



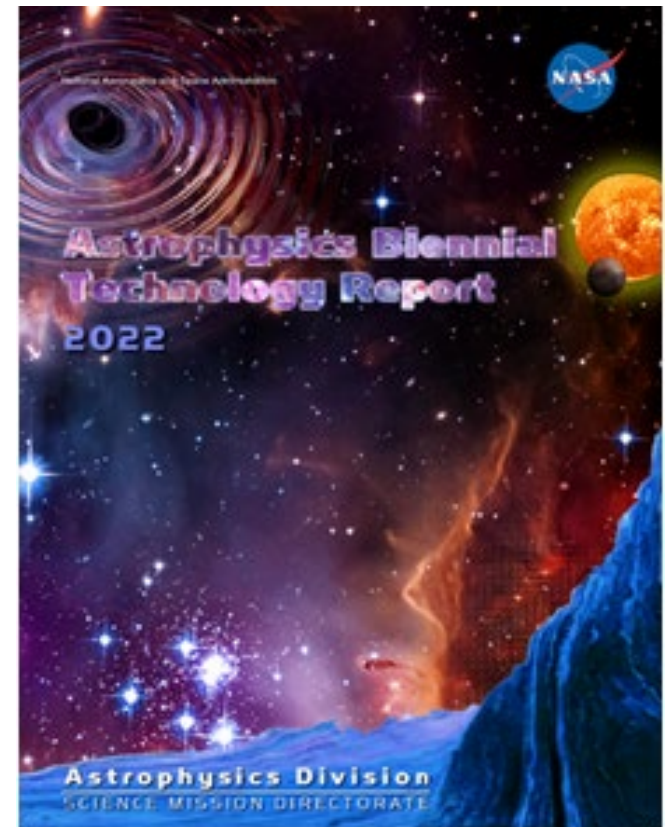
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# **2022 APD Technology Gaps: Prioritization Process and Results**

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- A technology gap is the difference between a capability needed to enable (or enhance) a future mission and the current state-of-the-art
- The Astrophysics Division maintains a prioritized Technology Gap List
- Program Office technologists carry out biennial technology gap prioritizations
  - **Identify technology gaps** applicable to Astrophysics strategic objectives
  - **Rank technology gaps** to prioritize for investment
  - **Inform the community** of Astrophysics technology needs through Astrophysics Biennial Technology Report (ABTR)  
[https://apd440.gsfc.nasa.gov/images/tech/2022\\_ABTR.pdf](https://apd440.gsfc.nasa.gov/images/tech/2022_ABTR.pdf)





# Technology Gap Solicitation and Prioritization Process



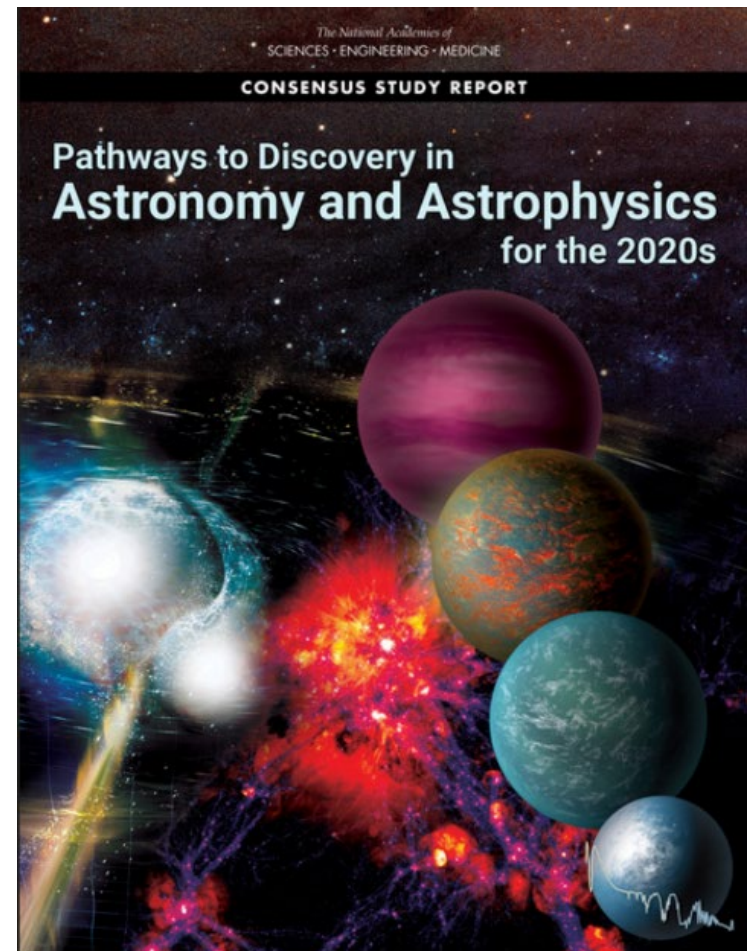
- 1. Technology gaps are solicited from the community, informed by the Decadal Survey**
  - Planned to collect by 6/1/2021, delayed to 1/3/2022 due to Astro2020 release schedule
- 2. Program Office (PO) staff review the collected gaps and assign each to the Program that would be most impacted by closing it**
- 3. Each PO consolidates the inputs for its Program and asks its community to review the gaps for accuracy and completeness before prioritization**
- 4. A Technology Management Board (TMB) reviews and prioritizes the resulting gaps**
  - TMB membership is diverse and includes senior members of Astrophysics Division, STMD, and the POs; technologists and scientists from the three POs; and subject matter experts
  - Prioritization is based on a published set of criteria (strategic alignment, benefits and impacts, urgency, and scope of applicability)
- 5. The lists from the three Programs is merged into a joint, prioritized Astrophysics technology gap list**

# How Did the Astro2020 Affect the 2022 Prioritization?

**1. The strategic missions recommended by Astro2020 replaced the previous set of strategic missions (the TDAMM program was added, despite no missions specified)**

- IROUV Flagship
- X-ray Flagship
- Far-IR Flagship
- X-ray Probe
- Far-IR Probe
- CMB Probe
- TDAMM

**2. The estimated launch timeline informed gap urgency**





# Astrophysics Technology Gap Priorities



[https://apd440.gsfc.nasa.gov/tech\\_gap\\_priorities.html](https://apd440.gsfc.nasa.gov/tech_gap_priorities.html)

## Tier 1 Technology Gaps

Advanced Cryocoolers  
Coronagraph Contrast and Efficiency  
Coronagraph Stability  
Cryogenic Readouts for Large-Format Far-IR Detectors  
Heterodyne Far-IR Detector Systems  
High-Performance, Sub-Kelvin Coolers  
High-Reflectivity Broadband Far-UV-to-Near-IR Mirror Coatings  
High-Resolution, Large-Area, Lightweight X-ray Optics  
High-Throughput Bandpass Selection for UV/VIS  
High-Throughput, Large-Format Object Selection Technologies for Multi-Object and Integral Field Spectroscopy

Large Cryogenic Optics for the Mid IR to Far IR  
Large-Format, High-Resolution Focal Plane Arrays  
Large-Format, Low-Darkrate, High-Efficiency, Photon-Counting, Solar-blind, Far- and Near-UV Detectors  
Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors  
Long-Wavelength-Blocking Filters for X-ray Micro-Calorimeters  
Low-Stress, High-Stability, X-ray Reflective Coatings  
Mirror Technologies for High Angular Resolution (UV/Vis/Near IR)  
Stellar Reflex Motion Sensitivity – Astrometry  
Stellar Reflex Motion Sensitivity – Extreme Precision Radial Velocity  
Vis/Near-IR Detection Sensitivity

## Tier 2 Technology Gaps

Broadband X-ray Detectors  
Compact, Integrated Spectrometers for 100 to 1000  $\mu\text{m}$   
Far-IR Imaging Interferometer for High-Resolution Spectroscopy  
Far-IR Spatio-Spectral Interferometry  
Fast, Low-Noise, Megapixel X-ray Imaging Arrays with Moderate Spectral Resolution  
High-Efficiency X-ray Grating Arrays for High-Resolution Spectroscopy  
High-Resolution, Direct-Detection Spectrometers for Far-IR Wavelengths  
Improving the Calibration of Far-IR Heterodyne Measurements  
Large-Aperture Deployable Antennas for Far-IR/THz/sub-mm Astronomy for Frequencies over 100 GHz

Large-Format, High-Spectral-Resolution, Small-Pixel X-ray Focal-Plane Arrays  
Polarization-Preserving Millimeter-Wave Optical Elements  
Precision Timing for Space-Based Astrophysics  
Rapid Readout Electronics for X-ray Detectors  
Starshade Deployment and Shape Stability  
Starshade Starlight Suppression and Model Validation  
UV Detection Sensitivity

## Tier 3 Technology Gaps

Advancement of X-ray Polarimeter Sensitivity  
Detection Stability in Mid-IR  
Far-UV Imaging Bandpass Filters  
High-Efficiency Far-UV Mirror  
High-Efficiency, Low-Scatter, High- and Low-Ruling-Density, High- and Low-Blazed-Angle UV Gratings

High-Quantum-Efficiency, Solar-Blind, Broadband Near-UV Detector  
Photon-Counting, Large-Format UV Detectors  
Short-Wave UV Coatings  
Warm Readout Electronics for Large-Format Far-IR Detectors

## Tier 4 Technology Gaps

Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry  
Improving the Photometric and Spectro-Photometric Precision of Time-Domain and Time-Series Measurements

UV/Opt/Near-IR Tunable Narrow-Band Imaging Capability  
Very-Wide-Field Focusing Instrument for Time-Domain X-ray Astronomy

## Tier 5 Technology Gaps

Complex Ultra-Stable Structures for Future Gravitational-Wave Missions  
Disturbance Reduction for Gravitational-Wave Missions  
Gravitational Reference Sensor  
High-Performance Spectral Dispersion Component/Device  
High-Power, High-Stability Laser for Gravitational-Wave Missions  
Laser Phase Measurement Chain for a Decihertz Gravitational-Wave Mission  
Micro-Newton Thrusters for Gravitational Wave-Missions  
Stable Telescopes for Gravitational Wave-Missions

You can find this priority list in the ABTR, or at the above URL



# Gaps by Program



Program	Tier 1	Tier 2	Tier 3	Tier 4	Tier 5	Total	2019
COR	11	7	7	2	1	<b>28</b>	23
ExEP	6	3	1	0	0	<b>10</b>	12
PCOS	3	6	1	2	7	<b>19</b>	13
<b>Total</b>	<b>20</b>	<b>16</b>	<b>9</b>	<b>4</b>	<b>8</b>	<b>57</b>	<b>48</b>



# Gaps by Relevant Strategic Mission



Mission	Tier 1	Tier 2	Tier 3	Tier 4	Tier 5	Total
IROUV	11	3	6	1		<b>21</b>
Far-IR Flagship	6	8	2			<b>16</b>
X-ray Flagship	3	5	1			<b>9</b>
Far-IR Probe	6	8	2			<b>16</b>
X-ray Probe	3	5	1	1		<b>10</b>
CMB Probe				1		<b>2</b>
TDAMM				2		<b>2</b>
None					8	<b>8</b>
<b>Total</b>	<b>29</b>	<b>30</b>	<b>12</b>	<b>5</b>	<b>8</b>	<b>84</b>



# What do These Priorities Mean?



- **The technology gap list informs the SAT solicitation and selections; historically the focus has been on the first two priority tiers**
- **However, gaps in lower tiers are not ignored**
  - Technologies that address any gap, whether solicited in SAT or not, may fit in APRA
  - Gaps in lower tiers have at times moved to higher tiers in later cycles (a new Astrophysics Implementation Plan, AIP, is expected to be released by the end of 2022, and will inform the next cycle)
  - Astrophysics Division may decide to direct-fund technologies they deem important enough after considering programmatic aspects, for example through the work of the Technology Strategy Team (TST)





# Where to find the 2022 Gap List

[https://apd440.gsfc.nasa.gov/tech\\_gap-descriptions.html](https://apd440.gsfc.nasa.gov/tech_gap-descriptions.html)



- Overview
- Technology
- Outreach
- 🔍

Astrophysics Program Offices

## 2022 Astrophysics Strategic Technology Gaps

[TECHNOLOGY GAPS: OVERVIEW](#) / [TECH GAP PRIORITIES](#) / [PRIORITIZATION PROCESS](#) / [TECH GAP DESCRIPTIONS](#)

[Download the Excel Spreadsheet](#)

Gap Name	Description	Current State-of-the-Art	TRL	Performance Goals and Objectives	Scientific, Engineering, and/or Programmatic Benefits	Applications and Potential Relevant Astrophysics Missions	Urgency
Cryogenic Readouts for Large-format Far-IR Detectors	Readout schemes including cryogenic multiplexing for arrays of large-format Far-IR detectors need to be developed.	Readout schemes using HEMT or SiGe amplifiers and frequency-multiplexed resonant circuits are in development. A few hundred channels per 1 mW HEMT amplifier have been demonstrated. Low power dissipation at 4 K is required. For TES-based detectors a microwave SQUID multiplexer using frequency division, time division or code division multiplexing is needed. Frequency division multiplexing is well advanced and can meet the needs of a far-IR flagship when scaled to 2000 pixels (resonators) per 4 GHz channel.	3	Near-term, this scheme should result in 2000 pixels per amplifier channel (enabling) 3000 pixels/channel (enhancing). HEMT amplifiers from Low Noise Factory can achieve 10 dB with 0.38 mW of dissipation at 4 K.	Sensitivity reduces observing times from many hours to a few minutes (~100x faster), while array format increases areal coverage by x10-100. Overall mapping speed can increase by factors of thousands. Sensitivity enables measurement of low-surface-brightness debris disks and protogalaxies with an interferometer. This is enabling technology. Suborbital and ground-based platforms can be used to validate technologies and advance TRL of new detectors.	Far-IR Flagship Far-IR Probe FIR detector technology is an enabling aspect of all future FIR mission concepts, and is essential for future progress. This technology can improve science capability at a fixed cost much more rapidly than larger telescope sizes. This development serves Astrophysics almost exclusively (with some impact on planetary and Earth studies). Many synergies exist with similar developments for x-ray microcalorimeters	Required TRL 6 by mission PDR. Extreme stretch Single tech

You can download the details of all gaps in Excel from this page



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# Backup



# Gap Submission Guidelines

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1. Focus on technology gaps associated with missions prioritized in Astro2020
2. Submit gaps directly applicable to Program objectives (not ones outside our purview (e.g., associated with launch vehicle, rover, avionics, s/c systems, etc.)
3. Don't include gaps that don't require technology development, that are not well defined, that are redundant (duplicate, similar, or subsets of existing gaps), or where solutions are at TRL 6 or higher for the relevant strategic mission(s)
4. Inputs should be submitted as gaps between the current state-of-the-art and what's required to achieve the science objective targeted, not specific solutions
5. Inputs should not endorse or advertise any organization, mission, or person
6. Inputs should not contain proprietary, or EAR/ITAR-restricted information

**Full details are provided in the gap submission form instructions**



# Four Prioritization Criteria

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- **Strategic Alignment:** How well does the technology align with astrophysics science and/or programmatic priorities set out in the Decadal Survey?
- **Benefits and Impacts:** How much impact does the technology have on applicable missions? To what degree does it enable and/or enhance achievable science objectives, reduce cost, and/or reduce mission risks?
- **Urgency:** Given the anticipated difficulty of maturing from current TRL of a full solution to TRL 6, assessed against the time available until the need-by year, how urgently does the gap need to be addressed?
- **Scope of Applicability:** How crosscutting is the technology? How many Astrophysics programs and/or mission concepts (strategic or other) would benefit by closing the gap?