

Rapid succession of SEP events associated with a series of EUV jets: Solar Orbiter, STEREO-A and near-Earth spacecraft observations

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T. S. Horbury¹², L. Rodriguez¹³, A.N. Zhukov¹³

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(3) Universidad de Alcalá

(4) The Johns Hopkins University. Applied Physics Laboratory

(5) University of Maryland, Baltimore County

(6) University of Maryland, College Park

(7) North-West University, Potchefstroom

(8) University of Turku

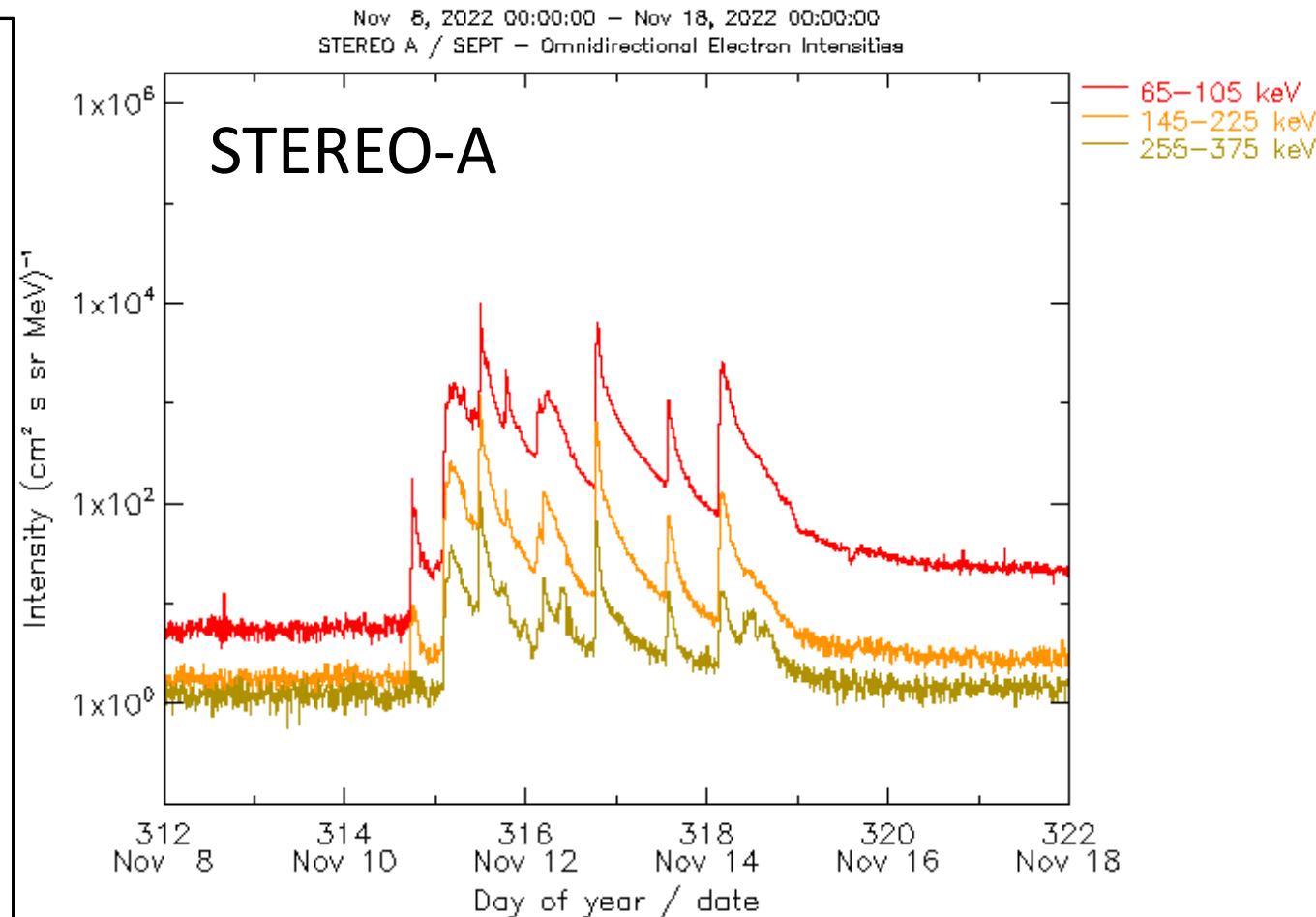
(9) Southwest Research Institute

(10) Christian-Albrechts-Universität zu Kiel

(11) The Catholic University of America

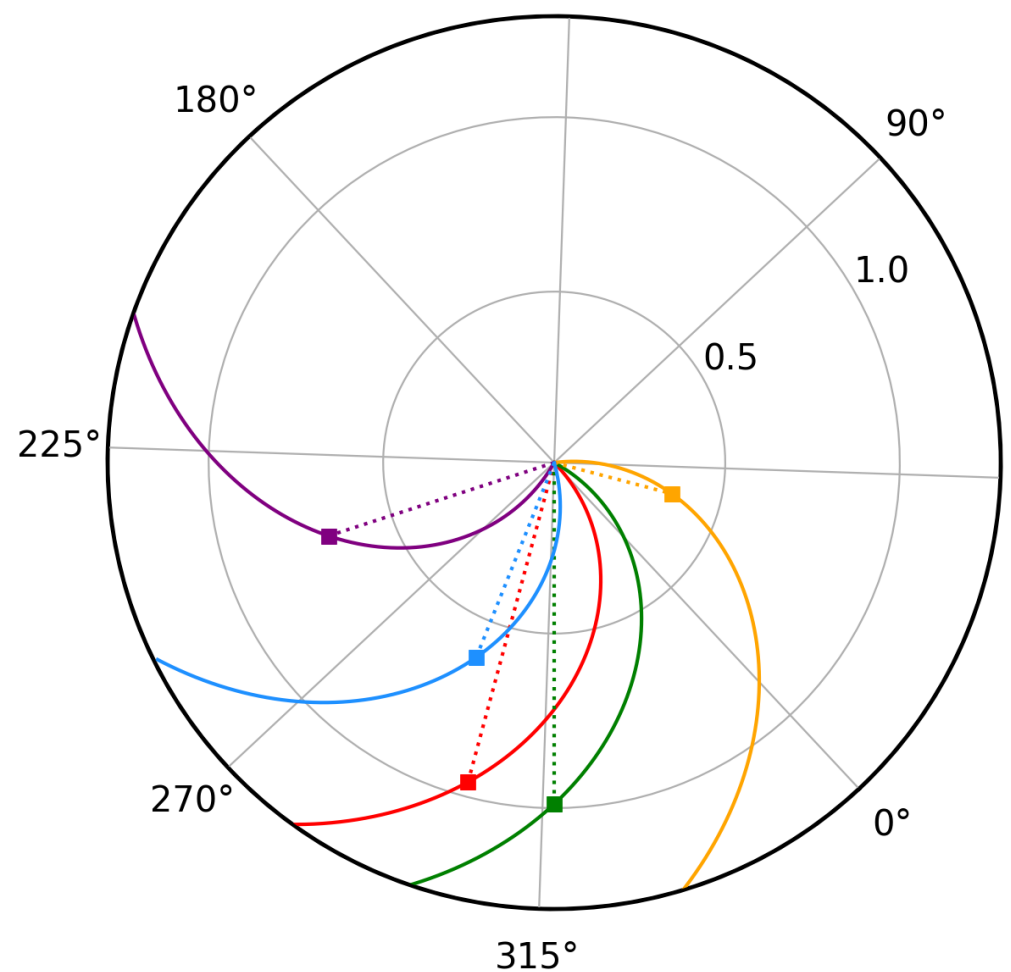
(12) Imperial College London

(13) SIDC, Royal Observatory of Belgium



Outline

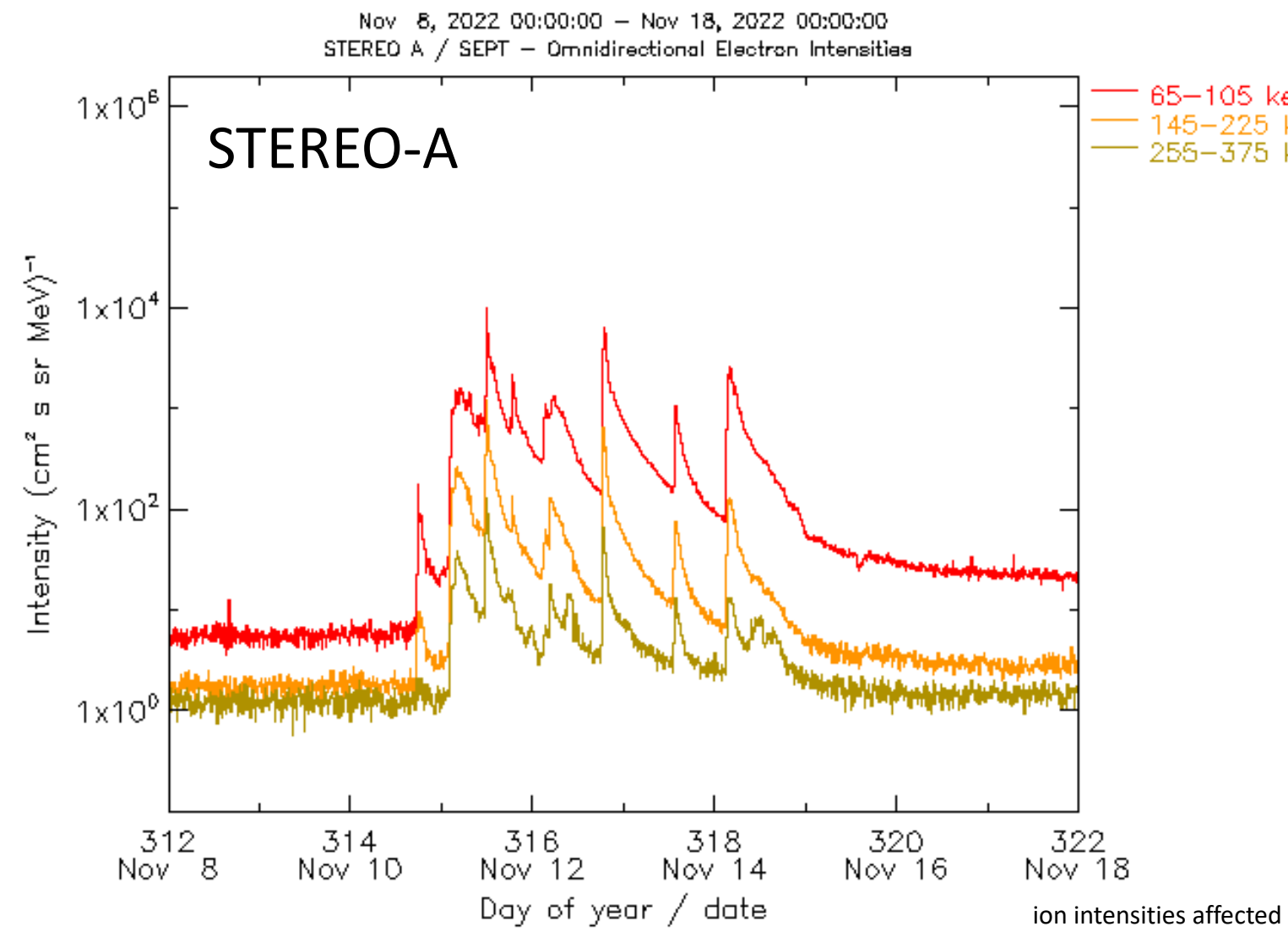
- Solar origin of these events
- The relationship between the multiple signatures of the parent solar eruptions (EUV jets, type III radio bursts, X-ray emissions, ...)
- Ion composition signatures of these events.
- Multi-spacecraft observations. Interplanetary context affecting the particle transport and spacecraft magnetic connection



9-14 Nov 2022 sequence of events

- STEREO A
- Earth
- BepiColombo
- Parker Solar Probe
- Solar Orbiter

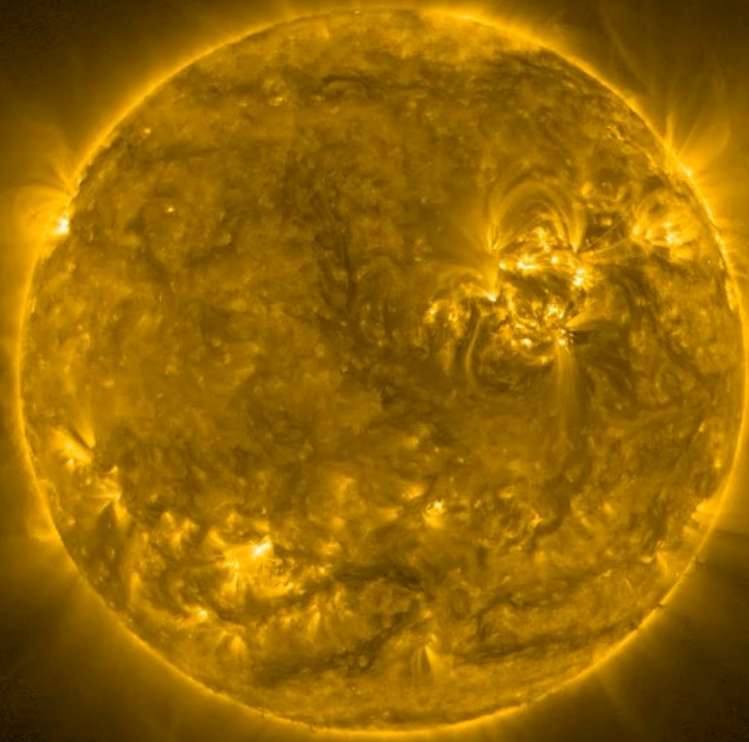
PSP ~0.69 AU -72 deg
 SoIo ~0.62 AU -21 deg Lat=7.89 deg
 STA ~0.96 AU -15 deg Lat=4.91 deg
 Earth ~0.99 AU 0 deg Lat=3.34deg



Solar Orbiter/EUI FSI 174

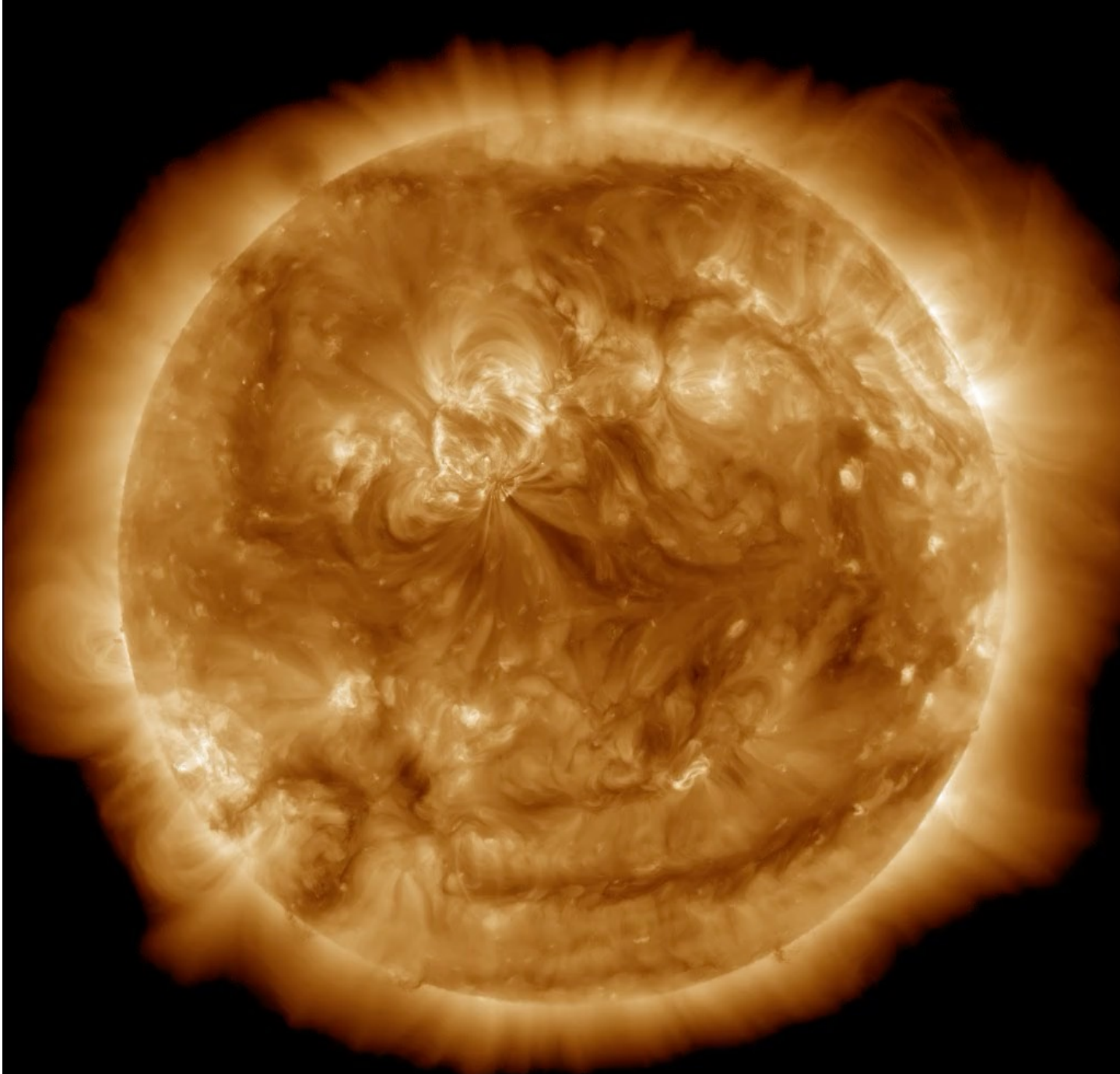
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crota -0.913 (deg) dsun_au 0.620 (AU)
crval1,2 88.42,123.80 (arcsec) v107_20221130_002+flown

2022-11-11 07:01:00 (UTC)



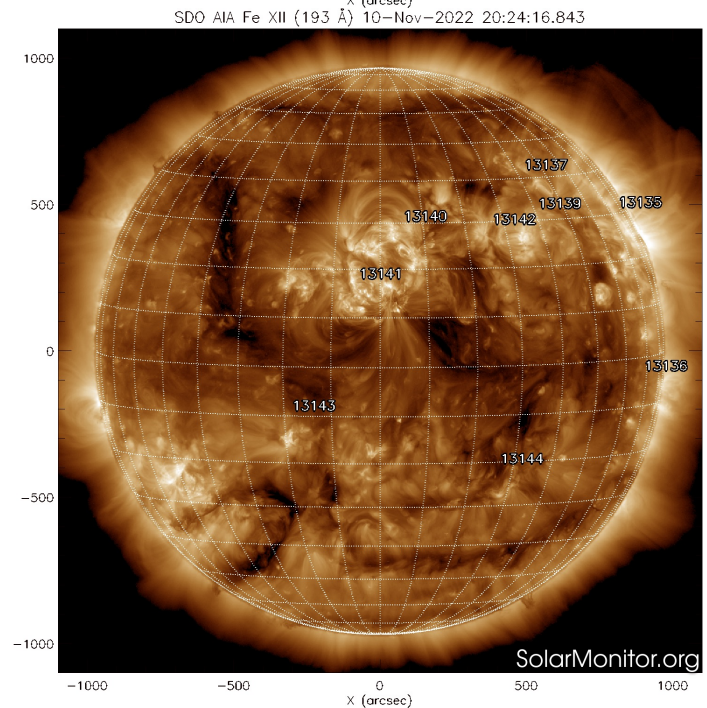
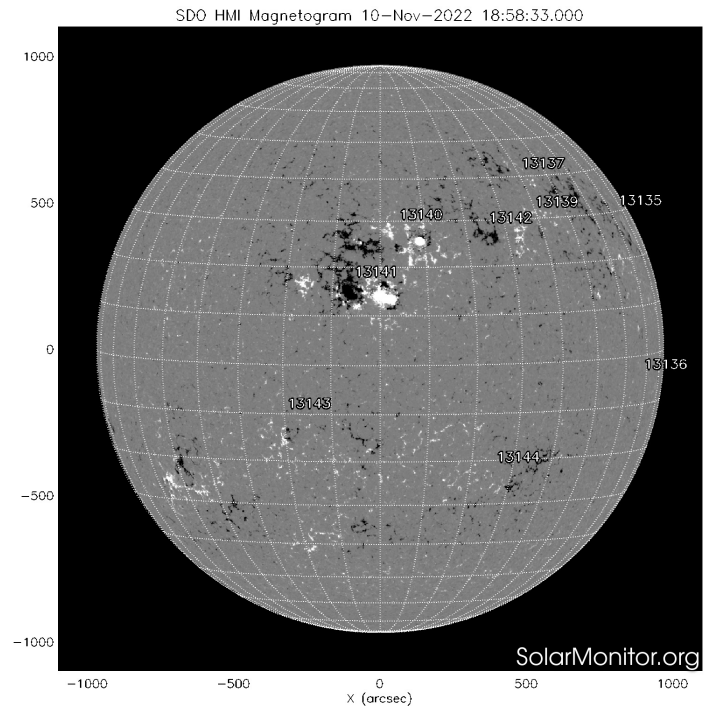
STEREO Ahead EUVI 195

2022-11-11 00:05:00

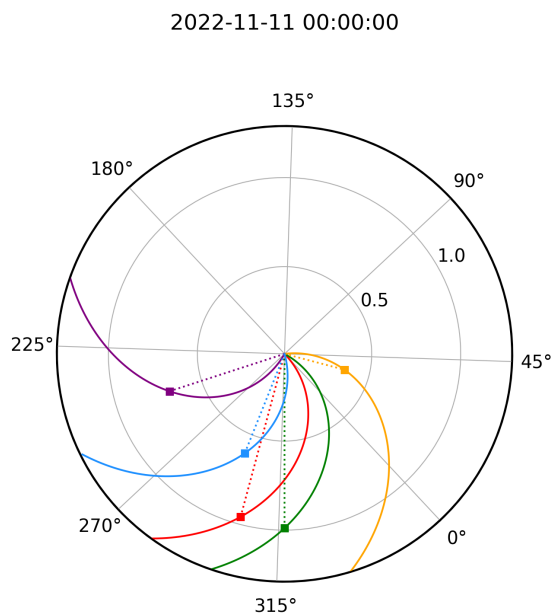


SDO/AIA 193 2022-11-10 00:00:17 UT

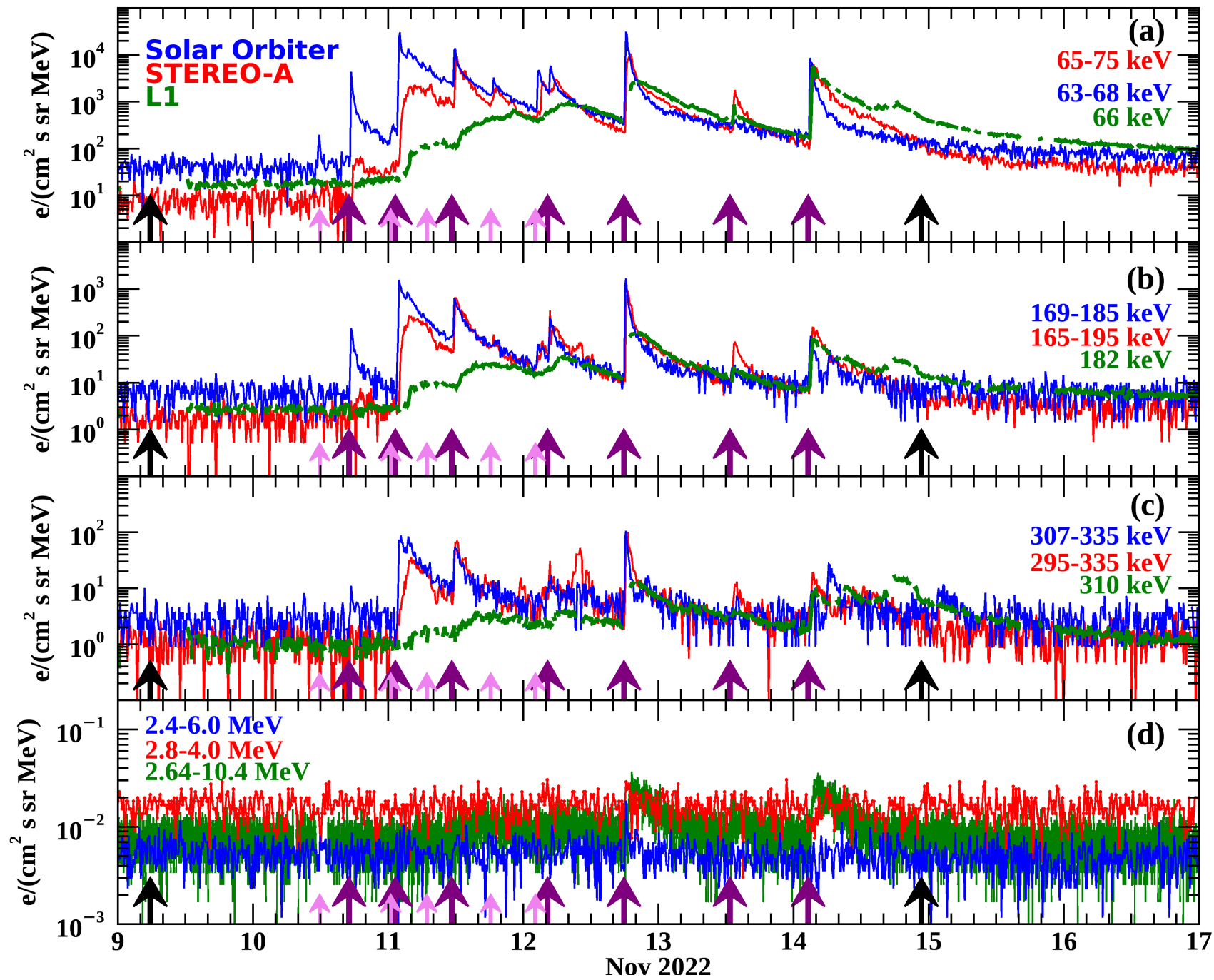
Multiple jets from AR 13141



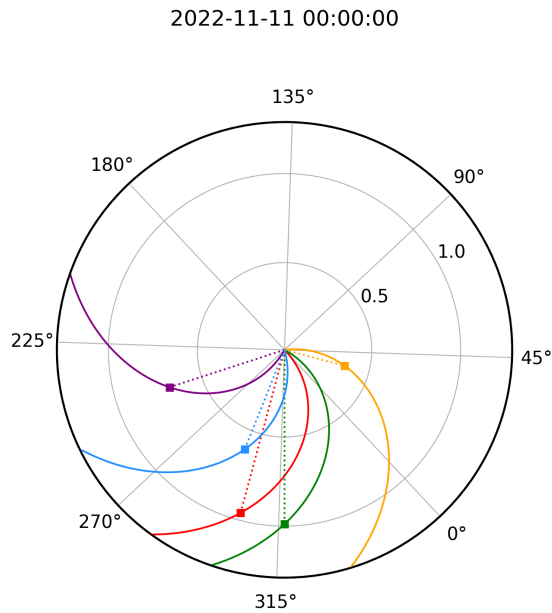
Electron observations



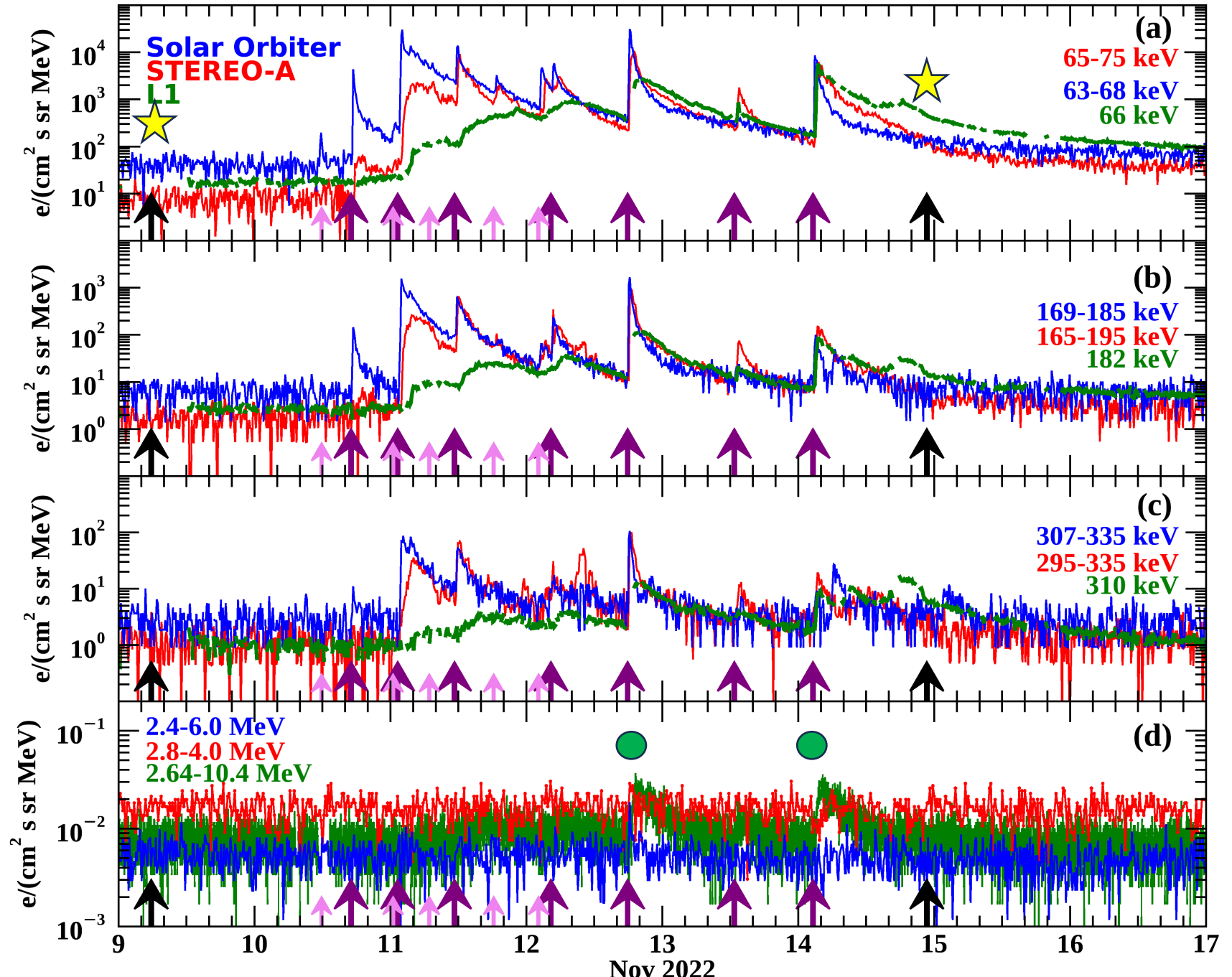
PSP ~0.69 AU -72 deg
 SoLo ~0.62 AU -21 deg
 STA ~0.96 AU -15 deg
 Earth ~1.02 AU 0 deg



Electron observations

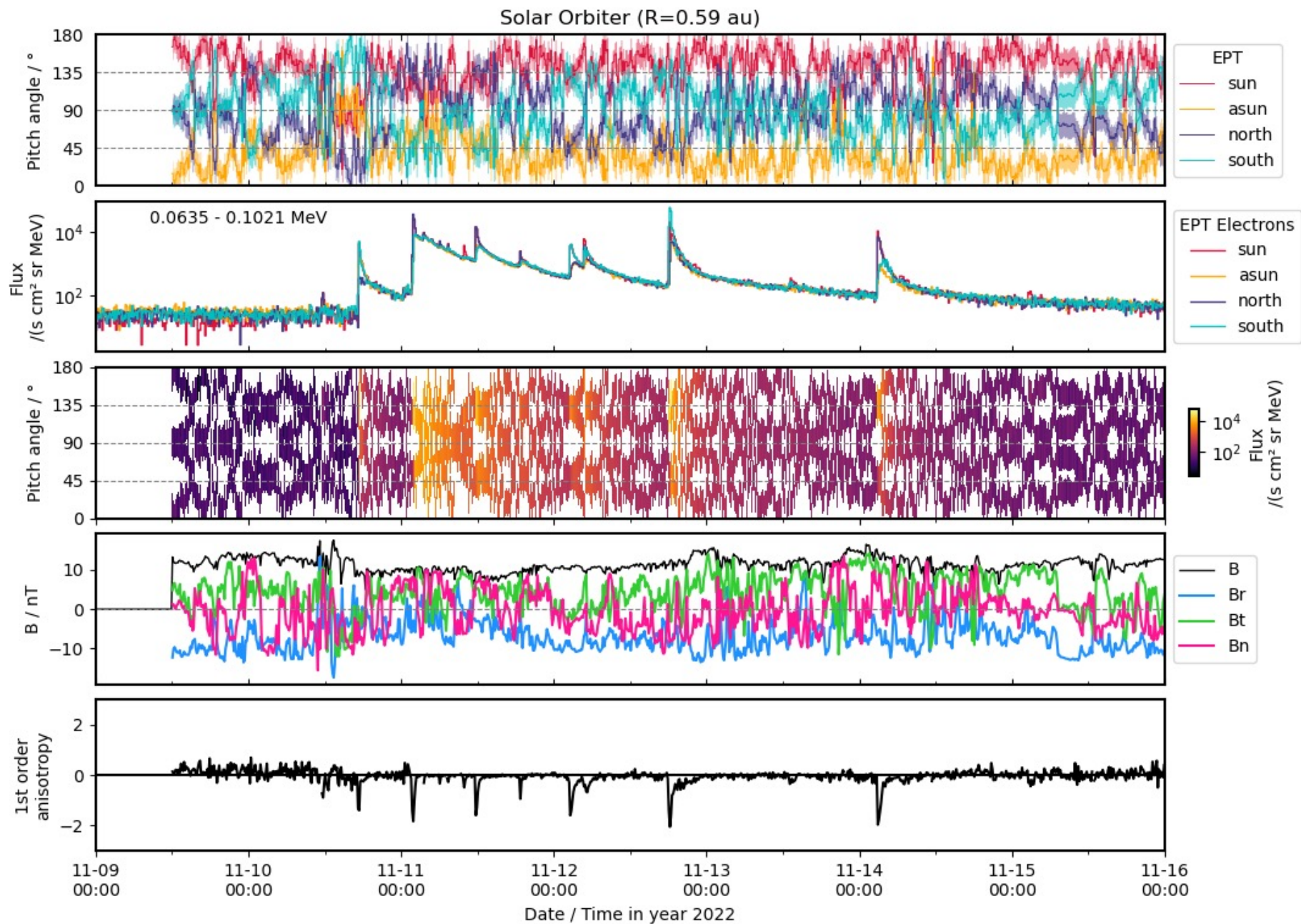


PSP ~0.69 AU -72 deg
 SoLO ~0.62 AU -21 deg
 STA ~0.96 AU -15 deg
 Earth ~1.02 AU 0 deg



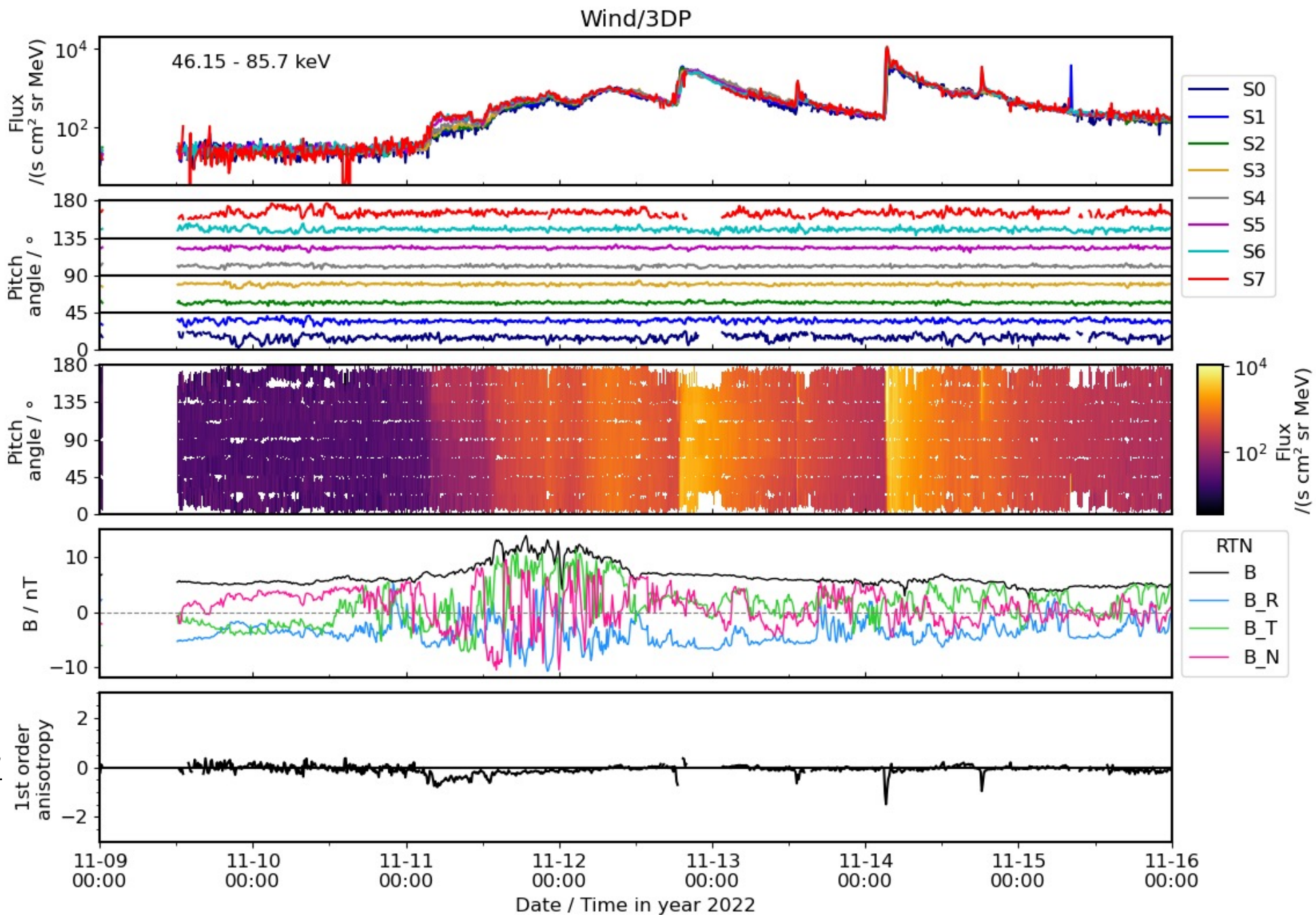
Solar Orbiter Anisotropy Information

$$A = \frac{3 \int_{-1}^{+1} I(\mu) \mu d\mu}{\int_{-1}^{+1} I(\mu) d\mu}$$

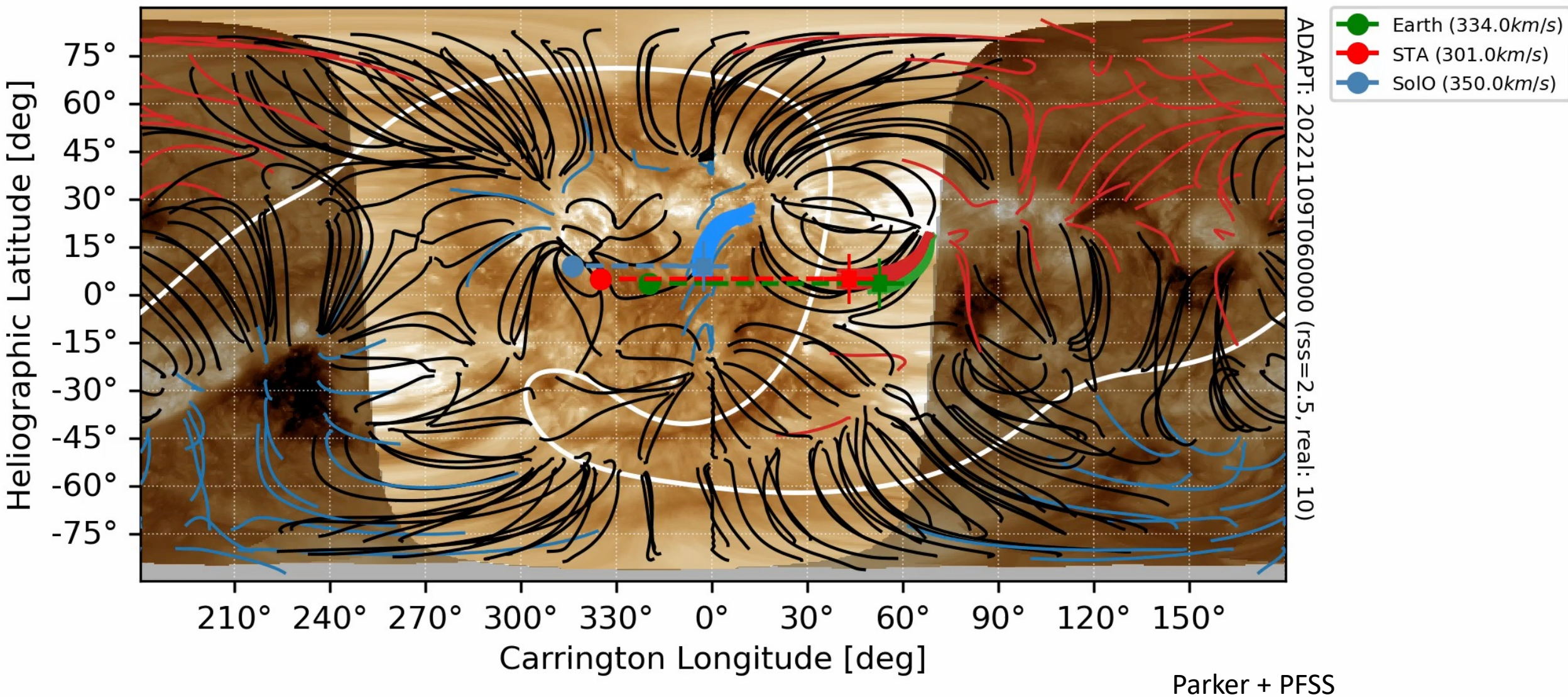


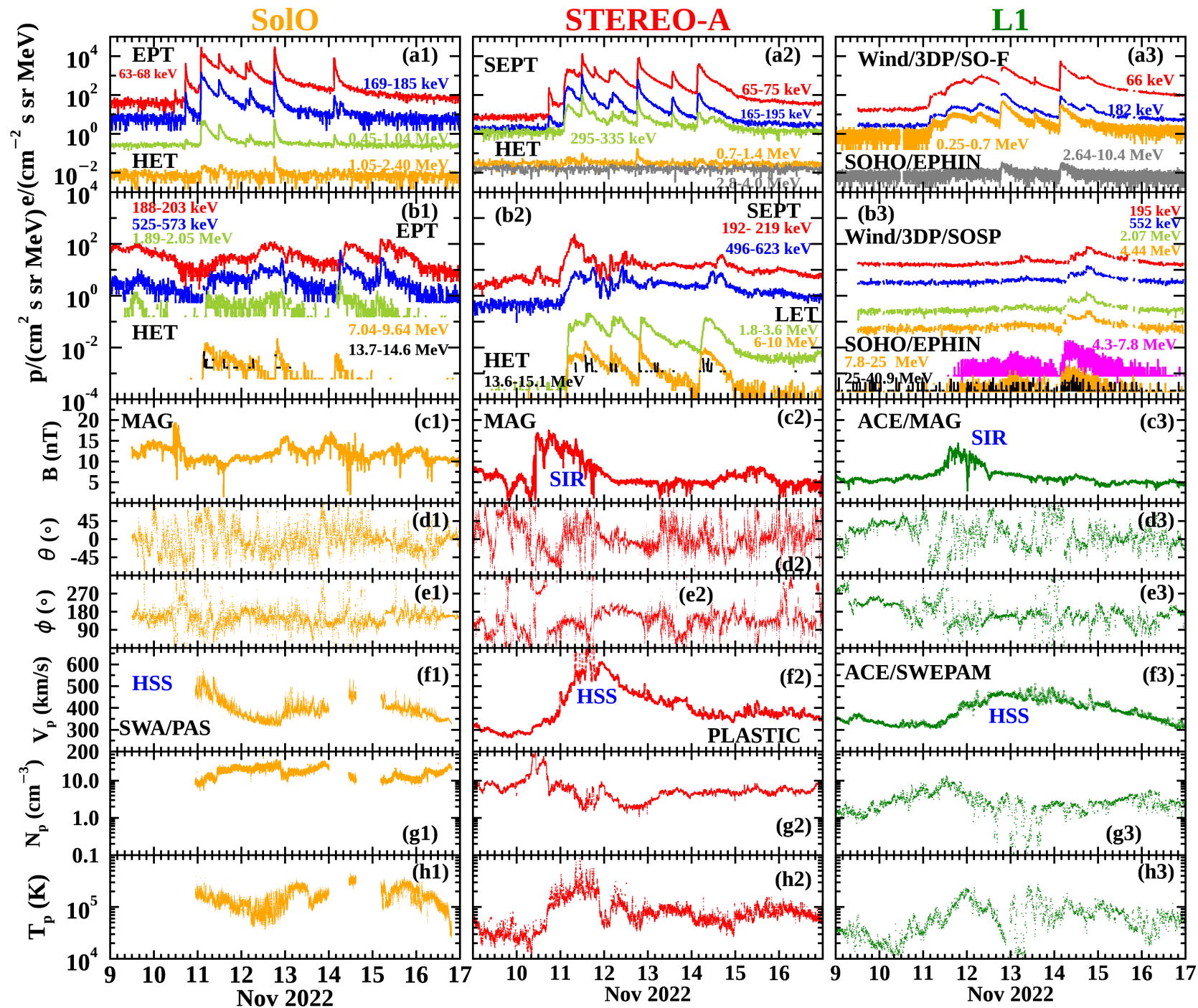
Wind/3DP Anisotropy Information

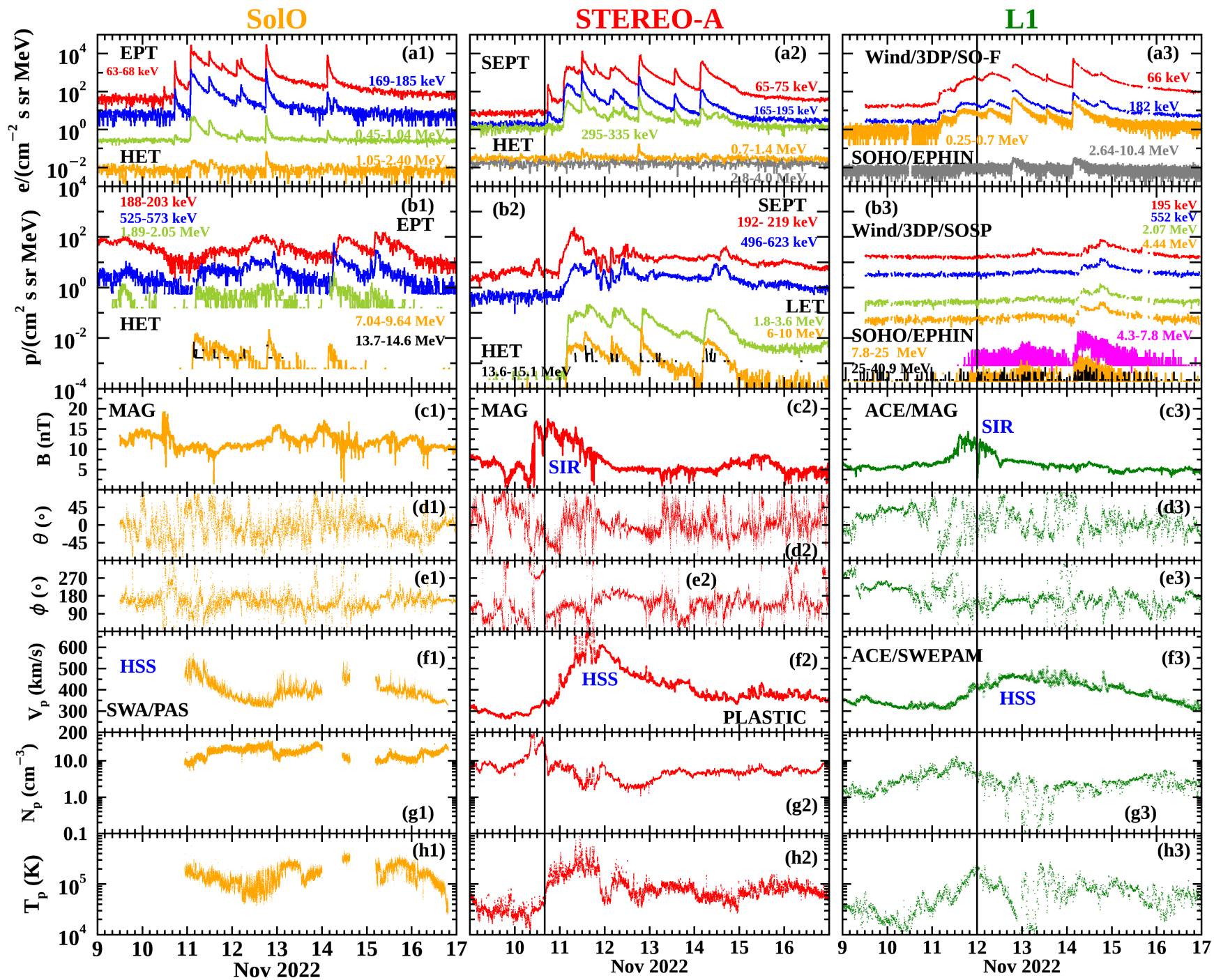
$$A = \frac{3 \int_{-1}^{+1} I(\mu) \mu d\mu}{\int_{-1}^{+1} I(\mu) d\mu}$$



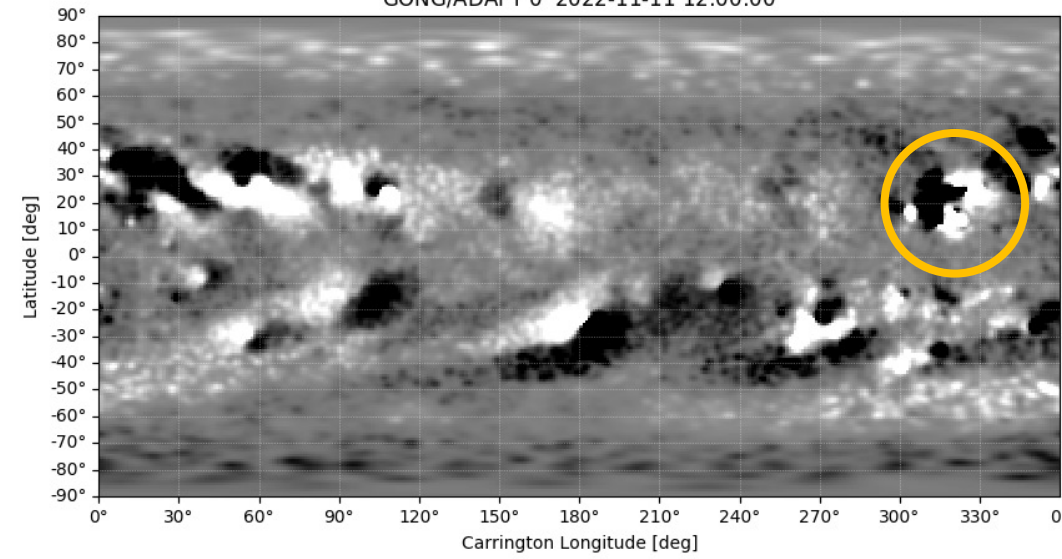
SDO/AIA-193 Carrington Synoptic Map: 20221109T0600



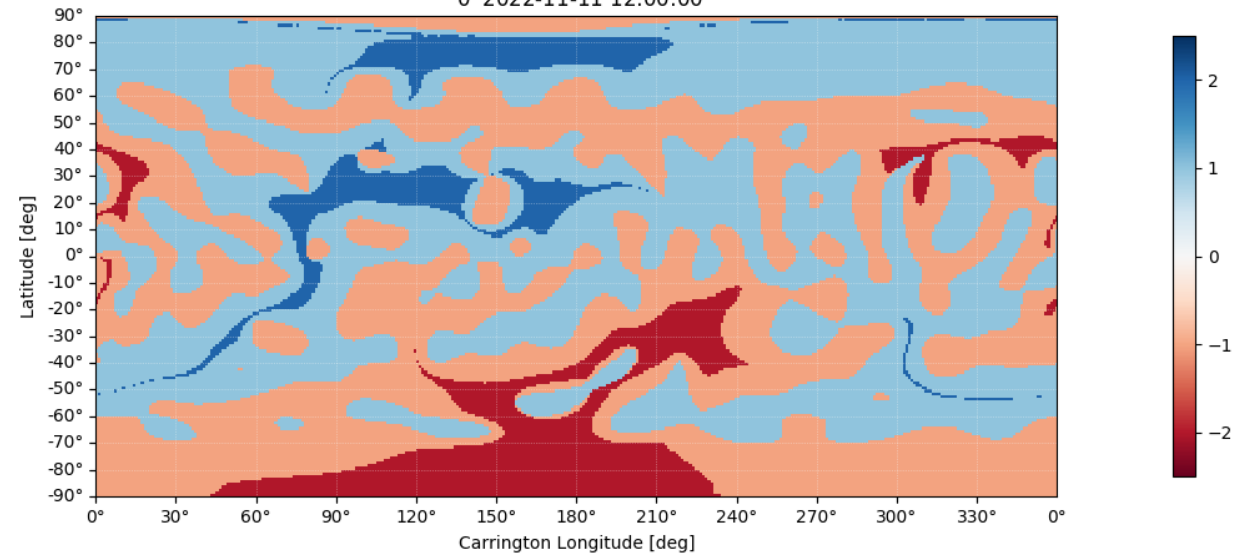




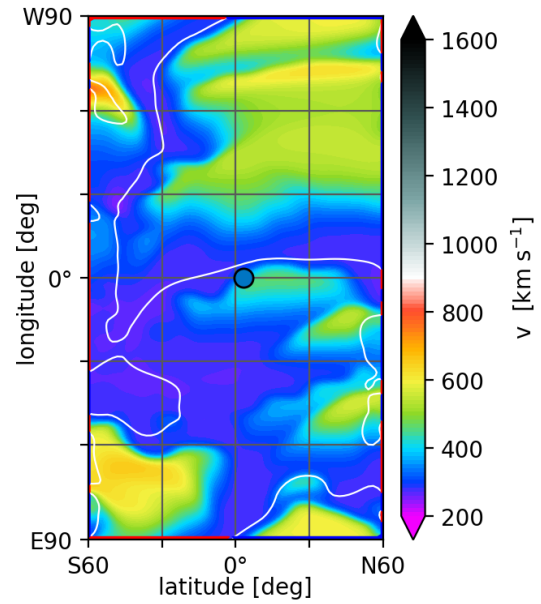
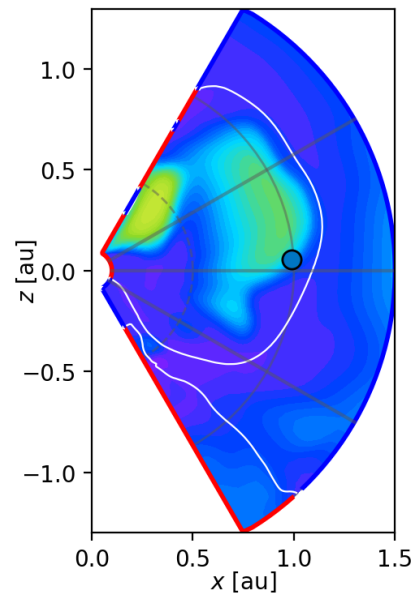
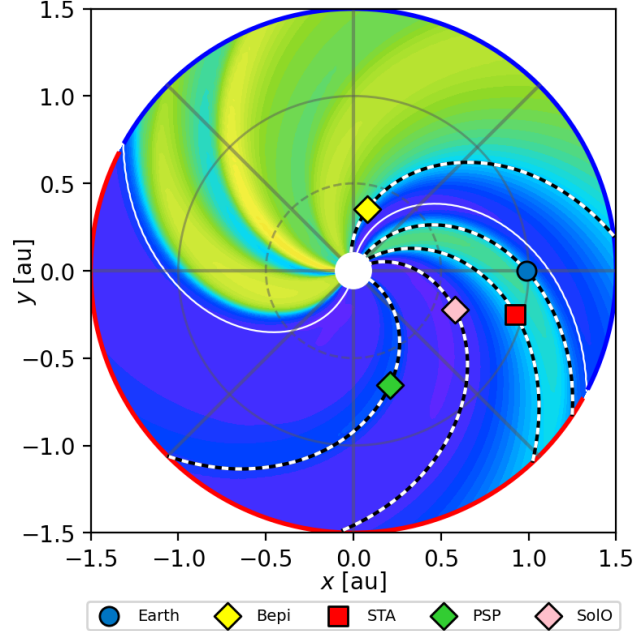
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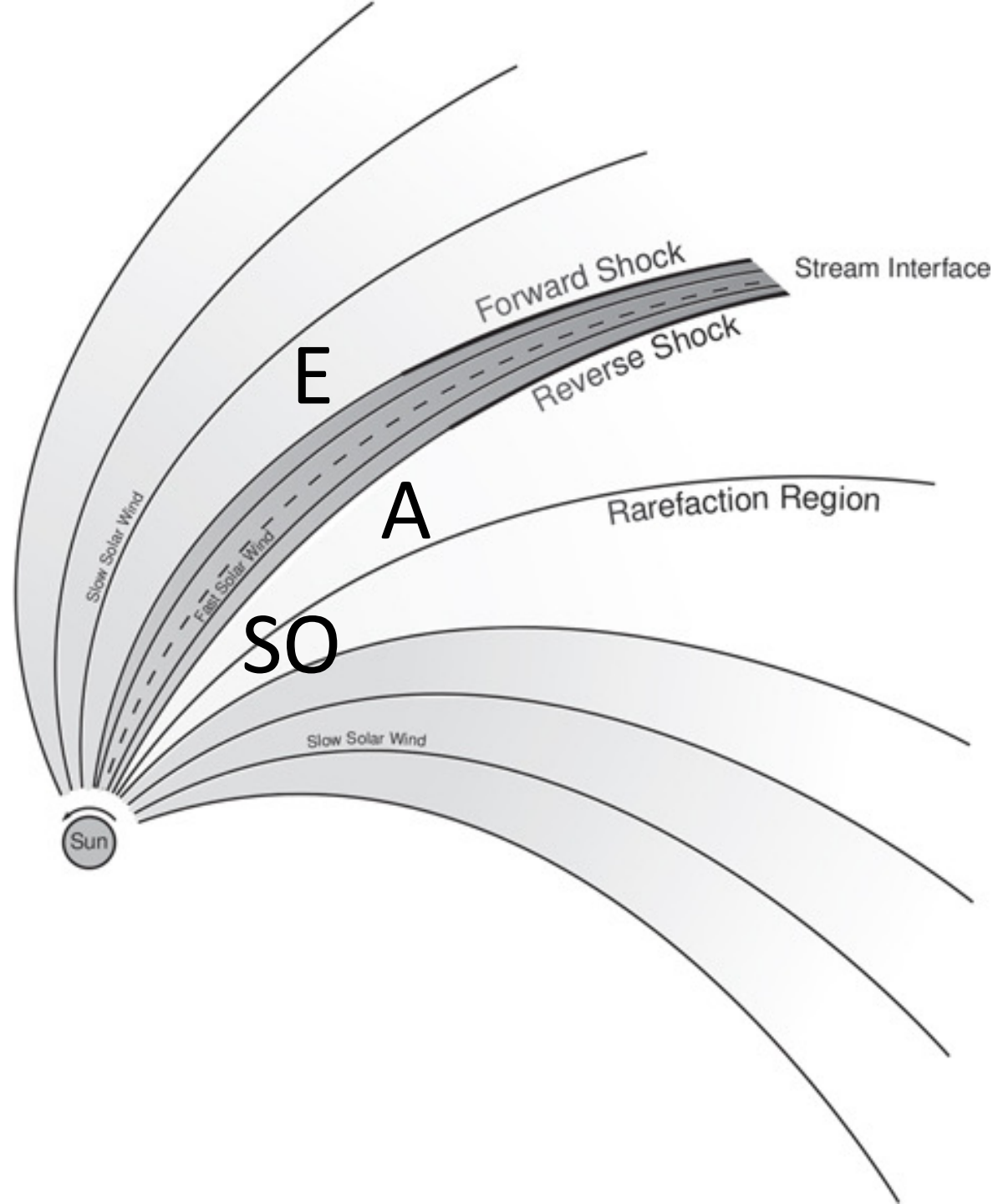
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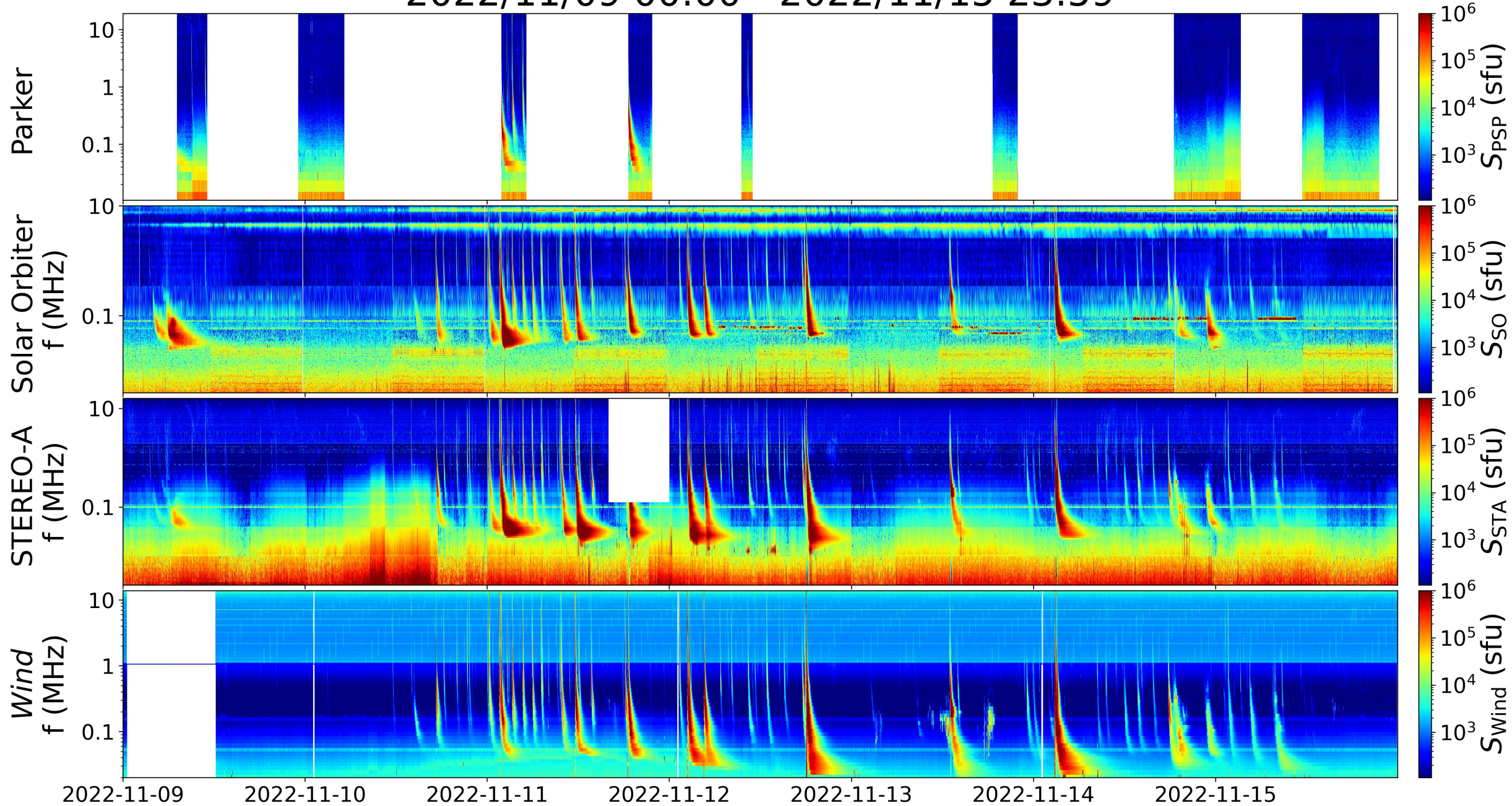
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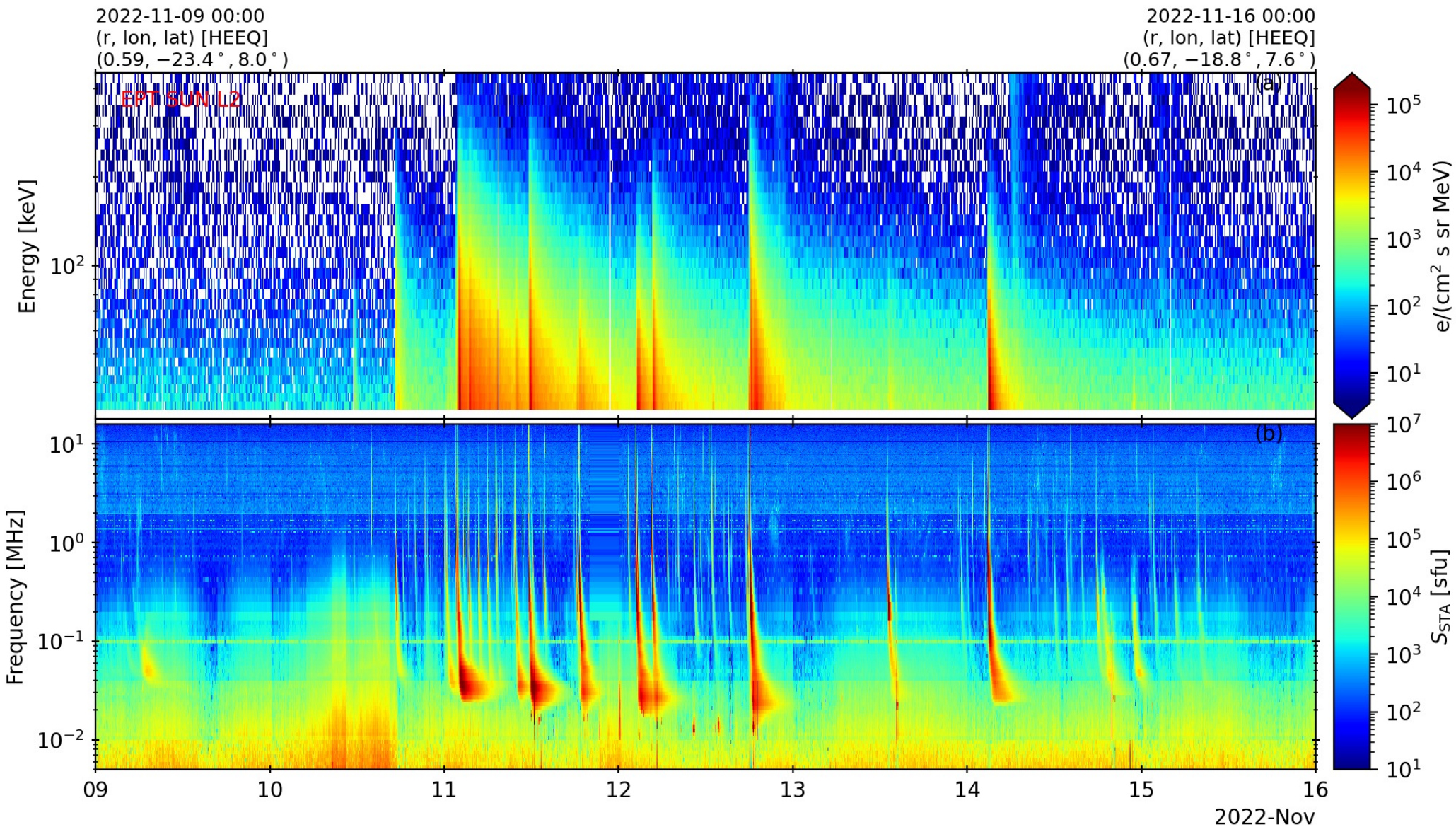
EUFORIA results



2022/11/09 00:00 - 2022/11/15 23:59



Solar Orbiter
EPT electrons
31-435 keV



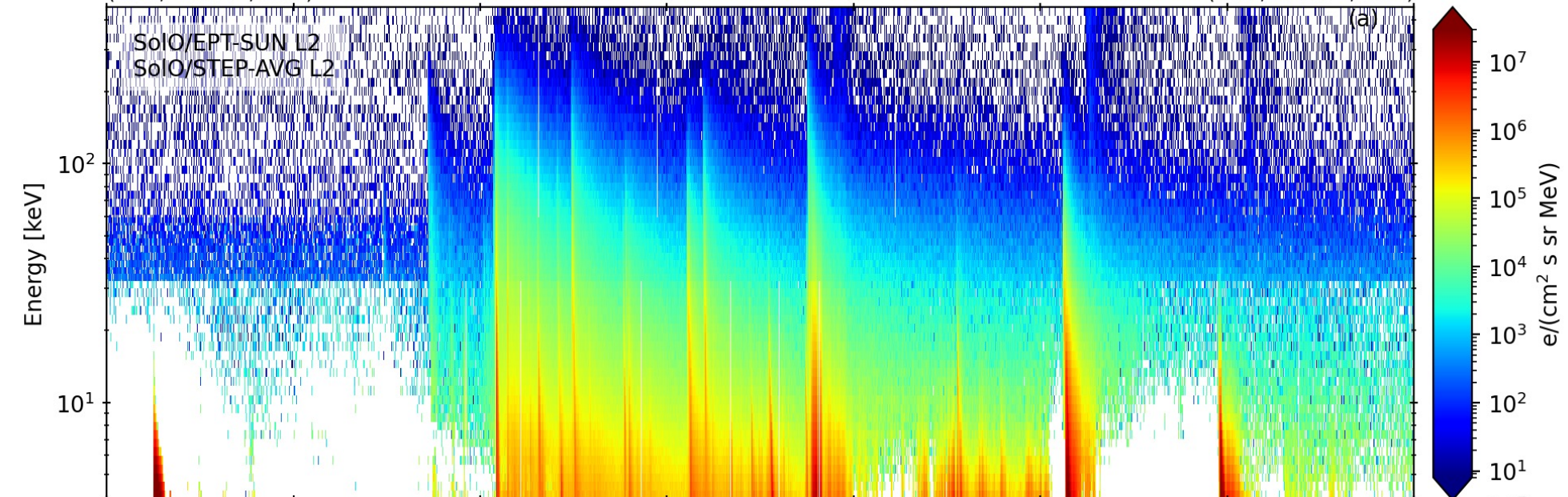
STEREO-A
Multiple Type III
Radio bursts

2022-11-09 00:00
(r, lon, lat) [HEEQ]
(0.59, -23.4°, 8.0°)

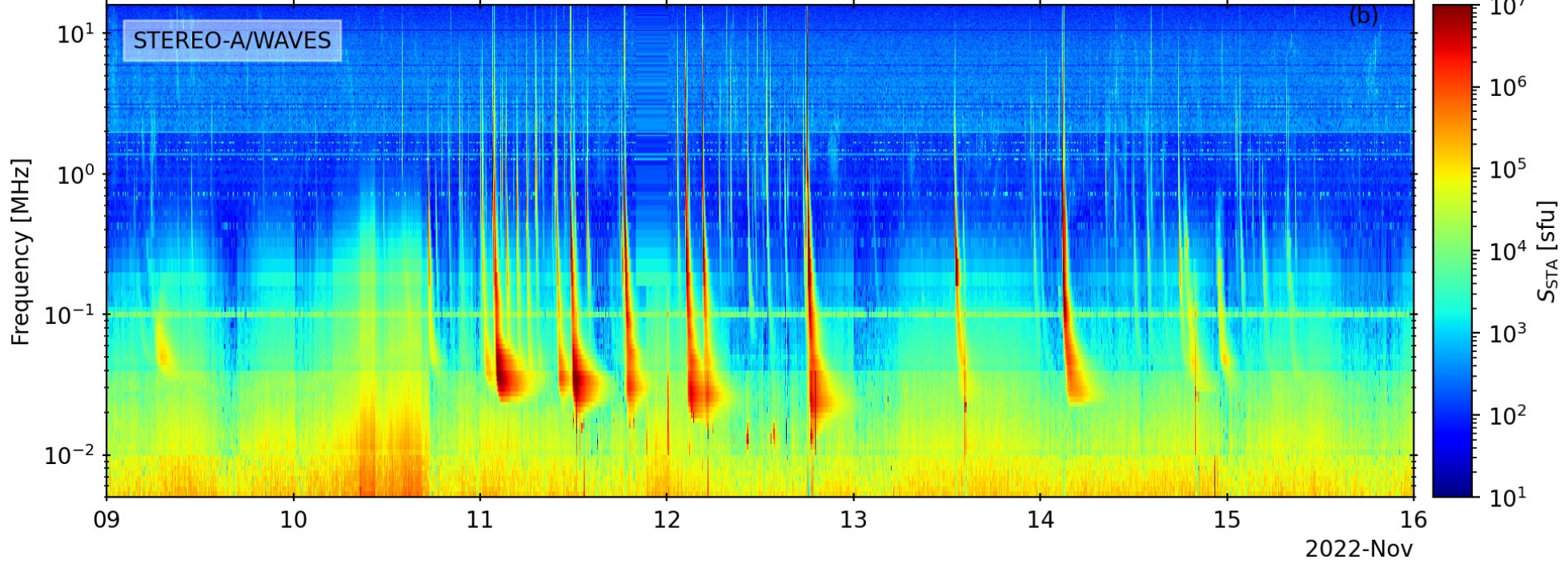
2022-11-16 00:00
(r, lon, lat) [HEEQ]
(0.67, -18.8°, 7.6°)

Solar Orbiter
EPT electrons
31-435 keV

STEP electrons
4-31 keV



STEREO-A
Multiple Type III
Radio bursts

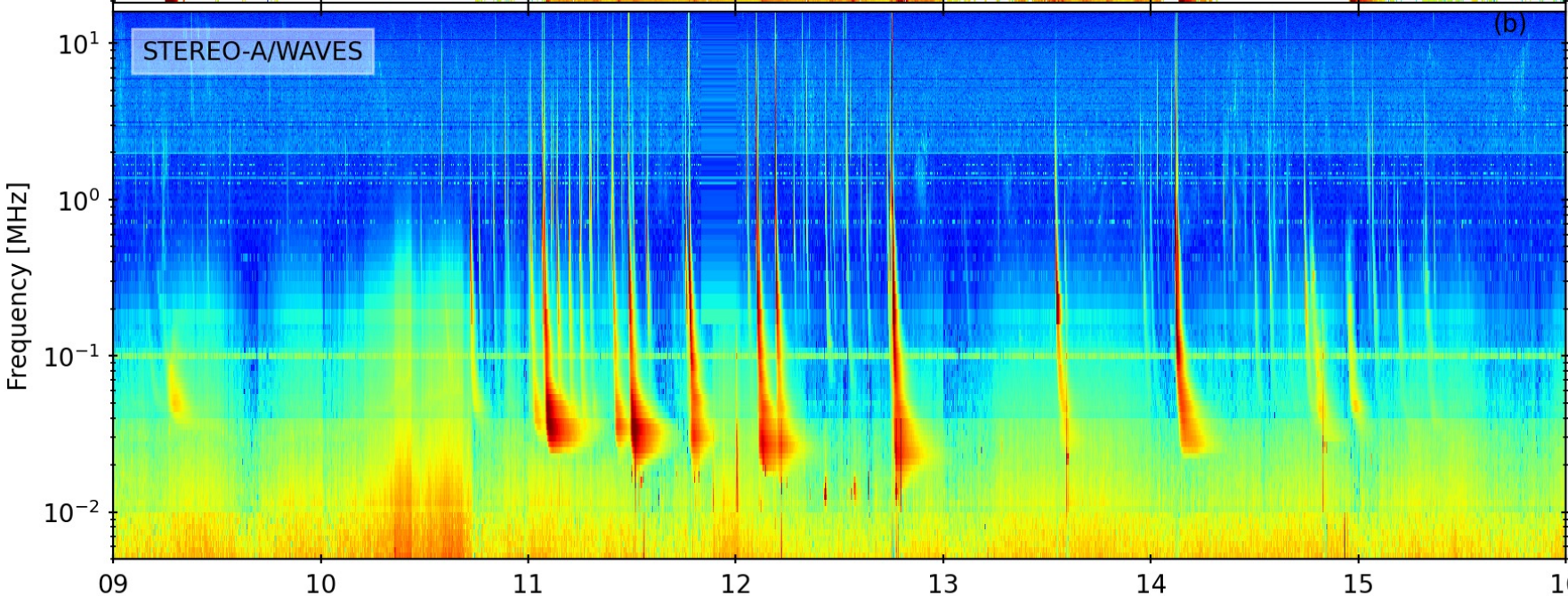
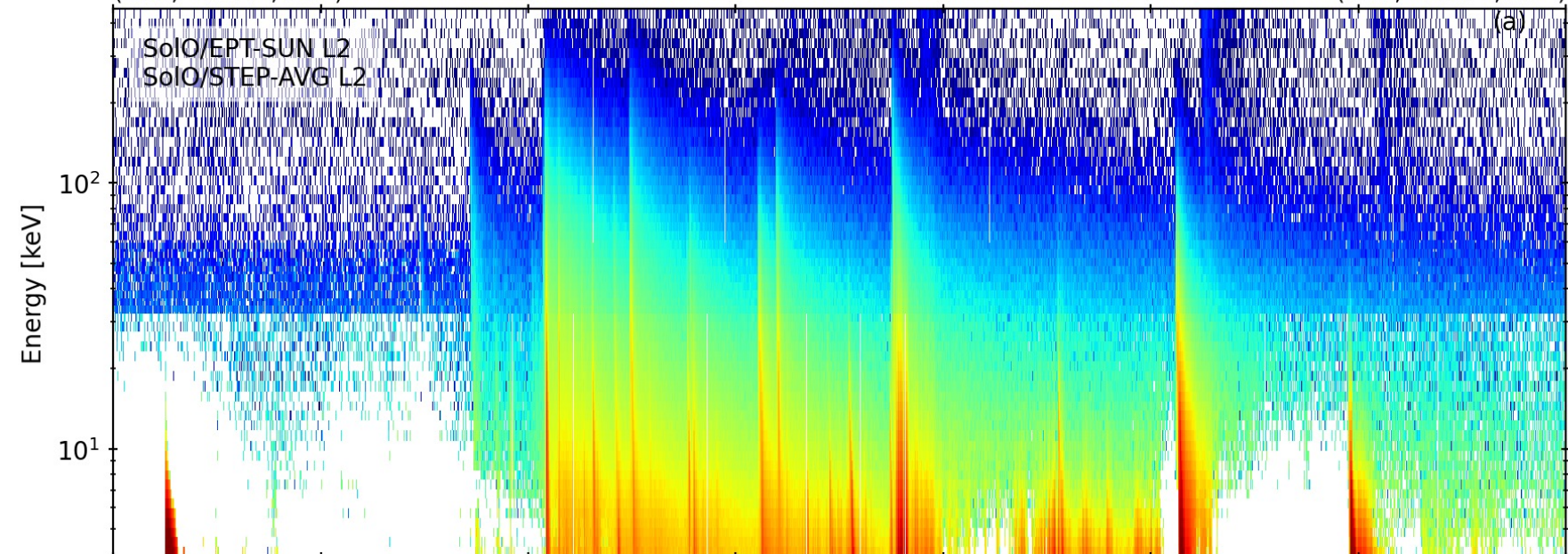


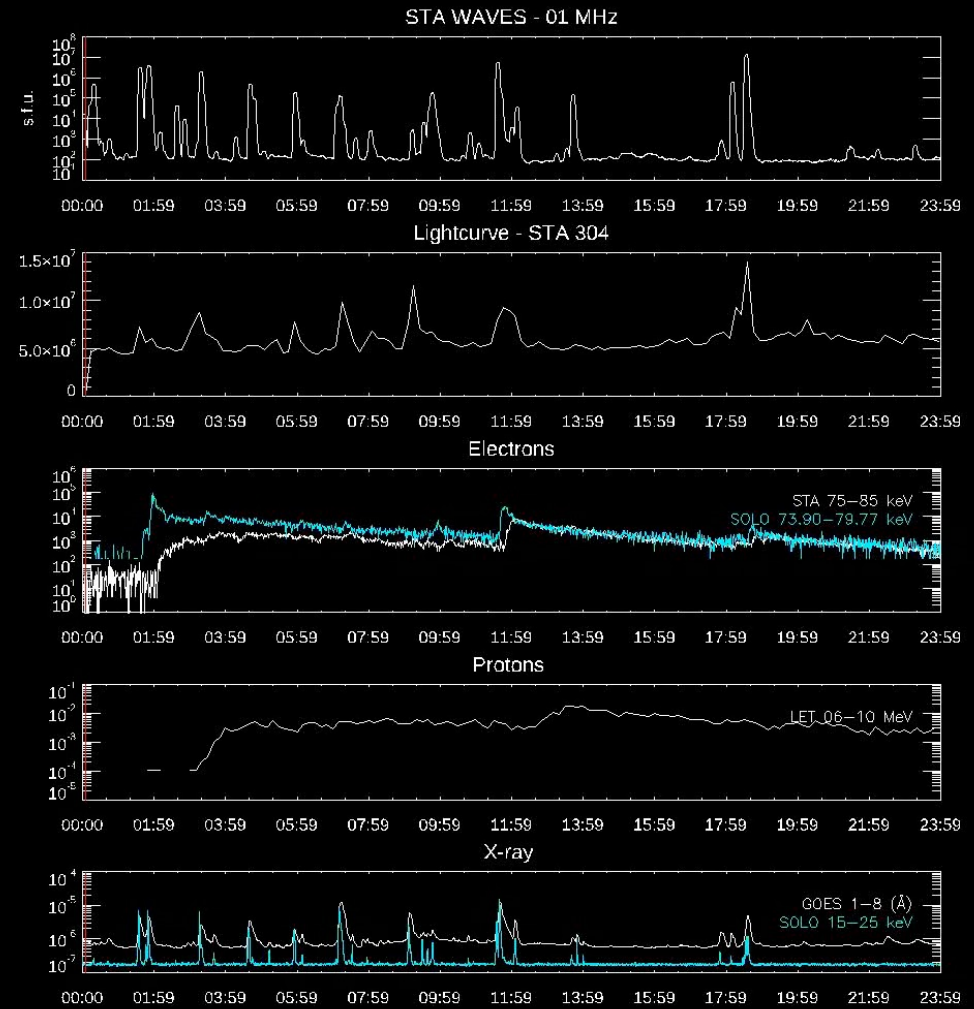
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(r, lon, lat) [HEEQ]
(0.59, -23.4°, 8.0°)

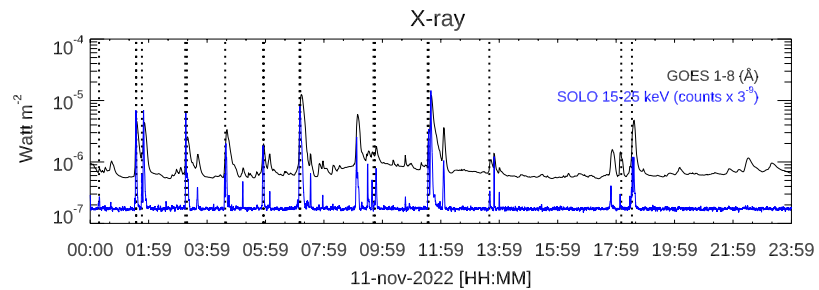
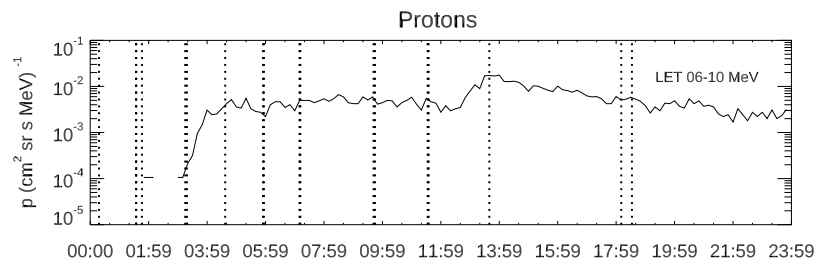
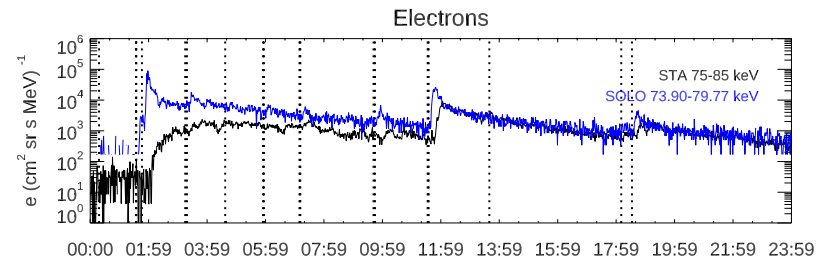
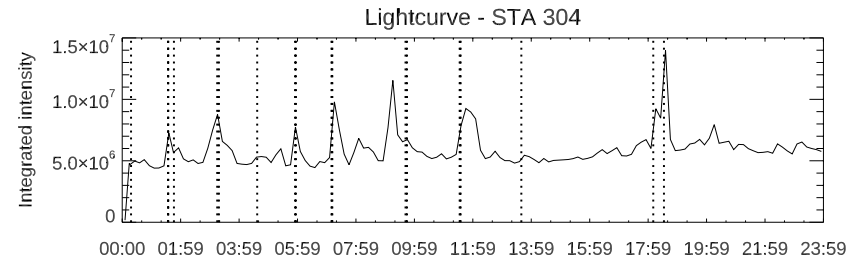
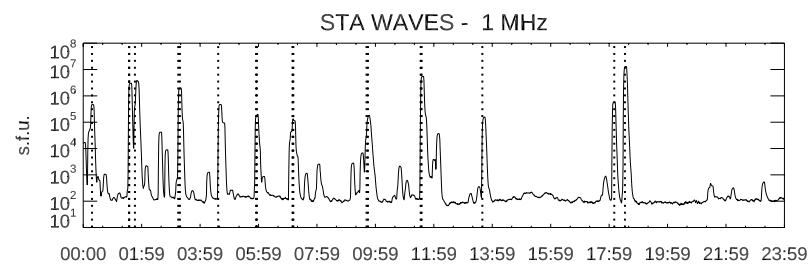
2022-11-16 00:00
(r, lon, lat) [HEEQ]
(0.67, -18.8°, 7.6°)

e- Release Times (Raúl)

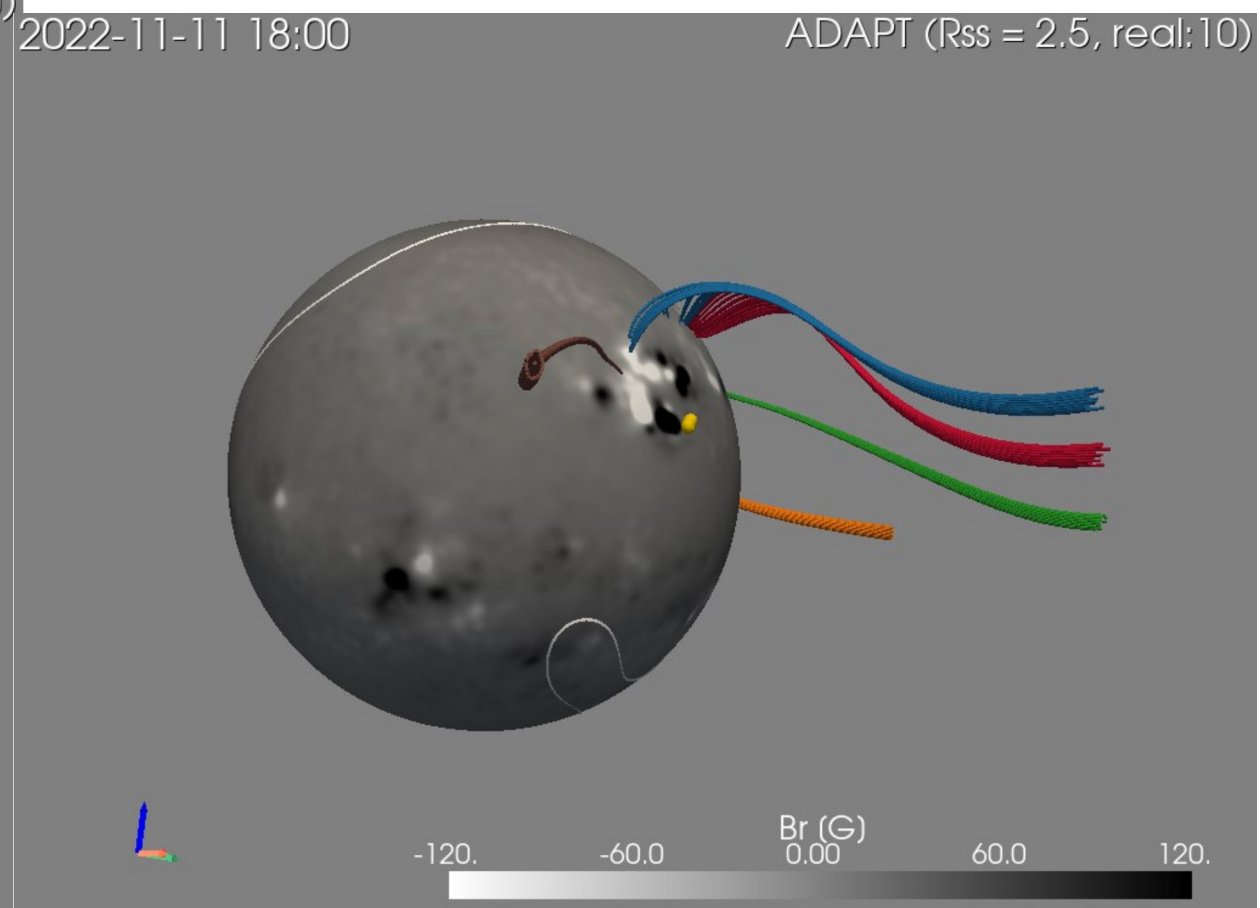
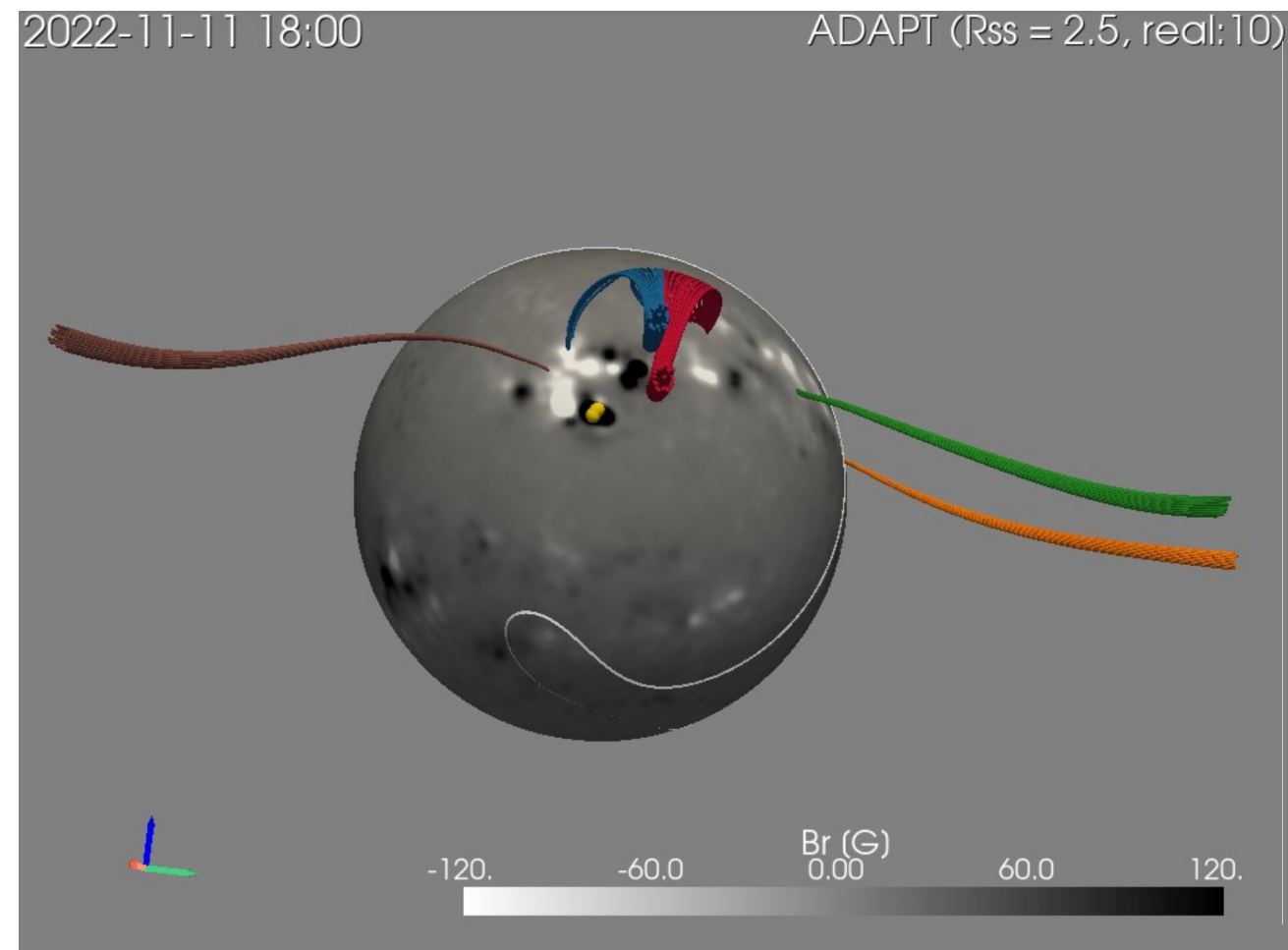
- 09 Nov 313/05:52:16 i
- 10 Nov 314/11:30:16
- 10 Nov 314/17:10:43
- 11 Nov 315/00:18:10
- 11 Nov 315/01:34:10 p
- 01:34:32(V)
- 11 Nov 315/01:46:10
- 11 Nov 315/03:15:44
- 03:18:37(V)
- 11 Nov 315/04:37:09
- 11 Nov 315/05:55:07
- 05:56:31(V)
- 11 Nov 315/07:09:47
- 07:11:31(V)
- 11 Nov 315/09:41:51
- 09:44:50(V)
- 11 Nov 315/11:32:58 p
- 11:35:00(V)
- 11 Nov 315/13:39:38
- 11 Nov 315/18:10:31
- 11 Nov 315/18:32:30
- 12 Nov 316/02:21:16 p
- 02:24:17(V)
- 12 Nov 316/04:33:13
- 04:34:47(V)
- 12 Nov 316/12:51:29
- 12 Nov 316/17:47:55 p
- 13 Nov 317/13:05:04
- 14 Nov 318/02:42:03 p
- 02:45:04(V)
- 14 Nov 318/22:36:53 i
- 22:31:37(V)





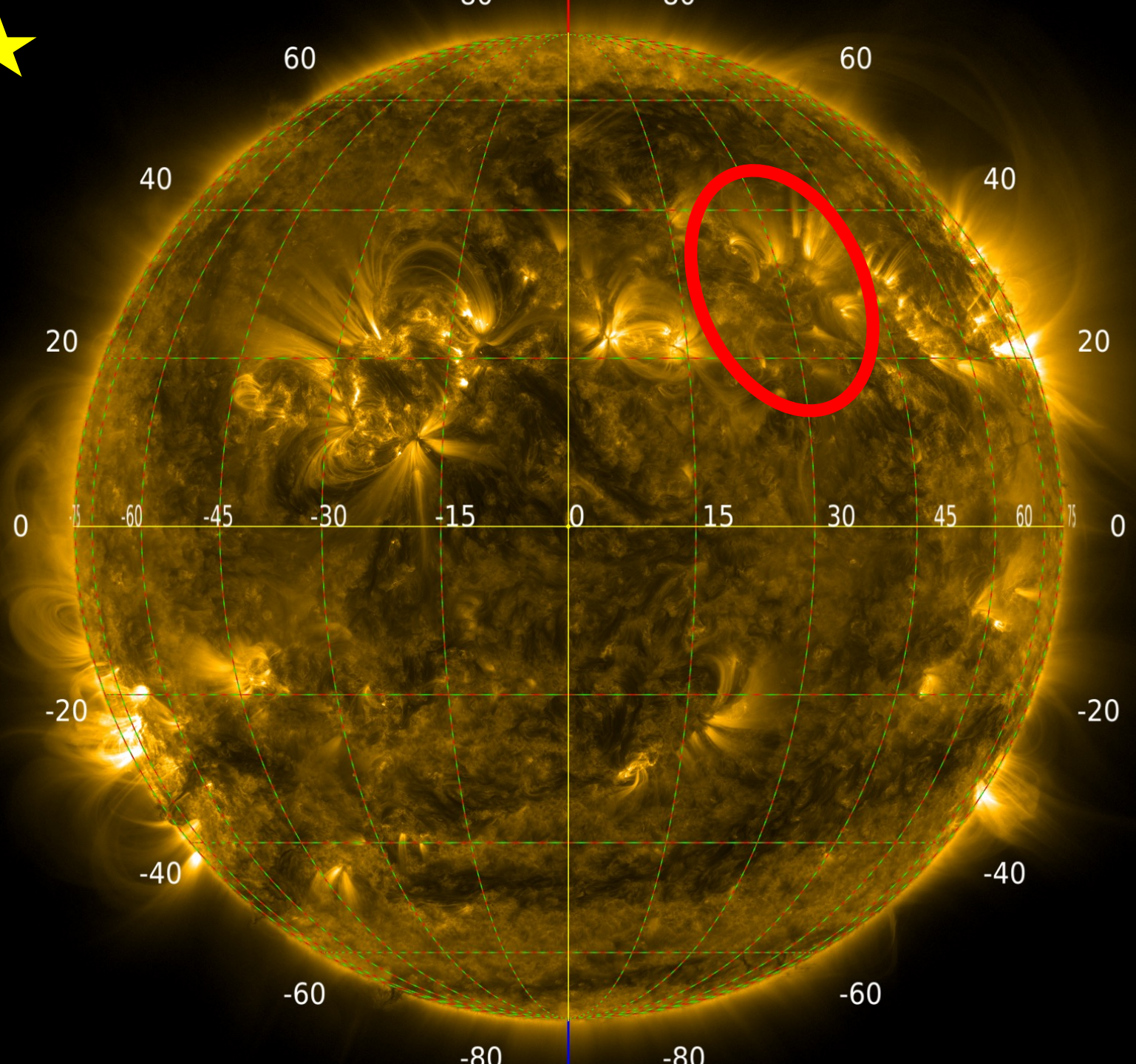


11-nov-2022 [HH:MM]

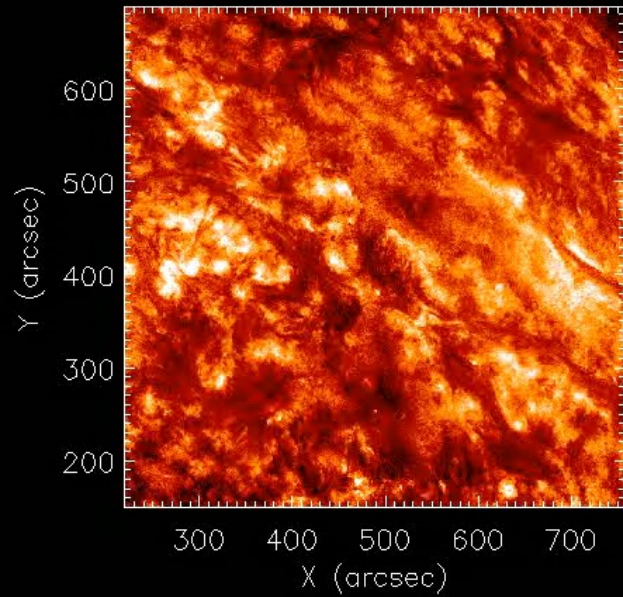


Parker + PFSS

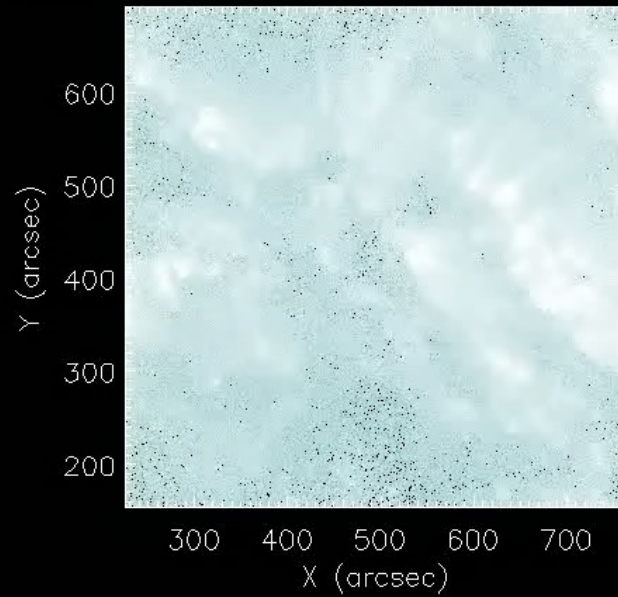
Nov 9 event ★



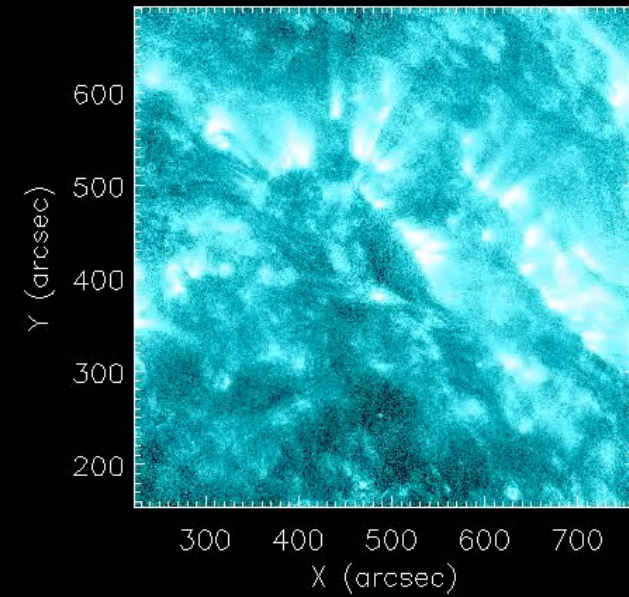
304



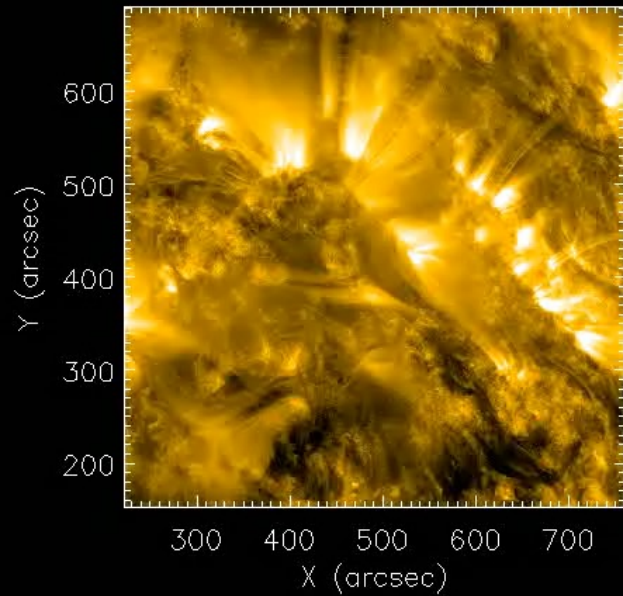
SDO AIA_4 94 9-Nov-2022 05:00:00.000 UT



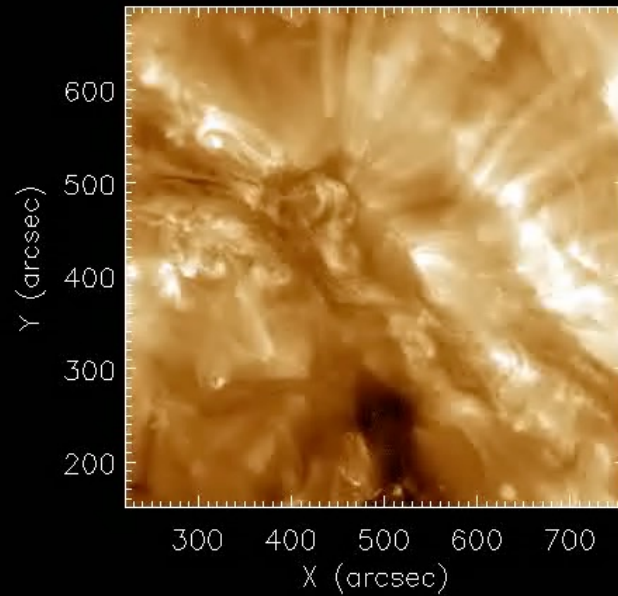
131



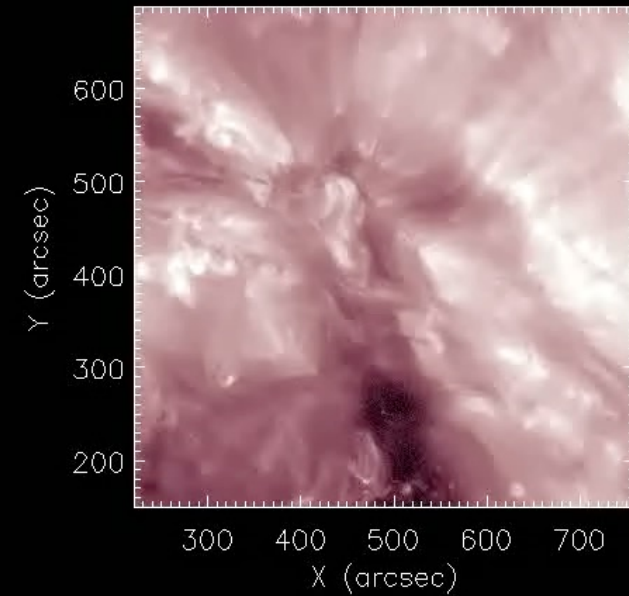
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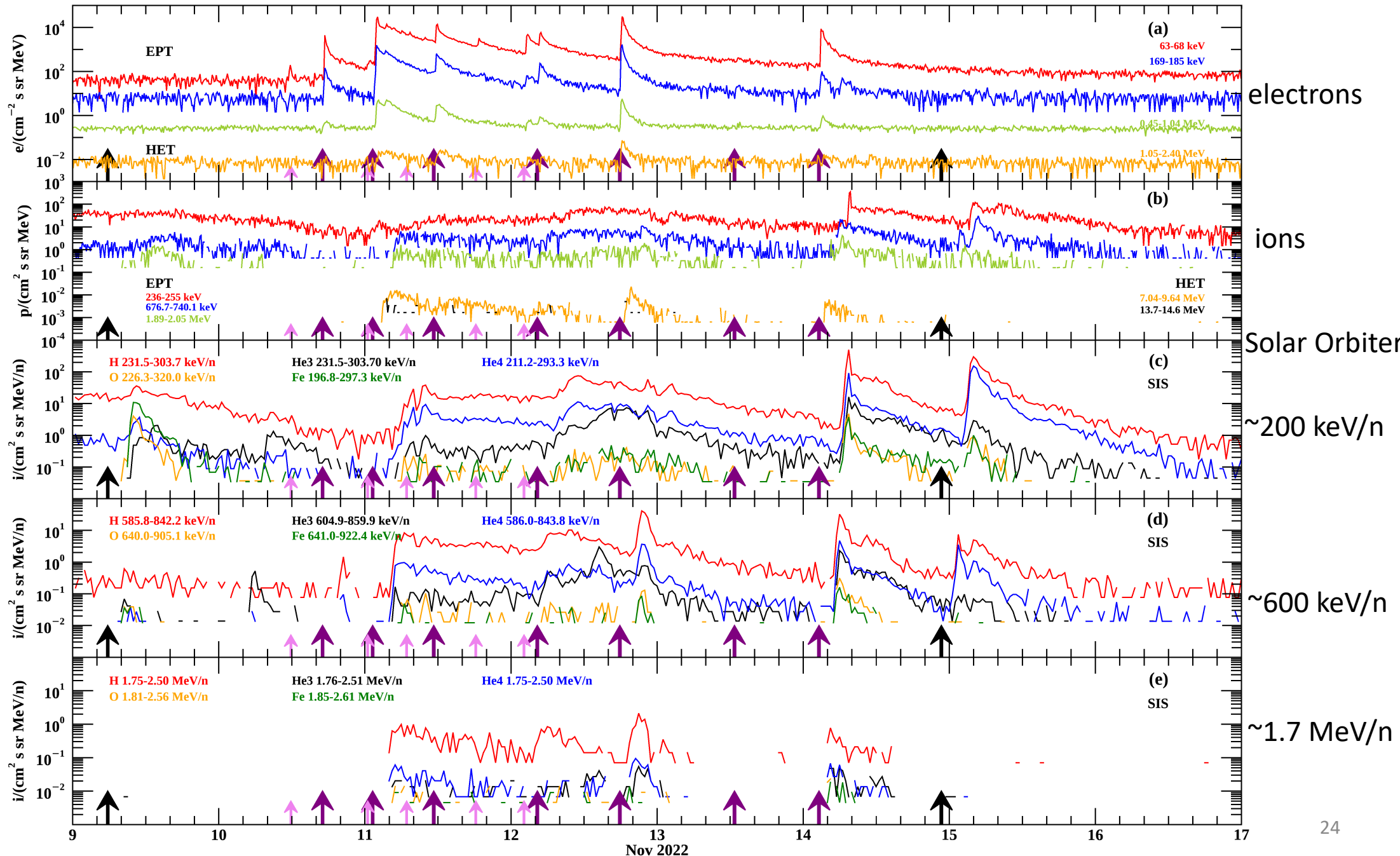
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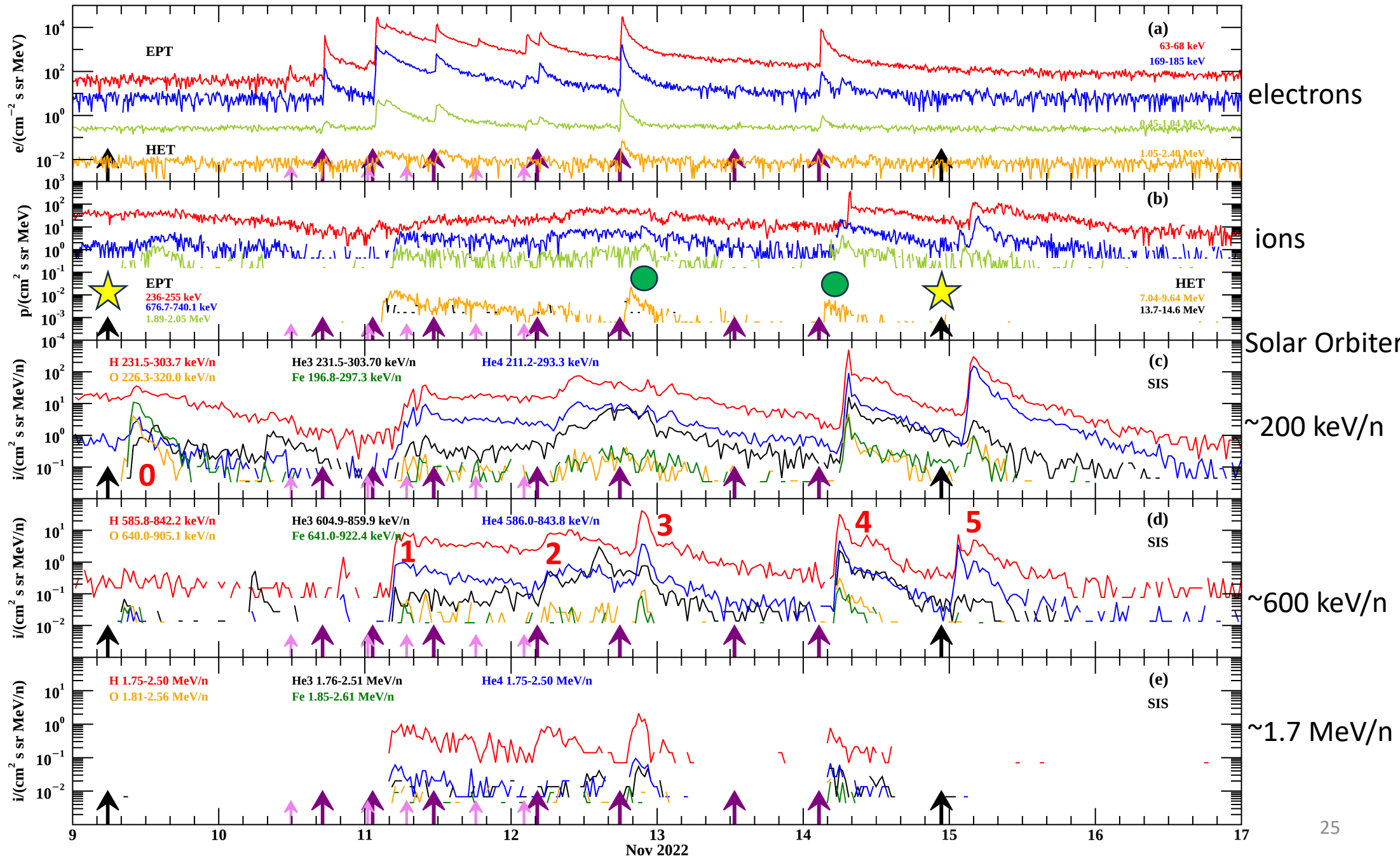
211



SoLO



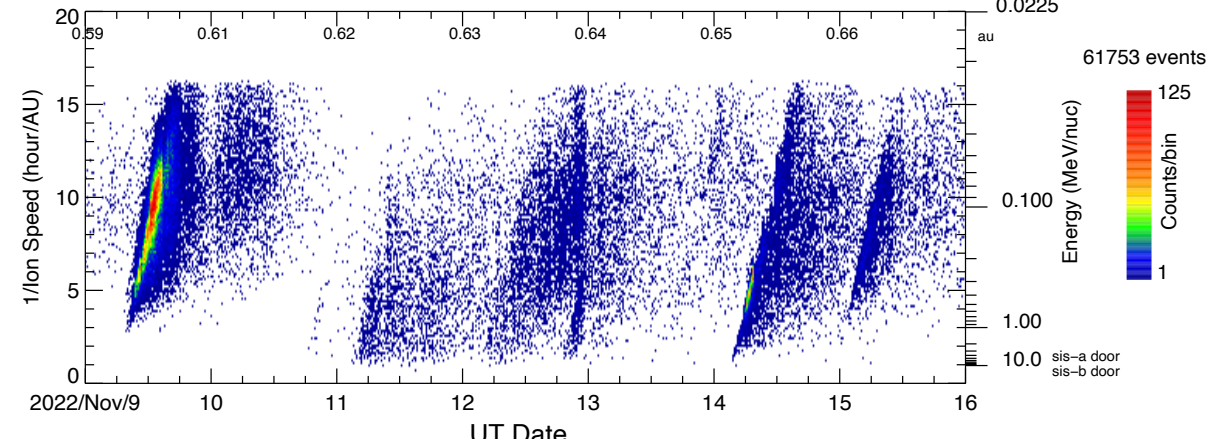
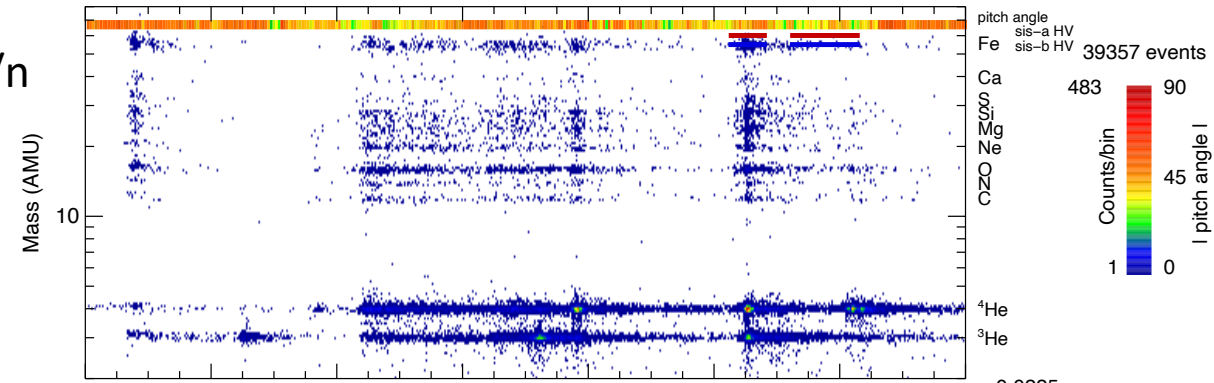
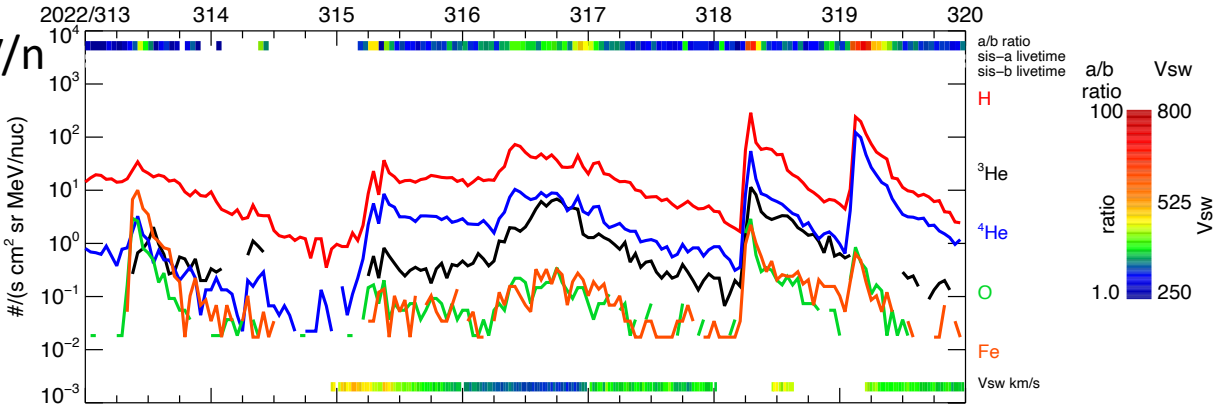
Solo



Top panel: SIS-a ~0.23–0.32 MeV/n; SIS-a/SIS-b ratio; Vsw speed from L2 data
 approx livetime: clear = >95%, blue = 80–95%, yellow = 60–80%, orange = 40–60%, red = 20–40%; black <20%

Middle panel: SIS-a energy range = 0.40–10 MeV/n; SIS-a l pitch angle from LL02_srf data

HV status: clear = in calibration; color bars = out of calibration: red = SIS-a; blue = SIS-b
 Bottom panel: SIS-a mass Range = 10.0–70 AMU; helio dist, lat; s/c angle from Earth; Mag footprint (400 km/s Vsw) at start of day
 approx door position: clear = 100%, blue = 26%, orange = 6%, red = 1%, black = 0%/unknown



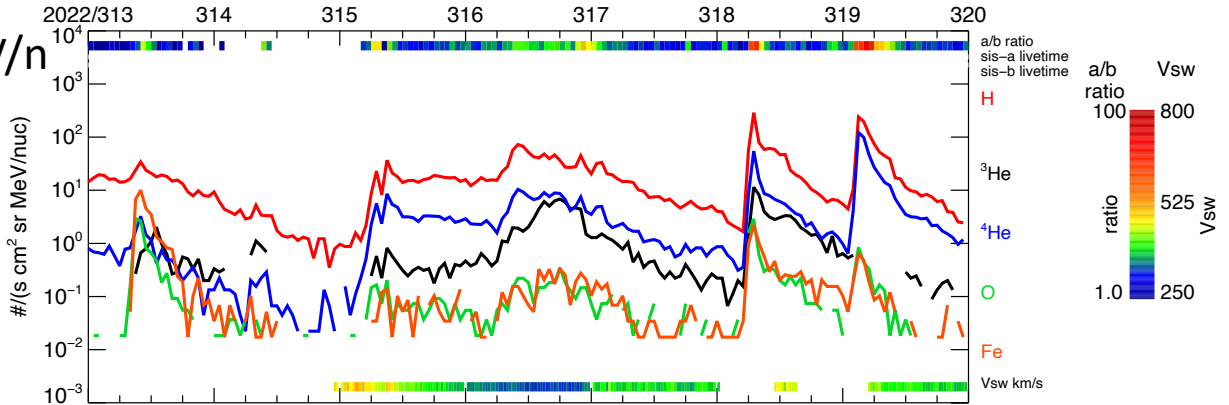
230-320 keV/n

0.4-10 MeV/n

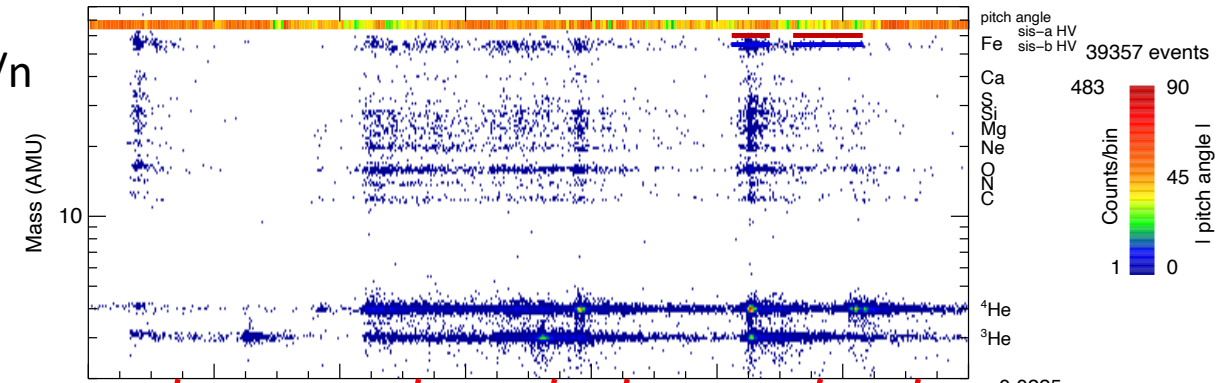
10-70 AMU

Top panel: SIS-a ~0.23–0.32 MeV/n; SIS-a/SIS-b ratio; Vsw speed from L2 data
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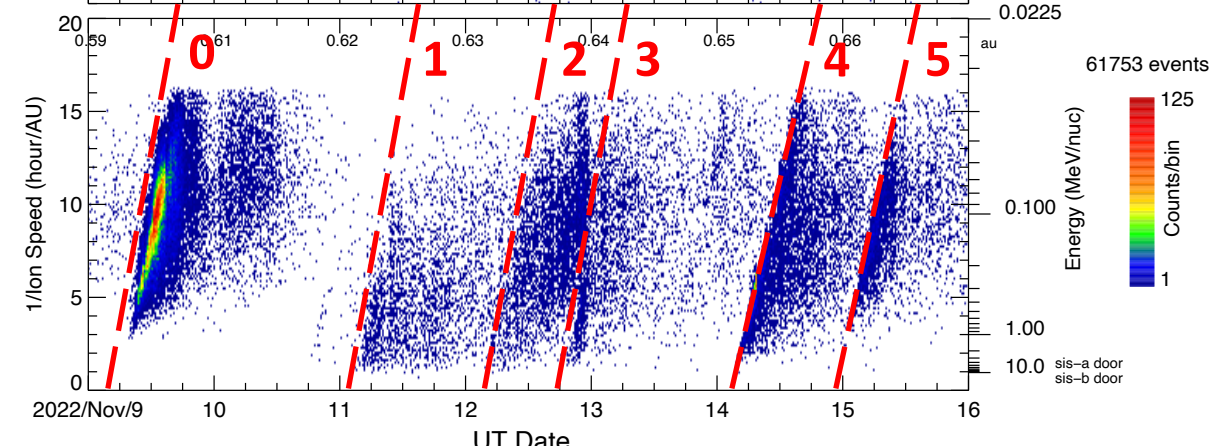
230-320 keV/n



0.4-10 MeV/n



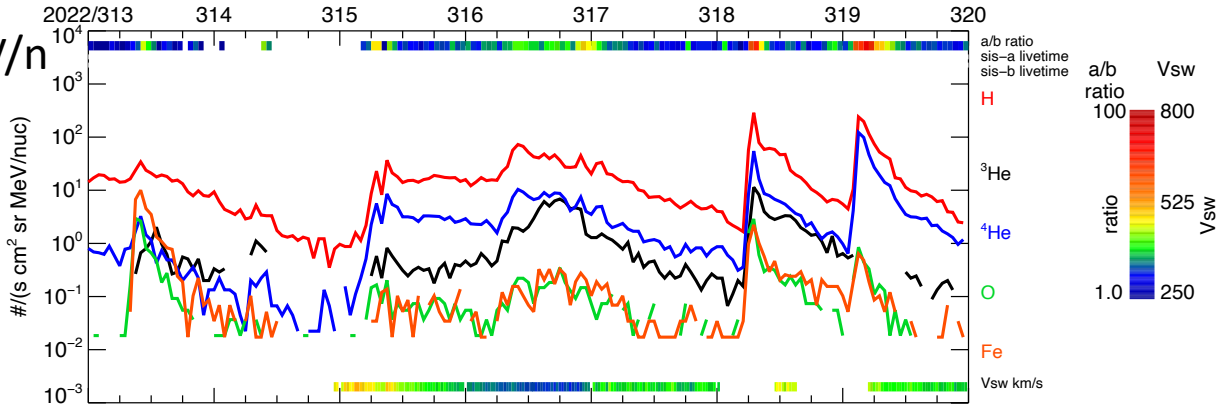
10-70 AMU



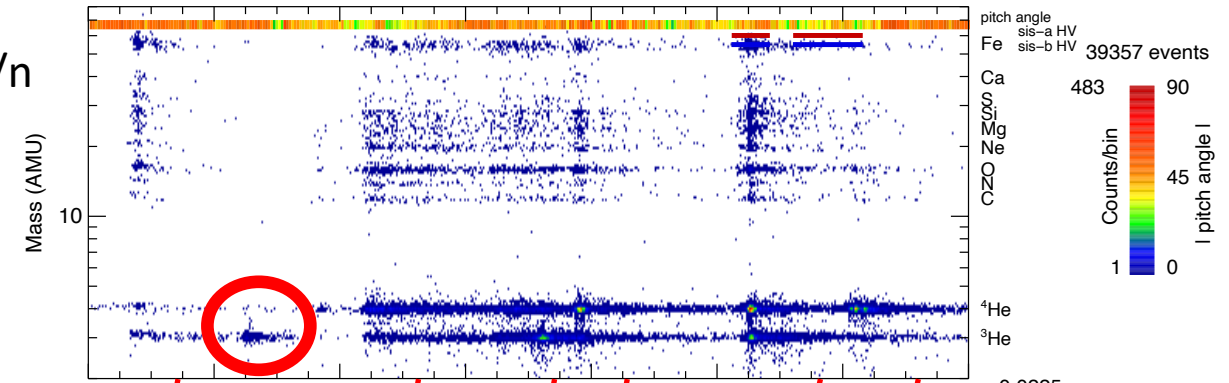
	3He/4He (0.5-2.0 MeV/n)	Fe/O (0.32-0.45 MeV/n)
Event #0	2.19±0.52	2.66±0.20
Event #1	0.22±0.01	0.62±0.10
Event #2	1.25±0.03	0.92±0.13
Event #3	0.36±0.02	0.93±0.24
Event #4	0.84±0.02	1.04±0.10
Event #5	0.12±0.01	1.00±0.26

Top panel: SIS-a ~0.23–0.32 MeV/n; SIS-a/SIS-b ratio; Vsw speed from L2 data
 approx livetime: clear = >95%, blue = 80–95%, yellow = 60–80%, orange = 40–60%, red = 20–40%; black <20%
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 HV status: clear = in calibration; color bars = out of calibration: red = SIS-a; blue = SIS-b
 Bottom panel: SIS-a mass Range = 10.0–70 AMU; helio dist, lat; s/c angle from Earth; Mag footprint (400 km/s Vsw) at start of day
 approx door position: clear = 100%, blue = 26%, orange = 6%, red = 1%, black = 0%/unknown

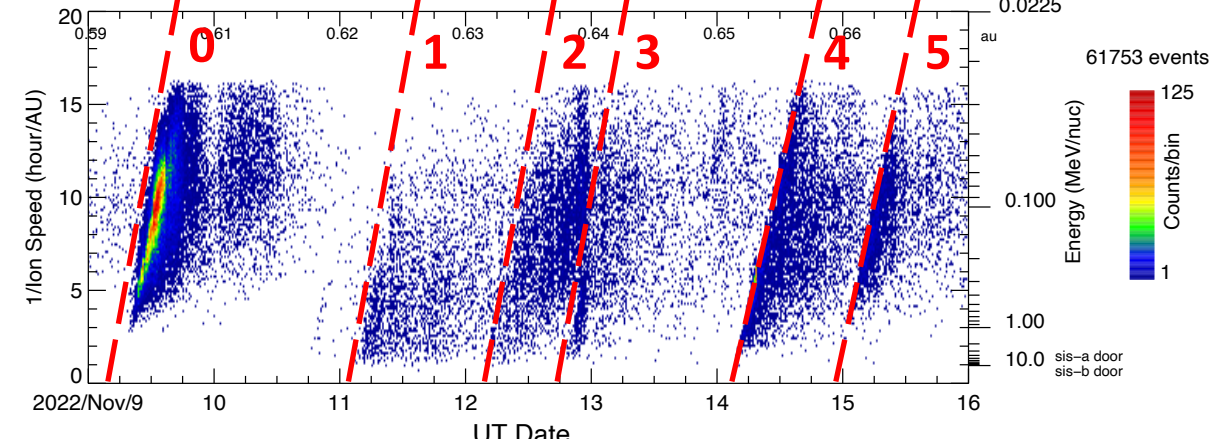
230-320 keV/n




0.4-10 MeV/n



10-70 AMU



	3He/4He (0.5-2.0 MeV/n)	Fe/O (0.32-0.45 MeV/n)
Event #0	2.19±0.52	2.66±0.20
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Event #4	0.84±0.02	1.04±0.10
Event #5	0.12±0.01	1.00±0.26
	24.33±9.12	--

Ho et al. (2019, Solar Phys. 294:33)

Conclusions

1- Statistics of the occurrence rate of impulsive events:

For 0.1-300 keV electron events, [Wang et al. \(2012\)](#) estimated ~ 190 events/yr at solar maximum (~ 0.5 events/day)
 ~ 12 events/yr at solar minimum (~ 1 event/month).

At an energy of ~ 40 keV, [Wang et al. \(2012\)](#) estimated over the whole Sun $\sim 10^4$ events/yr at solar maximum (~ 27 events/day);
 ~ 35 events/yr at solar minimum (< 3 events/month)

For 1.3-1.6 MeV/n 3He-rich events the quoted occurrence rate ~ 1000 3He-rich events/yr in the visible solar disk at maximum
(2.7 events/day) ([Reames et al. 1994](#))

In our case we have the record of 12 electron events on 11 November 2022 (five 3He-rich events at ~ 600 keV/n in 5 days)

2- Almost one-to-one coincidence of Type III radio, X-ray flares, and EUV jets at the origin of these events

Interesting enough the events on 9 Nov and late on 14 Nov rich in heavy ions at low energies (< 1 MeV/n) and lacking near-relativistic (> 30 keV) electrons did not show clear EUV jets, significant X-ray emission, and the Type III radio burst started at low frequencies (< 1 MHz)

3- CMEs (not the clearest CMEs in the world) able to generate significant ~ 2 MeV electron intensity increases (at least at Earth)

4- The complicated IP context medium, with a HSS and a SIR, does not prevent the arrival of particles at the 3 locations (especially at L1, although with more isotropic electron intensities)