



SAGE III /ISS

Stratospheric Aerosol and Gas Experiment

An Earth Science Mission on the International Space Station

An assessment of the SAGE III/ISS temperature and pressure research products

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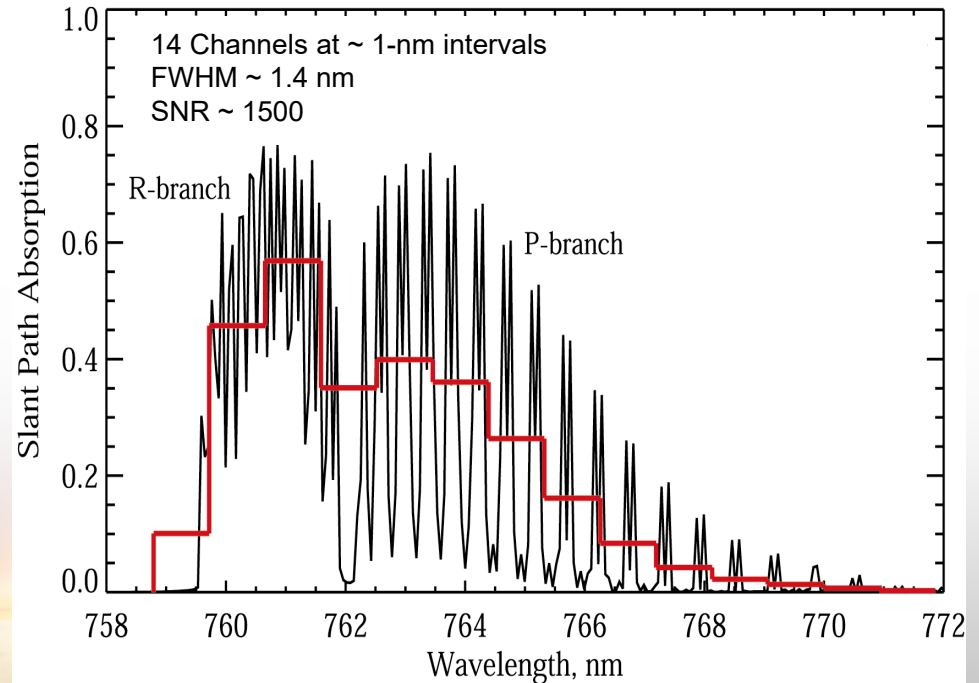
NASA Langley Research Center, Hampton, Virginia



- Iterative approach required due to non-linear nature of problem
- Uses non-linear least squares fitting routine (Levenberg-Marquardt)
 - Simultaneously fits measured A-band spectra from all 14 channels and 90 tangent altitudes
 - Includes O₂ absorption, Rayleigh scattering, aerosol extinction, and O₃ absorption components
 - Solves for successive adjustments to trial T,p profiles by minimizing residuals between measured and modeled absorption

$$\chi^2 = \sum_{i=1}^M \left[\frac{A_i^m - A_i^c(\mathbf{a})}{\sigma_i} \right]^2$$

- Includes non-rigid hydrostatic constraint



A-band contains ~290 individual absorption lines with the distinctive R-branch and P-branch structure. Broad features are still visible at SAGE III resolution (red line)

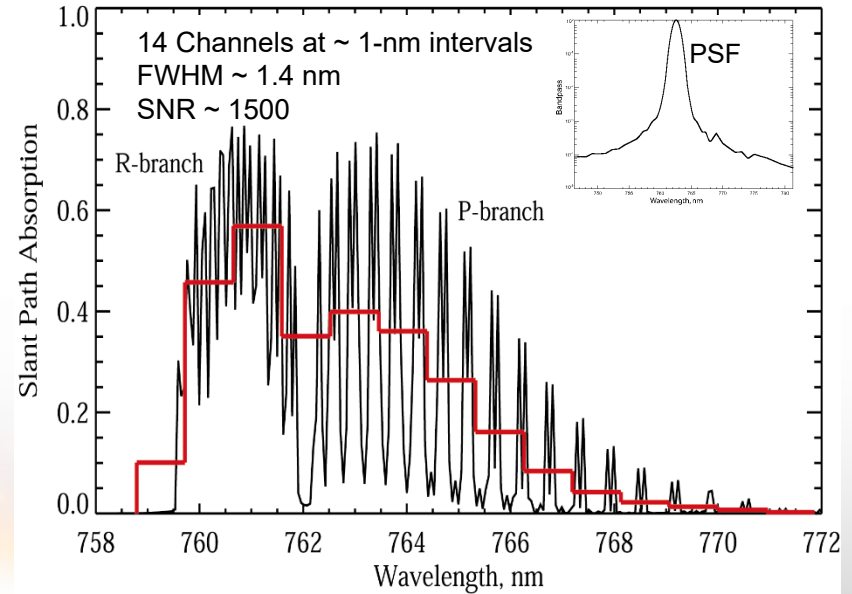
➤ **Forward model simulates SAGE III/ISS LOS transmission measurements in 14 O₂ A-band channels using LBL radiative transfer calculations**

➤ **Inputs:**

- Initial guess T/p profiles
- Pathlength matrix
- O₂ Spectroscopic parameters (HITRAN 2016)
- Initial guess aerosol (linear fit: slope + intercept)
- Initial guess ozone (cross sections + ozone number density)
- Channel wavelength map
- Channel point spread function (PSF)

➤ **Outputs**

- LOS transmission (O₂ + Rayleigh + aerosol + O₃)



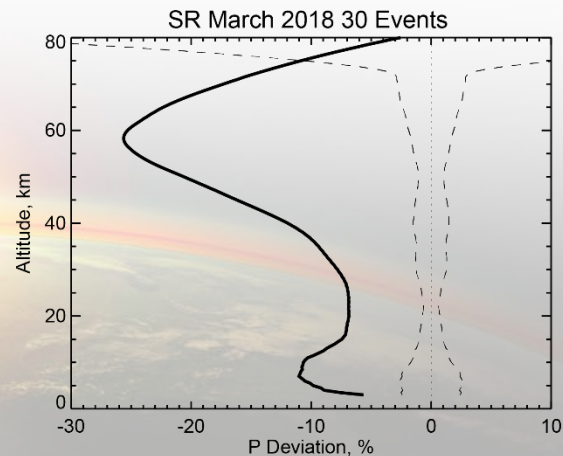
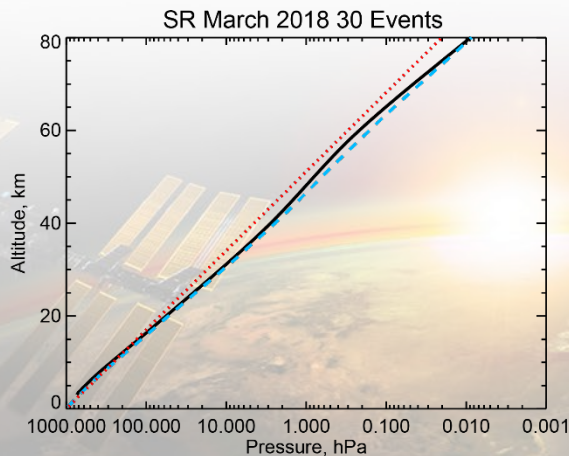
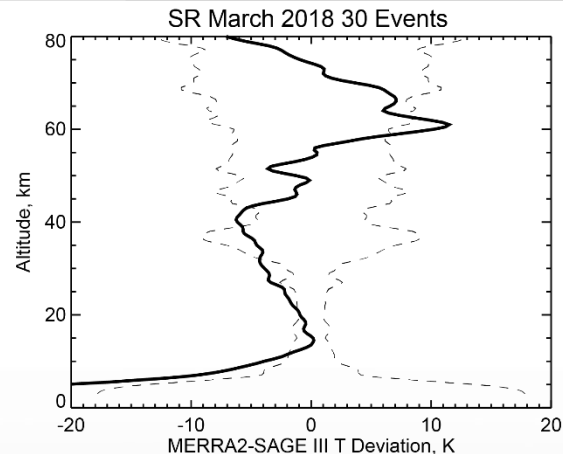
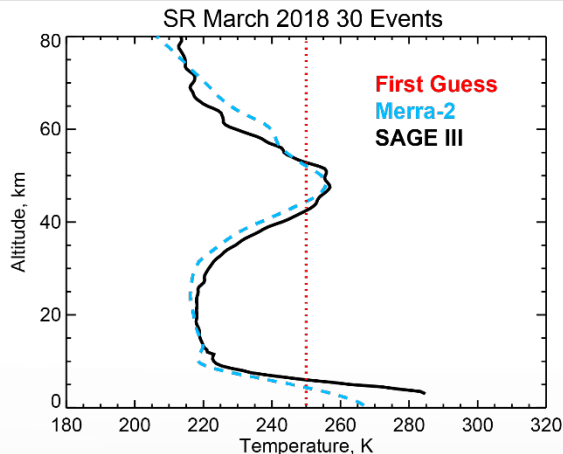
A-band contains ~290 individual absorption lines with the distinctive R-branch and P-branch structure. Broad features are still visible at SAGE III resolution (red line)



Baseline Retrievals (Sunrise Events)

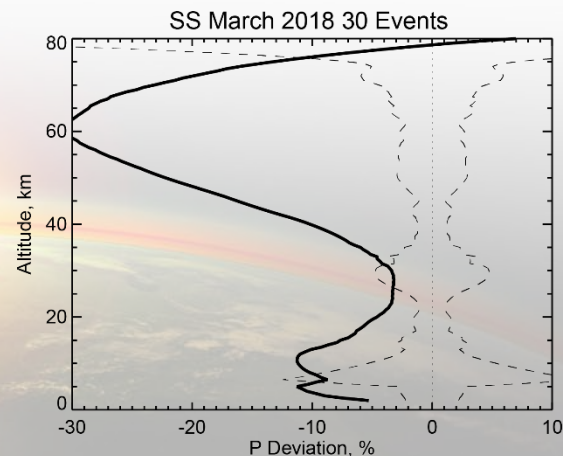
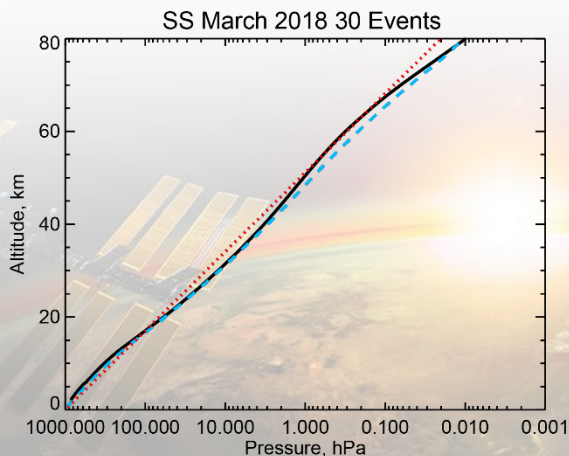
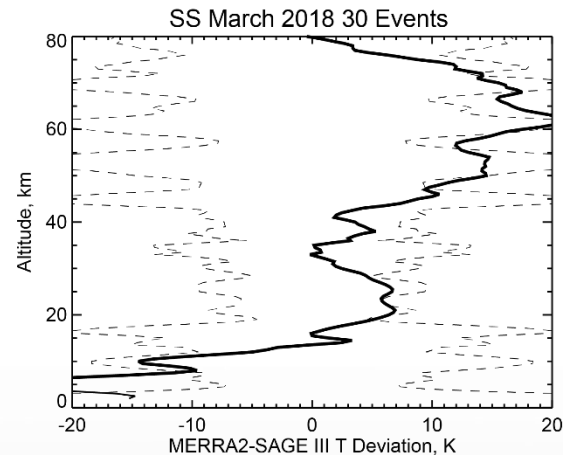
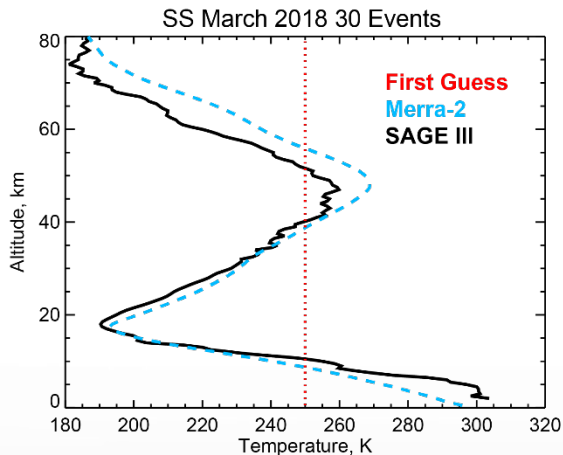


Mean SAGE III/ISS retrieved temperature and pressure profiles from 30 sunrise events during March 2018 with baseline wavelength map and PSF factor

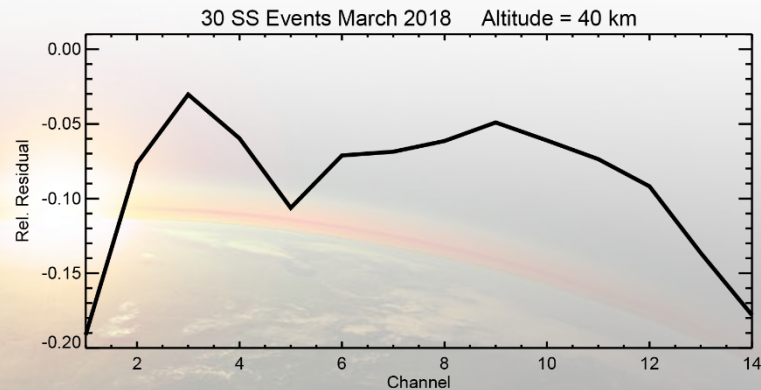
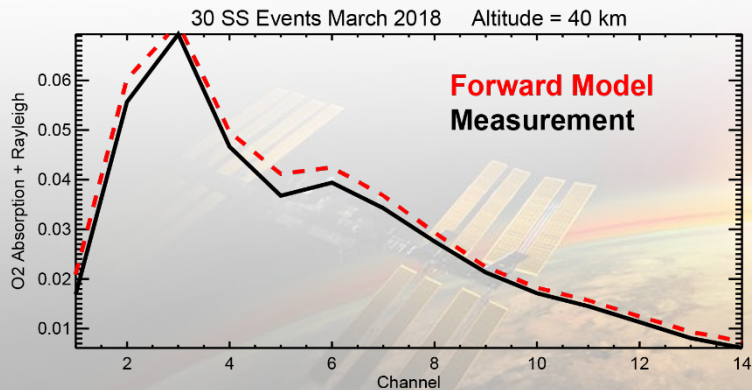
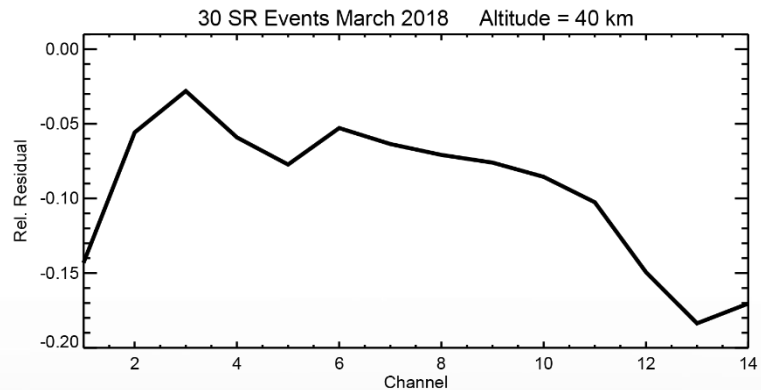
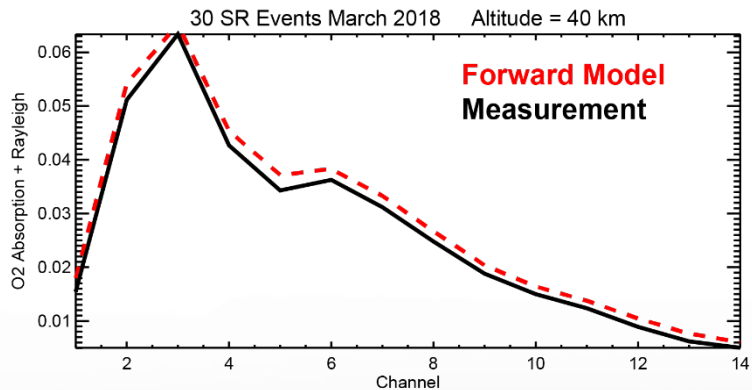


Baseline Retrievals (Sunset Events)

Mean SAGE III/ISS
retrieved temperature
and pressure profiles
from 30 sunset events
during March 2018
with baseline
wavelength map and
PSF factor

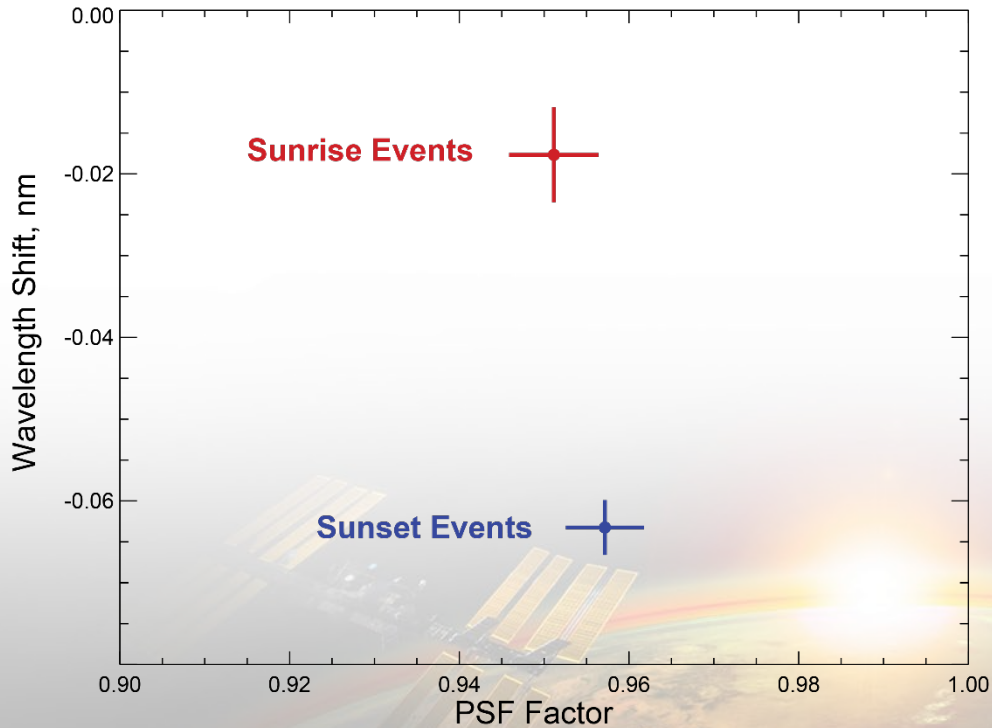


Baseline wavelength registration and instrument response functions

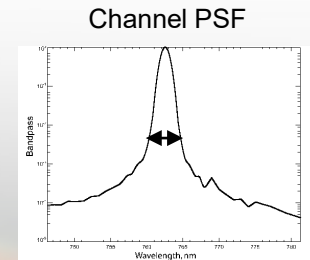


➤ **Modeled spectra too broad and misregistered in wavelength**

Wavelength and PSF Factor Adjustments



- Assume Merra-2 T/p profiles are truth
- Retrieved adjustments to baseline wavelength map and PSF widths (mean $\pm 1 \sigma$) required to best fit measured spectra

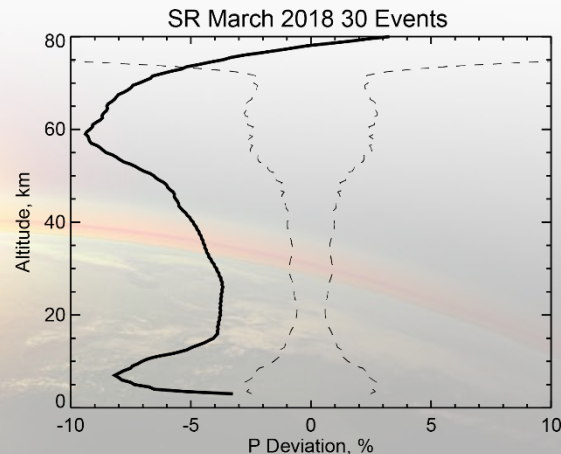
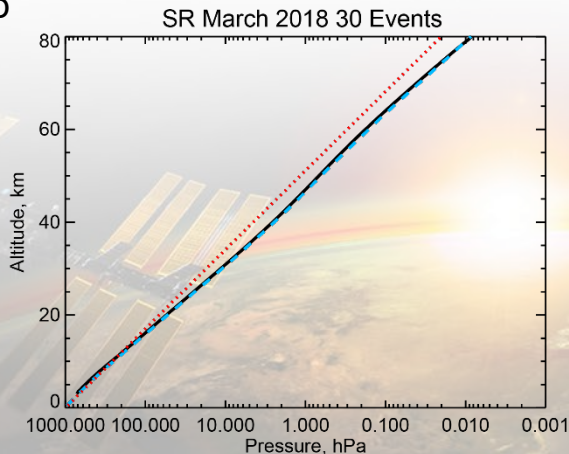
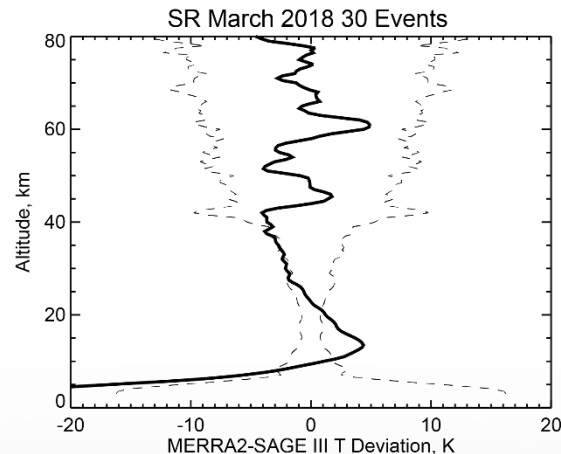
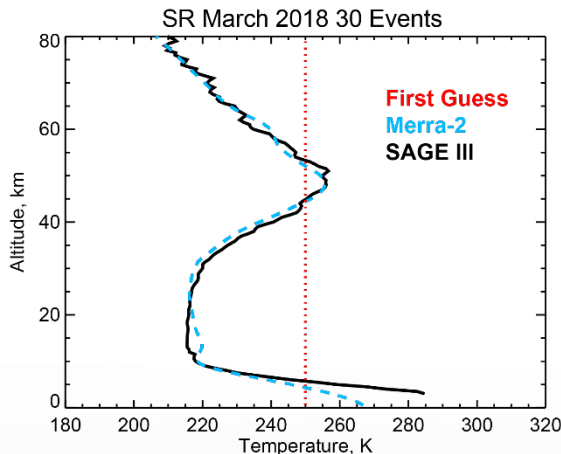




Retrievals Using Adjusted Wavelength & PSF



Mean SAGE III/ISS retrieved temperature and pressure profiles from 30 sunrise events during March 2018 with adjusted wavelength map (-0.018 nm) and PSF factor (0.95)

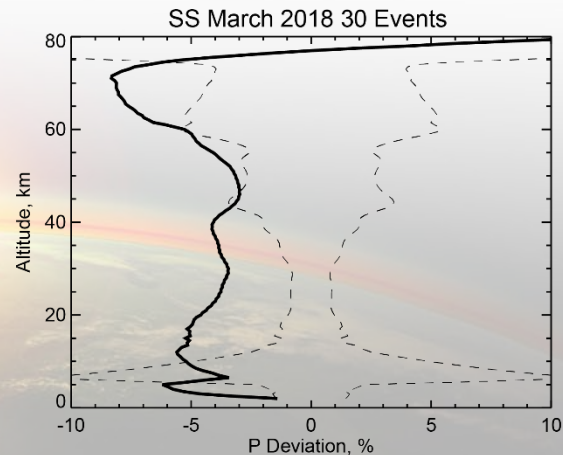
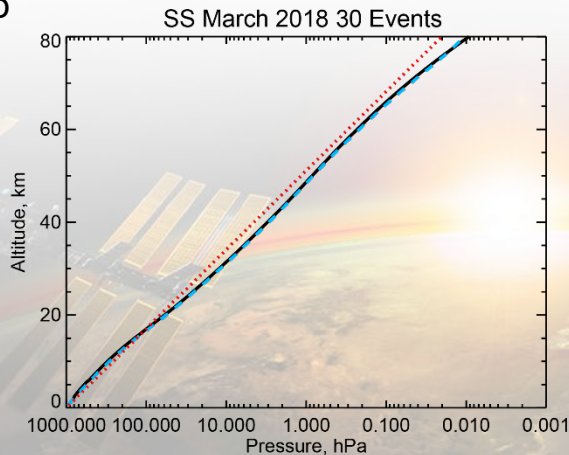
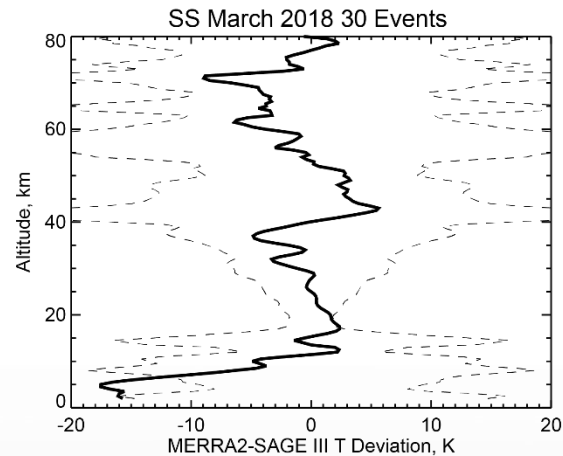
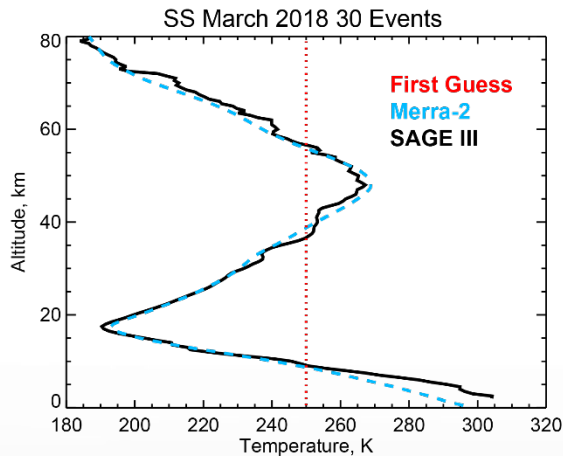


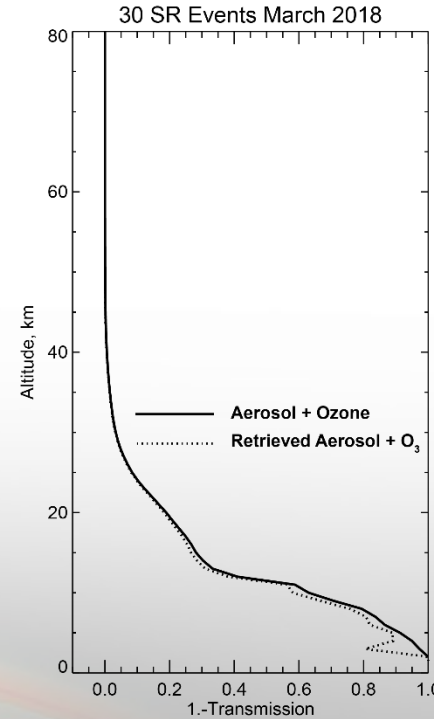
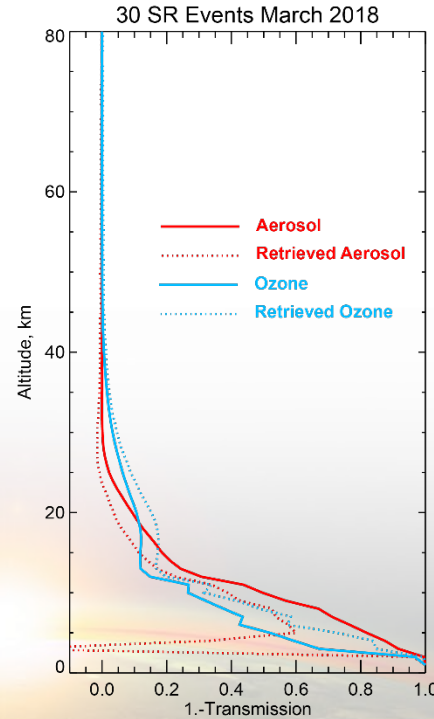
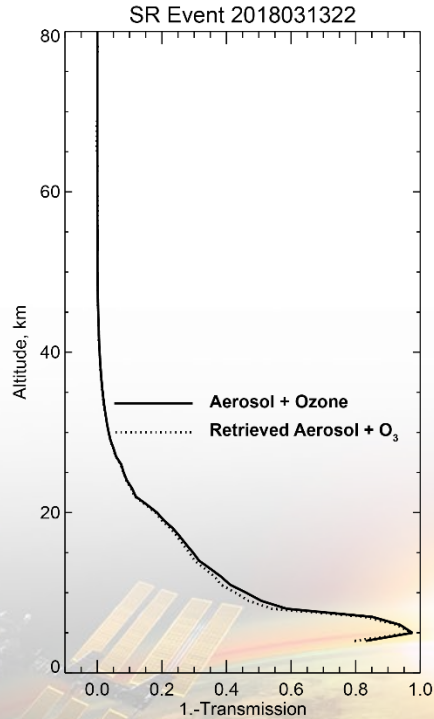
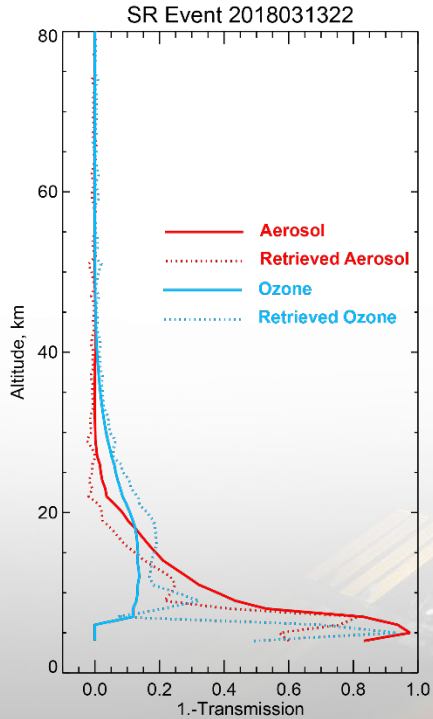


Retrievals Using Adjusted Wavelength & PSF



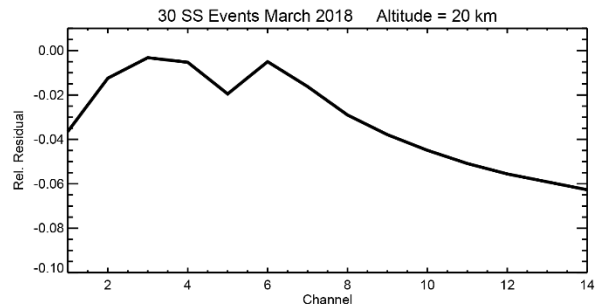
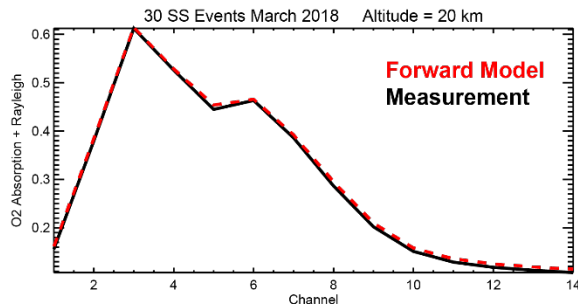
Mean SAGE III/ISS retrieved temperature and pressure profiles from 30 sunset events during March 2018 with adjusted wavelength map (-0.063 nm) and PSF factor (0.95)



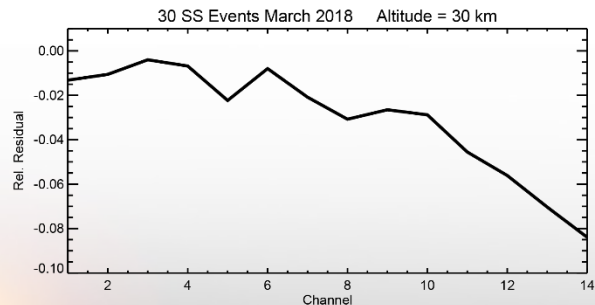
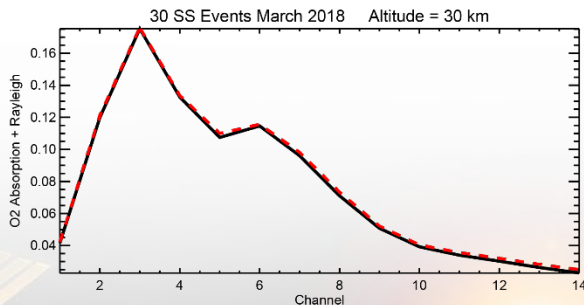


Retrieved aerosol and ozone components compared with measurements (aerosol interpolated from aerosol channels and AO3 ozone product)

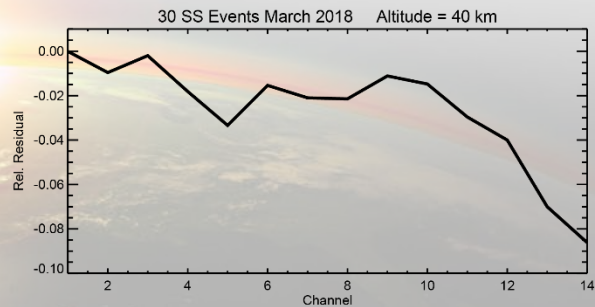
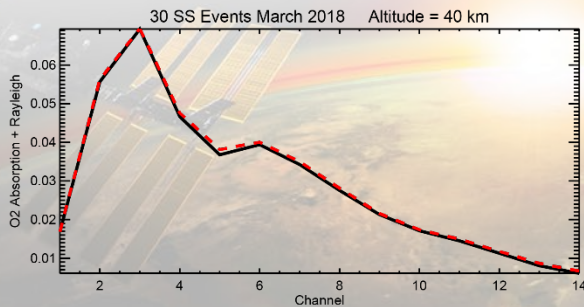
- Mean measured vs. modeled O₂ absorption + Rayleigh for 30 SS events



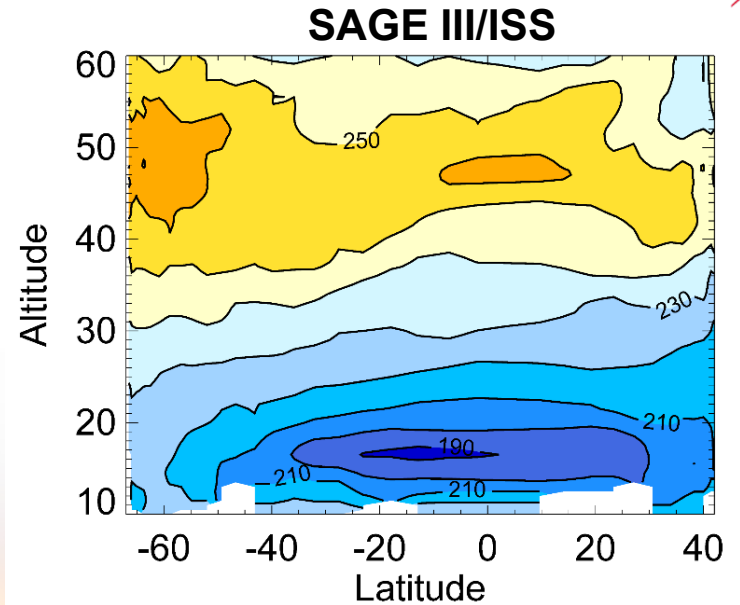
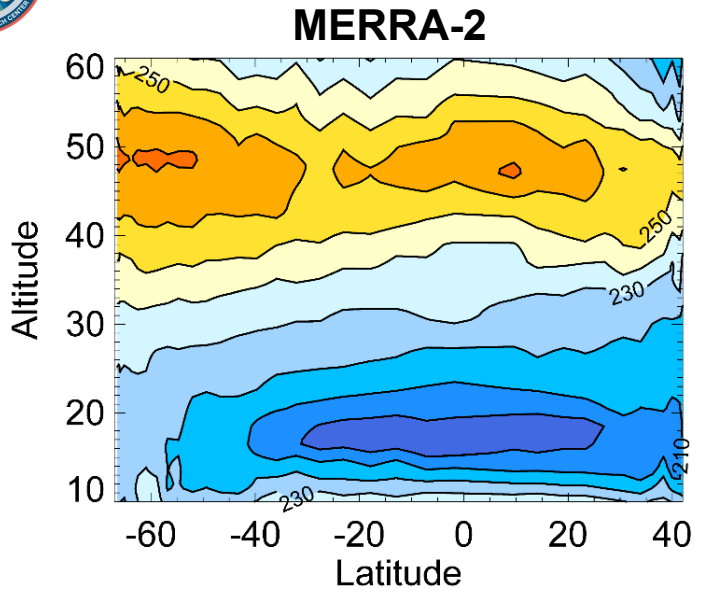
- Forward model parameters:
- $\Delta\lambda = -0.063$
- PSF Factor = 0.95



- Systematic shape of residuals may indicate that some component is not being properly fitted



SAGE III/ISS vs. Merra-2 zonal cross sections



MERRA-2 (left) and SAGE III/ISS (right) meridional temperature distributions compiled from sunrise events over the period 1 January -15 February 2018. The SAGE III/ISS temperature distribution captures the general temperature structure seen in MERRA-2, but is colder at the tropopause and stratopause, and somewhat noisier at higher altitudes.

- **Forward Model Assessment:**
 - Small wavelength shifts are apparent in A-band spectra relative to baseline registration (approximately -0.018 nm for sunset and -0.063 nm for sunrise events)
 - Baseline PSFs too broad- fits are improved using narrower (0.95 x baseline) widths
- **Retrievals with adjusted forward model parameters look encouraging with *near science quality* temperature products, but some issues remain, especially in pressure product**
- **Future Work:**
 - Continue examining sensitivity to forward model parameters
 - Bandpass (width, wings, etc.)
 - Spectroscopy (line intensities, line shape)
 - Potential altitude offset
 - Aerosol + O₃ components
 - Exclude weaker channels
 - Explore alternative retrieval approaches to L-M (e.g., Newton-Raphson)
 - Perform more detailed comparisons with correlative measurements (e.g., radiosondes, lidar, MLS, and global analyses) to quantify precision and accuracy of T/p products