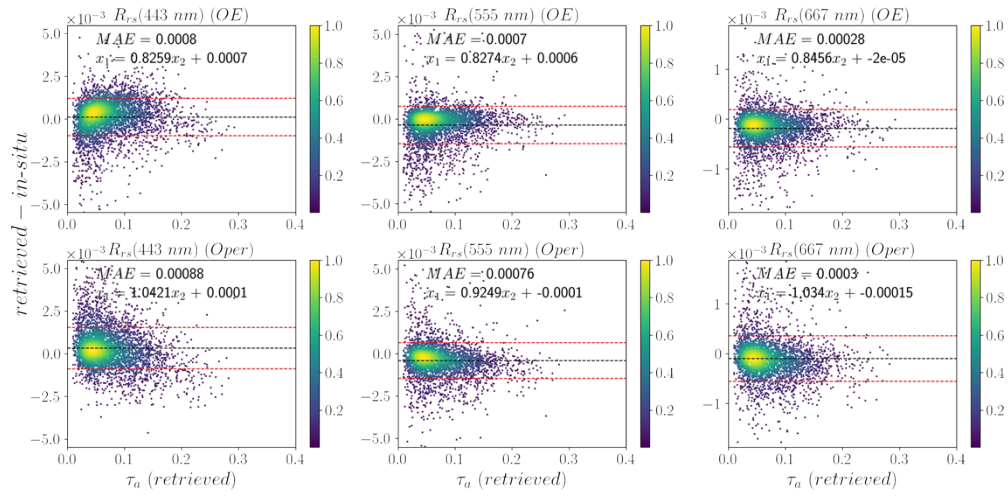


Fig. 8. Same as Fig. 6, but for the AERONET-OC dataset.

The detailed matchup statistics in Table 5 show better metrics for OE, with a smaller mean bias except for 667 nm and lower  $|\psi|_m$  and  $|\delta|$  across all bands.  $|\psi|_m$  was reduced by 1.5, 1.9, and 6.6% for 443, 555, and 667 nm, respectively. The centered metrics  $|\delta|_c$  and  $\Delta_c$  are consistently better for the OE algorithm, except for the red band, while the correlations are improved for all bands, except for the Pearson metric at 443 nm.

Table 5. Matchup statistics for the AERONET-OC dataset. OE statistics are in bold font-weight, while operational are in normal font. N is the number of negative  $R_{rs}$  retrievals.

	N (N <sup>-</sup> )	$\delta$	$ \psi _m$ (%)	$ \delta $	$ \delta _c$	$\Delta$	$\Delta_c$	$R^2$ (Pearson)	$R^2$ (Spearman)
<b><i>Rrs(443)</i></b>	4300 (16, 47)	<b><math>1.1 \times 10^{-4}</math></b> ( $3.2 \times 10^{-4}$ )	<b>37.7</b> (39.2)	<b><math>8 \times 10^{-4}</math></b> ( $8.8 \times 10^{-4}$ )	<b><math>8.3 \times 10^{-4}</math></b> ( $1 \times 10^{-3}$ )	<b><math>1.10 \times 10^{-3}</math></b> ( $1.25 \times 10^{-3}$ )	<b><math>1.12 \times 10^{-3}</math></b> ( $1.37 \times 10^{-3}$ )	<b>0.83</b> (0.84)	<b>0.82</b> (0.78)
<b><i>Rrs(555)</i></b>	3746 (0, 0)	<b><math>-3.5 \times 10^{-4}</math></b> ( $-4.1 \times 10^{-4}$ )	<b>12.0</b> (14.1)	<b><math>7 \times 10^{-4}</math></b> ( $7.6 \times 10^{-4}$ )	<b><math>8.6 \times 10^{-4}</math></b> ( $1 \times 10^{-3}$ )	<b><math>1.16 \times 10^{-3}</math></b> ( $1.14 \times 10^{-3}$ )	<b><math>1.31 \times 10^{-3}</math></b> ( $1.34 \times 10^{-3}$ )	<b>0.92</b> (0.91)	<b>0.92</b> (0.90)
<b><i>Rrs(667)</i></b>	3815 (13, 157)	<b><math>-1.8 \times 10^{-4}</math></b> ( $-1 \times 10^{-4}$ )	<b>30.8</b> (37.4)	<b><math>2.7 \times 10^{-4}</math></b> ( $3 \times 10^{-4}$ )	<b><math>4.1 \times 10^{-4}</math></b> ( $3.5 \times 10^{-4}$ )	<b><math>4.19 \times 10^{-4}</math></b> ( $4.68 \times 10^{-4}$ )	<b><math>5.31 \times 10^{-4}</math></b> ( $4.98 \times 10^{-4}$ )	<b>0.87</b> (0.85)	<b>0.82</b> (0.77)

Similar to the SeaBASS analysis, in Figure 9, we show the histogram of the matchup errors and the CDF of the absolute normalized error  $\Delta_N$  in the top row. The histogram shows a distribution similar to normal for both algorithms. The modes of the distributions are consistently close to zero, where at 555 nm, the OE algorithm is closer to zero than the operational algorithm. In the bottom row, the CDF comparison indicates a close agreement between the estimated uncertainty from the OE algorithm and the matchup errors for 443 nm. The 555 and 667 nm uncertainty are underestimated. This indicates that we did not account for all sources of errors in the algorithm.

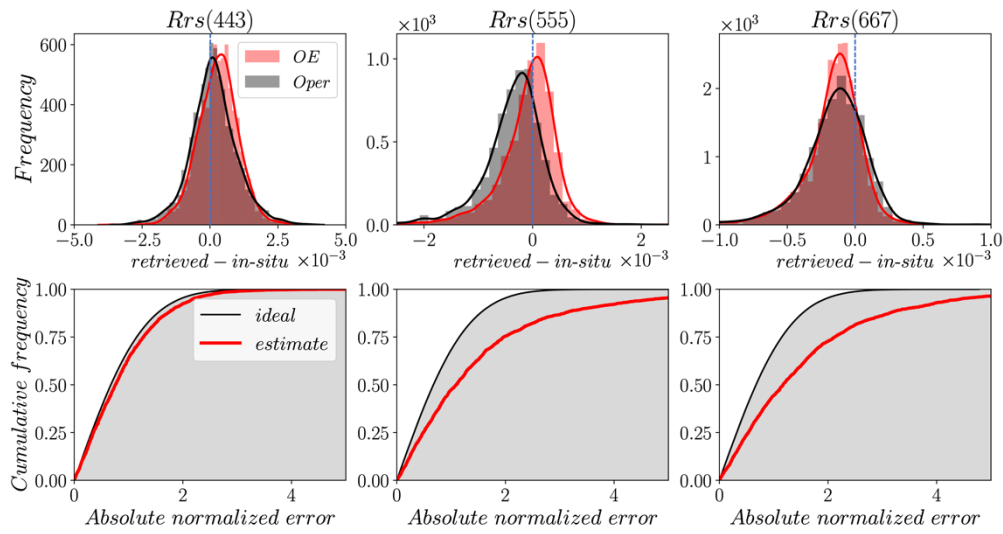


Fig. 9. Same as Fig. 7, but for the AERONET-OC matchups.

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 786 Stratification of data by location provides more insight into the performance of the OE  
 787 algorithm for different environmental conditions. Figure 10 presents results at a more granular  
 788 level by showing site-by-site CDFs of  $\Delta_N$  for the 9 sites with at least 250 matchups, plus the  
 789 remaining sites pooled together (labeled “Others”). The overall performance shows a good  
 790 agreement in the uncertainty estimate at 443 nm for all sites with slight underestimation. For  
 791 555 and 667 nm, the underestimation of the uncertainty is more significant, particularly for  
 792 MVCO (a highly productive region) and Palgrunden (an inland site). The best agreement was  
 793 for the Helsinki site, followed by Gustav, both characterized by their high CDOM  
 794 concentrations [86]. Although Venice provides the most significant volume of data, the  
 795 uncertainty was underestimated in the green and red bands.  
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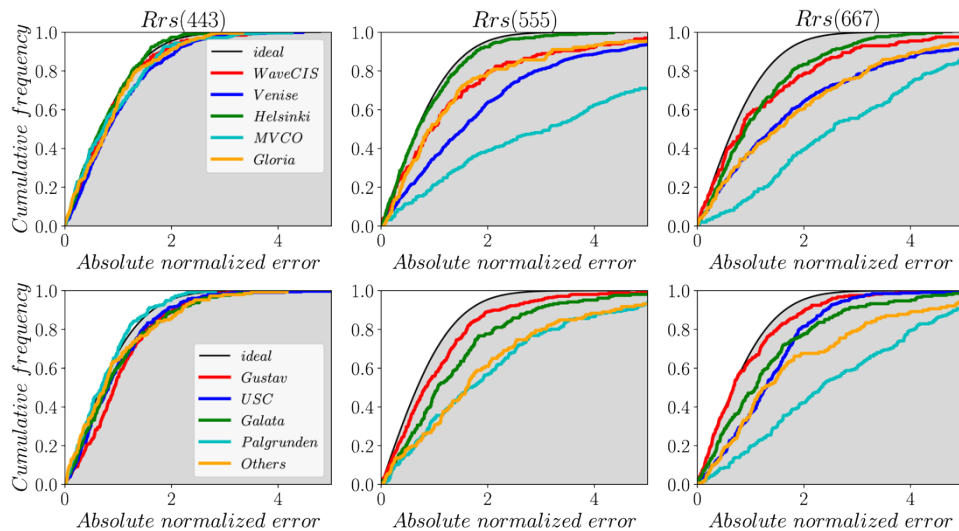


Fig. 10. Same as Fig. 9, but stratified for different AERONET-OC sites.

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800 It is important to note that  $\Delta_N$  shows a combined effect of retrieval bias and scatter; thus, a  
 801 highly biased retrieval that is not captured in the uncertainty estimate would lead to a significant  
 802 over- or underestimation of the normalized error.

803 We also provide (Table 6) the comparison between the 68<sup>th</sup> percentile of the error,  $\overline{\Delta_s}$ , and  
 804 the mean uncertainty from the retrieval,  $\overline{u_{sat}}$  for all sites and a breakdown of the 5 best-sampled  
 805 sites. Similar to Figure 10, there is a good agreement for all sites at 443 nm, however, the  
 806 uncertainty is underestimated for the green and red bands.  
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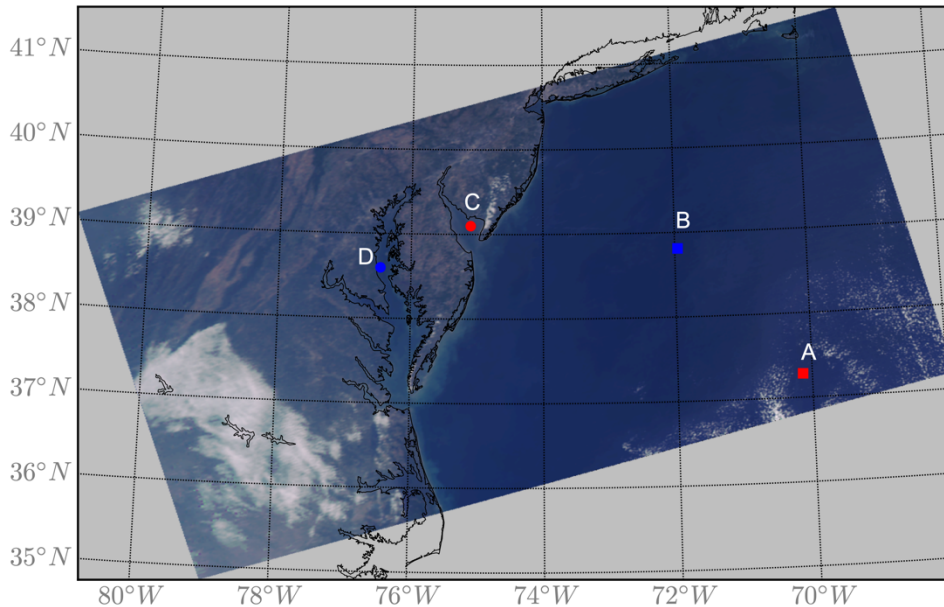
808 **Table 6.**  $\overline{\Delta_s}$  is the 68<sup>th</sup> percentile of the error between the truth and the retrieval and  $\overline{u_{sat}}$  is the mean  
 809 **uncertainty for AERONET-OC  $R_{rs}$  matchups.**

Site	$R_{rs}(443)$		$R_{rs}(555)$		$R_{rs}(667)$	
	$\overline{\Delta_s}$	$\overline{u_{sat}}$	$\overline{\Delta_s}$	$\overline{u_{sat}}$	$\overline{\Delta_s}$	$\overline{u_{sat}}$
All sites	0.000814	0.000724	0.000483	0.000257	0.000245	0.000137
Venise	0.000910	0.000720	0.000546	0.000241	0.000252	0.000125
Helsinki	0.000683	0.000721	0.000290	0.000272	0.000192	0.000150
MVCO	0.000846	0.000741	0.001150	0.000267	0.000425	0.000140
Gloria	0.000794	0.000726	0.000393	0.000253	0.000280	0.000134
Gustav	0.000854	0.000720	0.000297	0.000262	0.000154	0.000146

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811 **6.3.3 MODIS Aqua imagery analysis**

812 Figures 11 and 12 show results of the OE and NASA operational algorithms from a MODIS-Aqua  
 813 Aqua image over the eastern coast of the United States and extending into the Atlantic Ocean  
 814 on September 21<sup>st</sup>, 2010. The scene includes a wide range of water conditions, including coastal  
 815 waters such as the Chesapeake Bay region and open ocean low Chl-a regions further away from  
 816 the coast. Figure 11 shows a true-color composite, highlighting four pixels (labeled A-D),  
 817 representing different water conditions based on the Chl-a and the fitting residual  $\chi^2$  of the OE  
 818 algorithm.



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820 **Fig. 11.** True-color image composite of MODIS-Aqua from TOA reflectance over the eastern coast of the United  
 821 States from September, 21<sup>st</sup>, 2010.  
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823 Figure 12 shows the L2 image of  $R_{rs}$  at 443, 555, and 667 nm. Spatial patterns and  
 824 magnitudes are similar for both the OE (top row) and operational (middle) algorithms,  
 825 particularly in waters further away from the coast. However, there are differences in coastal

826 waters in regions where  $R_{rs}(555)$  are high, indicative of optically complex conditions with high  
827 particulate backscattering. The OE algorithm does not perform a cloud screening step similar  
828 to the operational approach. However, this caused artifacts in the OE  $R_{rs}$  retrieval around cloud  
829 edges, which can be mitigated by using additional cloud screening and masking approaches,  
830 such as limiting the retrieval with extremely high fitting error  $\chi^2$ .

831 In the third row of Figure 12, we show the pixel-level uncertainty produced by the OE  
832 algorithm for the three bands. On average, the magnitude of the uncertainty is higher for the  
833 blue bands than the green/red bands; it is mainly affected by the atmospheric correction of the  
834 aerosol optical depth and fine mode fraction, which are (in this scene) spatially smooth. The  
835 fourth row shows the  $R_{rs}$  relative uncertainty (%). These images show a more pronounced  
836 spatial structure; for example, in the optically-complex Chesapeake Bay waters, the water  
837 column's absorption coefficient is so significant that  $R_{rs}$  is small in the blue bands, thus the  
838 relative uncertainty is substantial. This is reversed in the bright blue waters further from the  
839 coast, where the uncertainty is smaller (5-10%). Similarly, in coastal waters, the  $R_{rs}$  in the  
840 green band is relatively large, so the uncertainty is smaller than in low Chl-a waters. This is  
841 also consistent for the red bands, where the low Chl-a conditions show very significant  
842 uncertainties (>50% and in some cases >100%), however, this is expected since the  $R_{rs}$  is near  
843 zero in the red band.



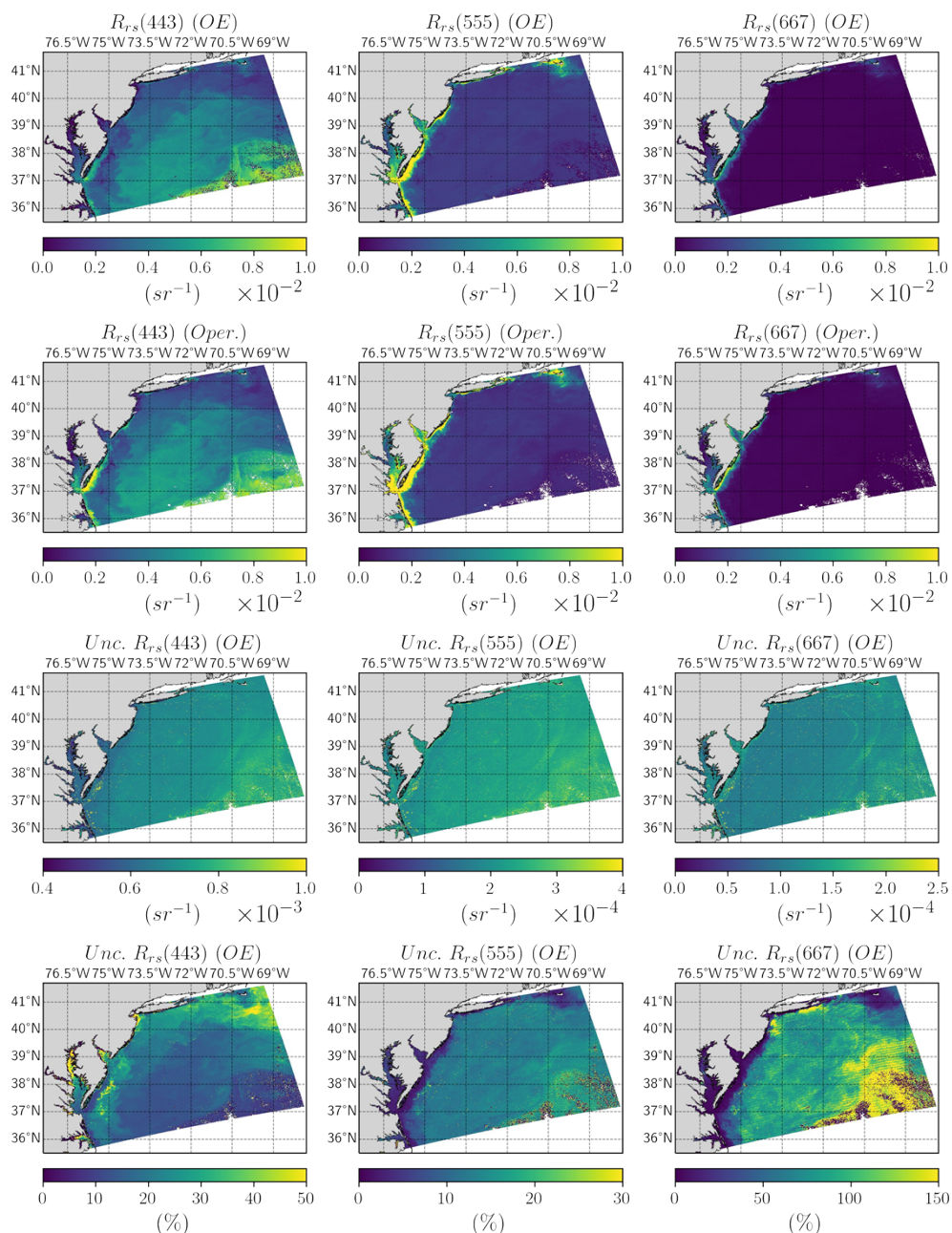


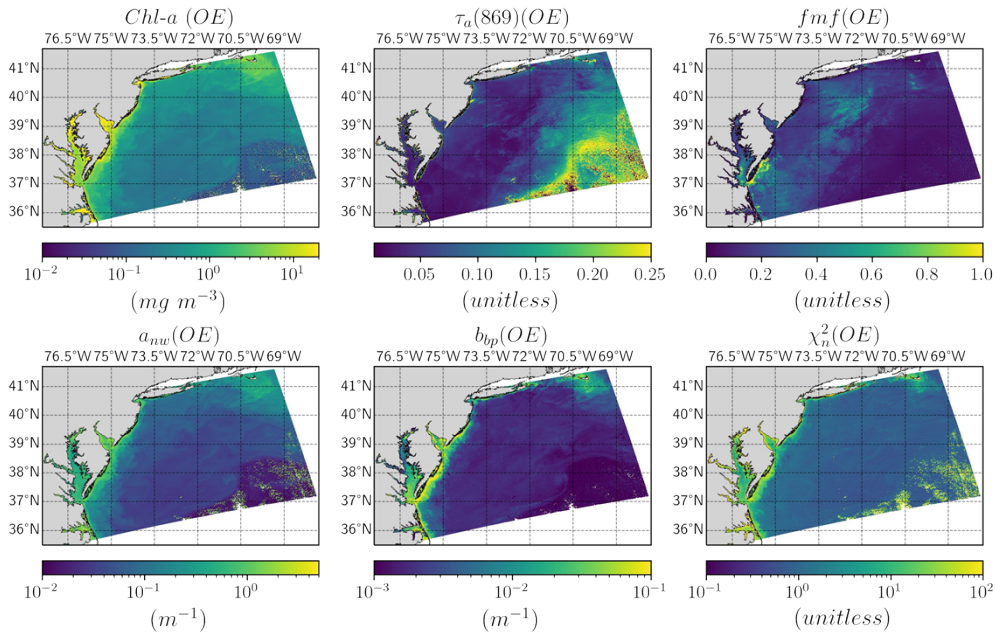
Fig. 12. MODIS-A image of  $R_{rs}$  retrieval at 443, 555, 667nm. The top row is the OE algorithm retrieval, and the middle is the operational algorithm. The third row is the absolute uncertainty estimated from the OE algorithm. The last row is for the relative percent uncertainty.

Figure 13 shows the Chl-a retrieval using the OC3 band ratio algorithm after performing the AC to retrieve  $R_{rs}$  [84]. The spatial distribution of Chl-a exhibits the typical spatial pattern in that region with high Chl-a values in the Chesapeake Bay (and its estuaries), Delaware Bay, Albemarle Sound, and low Chl-a values in offshore waters of the mid-Atlantic Bight.  $\tau_a$  and  $f_{mf}$  spatial distributions are smooth and do not show artifacts, particularly in very bright waters, where the non-negligible water-leaving radiance in the longer wavelengths can be

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855 erroneously attributed as an aerosol signal. However, there is a slight artifact near the mouth of  
 856 the Chesapeake Bay and adjacent to the southeast of the Delmarva Peninsula between 37° and  
 857 38° N in a region where  $R_{rs}$  values are relatively high. It is not clear if these are finer aerosol  
 858 values or retrieval errors due to the high water reflectance signal; however, some of these  
 859 artifacts are reflected as a higher uncertainty in  $R_{rs}$  and  $\chi_n^2$ , as shown in Figures 12 and 13,  
 860 respectively. Note that  $\chi_n^2$  is the normalized  $\chi^2$  where it is divided by the number of bands used  
 861 in the fitting, such that the theoretical  $\chi_n^2$  should have the mode close to 1.



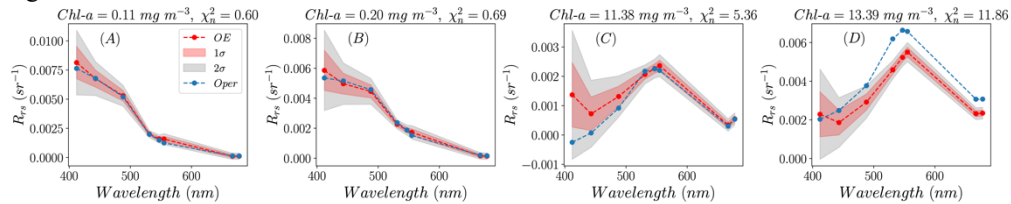
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 863 Fig. 13. The top row is the OE algorithm retrieval from MODIS-A of  $Chl - a$ ,  $\tau_a(869)$ ,  $fmf$ . The bottom row is  
 864 the OE algorithm retrieval of  $a_{nw}$  and  $b_{bp}$  both at 443nm and  $\chi_n^2$ .  
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866 As intermediate products of the OE algorithm, the absorption and scattering coefficients of  
 867 the GIOP model can be retrieved. The ocean non-water absorption,  $a_{nw}$ , and particulate  
 868 backscattering,  $b_{bp}$ , coefficients at 443 nm are shown in Figure 13 as well. The focus of this  
 869 algorithm is on the AC. Thus, any detailed evaluation of the IOPs retrieved we consider to be  
 870 beyond the scope of this manuscript. However, we show the spatial distribution of the IOPs  
 871 since the OE algorithm relies on a realistic estimate of the surface reflectance to better constrain  
 872 the AC process by utilizing more bands, including the visible bands. Both IOPs show realistic  
 873 spatial distributions with relatively high values in coastal waters, particularly within the  
 874 Chesapeake Bay, which is typically dominated by high CDOM absorption. Both coefficients  
 875 are smaller further away from the coast, indicating less presence of absorbing and scattering  
 876 matter in the open ocean.

877 The final panel of Figure 13 shows  $\chi_n^2$ , a good metric to indicate the performance of both  
 878 the forward model and the assumed TOA uncertainty estimate.  $\chi_n^2$  values around 1 show a good  
 879 match between the residual of the forward model at the solution and the uncertainty of the  
 880 signal; higher  $\chi_n^2$  values mean underfitting the forward model and vice versa for lower  $\chi_n^2$ .  
 881 Interestingly, in most of the scene,  $\chi_n^2$  is close to 1, particularly in pixels away from the coast.  
 882 However, in coastal waters,  $\chi_n^2$  values are, for example, higher than 5, indicating either more  
 883 difficulty fitting the observations with the forward model or underestimating the assumed  
 884 measurement/forward model uncertainty. This is expected as coastal waters are significantly

885 more challenging to model with only three parameters, while the atmosphere could also be  
 886 more complex in these regions (i.e., absorbing aerosols).

887 Lastly, Figure 14 compared  $R_{rs}$  from the OE and operational algorithms at the four locations  
 888 A-D in Figure 11. Cases A and B represent low Chl-a conditions with values of 0.11 and 0.2  
 889  $\text{mg m}^{-3}$ , and  $\chi_n^2$  of 0.6 and 0.69, respectively. There is an excellent agreement between both  
 890 retrievals as expected due to the simplicity of the environmental conditions in these waters.  
 891 This demonstrates that the OE algorithm does provide viable  $R_{rs}$  estimates, and low  $\chi_n^2$   
 892 indicates a good fit of the forward model to the measurements. Since the OE algorithm provides  
 893 the uncertainty estimate, we also show 1 and 2 standard deviations of  $R_{rs}$  estimated by the OE  
 894 algorithm.



895  
 896 Fig. 14. Spectral  $R_{rs}$  retrieval using the OE algorithm (red dashed lines) vs the Operational algorithm (blue  
 897 dashed lines) for 4 different cases (locations). The  $1\sigma$  and  $2\sigma$  envelope of the  $R_{rs}$  uncertainty estimated using the OE  
 898 algorithm is shown in red and grey shading, respectively. The 4 different cases (indicated in Fig. 11) highlight  $R_{rs}$  at  
 899 different water conditions, from low to high Chl-a and  $\chi_n^2$ .  
 900

901 Figure 14 panels C and D are the  $R_{rs}$  retrievals for the coastal sites with Chl-a values of  
 902 11.38 and 13.39  $\text{mg m}^{-3}$  and  $\chi_n^2$  of 5.36 and 11.86, respectively. There is a good agreement  
 903 between the two algorithms for the green-red bands with larger deviation in the shorter bands  
 904 for case C. Furthermore, in case C, the OE is higher than the operational algorithm, where  $R_{rs}$   
 905 is (unphysically) negative for the operational retrieval. The agreement between both retrievals  
 906 is mostly within one standard deviation of the OE algorithm, except for the 412 nm band, where  
 907 the OE retrieval appears unrealistically high, likely due to not applying the vicarious calibration  
 908 gain (which would have reduced TOA reflectance at 412 nm by approximately 2%, which is  
 909 significant). The last case (D) is from inland Chesapeake Bay waters that are typically highly  
 910 absorbing (high CDOM concentration) and highly scattering due to sediment discharge from  
 911 several estuaries in the region. Case D shows the worst mismatch between both algorithms. The  
 912 OE  $R_{rs}$  is lower across the whole spectrum than the operational algorithm, except for 412 nm,  
 913 where (similar to case C) it may be overestimated. Cases C and especially D show a high  $\chi_n^2$ ,  
 914 indicating the forward model likely is not fully capturing the radiative conditions of the  
 915 atmosphere and the ocean. Although it is challenging to conclude which algorithm provides a  
 916 more correct retrieval in this case, our previous in-situ matchup analysis indicates that the OE  
 917 algorithm performs better than the operational algorithm overall.

## 918 7. Discussion and conclusion

919 In this paper, we have developed a framework based on the optimal estimation algorithm as  
 920 presented in Rodgers 2000, which relies on Bayes' theorem to find the optimal solution to the  
 921 atmospheric correction problem given a representative model of the atmosphere-ocean system  
 922 and prior information on the state of that system. The advantage of this framework are as  
 923 follows:

- 924 • The ability to calculate pixel-level uncertainty estimates and fully consider the  
 925 covariance of the uncertainty in the system. Since the algorithm propagates the error  
 926 covariance, rather than just the diagonal elements of the covariance (i.e., without  
 927 correlation), it is possible to fully account for the correlation in the  $R_{rs}$  uncertainty  
 928 when further propagating the uncertainty in subsequent products such as IOPs and  
 929 Chl-a [61].

- 930 • Improved computational speed and differentiability through the NN forward model  
931 approach. The algorithm has been accelerated using a NN model that can accurately  
932 perform the forward calculations necessary for the iterative approach to find the  
933 optimal solution. The NN replaces the LUT interpolation of the AC and the analytical  
934 ORM, and also provides the Jacobian matrix needed for the optimization and error  
935 propagation.
- 936 • Potential for better utilization of the information-rich multi-angle polarimeter  
937 instruments for the PACE mission to improve the AC of the Ocean Color Instrument  
938 (OCI). OE can utilize prior information from external sources such as ancillary data  
939 sources. This knowledge about the state of the AO system can be fed into OE,  
940 improving and better constraining the AC problem.
- 941 • Because of the speed, differentiability of the algorithm, and its ability to process the  
942 full dynamic range of atmospheric and oceanic conditions, and it is operationally  
943 capable.
- 944 • The algorithm is flexible in its band set configurations since a spectral weight is  
945 assigned to the cost function, similar to the multi-band AC (MBAC) algorithm [18].  
946 This allows for the use of information from across the spectrum (i.e., using NIR and/or  
947 SWIR only, or using the entire spectrum, including the UV).

948  
949 Although this work demonstrates an improved framework for the AC problem, there are  
950 limitations. This OE framework is a research algorithm and has not been thoroughly tested on  
951 large-scale global data. Also, the OE algorithm requires an accurate uncertainty model of the  
952 TOA reflectance with a reasonable spectral dependence that influences the cost function. To  
953 the best of our knowledge, there has been no standardized approach to model the TOA  
954 uncertainty post-launch, including the covariance in uncertainties. In our work, we attempted  
955 to estimate the TOA uncertainty using MOBY matchups generated during system vicarious  
956 calibration, assuming that the most significant portion of the uncertainty budget is the  
957 instrument's systematic and forward modeling uncertainty. The assumption that the uncertainty  
958 estimates at MOBY can be applied to the global ocean is strong but may not be valid for the  
959 coastal AERONET-OC dataset, as evident from the underestimation of uncertainty relative to  
960 the error, particularly for 550 and 667nm.

961 In the synthetic data analysis, we found that the uncertainty estimate, compared to the truth,  
962 is slightly underestimated on average. The ratio between  $\overline{\Delta_s}$  and  $\overline{u_{sat}}$  of 1 indicates a perfect  
963 uncertainty estimate, and for a ratio >1, it indicates underestimation in the OE uncertainty,  
964 while <1 means overestimation. For the AC parameters, the ratio was 1.36 and 1.08 for  $f_{mf}$   
965 and  $\tau_a$ , respectively. The uncertainty in the IOPs showed an excellent agreement for  $a_{ph}$  with  
966 a ratio of 1.002 and a good agreement for  $a_{dg}$  with a ratio of 1.22, however, the uncertainty  
967 was severely underestimated for  $b_{bp}$  where the ratio is 2.22. On the other hand, it is important  
968 to note that the focus of this paper is to improve the estimate of  $R_{rs}$  and its associated  
969 uncertainty. The ratios between  $\overline{\Delta_s}$  and  $\overline{u_{sat}}$  for  $R_{rs}$  at 443, 555, and 667 nm are 0.98, 1.23, and  
970 0.94, respectively, showing a slight overestimation in the blue and red bands and  
971 underestimation in the green bands.

972 We tested the OE algorithm and its uncertainty estimation technique using the SeaBASS  
973 dataset, encompassing a large dynamic range of water conditions spanning coastal to open  
974 waters. While the overall validation statistics showed an improvement for the OE algorithm  
975 relative to the operational one, the improvement was not significant, where  $|\psi|_m$  was reduced  
976 by 1.7, 5, and 19.7% for 443, 555, and 667 nm, respectively, likely because both algorithms  
977 forward models rely on the same aerosol microphysical assumptions [22]. This is expected  
978 since a large portion of the uncertainty is likely from the modeling assumptions and the inherent  
979 limitations in the validation process that would apply to any newly developed algorithm.  
980 However, the OE algorithm shows an improved bias in the retrieval with fewer negative  $R_{rs}$

981 retrievals, particularly for 667 nm, where the error is the most reduced. This improvement is  
982 likely due to an improved AC and not an effect of lack of the vicarious calibration since the  
983 standard vicarious gain does increase the TOA reflectance by approximately 1%. The ratio  
984 between  $\Delta_s$  and  $\overline{u_{sat}}$  (1.27, 1.2, for 443 and 555 nm, respectively) indicates underestimated  
985 uncertainties at those wavelengths, while the ratio of 0.67 at 667 nm indicates an overestimate.  
986 These ratios show a relatively good agreement, given that we are not fully considering the  
987 uncertainty in the in-situ data and other error sources. Large outliers would significantly impact  
988 the analysis for small signals in the red. However, there is no clear explanation for why 667 nm  
989 uncertainty is overestimated, other than the retrieval error for the SeaBASS dataset is  
990 significantly smaller than that at MOBY (where the uncertainty at TOA is calculated).  
991

992 By extending this analysis to the AERONET-OC sites, we stratified the dataset by different  
993 locations. This is due to the large variability of environmental conditions, proximity to land,  
994 and water conditions [86]. Since the AERONET-OC sites are predominantly coastal, the  
995 validation process is expected to be more challenging. Similar to the SeaBASS dataset, all  
996 statistical metrics show an improvement in the matchups using the OE algorithm relative to the  
997 operational algorithm with a reduction in bias and improvement in error metrics. The matchups  
998 showed a significant decrease in negative  $R_{rs}$  retrievals, particularly for 443 nm, where it is  
999 reduced nearly three times and 12 times for 667 nm. This is a remarkable improvement and  
1000 shows that the simultaneous AO retrieval process using multiple bands for the AC provides a  
1001 valuable advantage over using only NIR bands in coastal waters. Moving to validating  
1002 uncertainties, in the case of using all available data, the agreement between  $\Delta_s$  and  $\overline{u_{sat}}$  is good  
1003 for 443 nm with a ratio of 1.12, showing a slight underestimation. However, for 555 and 667  
1004 nm, the ratio of 1.88 and 1.78 shows a significant underestimation of the uncertainty. This  
1005 underestimation happens at all sites, with the worst two performing sites being MVCO and  
1006 Palgrunden. Both are characterized by low aerosol loadings and smaller fine mode fractions  
1007 than average. MVCO showed a higher median wind speed (4.6 m/s compared to the median of  
1008 all cases of 2.5 m/s). Palgrunden is also a high latitude site where the solar angle is typically  
1009 larger than 40°. Some of these environmental conditions can impact the assessment of the  
1010 uncertainty validation due to underestimating the TOA reflectance uncertainty characterized at  
1011 MOBY and retrieval bias. At 443 nm, all sites showed a good agreement with a ratio that ranges  
1012 from 1.26 to 0.94 (Helsinki being the only site showing slight overestimation). Helsinki and  
1013 Gustav also showed the best agreement for 550 and 665nm; however, they are underestimated.  
1014

1015 There are a few theoretical and practical reasons that could explain the underestimation of  
1016 the satellite-derived uncertainty:

- 1017 • The OE algorithm relies on the assumption of a Gaussian posterior distribution, where  
1018 the variance of the distribution should capture the uncertainty estimate within one  
1019 standard deviation. This is not necessarily true for the atmosphere-ocean system, as  
1020 demonstrated using the grid approximation Bayesian inference method in [30], which  
1021 showed that the full posterior uncertainty is typically larger than the standard deviation  
1022 of a normal distribution.
- 1023 • The error propagation relies on the estimate of the Jacobian matrix (i.e., the first  
1024 derivative). This approximation would not hold for a highly nonlinear relationship  
1025 between the observations and the state parameters. This issue manifests in the  
1026 optimization procedure that relies on the derivative of the cost function, which could  
1027 lead to a local minima leading to a biased inversion.
- 1028 • The absolute normalized error metric requires complete knowledge of the uncertainty  
1029 in the in-situ data for each measurement. This encompasses instrument calibration and  
1030 radiometry knowledge, the effect of environmental conditions on the measurements  
1031 uncertainty, and spatio-temporal mismatch with the satellite retrieval. This is  
1032 consistent with the findings of Zibordi et al., 2022 [92], which found that when

1033 assuming 5% uncertainty in the satellite-derived water-leaving radiance, the absolute  
1034 normalized error metric consistently shows an underestimation of the uncertainty.  
1035 They attributed that to the overly optimistic 5% uncertainty typically set as a gold  
1036 standard for ocean color requirements. Additionally, ignoring complex spatio-  
1037 temporal uncertainties does play a significant role in the underestimation as well as  
1038 possible biases either in in-situ data or satellite retrievals due to, for example, land  
1039 adjacency effects [93].

- 1040 • We assumed ancillary data uncertainty based on fixed absolute and relative  
1041 uncertainties that do not vary with space or time. Recent work has shown that the  
1042 uncertainty varies geographically and could have a significant impact on the  $R_{rs}$   
1043 retrieval, particularly due to relative humidity and windspeed uncertainty which has a  
1044 large impact on the aerosol quantification [27].
- 1045 • We assumed that the uncertainty of the TOA observations estimated at the MOBY site  
1046 is representative of the global oceans. However, in coastal sites, the forward modeling  
1047 errors are likely larger than in the open ocean due to more complexity in the  
1048 atmosphere and ocean optical properties, such as the presence of strongly absorbing  
1049 aerosols and errors in the BRDF correction.

1050  
1051 Finally, in our analysis here of the OE algorithm performance, we did not apply the standard  
1052 vicarious gains that are otherwise applied to the input TOA radiances when operating the  
1053 standard NASA AC algorithm. Our justification is that these vicarious gains are tuned for the  
1054 standard algorithm, which relies on the black-pixel (NIR bands) assumption for the AC rather  
1055 than utilizing the entire visible spectrum as the OE algorithm does. As the OE approach relies  
1056 on all measurements simultaneously, it is less sensitive to measurement uncertainties than NIR-  
1057 based AC algorithms that typically use only two bands (unless there is a large systematic bias  
1058 in the observations). For MODIS-A, the standard vicarious gains are mostly close to 1, except  
1059 for water vapor bands near 645 and 869 nm (likely due to systematic uncertainty in the water  
1060 vapor correction) and 412 nm. The gain coefficient at 412 nm reduces the TOA reflectance by  
1061 ~2%, which is significant (and likely instrument-specific) and would be realized as a large bias  
1062 in  $R_{rs}$  for the OE algorithm. We noticed that  $R_{rs}$  at that band was consistently overestimated  
1063 relative to the operational algorithm, partially explaining less negative  $R_{rs}$  at that band.  
1064 However, negative  $R_{rs}$  at all other bands are significantly reduced, likely because of the better  
1065 constraint on the surface properties using the GIOP forward model. Future work will implement  
1066 a vicarious calibration procedure for the OE algorithm. Therefore, this research algorithm's  
1067 performance can only improve beyond what is presented here. That includes improving the  
1068 aerosol modeling, the RT accuracy, and the bio-optical modeling of the ocean. Our future work  
1069 plan includes the following steps:

- 1070 • Further investigate the impact of the prior information either from models or other  
1071 external sources on the reduction of  $R_{rs}$  uncertainty.
- 1072 • Assess the performance of the full error covariance matrix estimated from the OE  
1073 algorithm.
- 1074 • Develop and apply a system vicarious calibration (SVC) procedure for the OE  
1075 algorithm.
- 1076 • Develop an operational implementation of the OE algorithm for the PACE mission, to  
1077 fully exploit the combined capabilities of the OCI sensor and MAPs for ocean color  
1078 retrievals.

1079 In summary, this work presents a practical recasting of ocean color AC within a Bayesian  
1080 framework. It demonstrates slightly better quantitative retrieval performance than the current  
1081 standard approach, as well as quantitatively relevant pixel-level uncertainty estimates that have  
1082 been missing until now. The OE framework can be applied to current and heritage ocean color  
1083 sensors. Looking to the future, the Bayesian approach would allow the OCI instrument on

1084 PACE, for example, to utilize retrieval products from its companion instruments, the  
1085 information-rich MAPs, as informative priors to further constrain the AC process for OCI. In  
1086 a general sense, the OE framework provides a pathway to take advantage of complementary  
1087 instruments on the same satellite platform or atmospheric measurements from ancillary sources  
1088 to improve the quality of satellite ocean color retrievals.

1089

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1098 **Supplemental document.** See [Supplement 1](#) for supporting content.

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