

1 **Stratospheric temperature and ozone impacts of the Hunga Tonga-Hunga**  
2 **Ha’apai water vapor injection**

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15 Key Points:

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17 • The water vapor injection from the Hunga Tonga eruption impacts stratospheric temperature  
18 and ozone.

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20 • Largest additional ozone depletion is estimated to occur in Antarctic spring 2023. The ozone  
21 response may not be detectable in observations.

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23 • The impacts are projected to gradually diminish after 2024 as the excess water vapor is  
24 removed from the stratosphere.

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35 **Abstract.**

36 The January 2022 eruption of the Hunga Tonga-Hunga Ha’apai underwater volcano injected a  
37 large amount of water vapor into the mid-stratosphere. This study uses model simulations to  
38 investigate the resulting stratospheric impacts out to 2031. Maximum radiatively-driven model  
39 temperature changes occur in the Southern hemisphere (SH) subtropics in April-May 2022, with  
40 warming of ~1K in the lower stratosphere and cooling of 3K in the mid-stratosphere. The  
41 radiative cooling combined with adiabatic cooling driven by the quasi-biennial oscillation  
42 meridional circulation explains the near-record cold anomaly observed in the SH subtropical  
43 mid-stratosphere. Projected ozone responses maximize in 2023-2024 as the water vapor plume is  
44 transported globally throughout the stratosphere and mesosphere. The excess H<sub>2</sub>O increases the  
45 OH radical, causing a negative global ozone response (2-10%) in the upper stratosphere and  
46 mesosphere due to increased odd hydrogen-ozone loss, and a small positive ozone response (0.5-  
47 1%) in the mid-stratosphere due to interference of the NO<sub>x</sub> catalytic loss cycle by the additional  
48 OH. In the lower stratosphere, the excess H<sub>2</sub>O is projected to increase polar stratospheric clouds  
49 and springtime halogen-ozone loss, enhancing the Antarctic ozone hole by 25-30 DU in 2023.  
50 Arctic impact is small, with maximum additional ozone loss of 4-5 DU projected in spring 2024.  
51 These responses diminish after 2024 to be quite small by 2031, as the excess H<sub>2</sub>O is removed  
52 from the stratosphere with a 2.5-year e-folding time. Given the year-to-year variability of the  
53 stratosphere, the magnitudes of these ozone responses may be below the threshold of  
54 detectability in observations.

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59 Key Words: stratosphere, ozone depletion, Hunga Tonga Hunga Ha’apai eruption, water vapor,  
60 greenhouse gas

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67 **Plain Language Summary.**

68 Stratospheric ozone protects Earth’s biosphere from harmful ultraviolet radiation, and along with  
69 water vapor, are key components in determining temperature and chemistry of the atmosphere.

70 The January 2022 eruption of the Hunga Tonga-Hunga Ha’apai underwater volcano in the South  
71 Pacific injected water vapor into the atmosphere, increasing stratospheric water vapor by 10%. In  
72 this study, we use computer simulations of the stratosphere to project how this additional water  
73 vapor changed temperature and ozone in the months and years following the eruption. The water  
74 vapor cooled the middle stratosphere (roughly 14-25 miles above Earth's surface) and warmed  
75 the lower stratosphere (6-14 miles above the surface), with the largest changes of 2-5 degrees  
76 Fahrenheit in March-June 2022, several months after the eruption. The additional water vapor  
77 modified chemical processes that affect stratospheric ozone, leading to a projected 10-15%  
78 enhancement in the Antarctic ozone hole. This ozone hole enhancement is estimated to have  
79 maximized in October 2023, almost 2 years after the eruption, due to the slow circulation of  
80 stratospheric water vapor from the subtropics to the polar region. These impacts are expected to  
81 diminish after 2024 as the excess water vapor is slowly removed from the atmosphere by natural  
82 processes.

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## 98 **1. Introduction**

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100 Recent major volcanic eruptions have had large impacts on the stratosphere. The eruptions of El  
101 Chichón, Mexico in 1982, and Mt. Pinatubo, Philippines in 1991 injected substantial amounts of  
102 sulfur dioxide (SO<sub>2</sub>) into the stratosphere. The resulting increase in sulfate aerosol loading led to  
103 significant changes in stratospheric temperature (Pollack and Ackerman, 1983; Labitzke and  
104 McCormick, 1992; Angell, 1993; Randel et al., 1995; Rosenfield et al., 1997) and significant  
105 ozone depletion in the polar regions and globally (Hofmann and Solomon, 1989; Brasseur and  
106 Granier, 1992; Gleason et al., 1993; Randel et al., 1995; Solomon et al., 1996; Portmann et al.,  
107 1996; Rosenfield et al., 1997).

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109 The 15 January 2022 eruption of the Hunga Tonga-Hunga Ha'apai (HT) underwater volcano in  
110 the South Pacific (20.54°S, 175.38°W) was unique in that it injected a substantial amount of  
111 water vapor directly into the mid-stratosphere (Khaykin et al., 2022; Millán et al., 2022; Vömel  
112 et al., 2022), while only a modest amount of SO<sub>2</sub> was added (0.4-0.5 Tg, Witze, 2022; Taha et  
113 al., 2022). Satellite observations from the Aura microwave limb sounder instrument (MLS) show  
114 unprecedented stratospheric H<sub>2</sub>O concentrations following the eruption, with at least ~150 Tg of  
115 water vapor added to the stratosphere, a ~10% increase in the stratospheric burden (Millán et al.,  
116 2022).

117

118 Water vapor in the stratosphere is a greenhouse gas (Forster and Shine, 1999, 2002, Solomon et  
119 al., 2010), and is a source of the OH radical as well as polar stratospheric clouds, both of which  
120 strongly affect stratospheric ozone (Cohen et al., 1994; Jucks et al., 1998; Solomon et al., 2015).  
121 Because of these impacts, the enhanced H<sub>2</sub>O from the HT eruption significantly altered the  
122 radiative balance, dynamics, and photochemistry of the stratosphere in the months following the  
123 eruption (Vömel et al., 2022; Coy et al., 2022; Schoeberl et al., 2022, Sellitto et al., 2022;  
124 Schoeberl et al., 2023). Given the long lifetime of water vapor in the stratosphere, this  
125 perturbation is expected to have a multi-year impact.

126

127 In this paper, we use two-dimensional chemistry-climate model simulations to interpret some of  
128 the observed stratospheric responses to the HT H<sub>2</sub>O anomaly in 2022 through late 2023 and

129 estimate possible future impacts over the next decade. While the HT eruption also injected a  
130 modest amount of SO<sub>2</sub> which likely increased the stratospheric sulfate aerosol layer (Legras et  
131 al., 2022; Taha et al., 2022, Zhu et al., 2022), the focus of this study is the response due only to  
132 the water vapor injection. We quantify the projected temperature and ozone responses, as well as  
133 the impact on various chemical constituents important to stratospheric ozone chemistry. We also  
134 examine the dependence of the polar ozone response on background stratospheric conditions.

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## 137 **2. Model Simulations**

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### 139 **2.1 Model description**

140 Simulations in this study are conducted with the National Aeronautics and Space Administration  
141 (NASA)/Goddard Space Flight Center two-dimensional model (GSFC2D), which has been used  
142 in chemistry-climate coupling studies of the stratosphere and mesosphere as well as the World  
143 Meteorological Organization ozone assessments, including WMO (2022). The model has been  
144 described and evaluated previously and has been shown to provide realistic simulations of ozone,  
145 temperature, and transport-sensitive tracers for a variety of stratospheric perturbations  
146 (Bacmeister et al., 1995; Jackman et al., 1996, 2016; Rosenfield et al., 1997, 2002; Fleming et  
147 al., 2011; 2020). The model has very small internal variability, so that responses to small  
148 perturbations can be easily detected. In Appendix A, we provide a description of recent updates  
149 and model components important to the present study, including simulation of the quasi-biennial  
150 oscillation (QBO). We also provide an evaluation and comparison of the model with  
151 observations.

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### 153 **2.2 Water vapor simulation**

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155 For this study we include two water vapor tracers in the model for 2022. An unperturbed H<sub>2</sub>O  
156 includes all chemical production and loss in the stratosphere and mesosphere and is specified  
157 below the tropopause (seasonally and latitudinally dependent). In the troposphere, the 21-year  
158 average (1981–2001) of relative humidity data from the European Center for Medium-Range  
159 Weather Forecasts updated reanalysis (ERA-40) is used for the surface to 12 km, and the Upper

160 Atmosphere Research Satellite monthly reference atmosphere (Randel et al., 2001) is used for 12  
161 km to the tropopause. This unperturbed H<sub>2</sub>O does not interact with other model components.

162  
163 To simulate the HT water vapor perturbation in the model, we use MLS version 4 (v4) data as  
164 input. Previous analysis suggests that MLS v4 provides a more suitable data product to estimate  
165 the HT H<sub>2</sub>O anomaly compared to the most recent version 5 data (Millán et al., 2022). We  
166 determine the zonal mean MLS water vapor anomaly in the latitude-height domain for 15  
167 January 2022 through 31 December 2022 as the difference for each day from the 2005-2021  
168 average for a given month. The difference of the pre-eruption 1-14 January 2022 average from  
169 the 1-14 January 2005-2021 average is also removed at each latitude and altitude. This accounts  
170 for the long-term trend in water vapor and is generally +0.1-0.4 ppm depending on location. The  
171 MLS quality and convergence filters flagged some profiles during the first few weeks after the  
172 eruption, however, the water vapor enhancements were independent of the quality and  
173 convergence filtering after 8 February (Millán et al., 2022).

174  
175 This daily MLS-derived H<sub>2</sub>O anomaly is added to the unperturbed model H<sub>2</sub>O at each time step  
176 to create a separate perturbed H<sub>2</sub>O tracer that interacts with the model chemistry, radiation, and  
177 dynamics. This perturbed interactive H<sub>2</sub>O is updated in this manner on 15 January through 31  
178 December 2022. This simulates the evolving impacts of the water vapor anomaly through the  
179 end of 2022 using the MLS observations as input at each time step. Starting 1 January 2023, the  
180 evolution of the perturbed H<sub>2</sub>O is fully model-computed, with the MLS-based field on 31  
181 December 2022 serving as the initial condition.

182  
183 We also ran a parallel baseline simulation without the MLS-derived HT water vapor anomaly.  
184 The stratospheric response to the HT H<sub>2</sub>O perturbation is taken as the difference between the  
185 perturbation and baseline simulations.

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190 **3. Results**

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### 192 **3.1 Water vapor anomaly and temperature response in 2022**

193 The water vapor anomaly mixes and disperses throughout the global stratosphere during 2022, as  
194 seen in the MLS-derived observations for selected months (Figure 1, top panels). While the  
195 initial H<sub>2</sub>O injection at 20°S reached the upper stratosphere on 15 January (Carr et al., 2022), by  
196 April the anomalous water vapor was mainly confined to the mid-stratosphere in a latitude band  
197 from ~40°S to 25°N. By August and especially December, the upward tropical bulge and  
198 downward midlatitude bulge of the plume reflect transport by the Brewer-Dobson circulation  
199 (BDC) as noted previously in Schoeberl et al., 2022. The Southern hemisphere (SH) polar vortex  
200 remained strong and isolated through much of spring 2022, confining the water vapor anomaly to  
201 latitudes equatorward of ~60°S (Khaykin et al., 2022; Manney et al. 2023). The plume mixed  
202 into the polar region following the vortex break-up in late November and December (Figure 1e).

203

204 The corresponding model temperature response is significant and is highly correlated with the  
205 H<sub>2</sub>O anomaly (Figure 1, bottom panels). The increased stratospheric water vapor enhances the IR  
206 cooling above ~40 hPa (~23 km) and warming below. Maximum temperature changes occur near  
207 20°S in April-May 2022, with the largest cooling of -3.2K at 20 hPa and largest warming of ~1K  
208 at 54 hPa. Warming of up to several tenths of a degree K occurs around the tropical tropopause.  
209 The temperature response spreads vertically and horizontally and decreases in magnitude  
210 throughout 2022, following the dispersal of the water vapor plume in latitude and altitude  
211 (Figure 1).

212

### 213 **3.2 Interaction with the QBO circulation**

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215 In this section we examine how the model temperature response to the HT H<sub>2</sub>O anomaly  
216 compares with the NASA Modern-Era Retrospective Analysis for Research and Applications 2  
217 reanalysis (MERRA-2, Gelaro et al., 2017). Since the meridional circulation associated with the  
218 QBO has a significant impact in the tropical and subtropical stratosphere (Plumb and Bell, 1982),  
219 we also examine how this circulation impacts the HT temperature response. To isolate the HT  
220 and QBO-induced temperature responses, the long-term average seasonal cycle is removed from  
221 MERRA-2 and the model simulations in Figures 2-4.

222

223 MERRA-2 assimilated temperatures show that the SH mid-stratosphere cooled substantially in  
224 the months following the eruption (Coy et al., 2022). By May 2022, temperatures of nearly 5K  
225 below the seasonal average occur at 20°-25°S, 20 hPa (denoted by “A” in Figure 2a). This is  
226 nearly 4 standard deviations ( $\sigma$ ) colder than average at this location (denoted by “A” in Figure  
227 2b), with a significant area in the mid-stratosphere at 15°-40°S greater than  $2\sigma$  colder than  
228 average. This contrasts with most of the stratosphere outside of the polar regions which was  
229 generally within  $\pm 1\sigma$  of the long-term average in May 2022 (Figure 2b). We note that  
230 temperature anomalies derived from MLS version 5 data (not shown) are very similar to those  
231 shown in Figure 2.

232

233 The cooling in the mid-stratosphere is shown in the context of the MERRA-2 record since 2010  
234 in Figure 3a (black line). The cooling starts shortly after the 15 Jan 2022 eruption and increases  
235 for several months to a maximum of -5K ( $4\sigma$ ) in mid-May 2022. This is the second coldest  
236 period during the entire 1980-2022 MERRA-2 de-seasonalized data record, surpassed only by a  
237 shorter duration period of a few weeks during September 2019.

238

239 While much of the SH mid-stratospheric cooling was caused by the HT water vapor anomaly,  
240 some of the cooling was also likely due to the circulation associated with the QBO (Plumb and  
241 Bell, 1982, Baldwin et al., 2001). In May 2022, the QBO was in an easterly phase below ~20  
242 hPa, and the corresponding meridional circulation was characterized by sinking motion in the  
243 mid-stratosphere over the equator and rising motion in the subtropics and mid-latitudes in both  
244 hemispheres. This is depicted by the streamlines in Figure 2a-b which show the de-seasonalized  
245 residual mean meridional and vertical winds of the Transformed Eulerian-Mean formulation  
246 (e.g., Andrews et al., 1987). This circulation and the associated adiabatic heating and cooling  
247 contribute to the temperature anomalies at 40-10 hPa: warm over the equator and cold anomalies  
248 at ~10°-40° in both hemispheres.

249

250 To further examine the impact of the water vapor injection and QBO circulation on these  
251 observed temperature anomalies, we utilize model simulations with an interactively computed  
252 QBO included (details of the QBO simulation are provided in Appendix A). Here, the model



253 equatorial zonal wind is in roughly the same QBO phase as seen in the Singapore radiosonde  
254 observations and MERRA-2 reanalysis in May 2022, with easterlies below ~20 hPa ([https://acd-  
255 ext.gsfc.nasa.gov/Data\\_services/met/qbo/qbo.html#singau](https://acd-ext.gsfc.nasa.gov/Data_services/met/qbo/qbo.html#singau)). The corresponding model  
256 meridional QBO circulation and associated temperature anomalies are qualitatively consistent  
257 with MERRA-2. If the HT water vapor anomaly is not included, the SH cold anomaly is weaker  
258 than observed throughout the SH mid-latitudes and specifically at the location of maximum  
259 cooling at 20°S, 20 hPa (Figure 3a, blue line; “A” in Figure 4c). The temperature impact of the  
260 HT water vapor anomaly in isolation shows cooling at 40-10 hPa from the SH mid-latitudes to  
261 the Northern Hemisphere (NH) sub-tropics and warming below ~40 hPa (Figure 4d).

262  
263 Including both the QBO and the water vapor anomaly brings the model into very good  
264 agreement with the MERRA-2 mid-stratosphere warm and cold anomalies in the tropics and SH  
265 low-mid latitudes (Figures 4a-b). The model also captures quite well the MERRA-2 extreme cold  
266 anomaly at 20°S, 20 hPa following the eruption through May 2022, as well as the return to near  
267 normal seasonal temperatures by the end of August 2022 (Figure 3a, red line). For the cold  
268 temperature anomaly at 20°S, 20 hPa in May 2022 (~-5K), we estimate from the model that  
269 ~60% (-3K) is caused by IR cooling of the excess H<sub>2</sub>O, and ~40% (-2K) is caused by ascent of  
270 the QBO-induced circulation. The model with both the QBO and H<sub>2</sub>O anomaly included also  
271 qualitatively captures the warm and cold anomalies in MERRA-2 in the lower stratosphere  
272 below ~40 hPa at 30°S-30°N (Figure 4a-b). These model simulations confirm that the very cold  
273 mid-stratosphere anomaly in the SH subtropics in May 2022 is caused by the HT water vapor  
274 anomaly but is reinforced by the ascent and adiabatic cooling caused by the meridional  
275 circulation associated with the QBO being in an easterly phase in May 2022.

276  
277 By August 2022, the cold temperature anomaly deepened and shifted to mid-latitudes (40°S-  
278 60°S) and was prevalent throughout the SH stratosphere in the MERRA-2 data (Figure 2c-d).  
279 The MERRA-2 de-seasonalized temperatures at 40°S, 27 hPa are nearly 8K (6 $\sigma$ ) colder than  
280 average during mid-August 2022 (“B” in Figure 2c-d, Figure 3b). This is significantly colder  
281 than anytime during the entire 1980-2022 MERRA-2 data record (Coy et al., 2022). However,  
282 the corresponding model simulation with the QBO and water vapor anomaly (Figure 3b, red line)  
283 captures only ~50% of the magnitude of the MERRA-2 August cold temperature anomaly.

284

285 We note that throughout the SH winter, the mid-upper stratosphere at mid-latitudes was  
286 characterized by significantly weaker than normal wave-forced drag on the zonal mean flow, as  
287 represented by the Eliassen-Palm (E-P) flux divergence due to resolved waves (Andrews et al.,  
288 1987) derived from the MERRA-2 reanalysis (not shown). The corresponding mid-latitude  
289 residual vertical velocity field had significantly weaker than average descent at levels above ~40  
290 hPa, as indicated by the anomalous upward motion (streamlines) in this region in August  
291 (Figures 2c-d). It is therefore likely that this anomalous extratropical wave forcing contributed to  
292 the extreme cold temperature anomaly observed at midlatitudes during the SH winter. However,  
293 determining the source of this anomalous wave forcing is outside the scope of the 2D model used  
294 in this study.

295

### 296 **3.3 Long term water vapor and temperature response**

297

298 The excess H<sub>2</sub>O is transported to the SH mid-high latitudes in the months following the eruption,  
299 with values of 2-2.5 ppm reaching the SH polar region by late 2022 with the breakup of the polar  
300 vortex (Figure 5a). The model simulated H<sub>2</sub>O (starting 1 January 2023) maintains concentrations  
301 of ~2.5 ppm throughout the SH high latitudes in January 2023, before slowly diminishing to ~2  
302 ppm by March 2023. This decrease is due to the slow removal of the excess stratospheric water  
303 vapor to the troposphere as the anomaly reaches the SH mid-high latitude tropopause in early  
304 2023. Water vapor decreases to slightly less than the baseline in early June 2023 at SH high  
305 latitudes (Figure 5a) with the establishment of the polar vortex and onset of PSC formation. The  
306 negative anomalous mixing ratios are due to the increased sedimentation of ice PSCs with the  
307 enhanced water vapor relative to the baseline. Positive anomalous mixing ratios return to the SH  
308 polar region in spring 2023 with the breakup of the vortex and in-mixing of mid-latitude air.

309

310 In the NH, the MLS-derived H<sub>2</sub>O anomaly is primarily confined equatorward of ~30°N until late  
311 December 2022, when increased planetary wave activity mixes the plume into the Arctic (Figure  
312 5a). This process continues in the model simulation (starting in 2023), with anomaly values of ~1  
313 ppm poleward of 60°N throughout March and April 2023.

314

315 The model temperature response at 25 hPa (Figure 5b) is highly correlated with the H<sub>2</sub>O plume.  
316 The tropics and SH cool by 0.5-3 K during 2022-early 2023, with 0.4-1K cooling in the NH  
317 polar region in early 2023. In the lower stratosphere at 67 hPa (Figure 5c), warming of 0.4-1K is  
318 confined to the tropics and SH lower latitudes in 2022, with a small warming of 0.1-0.2K outside  
319 the polar regions (50°S-50°N) during 2023. The water vapor and temperature anomalies  
320 gradually diminish with time, with the H<sub>2</sub>O anomaly reduced to less than 0.1 ppm globally by  
321 mid-2029. Mostly small cooling of 0.1-0.2K occurs throughout the stratosphere after 2023,  
322 except in the SH polar region where larger temperature changes occur in response to the  
323 enhanced ozone hole. This will be discussed in section 3.5.

324  
325 For 1-3 years following the eruption, the water vapor anomaly is slowly transported upwards by  
326 the BDC, with mixing ratios of 1-1.5 ppm reaching the mesosphere by late 2023 (Figure 6a).  
327 Here we show the global average since the anomalies generally have similar patterns across most  
328 latitude zones (the ozone response in Figure 6c will be discussed in section 3.4). Global  
329 temperature changes again follow the H<sub>2</sub>O plume, with cooling of ~1K confined to the mid-  
330 stratosphere during 2022 and cooling of 1-1.5K in the mesosphere in mid-2023 through the end  
331 of 2024 (Figure 6b). Small global-mean warming of 0.1-0.15K occurs in the upper  
332 troposphere/lower stratosphere just after the eruption through late 2023. Starting in 2024, the  
333 global average temperature response is almost exclusively negative (cooling) throughout the  
334 middle atmosphere as the anomalies gradually diminish to be quite small by 2030.

335  
336 The total global burden (mass) of the H<sub>2</sub>O anomaly is 150-160 Tg shortly after the eruption  
337 through mid-2023 and is projected to decrease thereafter (Figure 7). The excess stratospheric  
338 water vapor is slowly removed by sedimentation of PSCs within the Antarctic vortex, as well as  
339 return of stratospheric air to the troposphere by the BDC at mid-high latitudes of both  
340 hemispheres. The combination of these processes leads to an exponential decay of the anomaly,  
341 with an estimated average e-folding time of 2.5 years from mid-2023 through 2031 (Figure 7, red  
342 dashed-dotted line). The burden is reduced to 4.7 Tg by the end of 2031, which is ~3% of its  
343 starting value in January 2022.

344

### 345 **3.4 Global profile ozone and related chemical responses**

346

347 Ozone is affected by the HT water vapor anomaly globally throughout the middle atmosphere  
348 due to changes in photochemistry, both directly, and indirectly via changes in the temperature-  
349 dependent ozone loss cycles (e.g., Dvortsov and Solomon, 2001; Brasseur and Solomon, 2005).  
350 MLS observations show that ozone in the mid stratosphere decreased significantly in the SH  
351 subtropics and midlatitudes starting in early winter 2022 (Wang et al., 2022). However, our  
352 model simulations suggest that the ozone response to the HT H<sub>2</sub>O anomaly in this region is quite  
353 small in 2022. This is discussed further in section A.4.3 of Appendix A, along with comparisons  
354 to the MLS data.

355

356 For globally averaged ozone, maximum projected changes occur in 2023-2024, ~1-2 years after  
357 the eruption (Figure 6c). During this time, the H<sub>2</sub>O anomaly is 20-25% above the background in  
358 the SH mid-stratosphere, and ~15% above the background globally throughout the mesosphere  
359 (Figure 8a). One of the direct consequences of excess H<sub>2</sub>O is the increase in odd hydrogen  
360 species ( $\text{HO}_x = \text{H} + \text{OH} + \text{HO}_2 + 2 * \text{H}_2\text{O}_2$ ), with the OH radical increasing by 5-10% throughout the  
361 stratosphere and mesosphere in 2023-2024 (Figure 8b). This enhances the HO<sub>x</sub> catalytic loss  
362 cycle, which is the major contributor to the total odd oxygen (O+O<sub>3</sub>) chemical loss above ~50  
363 km (Figure 9b). The resulting projected global ozone loss is >0.5% at altitudes above ~40 km  
364 throughout 2023-2026 (Figure 6c) and increases with altitude to 5-10% above ~60 km from mid-  
365 2023 to mid-2025.

366

367 Changes in the total odd nitrogen family

368 ( $\text{NO}_y = \text{N} + \text{NO} + \text{NO}_2 + \text{NO}_3 + 2 * \text{N}_2\text{O}_5 + \text{HNO}_3 + \text{HO}_2\text{NO}_2 + \text{HONO} + \text{ClONO}_2 + \text{BrONO}_2$ ; Figure 9a,  
369 orange line) and a subset of NO<sub>y</sub> directly involved in the odd nitrogen-ozone loss cycle

370 ( $\text{NO}_x = \text{N} + \text{NO} + \text{NO}_2 + \text{NO}_3 + 2 * \text{N}_2\text{O}_5$ ; Figure 9a, blue line) are mostly negative throughout the

371 stratosphere and mesosphere. Decreases in NO<sub>x</sub> are 5-10% in the SH polar lower stratosphere  
372 due mainly to increased sedimentation of nitric acid tri-hydrate (NAT) PSCs in the Antarctic  
373 vortex (e.g., Toon et al., 1986), with smaller NO<sub>x</sub> decreases of 2-4% in the Arctic (Figure 8d).

374 Decreases in global NO<sub>x</sub> maximize in the mid-stratosphere mainly due to increased OH which  
375 converts NO<sub>x</sub> to HNO<sub>3</sub> via the OH+NO<sub>2</sub> reaction (Figure 9a). There is also a small contribution  
376 to this NO<sub>x</sub> decrease due to a slight increase in the heterogeneous reaction

377  $\text{N}_2\text{O}_5 + \text{H}_2\text{O} \rightarrow 2 * \text{HNO}_3$  on sulfates. Although the model stratospheric sulfate aerosol surface area  
378 is specified and does not interact with the  $\text{H}_2\text{O}$  anomaly, the total rate of this reaction is slightly  
379 faster (1-3%) at 20-30 km at mid-high latitudes due to the increased water vapor.

380

381 In the mesosphere where the  $\text{HNO}_3$  concentration is very small and  $\text{NO}_x \approx \text{NO}_y$ , odd nitrogen  
382 decreases by 2-5% mainly due to the colder temperatures (Figures 8d, 9a). Here, the abundance  
383 of atomic nitrogen (N) is increased due to the reduced rate of the strongly temperature dependent  
384 reaction  $\text{N} + \text{O}_2 \rightarrow \text{NO} + \text{O}$  at lower temperatures. The increased N increases the loss of  $\text{NO}_y$  which  
385 is controlled by the reaction  $\text{N} + \text{NO} \rightarrow \text{N}_2 + \text{O}$  (Rosenfield and Douglass, 1998).

386

387 The model ozone changes in the mid-stratosphere (30-5 hPa, ~25-37 km) are projected to be  
388 predominantly positive, starting shortly after the eruption in early 2022 and lasting through 2029  
389 (Figures 6c). Maximum global ozone increases of 0.5-1% occur in mid-2022 through the end of  
390 2024. In this region, the  $\text{NO}_x$  catalytic cycle dominates the total odd oxygen chemical loss.

391 Because of reduced  $\text{NO}_x$ , anomalous  $\text{NO}_x$ -odd oxygen loss is positive throughout the  
392 stratosphere (Figure 9b, blue line), and the dominance of this loss cycle in the mid-stratosphere  
393 leads to a small positive total odd oxygen chemical tendency (Figure 9b, black line) and positive  
394 ozone change in this region (Figure 8e). This result is qualitatively consistent with previous  
395 studies that found reduced mid-stratospheric  $\text{NO}_x$ -ozone loss due to increased  $\text{HO}_x$   
396 concentrations from methane oxidation (Nevison et al., 1999; Randeniya et al., 2002).

397

398 In the lower stratosphere below ~23 km (40 hPa) the ozone response is projected to be mostly  
399 negative (Figures 6c and 8e), with the global response mainly reflecting a deepened Antarctic  
400 ozone hole. This will be discussed in the next section.

401

402

### 403 **3.5 Antarctic profile ozone and related chemical responses**

404

405 MLS data show that lower stratospheric Antarctic ozone during spring 2022 was significantly  
406 lower than the 2005-2021 average (Figure 10d). Previous model results show that this low ozone  
407 could be explained by the HT aerosol perturbation combined with colder than average

408 temperatures in the polar vortex (Wang et al., 2022). For the model simulations presented in this  
409 study, the SH planetary wave forcing (section A.2) was reduced during winter-spring 2022 to  
410 mimic the strong and isolated SH polar vortex that persisted well into November. However, the  
411 resulting additional ozone depletion due to the HT H<sub>2</sub>O anomaly was significantly smaller than  
412 shown in the de-seasonalized MLS data in spring 2022 (Figure 10c-d). The model shows some  
413 qualitative consistency with the MLS ozone during the first half of 2023, with a negative  
414 anomaly at 10-20 km and a positive anomaly at 20-25 km.

415

416 In late winter-spring 2023, the model simulates a deepened ozone hole, with a significant  
417 negative anomaly at ~10-25 km which is qualitatively consistent with MLS at most altitudes  
418 (Figure 10c-d). However, the persistence of a positive anomaly seen in MLS in a shallow layer  
419 near 20 km through mid-October 2023 is not captured in the model. Some of these differences  
420 are likely due to the background stratospheric variability which makes it difficult to detect the  
421 impact of the excess H<sub>2</sub>O in the de-seasonalized MLS ozone data. Model biases, at least for the  
422 first 1-1½ years after the eruption, also may be due in part to not including the HT aerosol  
423 perturbation in the model.

424

425 The processes that cause the ozone hole have been well established in past studies (e.g.,  
426 Solomon, 1999; Solomon et al., 2014, Solomon et al., 2015). To examine in more detail how the  
427 excess H<sub>2</sub>O impacts the model ozone hole, we focus on lower stratospheric anomalies for April-  
428 December 2023 of several constituents relevant to the chemistry driving the enhanced ozone  
429 hole. By mid-winter 2023, the water vapor anomaly is present within the Antarctic vortex  
430 throughout the depth of the stratosphere with corresponding anomalous cooling (Figures 10a-b).  
431 The combination of the H<sub>2</sub>O and temperature anomalies leads to projected: 1) increase in type I  
432 NAT PSCs in early winter with a corresponding decrease in gas-phase HNO<sub>3</sub>, and 2) increase in  
433 type II ice PSCs throughout the winter (Figures 11a-b). The increased sedimentation of NAT  
434 PSCs causes denitrification (e.g., Toon et al., 1986) starting in early June, with negative  
435 anomalous NO<sub>y</sub> persisting through winter and spring and negative anomalous NO<sub>x</sub>  
436 concentrations in October-November (Figure 11b).

437

438 The projected enhancement in PSCs and PSC surface area increases the heterogeneous  
439 conversion of chlorine and bromine from reservoir species (HCl, ClONO<sub>2</sub>, HOBr, BrONO<sub>2</sub>) to  
440 reactive forms that destroy ozone (Cl, ClO, Br, BrO). The anomalous heterogeneous chlorine  
441 activation on sulfate aerosols is also generally faster due to the lower temperatures and increased  
442 water vapor (e.g., Solomon, 1999; Burkholder et al., 2019), even though the model sulfate  
443 aerosol surface area is specified and does not interact with the H<sub>2</sub>O anomaly. The anomalous rate  
444 of a key heterogeneous reaction, ClONO<sub>2</sub> + HCl → HNO<sub>3</sub> + Cl<sub>2</sub> on PSCs and sulfates, is shown  
445 in Figure 11b (purple dotted line). The anomaly maximizes in June following the anomalies in  
446 PSC surface area and enhancement of the reaction on sulfates. The secondary anomaly maximum  
447 in September-October is caused by an increase in the sulfate reaction due to lower temperatures  
448 (Figure 11e), as the ice and NAT PSC and H<sub>2</sub>O anomalies are all quite small in the spring.  
449 Increased conversion of chlorine to reactive forms on sulfate aerosols under cold SH polar  
450 conditions was noted previously (e.g., Hanson et al., 1994).

451

452 The chemical loss of odd oxygen due to the chlorine and bromine catalytic cycles is enhanced  
453 with increased sunlight in early August through late October, and this controls the total chemical  
454 loss (Figures 11c-d). The additional ozone depletion (Figure 11d, black line) drives a significant  
455 reduction in the solar ultraviolet ozone heating and temperature starting in mid-late August as the  
456 solar elevation increases (Figure 11e). This further enhances the chemical ozone loss and  
457 reduction in heating, leading to a delay in the breakup of the polar vortex. This delays the  
458 increase of odd oxygen due to transport into the polar region associated with the vortex breakup  
459 (Figure 11d, green dashed-dotted line), which in turn further enhances the reduction of polar  
460 ozone prior to the vortex breakup. We note that the lower stratospheric Antarctic spring cold  
461 temperature anomaly occurs yearly throughout 2022-2029 (Figure 5c).

462

463 Anomalous ozone concentrations at 54 hPa reach a minimum of ~-5 DU/km in late October  
464 (Figure 11d, black line), before recovering to near-baseline values by mid-December. The  
465 recovery is driven mainly by transport associated with the vortex breakup as indicated by the  
466 transport tendency (Figure 11d, green dashed-dotted line), with a smaller contribution due to  
467 reduced chemical loss as indicated by the positive total chemical tendency (Figure 11d, magenta  
468 line). This positive chemical tendency is due to the reduced anomalous NO<sub>x</sub> concentrations and

469 a decrease in the NO<sub>x</sub> catalytic loss cycle in October-November (Figure 11b-c, blue lines). NO<sub>y</sub>  
470 and NO<sub>x</sub> return to near-baseline concentrations in mid-December following the vortex breakup  
471 and in-mixing of NO<sub>y</sub>-rich midlatitude air.

472  
473 In the middle stratosphere (~40-10 hPa), positive model ozone anomalies in the Antarctic occur  
474 from December 2022-June 2023 (Figure 10c), mainly due to the reduced NO<sub>x</sub> catalytic loss  
475 cycle (i.e., increased odd oxygen tendency, Figure 12 blue line). A negative ozone anomaly  
476 driven by the increased halogen loss cycles occurs during August-September 2023. Relatively  
477 small ozone changes occur during late September-October as the positive transport tendency  
478 largely offsets the anomalous chemical ozone loss. Positive ozone anomalies re-emerge in the  
479 mid-stratosphere during December 2023 through June 2024 (Figure 10c), driven by a  
480 combination of reduced NO<sub>x</sub> loss and positive anomalous transport tendency. This anomalous  
481 polar transport occurs during the SH late spring-summer and is a dynamical response to the  
482 enhanced ozone hole. This feature has been discussed in previous modeling studies (Kiehl et al.,  
483 1988; Mahlman et al., 1994; Stolarski et al., 2006), and has been seen in observations of  
484 temperature (Randel and Wu, 1999) and ozone (Stolarski et al., 2006). For the HT H<sub>2</sub>O response,  
485 the additional ozone depletion leads to a projected delay in 1) the spring vortex breakup, and 2)  
486 the corresponding wave-forced drag on the zonal mean flow and acceleration of the BDC. As a  
487 result, there is anomalous descent in the Antarctic mid-upper stratosphere and positive odd  
488 oxygen transport tendency relative to the baseline in late November-December 2023 (Figure 12).  
489 The anomalous descent and associated adiabatic warming cause the positive temperature  
490 anomaly in the polar mid-stratosphere above ~40hPa during November-December 2023 (Figure  
491 10b). This warm anomaly is a yearly recurring feature in the Antarctic summer mid-stratosphere  
492 throughout 2023-2030 (Figure 5b).

493  
494 In the very lower stratosphere below ~70 hPa (~18 km), small negative ozone anomalies persist  
495 through summer and fall 2024 (Figure 10c) as the transport processes associated with the delayed  
496 vortex breakup do not return ozone quite to the baseline value at these altitudes.

497  
498

### 499 **3.6 Total ozone response**



500

### 501 3.6.1 Baseline response

502

503 The additional ozone loss due to the HT water vapor injection results in a projected 25-30 DU  
504 enhancement of the Antarctic ozone hole during spring 2023 (Figure 13). The response  
505 diminishes to 20-25 DU in spring 2024, with decreasing severity in the ozone hole enhancement  
506 thereafter as the water vapor perturbation diminishes. In 2022, a strong and isolated SH polar  
507 vortex persisted well into November, with large observed ozone losses and a deeper than normal  
508 ozone hole relative to the historical data record (e.g., Newman and Lait, 2023). However as  
509 discussed in section 3.5, the H<sub>2</sub>O anomaly resulted in a relatively small model ozone hole  
510 enhancement of 5-7 DU in spring 2022 (Figure 13). This springtime ozone loss was limited since  
511 the additional H<sub>2</sub>O was confined to the SH mid-stratosphere during mid-winter and did not reach  
512 the polar lower stratosphere until late November 2022 (Figure 10a; see also Manney et al., 2023)  
513 following the polar vortex breakup, after the time of normal seasonal formation of PSCs.

514

515 In the Arctic, the anomalous water vapor and stratospheric cooling cause additional springtime  
516 ozone depletion due to processes that are generally like those in the Antarctic as discussed above.  
517 However, the projected Arctic ozone losses are much smaller, 4-5 DU, and do not start until  
518 spring 2024 (Figure 13). The anomalous Arctic ozone depletion diminishes with time to be < 1  
519 DU by the late 2020s. At mid-latitudes, maximum total ozone decreases are 4-6 DU in the SH  
520 and 1-2 DU in the NH, with changes of less than  $\pm 1$  DU in the tropics throughout the post-  
521 eruption period.

522

### 523 3.6.2 Sensitivity to stratospheric background conditions

524

525 The projected ozone responses shown in Figures 13 are from a model simulation that uses the  
526 standard planetary wave forcing which gives climatologically averaged stratospheric conditions  
527 (section A.2). This simulation is generally consistent with the observed long term average  
528 seasonal cycle in polar total ozone in each hemisphere (Figure 14, orange vs. black solid lines).  
529 However, the polar regions are characterized by large interannual variability in winter and spring  
530 caused primarily by variations in planetary wave driving in the stratosphere. This is depicted by

531 the gray shading in Figure 14 which shows the range in historical total ozone observations for  
532 1991-2022 (Newman and Lait, 2023).

533

534 To examine the dependence of the ozone response on the background stratospheric conditions,  
535 we ran a series of experiments with the stratospheric wave driving varied to mimic the observed  
536 range in total ozone shown in Figure 14. Here we focus on the model year with the largest  
537 anomalous ozone loss in each polar region: 2023 for the SH and July 2023-June 2024 for the  
538 NH. Substantially increased planetary wave forcing gives warm polar stratospheric conditions  
539 and total ozone at the upper end of the range in the historical data record in both hemispheres  
540 (Figure 14, red solid lines). Conversely, substantially reduced wave forcing results in a cold  
541 polar stratosphere and total ozone at the lower end of the data record (blue solid lines). For each  
542 wave driving case, the water vapor anomaly simulation is depicted by the dashed lines in Figure  
543 14.

544

545 In the SH, including the HT H<sub>2</sub>O perturbation under the strong wave forcing (warm) conditions  
546 results in somewhat less projected ozone depletion compared to the standard wave forcing, with  
547 polar cap average additional depletion of 13 DU vs. 20 DU in 2023 (Figure 14a). The stronger  
548 wave driving, and warmer stratospheric conditions also promote a faster return to the baseline  
549 total ozone in December 2023 compared to the standard case. Under the weak wave forcing  
550 (cold) conditions, the baseline already has very low ozone concentrations in the lower  
551 stratosphere, so that including the H<sub>2</sub>O anomaly has less of an impact than with the standard  
552 wave forcing, with additional polar cap average depletion of 14 DU (Figure 14a, blue dashed  
553 line). The largest impact occurs under conditions slightly warmer than the standard case (not  
554 shown). Here, the larger ozone concentrations available in the baseline combined with  
555 substantial anomalous chlorine and bromine activation resulted in additional polar cap depletion  
556 of 23 DU, compared to 20 DU for the standard wave forcing case.

557

558 In the Arctic, the projected model ozone response to the additional water vapor has a small  
559 dependence on the background stratospheric conditions in spring 2024 (Figure 14b). The colder  
560 conditions with weak wave driving result in 5 DU additional polar cap average depletion  
561 compared to the standard (3 DU) and strong (3 DU) wave driving cases. However, even under

562 the cold conditions of the weak wave driving case, the additional model ozone depletion in the  
563 Arctic is small compared to the Antarctic. This is due to a combination of factors, one being the  
564 smaller anomalous H<sub>2</sub>O concentrations transported to high NH latitudes. Just prior to the onset of  
565 PSC formation in early winter (November-December 2023), the excess water vapor is 0.2-1 ppm  
566 in the Arctic lower stratosphere, substantially less than the 1-2 ppm in the early winter Antarctic  
567 (May-June 2023). Another important factor is the generally warmer Arctic temperatures that  
568 limit additional ozone depletion (Solomon et al., 2014). For example, in the weak wave driving  
569 (cold) case, model Arctic temperatures throughout the lower stratosphere in February-April 2024  
570 are still 3-7K warmer compared to the Antarctic strong wave driving (warm) case in late winter –  
571 spring 2023. We note that these model NH ozone responses to the H<sub>2</sub>O anomaly are generally  
572 consistent with previous studies of the Arctic ozone response to stratospheric water vapor  
573 changes using 3-D chemical transport models driven by meteorological reanalysis (e.g., Vogel et  
574 al., 2011; Thölix et al., 2018).

575

576 Figure 14 suggests that the total ozone impact of the HT H<sub>2</sub>O anomaly is generally smaller than  
577 the year-to-year variability characteristic of the stratospheric polar regions during winter and  
578 spring. The standard deviation of the observed SH October-November polar cap total ozone is  
579 ~35 DU during 1991-2022, significantly larger than the maximum model estimated response to  
580 the H<sub>2</sub>O anomaly of 23 DU. The difference is even larger in the Arctic spring, with a March-  
581 April observed standard deviation of ~26 DU compared to a maximum model anomaly response  
582 of 5 DU. Therefore, it is possible that the response to the HT water vapor injection may not be  
583 easily detectable above the background variability in observational total ozone data.

584

585

#### 586 **4. Summary and Conclusions**

587

588 The January 2022 Hunga Tonga-Hunga Ha'apai volcanic eruption increased stratospheric water  
589 vapor by ~10% (~150 Tg) (Millán et al., 2022) which significantly altered the radiative balance,  
590 dynamics, and photochemistry of the stratosphere (Vömel et al., 2022; Coy et al., 2022;  
591 Schoeberl et al., 2022, Sellitto et al., 2022; Schoeberl et al., 2023). In this study, we examine

592 how this unique natural perturbation impacted stratospheric temperature and ozone in the first 1-  
593 2 years following the eruption and estimate possible future responses over the next decade.

594

595 The maximum radiatively-induced model temperature response occurs in March-June 2022,  
596 several months after the eruption, with a cooling of 2-3K in the SH mid-stratosphere, and ~1K  
597 warming in the lower stratosphere. This radiatively-driven warming is as much as several tenths  
598 of a degree K around the tropical tropopause, which may have important implications for the  
599 amount of water vapor entering the stratosphere. However, quantification of this effect is beyond  
600 the scope of the 2D model used in this study. We note that these impacts do not include sea  
601 surface temperature feedback, which may be important to fully quantify the response to the  
602 radiative forcing of the HT H<sub>2</sub>O anomaly.

603

604 The QBO was in an easterly phase in April-May 2022, and model simulations suggest that ascent  
605 and adiabatic cooling associated with the QBO circulation, combined with the radiative cooling  
606 of the H<sub>2</sub>O anomaly, can explain the near-record cold temperatures seen in the MERRA-2  
607 reanalysis in the SH subtropical mid-stratosphere during May 2022.

608

609 Transport of the water vapor plume to the Antarctic lower stratosphere is delayed until late 2022  
610 with the breakup of the polar vortex (Manney et al., 2023). The plume reaches the Arctic in  
611 winter 2022-2023 and is slowly transported to the mesosphere during 2023-2024 by the rising  
612 branch of the Brewer-Dobson circulation. Radiatively-induced temperature changes are projected  
613 to be small in the NH stratosphere ( $<\pm 0.5\text{K}$ ), with larger cooling in the mesosphere of 1-1.5K  
614 that peaks in late 2023-2024. The increased water vapor and cooling in the mesosphere may have  
615 implications for polar mesospheric cloud formation (e.g., Hervig et al., 2016; Lübken et al.,  
616 2018).

617

618 The anomalous water vapor impacts the chemistry in the middle atmosphere by increasing OH  
619 concentrations. This increases the odd hydrogen-ozone loss cycle, which is dominant in the  
620 mesosphere, causing a projected 5-10% reduction in mesospheric ozone globally from mid-2023  
621 to mid-2025. The additional OH also increases the conversion of odd nitrogen to HNO<sub>3</sub>, thereby  
622 reducing the NO<sub>x</sub>-ozone loss cycle throughout the stratosphere. This results in a small net global

623 ozone increase of 0.5-1% during 2023-2024 in the mid-stratosphere where the NO<sub>x</sub> catalytic  
624 cycle dominates the chemical loss of odd oxygen.

625  
626 In the lower stratosphere, the additional water vapor is projected to increase PSC surface area  
627 and the heterogeneous conversion of chlorine and bromine into reactive forms that destroy  
628 ozone. This causes a deepened ozone hole in the Antarctic spring. The decrease in solar ozone  
629 heating reduces temperatures and delays the break-up of the vortex, further enhancing the ozone  
630 loss. This effect is projected to maximize in 2023 with 25-30 DU enhancement of the ozone hole.  
631 Model sensitivity simulations suggest that this response will be somewhat dependent on the  
632 background stratospheric conditions, with an estimated polar cap average range of 13-23 DU  
633 additional ozone depletion. In the Arctic, additional ozone losses due to the excess water vapor  
634 are relatively small, with an estimated maximum additional depletion of 3-5 DU in spring 2024  
635 and slightly smaller losses of 2-4 DU in spring 2025 and 2026. The Arctic responses have a  
636 relatively small dependence on the background stratospheric conditions.

637  
638 By mid-2023 and beyond, the excess H<sub>2</sub>O is slowly removed by return to the troposphere at mid-  
639 high latitudes and by sedimentation of PSCs within the Antarctic vortex. The anomaly decays  
640 exponentially with a projected e-folding time of 2.5 years, and the corresponding temperature  
641 and ozone responses diminish slowly after 2024. By the end of 2031, the additional H<sub>2</sub>O is  
642 estimated to be ~3% of its initial value of ~150 Tg, with very small temperature and ozone  
643 responses.

644  
645 The focus of the present paper is on the response due only to the HT H<sub>2</sub>O anomaly in isolation so  
646 that sulfur impacts are not considered. The HT eruption injected a modest amount of SO<sub>2</sub> (0.4-  
647 0.5 Tg) which likely increased the stratospheric sulfate aerosol layer (Legras et al., 2022; Taha et  
648 al., 2022; Zhu et al., 2022). This could impact ozone chemistry via increased heterogeneous  
649 conversion of odd nitrogen to HNO<sub>3</sub>, and increased chlorine activation at lower temperatures  
650 (~195K). Stratospheric temperatures can also be affected via changes in ozone and aerosol  
651 absorption of infrared radiation. These impacts can occur globally and in the polar regions as  
652 shown for previous volcanic eruptions (e.g., Brasseur and Granier, 1992; Randel et al., 1995;  
653 Solomon et al., 1996; Portmann et al., 1996). For the HT eruption, Wang et al. (2022) showed

654 that significant additional ozone depletion occurred in the Antarctic lower stratosphere in spring  
655 2022 when including the SO<sub>2</sub> injection in model simulations. Therefore, the ozone response to  
656 the H<sub>2</sub>O injection presented here likely somewhat underestimates the full response to the HT  
657 eruption, at least in the first year following the eruption.

658

659 The model-projected ozone responses presented here are generally smaller than the natural  
660 variability of the stratosphere. The largest response is likely the enhancement of the Antarctic  
661 ozone hole. However, the additional model ozone loss is at most 23 DU (2023 polar cap  
662 average), which is smaller than the standard deviation of springtime polar total ozone  
663 observations in the SH (~35 DU). Therefore, it is possible that any ozone response to the HT  
664 water vapor injection may not be detectable above the background variability in observational  
665 data sets.

666

667

668

## 669 **Appendix A: GSFC2D Model Description and Evaluation**

670

671 In this appendix, we provide a description and evaluation of the GSFC2D model, focusing on  
672 recent updates and model components important to the present study. We also provide some  
673 model evaluation via comparisons with observations of age of air, H<sub>2</sub>O, and ozone.

674

### 675 A.1 Model chemistry and radiation

676

677 The model has full stratospheric chemistry, with a diurnal cycle computed for all constituents  
678 each day. Transport of all constituents follows the chemistry calculations at each time step within  
679 the diurnal cycle. This is updated from the previous scheme in which only the diurnal averages  
680 were transported at the end of the diurnal cycle. The resulting changes are most significant at  
681 mid-high latitudes during late autumn through early spring when the photochemical time scales  
682 are long, and the new methodology improves the simulations of polar ozone compared with  
683 observations (see Figure 14). The model domain extends from the surface to ~92 km (.002 hPa)  
684 with a grid spacing of 4° latitude and 1 km in altitude.

685

686 Time dependent surface mixing ratio boundary conditions are taken from WMO (2022) for the  
687 major ozone depleting substances and Meinshausen et al. (2020) for the major greenhouse gases  
688 CO<sub>2</sub>, CH<sub>4</sub>, and N<sub>2</sub>O. The latest JPL-2019 recommendations (Burkholder et al., 2019) are used for  
689 the kinetic reaction rates, photolysis cross sections, and heterogeneous reactions on the surfaces  
690 of polar stratospheric clouds (PSCs) and stratospheric sulfate aerosols. The stratospheric aerosol  
691 surface area density is specified from the CMIP6 dataset based on Kovilakam et al. (2020).

692

693 A parameterization to simulate type Ia (solid nitric acid trihydrate, NAT) and type II (Ice) PSC  
694 formation follows Considine et al. (1994) and includes sedimentation of NAT and ice aerosols.  
695 Calculations of PSC occurrence frequency and surface area density use a combination of the  
696 zonal mean temperatures computed in the model and longitudinal temperature probability  
697 distributions (deviations from the zonal mean, T') obtained from the MERRA-2 reanalysis. Here,  
698 the MERRA-2 climatological T' distribution is based on daily data averaged over 2000-2020 and  
699 is added to the model-computed zonal mean temperatures at each time step. This hybrid  
700 methodology allows the PSCs to respond to both observational-based longitudinal variations and  
701 the evolving model zonal mean temperature. To convert the PSC concentration to surface area  
702 density, a lognormal particle size distribution is assumed (Considine et al., 1994). The uptake  
703 coefficients for heterogeneous reactions on PSCs are specified to be constant (Burkholder et al.,  
704 2019). We note that supercooled ternary solution (STS, H<sub>2</sub>SO<sub>4</sub>/HNO<sub>3</sub>/H<sub>2</sub>O) type Ib PSCs are not  
705 computed in the current model configuration.

706

707 Effects of spherical geometry in the photolysis and solar heating rate calculations are  
708 approximated by use of the Chapman function (McCartney, 1976), accounting for twilight  
709 conditions for solar zenith angles up to 94°. For the infrared (IR) parameterization, the model  
710 uses the Rapid Radiative Transfer Model for GCM Applications (RRTMG), a state-of-the-art  
711 algorithm that is used in various climate and weather forecast models (Mlawer et al., 1997;  
712 Clough et al., 2005; Hurwitz et al., 2015).

713

714 A.2 Model dynamical parameterizations

715

716 The planetary wave parameterization (Bacmeister et al., 1995; Fleming et al., 2011) uses lower  
717 boundary conditions at 750 hPa (~2 km) of geopotential height amplitude and phase for zonal  
718 wave numbers 1-4. These are derived as a function of latitude and season using a 30-year  
719 average (1991-2020) of MERRA-2 data for the standard model wave forcing. These boundary  
720 conditions can be adjusted to modify the planetary wave forcing in the stratosphere, thereby  
721 giving colder or warmer than average conditions during the winter and spring in each  
722 hemisphere. In section 3.6.2, we examine the sensitivity of the HT H<sub>2</sub>O anomaly on the  
723 background stratospheric conditions using this modified wave forcing.

724

725 Momentum deposition and vertical eddy diffusion from breaking gravity waves in the  
726 stratosphere and mesosphere are computed following the parameterization originally developed  
727 by Lindzen (1981) and modified by Holton and Zhu (1984). The parameterization solves for a  
728 general spectrum of monochromatic waves with phase speeds covering the range of  $\pm 40$  m/sec at  
729 intervals of 10 m/sec. The momentum flux is specified in the upper troposphere (~325 hPa) as a  
730 function of latitude and season, with the slower phase speed waves having larger momentum flux  
731 (Holton and Zhu, 1984). The parameterization also solves separately for a single stationary  
732 gravity wave generated by flow over orography. Here, vertical profiles of momentum deposition  
733 are computed on a longitude-latitude grid ( $10^\circ \times 4^\circ$  grid spacing), with the model zonal mean  
734 zonal wind used for each longitude. The orographic surface forcing is based on a multi-year  
735 average of monthly zonal gravity wave surface stress from the earlier version 3 of the Whole  
736 Atmosphere Community Climate Model (Garcia et al., 2007; see also McFarlane, 1987).

737

738 The zonally averaged momentum deposition from gravity waves and planetary waves is used in  
739 the 2D model zonal wind and meridional circulation calculations. The resulting model zonal  
740 wind and temperature distributions compare well with multi-year averaged monthly  
741 meteorological reanalysis (e.g., Fleming et al., 2011).

742

### 743 A.3 Model QBO simulation

744

745 Simulation of the QBO in equatorial zonal wind is based on previous 2D model studies which  
746 parameterize the momentum deposition from thermally damped large scale, long period Kelvin



747 and Rossby-gravity waves (e.g., Plumb and Bell, 1982; Gray and Pyle, 1989; Dunkerton, 1997).  
748 We include two slow Kelvin waves (zonal wavenumber 2) with phase speeds of +20 m/sec and  
749 +30 m/sec, and a Rossby-gravity wave (zonal wavenumber 4) with a phase speed of -40 m/sec. A  
750 fast Kelvin wave with phase speed of +60 m/sec is also included to simulate the westerly phase  
751 of the semiannual oscillation in the upper stratosphere and lower mesosphere.

752

753 Previous work has shown the importance of including the momentum flux from small scale  
754 gravity waves in generating a realistic QBO (Dunkerton, 1997; Geller et al., 2016). We use the  
755 parameterization described in section A.2 with high vertical resolution (250 meters) to compute  
756 the momentum flux from a spectrum of equatorial gravity waves with phase speeds covering the  
757 range of  $\pm 40$  m/sec at intervals of 2 m/sec (41 waves).

758

759 For each large scale and gravity wave component, the momentum flux is specified at the bottom  
760 boundary in the upper troposphere ( $\sim 325$  hPa), with the slower phase speed gravity waves  
761 having larger momentum flux (Dunkerton, 1997). As discussed in previous studies (Geller et al.,  
762 2016), the input parameters of wave phase speed and lower boundary momentum flux,  
763 respectively, are adjusted to obtain a QBO with realistic amplitude and period. The model  
764 equatorial zonal wind shows good agreement with Singapore radiosonde data and the MERRA-2  
765 reanalysis (Coy et al., 2016) in reproducing the general features of the QBO, including a period  
766 of  $\sim 28$  months and similar rate of downward phase progression. This is shown in a representative  
767 comparison with MERRA-2 in Figure A1. Note that the model is free running and does not  
768 correspond to a specific year as is represented in the reanalysis. The maximum model zonal wind  
769 QBO amplitude of  $\sim 20$  m/sec occurs at 10-30 hPa, decreases to  $\sim 5-6$  m/sec at 70 hPa, and is near  
770 zero at the tropical tropopause.

771

772 Momentum forcing from the different wave components is specified to decrease rapidly away  
773 from the equator, with a latitudinal dependence for the large-scale waves as in Gray and Pyle  
774 (1989). The resulting model QBO amplitude has a latitudinal variation consistent with  
775 observations (Wallace, 1973), with a half width of 10-15 degrees. The meridional circulation  
776 associated with the QBO (Plumb and Bell, 1982) is also consistent with observations, as seen in  
777 the circulation-induced temperature changes over the equator and in the subtropics discussed in

778 section 3.2 (Figure 4). This circulation is also important for tracer transport in this region (Trepte  
779 and Hitchman, 1992; Randel et al., 1998; Baldwin et al., 2001).

780  
781

#### 782 A.4 Comparison of model tracers with observations

783

784 In this section, we provide an evaluation of the model transport fields by comparing the age of  
785 air and H<sub>2</sub>O simulations with observations. We also show the model ozone response to the HT  
786 H<sub>2</sub>O anomaly compared with MLS observations at SH low-middle latitudes during 2022-2023.

787

##### 788 A.4.1 Age of Air

789

790 Stratospheric mean age of air is a widely used diagnostic that tests the overall fidelity of model  
791 transport. Figure A2 shows the age of air derived from measurements of SF<sub>6</sub> and CO<sub>2</sub> at 20 km  
792 (~50 hPa) and vertical profiles for three latitudes zones during the 1990s (Hall et al., 1999),  
793 along with the model simulation. There are differences in the observations at middle and higher  
794 latitudes that may reflect photochemical influences on SF<sub>6</sub> which would cause an overestimation  
795 in the inferred ages (Hall and Waugh, 1998). Some of the older age measurements at 65°N may  
796 also reflect remnants of the polar vortex (Ray et al., 1999).

797

798 For the most part GSFC2D compares generally well with the observations in reproducing the  
799 absolute values and the latitudinal and vertical gradients. The model slightly underestimates the  
800 observations at 20 km at 30°N-45°N, and at NH high latitudes above 30 km. However, the good  
801 overall agreement illustrates that the model stratospheric transport rates, including the relative  
802 magnitudes of vertical motion and horizontal mixing, are generally realistic. This is also  
803 important to provide a reasonable simulation of the H<sub>2</sub>O anomaly decay rate (Figure 7).

804

##### 805 A.4.2 Water vapor

806

807 Here we compare the full water vapor field (background plus HT anomaly) for 2022-2023 from a  
808 simulation in which the model is forced with the MLS-derived H<sub>2</sub>O anomaly (section 2.2), but

809 only through the end of February 2022. Starting 1 March 2022, evolution of the full H<sub>2</sub>O field is  
810 model computed.

811

812 The model simulates the full water vapor field generally well in the latitude-height domain  
813 compared with MLS version 5 (v5, Figure A3), as well as the time evolution through November  
814 2023 in the mid-stratosphere (Figure A4). The model reproduces the observed transport of large  
815 H<sub>2</sub>O concentrations (> 7 ppm) associated with the HT anomaly from the initial injection to SH  
816 mid-high latitudes through November 2023. The H<sub>2</sub>O anomaly appears in the Arctic as a small  
817 enhancement in the MLS data starting in early 2023, and this feature is qualitatively similar in  
818 the model. The model underestimates the MLS observations by ~0.5 ppm throughout much of  
819 the stratosphere, likely reflecting the amount of H<sub>2</sub>O entering through the tropical tropopause in  
820 the model which is specified from the Upper Atmosphere Research Satellite monthly reference  
821 atmosphere (Randel et al., 2001). Some of the model differences with MLS are also due to  
822 atmospheric variability not resolved in the simulation. Isolation of the tropical stratosphere  
823 below ~20 hPa tends to be somewhat overestimated in the model, as indicated by the stronger  
824 latitudinal gradients at 15°S-15°N compared with MLS, especially during 2022 (Figure A4).

825

826 The model simulates the Antarctic H<sub>2</sub>O distribution generally well compared with MLS (Figure  
827 A5), although the model has a small high bias in the late winter-spring below ~23 km. Some of  
828 this may be due to interannual variability and the model not fully resolving the observed  
829 isolation of the SH polar region through late spring 2022 (Manney et al., 2023) and through  
830 September 2023, as well as possible underestimation of ice sedimentation.

831

832 Both the MLS data and model show a return to more typical low H<sub>2</sub>O values (< 4-5 ppm) in the  
833 equatorial mid stratosphere by early 2023 (Figure A4). This reflects the isolation of the tropics  
834 from mid-latitudes and the upward transport of drier air from the tropopause, as seen in the  
835 equatorial time-height sections (Figure A6a-b). As discussed in section 3.2, the model zonal  
836 wind QBO in Figure A6b is in roughly the same phase as the Singapore radiosonde data, with  
837 easterlies at 30 hPa at the time of the eruption through mid-2022 (“E” in Figure A6a-b). The  
838 corresponding QBO-induced circulation, with relative descending motion over the equator, slows  
839 the overall ascent in the tropical mid-stratosphere. As a result, upward transport of the enhanced

840 H<sub>2</sub>O is relatively slow throughout most of 2022, and the model compares well with the MLS data  
841 in this regard (Figure A6a-b). By late 2022 – early 2023, the QBO is in a westerly phase (“W” in  
842 Figure A6a-b) so that the QBO-induced circulation enhances the overall ascent in the tropics.  
843 This results in more rapid upward transport of the H<sub>2</sub>O plume with large water vapor  
844 concentrations (> 7 ppm) in the equatorial upper stratosphere through November 2023 in both  
845 MLS and the model.

846  
847 A model sensitivity test reveals that this upward transport of the H<sub>2</sub>O plume is somewhat  
848 dependent on the phase of the QBO at the time of the eruption. With the model QBO in a  
849 westerly phase in January 2022 (“W” in Figure A6c), the accompanying circulation causes more  
850 rapid upward transport of the plume during the first half of 2022 compared to the easterly phase  
851 in Figure 6a-b. This upward transport during the westerly phase slows in late 2022 and into 2023  
852 as the QBO shifts to an easterly phase (“E” in Figure A6c). This simulation suggests that  
853 different phases of the QBO at the time of the eruption can cause a H<sub>2</sub>O variation of  $\pm 1$ -3 ppm in  
854 the equatorial mid-upper stratosphere, at least during the first  $\sim 1\frac{1}{2}$  years after the eruption  
855 (Figure A6d). This dependence gradually fades after 2023 (not shown) as the excess H<sub>2</sub>O decays  
856 and mixes throughout the global stratosphere.

#### 857 858 A.4.3 Ozone

859  
860 MLS observations show that ozone in the mid-stratosphere decreased significantly in the SH  
861 subtropics and midlatitudes starting in early winter 2022. Mixing ratios decreased by up to 0.4-  
862 0.5 ppm ( $\sim 10$ -12%) at 30-40 hPa in August 2022 (Figure A7a). Wang et al. (2022) suggest that  
863 this anomalous low ozone was mainly due to circulation impacts (i.e., slowing of the BDC)  
864 rather than chemical effects due to the HT SO<sub>2</sub> and H<sub>2</sub>O injection. GSFC2D model simulations  
865 with the QBO and HT H<sub>2</sub>O anomaly included qualitatively reproduce many of the positive and  
866 negative ozone anomalies seen in the MLS data averaged over 15°S-40°S (Figure A7b).  
867 However, the anomaly magnitudes are generally underestimated in the model, and some of this  
868 bias may be due to not including the HT sulfur injection in the simulation. Most of the model  
869 anomalies are driven by circulation effects associated with the QBO (Figure A7c). However, the  
870 impact of the HT H<sub>2</sub>O anomaly in isolation is quite small, with ozone decreases of at most 0.03

871 ppm (-0.5%) at 30-40 hPa (Figure A7d). This result is generally consistent with Wang et al.  
872 (2022) as most of the model ozone decrease at 15°S-40°S in winter 2022 is due to circulation  
873 effects.

874  
875

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884 **Data Availability Statement.** Total ozone satellite observations used in this manuscript  
885 (Newman and Lait, 2023) are available on the NASA ozone watch website:

886 <https://ozonewatch.gsfc.nasa.gov/>.

887 The MERRA-2 reanalysis data products (GMAO, 2015) are available from the NASA Goddard  
888 Earth Sciences Data and Information Services Center (GES DISC):

889 <https://disc.gsfc.nasa.gov/datasets?project=MERRA-2>.

890 The Aura/MLS data products are available from the NASA Goddard Earth Sciences Data and  
891 Information Services Center (GES DISC) for H<sub>2</sub>O version 4 (Lambert et al., 2015) at:

892 <https://doi.org/10.5067/Aura/MLS/DATA2009>; H<sub>2</sub>O version 5 (Lambert et al., 2020) at:

893 <https://doi.org/10.5067/Aura/MLS/DATA2508>; and ozone version 5 (Schwartz et al., 2020) at:

894 <https://doi.org/10.5067/Aura/MLS/DATA2516>.

895 GSFC2D model description, configuration, input parameters and forcing datasets, and associated  
896 references are provided in section 2 of the main text and Appendix A. GSFC 2D model output  
897 used in the generation of the figures of this paper are available at

898 [https://portal.nccs.nasa.gov/datashare/trop-str/pub/HT\\_H2O](https://portal.nccs.nasa.gov/datashare/trop-str/pub/HT_H2O)

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1237 **Figure captions**

1238  
1239 **Figure 1.** *Top panels:* water vapor anomaly derived from MLS v4 observations for the months  
1240 indicated. Contours are 0.1, 0.5, 1, 2, 4, 6, and 8 ppm. *Bottom panels:* model temperature  
1241 response, taken as the difference between simulations with and without the water vapor  
1242 anomaly. Negative contours are -3, -2, -1, -0.5, and -0.2 K, and positive contour intervals are in  
1243 0.1K increments. The thick black solid line indicates the tropopause.

1244  
1245 **Figure 2.** *Top panels:* monthly mean de-seasonalized temperature (color shading) and residual  
1246 circulation (white streamlines) from the Modern-Era Retrospective Analysis for Research and  
1247 Applications, version 2 (MERRA-2) for May (left) and August (right) 2022. *Bottom panels:* de-  
1248 seasonalized temperature normalized by the standard deviation (color shading). See text for  
1249 details. The letters show the locations of the time series in Figure 3.

1250  
1251 **Figure 3.** Daily de-seasonalized temperature for 2010 through 2022 from the MERRA-2  
1252 reanalysis (black lines) at 20°S, 20 hPa (top) and 40°S, 27 hPa (bottom), the locations of “A” and  
1253 “B”, respectively, shown in Figures 2 and 4. Two model simulations are included: 1) with an  
1254 interactive QBO and no water vapor anomaly (blue line); and 2) with both the QBO and water  
1255 vapor anomaly (red line). The vertical magenta lines denote the date of the eruption (15  
1256 January 2022). The model is free running and does not correspond to a particular year, except  
1257 following the eruption when the model zonal wind QBO phase is similar to the Singapore  
1258 radiosonde observations and MERRA-2. The right-hand axes show the de-seasonalized  
1259 temperature in terms of the number of standard deviations.

1260

1261 **Figure 4.** May 2022 average de-seasonalized temperature (color shading) and de-seasonalized  
1262 residual circulation (red streamlines). Shown are (a) MERRA-2 reanalysis (repeated from Figure

1263 2a), and (b)-(d) model simulations that include an interactive QBO: (b) with the water vapor  
1264 anomaly; (c) no water vapor anomaly; and (d) the difference, (b) minus (c). The model with  
1265 both the QBO and water vapor anomaly (b) best matches MERRA-2 (a). The letter "A" shows  
1266 the location of the time series in Figure 3a.

1267  
1268 **Figure 5.** Time-latitude cross-sections of the model daily water vapor and temperature  
1269 anomalies for the altitudes indicated. The vertical dotted line in panel (a) denotes the change  
1270 from the MLS v4 H<sub>2</sub>O anomaly to model output on 1 January 2023. Contour intervals are: (a)  
1271 .02, .05, .1, .2, .5, 1, 2, 5, 10, and 15 ppm; (b)-(c):  $\pm 1$ K and includes the  $\pm 0.4$ K contours.

1272  
1273 **Figure 6.** Time-altitude cross-sections of the model daily globally averaged (a) water vapor, (b)  
1274 temperature, and (c) ozone anomalies. The vertical dotted line in the top panel denotes the  
1275 change from the MLS v4 water vapor anomaly to model output on 1 January 2023. Contour  
1276 intervals are: (a) .05, .1, .2, .5, 1, and 2 ppm; (b):  $\pm 0.1$ , -0.5, and -1K; (c):  $\pm 0.5$ , -1, -2, -5, and  
1277 -10%.

1278  
1279 **Figure 7.** Daily total global water vapor burden anomaly (teragrams, Tg, black line), and a fitted  
1280 decay of the simulated burden from 1 July 2023 assuming a constant first order loss with a  
1281 global lifetime of 2.5 years (red dashed-dotted line). The vertical dotted line denotes the  
1282 change from the MLS v4 anomaly to model output on 1 January 2023.

1283  
1284 **Figure 8.** Model anomalies of (a) H<sub>2</sub>O, (b) OH, (c) temperature, (d) NO<sub>x</sub>, and (e) ozone, averaged  
1285 over a two-year (2023-2024) period. Contour intervals are: (a) +5%; (b): +2%; (c): -1, -0.5, 0.05,  
1286 0.1K; (d) and (e): -5, -2, +1%. The thick black solid line indicates the tropopause.

1287  
1288 **Figure 9.** Model global mean vertical profiles of anomalies averaged over a two-year (2023-  
1289 2024) period. Shown are: (a) concentrations of the constituents indicated; (b) tendency of odd  
1290 oxygen (O+O<sub>3</sub>) due to the chemical loss cycles of the odd hydrogen (HO<sub>x</sub>), odd nitrogen (NO<sub>x</sub>),  
1291 odd oxygen (O<sub>x</sub>), and chlorine (ClO<sub>x</sub>) families, and the total chemical tendency.

1292

1293 **Figure 10.** Time-altitude cross-sections of daily model (a) water vapor, (b) temperature, and (c)  
1294 ozone anomalies averaged over the Antarctic polar cap (65°S-90°S), and (d) daily MLS v5 ozone  
1295 through 30 November 2023 averaged over 65°S-80°S with the 2005-2021 average seasonal  
1296 cycle removed. The vertical dotted line in panel (a) denotes the change from the MLS v4 water  
1297 vapor anomaly to model output on 1 January 2023. Contours are: (a) 0.1, 0.2, 0.5, 1, and 2 ppm;  
1298 (b): intervals of  $\pm 0.5$ K and includes the  $\pm 0.1$  and  $\pm 0.2$ K contours; (c) and (d): intervals of  $\pm 1$   
1299 DU/km and includes the  $\pm 0.2$  and  $\pm 0.5$  DU/km contours.

1300  
1301 **Figure 11.** April-December 2023 daily model anomalies averaged over the Antarctic polar cap  
1302 (65°S-90°S) at 54 hPa (20 km) for (a) water vapor and ice PSCs, (b) nitrogen species (ppb, solid  
1303 lines) and the combined rate of the heterogeneous reaction  $\text{ClONO}_2 + \text{HCl} \rightarrow \text{HNO}_3 + \text{Cl}_2$  on NAT  
1304 and Ice PSCs and sulfate aerosols (1/day, dotted line), (c) odd oxygen loss (mainly ozone), (d)  
1305 ozone, odd oxygen loss and transport, and (e) temperature and solar ozone heating. The right-  
1306 hand axis indicates the odd oxygen tendencies in panel (d), and the solar ozone heating rate in  
1307 panel (e).

1308  
1309  
1310 **Figure 12.** Model anomalous odd oxygen ( $\text{O} + \text{O}_3$ ) tendencies over the Antarctic polar cap (65°S-  
1311 90°S) at 28 hPa (25 km) for 2023. Shown are the daily chemical tendencies due to the odd  
1312 hydrogen ( $\text{HOx}$ , green), odd nitrogen ( $\text{NOx}$ , blue), chlorine ( $\text{ClOx}$ , red), and bromine ( $\text{BrOx}$ ,  
1313 orange) families. Also shown is the tendency due to the transport of odd oxygen (black dash dot  
1314 line).

1315  
1316 **Figure 13.** Daily model total ozone anomalies over the 2022-2031 period. Contour intervals are  
1317 -5 DU and include the  $\pm 1$  DU contours to show minor effects.

1318  
1319 **Figure 14.** Daily total ozone seasonal cycle for (a) 2023 Antarctic (63°S-90°S) and (b) July 2023 -  
1320 June 2024 for the Arctic (63°N-90°N). Shown are model simulations under different planetary  
1321 wave forcing conditions (colored lines) for the baseline (solid) and including the HT water vapor  
1322 anomaly (dashed). The gray shades depict satellite climatology (1991-2022, Newman and Lait,  
1323 2023).

1324  
1325 **Figure A1.** Time series of equatorial stratospheric zonal mean zonal wind from the (a) MERRA-2  
1326 reanalysis for 1993-2005, and (b) model simulation. Contour intervals are  $\pm 10$  m/sec, with the  
1327 westerlies in red and the easterlies in blue.

1328  
1329 **Figure A2.** Age of air at: (a) 20 km ( $\sim 50$  hPa) derived from ER-2 aircraft observations of SF<sub>6</sub> (blue  
1330 asterisks) and CO<sub>2</sub> (red triangles), and (b)-(d) vertical profiles of the age of air derived from  
1331 balloon measurements of SF<sub>6</sub> (blue asterisks, green plus signs) and CO<sub>2</sub> (red triangles) at the  
1332 latitudes indicated (adapted from Hall et al., 1999). The observations are taken during the  
1333 1990s. Also shown is the model age of air annually averaged over the 1990s (black line). The  
1334 age is taken relative to the tropical tropopause.

1335  
1336 **Figure A3.** Monthly mean water vapor (ppm) for June and October 2022 from MLS v5 (top  
1337 panels) and model simulation (bottom panels). The contour interval is 1 ppm.

1338  
1339 **Figure A4.** Time-latitude cross-sections of daily water vapor (ppm) at 28 hPa for January 2022-  
1340 November 2023 from (a) MLS v5 and (b) model. The model is forced with the MLS-derived H<sub>2</sub>O  
1341 anomaly prior to 1 March 2022 (vertical dotted line in panel (b)). The contour interval is 1 ppm.

1342  
1343 **Figure A5.** Time-altitude cross-sections of daily water vapor (ppm) averaged over 65°S-80°S for  
1344 January 2022-November 2023 from (a) MLS v5 and (b) model. The contour interval is 1 ppm.

1345  
1346 **Figure A6.** Time-altitude cross-sections of daily water vapor (ppm) averaged over 10°S-10°S for  
1347 January 2022-November 2023 from: (a) MLS v5, and model simulations with the zonal wind  
1348 QBO in the same phase (b) and roughly 180° out of phase (c) with the Singapore radiosonde  
1349 observations. Panel (d) shows the model difference, (c) minus (b). Maximum easterly (“E”) and  
1350 westerly (“W”) zonal winds at 30 hPa are from the Singapore observations in panel (a) and the  
1351 respective model simulations in panels (b) and (c). The model is forced with the MLS-derived  
1352 H<sub>2</sub>O anomaly prior to 1 March 2022 (vertical dotted line in panels (b)-(d)). The contour interval  
1353 is 1 ppm.

1354

1355 **Figure A7.** Time-altitude cross-sections of daily de-seasonalized ozone (ppm) averaged over  
1356 15°S-40°S. Shown are (a) MLS v5 data, and (b)-(d) model simulations that include an interactive  
1357 QBO: (b) with the water vapor anomaly; (c) no water vapor anomaly; and (d) the difference, (b)  
1358 minus (c). The contour interval is  $\pm 0.1$  ppm. The vertical magenta lines denote the date of the  
1359 eruption (15 January 2022).