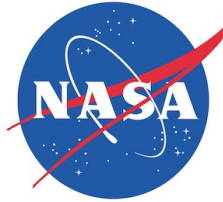


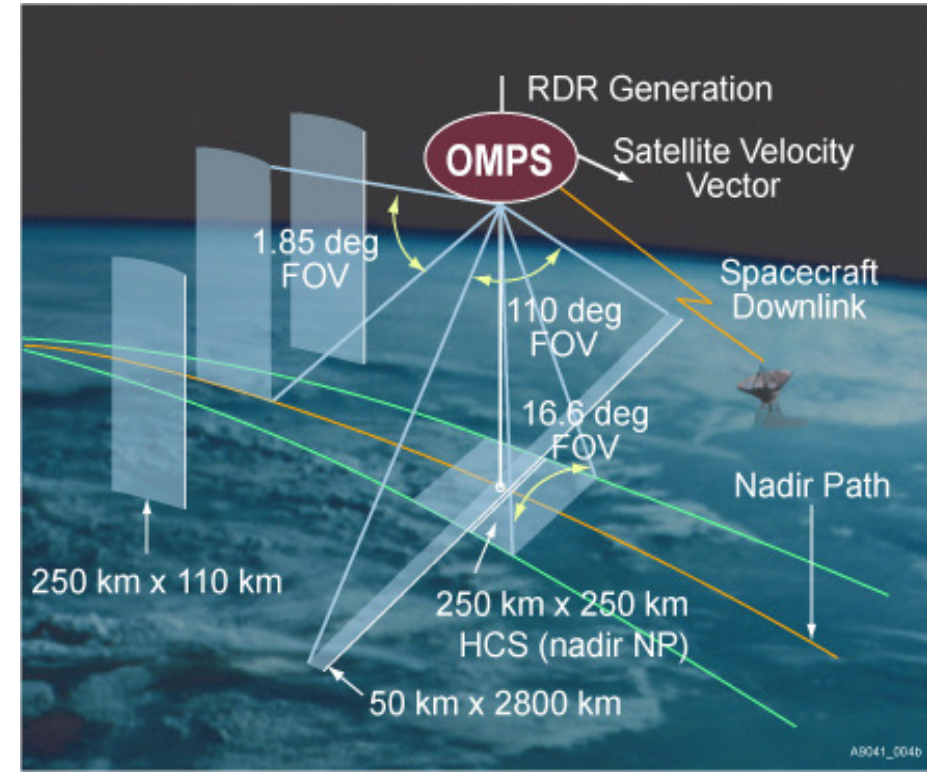
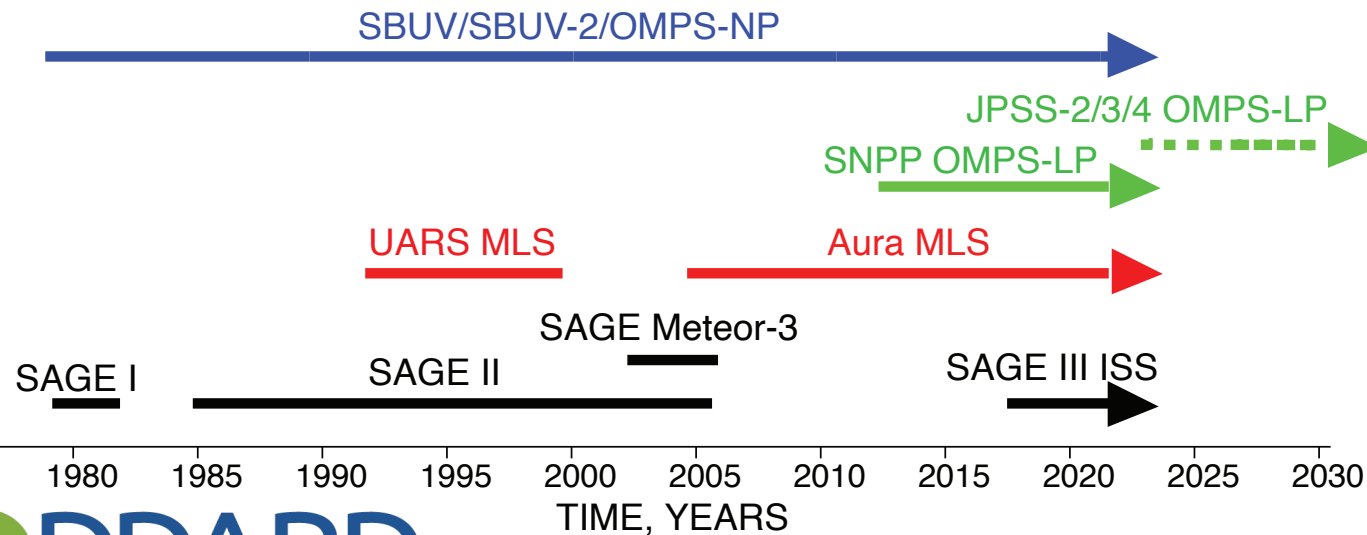
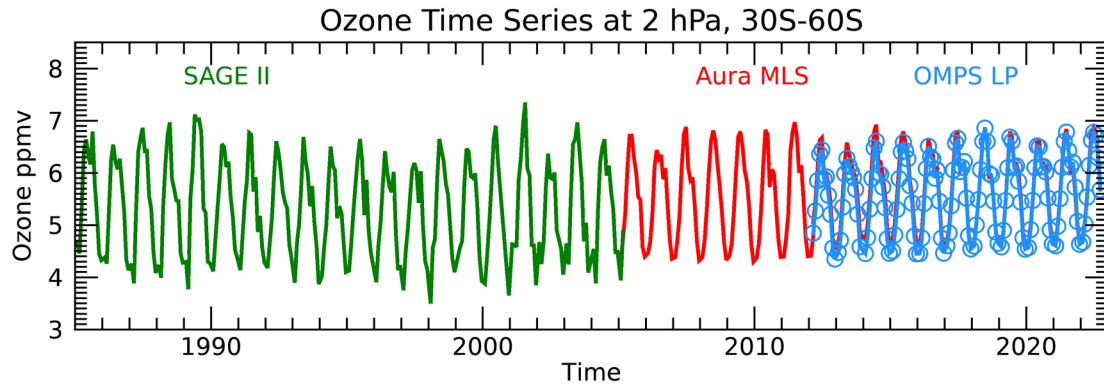
# Ten Year Ozone Profile Record from Suomi NPP OMPS Limb Profiler

Natalya Kramarova<sup>1</sup>, Philippe Xu<sup>2</sup>, Jungbin Mok<sup>3</sup>, P.K. Bhartia<sup>1,\*</sup>, Glen Jaross<sup>1</sup>, Leslie Moy<sup>3</sup>,  
Clark Weaver<sup>4</sup>, Stacey Frith<sup>3</sup>, Jerald Ziemke<sup>5</sup>, Zhong Chen<sup>3</sup>, Daniel Kahn<sup>3</sup>, Ernest Nyaku<sup>3</sup>,  
Jason Li<sup>3</sup>, Sean Davis<sup>6</sup>, and Yue Jia<sup>6</sup>

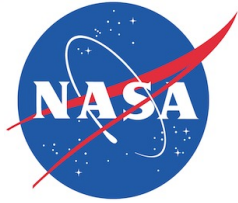
1-NASA GSFC, Greenbelt, MD; 2- SAIC, Greenbelt, MD; 3- SSAI, Greenbelt, MD; 4-ESSIC, College Park, MD; 5-Morgan State University,  
Baltimore, MD; 6-NOAA CSL, Boulder, CO; \*- Emeritus



## SATELLITE OZONE PROFILE MEASUREMENTS



SCIENCE LEADS THE FUTURE

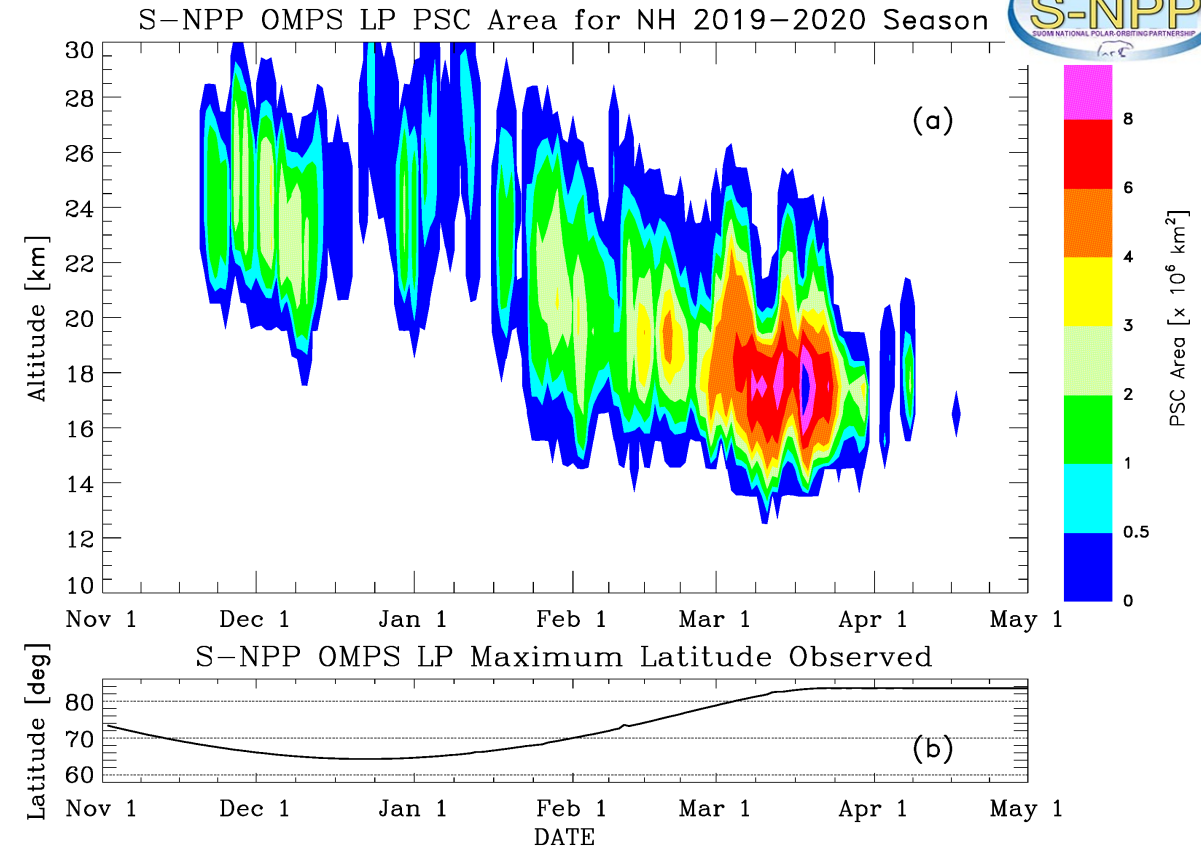
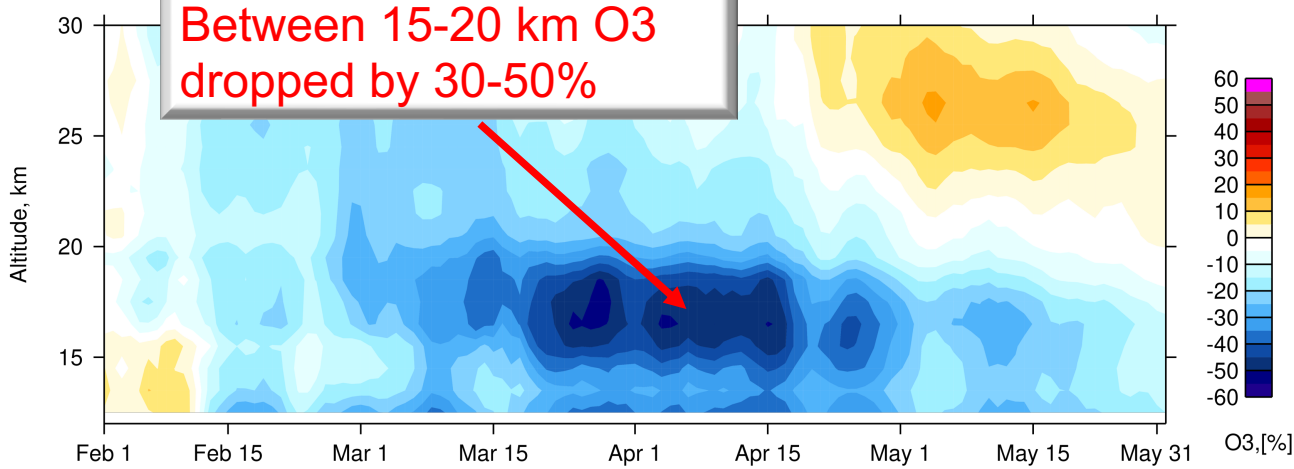


## MONITORING POLAR OZONE DEPLETION WITH THE SNPP OMPS

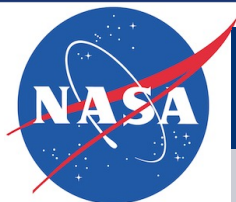


- Feb-Mar 2020: Lack of planetary wave forcing → stratospheric T below average;
- Cold temperatures → increased volume of PSCs over the Arctic;
- The Arctic ozone depletion in March 2020 was the worst since 1979.

2020 OMPS LP ozone anomalies, 60N-82N



[DeLand et al., JRL, 2020]



## Level 1

## Level 2

### Altitude registration:

- Static correction update (1.58 km or +200 m);
- Remove the second 100m step in Sep. 2014;
- Simplified intra-orbital correction (~ 650 m);

### Update Stray Light correction:

- Slit image increased by a factor of 1.5 for VIS;
- 12% increase in the tails for PSF for OOR;

### Static radiometric calibrations updates:

- Smoothed albedo pre-launch;
- Goniometric Day-1 + seasonal component;
- Wavelength scale Day 1 assignment

### Wavelength-shift correction (time-dependent);

### Radiometric calibration drift;

### Update $O_3$ and $NO_2$ absorption cross sections and climatologies:

- Brion-Dumont-Mallicet (BDM) in UV (290-355 nm);
- Serdyuchenko-Gorshchev (SG) in VIS (500-700 nm);
- Update  $O_3$  and  $NO_2$  climatological profiles;

### Combine UV and VIS radiances to retrieve a single $O_3$ profile;

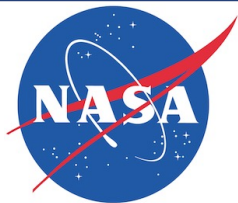
### Increase number of wavelengths (6 UV pairs and 1 VIS triplet);

### Dynamical vertical range for each pair/triplet contribution (based on max sensitivity to $O_3$ );

### Implement Tikhonov regularization to vertically smooth retrieved profiles;

### Update convergence criteria and quality filters;

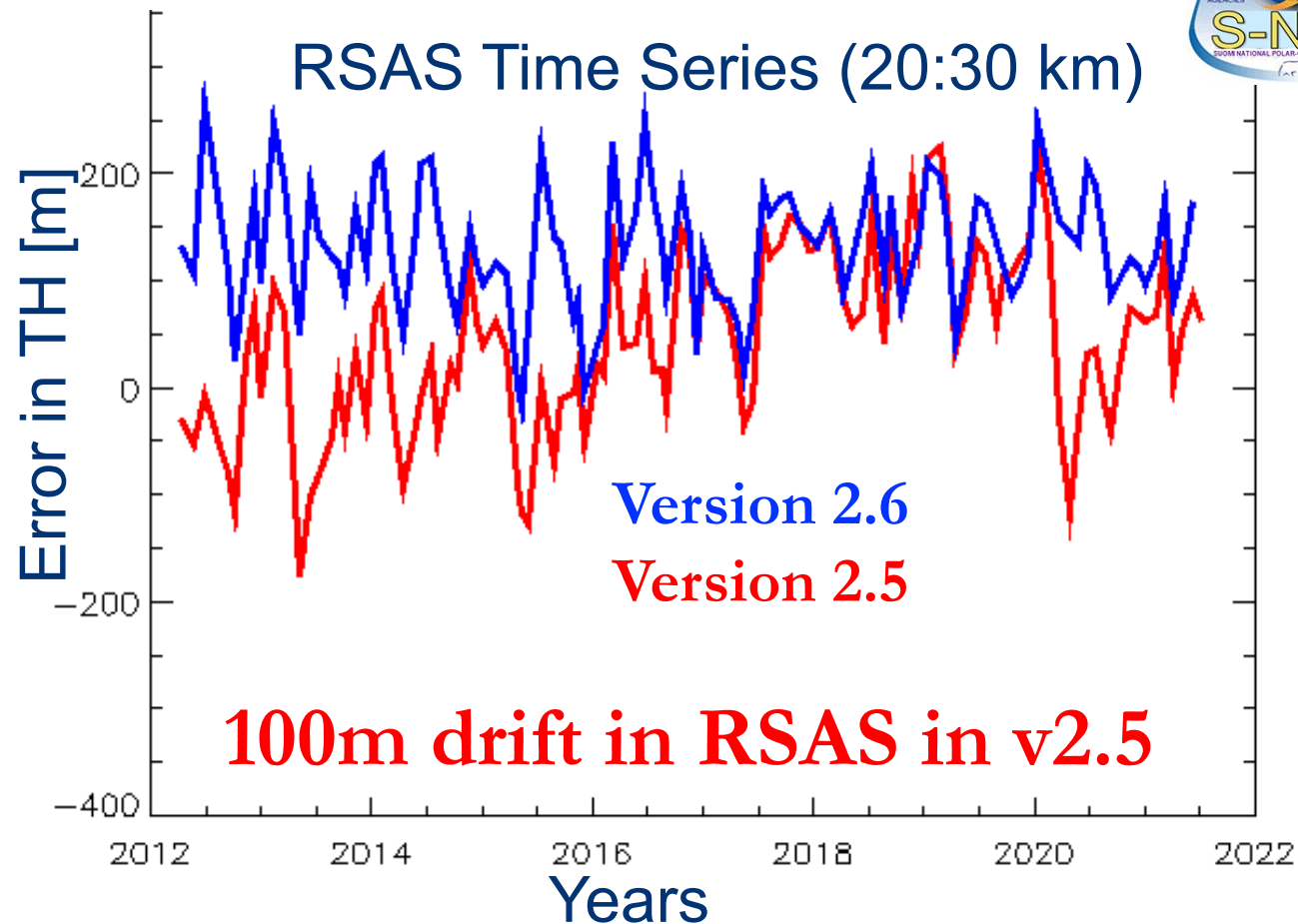
### Switch to gamma-function aerosol size distribution;

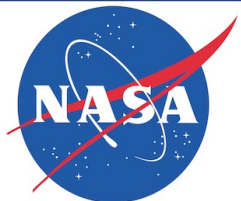


# OMPS LP ALTITUDE REGISTRATION



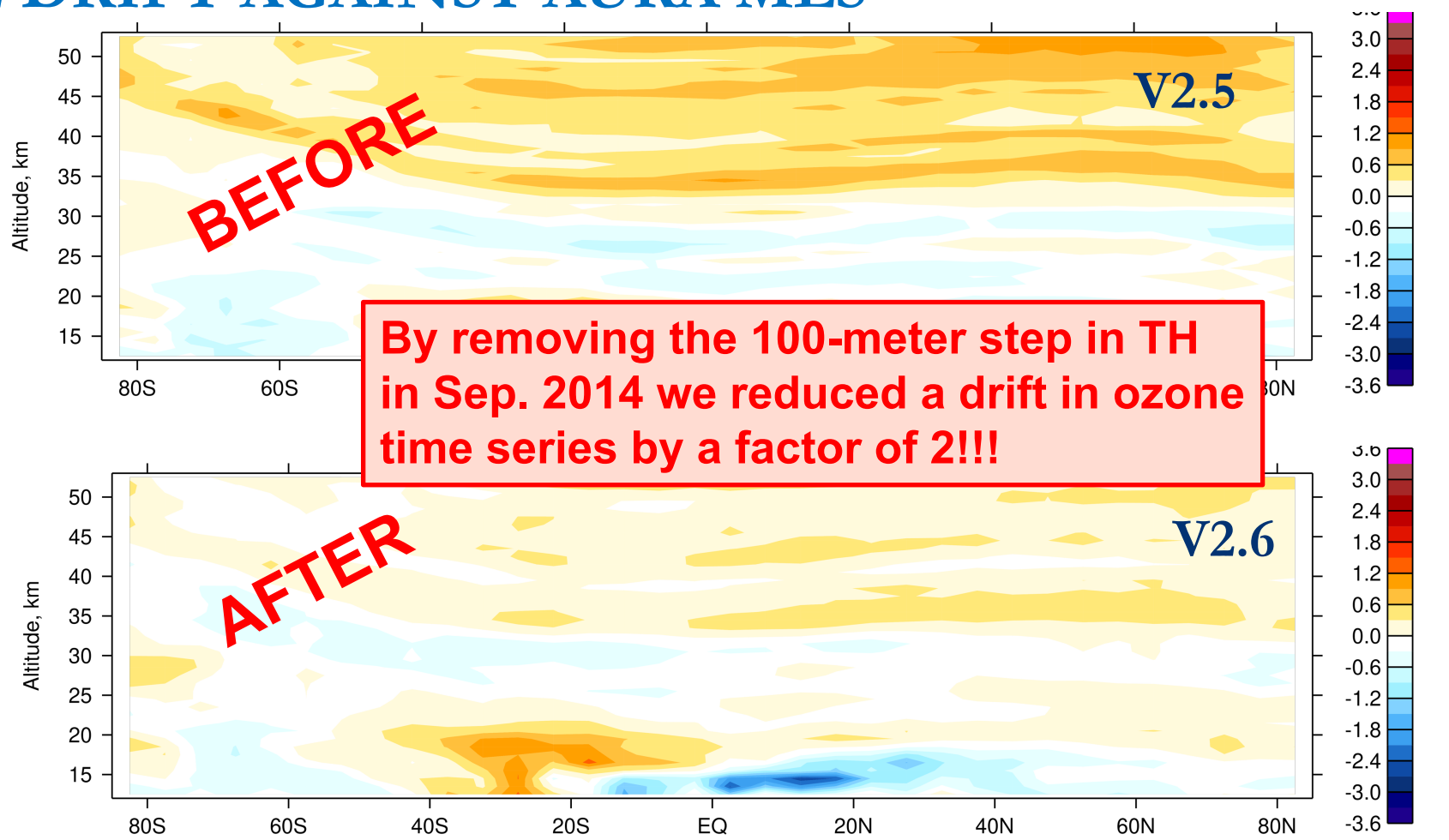
- Uncertainties in the altitude registration are the main source of errors in the limb scattering technique (200 m  $\rightarrow$  5% error in  $O_3$ );
- Preliminary analysis of v2.6 data shows no significant drift in the center slit;
- Radiometric calibrations, improvements in forward model calculations and updated aerosol PSD led to improvements in RSAS time series in v2.6.

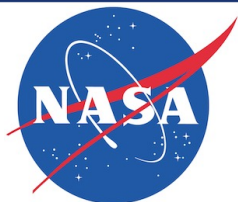




## RELATIVE DRIFT AGAINST AURA MLS

- Several studies found a drift in v2.5 LP O<sub>3</sub> time series with the patterns consistent with a possible drift in altitude registration;
- Preliminary results demonstrate a factor of 2 reduction in relative drifts.



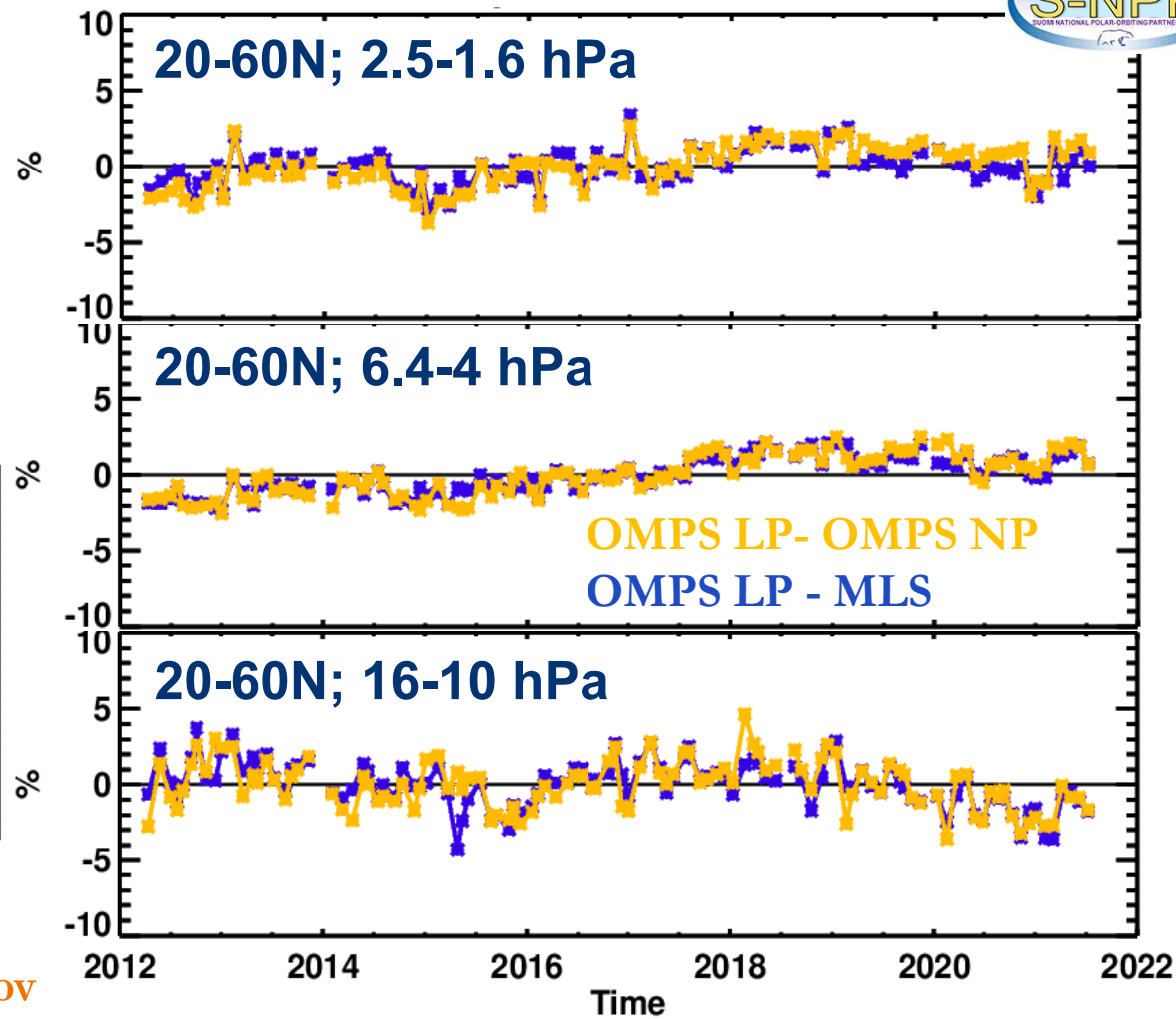
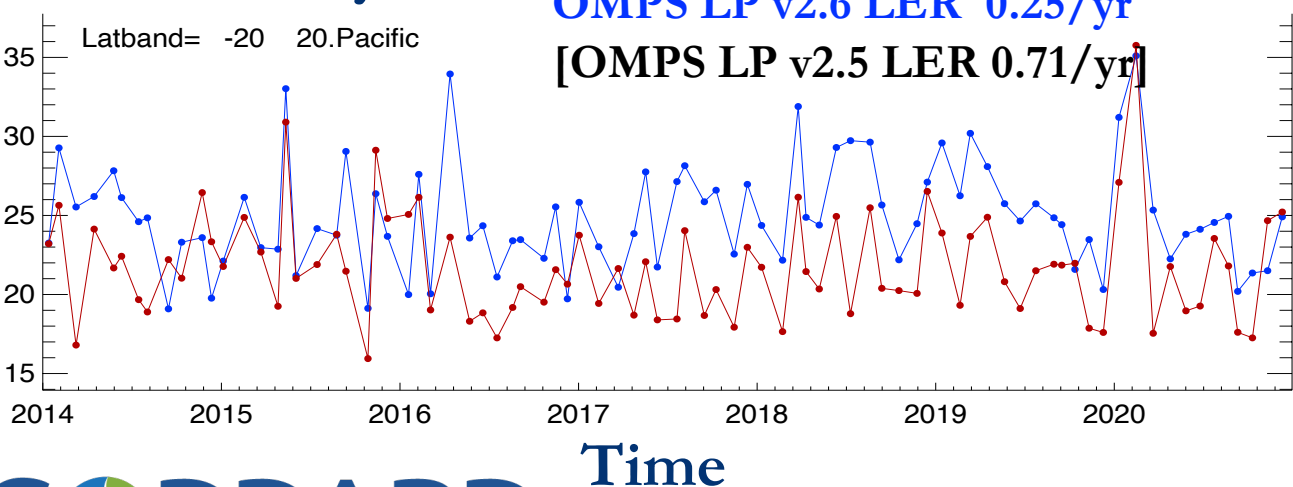


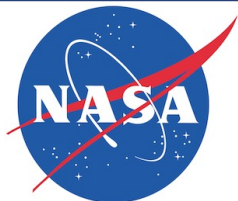
## EVALUATION OF OMPS LP ALTITUDE REGISTRATION USING OMPS NADIR



We use reflectivity and ozone profiles derived from OMPS nadir to evaluate the accuracy of the LP altitude registration.

### Reflectivity

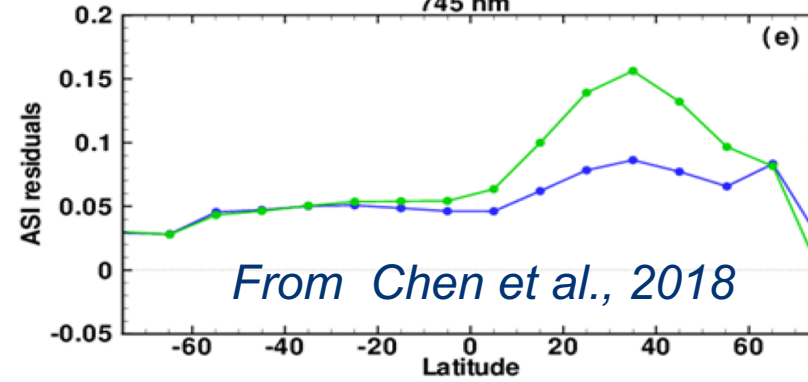
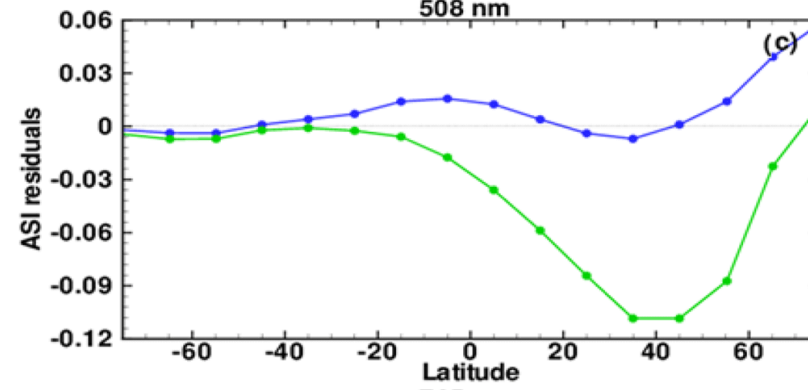
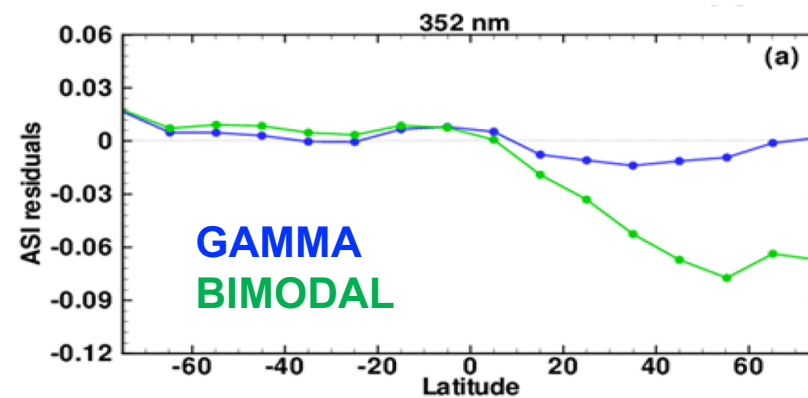
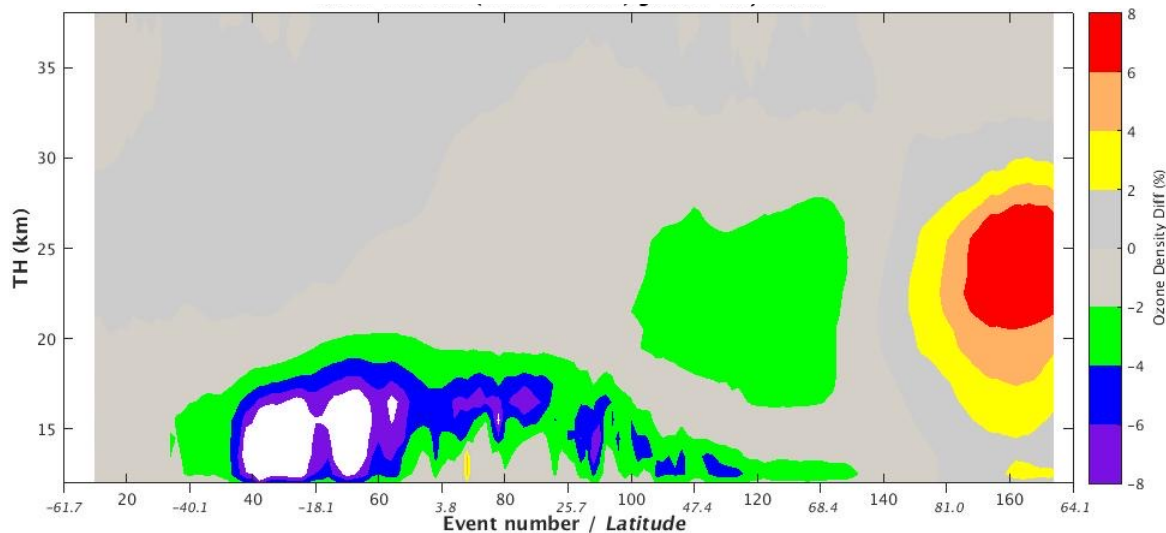




## AEROSOL CORRECTION

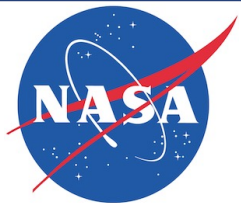
In v2.5 we assumed the bimodal lognormal particle size distribution (PSD) which we replaced with the gamma function PSD in v2.6.

O3 differences, (Gamma-Bimodal)



From Chen et al., 2018



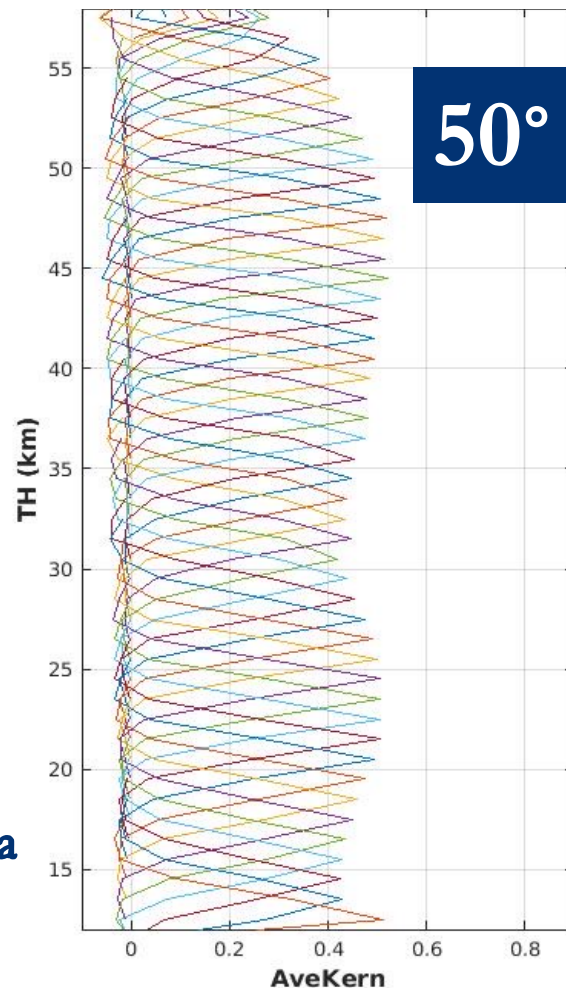


## ALGORITHMIC CHANGES

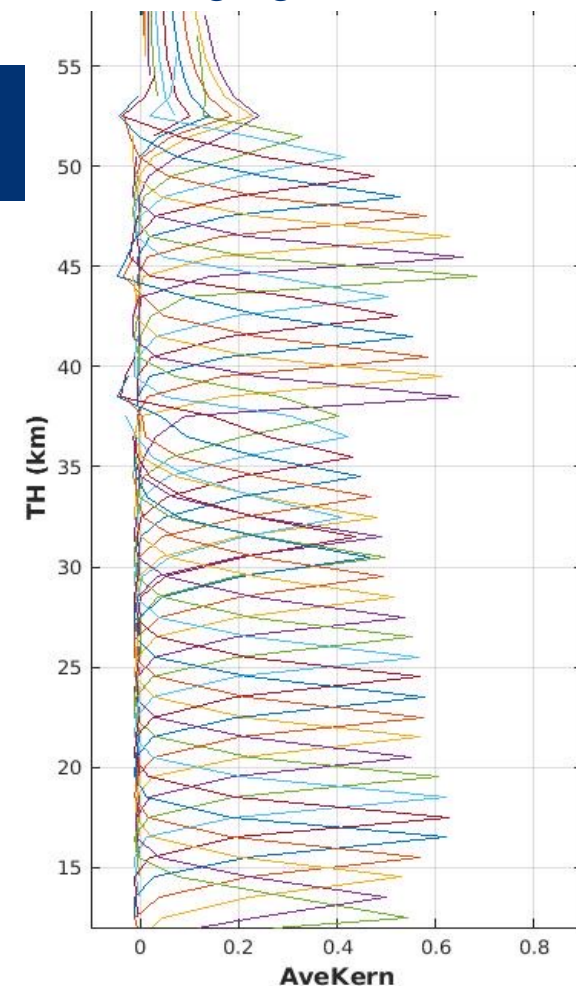
Key changes include:

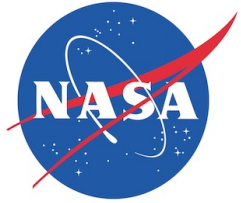
- Merged UV and VIS → combined O<sub>3</sub> profile;
  - Increased number of UV pairs with dynamically controlled contributions;
  - Replaced Optimal Estimation with Tikhonov regularization algorithm.
- The averaging kernels have consistent widths in the stratosphere leading to a vertical resolution of ~ 2 km
  - The estimated precision reduced to 3-4% between 20 and 52 km (compared to 6-8% in V2.5).
  - Updated convergence criteria and quality flags for data screening.

Averaging Kernels v2.6

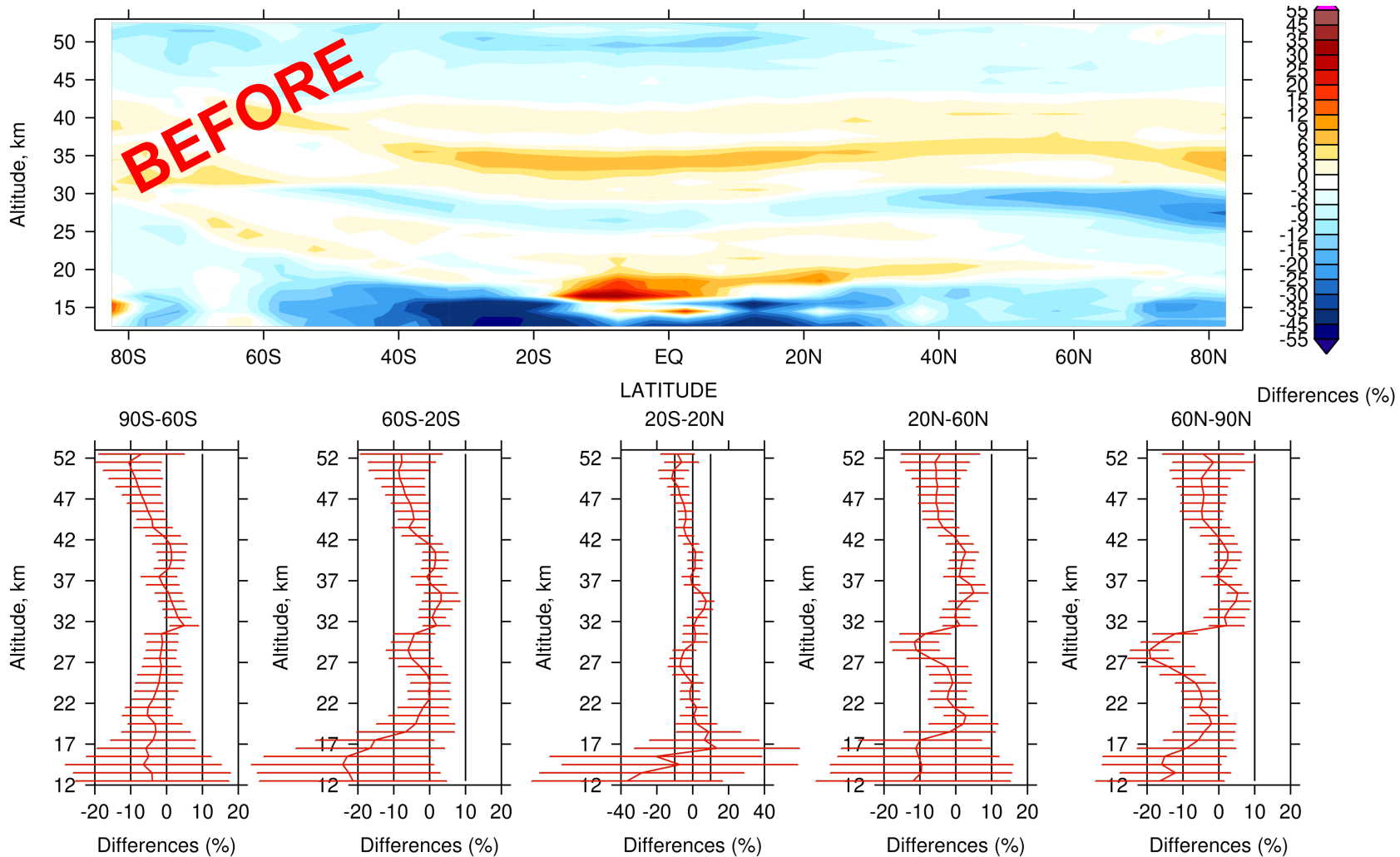


Averaging Kernels v2.5

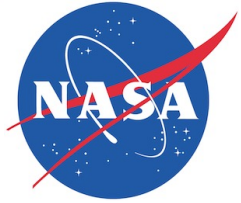




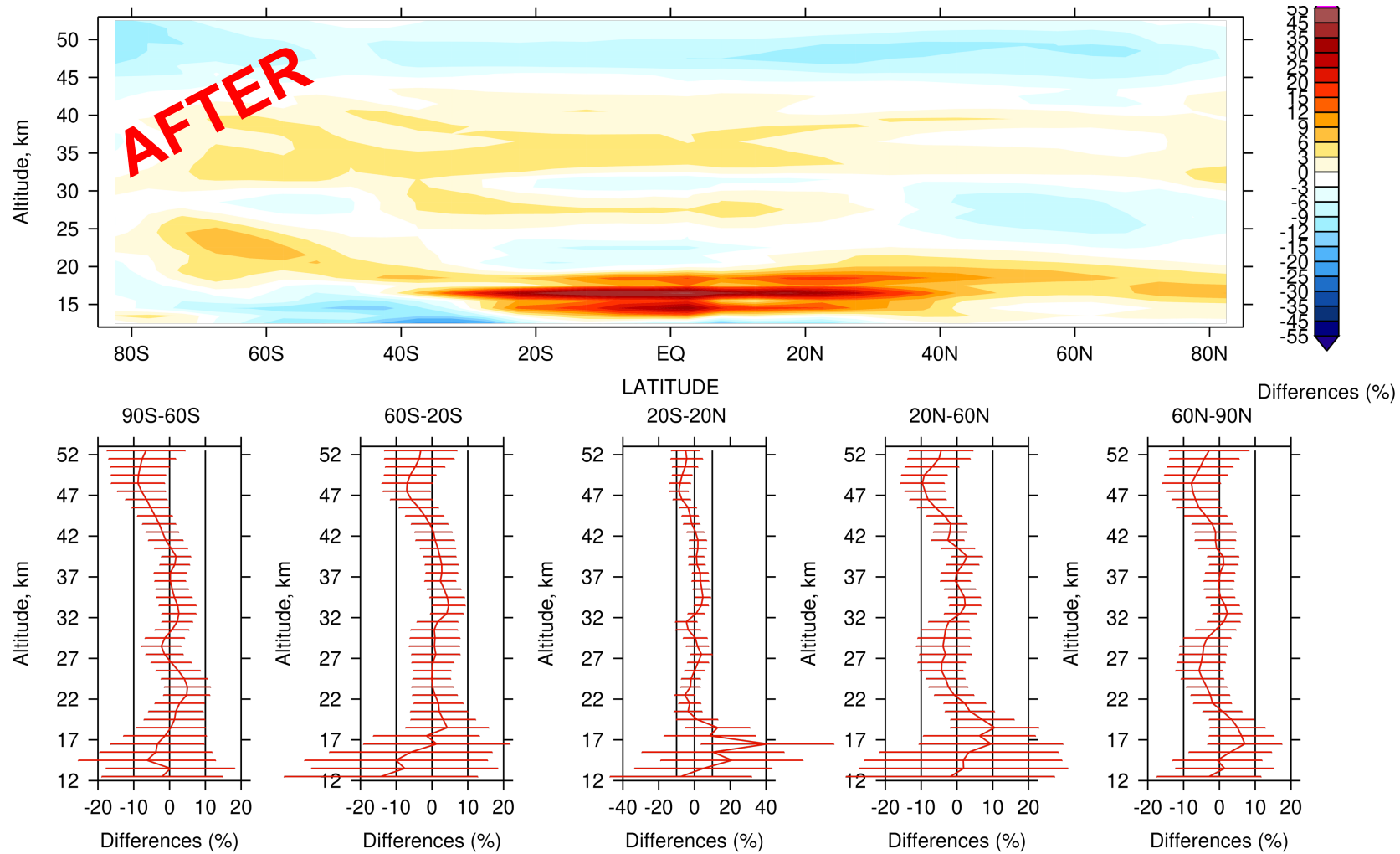
## Mean differences OMPS LP v2.5 and MLS v5

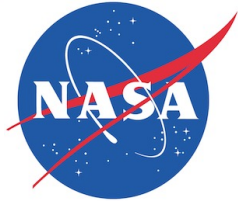


Natalya.a.Kramarova@nasa.gov

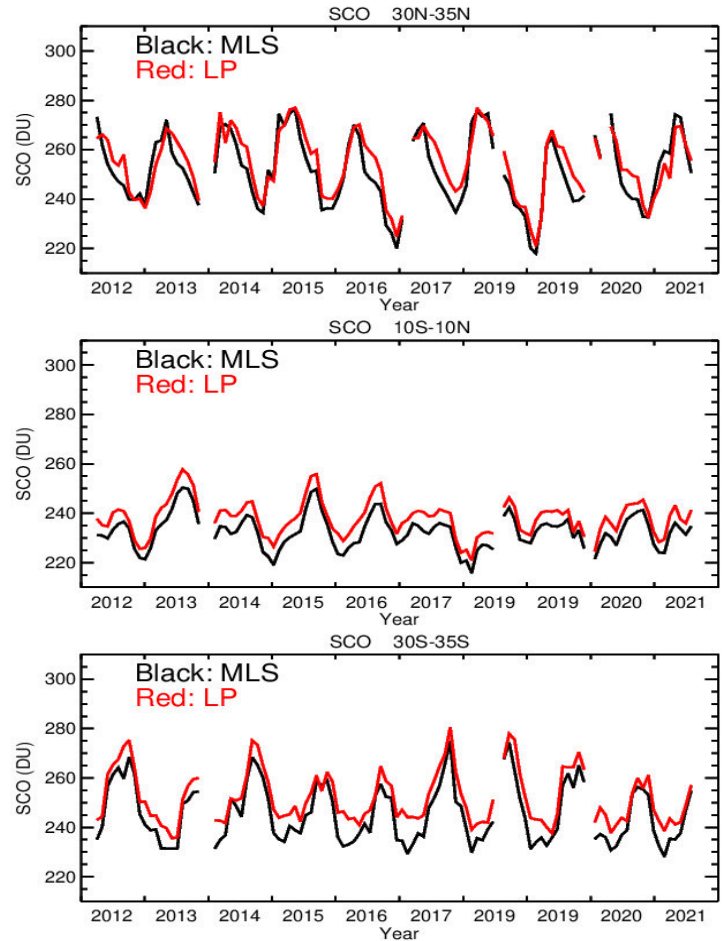


## Mean differences OMPS LP v2.6 and MLS v5



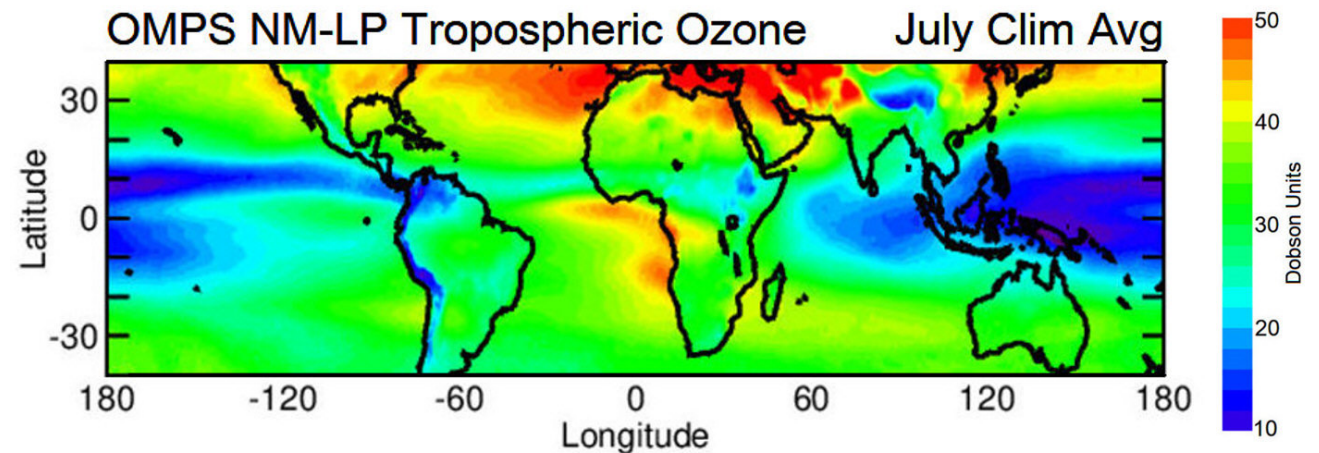


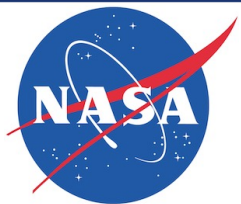
## TROPOSPHERIC OZONE FROM OMPS LIMB AND NADIR



For daily tropospheric ozone maps, stratospheric ozone columns (SCO) from LP are **highly consistent with MLS**

However, the current **12.5 km low altitude cutoff** for LP ozone profiles limits global coverage to about  $\pm 40^\circ$  latitudes





## SUMMARY:

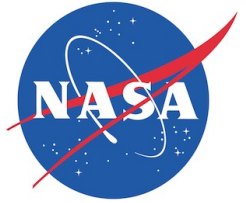
Improvements in v2.6 include:

- ✓ Updated, traceable cross-sections and climatologies for  $O_3$  and  $NO_2$ .
- ✓ Consistent sensitivity and vertical resolution from lower stratosphere to lower mesosphere, realistic precision.
- ✓ Combined UV/VIS profile enables smoother transition between UV and VIS.
- ✓ Substantial reduction in relative drift against MLS.

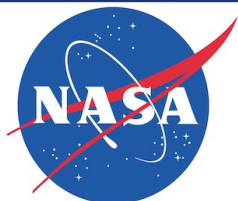
<https://ozoneaq.gsfc.nasa.gov/data/ozone/>

JPSS-2 (NOAA-21) successfully lifted off from Vandenberg Space Force Base on November 10, 2022 at 1:49 a.m.



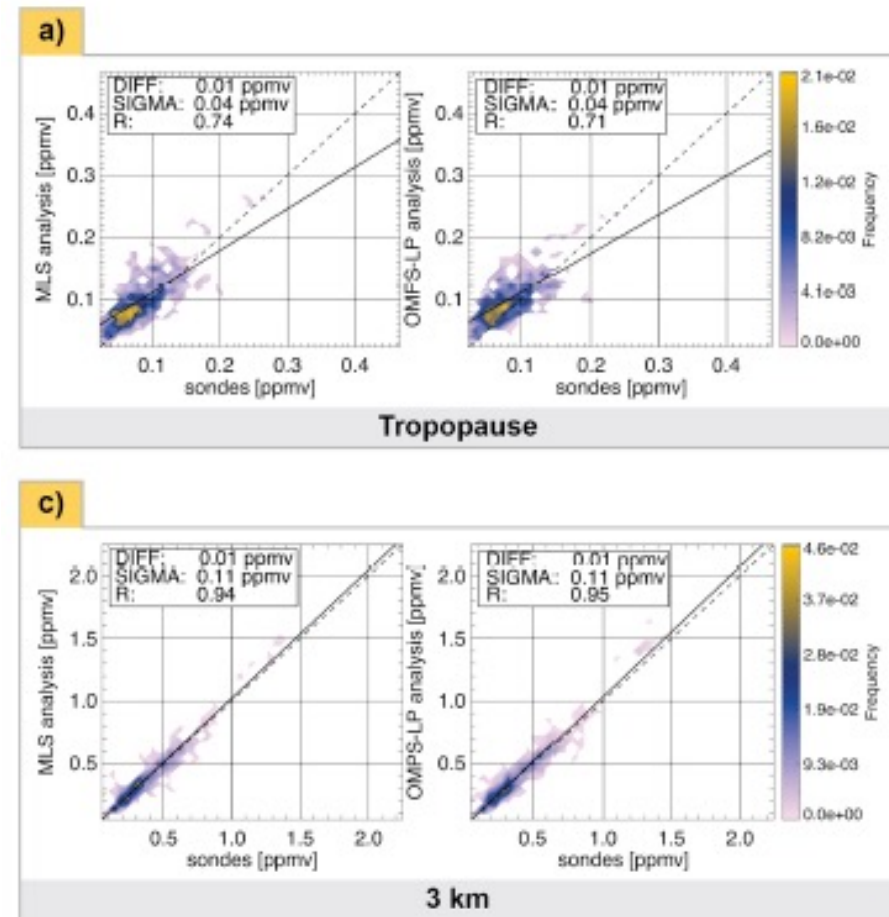


# BACK-UP SLIDES



## ASSIMILATION OF OMPS LP OZONE

- *MLS and OMPS-LP assimilation experiments are in good agreement with independent data and with each other*
- However, further work is needed to reduce a long-term spurious drift in OMPS-LP data

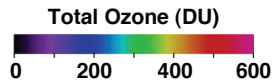
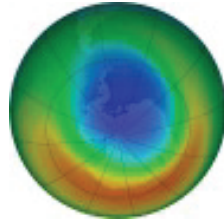


[Wargan et al., JGR, 2020]

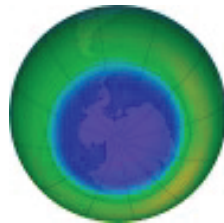
## NASA MONITORING POLAR OZONE DEPLETION WITH THE SUOMI NPP OMPS

- **2019:** strong stratospheric warming weakened the polar vortex in earlier September resulting in the smallest ozone hole since early 1980s.
- **2020:** persistently cold temperatures and strong circum-polar winds led to formation of a large, deep and long-lasting ozone hole.
- Continuation of ozone observations with OMPS instruments on board of Suomi NPP, NOAA-20 and upcoming JPSS-2, -3 and -4 is critical for the next decades to monitor the Antarctic ozone recovery.

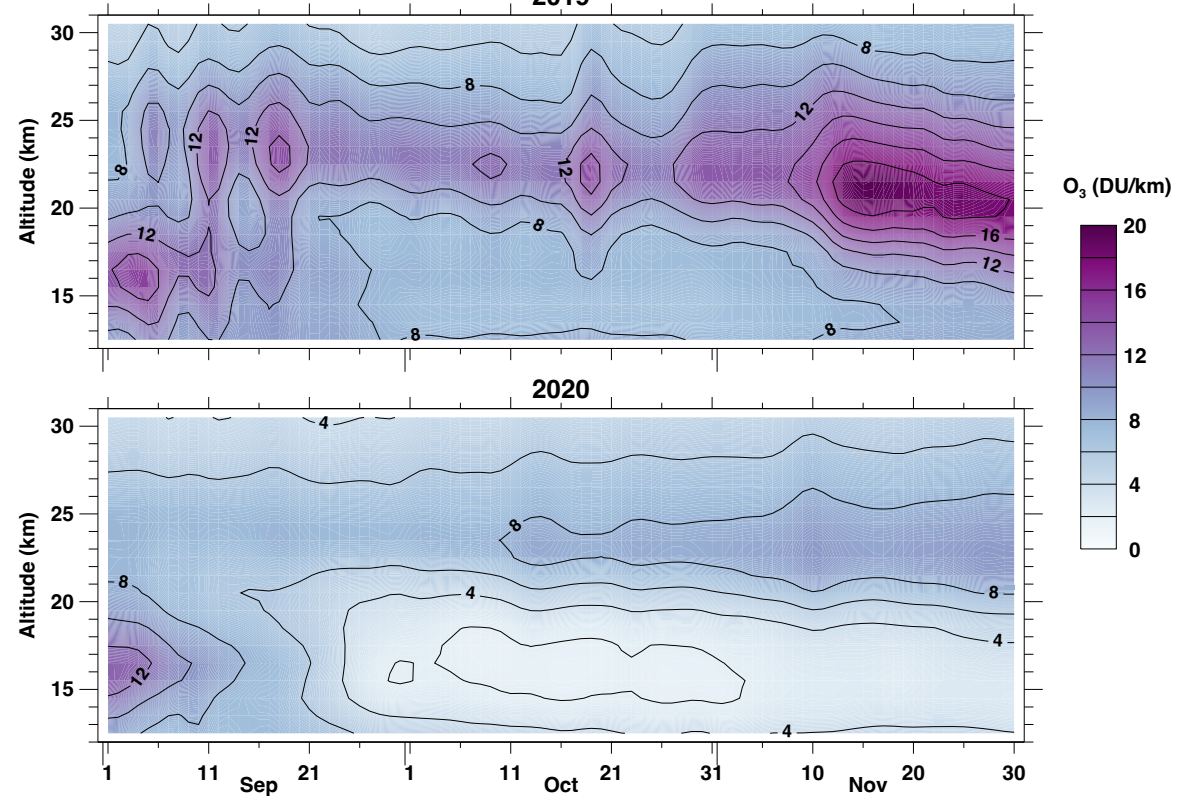
September 2019

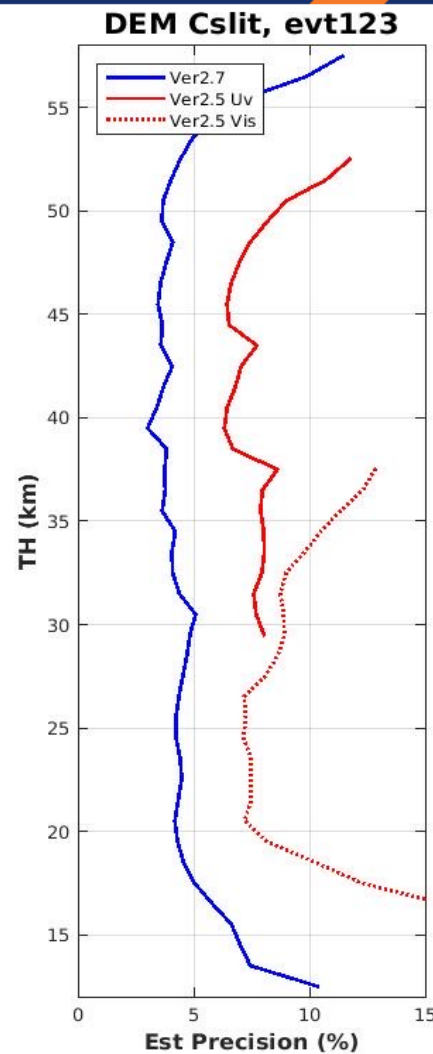
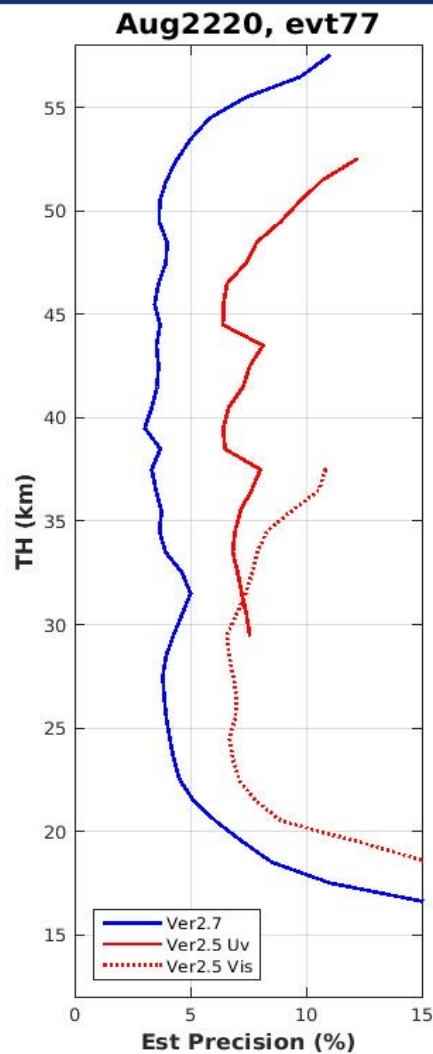
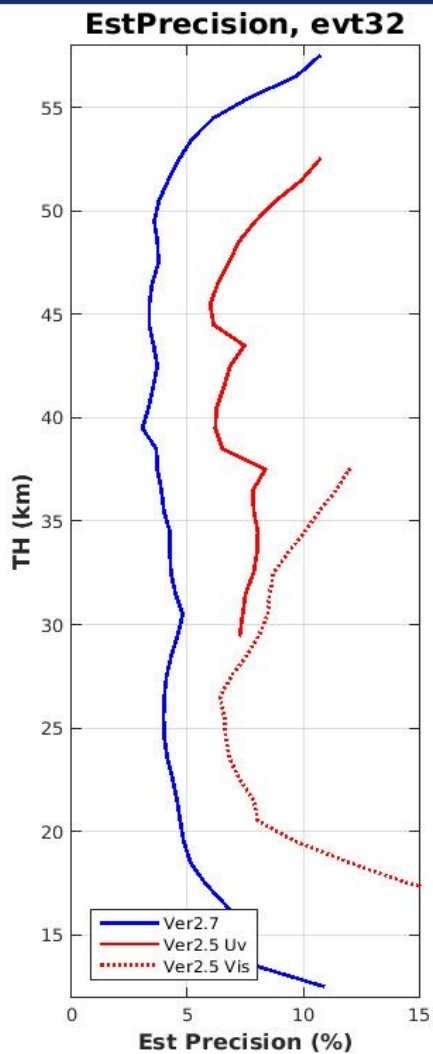
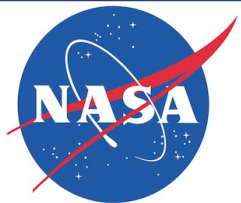


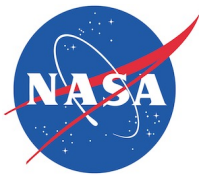
September 2020



OMPS LP ozone, 70–82°S  
2019





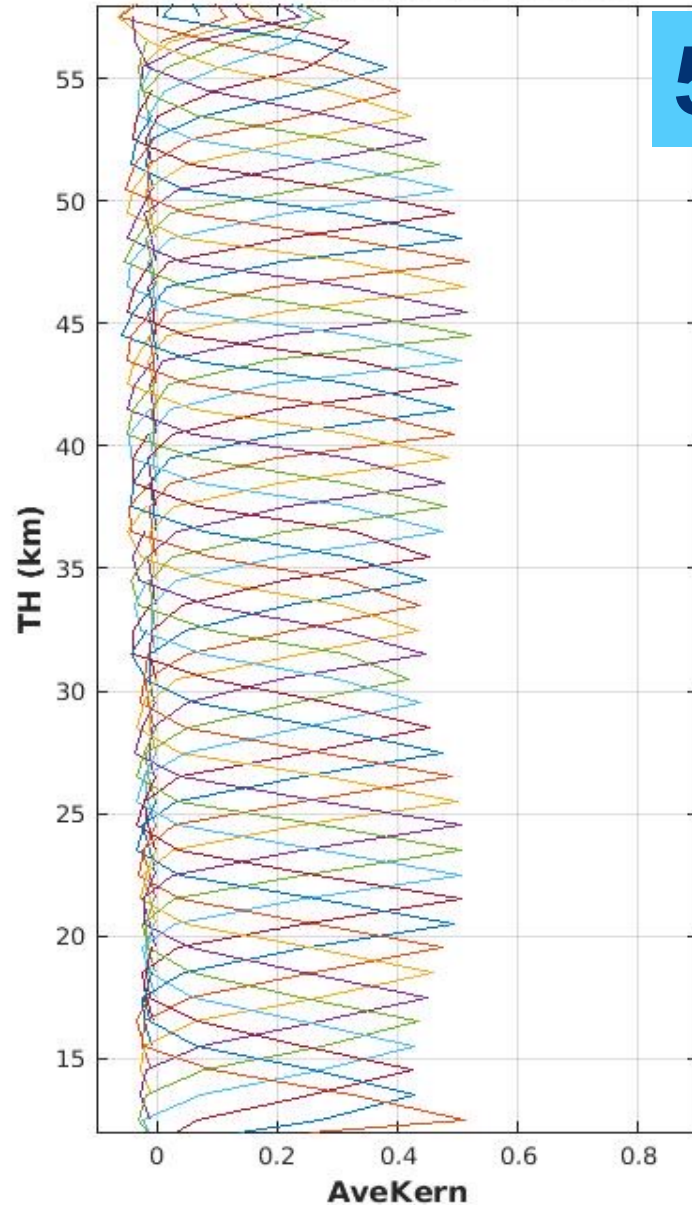


# Algorithmic changes in v2.6

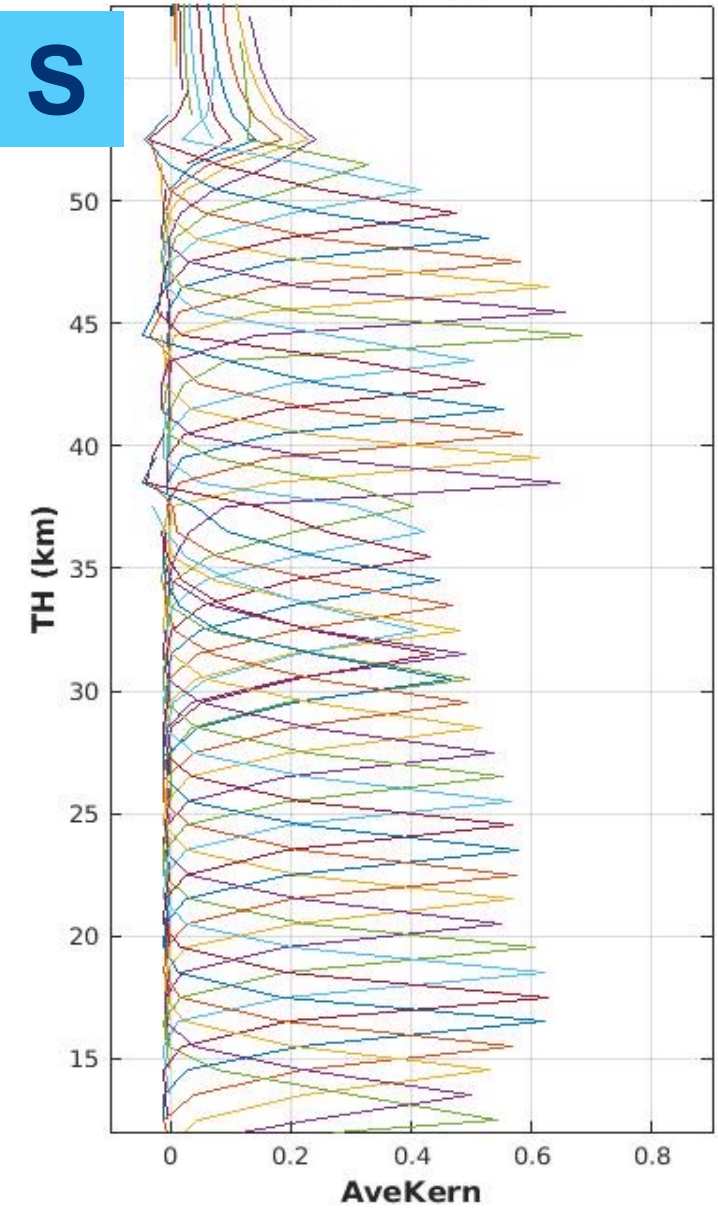


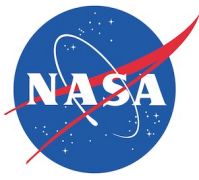
Aug2220, AveK, Ver2.7

CslitEvt32, Ver2.5 Uv Vis



50 S

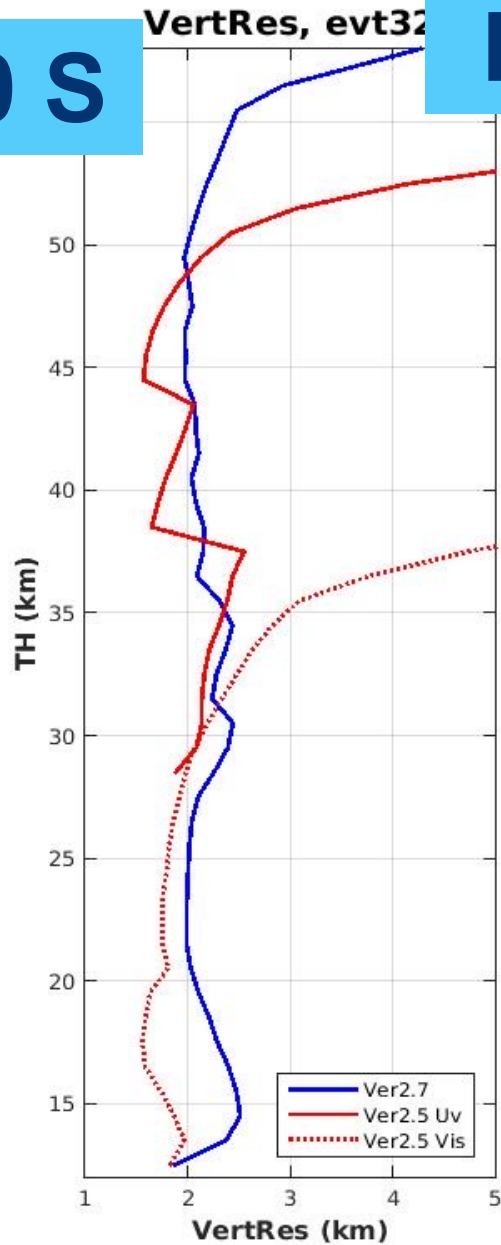




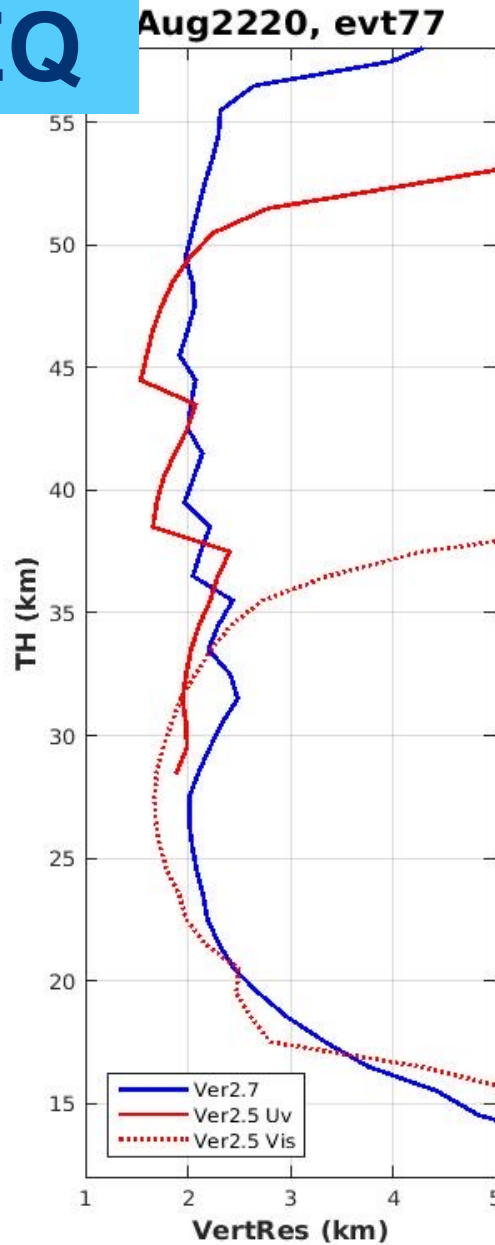
# Algorithmic changes in v2.6



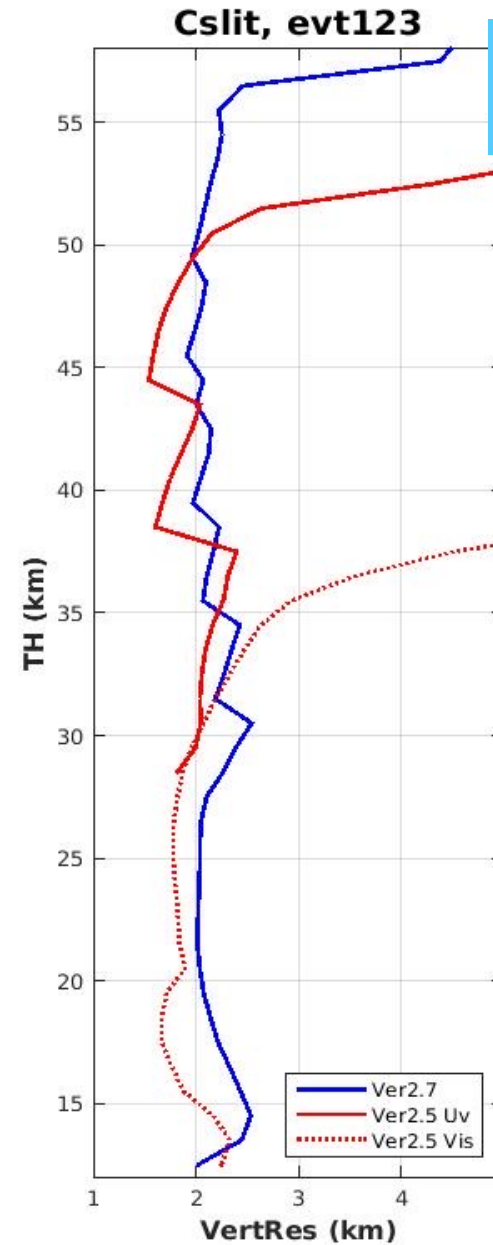
50 S

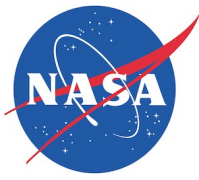


EQ



50 N

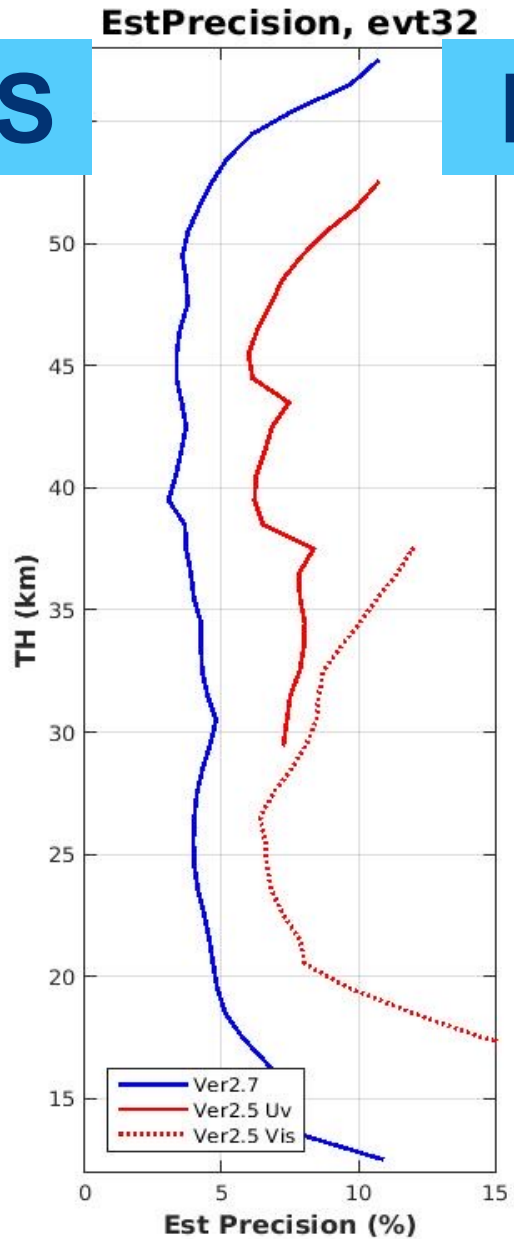




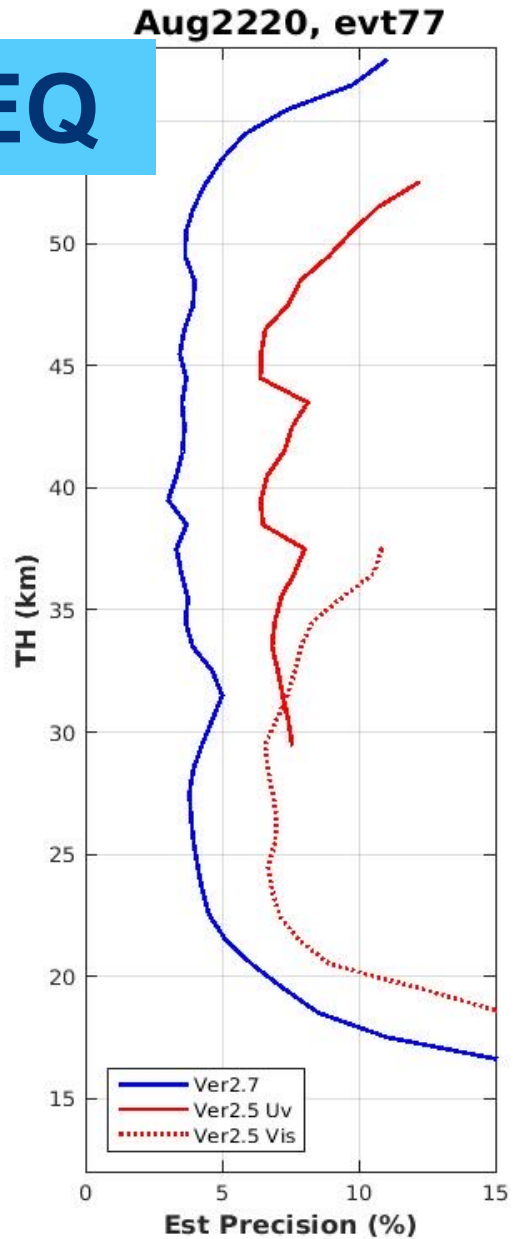
# Algorithmic changes in v2.6



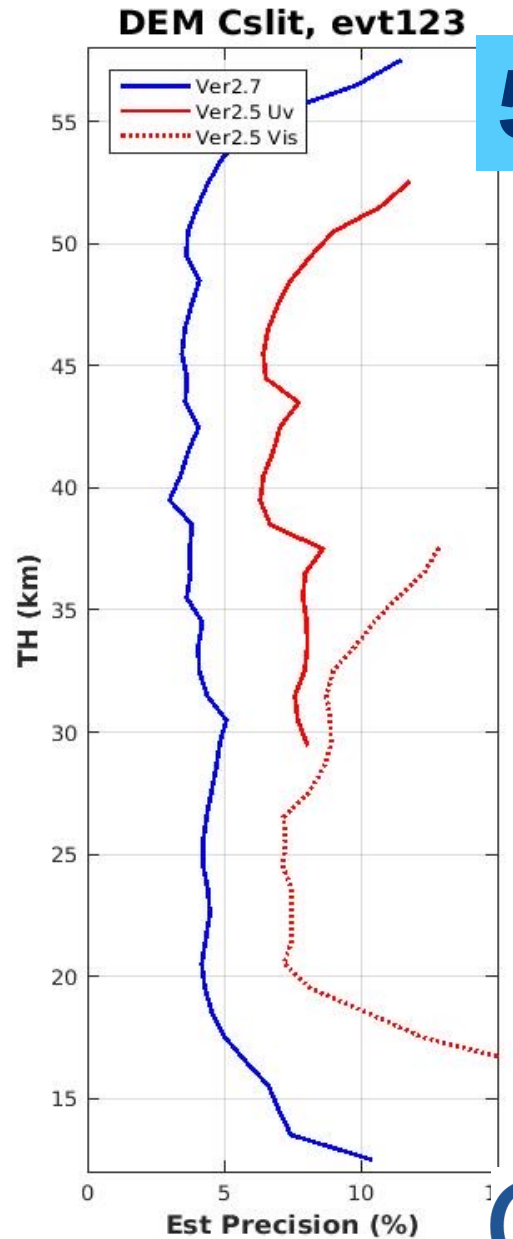
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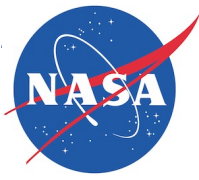


EQ



50 N





# Ozone Profiles. OMPS LP v2.5, v2.6 and MLS v5

Ozone mzm,05/2020, lat37S, Center

Ozone mzm differences [%],05/2020, lat37S



## May 40 S

