

1 **Analysis and impact of the Hunga Tonga-Hunga Ha'apai Stratospheric Water Vapor**

2 **Plume**

3 M. R. Schoeberl & Yi Wang, Science and Technology Corporation, Columbia, MD, USA

4 R. Ueyama, NASA Ames Research Center, Moffett Field, CA, USA

5 G. Taha, Morgan State University, Baltimore, MD, USA

6 E. Jensen, CIRES, Univ. of Colorado, Boulder, CO, USA

7 W. Yu, Hampton University, Hampton, VA, USA

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9 Corresponding Author: Mark Schoeberl (mark.schoeberl@mac.com)

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## 11 **Key Points**

- 12 • Hunga Tonga-Hunga Ha'apai eruption produced vertically overlapping but slightly  
13 displaced mid-stratospheric enhancements in H<sub>2</sub>O and aerosols.
- 14 • IR cooling by enhanced H<sub>2</sub>O layer explains the observed 4.1 K mid-stratospheric  
15 temperature decrease following the eruption.
- 16 • A simple model of the eruption H<sub>2</sub>O enhancement combined with spreading of the plume  
17 explains the observations.

18

## 19 **Plain Language Summary**

20 The Hunga Tonga-Hunga Ha'apai submarine volcanic eruption on January 15, 2022, injected up  
21 to 150 Tg of water into the stratosphere. A month after eruption, a distinct aerosol and water  
22 vapor layer formed in the tropical southern hemisphere (SH) stratosphere. The water vapor layer  
23 is slightly displaced above the aerosol layer at 26 km. These two layers continued to persist in  
24 the tropical SH stratosphere until the end of June while slowly moving apart in altitude. The  
25 isolation of the layers and their separate motion are consistent with our understanding of tropical  
26 stratospheric dynamics. A cold temperature anomaly forms coincident with the water vapor  
27 layer, which we show to be due to enhanced IR radiative cooling by water vapor. Using a simple  
28 model, we show how the water vapor layer forms slightly above the aerosol layer.

29

## 30 **Abstract**

31 On Jan. 15, 2022, the Hunga Tonga-Hunga Ha'apai eruption injected SO<sub>2</sub> and H<sub>2</sub>O into the  
32 middle stratosphere. The eruption produced a persistent mid-stratospheric sulfate aerosol and  
33 H<sub>2</sub>O layer mostly confined to Southern Hemisphere (SH) tropics (Eq. to 30°S). These layers are  
34 still present in the tropics 5 ½ months after the eruption. The SH tropical confinement is  
35 simulated using a trajectory model. Measurements following the eruption show that the H<sub>2</sub>O  
36 layer is slowly rising while the aerosol layer is descending. The H<sub>2</sub>O layer's upward movement  
37 is consistent with the residual vertical velocity. Gravitationally settling explains the descent of  
38 the aerosol layer. A -4K temperature anomaly coincident with the H<sub>2</sub>O enhancement is observed  
39 and is caused by thermal adjustment to the additional H<sub>2</sub>O IR cooling. A simple model of  
40 volcanic water injection at the time of the eruption simulates the observed vertical distribution  
41 H<sub>2</sub>O.

42

## 43 **Index Terms**

44 0340 Middle atmosphere dynamics

45 0341 Middle atmosphere: constituent transport and chemistry

46 0370 Volcanic effects

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## 50 1. Introduction

51 The Hunga Hunga-Tonga Ha'apai (HT) (20.54°S, 178.3°W) submarine volcano violently  
52 erupted on Jan. 15, 2022. The volcanic explosivity index (VEI) was 5, comparable to Krakatau  
53 eruption in 1883. Since HT was a submarine volcano, it appears to have lofted a significant  
54 amount of water into the stratosphere. Indeed, Microwave Limb Sounder (MLS) measurements  
55 show that HT water enhancement was quite high relative to SO<sub>2</sub> (Millán et al., 2022) – hereafter  
56 M22. The MLS estimated water injection was up to 146 Tg (M22). The eruption plume was  
57 detected up to 57 km on January 15, 2022 (Carr et al., 2022; Proud et al., 2022). The Ozone  
58 Mapping and Profile Suite – Limb Profiler (OMPS-LP) detected extinction enhancements above  
59 45 km (Taha et al., 2022).

60  
61 In this paper we will examine at the evolution of the water vapor and aerosol enhancements that  
62 followed the HT eruption. M22 noted that the amount of water deposited in the stratosphere by  
63 HT was unprecedented in the modern history of volcanic eruption observations. Several MLS  
64 water vapor profiles made shortly after the eruption show concentrations exceeding 300 ppmv  
65 against a normal stratospheric concentration of ~4 ppmv. As the eruption evolved, MLS water  
66 vapor maps show that above about 2 hPa (~43 km), the plume quickly spreads and that the water  
67 vapor enhancement disperses. A secondary maximum at about 25 hPa (~26 km) persists (M22).  
68 The aerosol field shows similar behavior with rapid dispersal at higher altitudes but persistent  
69 high levels of aerosol extinction below ~ 25 hPa (~26 km) (Taha et al., 2022). The aerosol  
70 extinction in this layer grows over the 30 days following the eruption presumably due to the  
71 conversion of SO<sub>2</sub> to sulfate aerosols (e.g. Zhu et al., 2020).

72  
73 There are several key questions concerning the HT eruption: Why did the unusual water vapor  
74 layer form and persist? How is it related to the aerosol layer? Below we show that the water  
75 vapor enhancement overlaps the top of the extinction anomaly, but they are distinct, and  
76 furthermore the two enhancements vertically separate over time. We have also discovered a  
77 temperature anomaly in the 25-28 km region. We provide an explanation for the temperature  
78 anomaly as well as for the formation and evolution of the water vapor and aerosol layers.

## 79 80 2. Data sets

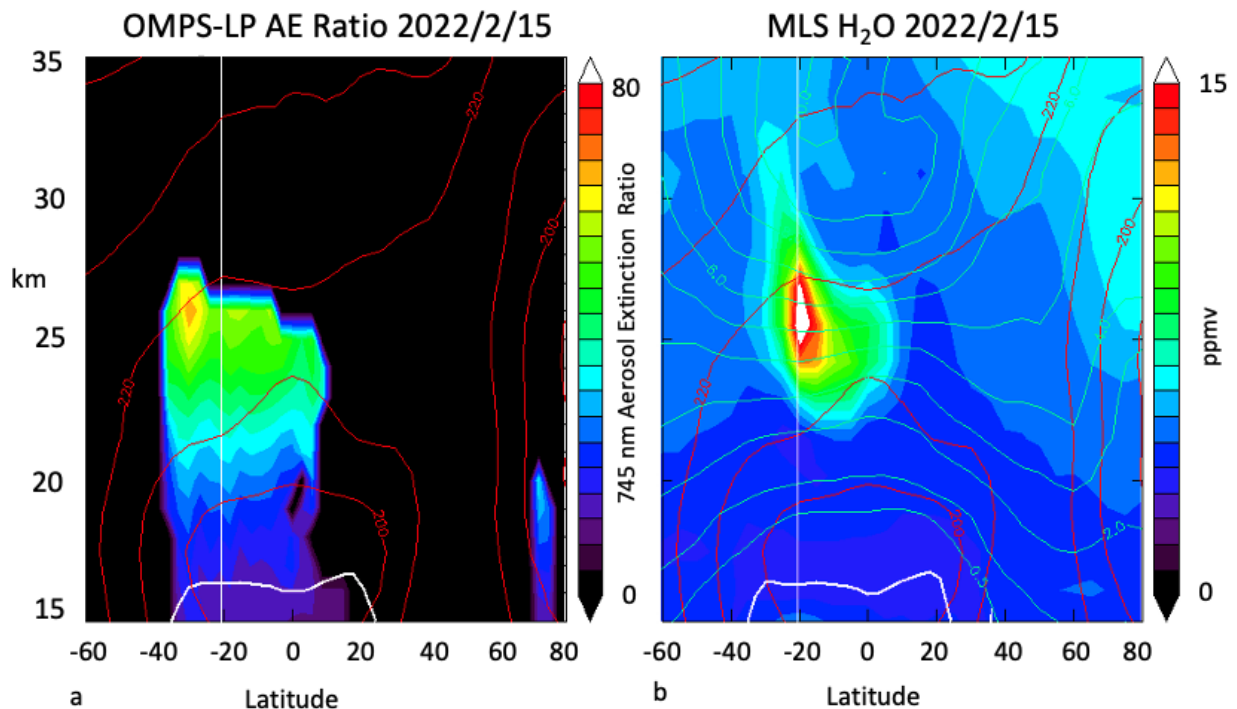
81  
82 Generally, we use MLS v5 for ozone, temperature and H<sub>2</sub>O where the data where the quality and  
83 convergence flags are not set. However, the MLS algorithm quality flags and convergence alerts  
84 were set for some plume profiles in the week or so after the eruption. However, even with the  
85 quality flag and convergence filters set, the data look reasonable and generally agrees with sonde  
86 and other validation data so we used the immediate post eruption data. The data quality for the  
87 HT anomaly is detailed in M22 and MLS data is described in Livesy et al. (2021). For aerosols,  
88 we use OMPS-LP level-2 V2.1 745 nm extinction-to-molecular ratio data (AE) from all three  
89 OMPS-LP slits (see Taha et al., 2021). Taha et al. (2022) indicated that the standard V2.1  
90 released data (used in this study) provide the most accurate aerosol retrieval up to 36 km. Thus,  
91 we restrict our constituent analysis to below 35 km which contains the main locus aerosol plume  
92 (Taha et al. 2022; Fig. 4). The MLS and OMPS-LP extinction data sets are averaged over 4 days  
93 and then averaged onto a 5° latitude-longitude grid.

94

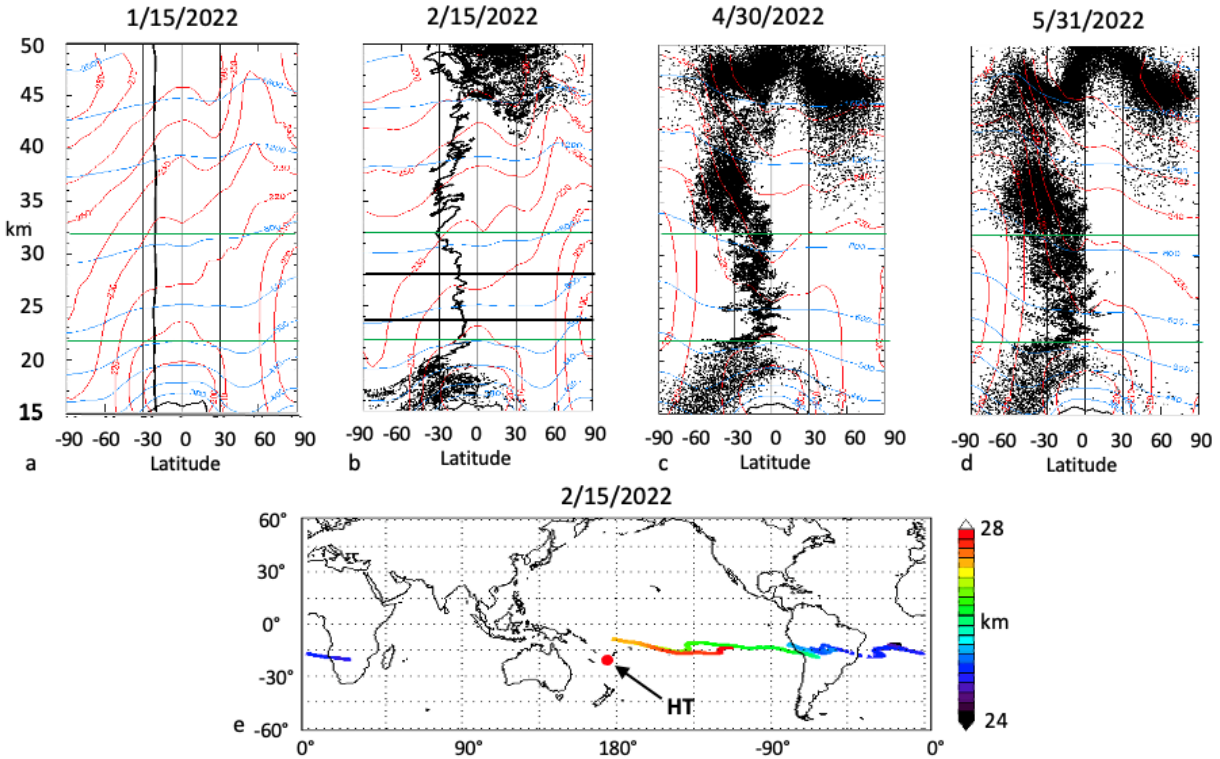
95 To simulate the dispersal of the water vapor/aerosol plumes, we use the Forward Domain Filling  
96 (FDF) trajectory model (Schoeberl et al., 2018) modified to inject a dense uniform column of  
97 parcels over the HT location on Jan 15, 2022. This simulation uses MERRA-2 reanalysis winds,  
98 temperatures, and heating rates (Gelaro, et al., 2017).  
99

### 100 3. Analysis

101  
102 Figure 1 shows the zonal mean distribution of water vapor and aerosol extinction ratio on Feb  
103 15, 2022, a month after the eruption. The HT aerosol plume reaches 26 km in the region 30°S to  
104 about 5°N. The extinction data are quite sensitive to plumes extending outward from the tropics  
105 thus tend to show a wider distribution than the water vapor field. The water vapor plume is  
106 centered at 26 km and extends up to 30 km in the SH tropics. The water vapor plume mostly  
107 overlaps the aerosol plume while extending slightly above it.  
108  
109



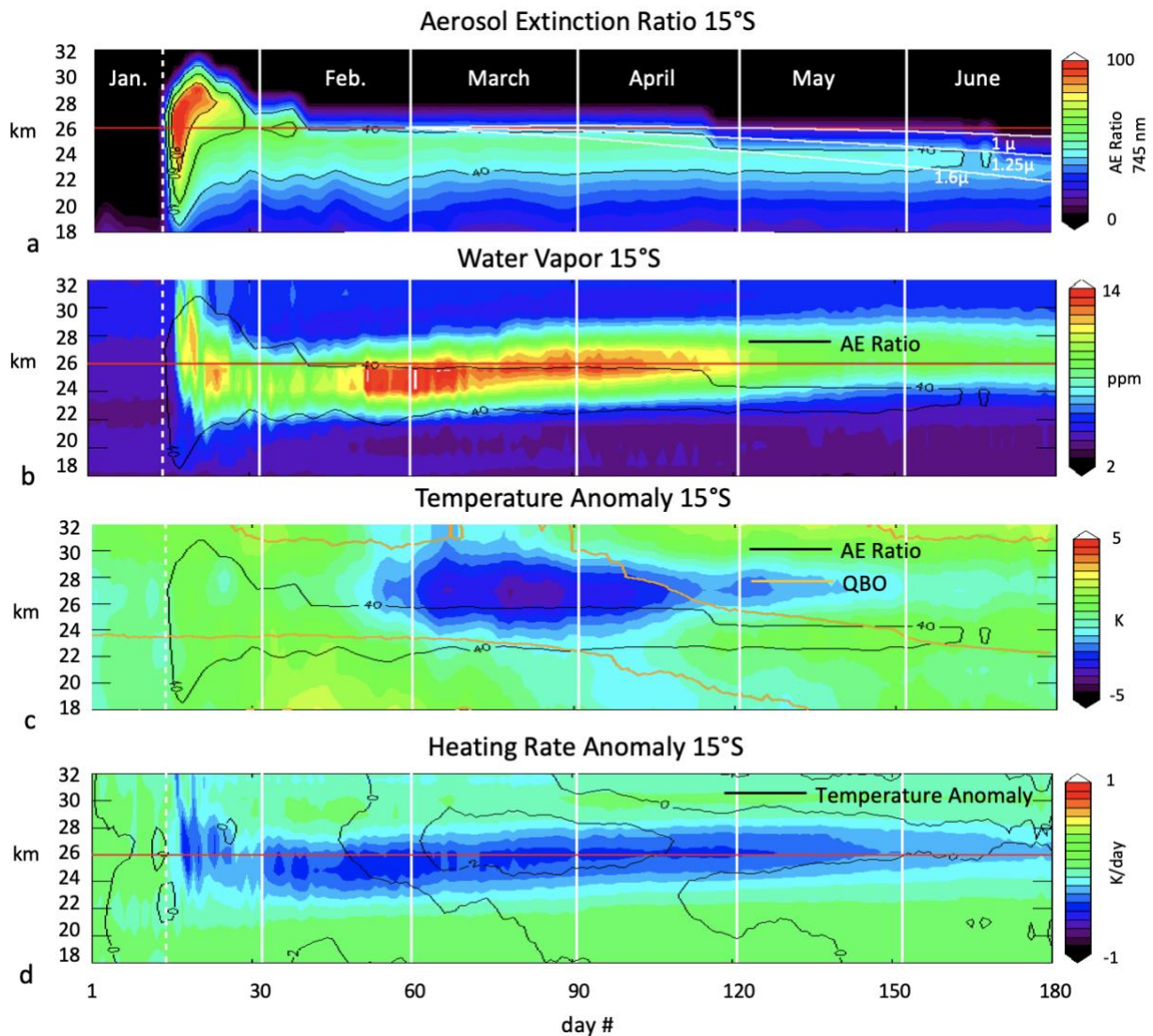
110  
111 *Figure 1 The zonal mean OMPS-LP 745 nm aerosol extinction/ molecular extinction ratio (Part*  
112 *a) and MLS water vapor (ppmv) (Part b) on Feb. 15, 2022. The red contours show the MLS*  
113 *temperature field. The thick white line is the zonal mean tropopause. The green contours are*  
114 *MLS ozone mixing ratio (ppm). The vertical white line denotes the latitude of the HT volcano.*



115  
 116 *Figure 2 Dispersal of HT plume simulated by the FDF model: (Part a) shows the initial parcel*  
 117 *distribution on Jan 15, (Part b) parcel distribution of Feb. 15, 2022, (Part c) shows the*  
 118 *distribution at the end of April, and (Part d) shows the distribution at the end of May. Along the*  
 119 *bottom (Part e), a map of parcels between 24-28 km with color scale indicating altitude. In Parts*  
 120 *a,b,c,d red contours are MERRA-2 temperatures, blue contours are potential temperature.*  
 121 *Horizontal green lines show the isolation region 22-32km. Horizontal black lines in Part b*  
 122 *indicate the domain in Part e. The red dot locates HT on the map.*

123 Figure 2 shows the dispersal of the plume using the FDF trajectory model. From the initial  
 124 uniform altitude distribution (Fig. 2a), the plume evolves slowly and is mostly confined to the  
 125 region between the equator and 30°S in the height range 22-32 km. This confinement is still  
 126 somewhat evident at the end of April. The isolation of the Northern Hemisphere (NH) tropics  
 127 from the Southern Hemisphere (SH) tropics in this region was first noted by Stolarski et al.  
 128 (2014) when analyzing the interhemispheric phasing of the tropical ozone concentration. Below  
 129 ~20 km parcels are dispersing mostly to the SH extra-tropics along the isentropes. Above about  
 130 35 km parcels are also dispersing and drifting poleward together into the SH. At highest  
 131 altitudes, parcels are moving out of the tropics into the NH extra-tropics.

132  
 133 Timeseries of the zonal mean aerosols and water vapor at 15°S±2.5° are shown in Fig. 3a,b. We  
 134 also plot the temperature anomaly (Fig. 3c) as a departure of the zonal mean MLS temperature  
 135 from the 2021-2016 MERRA2 climatology. The perturbation heating rate shown in Fig. 3d is  
 136 computed using the AER longwave radiative transfer model (Mlawer et al., 1997). The heating  
 137 rate calculation uses observed MLS ozone, temperature, and water vapor. The heating rate  
 138 anomaly is computed by fixing the water vapor to the pre-eruption profile and computing the  
 139 heating rates over the period. We then subtract those heating rates from the heating rates  
 140 computed using MLS observed water vapor data.



142  
143 *Figure 3 Times series of 2022 aerosol extinction ratio (AE) (Part a), water vapor (Part b),*  
144 *temperature anomaly (Part c), and heating rate anomaly (Part d) at 15°S. Parts b & c show*  
145 *black contours of 40 AE ratio that outline the aerosol anomaly. In Part d, the heating rate*  
146 *anomaly has the Part c temperature contours superimposed. White lines in Part a represent the*  
147 *downward gravitational settling of aerosols of different diameters ( $\mu\text{m}$ ) as labeled. Orange*  
148 *contours in Part c (QBO) indicates the altitude of the zero zonal wind lines at the equator*  
149 *showing the descent of the QBO. The red line in all parts is at 26 km. Vertical white lines show*  
150 *month boundaries, months labeled in Part a.*

151 Comparing the aerosol extinction field (Fig. 3a) with the water vapor (Fig. 3b), we see that the  
152 water vapor anomaly is slowly ascending whereas the aerosol concentration is descending. The  
153 simple explanation for this effect is that the water vapor is transported upward with the diabatic  
154 circulation that gives rise to the tropical trace gas tape recorders (Schoeberl et al., 2008a)  
155 whereas the aerosols are gravitationally settling. The 26 km water vapor anomaly ascent rate is  
156  $\sim 2$  km over 45 days (after March 1) or  $\sim 0.044$  km/day. We have computed the residual

157 circulation over the same period, and it averages to 0.045 km/day consistent with the estimate  
158 from water vapor. Using  $w^*$  as the ascent velocity, Fig 3a shows the net settling rate for aerosols  
159 with different sizes after day 60. The settling rate is computed from Stokes formulas in  
160 Pruppacher et al. (1998). The change in the aerosol height appears to match the settling for  
161 aerosol modal diameter of  $\sim 1.2 \mu\text{m}$ . Smaller particles would be carried upward by the  
162 circulation into warmer, lower relative humidity environment, and would evaporate  
163 (Tsagkogeorgas et al., 2017).

164

165 By mid-March, the descending QBO circulation weakens the background upward residual  
166 circulation to  $\sim 0.02$  km/day which slows the ascent of the water vapor anomaly as is evident in  
167 Fig. 2b. The equatorial zero wind line altitude is superimposed on Fig. 3c to show the descent  
168 (see [https://acd-ext.gsfc.nasa.gov/Data\\_services/met/qbo/qbo.html](https://acd-ext.gsfc.nasa.gov/Data_services/met/qbo/qbo.html)).

169

170 Fig. 3c shows a cold temperature anomaly that begins to appear in early to mid-February, and the  
171 anomaly magnitude is consistent with radiosonde measurements (Vömel et al., 2022). This  
172 temperature anomaly, which exceeds  $-4\text{K}$ , is approximately coincident with the change in the  
173 cooling rate (Fig. 3d; correlation of  $r = 0.77$  for the period Feb. 15-July 1) due to enhanced water  
174 vapor. If we assume in the thermodynamic equation that the temperature change ( $\Delta T$ ) balances  
175 the change radiative heating ( $\Delta H$ ),  $\Delta T \sim a\Delta H$ , then we compute a Newtonian cooling time scale  
176 ( $\text{a}^{-1}$ ) of 3.3 days at 26 km. This time constant is consistent other estimates of the Newtonian  
177 cooling rate for this region (e.g. Newman and Rosenfield, 1997). Thus, the temperature changes  
178 observed in the mid stratosphere are part of the thermal adjustment to the increased IR cooling,  
179 and we expect circulation changes as well (Coy et al., 2022).

180

181 Note that volcanic aerosols can heat the stratosphere (Aubry et al., 2021 and references therein)  
182 and this heating would oppose the water vapor cooling. Shortly after the eruption, sonde  
183 measurements show a  $< 2\text{K}$  increase in temperatures below 25 km that disappears by early  
184 February (Vömel et al., 2022). After February we see no evidence of a temperature change co-  
185 located with the aerosol layer probably because the dispersed aerosol layer is too attenuated.

186

187 What governs the vertical structure of the water vapor anomaly? To explore this problem, we  
188 have constructed a very simple model of the HT plume based on observations. Initially, the  
189 eruption is propelled upward by the explosion and latent heat release through condensation of  
190 water vapor at lower altitudes. The initial plume temperature is likely well above stratospheric  
191 ambient temperatures. Within days to weeks the plume shears out and plume temperature cools  
192 to ambient. Estimates of aerosol radiative heating by Silletto et al. (2022) show that longwave  
193 aerosol cooling and shortwave aerosol heating nearly cancel leaving water vapor cooling as the  
194 major radiative component. We assume, for simplicity, that the amount of water vapor available  
195 is now limited by the saturation mixing ratio over ice i.e., excess water forms ice particles that  
196 quickly fall out until the relative humidity is reduced to 100%. The falling ice evaporates,  
197 saturating any unsaturated layers below. We then assume that the amount of HT water lofted  
198 decreases above the eruption top centroid height in mid-February – prior to mid-February the  
199 system is still in adjustment (Legras et al., 2022). These two principles define the available  
200 water. Based on GOES images, the 16 km volcanic cloud is  $\sim 500\text{-}1000$  km in diameter. This  
201 area must expand with altitude, to conserve mass. As in Fig. 2e, the eruption cloud stretches out  
202 in longitude. Thus, by mid-February MLS zonal mean water vapor is the available water in the



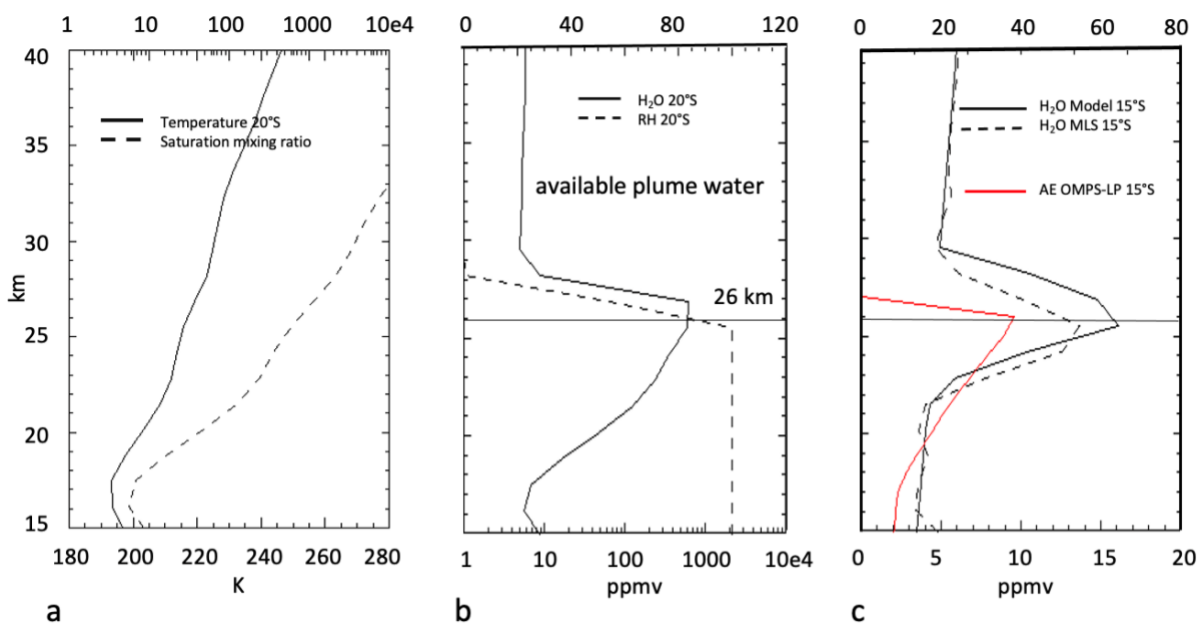
203 eruption cloud reduced by the ratio of the initial eruption cloud area to the tropical zonal mean  
 204 area.

205  
 206 The model uses the OMPS-LP mid-February aerosol extinction profile (Fig. 1a) to set the  
 207 eruption top centroid height,  $z_{cent}$ ; the model assumes the cloud is roughly mixed zonally. We  
 208 center a vertical Gaussian-type distribution around the centroid. The temperature profile at 20°S  
 209 is shown in Fig. 4a along with the saturation mixing ratio over ice (Murphy and Koop, 2005).  
 210 Below the eruption centroid, the ice amount is equal to the saturation mixing ratio; above the  
 211 centroid, the amount of water available is the saturation mixing ratio decreasing with altitude as  
 212  $\exp(-(z-z_{cent})^2/2L^2)$  where  $L=0.65\text{km}$  and  $z_{cent} = 26 \text{ km}$ . We add the observed background pre-  
 213 eruption zonal mean MLS water vapor profile for realism. Fig. 4b shows the assumed eruption  
 214 available water vapor profile and relative humidity on Feb 15. The available water reaches 600  
 215 ppmv at 26 km. MLS did observe the water vapor mixing ratios over 300 ppm at 26 km Jan. 16  
 216 (M22) and there are sonde measurements of even higher mixing ratios in this stratospheric region  
 217 (Vömel et al., 2022).

218  
 219 The Feb. 15 zonal mean water vapor field (Fig. 4c) is assumed to be 15° wide from 5°S to 20°S  
 220 and consists of the diluted plume shown in Fig. 3b. Fig. 4c also shows the observed aerosol  
 221 extinction profile. The aerosol extinction profile is only used to verify the height of the eruption  
 222 centroid and its width. The water vapor profile shows good agreement with zonal mean MLS  
 223 data at 15°S. The extension of the water plume above the aerosol plume is also reproduced. The  
 224 Fig 4c column water vapor mass above 100 hPa is 32.8 Tg; the MLS mass is 31.2 Tg.

225  
 226 In summary, the simple model requires three factors to explain the water vapor anomaly that  
 227 extends above the aerosol anomaly: (1) the change in the saturation mixing ratio with altitude as  
 228 controlled by the tropical temperature profile, (2) a decrease in volcanic water injection above  
 229 the eruption top, and (3) spatial dilution of the eruption plume.

230



231  
 232 *Figure 4 Model of HT water vapor injection. Part a shows Jan 15, 2022 temperature and*  
 233 *saturation mixing ratio profile at the location of HT (20°S). Part b shows a model of the eruption*

234 *available water vapor to be mixed with the environment, and relative humidity profile, the*  
235 *eruption top is at 26 km. Part c shows the zonal mean aerosol extinction ratio profile for Feb.*  
236 *15, 2022 (red). Zonal mean water vapor profile (black) for the model and MLS zonal mean water*  
237 *vapor (dashed).*

238

## 239 **5. Summary and Discussion**

240 The HT volcanic eruption produced stratospheric enhancements of both water and aerosols  
241 (sulfate after SO<sub>2</sub> oxidation). Our analysis shows that the aerosol and water vapor enhancements  
242 persisted from Jan 15 to July 1, 2022. Between 22-32 km the enhancements are confined mostly  
243 to the SH tropics as is evident from observations and consistent with a trajectory analysis. This  
244 isolation of the stratospheric SH tropics from the NH tropics is consistent with tropical ozone  
245 observations (Stolarski et al., 2014). Below about 20 km, the aerosol observations and trajectory  
246 analyses show that aerosols and water mostly disperse out of the SH tropics. The trajectories  
247 suggest that most of the aerosols move to the SH with a smaller amount moving into the NH.  
248 Above 40 km the trajectory model suggests that eruption material moves into the Northern  
249 Hemisphere as part of the cross-hemispheric upper stratospheric circulation (Schoeberl and  
250 Strobel, 1978; Holton and Wehrbein, 1980).

251

252 By mid-February, the tropical mid-stratosphere aerosol and water vapor enhancements are  
253 slightly offset from each other, with the water vapor anomaly about 1 km higher. The two  
254 distinct layers continue separate over the 5½ month period following the eruption. The ascent  
255 speed of the water vapor anomaly is consistent with the magnitude of the upward branch of the  
256 large scale residual circulation. The descent of the aerosol layer is consistent with the  
257 gravitational settling of particles ~ 1.2 μm consistent with an independent analysis by Legras et  
258 al. (2022). Smaller particles will be carried upward by the circulation and evaporate in the  
259 warmer layers above.

260

261 Tropical temperatures at 26 km, 15°S show anomalous decreases about a month after the  
262 eruption and are coincident with the water vapor enhancement at that altitude. This temperature  
263 decrease is also seen in sonde measurements (Vömel et al., 2022). IR radiative transfer  
264 computations show that the temperature decrease is correlated with enhanced water vapor IR  
265 cooling as might be expected (de F. Forster and Shine, 1999). The short-wave heating and long  
266 wave cooling by aerosols appear to roughly cancel (Silleto, 2022). Thus, the temperature  
267 change appears to be part of the dynamical response to the increased H<sub>2</sub>O IR cooling. The other  
268 part of the response will be a circulation adjustment (Coy et al., 2022). The Newtonian cooling  
269 rate calculated from observed temperature and cooling rate changes is consistent with previous  
270 computations (Newman and Rosenfield, 1997).

271

272 To explore the formation of the water vapor anomaly, we use a simple model of the eruption. In  
273 the model we define an eruption top altitude, we assume that there is a decreasing amount of  
274 water injected above that altitude and the relative humidity below that altitude is 100%. The  
275 water vapor then disperses zonally. Our model water vapor matches the zonal mean MLS  
276 measurements one month after the eruption and is consistent with the range of MLS H<sub>2</sub>O  
277 measurements made shortly after the eruption (M22).

278

279 Our simple model suggests that even larger water vapor anomalies would have formed if the  
280 volcanic eruption had lofted water into higher, warmer stratospheric air. On the other hand,  
281 smaller water vapor anomalies would have occurred for lower altitude injections or higher  
282 latitude injections into colder stratospheric air. This, along with the fact that most volcanic  
283 eruptions in the recent past were not submarine may explain why water vapor enhancements  
284 have not been as large in previous eruptions (e.g. St. Helens - Murcray et al., 1981; Calbuco -  
285 Sioris et al. 2016; Kasatochi - Schwartz et al., 2013).

286

### 287 **Acknowledgements**

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290 suggestions.

291

### 292 **Open Research**

293 MERRA-2 Reanalysis data. Gelaro et al. (2017). MERRA-2 data are obtained from the Global  
294 Modeling and Assimilation Office (GMAO), *inst3\_3d\_asm\_Cp: MERRA-2 3D IAU State,*  
295 *Meteorology Instantaneous 3-hourly (p-coord, 0.625x0.5L42), version 5.12.4* at [https://doi.org/](https://doi.org/10.5067/WWQSXQ8IVFW8)  
296 [10.5067/WWQSXQ8IVFW8](https://doi.org/10.5067/WWQSXQ8IVFW8). Data is public, unrestricted access (registration required).

297

298 OMPS-LP data, Taha et al. (2021), is available at

299 [https://disc.gsfc.nasa.gov/datasets/OMPS\\_NPP\\_LP\\_L2\\_AER\\_DAILY\\_2/summary](https://disc.gsfc.nasa.gov/datasets/OMPS_NPP_LP_L2_AER_DAILY_2/summary),

300 DOI: <https://doi.org/10.5067/CX2B9NW6FI27>. The algorithm is documented in Taha et al.  
301 (2021). Data is public, unrestricted access (registration required).

302

303 Aura MLS Level 2 data, Livesey et al. (2021) JPL D-33509 Rev. C, is available at

304 <https://disc.gsfc.nasa.gov/datasets?page=1&keywords=AURA%20MLS>

305 The temperature data is available at

306 [https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura\\_MLS\\_Level2/ML2T.004/](https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2T.004/)

307 The V4 water vapor data is available at

308 [https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura\\_MLS\\_Level2/ML2H2O.004/](https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2H2O.004/)

309 The V5 water vapor data is available at

310 [https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura\\_MLS\\_Level2/ML2H2O.005/](https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2H2O.005/)

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