

# Photochemical Evolution of the 2013 California Rim Fire: Synergistic Impacts of Reactive Hydrocarbons and Enhanced Oxidants

5 Glenn M. Wolfe<sup>1</sup>, Thomas F. Hanisco<sup>1</sup>, Heather L. Arkinson<sup>2</sup>, Donald R. Blake<sup>3</sup>, Armin Wisthaler<sup>4,5</sup>,  
Tomas Mikoviny<sup>5</sup>, Thomas B. Ryerson<sup>6,7,\*</sup>, Ilana Pollack<sup>7,\*\*</sup>, Jeff Peischl<sup>7</sup>, Paul O. Wennberg<sup>8,9</sup>, John D.  
Crouse<sup>8</sup>, Jason M. St. Clair<sup>8,\*\*\*</sup>, Alex Teng<sup>8,\*\*\*\*</sup>, L. Greg Huey<sup>10</sup>, Xiaoxi Liu<sup>10,\*\*\*\*\*</sup>, Alan Fried<sup>11</sup>, Petter  
Weibring<sup>11</sup>, Dirk Richter<sup>11</sup>, James Walega<sup>11</sup>, Samuel R. Hall<sup>12</sup>, Kirk Ullmann<sup>12</sup>, Jose L. Jimenez<sup>7,13</sup>, Pedro  
Campuzano-Jost<sup>7,13</sup>, T. Paul Bui<sup>14</sup>, Glenn Diskin<sup>15</sup>, James R. Podolske<sup>14</sup>, Glen Sachse<sup>15,16,†</sup>, and Ronald  
C. Cohen<sup>17,18</sup>

<sup>1</sup>Atmospheric Chemistry and Dynamics Laboratory, NASA Goddard Space Flight Center, Greenbelt, MD, USA

10 <sup>2</sup>Department of Oceanic and Atmospheric Science, University of Maryland, College Park, MD, USA

<sup>3</sup>Department of Chemistry, University of California Irvine, Irvine, CA, USA

<sup>4</sup>Institute for Ion Physics and Applied Physics, University of Innsbruck, Innsbruck, Austria

<sup>5</sup>Department of Chemistry, University of Oslo, Oslo, Norway

<sup>6</sup>Chemical Sciences Laboratory, NOAA, Boulder, CO, USA

15 <sup>7</sup>Cooperative Institute for Research in Environmental Sciences, University of Colorado Boulder, Boulder, CO, USA

<sup>8</sup>Division of Geological and Planetary Sciences, California Institute of Technology, Pasadena, CA, USA

<sup>9</sup>Division of Engineering and Applied Science, California Institute of Technology, Pasadena, CA, USA

<sup>10</sup>School of Earth and Atmospheric Sciences, Georgia Institute of Technology, Atlanta, GA, USA

<sup>11</sup>Institute of Arctic and Alpine Research, University of Colorado, Boulder, CO, USA

20 <sup>12</sup>Atmospheric Chemistry Observations and Modeling Laboratory, National Center for Atmospheric Research, Boulder, CO, USA

<sup>13</sup>Department of Chemistry, University of Colorado Boulder, Boulder, CO, USA

<sup>14</sup>Atmospheric Sciences Branch, NASA Ames Research Center, Moffett Field, CA, USA

<sup>15</sup>NASA Langley Research Center, Hampton, VA, USA

25 <sup>16</sup>National Institute of Aerospace, Hampton, VA, USA

<sup>17</sup>Department of Earth and Planetary Sciences, University of California, Berkeley, CA, USA

<sup>18</sup>College of Chemistry, University of California, Berkeley, CA, USA

\*Now at Scientific Aviation, Boulder, CO, USA

\*\*Now at Department of Atmospheric Science, Colorado State University, Fort Collins, CO, USA

30 \*\*\*Now at Joint Center for Earth Systems Technology, University of Maryland Baltimore County, Baltimore, MD, USA

\*\*\*\*Now at Fifty Years, San Francisco, CA, USA

\*\*\*\*\*Now at California Air Resource Board, Los Angeles, CA, USA

†Deceased

*Correspondence to:* Glenn M. Wolfe (glenn.m.wolfe@nasa.gov)

## 35 Abstract

Large wildfires markedly alter regional atmospheric composition, but chemical complexity challenges model predictions of downwind impacts. Here, we elucidate key facets of gas-phase photochemistry and assess novel chemical processes via a case study of the 2013 California Rim Fire plume. Airborne in situ observations, acquired during the NASA Studies of Emissions, Atmospheric Composition, Clouds and Climate Coupling by Regional Surveys (SEAC<sup>4</sup>RS) mission, illustrate the evolution of

40 volatile organic compounds (VOC), oxidants, and reactive nitrogen over 12 hours of atmospheric aging. Measurements show rapid formation of ozone and peroxyacyl nitrates (PNs), sustained peroxide production, and prolonged enhancements in oxygenated VOC and nitrogen oxides (NO<sub>x</sub>).

Measurements and Lagrangian trajectories constrain a 0-D puff model that approximates plume photochemical history and provides a framework for evaluating key processes. Simulations examine the effects of 1) previously-unmeasured  
45 reactive VOC identified in recent laboratory studies, and 2) emissions and secondary production of nitrous acid (HONO). Inclusion of estimated unmeasured VOC leads to a 250% increase in OH reactivity and a 70% increase in radical production via oxygenated VOC photolysis. HONO amplifies radical cycling and serves as a downwind NO<sub>x</sub> source, although two different HONO production mechanisms (particulate nitrate photolysis and heterogeneous NO<sub>2</sub> conversion) exhibit markedly different effects on ozone, NO<sub>x</sub>, and PNs. Analysis of radical initiation rates suggests that oxygenated VOC photolysis is a  
50 major radical source, exceeding HONO photolysis when averaged over the first 2 hours of aging. Ozone production chemistry transitions from VOC-sensitive to NO<sub>x</sub>-sensitive within the first hour of plume aging, with both peroxide and organic nitrate formation contributing significantly to radical termination. To simulate smoke plume chemistry accurately, models should simultaneously account for the full reactive VOC pool and all relevant oxidant sources.

## 1 Introduction

55 Biomass burning accounts for at least 30% of global emissions of non-methane volatile organic compounds (VOC) (Akagi et al., 2011; Andreae, 2019; Yokelson et al., 2008). Pyrogenic VOC fuel the production of ozone and secondary organic aerosol, with significant consequences for regional and global air quality and climate (Jaffe and Wigder, 2012; McClure and Jaffe, 2018; Buysse et al., 2019; Val Martin et al., 2015; Hodshire et al., 2019). Pyrogenic emissions consist of thousands of unique compounds. Existing emission inventories include over 100 individual VOC (Andreae, 2019; Akagi et al., 2011). These  
60 compilations are incomplete, and it is estimated that previously “unidentified” VOC account for ~50% of total pyrogenic VOC mass (Akagi et al., 2011; Yokelson et al., 2013; Gilman et al., 2015). Recent advances in instrumentation have broadened the suite of detectable VOC to over 500 species (Sekimoto et al., 2018; Hatch et al., 2019, 2017; Koss et al., 2018), although significant uncertainty remains regarding the speciation, reactivity, and fate of this extended VOC pool.

The photochemistry of biomass burning plumes is perhaps less well understood than emissions, especially within the  
65 context of total reactive VOC. Reported observations of ambient smoke typically encompass a few hours of physical age and include a (relatively) limited set of compounds (Akagi et al., 2012, 2013; Alvarado et al., 2010; Liu et al., 2016; Müller et al., 2016). Detailed simulations may provide some insight into the chemistry of unidentified VOC, though models are often under-constrained and both model and measurement uncertainties can be large (Lonsdale et al., 2019; Alvarado et al., 2015; Müller et al., 2016; Mason et al., 2001; Liu et al., 2016). With extended VOC inventories from recent laboratory work (Hatch et al.,  
70 2017; Koss et al., 2018), several studies have begun characterizing the impacts of previously unidentified VOC. Analyzing a series of laboratory burns, Coggon et al. (2019) estimate that species included in the Master Chemical Mechanism

(MCMv3.3.1) account for ~60% of the primary hydroxyl radical (OH) reactivity measured via Proton Transfer Time-of-flight Mass Spectrometry (PTR-MS). This study also demonstrates that furans, a previously unconsidered class of reactive VOC, may increase O<sub>3</sub> production within an agricultural fire plume by ~10% over the first hour of aging. Decker et al. (2019) estimate  
75 that the MCM accounts for ~30% of the observed nitrate radical (NO<sub>3</sub>) reactivity in the same laboratory experiments.

Despite improved knowledge of pyrogenic VOC speciation, it remains unclear how to best synthesize this information within existing regional/global model frameworks. It is not feasible to add a multitude of new species to chemical mechanisms; indeed, many global models do not explicitly account for all “known” VOC, let alone the large fraction of unidentified VOC (Duncan et al., 2007; Wiedinmyer et al., 2011). Nor is it evident that such modifications would improve model results, as  
80 plume chemistry may be a “sub-grid scale” process more suited to parameterization than explicit simulation. On the other hand, reactive VOC chemistry can persist for days downwind of a fire (Mauzerall et al., 1998; Alvarado et al., 2020; Forrister et al., 2015), and an improved representation of such processes may alter regional budgets of ozone, CO, oxidized VOC, reactive nitrogen, and organic aerosol.

Oxidant sources also remain poorly understood in this environment. Smoke plumes are photochemically complex due to  
85 spatial and temporal variability in radical precursors, radiation, and other factors (Wang et al., 2021). Recent work has highlighted the importance of nitrous acid (HONO) emissions as a radical source in nascent smoke plumes (Peng et al., 2020; Theys et al., 2020; Robinson et al., 2021). Secondary HONO production via heterogeneous processes may sustain downwind radical production in some, but not all, smoke plumes (Alvarado and Prinn, 2009; Alvarado et al., 2015). Multiple mechanisms have been proposed to explain observed HONO in other environments (Zhang et al., 2019), but their controlling factors and  
90 potential impacts are not well-characterized.

Here, we utilize a case study of the 2013 California Rim Fire to examine the impacts of newly identified reactive VOC and HONO on gas-phase chemistry. Airborne in situ observations from the NASA Studies of Emissions, Atmospheric Composition, Clouds and Climate Coupling by Regional Surveys (SEAC<sup>4</sup>RS) mission constrain the evolution of oxidants and oxidation products over ~12 hours of atmospheric aging. We combine these observations with air mass trajectories to drive a  
95 0-D puff model that approximates plume aging. With this framework, we illustrate how an extended VOC pool and various HONO sources alter the chemistry of oxidants, oxidized VOC, reactive nitrogen, and ozone. We evaluate model results through comparison with observations and use the model to quantify time and mechanism-dependent changes in OH reactivity, radical production/termination, and ozone production sensitivity to VOC and NO<sub>x</sub>.

## 2 Methods

### 100 2.1 Rim Fire and SEAC<sup>4</sup>RS

The Rim Fire was an extreme wildfire in the central Sierra Nevada Mountains of California (37.85°N, 120.08°W). Ignited by an illegal campfire on 17 August 2013, the fire was not fully contained until 24 October 2013 and consumed a total area exceeding 1000 km<sup>2</sup>. At the time of SEAC<sup>4</sup>RS flights, the fire front spanned a width of ~40 km in both the N-S and E-W

directions. Fuels consisted mostly of mixed conifer forest (Lydersen et al., 2014). Previous Rim Fire studies have investigated  
105 fire meteorology (Peterson et al., 2015), emissions (Liu et al., 2017; Yates et al., 2016; Saide et al., 2015), aerosol properties  
(Forrister et al., 2015; Adler et al., 2019; Perring et al., 2017), radiative effects (Yu et al., 2016a), and regional chemistry model  
performance (Baker et al., 2018). We focus here on downwind gas-phase photochemistry.

SEAC<sup>4</sup>RS was a NASA-funded effort to characterize the processes controlling atmospheric properties in the summertime  
U.S. (Toon et al., 2016). We utilize in situ observations acquired from the NASA DC-8 aircraft, which sampled Rim Fire  
110 outflow on 26 and 27 August 2013. Analysis focuses on the “long-axis” portion of the 26 August flight, which extends from  
directly over the fire to 470 km downwind (Fig. 1). The actual sample time window for this leg is UTC 23:06 to 23:57, and  
the physical smoke age ranges from 0 to ~12 hours (Sect. 2.2). The aircraft maintained a pressure altitude of 4.3 km until 23:32  
(smoke age ~5 h), when it descended to 3.6 km (Supplementary Information (SI) Fig. S1). Terrain was mountainous, and  
corresponding altitude above ground level ranged from 1.3 to 3.2 km. Sampling was predominantly in the upper edge of the  
115 smoke plume, which mostly resided in the lower free troposphere according to boundary layer depth derived from trajectory  
meteorological fields (Fig. S1c). The modified combustion efficiency (MCE) near the source ranged from 0.91 to 0.94,  
indicating a mix of flaming and smouldering combustion. While the DC-8 also sampled Rim Fire smoke as old as 2 – 3 days,  
influence from other fires, surface-atmosphere exchange, and changing background concentrations complicates analysis of  
that data.

120 SI Table S1 lists the instruments and measurement accuracy for observations used in this study; the payload is further  
detailed in Toon et al. (2016). Fast (1 Hz) data are averaged to the 40-second collection window of the Whole Air Sampler  
(WAS). Normalized excess mixing ratios (NEMRs) are calculated using carbon monoxide (CO) from the Differential  
Absorption Carbon Monoxide Measurement (DACOM) as the dilution tracer.

$$NEMR(X) = \Delta X / \Delta CO = (X - X_b) / (CO - CO_b) \quad (1)$$

125 Here,  $X$  is the observed mixing ratio and  $X_b$  is the background mixing ratio. Fast observations fluctuate rapidly in fire plumes  
and can sometimes contain gaps over a portion of a WAS sampling interval. We adopt a custom averaging procedure to ensure  
rigorous NEMR calculation for such data. First, each 1 Hz variable  $X$  is time-aligned to DACOM CO via optimized lag-  
correlation. Second, CO is filtered to exclude points where  $X$  is missing. Finally,  $X$  and the filtered CO are averaged to the  
WAS time base. Executing this procedure over all data effectively creates a unique, gap-filtered, WAS-averaged DACOM CO  
130 for each variable. Background mixing ratios are averaged over a single WAS sample collected east of the plume (orange star  
in Fig. 1). We also explored using observations just upwind of the fire for background estimation but found poor agreement  
between the modelled and observed O<sub>3</sub> NEMR in this case due to higher upwind O<sub>3</sub> (78 vs. 50 ppbv). Figure S2 shows the  
dilution factor as a function of plume age, which ranges from 1 to 24.

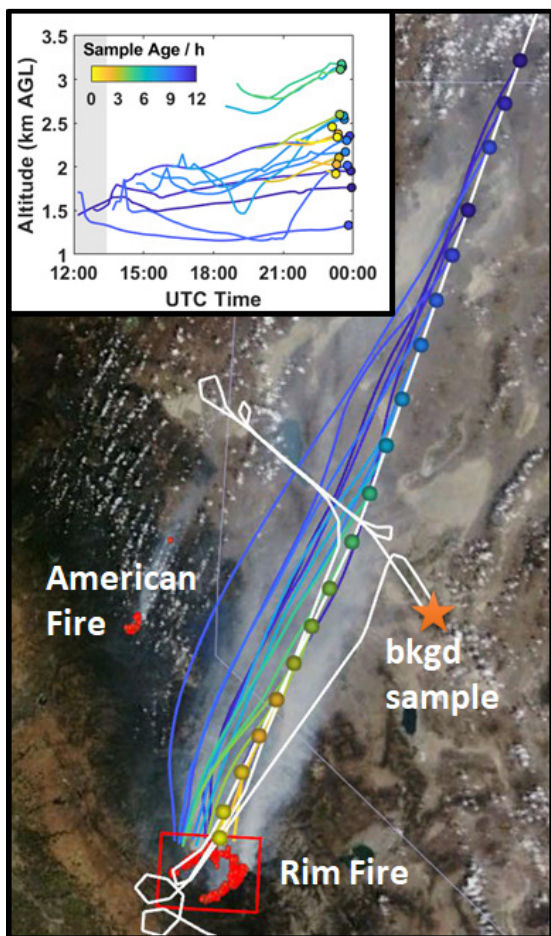


Figure 1. Map of the study region. The background shows visible imagery and fire counts (red dots) from Terra/MODIS (<https://worldview.earthdata.nasa.gov/>). The white line is the DC-8 flight track. The aircraft entered from the south, flew over the fire, proceeded NE to do a missed approach at Fallon airport (near orange star), executed a “wall” pattern of perpendicular transects at several altitudes, and returned to the fire before proceeding NE along the long axis leg. Colored circles denote WAS sample locations and colored lines are the corresponding median trajectories. Yellow-to-blue shading indicates Lagrangian plume age at the time of sampling (see color bar in inset). The orange star is the location of the background sample. The red box is the “Rim Fire” box used for trajectory filtering. (Inset) Time series of median trajectory altitudes with respect to ground level. Colored circles denote the observation time. Gray shaded area indicates night time.

## 135 2.2 Trajectory Analysis

Air mass trajectories provide an estimate of smoke age and inputs for the puff model described below. Trajectories are computed with the Hybrid Single-Particle Lagrangian Integrated Trajectory (HYSPLIT) Model (Stein et al., 2015), interfaced with custom MATLAB software for input and output handling (<https://github.com/AirChem/HYSPLITcontrol>, last accessed 1 Nov 2021). Trajectories are initialized at the midpoint of each WAS sample and run backward in time for 48 hours. Outputs include both position and meteorology (temperature, pressure, relative humidity, solar zenith angle (SZA), and boundary layer depth). To build statistics, we utilize archive meteorological fields from the North American Mesoscale Forecast System (NAMs, 12 km, 1-hour) and North American Regional Reanalysis (NARR, 32 km, 3-hour) datasets (<https://ready.arl.noaa.gov/archives.php>, last accessed 24 March 2020). Trajectories run as meteorological ensembles with HYSPLIT default settings (27 members, 1 grid-point horizontal shift, 0.01 sigma unit vertical shift), giving 54 trajectories per initialization. Next, we define a geographic box for the Rim Fire based on satellite imagery and fire counts (Fig. 1). Trajectories are filtered to exclude those not passing through this box. Filtered trajectories are spatially averaged using a geometric median,

resulting in one “median trajectory” for each observation (Fig. 1, inset). Meteorological quantities are averaged via arithmetic mean.

We define Lagrangian smoke age from the time when an averaged trajectory first intersected the Rim Fire box out to the downwind time of sampling. The standard deviation of the smoke age from individual trajectories is typically ~9% of the mean for each ensemble, and we take this as a  $1\sigma$  precision estimate. Smoke at our nominal starting point (the North end of the box) is actually a mixture of ages. Based on observed wind speed, we estimate a transit time of ~1 hour for emissions from the Southern-most fire front to reach the North end of the box. We take half of this value as an approximation for variability in actual smoke age. The total estimated uncertainty in average smoke age is then  $9\% + 0.5\text{ h}$ .

### 155 2.3 Puff Model

Simulation of any fire plume is challenging due to high concentrations, strong dilution, aerosol radiative perturbations, and other factors. SEAC<sup>4</sup>RS did not sample the core of the Rim Fire plume, and smoke ages of 0 – 12 hours were probed over a single hour of flight. Thus, if we wish to compare with observations, it is not adequate to represent the Rim Fire as a single Lagrangian plume. Detailed fire plume models have improved in recent years (Lonsdale et al., 2019), but observational constraints are limited compared to the complexity of such a model. The goal of our simulation is to obtain a meaningful comparison against observations without over-elaboration (Box, 1976).

In our model framework, the plume is approximated as a series of 0-dimensional “puffs.” The puff model is developed within the Framework for 0-D Atmospheric Modelling (F0AMv4.2.1, available at <https://github.com/AirChem/F0AM/>) (Wolfe et al., 2016). One puff is simulated for each WAS sample/trajectory pair shown in Fig. 1. Each puff evolves in real time along the “average trajectory” with meteorological constraints updated every 10 minutes. For each puff, along-trajectory pressure, temperature, relative humidity, and solar zenith angle are constrained with averaged trajectory output. Trajectory meteorology is rescaled by multiplicative factors (typically within a few percent of unity) based on the ratio of observed-to-trajectory values at trajectory endpoints. Photolysis frequencies are semi-constrained to observations using a two-part scaling that helps account for smoke radiative perturbations (Appendix A). Variations in scaling factors suggest stronger attenuation in the UV than the visible (compare  $J(\text{O}^1\text{D})$  and  $J(\text{NO}_2)$  in Fig. A1b). Chemical concentrations from the endpoint of each puff are extracted for comparison with observations.

Initial concentrations are the same for all puffs. These are estimated by combining Rim Fire emission ratios (ERs) from Liu et al. (2017) with excess DACOM CO mixing ratio from the first WAS sample of the long-axis leg. Normalized excess mixing ratios for the near-source WAS sample are generally within  $\pm 50\%$  of the ERs reported by Liu et al. (2017) for the Rim Fire (Fig. S3). The Liu et al. (2017) ERs incorporate multiple near-source intercepts and are more representative of average fire conditions. Actual ERs likely vary among the plume samples; in particular, we might expect older samples to show an increased signature of smouldering relative to flaming due to typical wildfire diurnal progression (Wiggins et al., 2020). Observations, however, do not conclusively indicate an age-dependent trend in ERs. MCE generally declines with age (Fig. S4a), but frequent deviation from the expected wildfire value range of 0.8 – 1.0 (Akagi et al., 2011) suggests non-emission

180 influence (e.g., background CO<sub>2</sub> variability) that degrades this metric as a combustion phase tracer at later ages. NEMRs for the conserved fire tracers HCN and CH<sub>3</sub>CN exhibit some variability (mean/standard deviation = 10 – 15%) but no trend (Fig. S4b), suggesting no systematic change in fire-average ERs (Roberts et al., 2020). Conversely, the trend in total observed NO<sub>y</sub> may indicate time-varying emissions (Sect. 3.1.4). Given these ambiguities, constant initial concentrations is a reasonable assumption.

185 Dilution is treated in analogy with Gaussian plume dispersion (Alvarado et al., 2015):

$$\left(\frac{\partial X}{\partial t}\right)_{dil} = \frac{-1}{\tau_g + 2t}(X(t) - X_b) \quad (2)$$

Background concentrations ( $X_b$ ) are the same as those used for dilution normalization. The Gaussian timescale,  $\tau_g$ , is constant for each puff and calculated using the analytical solution of the integral of Eq. (2) with CO concentrations at start and endpoints. The Gaussian timescale varies from 28 to 760 s for individual puffs (Fig. S5). Such variability is not surprising given  
190 the horizontal extent of the fire and differences among trajectories for each puff.

Model chemistry utilizes the Master Chemical Mechanism (MCMv3.3.1) (Jenkin et al., 2015, 1997; Saunders et al., 2003) with modifications. Additional reactions include photolysis of pernitric acid (Atkinson et al., 2004), reaction of methyl peroxy radical with OH (Assaf et al., 2016; Caravan et al., 2018), reaction of hydroxymethyl hydrogen peroxide (HMHP) with OH (Allen et al., 2018), and oxidation of propadiene (C<sub>3</sub>H<sub>4</sub>). For the latter, we use the OH reaction rate coefficient of Atkinson  
195 and Arey (2003) rather than the 1.8-times slower rate coefficient of Daranlot et al. (2012), based on the similar observed decay rates of propadiene and ethene. Subsequent propadiene chemistry follows the mechanisms of Daranlot et al. (2012) and Xu et al. (2019). We also update rate coefficients for reaction of peroxyacetic acid with OH (Berasategui et al., 2020) and peroxyacyl radicals with HO<sub>2</sub> (Jenkin et al., 2019), which are slower/faster than MCM default values by factors of 123/1.33, respectively. Oxidation of some additional biomass burning VOC (furans, syringol, and guaiacol) is incorporated using an MCM extension  
200 developed following recent laboratory and field studies (Coggon et al., 2019; Decker et al., 2019, 2021; Robinson et al., 2021).

There are multiple potential error sources in the puff model, and many are not easily quantified. We assume constant initial concentrations, but the smoke is a heterogeneous mixture of multiple ages and burning phases. Parameterizations for photolysis and dilution make assumptions about the history of each puff based on the observed evolution along the flight path. Heterogeneous chemistry is explicitly included. Despite all these issues, results will illustrate that the puff model is a reasonable  
205 approximation and a useful testbed for probing plume chemistry. Measurement accuracy (Table S1) is typically the dominant uncertainty in observations at high signal/noise ratios, thus we use this to define the uncertainty for model-measurement comparisons.

## 2.4 Model Scenarios

Model scenarios systematically characterize the effects of varying emissions and chemistry (Table 1). In the base simulation  
210 (M0), initial species are limited to observations. This includes three species that are not in the MCM: furan, HMHP, and propadiene. Additional simulations incorporate unmeasured reactive VOC and various HONO sources as detailed below.

**Table 1. Summary of model simulations.**

Simulation	Description
M0	Base simulation using only measured VOC
M1	M0 + Unmeasured VOC
M2a,b,c	M1 + primary HONO (5, 15, 25 ppbv)
M3a,b,c	M1 + secondary HONO via $\text{pNO}_3^-$ photolysis ( $J(\text{pNO}_3^-)$ scaling factors of 0.5, 1, 2)
M4a,b	M1 + secondary HONO via $\text{NO}_2$ + aerosol ( $\square(\text{NO}_2)$ scaling factors of 1, 1000)

#### 2.4.1 Addition of Unmeasured VOC

Simulation M1 incorporates unmeasured VOC using the fire laboratory emissions data of Koss et al. (2018), which includes over 500 compounds and 20 Western US fuel types. We restrict this dataset to 152 compounds with specific molecular assignments and estimated OH reaction rate coefficients (Table S5 of Koss et al. (2018)). Incorporating this data requires 1) estimating initial VOC concentrations, and 2) allocating these VOC to model species. For the first step, we create a composite ER profile for the Rim Fire by comparing the ERs of 11 species reported by both Liu et al. (2017) and Koss et al. (2018): acetylene, ethene, propene, methanol, formaldehyde, acetaldehyde, furan, benzene, toluene, HCN, and  $\text{CH}_3\text{CN}$ . Specifically, we optimize fractional fuel contributions by minimizing the sum square relative error of log-transformed ERs between Liu et al. (2017) and the composite profile. This procedure suggests significant contributions from Jeffrey Pine (34%), Ceanothus (21%), and Manzanita (15%), with lesser contributions from other fuels. Estimated fuel composition is sensitive to which VOC are included in the optimization, as ER profiles are highly correlated among different fuels. For the same reason, model results are not especially sensitive to the choice of fuel composition. Combining composite ERs with observed CO in the first plume sample gives initial mixing ratios. We assume zero background for all unmeasured VOC. ERs for acrolein and biacetyl are reduced by a factors of 2.3 and 10, respectively, due to known calibration issues (SI Text S1).

Unmeasured VOC are assigned to model species in two steps. 51 of these gases appear in the MCM or the extended NOAA biomass burning mechanism. The remaining 101 compounds lack direct MCM analogues. These VOC are allocated to MCM proxies based on OH reaction rate coefficients ( $k_{\text{OH}}$ ) and molecular formulae. For each VOC, we first identify MCM species that are within some threshold (nominally 20%) of the VOC's estimated  $k_{\text{OH}}$ . We then attempt to filter the MCM species list to include only those containing a similar number of carbon and oxygen atoms (within  $\pm 1$  the number of each atom). Specific functional groups are not considered. If the molecular formulae criterion is too restrictive (i.e., no species are identified), we only use the  $k_{\text{OH}}$  criterion. This procedure, though imperfect, optimizes initial OH reactivity and total VOC mass with the available information. Table S2 lists the MCM assignments and ERs for all unmeasured VOC.



## 2.4.2 Addition of HONO

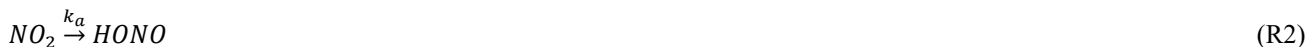
Several sets of sensitivity simulations explore the impacts of HONO from both primary (emissions) and secondary sources. All of these simulations include the same initial unmeasured VOC as simulation M1. Simulations M2a, b, and c incorporate primary HONO at initial mixing ratios of 5, 15, and 25 ppbv, respectively. The upper end of this range is based on recent work reporting an average  $\Delta\text{HONO}/\Delta\text{NO}_x$  emission ratio of  $0.7 \pm 0.3$  ppbv ppbv<sup>-1</sup> and an average  $\Delta\text{HONO}/\Delta\text{CO}$  emission ratio of  $5.3 \pm 5.2$  pptv ppbv<sup>-1</sup> for Western U.S. wildfires (Peng et al., 2020).

Simulations M3a, b, and c incorporate photolysis of particulate nitrate (pNO<sub>3</sub><sup>-</sup>):



The photolysis frequency for reaction R1 is rescaled from that of gas-phase HNO<sub>3</sub> by a factor of 286 following Ye et al. (2017, 2018). The mechanism for this reaction is not well understood (Baergen and Donaldson, 2013), and the efficacy of pNO<sub>3</sub><sup>-</sup> photolysis likely depends on aerosol composition (Ma et al., 2021). Other studies have estimated photolysis frequencies an order of magnitude or more lower in non-biomass burning environments (Romer et al., 2018; Shi et al., 2021). Particulate nitrate concentrations are constrained by aerosol mass spectrometer (AMS) observations (Fig. S6a). Comparison with other aerosol composition observations suggests minor pNO<sub>3</sub><sup>-</sup> contained in coarse mode aerosol excluded by the AMS (Fig. S7). It is not possible to partition AMS pNO<sub>3</sub><sup>-</sup> between organic and inorganic forms for SEAC<sup>4</sup>RS (Ulbrich et al., 2009; Day et al., 2021); however, the nature of particulate nitrate participating in reaction (R1) is also unclear. The model treats pNO<sub>3</sub><sup>-</sup> as non-reactive (no chemical production or loss), but it dilutes with a puff-dependent background concentration, chosen such that concentrations match observations at the start and end of each puff. This is a workaround to ensure the model carries reasonable pNO<sub>3</sub><sup>-</sup> concentrations throughout the simulation in the absence of explicit model aerosol chemistry. Simulations M3a, b, and c scale the nominal rate of reaction (R1) by factors of 0.5, 1, and 2, respectively.

Simulations M4a and b incorporate reaction of NO<sub>2</sub> on aerosol surfaces:



$$k_a = 0.25v_{\text{NO}_2}S_a\gamma \quad (3)$$

Here,  $k_a$  is the first-order rate coefficient,  $v_{\text{NO}_2}$  is the mean molecular speed of NO<sub>2</sub>,  $S_a$  is particle surface area density, and  $\gamma$  is the reactive uptake coefficient. We assume that gas diffusion is not a limiting factor for small values of  $\gamma$  ( $10^{-3}$  to  $10^{-6}$ ). For  $S_a$ , we use Laser Aerosol Spectrometer observations (Yu et al., 2016b) scaled up by a factor of 1.7 to account for calibration bias (Fig. S7) and linearly interpolated over plume age. Values of  $S_a$  range from 800 to 6600  $\mu\text{m}^2 \text{cm}^{-3}$  (Fig. S6b). Parameterization of  $\gamma$  follows a laboratory-derived relationship with solar radiation (Stemmler et al., 2006; Zhang et al., 2019):

$$\gamma = \begin{cases} \alpha J_{\text{NO}_2}, & \text{Rad} \leq 400 \text{ W m}^{-2} \\ \alpha J_{\text{NO}_2}(\text{Rad}/400)^2, & \text{Rad} > 400 \text{ W m}^{-2} \\ 10^{-6}, & \text{minimum} \end{cases} \quad (4)$$

Here,  $J_{\text{NO}_2}$  is the NO<sub>2</sub> photolysis frequency,  $\alpha = 2.5 \times 10^{-4}$  is a scaling factor, and  $\text{Rad}$  is total solar irradiance. The latter is estimated using trajectory-dependent SZA and the linear relationship between SZA and observed solar irradiance (Fig. S8).

The minimum value of  $\alpha$  follows Aumont et al. (2003). Figures S6c-d show calculated  $\alpha$  and  $k_a$  at the end point of each puff, with values in the range of  $4 - 11 \times 10^{-6}$  and  $0.3 - 4.9 \times 10^{-6} \text{ s}^{-1}$ , respectively. Simulation M4a uses this default parameterization, while in simulation M4b  $k_a$  is multiplied by a factor of 1000. This range spans observed uptake on humic acid ( $\alpha = 2 - 8 \times 10^{-5}$  (Stemmler et al., 2006)) and soot ( $\alpha = 3.7 - 11 \times 10^{-3}$  (Ammann et al., 1998)) surfaces.

We do not consider ground surface HONO sources (Chai et al., 2021), as plume sampling is predominantly in the free troposphere and uppermost mixed layer.

### 2.4.3 Parameter Optimization

In a final set of simulations, we simultaneously tune initial unmeasured VOC, initial HONO, and  $\text{pNO}_3^-$  photolysis. For these runs, default values for initial unmeasured VOC concentrations (M1), initial HONO (M2c), and particulate nitrate photolysis rate (M3b) are each independently scaled by factors of 0, 0.25, 0.5, 0.75, and 1. Iterating over all combinations yield 125 simulations. Results from these simulations are analyzed in terms of normalized mean bias (NMB) (Gustafson and Yu, 2012) relative to observed NEMR age profiles.

## 3. Results

### 3.1 Observations and Base Simulation

We first examine the observed evolution of trace gases in the Rim Fire. Comparison to base simulation output (M0) benchmarks the representation of this chemistry with MCMv3.3.1 and available constraints.

#### 3.1.1 VOC

The downwind evolution of primary VOC illustrates the differing effects of emissions, dilution, and oxidation. We group VOC into four categories: long-lived alkanes, aromatics/intermediate-lived alkenes, short-lived alkenes, and biogenic terpenes. Groups reflect photochemical lifetimes and similarities in model NMB. Figures 2a-d show an example from each category, Fig. S9 shows the time series of all observed VOC, and Fig. S10 shows the normalized mean bias for each VOC calculated over the full simulation.

Long-lived alkanes such as ethane, propane, and butanes do not decay monotonically with plume age (Figs. 2a and S9a-d). NEMRs for these gases peak downwind of the Rim Fire; for example, the highest propane NEMR occurs at an age of 4 h and is 50% higher than the initial value. All model simulations exhibit negative NMB of -26% or less for these gases (Fig. S10). With lifetimes of days or more against OH oxidation, these gases are relatively sensitive to variability in background levels and regional emissions (e.g., anthropogenic or oil/natural gas activities). For example, at an age of 4 h the ethane enhancement is 4 times its background value, whereas the ethene enhancement is a factor of 70. Variability in fire emissions is a less likely explanation for model-measurement mismatch. Reconciliation with the observed increase in NEMRs at an age of 2 h would require a VOC-to-CO emission ratio increases of 50 - 100%, and we would expect similar changes in other VOC

emission ratios (Liu et al., 2017; Permar et al., 2021) that are not observed. These VOC are minor contributors to plume photochemistry; however, this comparison underscores the challenge of accounting for background variability in Lagrangian or pseudo-Lagrangian simulations.

300           The second group, including most aromatics (benzene, toluene, ethyl benzene and xylenes) and several alkenes (ethene and propadiene), exhibits stronger NEMR decays with age (Figs. 2b and S9e–k). Photochemical lifetimes for these gases range from 11 – 48 hours (except for benzene, with a lifetime of 4 – 10 days), thus we expect both emissions and chemistry to influence NEMR variability. Similar to alkanes, longer-lived aromatics display enhancements at 2 – 6 hours of age. Xylenes, ethene and propadiene exhibit a more monotonic decay. The base model NMB is below 8% for the latter three  
305 compounds.

          Oxidation controls the trend of short-lived alkenes, including propene, butenes, butadiene, pent-1-ene, and furan (Figs. 2c and S9l–s). These gases have near-zero background mixing ratios and react primarily with OH. Loss via O<sub>3</sub> reaction is < 15% of the total loss rate, and nitrate radical reaction is negligible within these air masses. NEMRs decay rapidly, and the most reactive gases fall below measurement detection limits at later ages. The base model simulation exaggerates this decay,  
310 resulting in negative NMB as large as -233% for furan.

          Terpenes (isoprene and  $\alpha$ -pinene) are also highly reactive, and emissions of these compounds are lost rapidly (Figs. 2d and S9t,u). Around an age of 2 h, however, NEMRs increase beyond those in the near-source sample. A similar pattern appears in the isoprene oxidation products methyl vinyl ketone (MVK) and methacrolein (MACR) (Fig. S12d). Furthermore, this particular sample is uncharacteristically dilute relative to the Lagrangian age (Fig. S2). This evidence  
315 suggests that local biogenic emissions mixed with the Rim Fire smoke samples at Lagrangian ages of 2 – 4 hours.

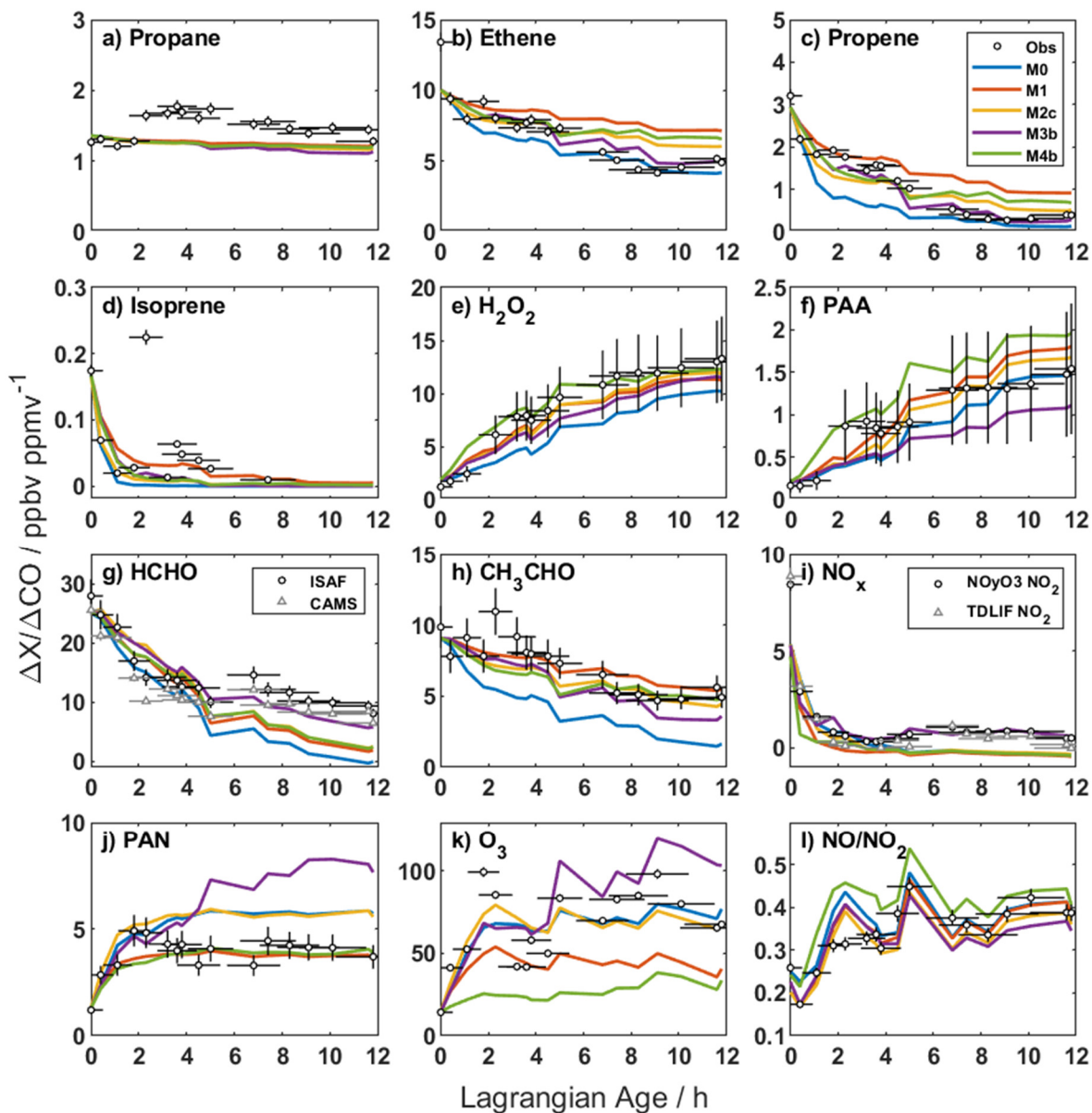


Figure 2. Age evolution of NEMRs for reactive gases (a-k) and the NO/NO<sub>2</sub> ratio (l). Black circles are observations with their corresponding uncertainty due to measurement accuracy and age. Colored lines are model output from the base simulation (M0, blue), addition of unmeasured VOC (M1, red), and addition of unmeasured VOC plus initial HONO (M2c, yellow), pNO<sub>3</sub><sup>-</sup> photolysis (M3b, purple), or NO<sub>2</sub> heterogeneous uptake (M4b, green). In (g) and (i), circles and triangles represent two independent measurements of HCHO and NO<sub>2</sub>, respectively (Table S1). NO and NO<sub>2</sub> observations in (l) are from the NOyO<sub>3</sub> instrument.

### 3.1.2 HO<sub>x</sub> and Peroxides

The SEAC<sup>4</sup>RS DC-8 payload did not include observations of HO<sub>x</sub> (= OH + HO<sub>2</sub>). The loss and production of other compounds, however, indirectly constrains HO<sub>x</sub> abundance. The decay rate of short-lived alkenes provides a benchmark for OH, while peroxide production is an indicator for HO<sub>2</sub> and, to some extent, RO<sub>2</sub>.

320 As discussed above, short-lived alkene NEMRs decay faster than observed NEMRs in the base simulation, especially in the first few hours. The discrepancy between modelled and observed short-lived alkenes suggests over-prediction of OH in the young plume. Base model OH starts at  $4.9 \times 10^6 \text{ cm}^{-3}$  and declines to  $\sim 1.2 \times 10^6 \text{ cm}^{-3}$  after  $\sim 4$  hours (Fig. S11a).

325 Self-reaction of HO<sub>2</sub> produces hydrogen peroxide (H<sub>2</sub>O<sub>2</sub>), and reaction of HO<sub>2</sub> with peroxyacetyl radical (PA) produces peroxyacetic acid (PAA). NEMRs for both peroxides start low and increase with plume age (Fig. 2e,f). Initial mixing ratios are near background levels, at odds with a previous study suggesting significant primary emissions (Lee et al., 1997). The base simulation exhibits an upward trend but with a slower growth rate, especially over the first 4 hours. Model H<sub>2</sub>O<sub>2</sub> NEMRs skirt the lower edge of measurement accuracy, while PAA agreement generally improves with age. This comparison suggests under-prediction of HO<sub>2</sub> and possibly RO<sub>2</sub> at early ages. Base model HO<sub>2</sub> peaks at  $\sim 85$  pptv and declines to values as low as 23 pptv (Fig. S11b).

### 330 3.1.3 Oxygenated VOC

Formaldehyde (HCHO), as an oxidation product of numerous VOC, serves as a top-down constraint on in-plume VOC processing. The observed HCHO NEMR decreases by a factor of three in the first 5 hours before levelling out at later ages (Fig. 2g). This behavior is consistent with loss of primary emissions alongside secondary production (Liao et al., 2021). The HCHO lifetime due to photolysis and OH oxidation is  $\sim 3.5$  h at the time of sampling, but accumulation of multi-generation oxidation products may sustain downwind HCHO production (Alvarado et al., 2020). The sharp rise in the HCHO NEMR between 5 and 7 hours coincides with descent to a lower sampling altitude (Fig. S1). The base simulation follows observations early on but overestimates the decay at intermediate and later ages.

340 The acetaldehyde (CH<sub>3</sub>CHO) NEMR decay exhibits a more constant slope (Fig. 2h). The small post-emission peak in the acetaldehyde NEMR at an age of 2 h coincides with the aforementioned sharper peaks in biogenic markers and enhancements in long-lived alkanes, which may indicate influence from non-fire surface emissions. Base model NEMRs initially decay more rapidly than observed, but the slope matches observations after the first few hours. This discrepancy is consistent with over-prediction of OH in the young plume, as OH oxidation accounts for 80% of acetaldehyde loss. Conversely, OH only accounts for 15 – 50% of HCHO loss (photolysis is the remainder). Thus, as will be evident in sensitivity simulations discussed later, HCHO and CH<sub>3</sub>CHO exhibit opposite responses to varying OH.

345 Text S2 and Figure S12 present several other observed oxygenated VOC (oVOC), including methanol, acetone + propanal, hydroxyacetone, and MVK + MACR.

### 3.1.4 Reactive Nitrogen

Major wildfire reactive nitrogen emissions include  $\text{NO}_x$ , HONO, and  $\text{NH}_3$  (Lindaas et al., 2020; Roberts et al., 2020). Observations of the latter two are not available for SEAC<sup>4</sup>RS, so here we focus on  $\text{NO}_x$  and observed reservoir species (collectively,  $\text{NO}_y$ ).

The base simulation reasonably reproduces the initial loss of  $\text{NO}_x$  but does not capture later behavior (Fig. 2i). In the first few hours, the  $\text{NO}_x$  NEMR decays with an e-folding timescale of 20 – 40 minutes. After reaching a minimum at  $\sim 3.6$  h, the  $\text{NO}_x$  NEMR rises to a sustained enhancement at  $\sim 10\%$  of the initial value. An additional  $\text{NO}_x$  source of  $200 \text{ pptv h}^{-1}$  is required to close the  $\text{NO}_x$  budget in the period between 6 and 12 h. This discrepancy is explored further in Sect. 3.3.

Peroxyacetyl nitrate (PAN) typically comprises the majority of total peroxy nitrates (Wooldridge et al., 2010) and is produced via the reversible reaction of  $\text{NO}_2$  with PA. The PAN NEMR rises rapidly in the first two hours before stabilizing (Fig. 2j), comparable to previous observations (Alvarado et al., 2010, 2015). The base simulation captures the early rise in PAN but overshoots the asymptote by  $\sim 20\%$ . SI Text S3 and Figures S13-S14 compare model output to measurements of other speciated peroxy nitrates, total peroxy nitrates, and alkyl nitrates, which show varying levels of model-measurement agreement.

Reaction of OH with  $\text{NO}_2$  primarily forms nitric acid ( $\text{HNO}_3$ ). The observed  $\text{HNO}_3$  NEMR is negative and trends downward with age, indicating that in-plume  $\text{HNO}_3$  is below background levels (Fig. S14c). Excess ammonia/ammonium (Perring et al., 2017) likely drives rapid formation of particulate ammonium nitrate (Lindaas et al., 2021). The puff model lacks aerosol chemistry and thus predicts net growth of the  $\text{HNO}_3$  NEMR. The difference between the base model and observed  $\text{HNO}_3$  NEMR implies an effective  $\text{HNO}_3$  lifetime of less than 1 hour.

Figure 3 shows the evolution of the sum of observed  $\text{NO}_y$ . We refer to this quantity as  $\square \text{NO}_{y,\text{obs}}$  to acknowledge missing observations of some  $\text{NO}_y$  species such as HONO,  $\text{HO}_2\text{NO}_2$ , and nitroaromatics. The presence of significant PNs within the first sample indicates some chemical aging, as expected for a mixture of smoke ages. After the first hour, PNs comprise more than half of  $\square \text{NO}_{y,\text{obs}}$ ,  $\text{pNO}_3^-$  comprises another 20 – 30%, and  $\text{NO}_x$  and alkyl nitrates contribute 10 – 15% each. The  $\square \text{NO}_{y,\text{obs}}$  NEMR decreases from 18 to  $6.3 \text{ ppbv ppmv}^{-1}$  over 12 hours. In the absence of emission changes and deposition, we expect conservation of the total  $\text{NO}_y$  NEMR (Juncosa Calahorrano et al., 2020). It seems unlikely that unmeasured  $\text{NO}_y$  can explain a 2.6-fold decrease over this period (Roberts et al., 2020). Decreasing  $\square \text{NO}_{y,\text{obs}}$  may indicate that  $\text{NO}_x$  emissions are changing over time or that we are sampling different parts of the Rim Fire with differing  $\text{NO}_x$  emission ratios. As discussed in Sect. 2.3, evidence regarding emission ratio trends over the sample period is inconclusive. Furthermore, a model simulation assuming declining  $\text{NO}_x$  emissions with age would further degrade model prediction of observed  $\text{NO}_x$  at later ages. For comparison, modeled  $\square \text{NO}_{y,\text{obs}}$  (which excludes  $\text{pNO}_3^-$ ) is  $\sim 7 \text{ ppbv ppmv}^{-1}$  throughout simulation M0.

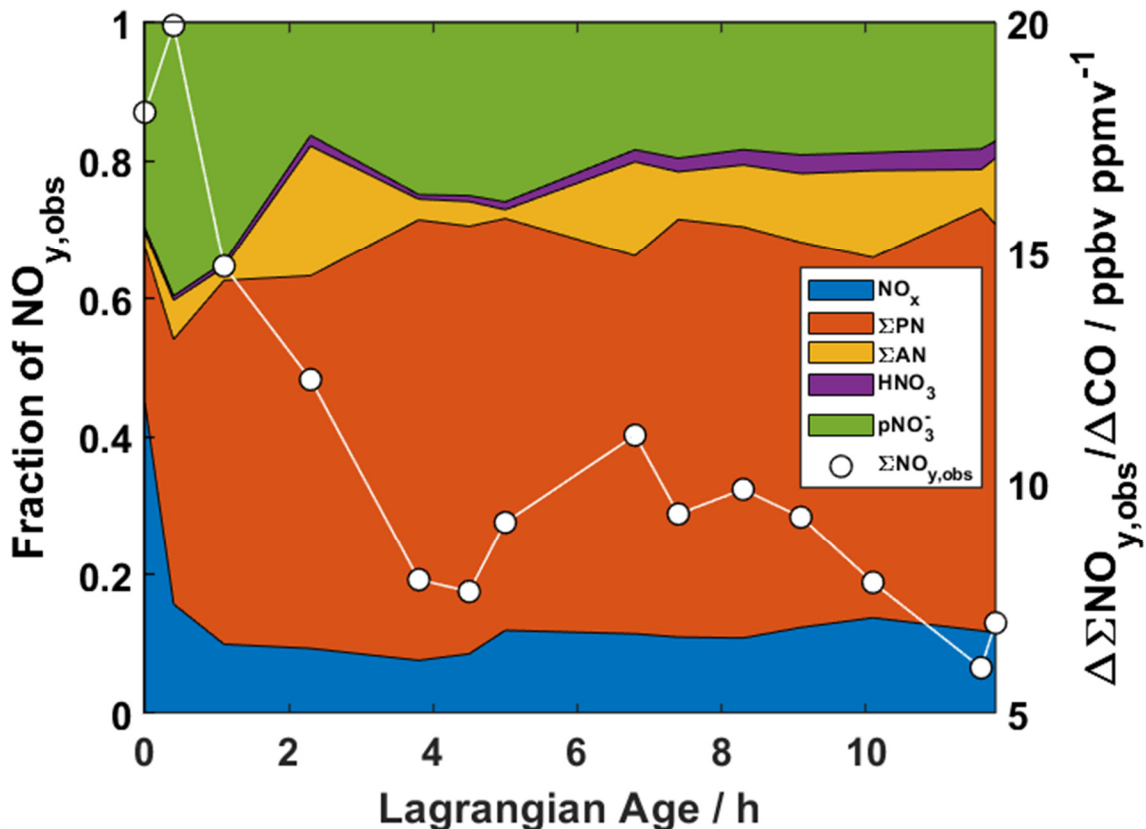


Figure 3. Age evolution of observed NO<sub>y</sub> speciation (solid colors, left axis) and the  $\square$ NO<sub>y,obs</sub> NEMR (white circles/line, right axis).  $\square$ NO<sub>y,obs</sub> is the bottom-up sum of observed NO<sub>y</sub>. ΣPN represents TDLIF observations.

### 3.1.5 Ozone

The observed O<sub>3</sub> NEMR grows rapidly within the first few hours before slowing down (Fig. 2k). The peak value at 2 h coincides with the aforementioned peak in biogenic VOC. The base simulation reproduces the general trend of the observed O<sub>3</sub> NEMR time profile but misses the maximum and under-predicts after 5 h. Significant point-to-point variability in the observed O<sub>3</sub> NEMR mostly reflects variability in the CO excess mixing ratio rather than O<sub>3</sub> itself, as O<sub>3</sub> is relatively close to its background value. Absolute O<sub>3</sub> reaches a peak of 130 ppbv within the first hour before diluting rapidly to 65 – 75 ppbv, remaining above the background estimate of 50 ppbv (Fig. S15). Mean absolute bias for base simulation O<sub>3</sub> is -3.2 ppbv averaged over all observations and +0.5 ppbv for ages greater than 2 h.

The ratio of NO to NO<sub>2</sub> relates closely to radical turnover and O<sub>3</sub> production. In the absence of strong NO<sub>x</sub> sources or sinks, photolysis of NO<sub>2</sub> and oxidation of NO establishes a photostationary state that controls this ratio:

$$\frac{[NO]}{[NO_2]} = \frac{J_{NO_2}}{k_{NO+O_3}[O_3] + k_{NO+HO_2}[HO_2] + \sum k_{NO+RO_2}[RO_2]} \quad (5)$$

Here,  $J_{NO_2}$  is the  $NO_2$  photolysis frequency and  $k_{X+Y}$  are reaction rate coefficients. The observed  $NO/NO_2$  ratio doubles over the course 12 hours (Fig. 2l), consistent with the decline of peroxy radicals and ozone mixing ratios.  $J(NO_2)$  does not exhibit a clear trend over this period (Fig. A1). The base simulation over-predicts this ratio in the first few hours and agrees within uncertainties afterward.

### 3.2 Accounting for Unmeasured VOC

Differences between observations and the base simulation are consistent with missing reactive VOC in the model. The decay of reactive alkenes is faster than observed, suggesting that model OH is too high. Peroxide and oVOC production is too slow, indicating missing sources of peroxy radicals and organic carbon. Simulation M1 approximates the effects of unmeasured VOC, incorporated following the procedures outlined in Sect. 2.4.1.

Total OH reactivity – the inverse of the OH lifetime – increases significantly upon addition of unmeasured VOC (Fig. 4a). Initial OH reactivity grows from 77 to 182  $s^{-1}$ . The top five components of OH reactivity in the observations and simulation M0 are HCHO, CO,  $CH_3CHO$ , furan, and propene (Fig. 4c). Species included in simulation M0 comprise ~45% of the OH reactivity in simulation M1. The top 5 additional contributors in simulation M1 are aromatics, consistent with Coggon et al. (2019). Enhancements persist as the plume ages, with OH reactivities of 7  $s^{-1}$  (M1) versus 3  $s^{-1}$  (M0) at an age of 12 h. After 12 h, 32% of M1-simulated OH reactivity is comprised of nearly 2200 species, mostly oxygenated VOC.

Normalizing for dilution reveals a more modest decline in OH reactivity due to photochemistry alone (Fig. 4b). Normalized excess OH reactivity declines by 61% and 45% for M0 and M1, respectively. The relatively slower decline in M1 reflects less OH in this simulation. Modelled declines in OH reactivity over 12 h of aging imply a pseudo-first order lifetime of 12.7 h and 21.4 h for OH reactivity in simulations M0 and M1, respectively. OH reactivity estimates from the model may be an upper limit, as some reactive carbon will partition to the aerosol phase (Palm et al., 2020).

Additional VOC reactivity markedly alters downwind chemistry. OH decreases to  $5 \times 10^5 \text{ cm}^{-3}$  for most of the simulation (Fig. S11a), reducing the decay of reactive alkenes to a shallower slope than observed (Fig. 2b–d). Maximum  $HO_2$  increases by 35% (Fig. S11b), and  $H_2O_2$  agrees better with observations (Fig. 2e). PAA increases due to enhanced production of  $HO_2$  and PA (Fig. 2f). Peroxyacetyl radical is also a PAN precursor, but PAN actually decreases in simulation M1 due to lower  $NO_2$  (Fig. 2j) and competing formation of larger peroxy nitrates (Fig. S13, S14). Rapid sequestration of  $NO_x$  also suppresses ozone formation in the young plume (Fig. 2k). HCHO and  $CH_3CHO$  increase (Fig. 2g–h), but the model still under-predicts HCHO at later ages.



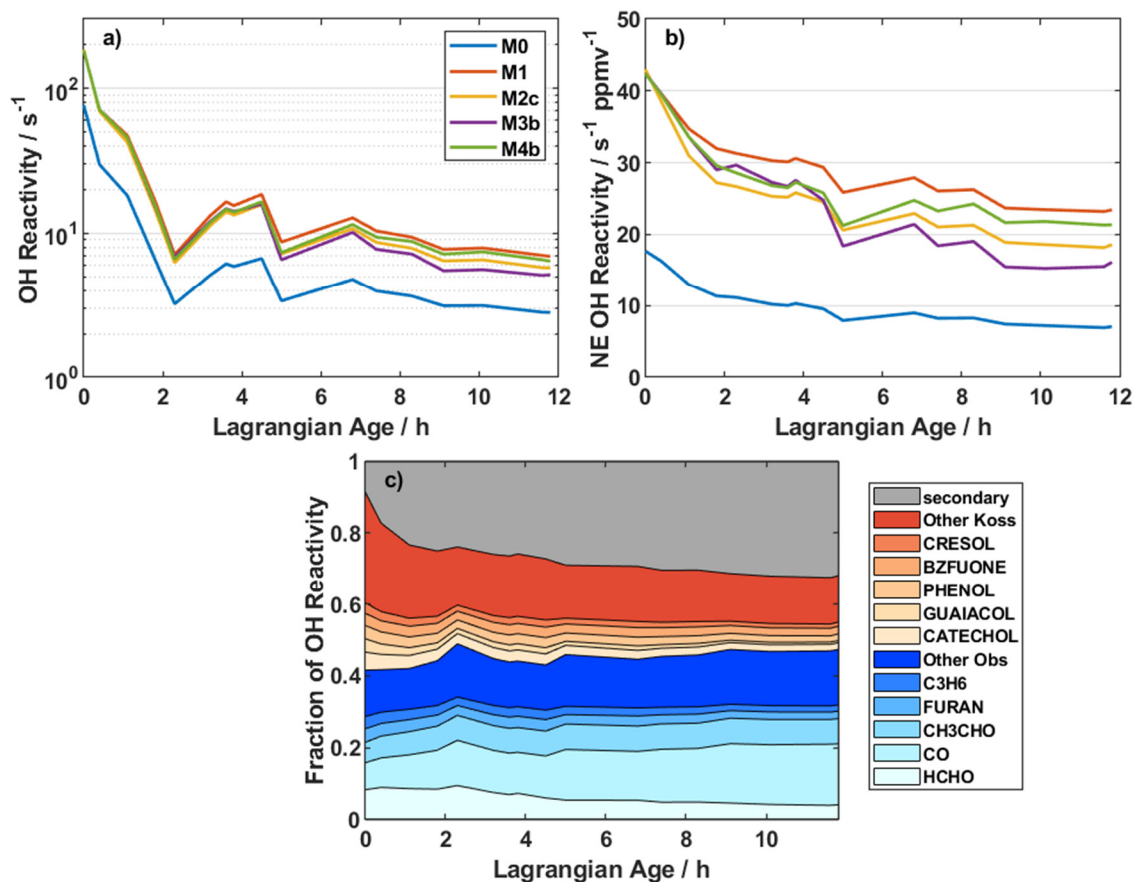


Figure 4. (a) Age evolution of total model OH reactivity, (b) normalized excess OH reactivity, and (c) fractional contributions of individual reactants to total OH reactivity in simulation M1. In (b), normalization is analogous to NEMR calculation (Eq. 1), and the OH reactivity background is 1.1 s<sup>-1</sup> based on summation over observed OH reactants in the background sample. In (c), blue shades denote measured compounds, red shades are species added from the Koss et al. (2018) inventory, and grey includes all other MCM species reacting with OH.

### 3.3. Accounting for Primary and Secondary HONO

The contrast of observations and simulation M1 in Fig. 2 implies insufficient OH and NO<sub>x</sub>. The decay of short-lived alkenes is now slower than observed owing to the large increase in OH reactivity, while the decay of NO<sub>x</sub> is faster than observed due to rapid peroxy nitrate formation. Both low OH and low NO<sub>x</sub> contribute to under-prediction of O<sub>3</sub>.

420 Potential explanations for the model-measurement discrepancy in NO<sub>x</sub> include NO<sub>2</sub> measurement artefacts, emissions variability, unaccounted-for recycling via NO<sub>x</sub> reservoirs, and missing NO<sub>x</sub> sources. NO<sub>2</sub> measurement artefacts are unlikely to play a significant role, given the excellent correlation between two independent NO<sub>2</sub> measurements and the magnitude of model-measurement disagreement (Text S3). Doubling initial NO<sub>x</sub> has a minor effect on the simulated NO<sub>x</sub> NEMR after ~5 h (Fig. S16), and total observed NO<sub>y</sub> decreases with Lagrangian age (Fig. 3); thus, NO<sub>x</sub> emissions variability also cannot explain

425 this difference. Recycling via decomposition or oxidation of  $\text{NO}_x$  reservoirs, such as organic nitrates, should be adequately  
captured by simulation M1 as the NEMRs of total peroxy nitrates and alkyl nitrates are simulated well (Fig. S13 and S14). We  
lack observations of  $\text{HO}_2\text{NO}_2$  or  $\text{CH}_3\text{O}_2\text{NO}_2$ , but lifetimes for these gases are a few minutes for our conditions and modelled  
mixing ratios are less than 10 pptv after 5h. Nitroaromatic photolysis may generate HONO (Sangwan and Zhu, 2016), but  
430 these compounds do not build up to sufficient levels to act as a major  $\text{NO}_x$  source in the model. Therefore, a missing  $\text{NO}_x$   
source is the most likely explanation.

Previous work has demonstrated the important role of HONO in smoke plume chemistry (Alvarado et al., 2015;  
Alvarado and Prinn, 2009; Peng et al., 2020; Theys et al., 2020). Direct HONO emissions amplify radical production in the  
nascent plume, while secondary HONO formation may sustain chemistry as the plume ages. As discussed in Sect. 2.4.1 and  
summarized in Table 1, we implemented both primary emissions (simulations M2a,b,c) and secondary production via  $\text{pNO}_3^-$   
435 photolysis (M3a,b,c) or heterogeneous  $\text{NO}_2$  reaction (M4a,b). Figure S17 shows the HONO NEMR and absolute mixing ratio  
for several representative simulations. The HONO photolysis lifetime, based on observed photolysis frequencies, is 10 – 20  
minutes. HONO mixing ratios in simulations with secondary production (M3b and M4b) are 10's of pptv after a few hours.  
Most figures in the main text and supplement display results from simulations with high primary HONO (M2c), moderate  
 $\text{pNO}_3^-$  photolysis (M3b) and fast  $\text{NO}_2$  heterogeneous uptake (M4b), while Figures S18 – S20 show full results for each  
440 sensitivity series.

Initial HONO stimulates chemistry in the first several hours. Simulation M2c starts with 25 ppbv of HONO,  
comparable to levels observed in other Western U.S. wild fires (Peng et al., 2020). Photolysis of HONO increases OH and NO  
production, leading to faster VOC decay and product formation (Fig. 2 and S11). Intensification is limited to the first few  
hours, and the decay of short-lived alkenes, like propene and butenes, is too slow at later ages (Fig. 2c).  $\text{NO}_x$  agreement also  
445 improves at early times, but under-prediction persists after 3 h of aging (Fig. 2i). PAN and  $\text{O}_3$  profiles are close to those in the  
base simulation (Fig. 2j-k).

Photolysis of  $\text{pNO}_3^-$  leads to more sustained impacts. Simulation M3b uses the literature-derived rate for reaction R1  
(Ye et al., 2017). For this case, median OH increases by a factor of 3.5 relative to simulation M1 (Fig. S11), improving model  
agreement with short-lived VOC (Fig. 2b-c).  $\text{HO}_2$  is mostly unchanged, reflecting a counterbalance of faster production via  
450 VOC oxidation and faster loss via reaction with NO. Unique to the mechanisms tested in this study,  $\text{pNO}_3^-$  photolysis  
reproduces the enhancements in HCHO and  $\text{NO}_x$  NEMRs observed at later ages (Fig. 2g, 2i). On the other hand, PAN and  
ozone are now over-predicted (Fig. 2j, 2l), as are  $\square\text{PN}$  and some speciated PNs (Fig. S13, S14). Halving the  $\text{pNO}_3^-$  photolysis  
rate improves agreement with PAN and  $\text{O}_3$  at the expense of  $\text{NO}_x$  and HCHO (Fig. S19, compare simulations M3a and M3b).  
Over-prediction of PAN and its analogues in these simulations may reflect errors in the VOC distribution and/or chemical  
455 kinetics (Sect. 4.2). These simplified simulations also do not capture age-dependent variability in aerosol composition (e.g.,  
organic versus inorganic  $\text{pNO}_3^-$ ), which could influence the effective  $\text{pNO}_3^-$  photolysis rate and product yields.

In contrast to the above, HONO production via heterogeneous reaction of  $\text{NO}_2$  generally degrades agreement with  
observations (Fig. S19). Results from simulation M4a are nearly identical to those from M1, while in simulation M4b ( $\square$

multiplied by 1000) ozone and  $\text{NO}_x$  under-prediction worsens. With rapid HONO photolysis, this process effectively converts  
460  $\text{NO}_2$  to NO while producing OH. This increases the loss of  $\text{O}_3$  to reaction with NO (reflected in the NO/ $\text{NO}_2$  ratio), reducing  
net  $\text{O}_3$  production.  $\text{NO}_2$  conversion to HONO is nearly  $\text{NO}_x$ -neutral, whereas  $\text{pNO}_3^-$  photolysis is effectively a  $\text{NO}_x$  source.

### 3.4 Co-optimization of VOC and HONO

Additional reactive VOC and HONO chemistry act together to improve model-measurement agreement for most species.  
HONO sensitivity simulations presented in Sect. 3.3 utilize the same initial VOC as simulation M1. Uncertainties in initial  
465 VOC concentrations stem from the age distribution and history of sampled smoke, adaptation of laboratory-derived emission  
factors to ambient conditions, potential errors in the emission factors themselves (SI Text S1), and translation of unmeasured  
VOC to MCM species. Uncertainties in primary and secondary HONO are also significant, as we lack a HONO measurement  
for comparison and putative aerosol mechanisms are not well understood.

To illustrate the co-dependence of VOC and HONO, we simultaneously scale initial unmeasured VOC, initial HONO,  
470 and  $\text{pNO}_3^-$  photolysis as described in Sect. 2.4.3. We ignore  $\text{NO}_2$  heterogeneous uptake here and focus analysis on  $\text{O}_3$ . For  
each of the 125 simulations, we first calculate the NMB of the  $\text{O}_3$  NEMR. Consistent with the examples presented above, this  
bias trends negative with increasing VOC and positive with increasing HONO (Fig. S21a-e). Multiple scaling combinations  
produce a net  $\text{O}_3$  NEMR NMB near zero (Fig. 5). In general, increasing initial unmeasured VOC concentrations necessitates  
higher HONO to maintain agreement with observed  $\text{O}_3$ . Initial and secondary HONO similarly influence the overall  $\text{O}_3$  NEMR  
475 NMB, though this simplified metric masks age-dependent differences (Fig. 5 inset).

Figure S22 shows similar minimum NMB isopleths for  $\text{NO}_x$  and PAN NEMRs. Overall behavior is relatively similar.  
Compared to  $\text{O}_3$ , higher HONO is required to close the  $\text{NO}_x$  budget. PAN bias is minimized at higher VOC, but only with low  
initial HONO and relatively slow  $\text{pNO}_3^-$  photolysis. No combination of scaling factors optimizes agreement among all  
observations.

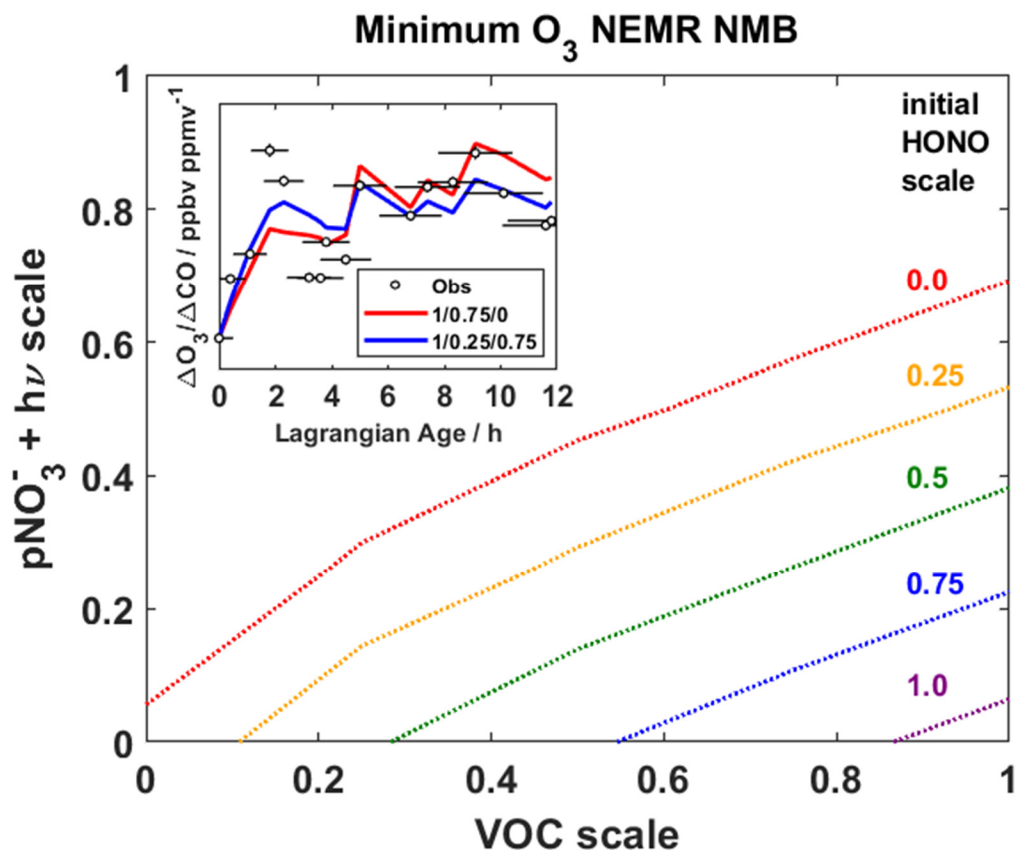


Figure 5. Isopleths for zero values of the  $O_3$  NEMR NMB. Each colored dotted line represents a fixed scaling factor for initial HONO mixing ratios. The x-y coordinates for a point on a given line represent a combination of VOC and  $pNO_3^- + hV$  photolysis scaling factors that minimize the  $O_3$  NEMR NMB. Isopleths are based on interpolation of results from the optimization simulations (Fig. S21). The inset compares observed  $O_3$  NEMRs with those from two simulations with near-zero NMB (red: VOC scale = 1,  $pNO_3^- + hV$  scale = 0.75, initial HONO scale = 0; purple: VOC scale = 1,  $pNO_3^- + hV$  scale = 0.25, initial HONO scale = 0.75).

## 480 4 Discussion

### 4.1 Radical Production and Fate

Inclusion of additional reactive VOC and HONO significantly accelerates radical throughput. Figures 6a-b summarize rates of key  $RO_x$  ( $OH + HO_2 + RO_2 + RO$ ) production and loss pathways integrated over the first 2.3 hours of plume aging, where NEMRs change most rapidly. Compared to the base simulation, total  $RO_x$  initiation and termination doubles in simulation M1 (VOC addition) and more than triples in M2c (initial HONO). As illustrated in Sect. 3, these changes influence the lifetime of reactive gases and the production of secondary compounds such as ozone, peroxides, organic nitrates, and oVOC. Comparison of  $O_3$  NEMRs in Fig. 2k, however, also demonstrates simulation of an ozone profile that reasonably matches observations

even with missing processes. Constraints on other aspects of the chemical system facilitate holistic evaluation of additional chemistry.

490            Photolysis sources dominate radical initiation (Fig. 6a). Photolysis of  $O_3$  to  $O^1D$  and  $H_2O_2$  are each less than 2% of total production in all simulations. In the base simulation, photolysis of HCHO and other oVOC comprises 75% of the radical source. Enhanced reactive VOC in simulation M1 doubles the initiation rate, mainly via further oVOC photolysis and alkene ozonolysis. The largest contributors in the former category are glyoxal (19%) and methyl glyoxal (15%). HONO photolysis in  
495 the plotted simulations, with HONO photolysis comprising 33% of initiation. The relative contribution of HONO is smaller here than in other recent studies (Peng et al., 2020; Robinson et al., 2021) for two reasons. First, we average over the first 2 hours of aging, and the lifetime of HONO is tens of minutes. Second, incorporation of an extended VOC pool greatly enhances oVOC photolysis. There is uncertainty in this contribution due to MCM mapping of VOC (Table S2). Nonetheless, studies failing to account for all oVOC might under-predict radical initiation and over-estimate the relative importance of other radical  
500 sources.

Radical termination includes significant contributions from both  $RO_x-RO_x$  and  $RO_x-NO_x$  reactions (Fig. 6b). Formation of peroxides comprises most of the former group, with equal contributions from  $HO_2 + HO_2$  and  $RO_2 + HO_2$  in simulation M0. Addition of reactive VOC in M1 doubles the rate of  $HO_2 + HO_2$  and triples the rate of  $RO_2 + HO_2$ . PN formation comprises 65 – 86% of  $RO_x-NO_x$  termination, with larger contributions at higher VOC and higher initial HONO. Pernitric acid  
505 formation ( $HO_2 + NO_2$ ) is 18% of  $RO_x-NO_x$  termination in simulation S2c, reflecting fast formation in the concentrated plume followed by rapid dilution (i.e., dilution outpaces thermal decomposition). Contributions of nitric acid ( $OH + NO_2$ ) and nitroaromatic ( $ArNO_2$ ) formation are 3 – 15% and 0.1 – 6% of  $RO_x-NO_x$  termination, respectively. The “OH + X” group includes reaction of OH with HONO and PNs to form  $NO_2$  and other products. These reactions remove  $RO_x$ , but they are not strictly radical termination.

510            The balance of  $RO_x$  loss via  $NO_x$  and  $RO_x$  sinks, typically represented as the ratio of  $NO_x$ -related loss ( $L_N$ ) to total radical loss or production ( $Q$ ), quantifies the sensitivity of ozone production to VOC and  $NO_x$  availability (Kleinman et al., 1997; Kleinman, 2005). Here we define  $L_N$  as the sum of losses via formation of ANs, PNs, nitroaromatics, nitric acid, and pernitric acid (blue-shaded reactions in Fig. 6b). Traditionally, the assumption that  $HNO_3$  formation is the dominant  $NO_x$  sink implies a transition from VOC to  $NO_x$ -sensitive  $O_3$  production at  $L_N/Q = 0.5$  (Kleinman et al., 1997). Significant organic nitrate  
515 formation alters this threshold (Robinson et al., 2021; Schroeder et al., 2017). We adopt a threshold of  $L_N/Q = 0.41$  based on a recent study of Western U.S. wildfires (Robinson et al., 2021).

Ozone production in the young plume is sensitive to both VOC and  $NO_x$  (Fig. 6c). All plotted simulations start with a ratio around 0.8 (VOC-sensitive) and transition to  $NO_x$ -sensitive within the first hour. This rapid shift is consistent with afternoon aging in other Western U.S. wildfires (Robinson et al., 2021) and reflects sequestration of  $NO_x$  within PNs and other  
520 organic nitrates. Addition of VOC in simulation M1 accelerates this transition (more  $RO_x$ ), while HONO slows the transition

in M2c and M3b. Simulation M4b exhibits the fastest transition, where  $\text{NO}_2$  conversion to HONO amplifies  $\text{RO}_x$  while mitigating PAN formation.

The consequences of secondary chemistry assumptions become more apparent at later ages. Most simulations decay to steady  $L_N/Q$  values of 0.03 – 0.06 after 5 hours of aging. Simulation M3b is the exception, maintaining  $L_N/Q$  around 0.2 due to a sustained  $\text{NO}_x$  source via  $\text{pNO}_3^-$  photolysis. The accuracy of this simulation is dubious due to a lack of HONO observations; however, the contrast between M3b and M4b illustrates divergent effects of HONO production mechanisms on radical chemistry. Such differences can affect sensitivity to downwind perturbations, such as when mixing with urban (high  $\text{NO}_x$ ) or biogenic (high VOC) air.

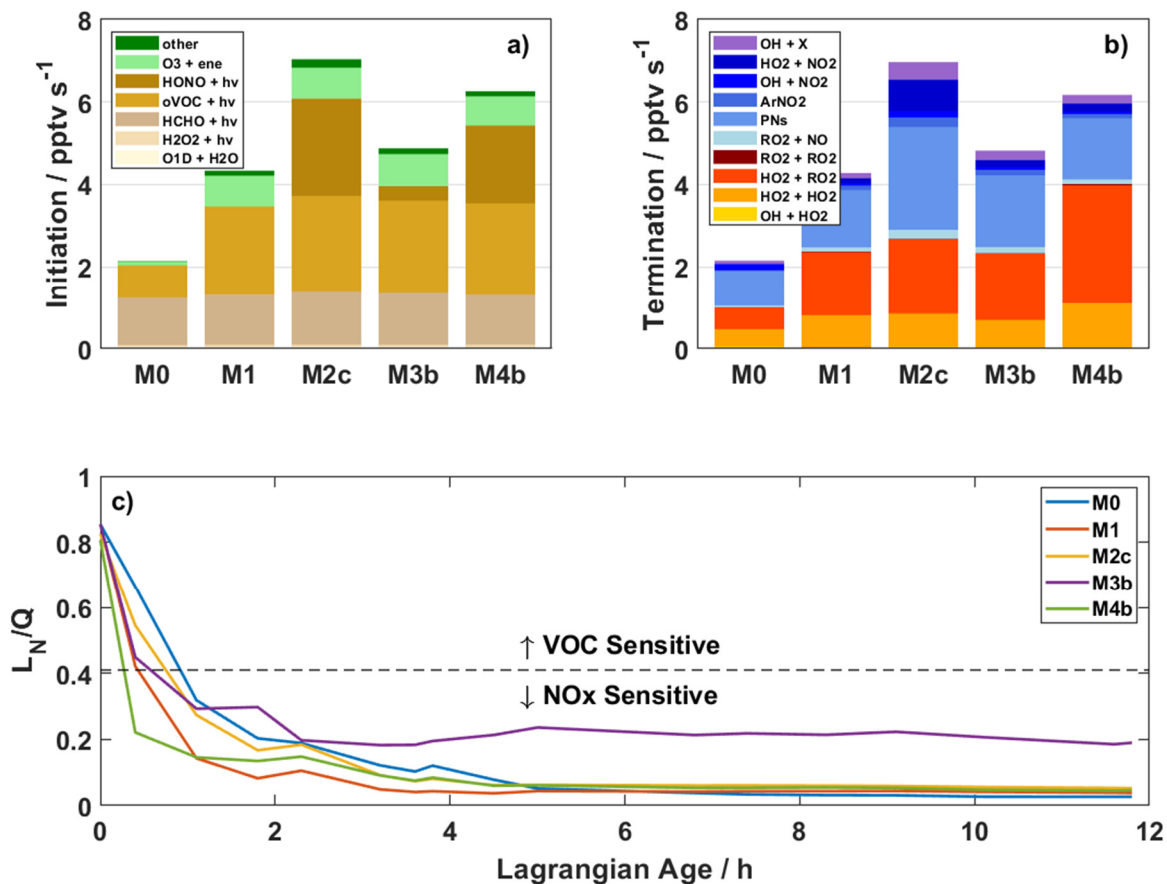


Figure 6. Upper: rates of  $\text{RO}_x$  initiation (a) and termination (b) integrated over the first five data points (Lagrangian age 0 – 2.3 hours) via trapezoidal integration. In (a), “oVOC + hv” includes all oVOC photolysis reactions other than HCHO and “other” includes  $\text{NO}_3 + \text{VOC}$  and other minor reactions. In (b), yellow-red and blue shades represent  $\text{RO}_x\text{-RO}_x$  and  $\text{RO}_x\text{-NO}_x$  reactions, respectively. “ArNO<sub>2</sub>” represents formation of nitroaromatics and “PNs” is net formation of peroxy nitrates. “OH + X” (purple) represents reactions of OH with organic compounds (typically PNs), which result in formation of  $\text{NO}_2$  and other non- $\text{RO}_x$  products. Lower: Fraction of  $\text{RO}_x$  radicals lost via reactions with  $\text{NO}_x$  (c). Model simulations are as described in Fig. 2 and Table 1. The dashed line denotes the approximate transition between  $\text{NO}_x$ -sensitive and VOC-sensitive ozone production for biomass burning chemistry as suggested by Robinson et al. (2021).

## 4.2 Implications for Modelling Biomass Burning Chemistry

530 Box model deficiencies temper the above results and underscore the challenges of simulating smoke plume chemistry. Simulations including both unmeasured reactive VOC and initial HONO or  $\text{pNO}_3^-$  photolysis (M2c or M3b) reasonably match observed NEMRs in the first 2 hours of aging (Fig. 2); however, none of the scenarios explored here simultaneously reconcile  $\text{O}_3$ ,  $\text{NO}_x$ , and PAN throughout the whole sample period. Errors due to assumptions about emissions and air mass history may become more significant further downwind; for example, the model does not capture the large decline in total observed  $\text{NO}_y$  (Fig. 3). Better observations of emission and background variability would reduce uncertainty, although this can be logistically challenging for large smoke plumes like the Rim Fire.

Errors in kinetics or VOC speciation may also contribute to model uncertainties. For example, our mechanism includes updated reaction rate coefficients for  $\text{PA} + \text{HO}_2$  (faster by a factor of 1.33) and  $\text{PAA} + \text{OH}$  (slower by a factor of 123) based on recent laboratory results (Berasategui et al., 2020; Jenkin et al., 2019). Using MCM default values instead reduces model PAA NEMRs by half (results not shown). Analogous changes to other aspects of the mechanism, due to yet-unrecognized systematic errors, could influence other species. Over-prediction of PAN in simulations M2c and M3b may reflect errors in the thermal equilibrium, the  $\text{NO}/\text{NO}_2$  ratio (under-predicted at later ages in simulation M3b), and/or VOC speciation. Regarding the last issue, acetaldehyde oxidation comprises half of PA production in our simulations, while the other half stems from mostly unmeasured precursors like methyl glyoxal (24% of production) (Fig. S23). Observations of major PAN precursors are necessary to close the PAN budget. Observations of PNs with specific precursors also afford complementary information; for example, over-prediction of APAN in most scenarios (Fig. S13) may imply that initial acrolein is too high. Observations of HONO,  $\text{HO}_2\text{NO}_2$ , and total  $\text{NO}_y$  would also help to fully constrain radical sources, cycling, and fate.

For regional and global models, whether and how to account for the full VOC distribution remains an open challenge. Advances in instrumentation have facilitated quantification of myriad reactive gases (Heald and Kroll, 2020). Cumulatively, individually minor species can comprise a significant fraction of OH reactivity and potential organic aerosol mass (Gilman et al., 2015). Condensed mechanisms cannot represent this level of speciation. One option to reduce this complexity, similar to our methodology, is to identify proxy or surrogate species within a given mechanism based on mass, molecular formula, reactivity, volatility, or other metrics. Given the variability of chemistry within smoke plumes, machine learning techniques may also prove useful for condensing VOC into a manageable framework (Kelp et al., 2020). Multifaceted observations can help ensure accurate representation of key chemical processes that influence the spatial and temporal extent of air quality and climate impacts.

Many atmospheric chemistry models do not incorporate primary emissions or heterogeneous production of HONO in standard simulations, including GEOS-Chem (L. Hu, personal communication, 2021). Primary HONO is a key oxidant source in young fire plumes (Peng et al., 2020). The HONO photolysis lifetime of 10 – 20 minutes can be on the order of, or much shorter than, an advection timescale for a single grid cell (e.g., a 12 km grid cell at a wind speed of 10 m/s has an advection timescale of 20 minutes). Thus, “instantaneous dilution” may be a major hurdle to simulating this fast chemistry. Plume

heterogeneity (e.g., a darker core) also confounds efforts to simulate chemistry in an average sense (Wang et al., 2021). Putative secondary HONO sources are manifold (Zhang et al., 2019), as are associated uncertainties. A recent lab study of  $\text{pNO}_3^-$  photolysis on inorganic aerosol suggests a ten-times slower photolysis rate than assumed in simulation M3b (Shi et al., 2021), and a similar reduction was derived from an analysis of  $\text{NO}_y$  partitioning in polluted marine air (Romer et al., 2018). Further lab work has shown that the rate and product yield of this reaction is sensitive to surface composition (Ma et al., 2021). Robust parameterizations will require continued systematic study of controlling factors, preferably under environmentally relevant and near-ambient conditions.

Wildfire emissions may have a more sustained influence on regional background chemistry than is currently appreciated. Synergistic increases in OH reactivity and oxidants in our sensitivity simulations stimulate production of multi-generation oxygenated VOC and organic  $\text{NO}_x$  reservoirs like PAN, altering the spatiotemporal scale of ozone and organic aerosol production. The modest decline of normalized excess OH reactivity (Fig. 4b) further implies that active chemistry persists far downwind. For example, the 62% decrease over 11.8 h of aging in simulation M3b implies an effective pseudo-first-order photochemical lifetime of 12.2 h for the total OH sink. This aligns with a recent analysis of satellite observations suggesting an effective lifetime of 20 hours or more for formaldehyde and glyoxal in aging smoke plumes (Alvarado et al., 2020). Space-based remote sensing of atmospheric composition may aid evaluation of model biomass burning emissions and chemistry on these scales.

## 5 Conclusions

Using a 0-D puff model constrained with SEAC<sup>4</sup>RS in situ observations, we have examined the gas-phase chemical evolution of the 2013 California Rim Fire plume and illustrated the sensitivity of chemistry to unmeasured reactive VOC and HONO. The rich measurement suite permits a holistic evaluation of the various components of the chemical system, including VOC,  $\text{HO}_x$ ,  $\text{NO}_y$ , and  $\text{O}_3$ . Initializing with observed gases only, the model reasonably reproduces the evolution of  $\text{O}_3$  over 12 hours of aging but fails in other key aspects, including over-prediction of reactive VOC decay rates and under-prediction of  $\text{NO}_x$  and HCHO. Accounting for additional VOC identified by recent laboratory studies increases OH reactivity more than twofold throughout the simulation, drawing down model OH and generally degrading performance. Subsequent addition of HONO amplifies radical production and cycling. Addition of initial HONO (assumed to be primary emissions) or secondary production via  $\text{pNO}_3^-$  photolysis improves predictions of ozone, oVOC, and  $\text{NO}_x$ , while HONO production via  $\text{NO}_2$  heterogeneous conversion generally degrades model performance. Further “optimization” simulations demonstrate that multiple combinations of enhanced VOC and primary/secondary HONO can minimize model-measurement bias with respect to  $\text{O}_3$  and  $\text{NO}_x$ , although we cannot reconcile the model with all observations simultaneously. Examination of model reaction rates over the first 2.3 hours of aging demonstrates the potentially dominant contribution of oVOC photolysis to radical initiation, the importance of PN formation as a  $\text{NO}_x$  sink, and the competitive roles of  $\text{RO}_x\text{-RO}_x$  and  $\text{RO}_x\text{-NO}_x$  reactions in radical termination. Ozone production is sensitive to both VOC and  $\text{NO}_x$  in the young plume for all sensitivity simulations.



Ozone production transitions from VOC-sensitive to NO<sub>x</sub>-sensitive after ~1 hour of aging. The timing of this transition depends on VOC and HONO. Downwind NO<sub>x</sub> sensitivity depends on the nature and efficacy of assumed secondary HONO/NO<sub>x</sub> sources.

A primary finding of this study is that reactive VOC and oxidant sources are complementary in the wildfire-influenced atmosphere, and consideration of both is necessary for accurate simulation of near- and far-field impacts on atmospheric composition, air quality, and radiative forcing. Future efforts must focus on an efficient solution for incorporating this somewhat novel chemistry in regional and global models, or else quantify the uncertainty associated with neglecting it.

This work also demonstrates the value of a near-comprehensive payload afforded by a heavy-lift aircraft like the NASA DC-8. The breadth of information in the SEAC<sup>4</sup>RS dataset enables a holistic examination of individual aspects of the chemical system within the context of the whole. Ongoing analysis of data from FIREX-AQ, WE-CAN, and similar missions will illuminate the factors driving variability in reactive VOC, HONO, and other key aspects of fire plume chemistry. This same approach is beneficial to investigation of other complex environments (urban, biogenic, etc.) that comprise the lower troposphere.

## Appendix A. Photolysis Parameterization

For each puff, photolysis frequency  $J$  at solar time  $t$  (and a corresponding Lagrangian plume age) is calculated as

$$J(t, age) = sJ_{cs}(t)/r(age) \quad (A1)$$

Here,  $J_{cs}(t)$  is the clear-sky photolysis frequency and  $s$  and  $r(age)$  are observation-based scaling factors. Clear-sky photolysis frequencies stem from FOAM's hybrid parameterization, which combines solar spectra from the tropospheric ultraviolet and visible radiative transfer model (TUVv5.2, available at <https://www2.acom.ucar.edu/modeling/tropospheric-ultraviolet-and-visible-tuv-radiation-model>) with literature-recommended cross sections and quantum yields. This calculation uses actual SZA and measurement altitude and assumes an overhead ozone column of 290 DU and a surface albedo of 0.01. The age-dependent scaling factor  $r(age)$  is determined by fitting the ratio of clear-sky to observed photolysis frequencies as a function of Lagrangian age using a modified exponential:

$$r(age) = r_0 e^{-age/\tau} + r_\infty (1 - e^{-age/\tau}) \quad (A2)$$

Here,  $r_0$ ,  $r_\infty$ , and  $\tau$  are fitting coefficients that depend on  $J$ . Figure A1 illustrates this fit for ozone and NO<sub>2</sub> photolysis frequencies, which are suppressed by as much as a factor of 3.2 and 1.9, respectively, in the fresh plume. Fits capture 83 – 92% of the variability in the clear-sky to observed ratio for all  $J$ -values, with the exception of two points with ages of 1 – 2 hours. These points are not included in the fit as they appear to be outliers, especially for  $J(O^1D)$ , and their inclusion would significantly degrade overall fit quality. It is unclear why these points differ from the overall trend, as no other dilution or chemical markers show exceptional behavior at these times. The age correction is the same for all puffs. The other parameter,  $s$ , is a scalar multiplicative factor that adjusts model  $J$ -values to agree with observations at trajectory endpoints (analogous to the scaling applied for trajectory meteorology) and is different for each puff.

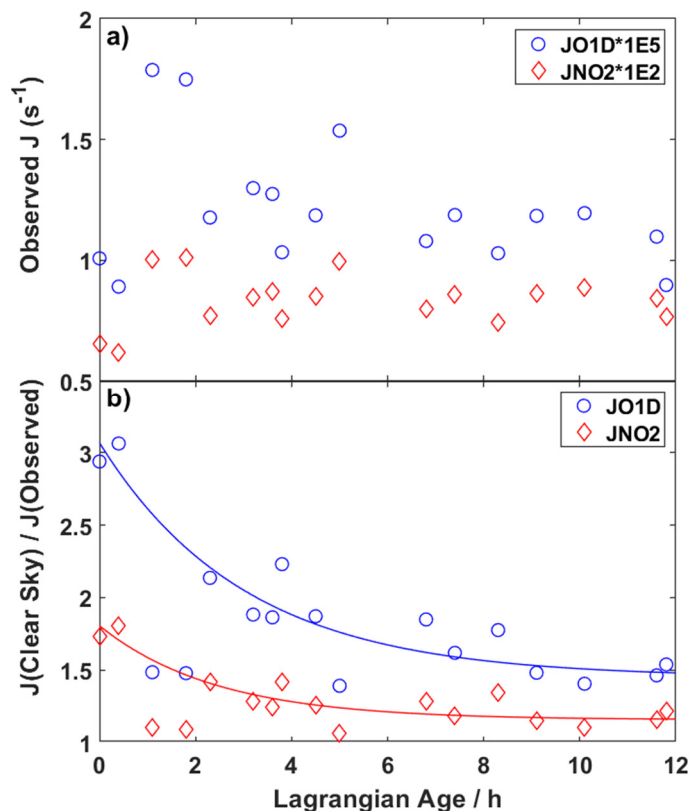


Figure A1. (a) Average observed photolysis frequencies for each WAS sample for  $\text{O}_3 \rightarrow \text{O}(^1\text{D}) + \text{O}_2$  (blue circles) and  $\text{NO}_2 \rightarrow \text{NO} + \text{O}(^3\text{P})$  (red diamonds). (b) Ratio of clear-sky to observed photolysis frequencies. Clear-sky values are calculated using TUV solar spectra and F0AM parameterizations. Lines represent least-squares fits to Eq. (A2). The two points between ages of 1 – 2 hours are excluded from the fit.

### Data and Code Availability

Data used in this study is archived at <http://doi.org/10.5067/Aircraft/SEAC4RS/Aerosol-TraceGas-Cloud>. The F0AM box model is available at <https://github.com/AirChem/F0AM>. Model setup code is available from the contact author upon request.

### Author Contributions

630 GMW conceptualized the study, conducted the modelling and analysis, and wrote the publication. All authors contributed to data curation and review of the manuscript.

## Competing interests

Some authors are members of the editorial board of ACP. An independent editor guided the peer review process, and the authors have no other competing interests to declare.

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