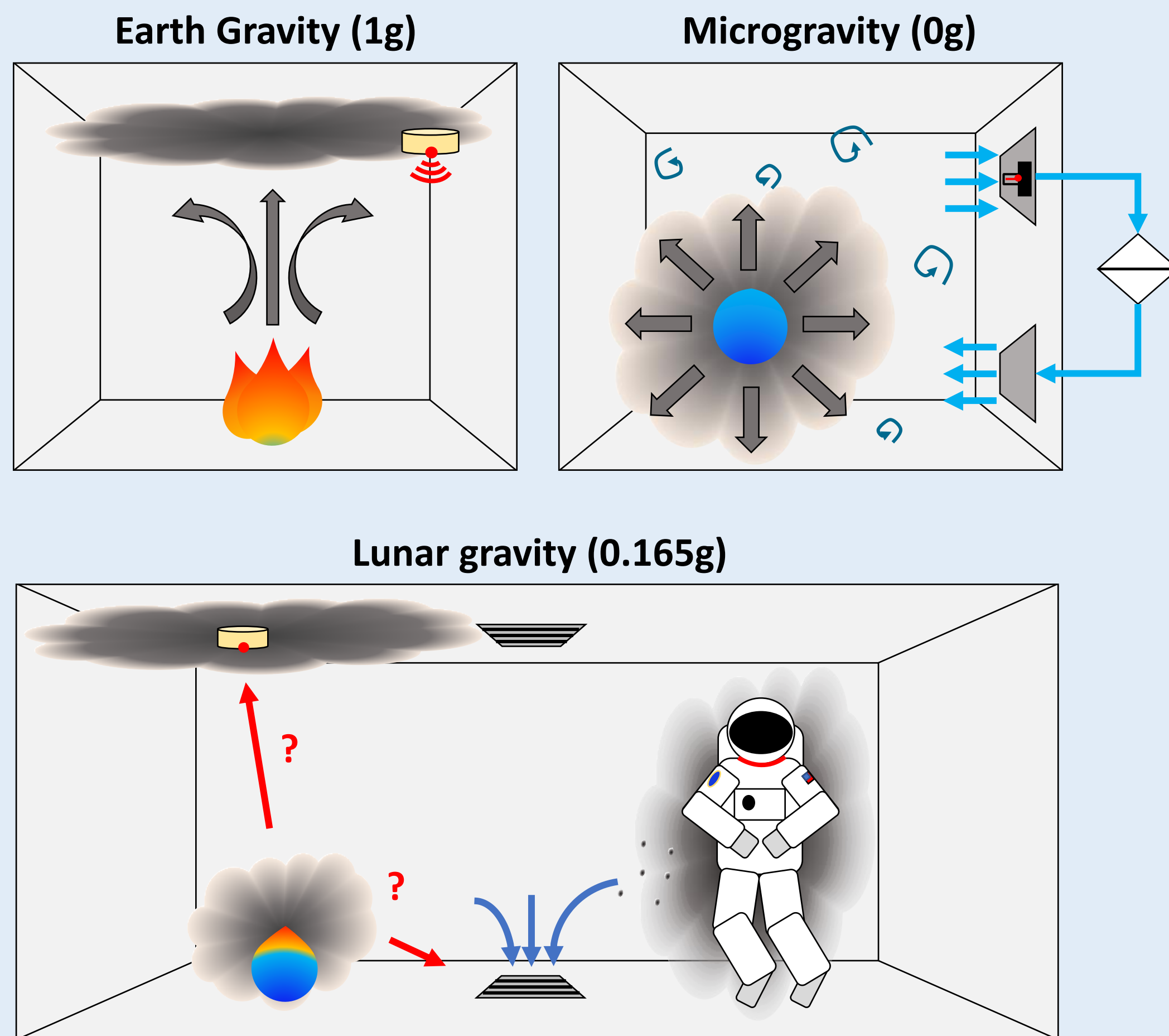


## Background and Motivation



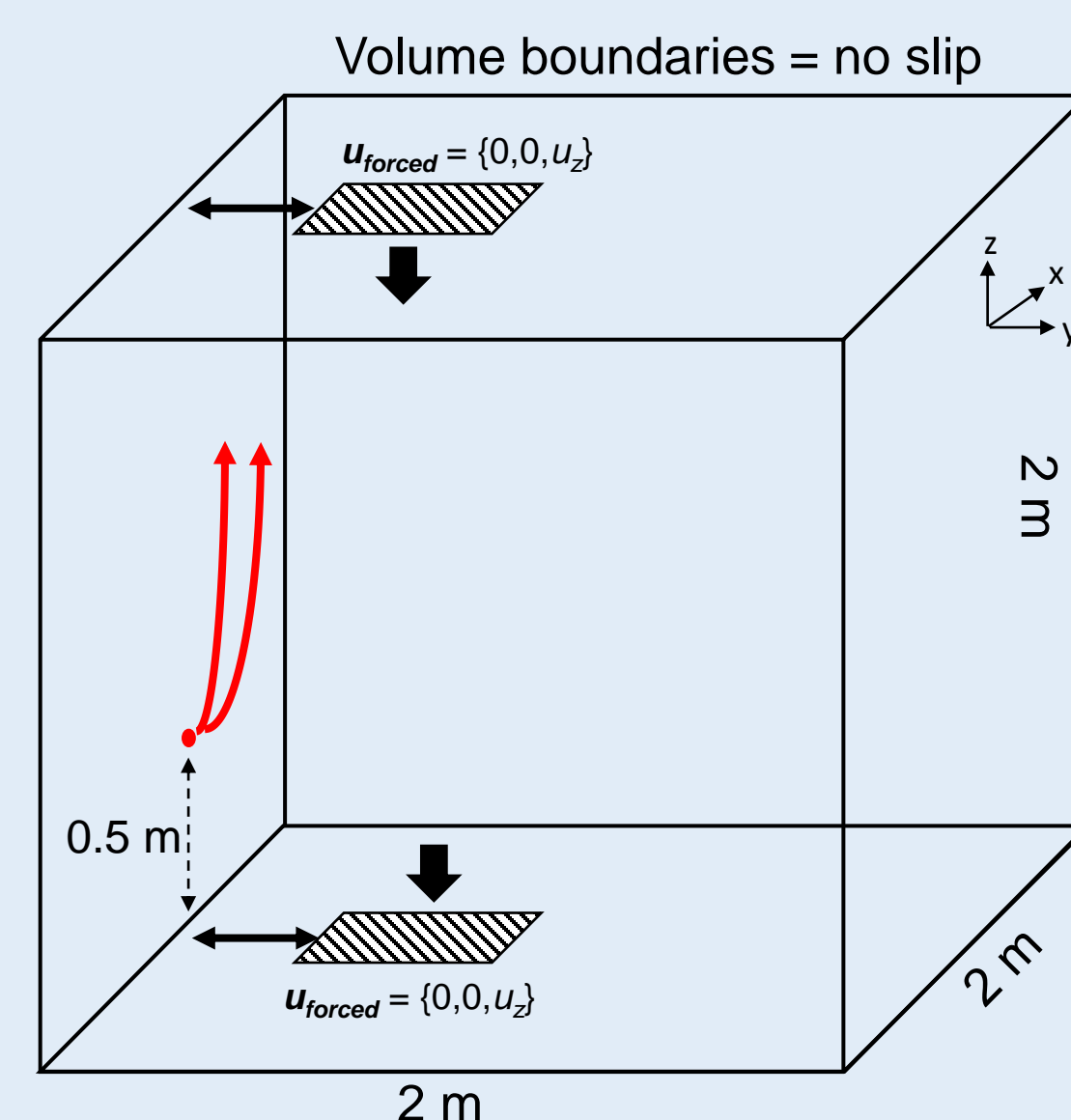
- Earth gravity: smoke stratifies and preconcentrates at ceiling
- Microgravity: no buoyancy, mixing driven by low-speed turbulent flow and high filtration rates
- Some buoyant plume flow expected in lunar gravity
- Plume transport will be governed by combination of buoyant flow and forced flow, which may be similar orders of magnitude
- How will lunar dust mitigation strategy affect smoke transport?
- Where should smoke detectors be placed in a lunar habitat?
- What should smoke detectors be set to? Times to alarm?

## Modeling Strategy

### Computations performed using COMSOL Multiphysics:

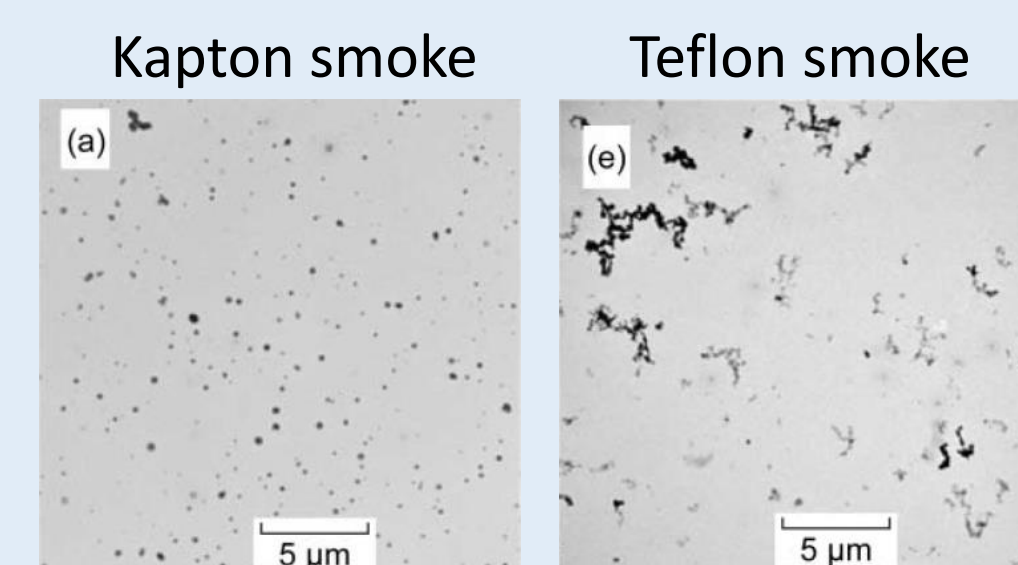
#### Study 1: Calculate flow field

- Fluid flow model: large eddy simulation (LES)
  - $P_{\text{ambient}} = 8.2 \text{ psia}$ ,  $T_{\text{ambient}} = 293.15 \text{ K}$
  - Gravity  $\mathbf{g} = \{0, 0, g_z\}$  where  $g_z = g_{\text{Lunar}}$  or  $g_{\text{Earth}}$
- Heat transfer in fluids
  - Walls = open boundaries,  $T = T_{\text{ambient}}$
  - Overheating 2 cm circular patch on wall
  - $T_{\text{fuel}}$  ranging from 600-1000K  $\pm 1 \text{ K}$  fluctuation
- Mass and heat transfer coupled through nonisothermal fluid flow multiphysics node
- Forced flow introduced from supply on/near ceiling and return on floor



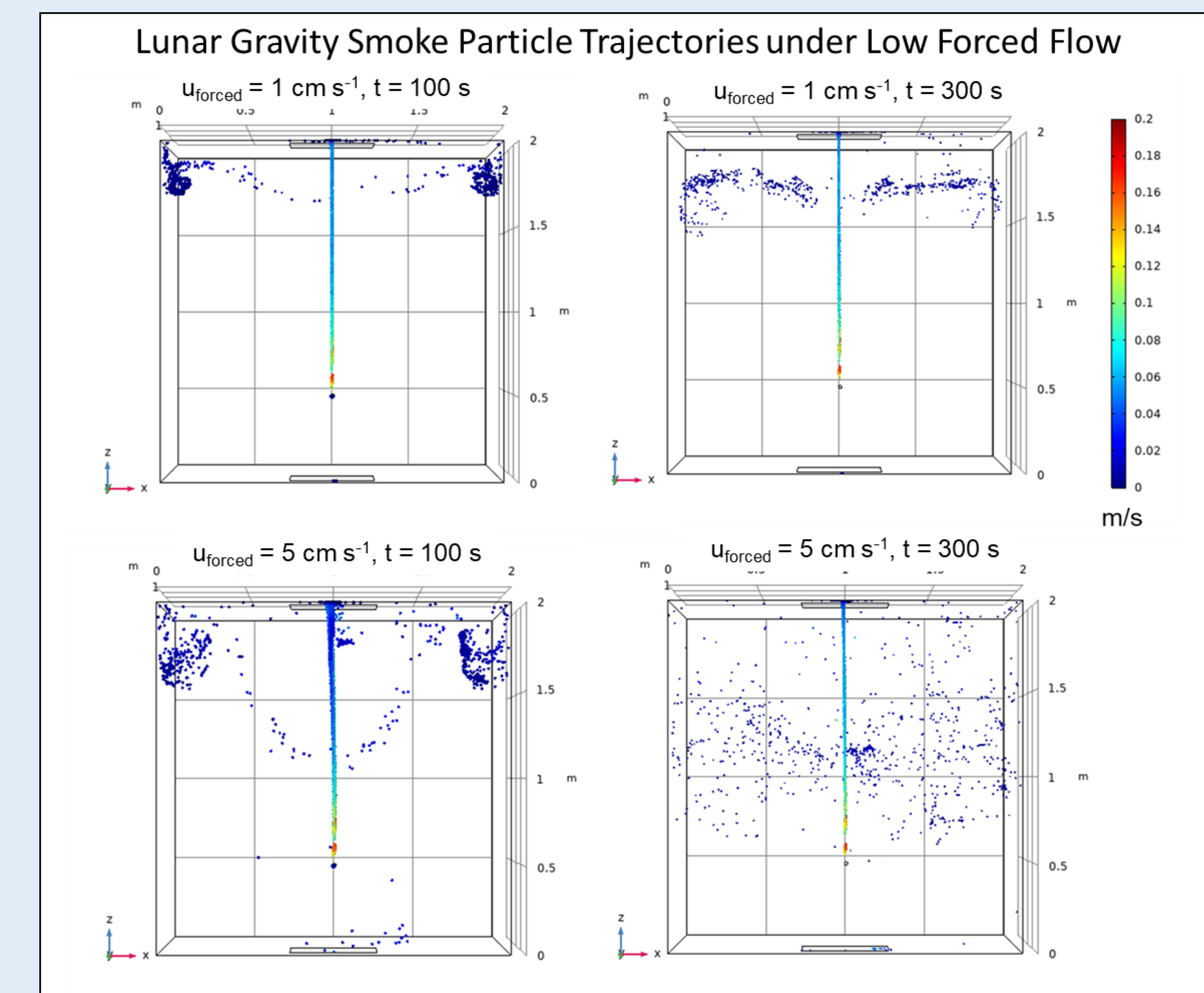
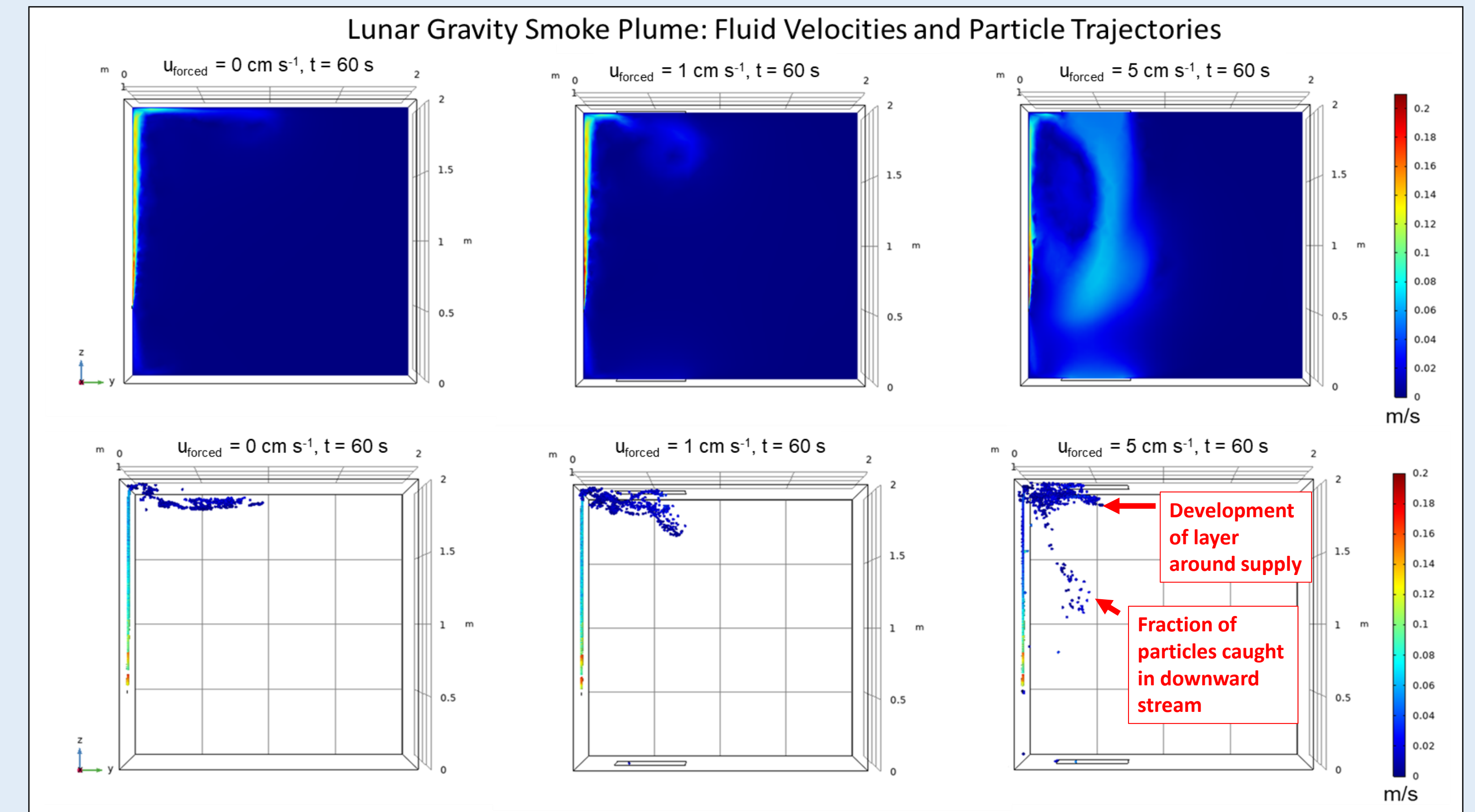
#### Study 2: Calculate particle trajectories

- Newtonian formulation:  $F_{\text{Total}} = F_{\text{Drag}} + F_{\text{Brownian}} + F_{\text{Gravity}}$
- Massless formulation: particle velocities are the same as velocity of surrounding fluid; **dramatically reduces computation time**
- Particle assumptions:  $d < 1 \mu\text{m}$ , spherical, density  $\sim 1\text{-}2 \text{ g/cm}^3$
- With these assumptions,  $Re_p < 0.5$  with external velocities of 100 cm/s



Meyer et al., Aerosol Sci. Tech., 49, 5, 299-309, 2015.

## Results and Discussion



- In lunar gravity, stratified plume layer is somewhat retained at  $u_{\text{forced}} = 1 \text{ cm/s}$  condition, before forced flow path develops
- At  $u_{\text{forced}} = 5 \text{ cm/s}$ , particles are mixed into volume at 5 minutes
- **US ECLSS velocity requirements for intramodule circulation: 5.1-20 cm/s (3.6-102 cm/s as lower/upper limits)**
- In a scenario where an early fire occurs near vent flow path, even low forced flow velocities (5 cm/s) can impede development of smoke layer at ceiling

## Conclusions and Future Work

### What does this mean for future Lunar habitats?

- In Lunar gravity, even very low forced flow velocities can disrupt smoke layer formation
- Smoke detectors **could** be placed on ceilings to facilitate earlier smoke detection, but detector placement relative to supplies and returns must be carefully considered

### What are the next steps?

- Evaluate different supply/return placements and range of tested flow velocities
  - How does smoke layer development change?
  - Is there an alternative to ceiling/floor placement?
  - **Simplifying assumptions about particle transport must be considered with each scenario**
- Explore different, more realistic early fire scenarios