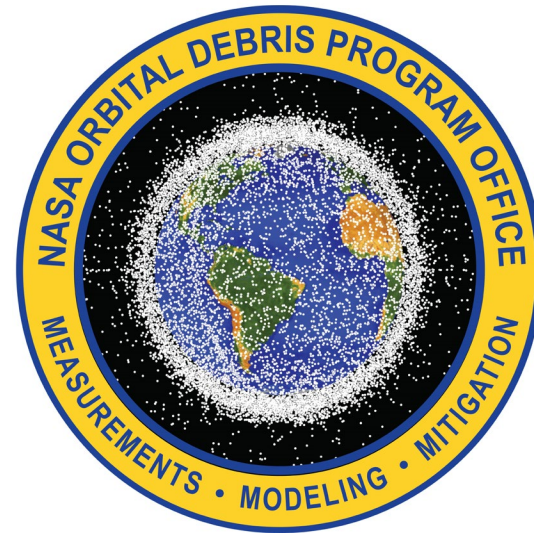




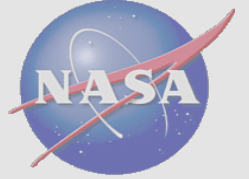
NASA Orbital Debris Radar Measurements by the Haystack Ultra-wideband Satellite Imaging Radar (HUSIR)



Mark Matney¹, Jessica Arnold Headstream², Alyssa Manis¹

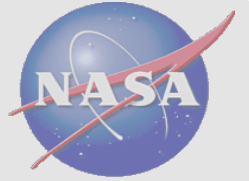
¹NASA Orbital Debris Program Office, ²Jacobs

Space Control Conference
08 May 2024



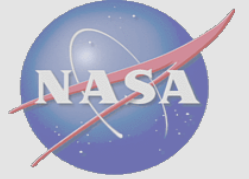
The Debris Environment

- **The DOD is tasked with tracking and cataloging satellites in Earth orbit using the Space Surveillance Network (SSN)**
 - Intact spacecraft, rocket bodies, satellite breakup debris
- **Tracking = determining an orbit of sufficient accuracy to recover the object later and to positively identify it as the same object**
 - Tracking also allows position predictions for collision avoidance
- **Due to limitations of geographic coverage and sensor sensitivity, the tracked catalog is only complete down to some limiting size**
 - Historically, this limit is >10 cm, but the Space Fence has reduced that limit somewhat



NASA's Approach to the Debris Environment

- **Just because we cannot easily track objects smaller than 10 cm, doesn't mean that we cannot extract useful information on their properties**
 - Number
 - Altitude distribution
 - Size distribution
 - Sources
 - Inclination distribution
- **NASA uses statistical sampling to estimate these numbers**
 - This consists of staring at a region of sky and counting objects that pass through
 - Using knowledge of how orbiting objects behave, a surprising amount of information can be extracted
- **The statistical data is used to build NASA's Orbital Debris Engineering Model (ORDEM)**
 - ORDEM provides estimates of the orbital debris flux environment for spacecraft operators and designers to plan missions and spacecraft shielding



Radar and Optical Measurements in ODPO

Goal: To statistically sample the orbital debris (OD) environment below the threshold of the SSN to build and validate orbital debris models



Credit: Reprinted with permission
Courtesy of MIT Lincoln Laboratory
(MIT/LL), Lexington, Massachusetts



Credit: Courtesy
NASA/JPL-Caltech,
<https://deepspace.jpl.nasa.gov/galleries/goldstone/#gallery>



Credit: Ben Hanna

Haystack Ultrawideband
Satellite Imaging
Radar (HUSIR)

Goldstone Orbital
Debris Radar

Eugene Stansbery
Meter-Class Autonomous
Telescope (ES-MCAT)

Space Surveillance Network
(SSN), Space Fence

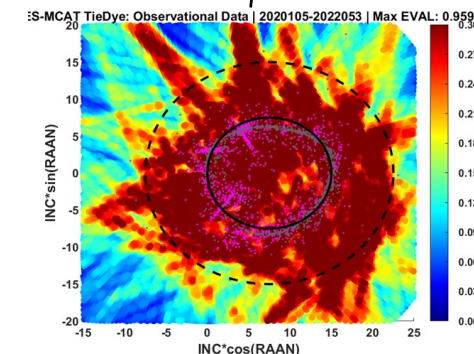
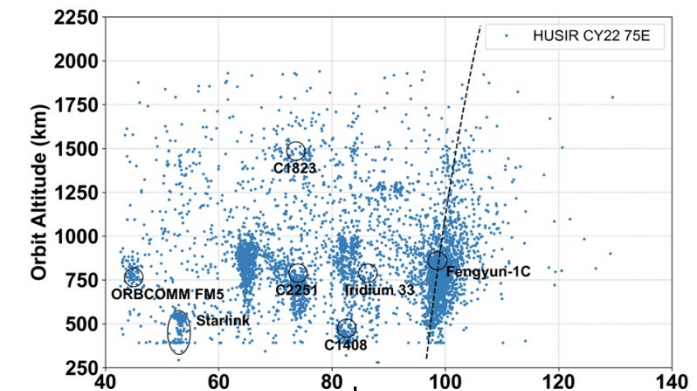
Space Surveillance
Telescope (SST)

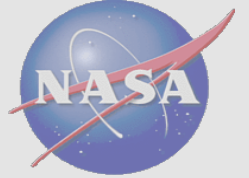
Provide data for
building/validating OD models

Measurements

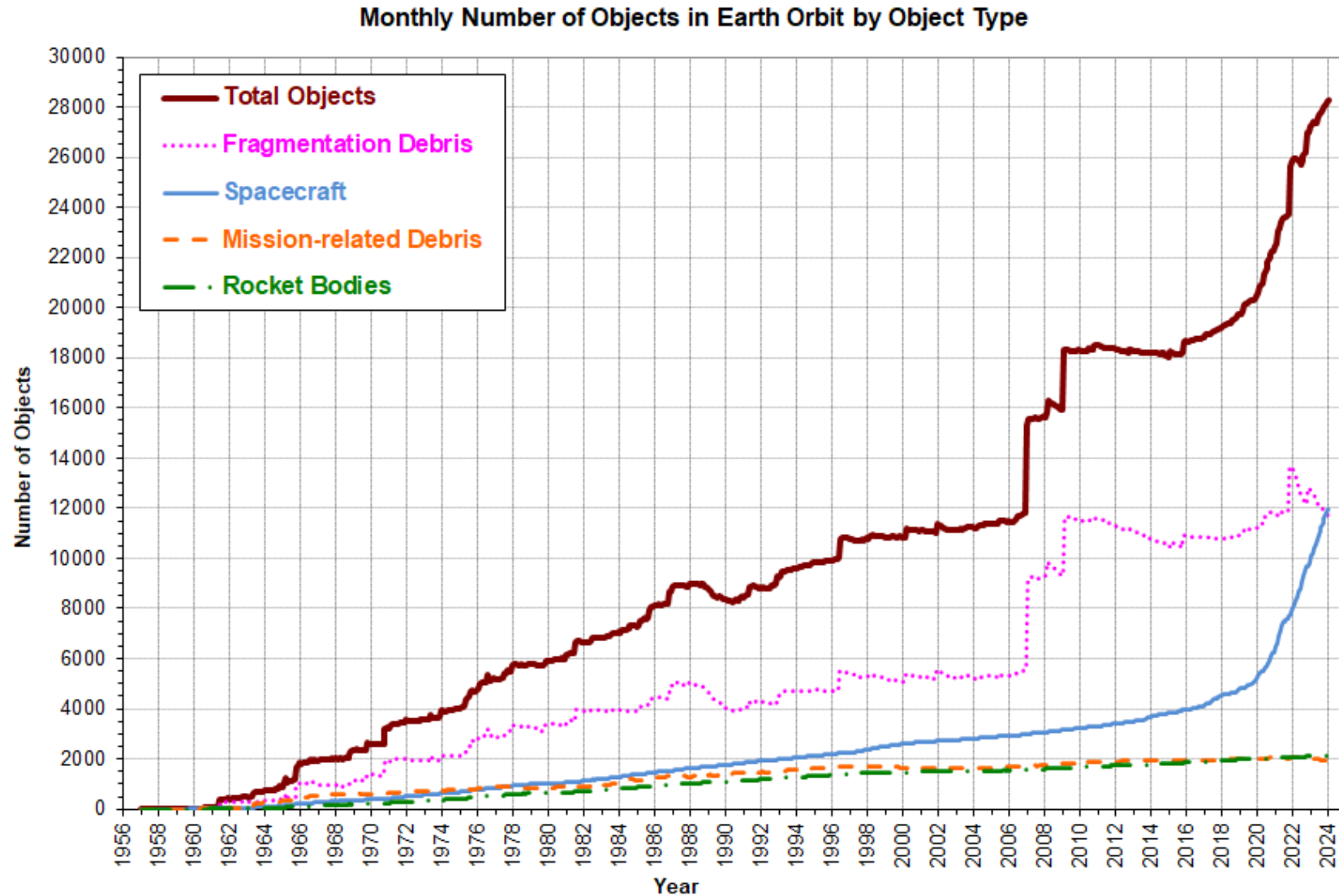
Modeling

Provide feedback on how to
improve measurements

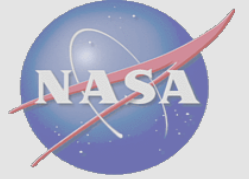




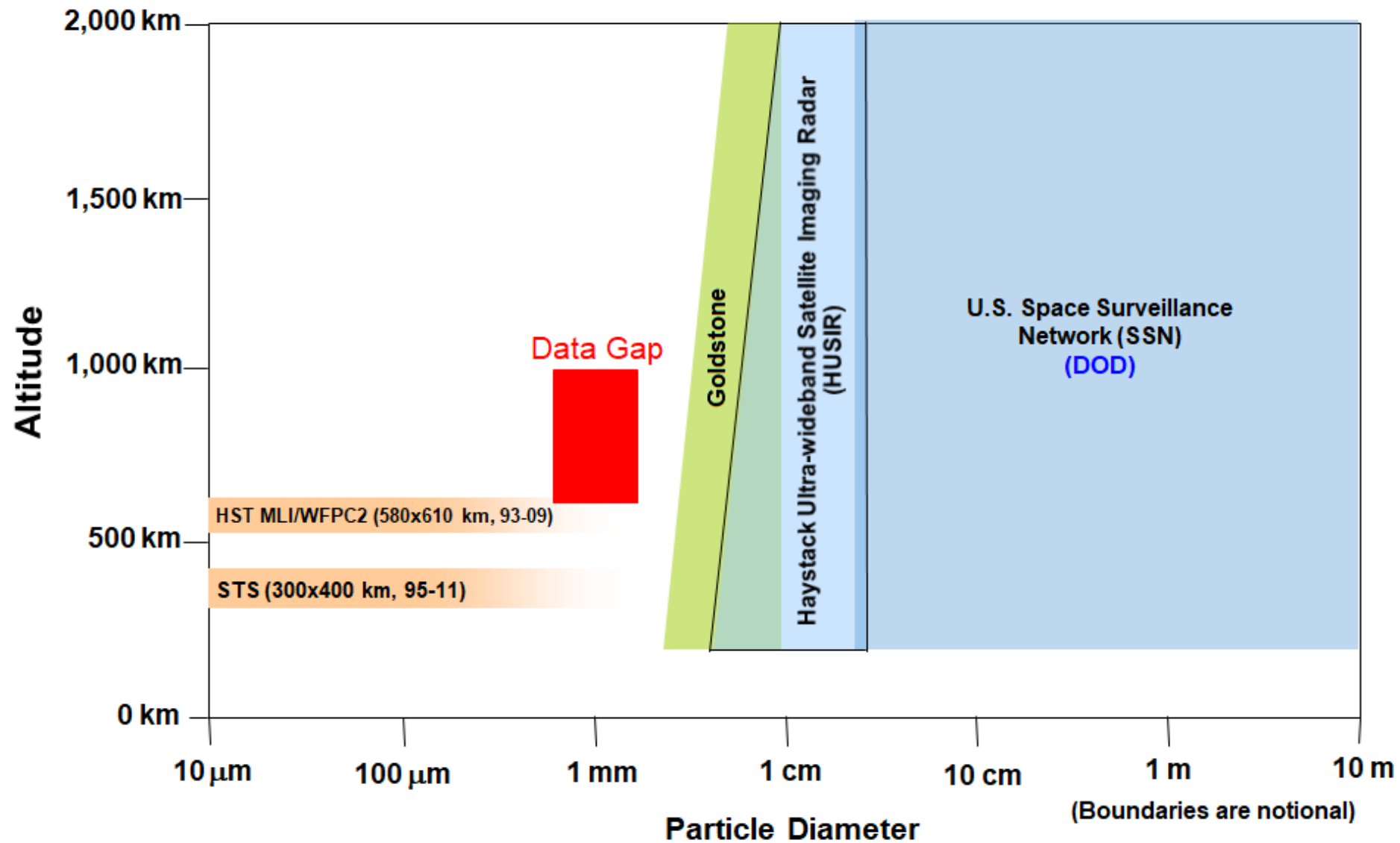
Growth of Cataloged Space Object Population

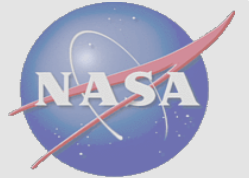


**Environment is
dynamic, so constant
monitoring is necessary**



Data Coverage



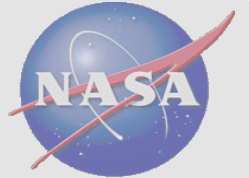


HUSIR Radar

- **Located in Massachusetts – 42.6° latitude, operated by Massachusetts Institute of Technology, Lincoln Laboratory (MIT/LL)**
- **Haystack Ultrawide Satellite Imaging Radar (HUSIR – previously known as Haystack)**
 - 36 m diameter dish
 - 3 cm wavelength (X-band – 10/10.1 GHz)
 - Can detect debris >5 mm in LEO
 - Narrow beam – 0.058°
- **The radar accurately measures RCS, polarization, range, and Doppler velocity along line of sight, but has trouble measuring other velocity components, so precision orbit not possible**



Credit: Reprinted with permission Courtesy of MIT Lincoln Laboratory, Lexington, Massachusetts



Estimating Debris Parameters

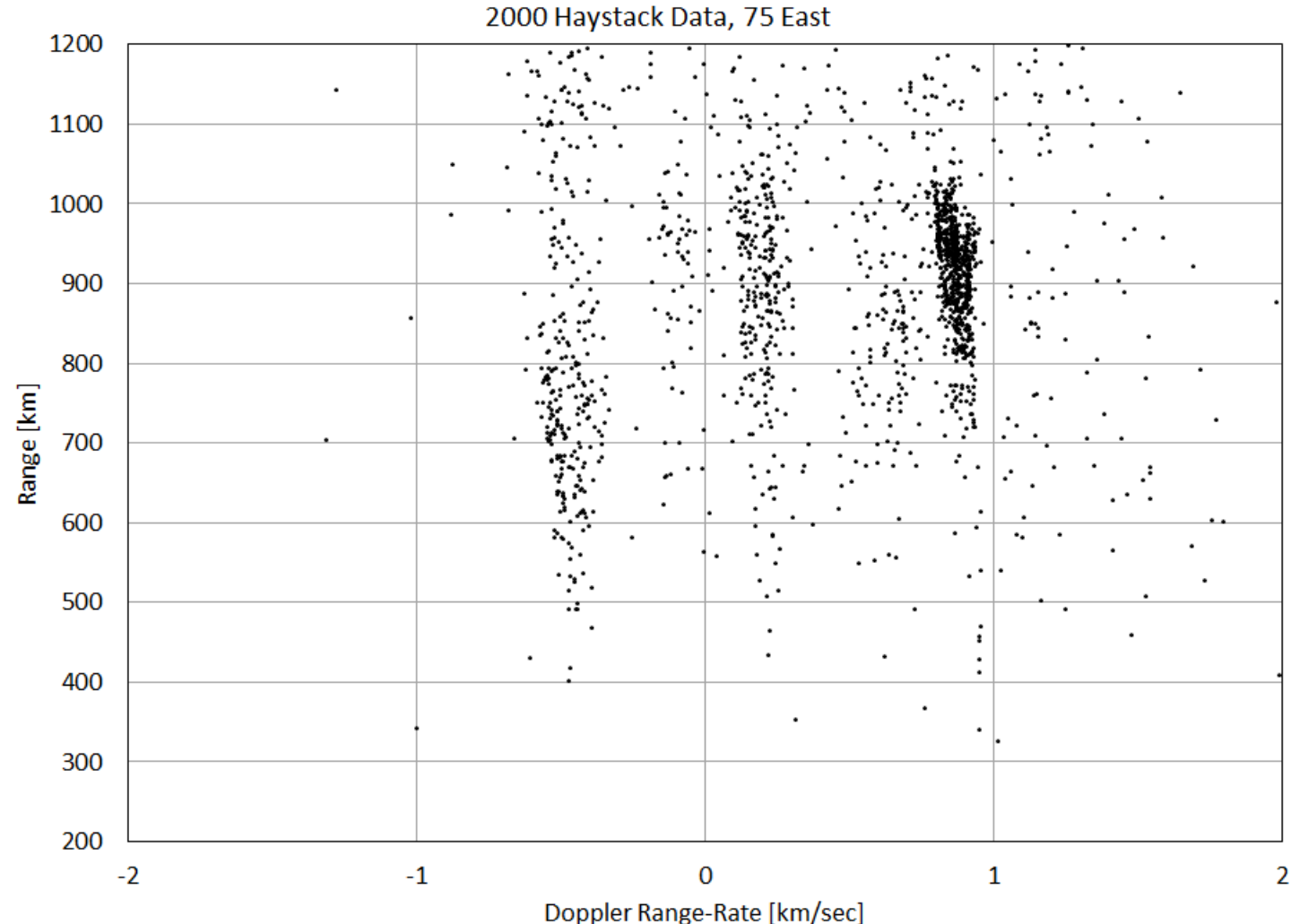
- We receive surrogate measurements for the OD parameters we really want for modeling

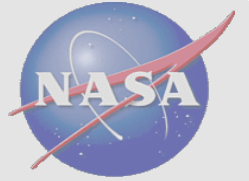
Range / Range-Rate (Doppler Velocity)



Altitude / Inclination

- Doppler inclination calculation assumes circular orbits
- Requires a special pointing mode – 75° elevation, 90° azimuth (East)





Estimating Debris Parameters

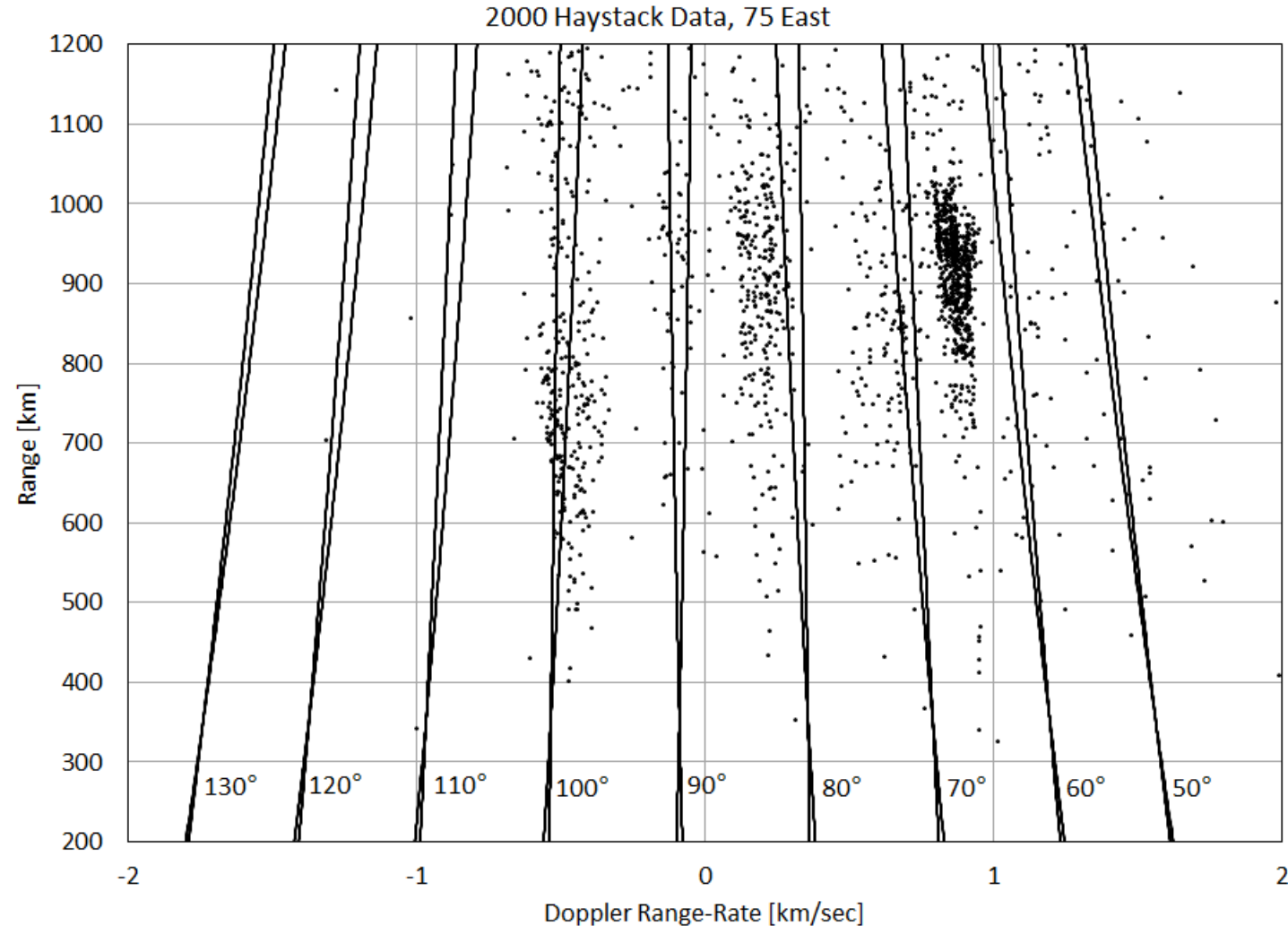
- We receive surrogate measurements for the OD parameters we really want for modeling

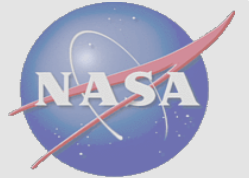
Range / Range-Rate (Doppler Velocity)



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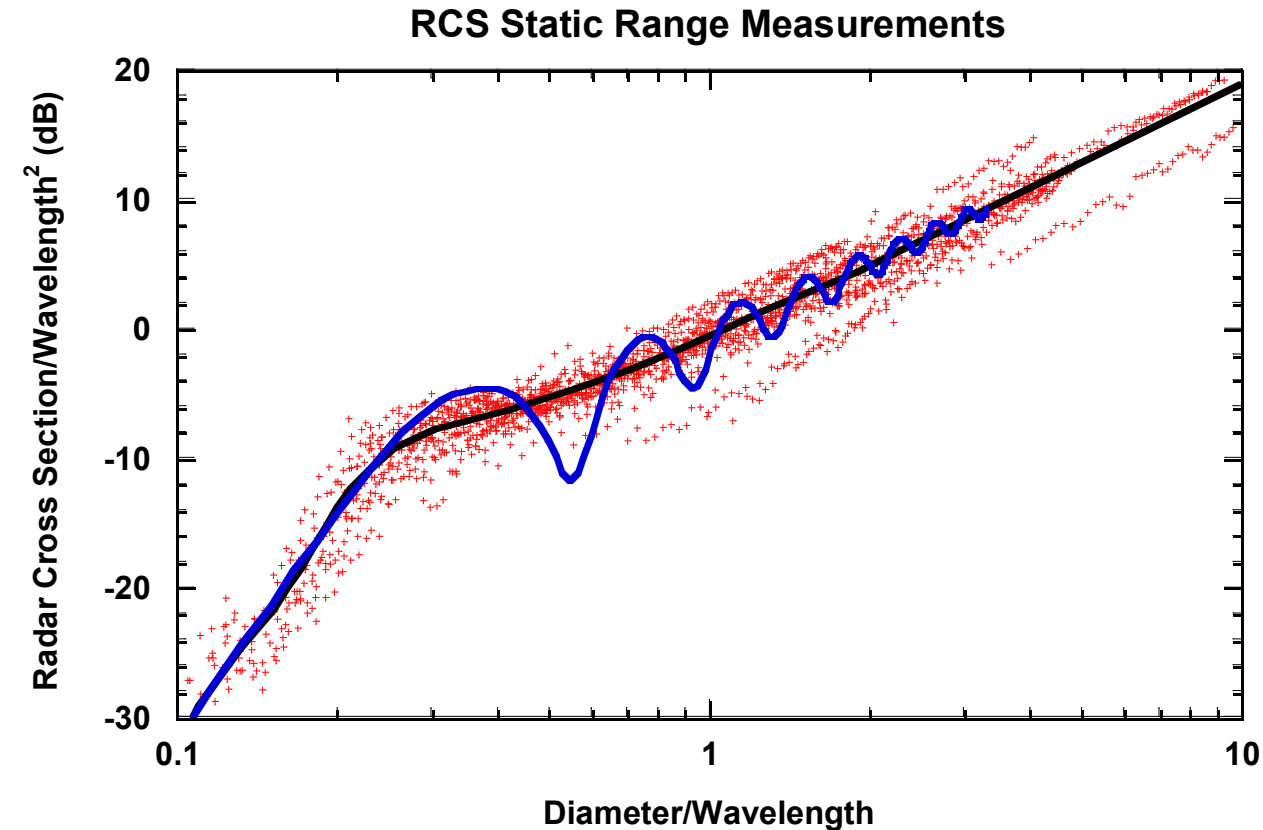


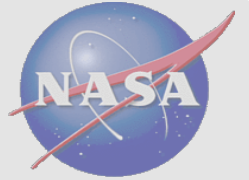
Estimating Debris Parameters

- We receive surrogate measurements for the OD parameters we really want for modeling

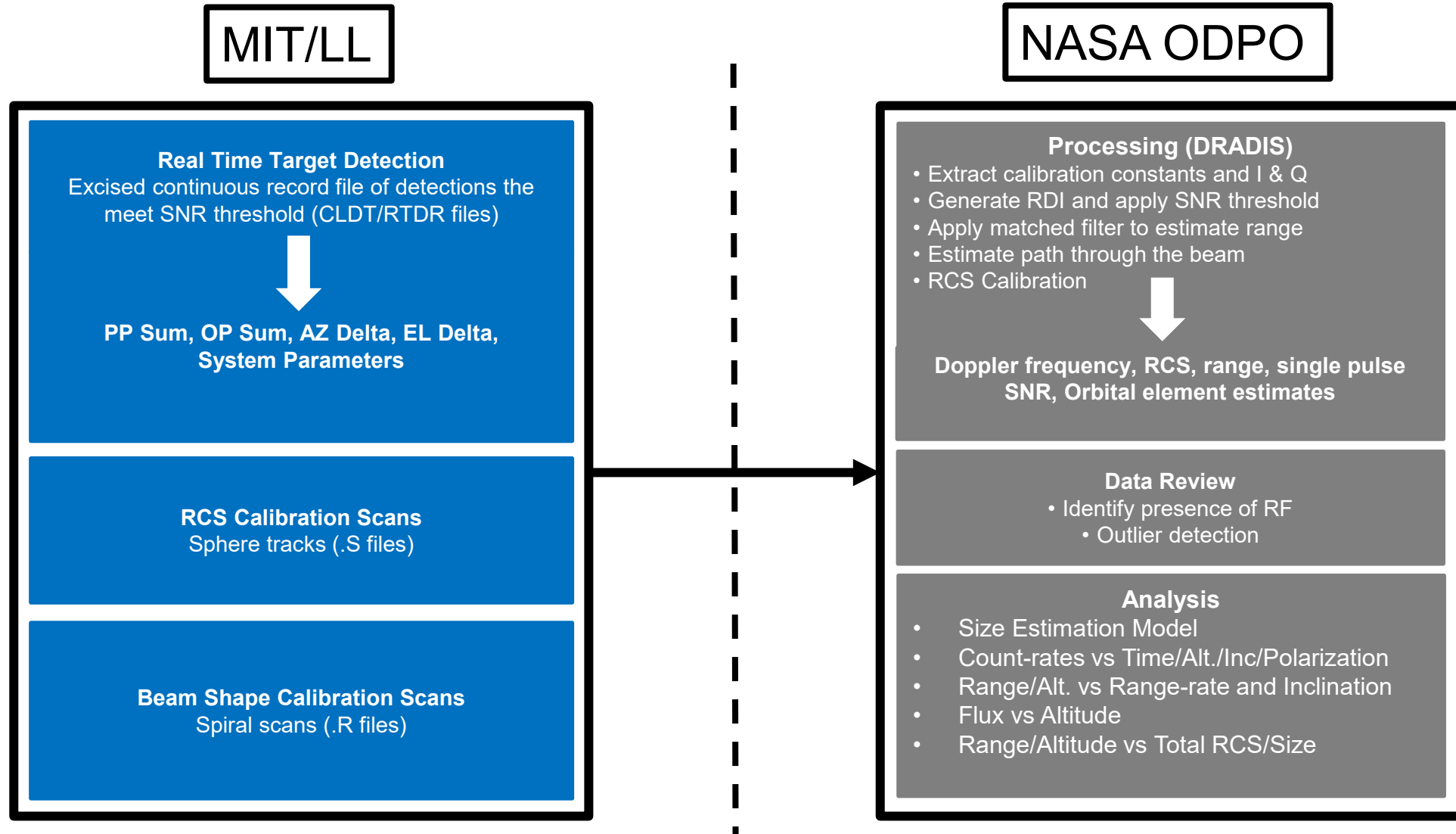
RCS → size

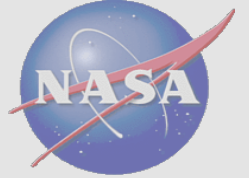
- NASA size estimation model
 - Empirical model based on measurements of 39 representative debris objects averaged over many orientations at specific frequencies



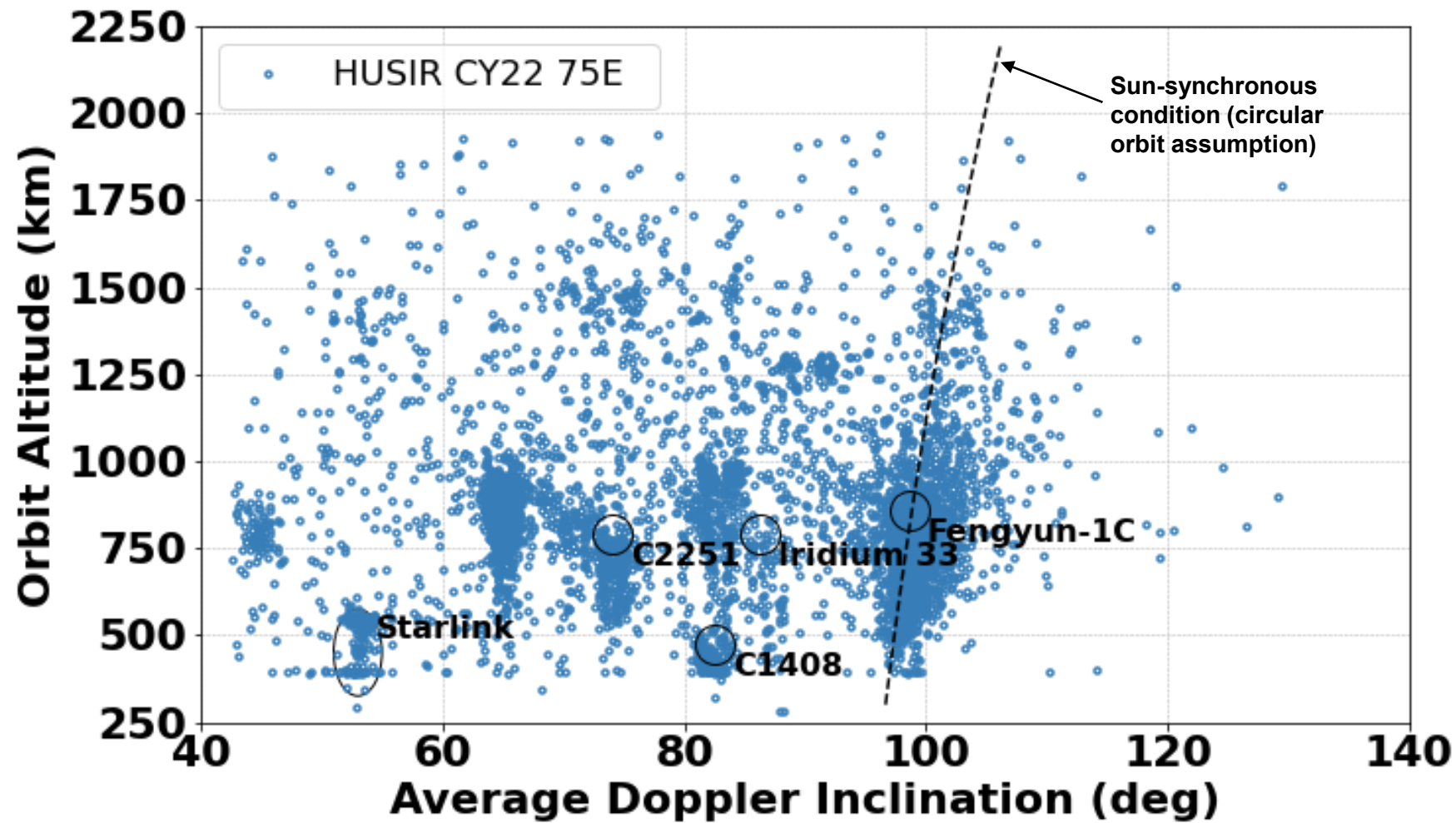


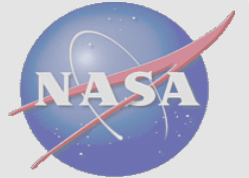
HUSIR Data Collection and Processing



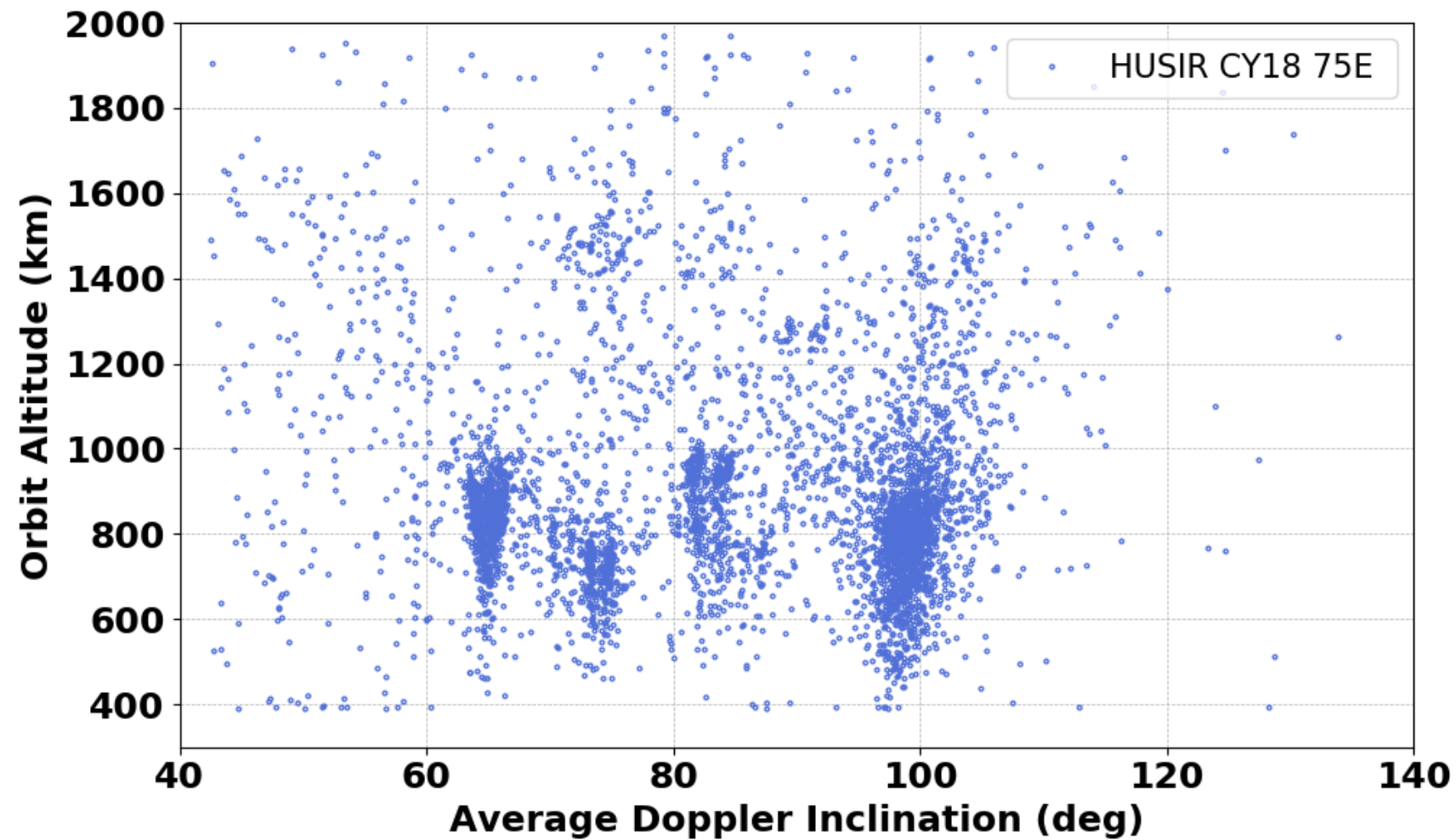


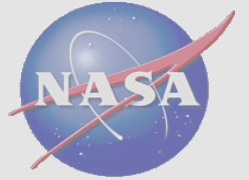
Altitude vs. Inclination, HUSIR 75E



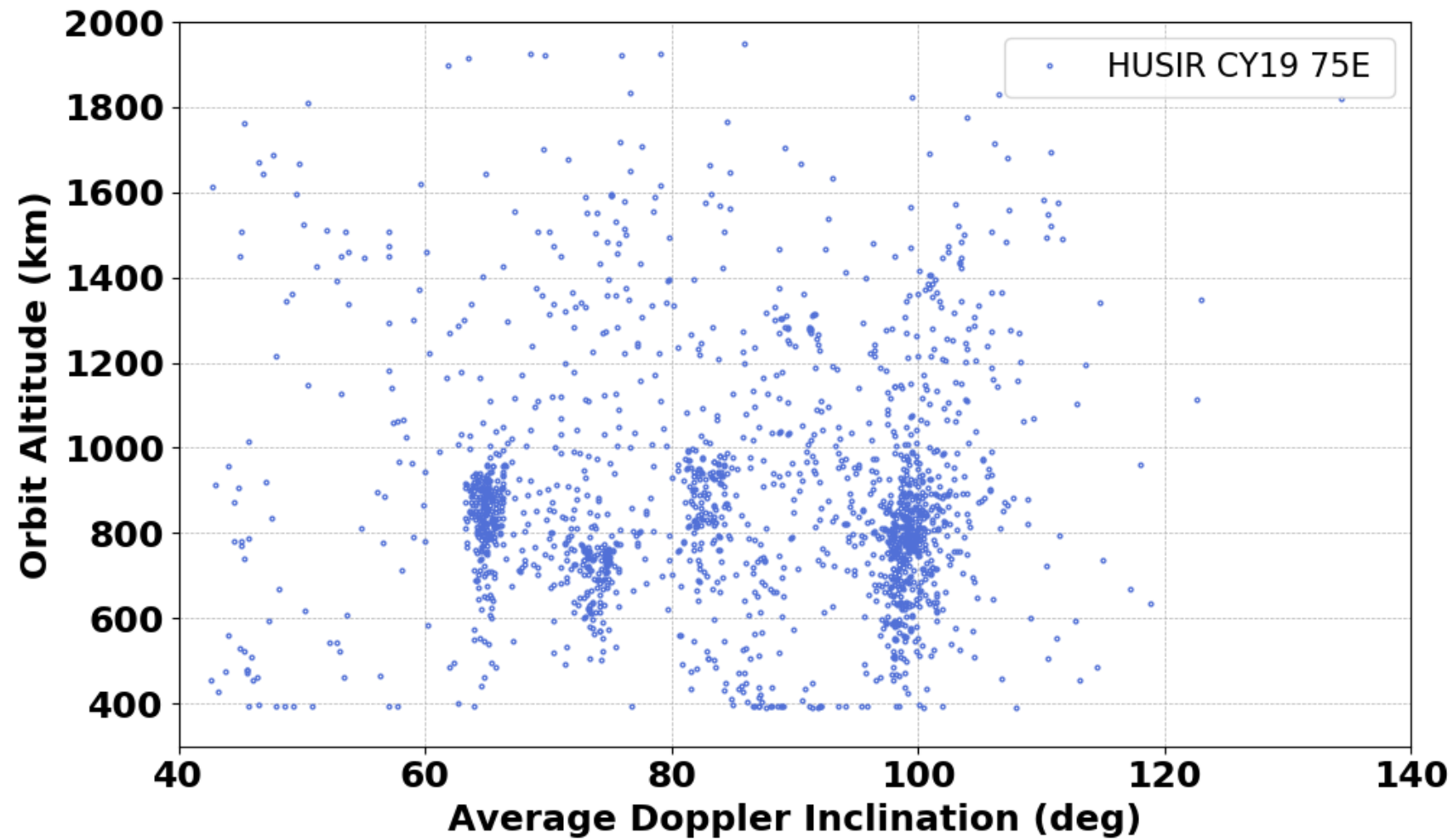


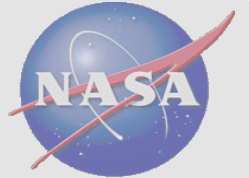
Altitude vs. Inclination, 2018



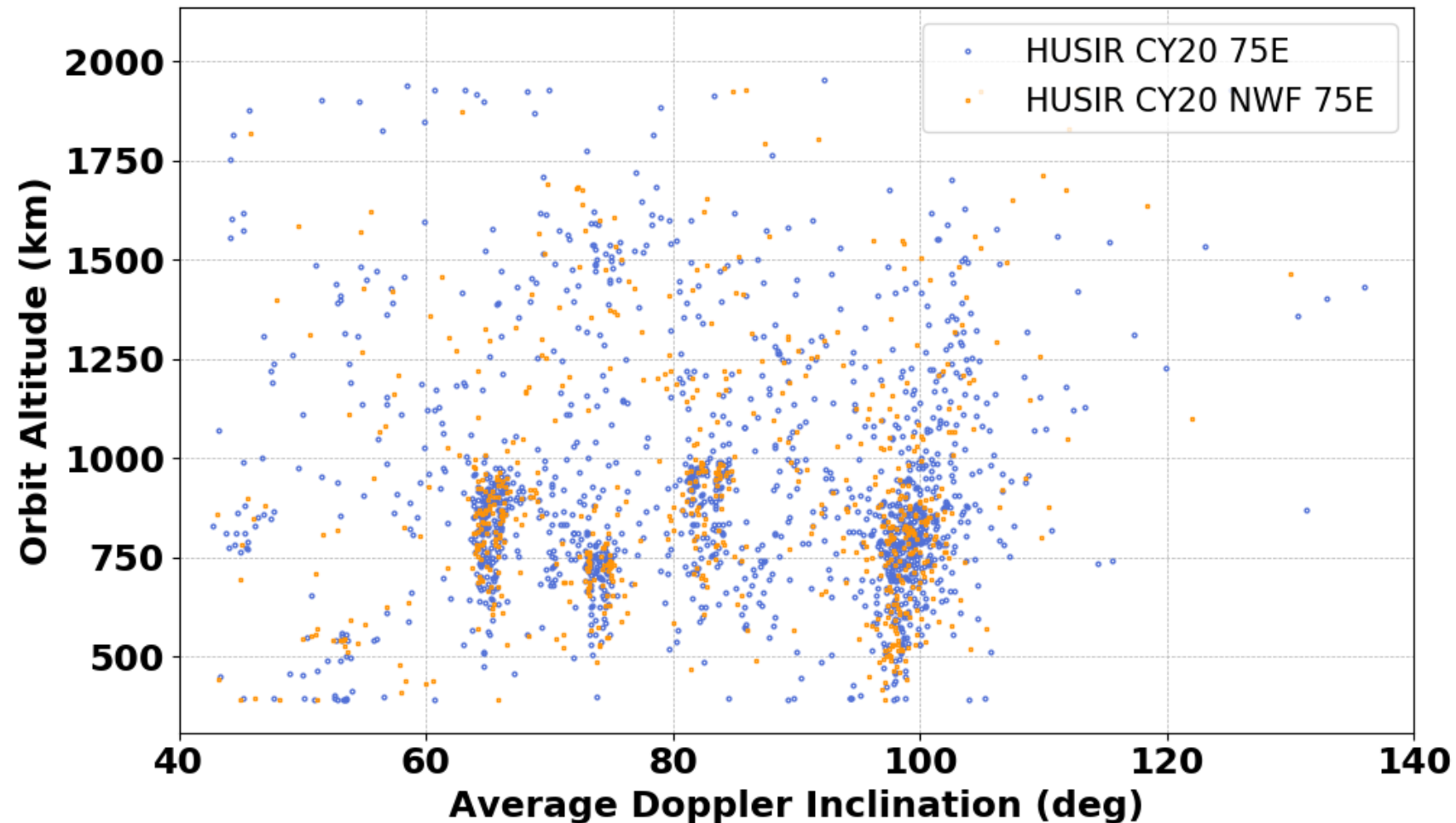


Altitude vs. Inclination, 2019

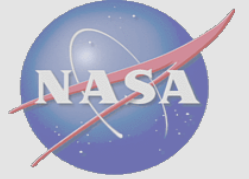




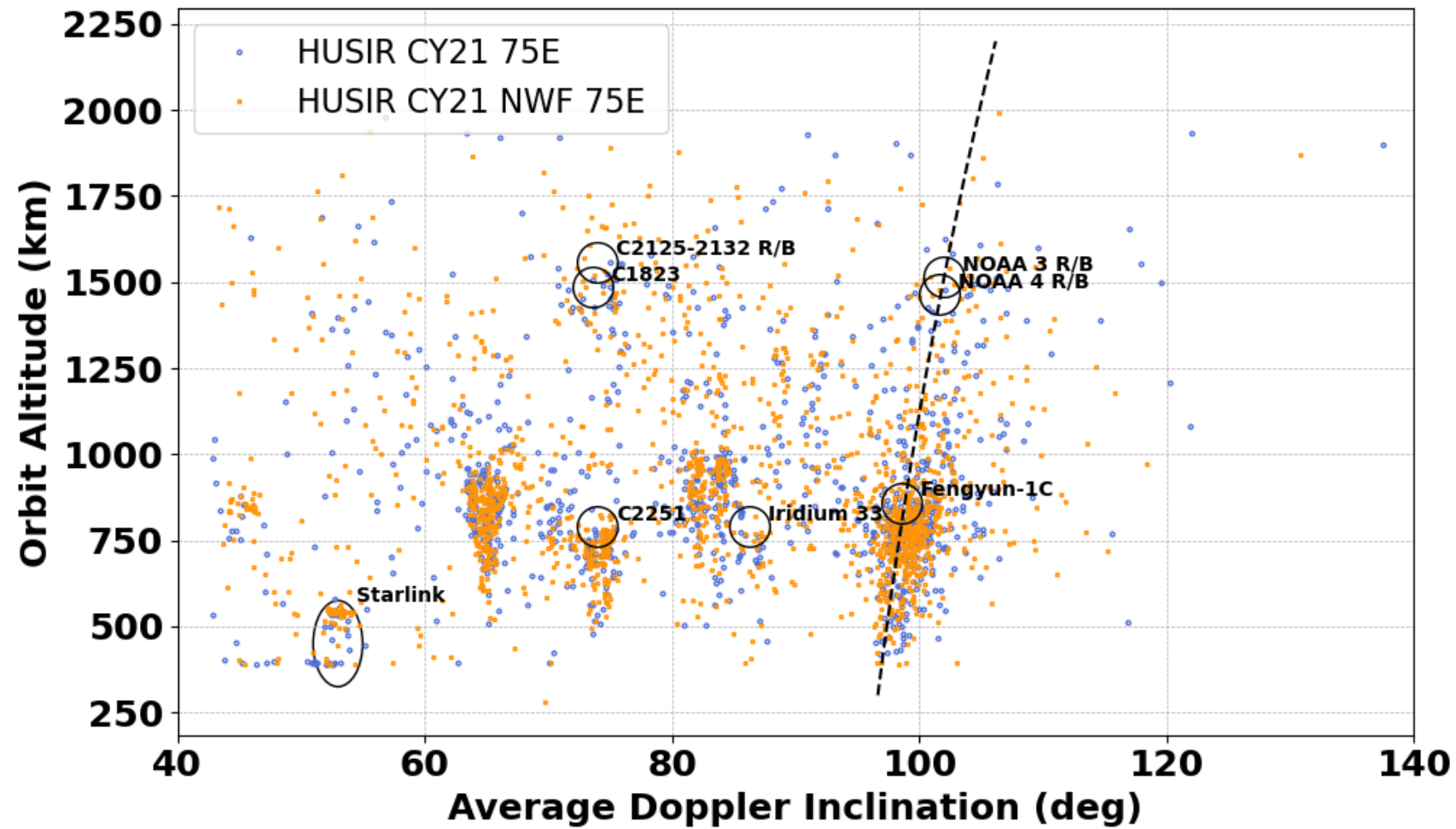
Altitude vs. Inclination, 2020

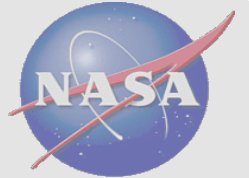


In 2020, we began experimenting with using 10.1 GHz instead of 10 GHz to minimize RFI

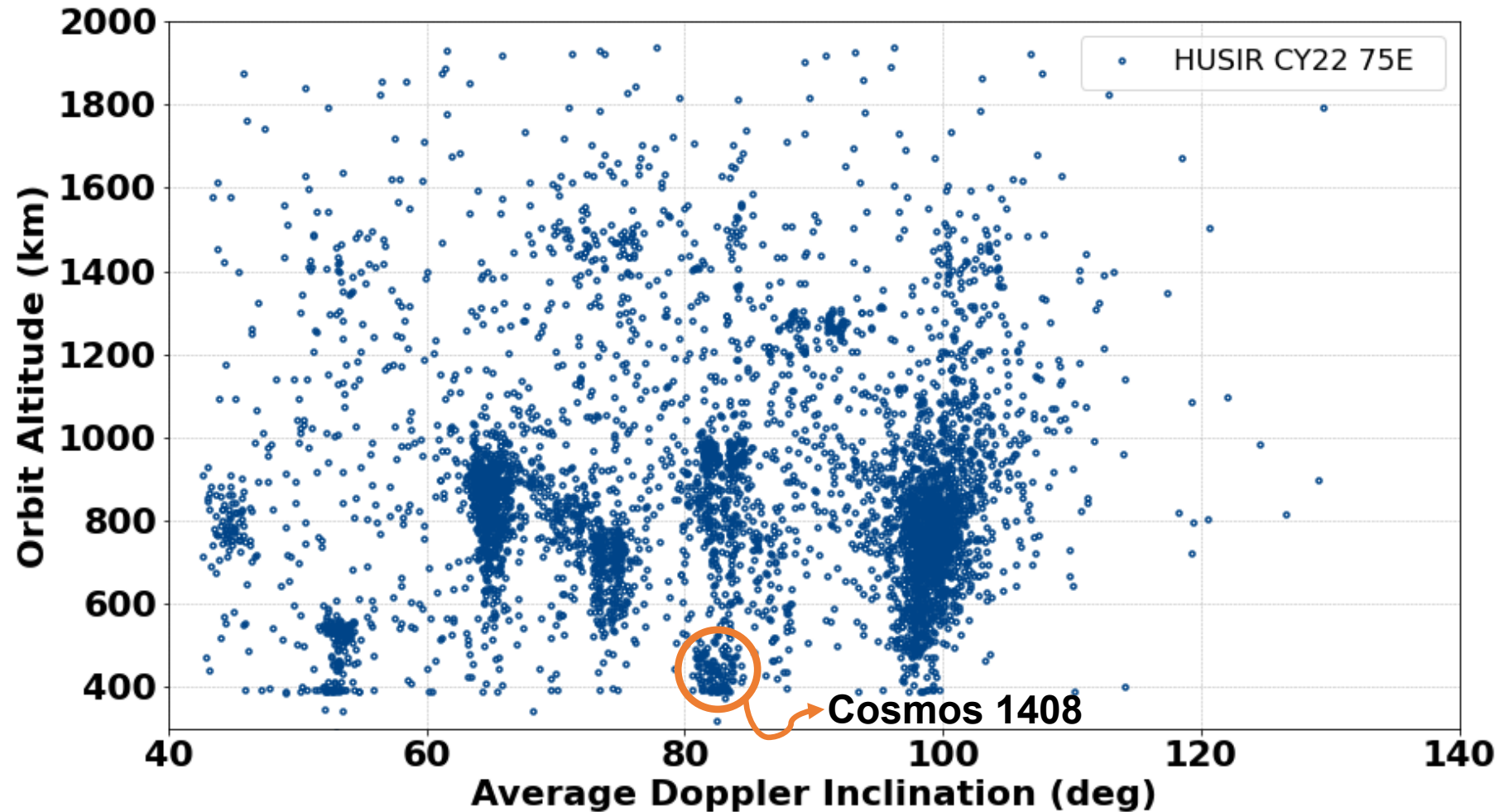


Altitude vs. Inclination, 2021





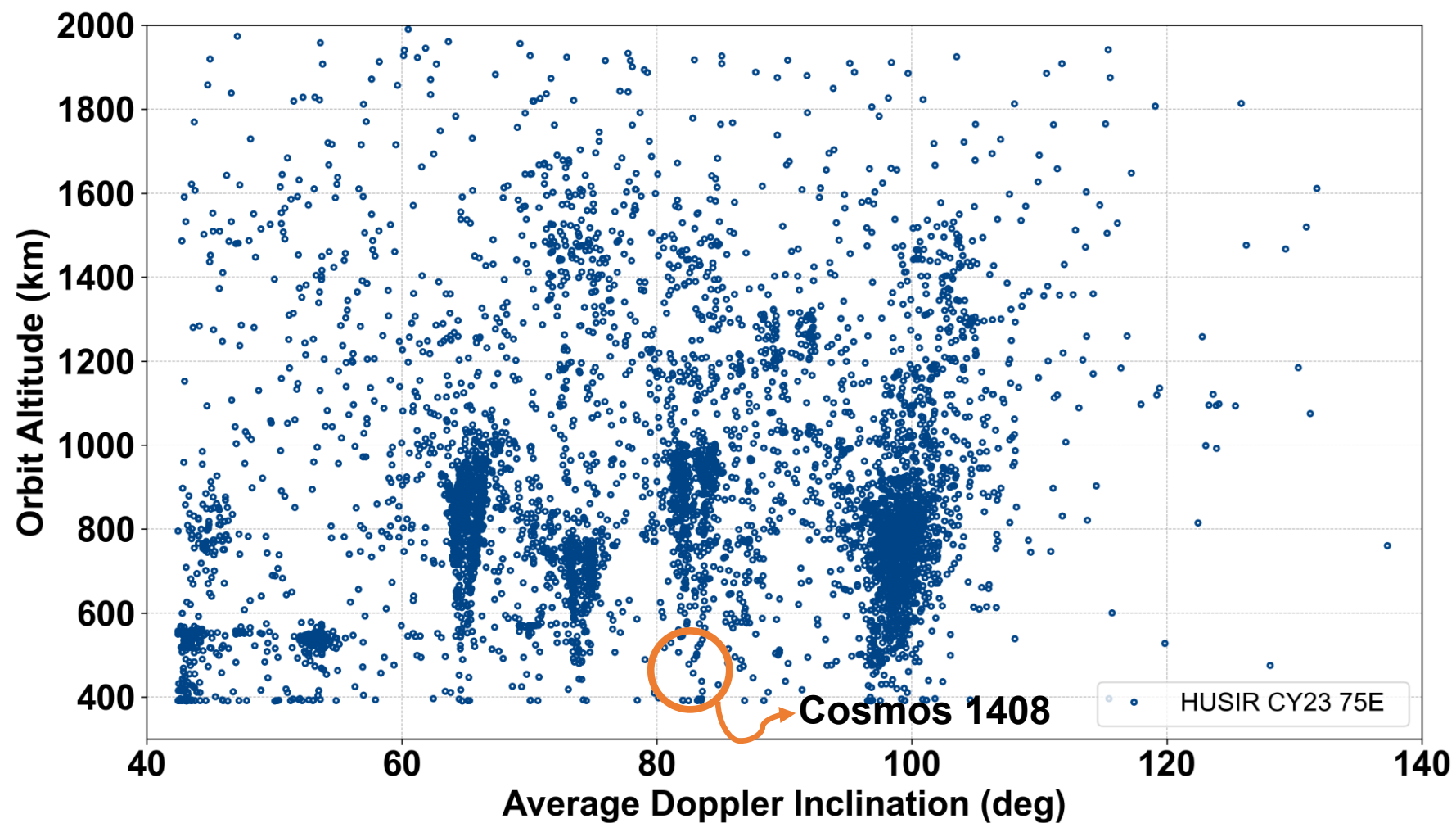
Altitude vs. Inclination, 2022

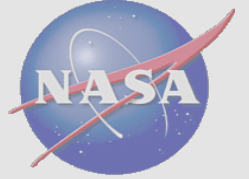


Cosmos 1408 was the target of the Russian ASAT test in November 2021

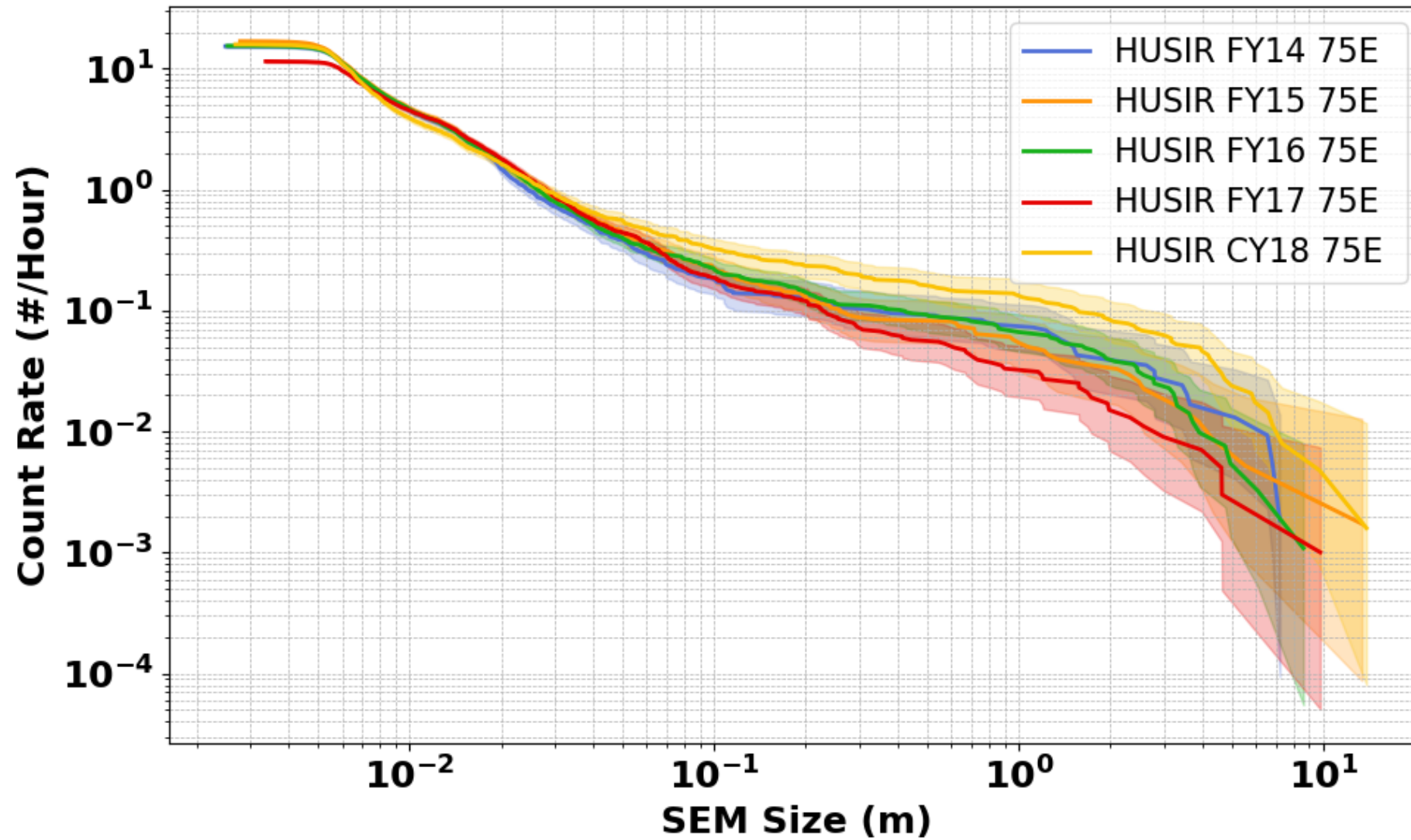


Altitude vs. Inclination, 2023



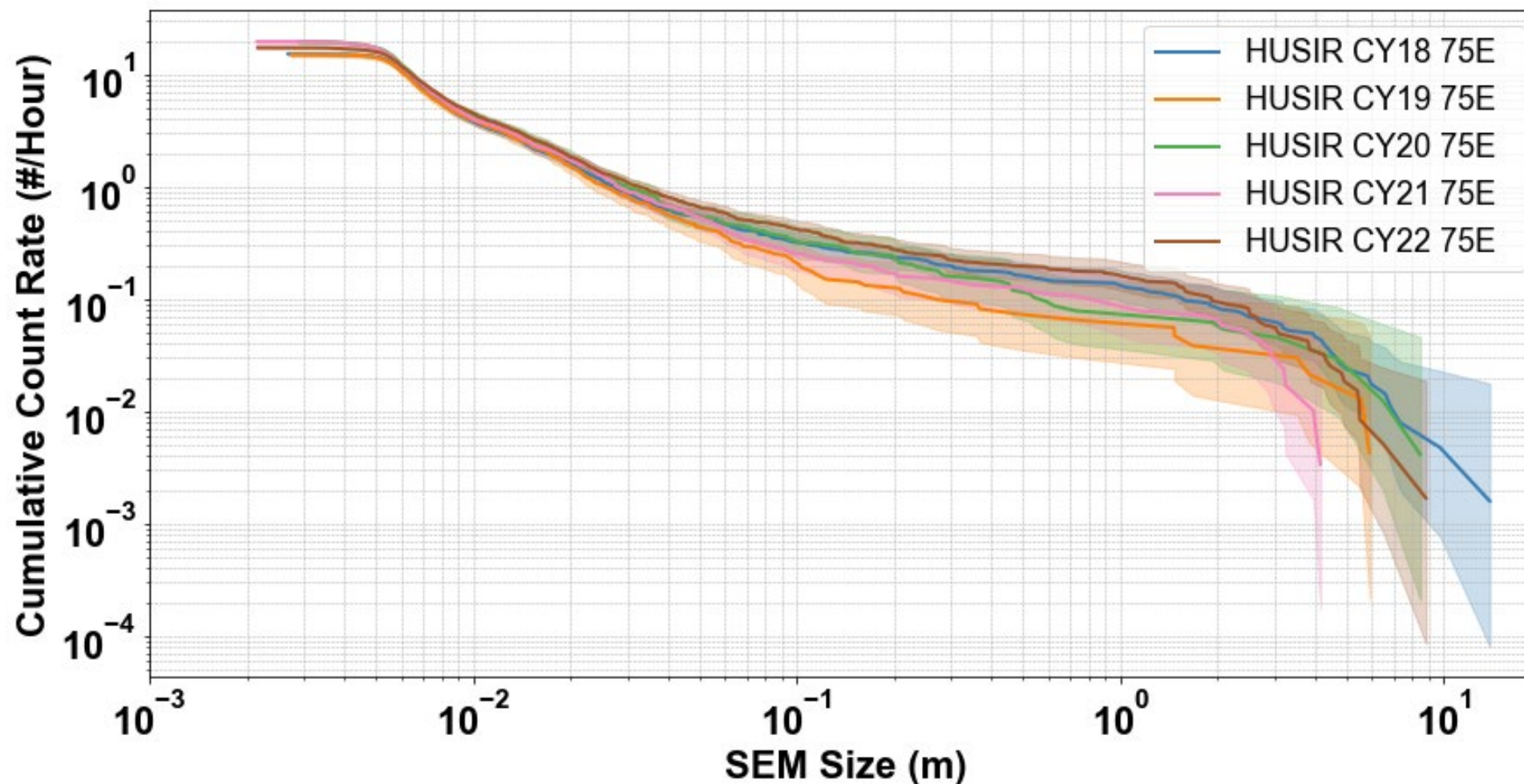


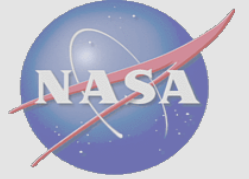
Count Rates over Time, HUSIR 75E



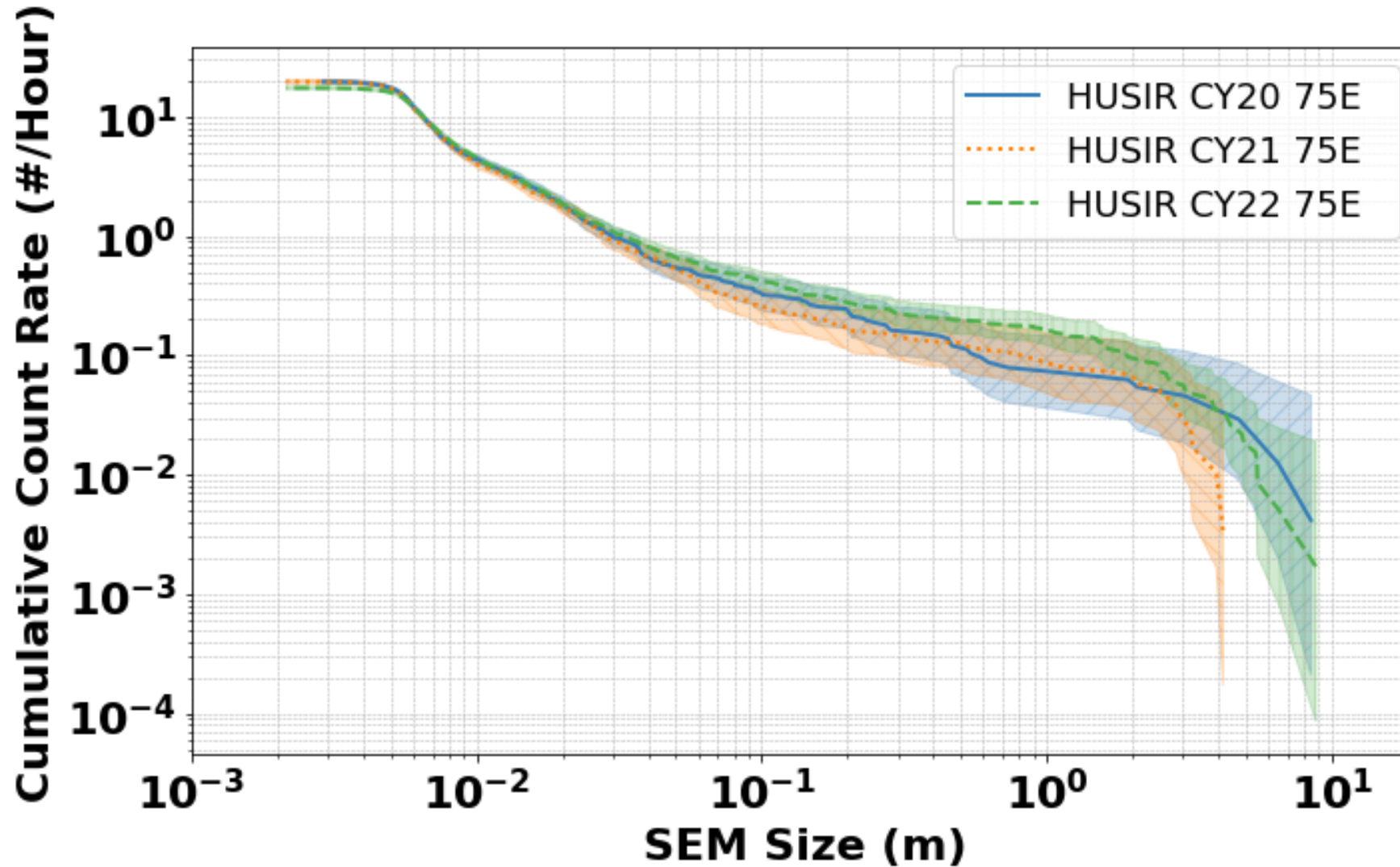


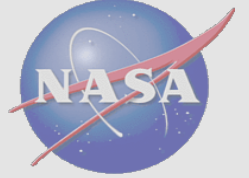
Count Rates over Time, HUSIR 75E (continued)



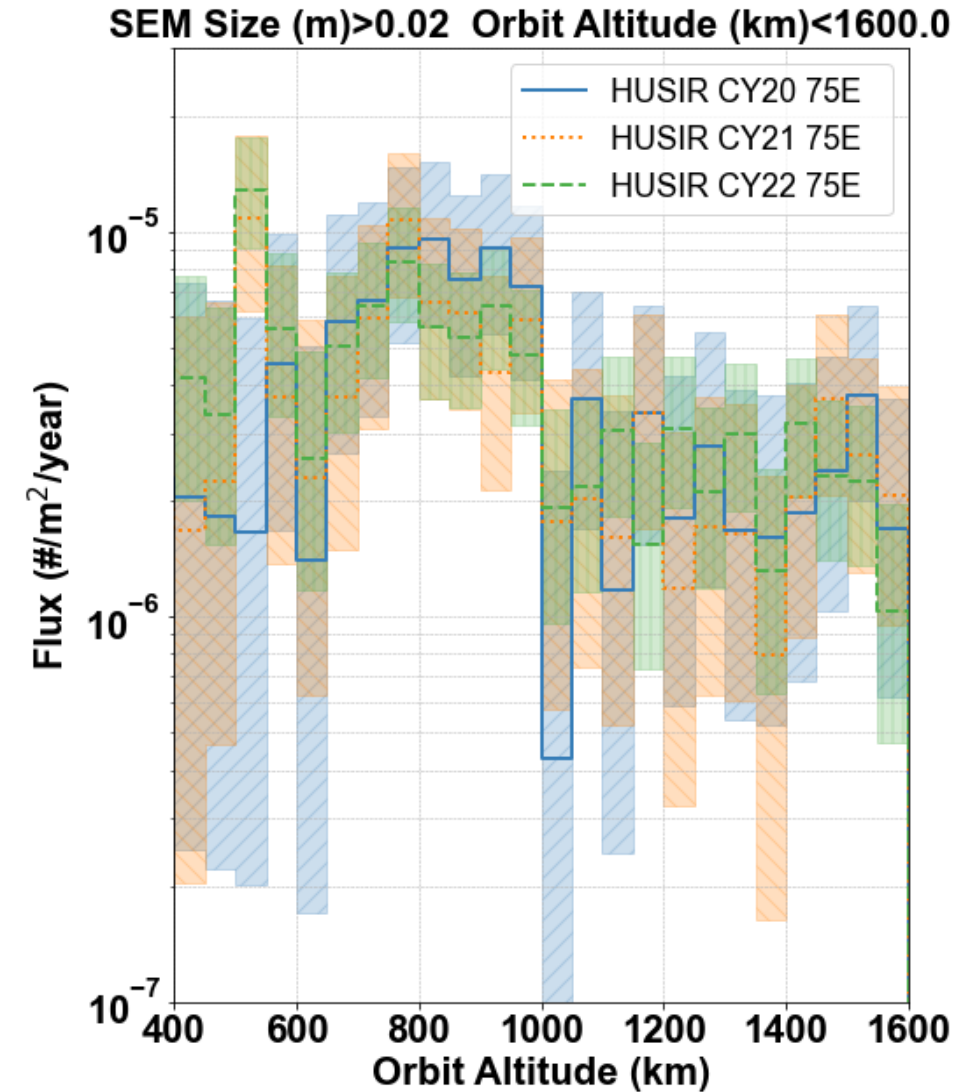
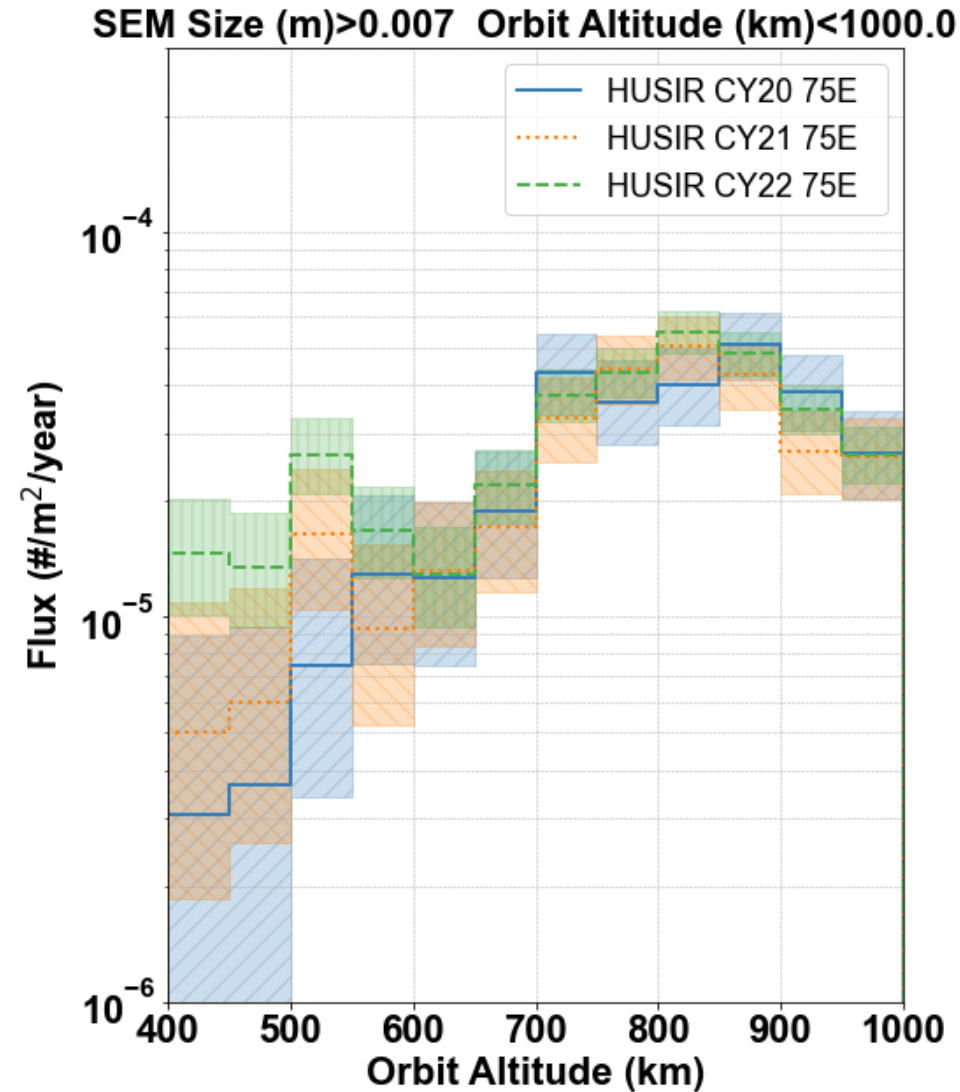


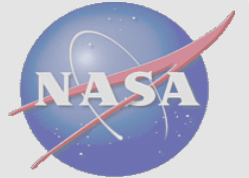
Count Rate vs. Size, 2020-2022



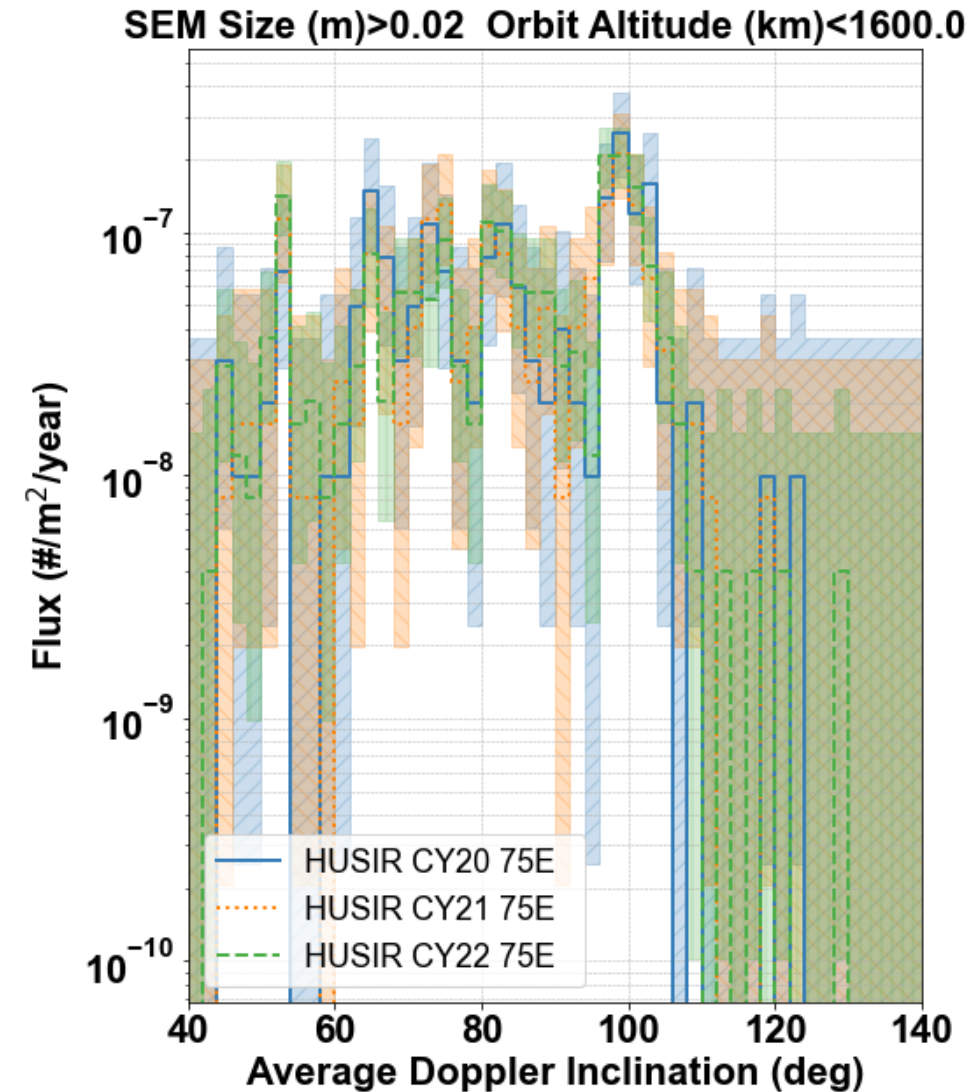
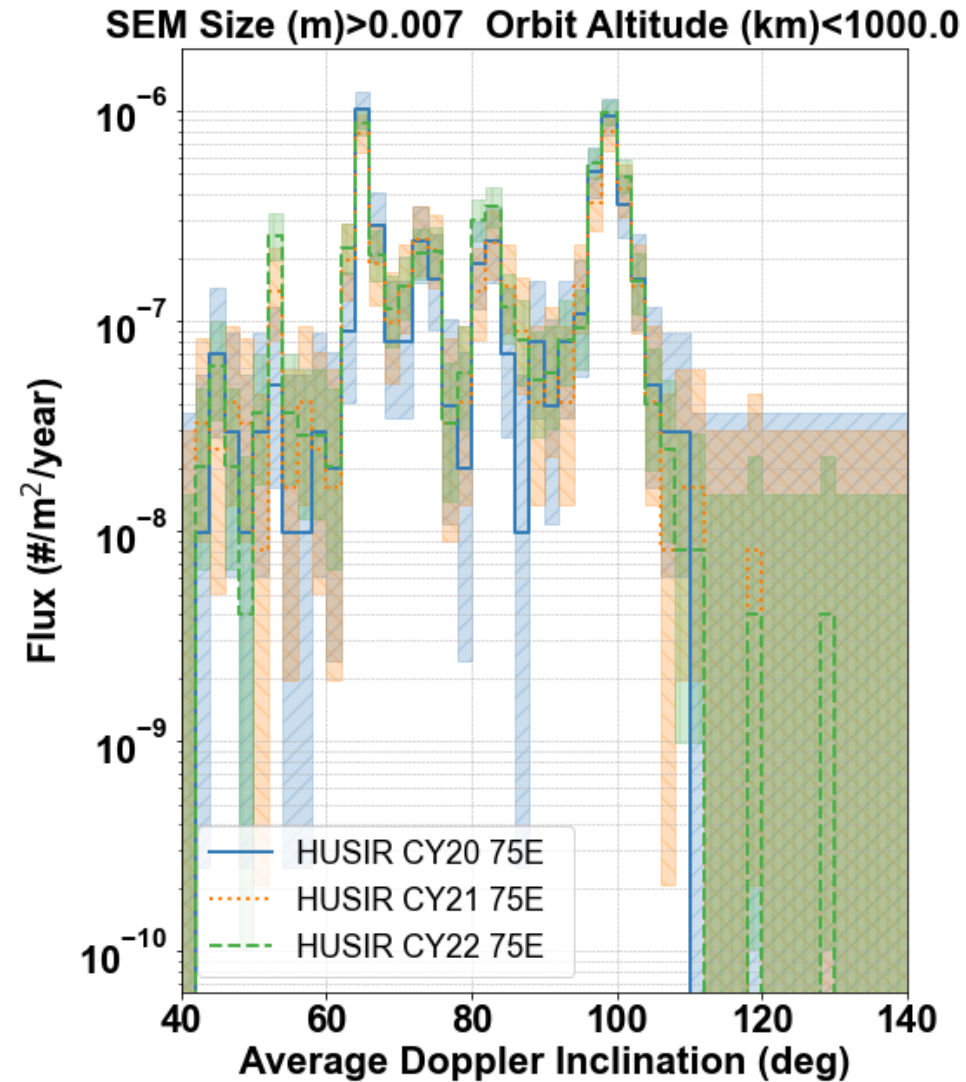


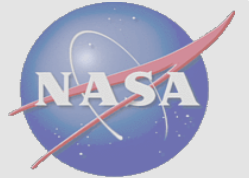
Flux vs. Altitude



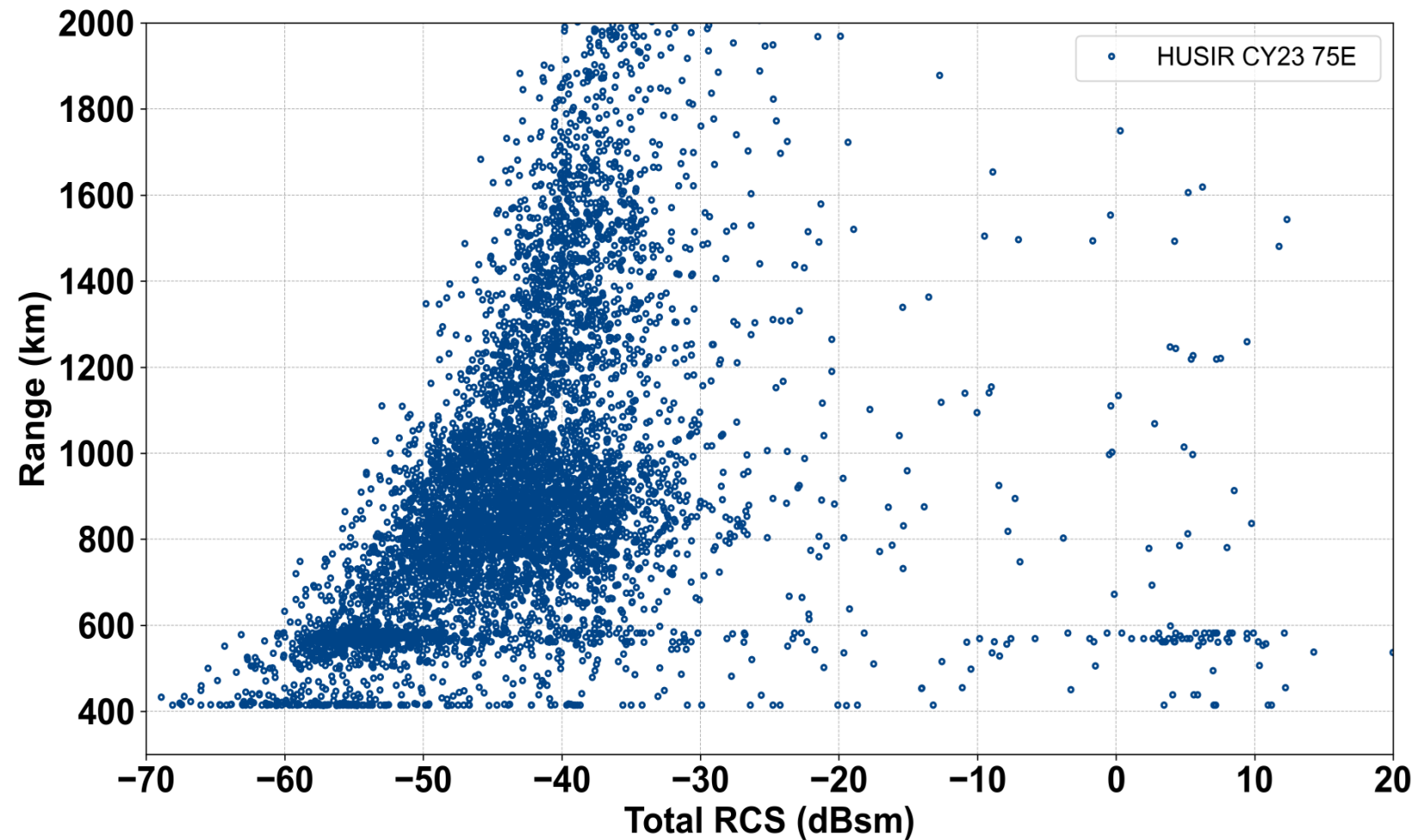


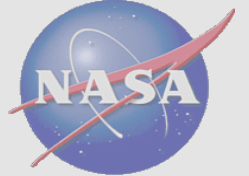
Flux vs. Inclination



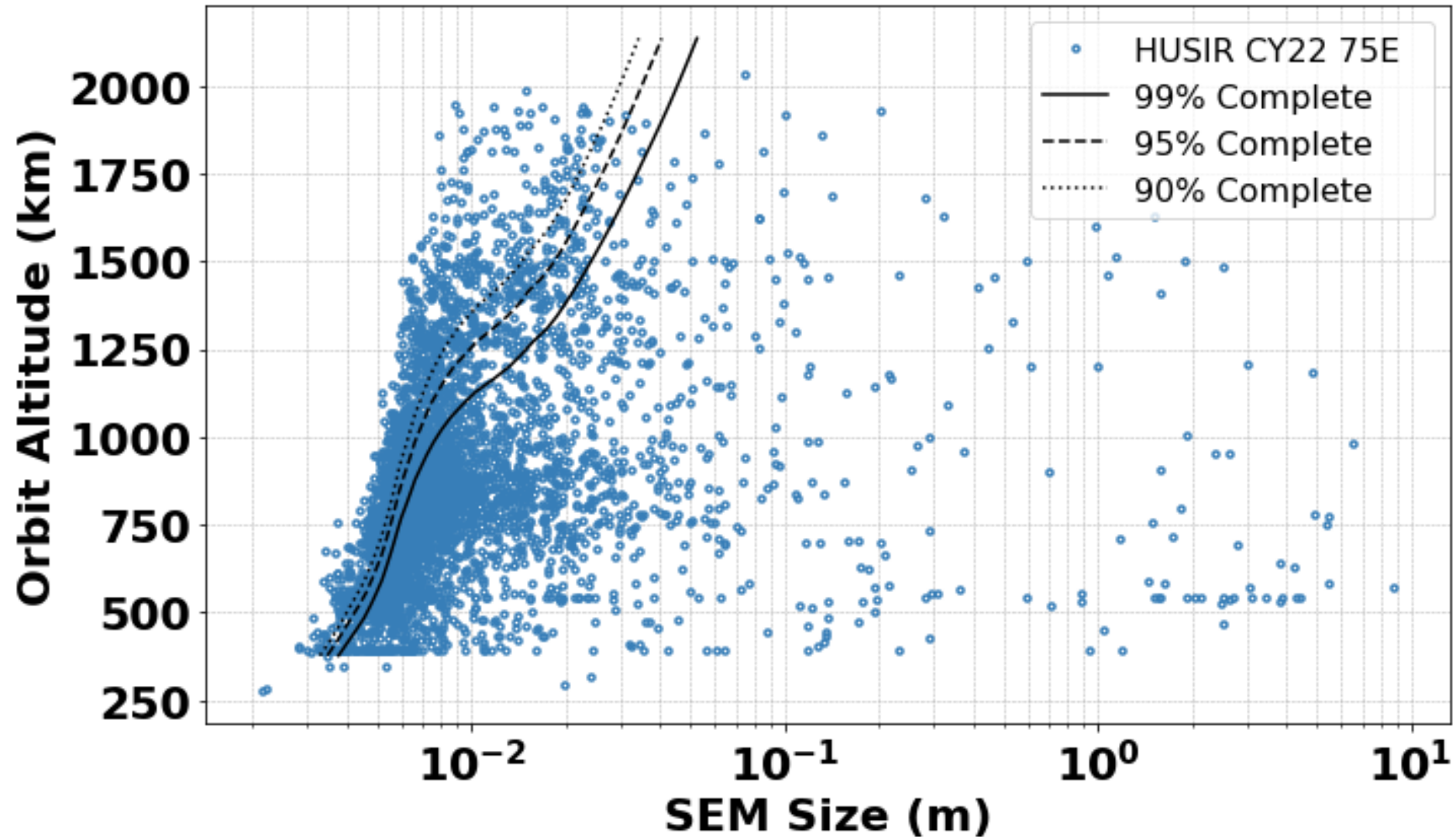


Range vs. RCS, HUSIR 75E



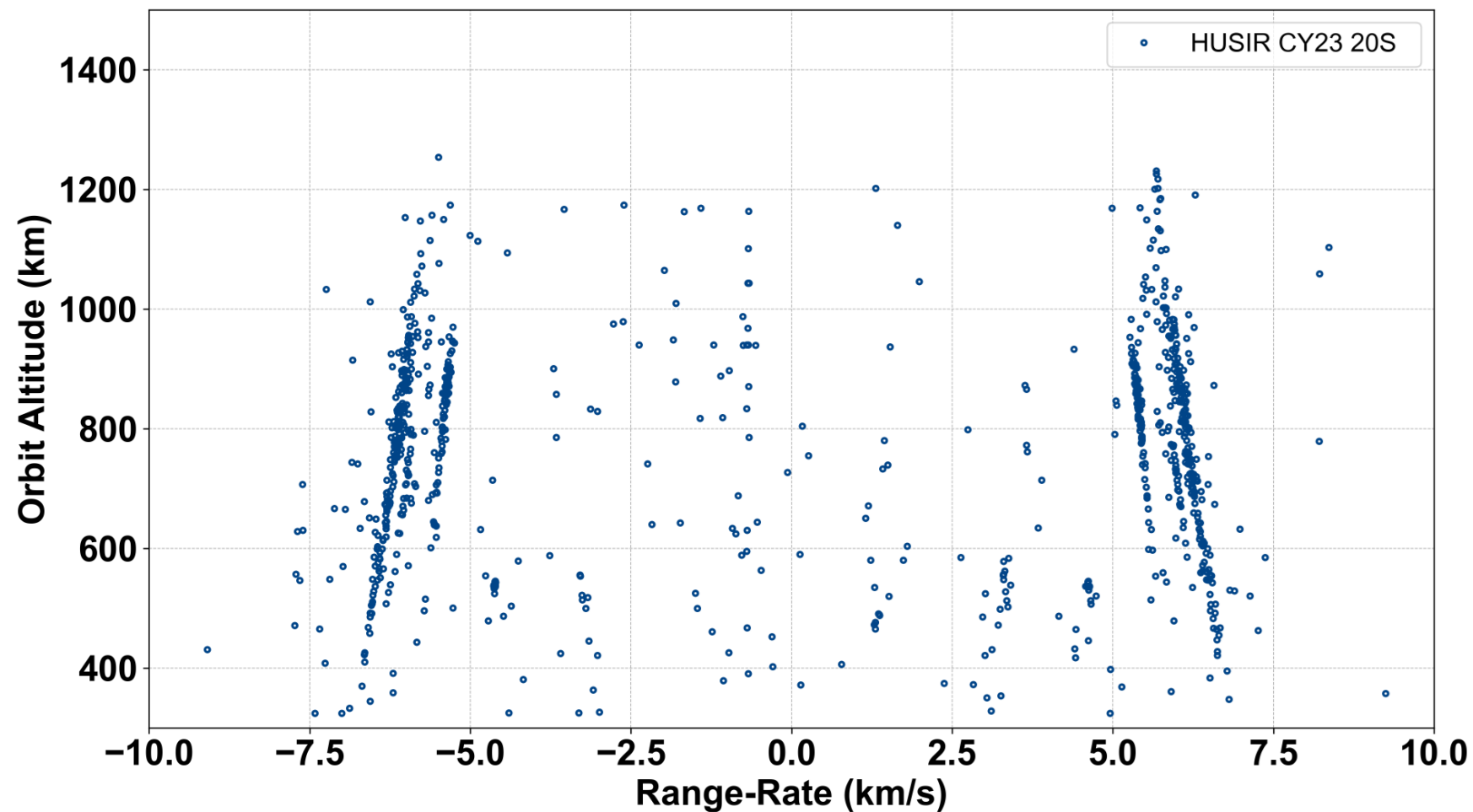


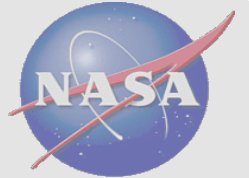
Altitude vs. Size, HUSIR 75E



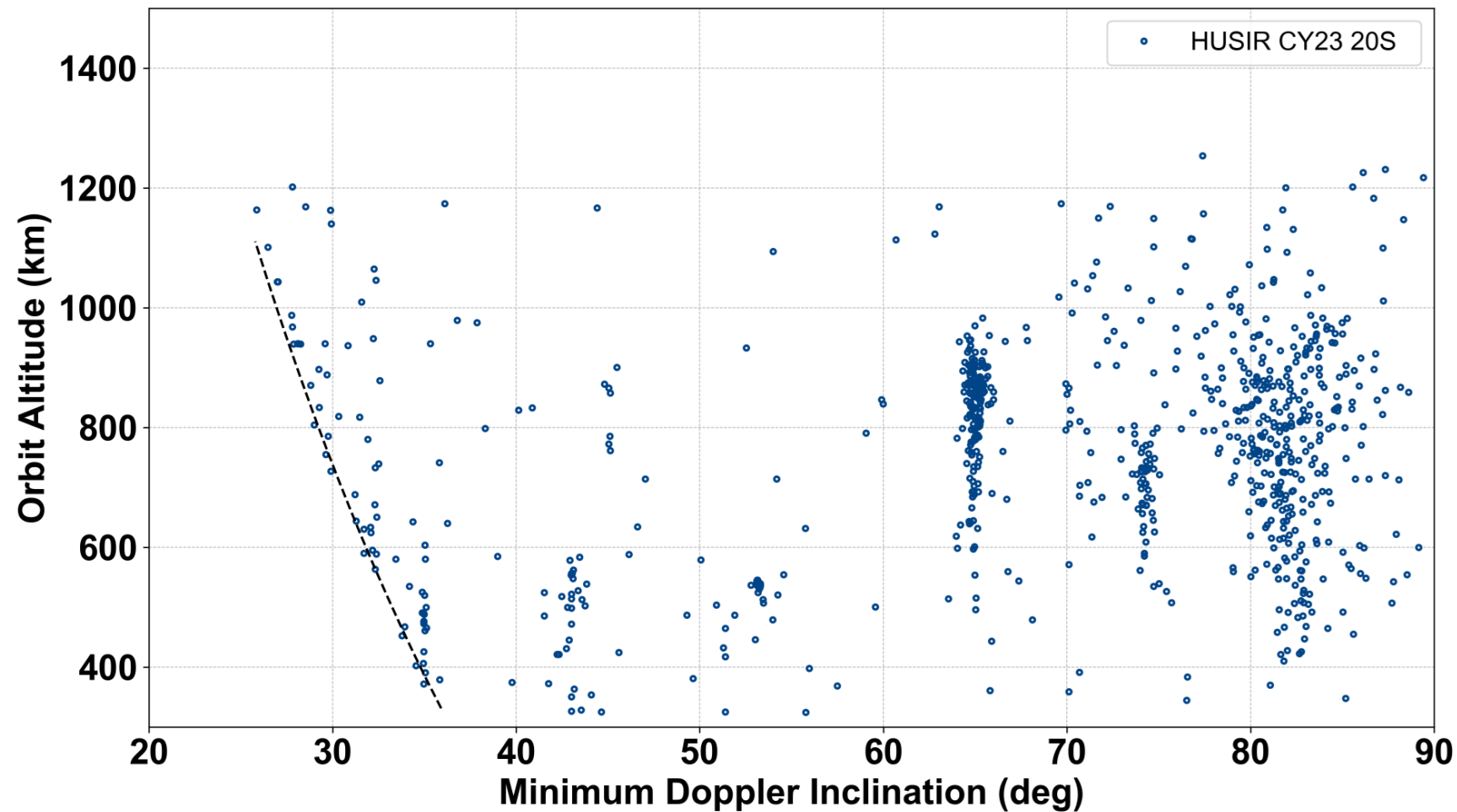


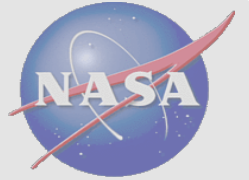
Range vs. Range-Rate, HUSIR 20S



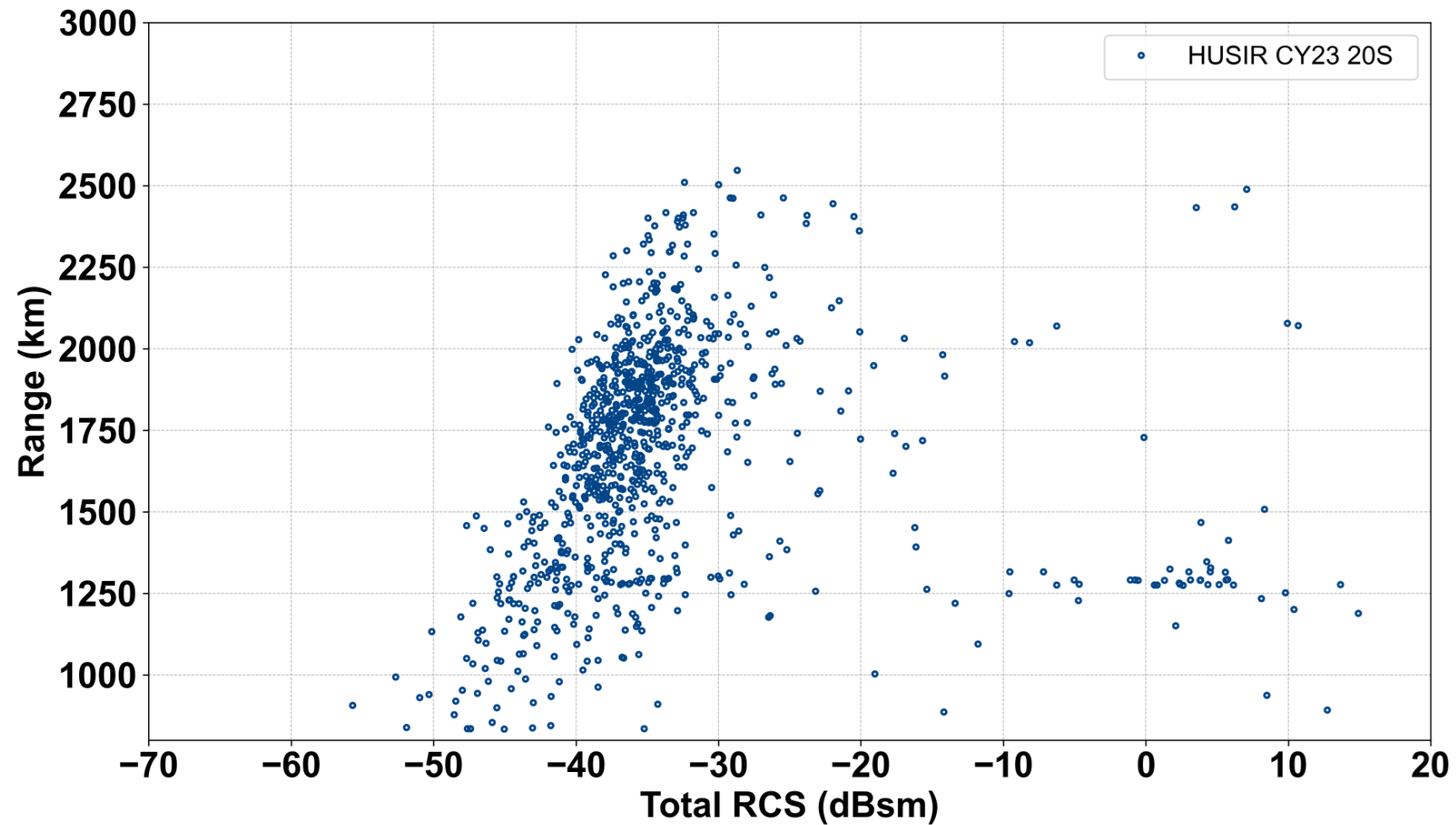


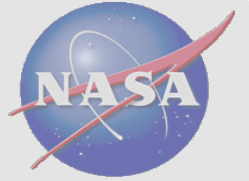
Altitude vs. Inclination, HUSIR 20S





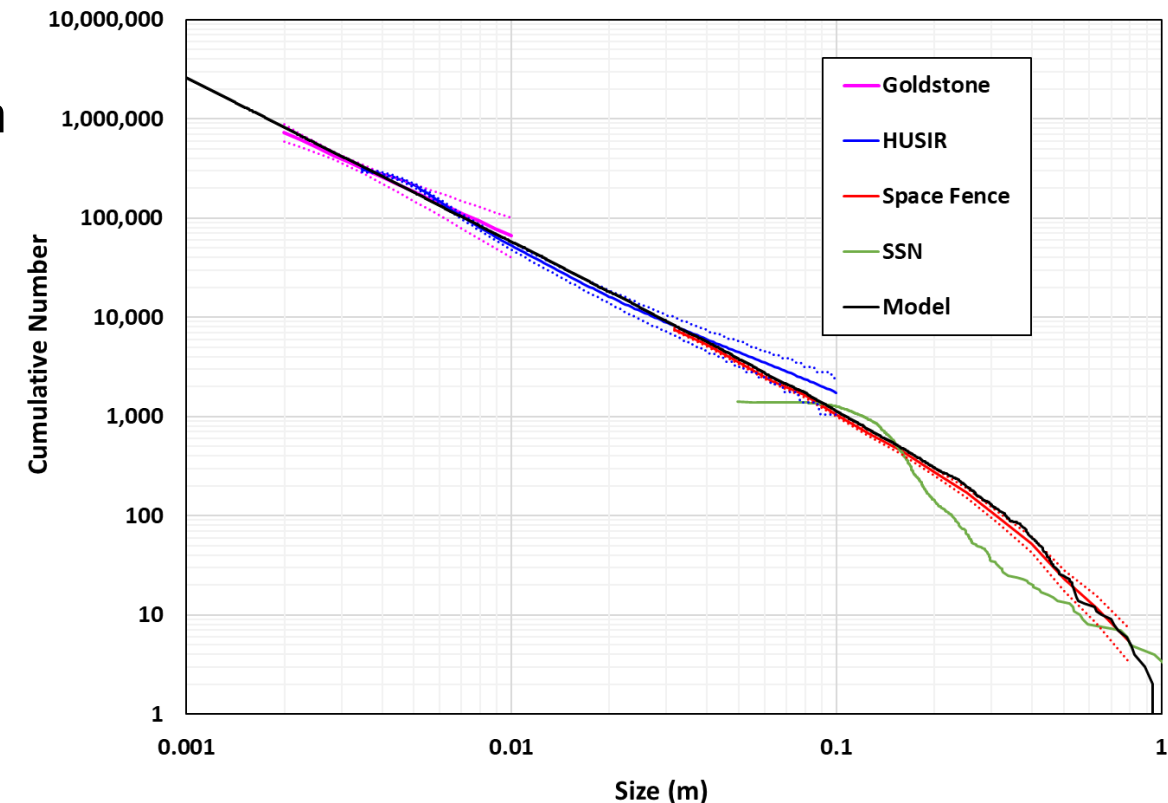
Range vs. RCS, HUSIR 20S





Special Observations – Cosmos 1408

- HUSIR and Goldstone are also used for special breakup observations
- Following the ASAT test conducted by the Russian Federation on Cosmos 1408 in November 2021, the ODPO partnered with MIT/LL and NASA JPL to collect special radar measurement data on small Cosmos 1408 fragments using HUSIR and Goldstone
- MIT/LL, in coordination with the 18 SPCS, also shared a unique Space Fence data set consisting of the initial two passes of the Cosmos 1408 fragment cloud
- Model predictions from the NASA Standard Satellite Breakup Model (SSBM) match special datasets very well

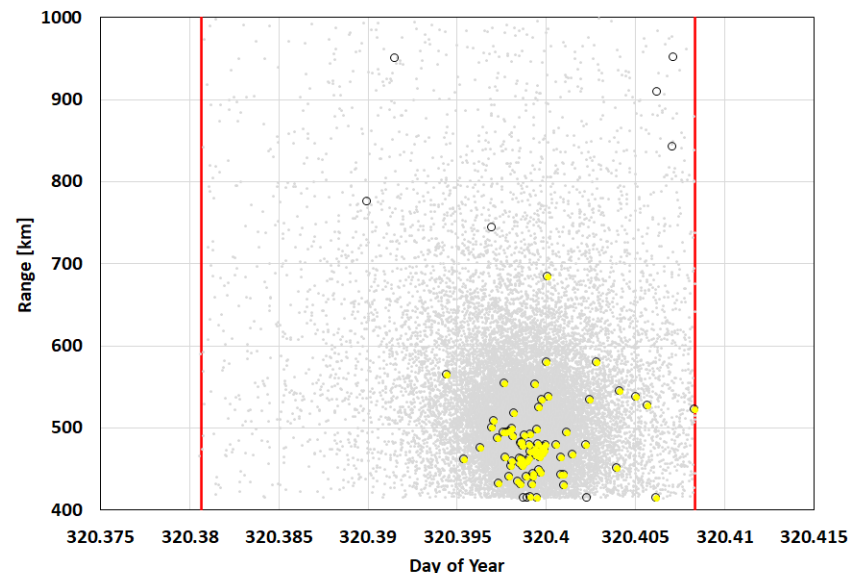


Cumulative size distribution of the Cosmos 1408 fragments based on special radar measurement data - the black line is the NASA SSBM prediction

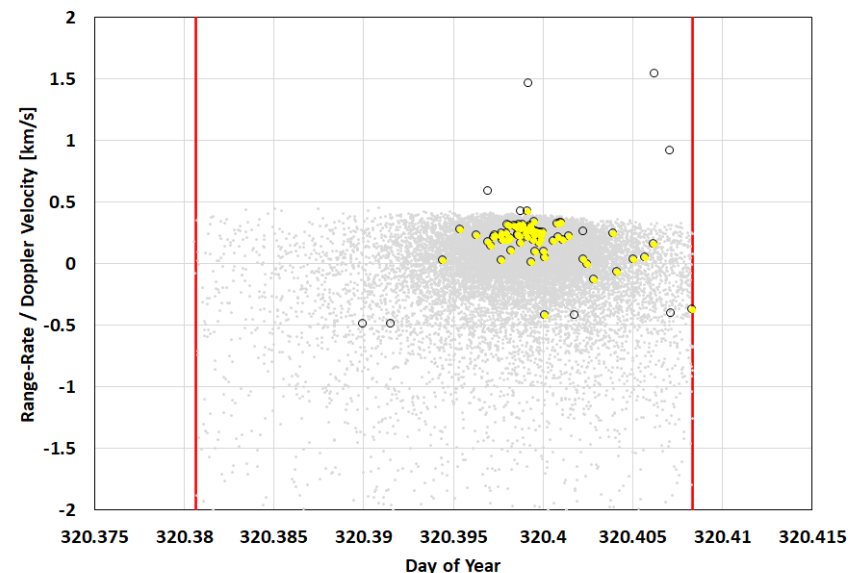


Special Cosmos-1408 Breakup Cloud Observation Example

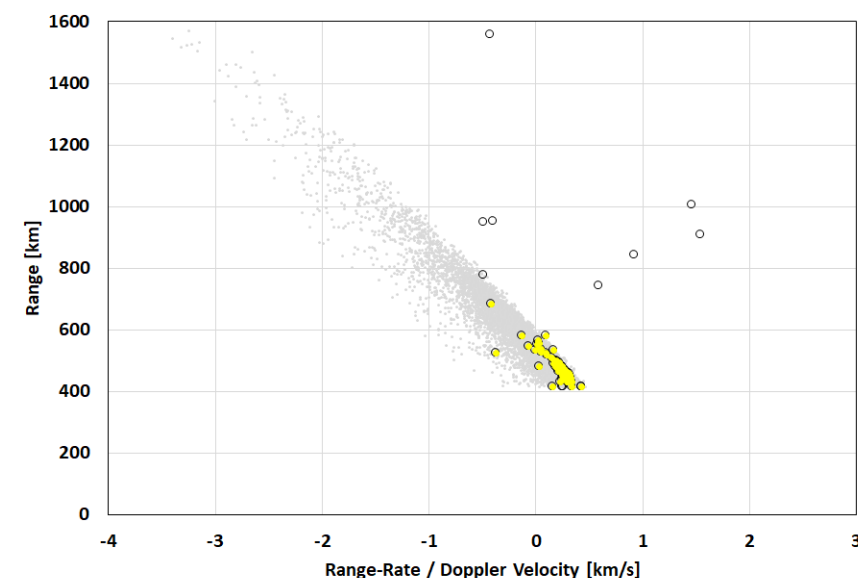
Day of Year 320A HUSIR Measurements



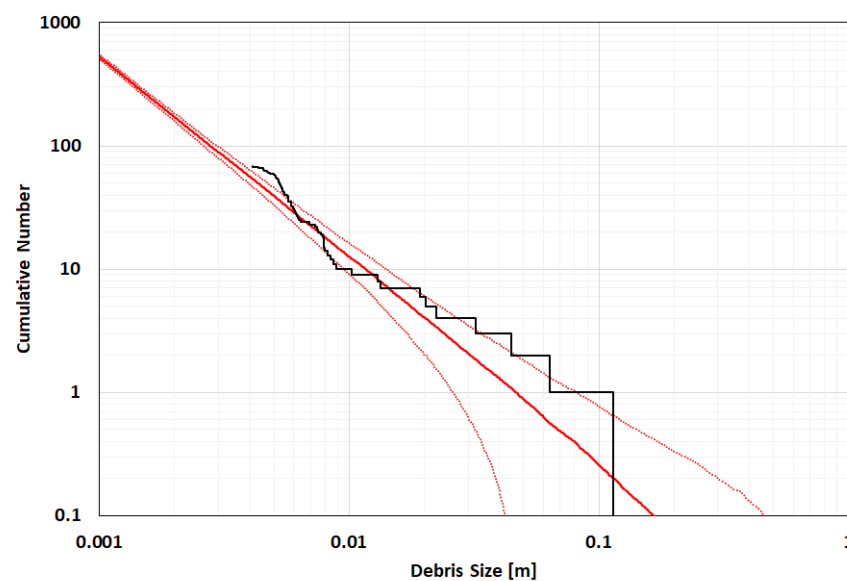
Day of Year 320A HUSIR Measurements



Day of Year 320A HUSIR Measurements

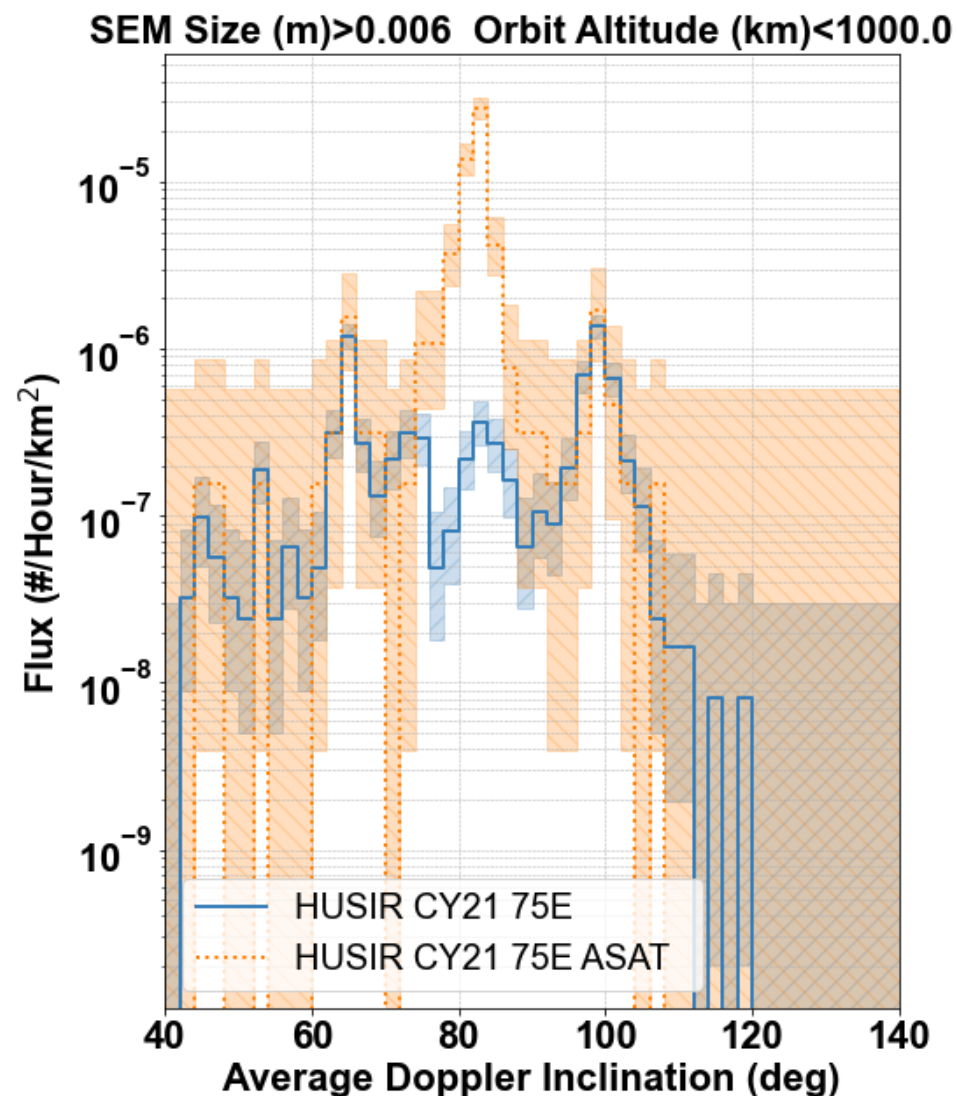
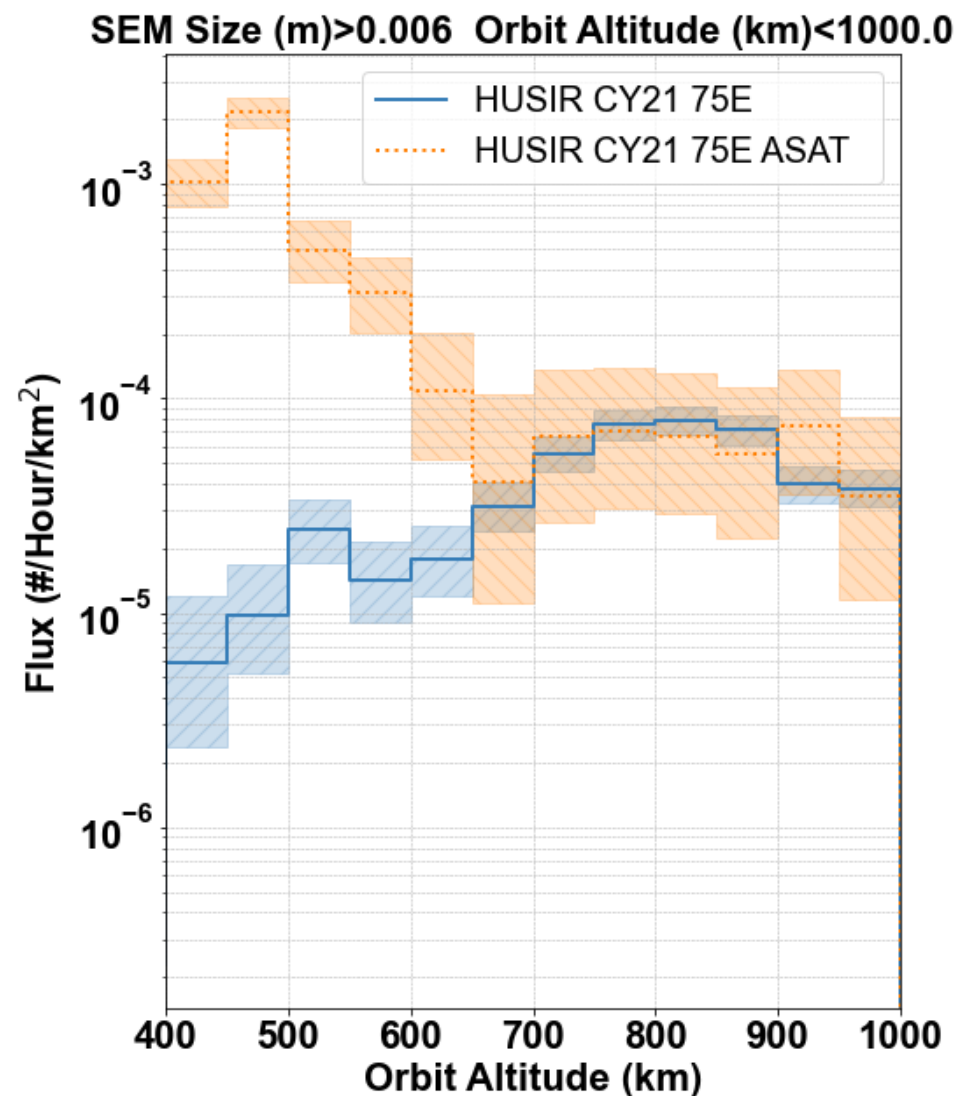


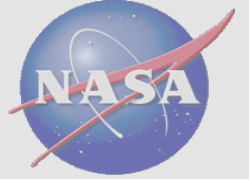
Size Distribution for Day of Year 320A





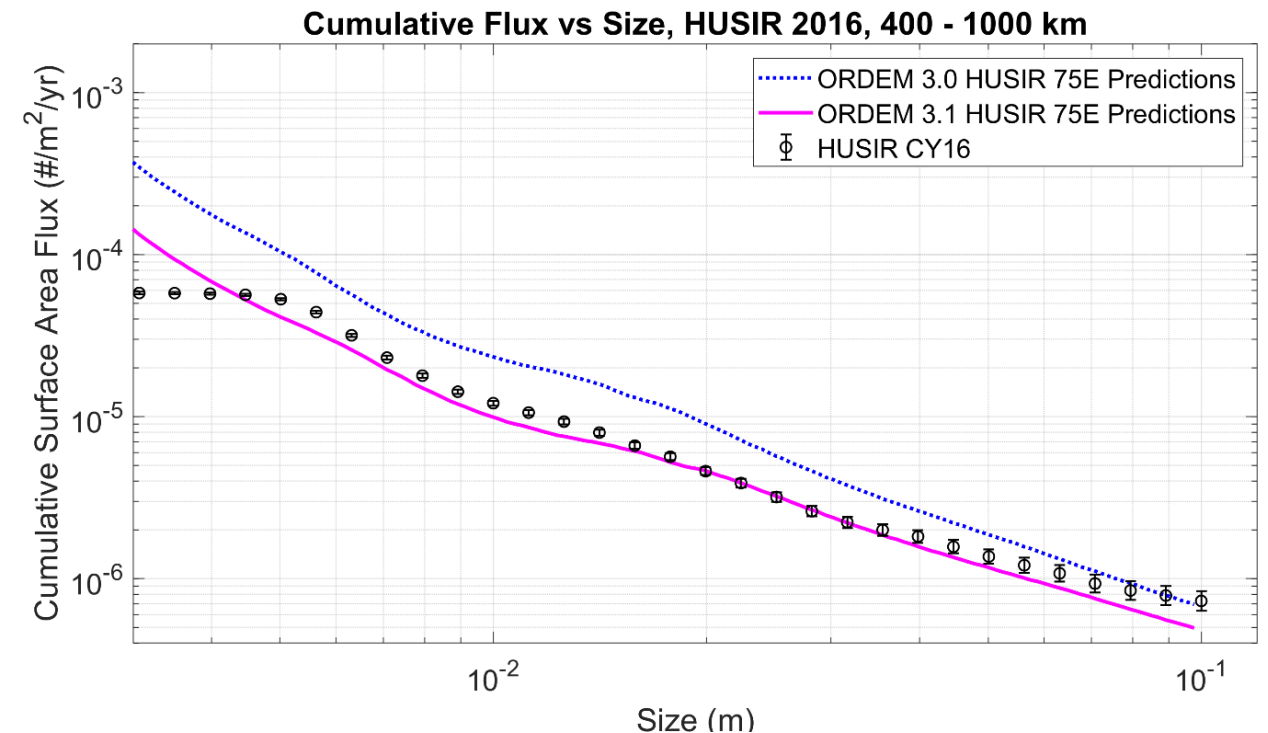
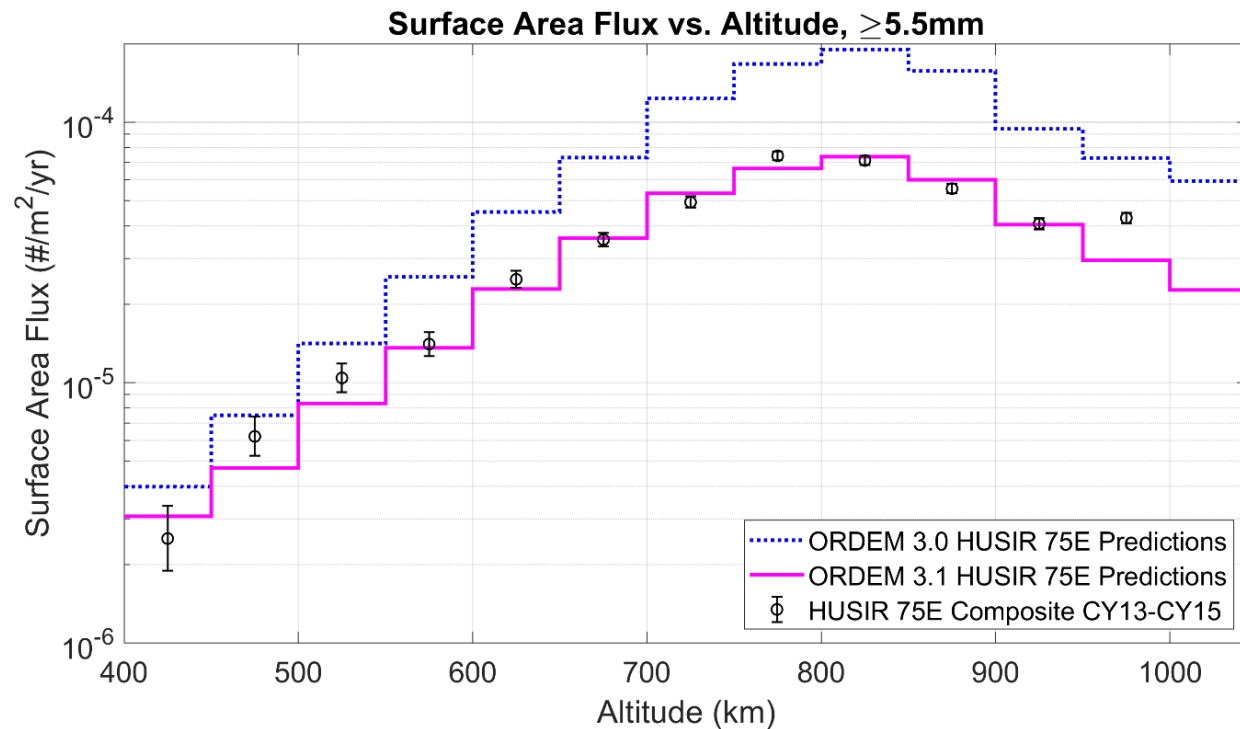
Flux vs. Altitude, Cosmos 1408

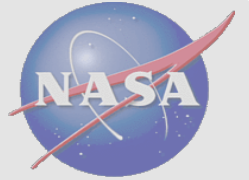




Developing OD Models

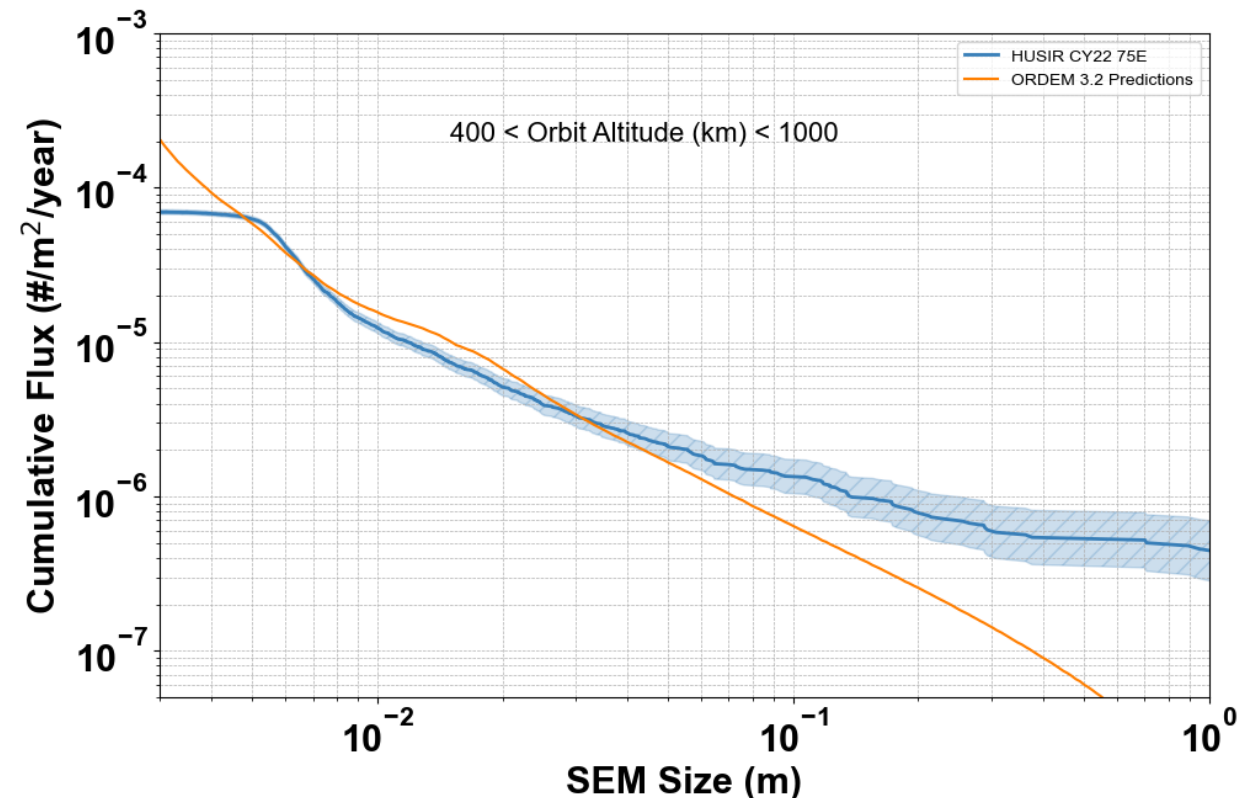
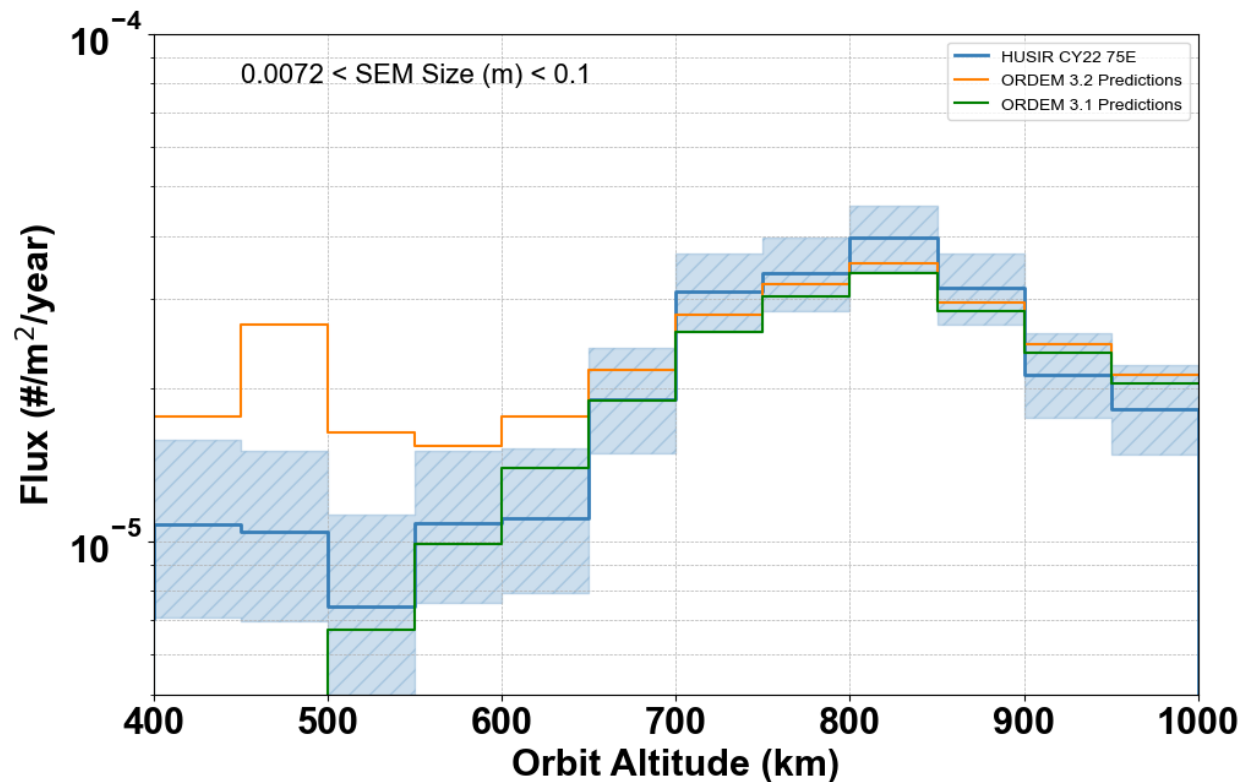
- HUSIR data is used to scale initial model populations for building and validating the Orbital Debris Engineering Model (ORDEM)**

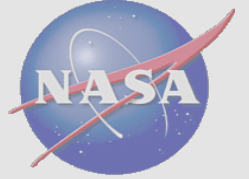




Developing OD Models

- HUSIR data is used to empirically scale model populations for building and validating the Orbital Debris Engineering Model (ORDEM)
- Work is ongoing on the next generation ORDEM 4.0, and HUSIR data will play a very important part in creating that model debris population





Summary

- **The SSN catalog is the tip of the proverbial iceberg when it comes to monitoring OD**
 - HUSIR/Haystack has been used to discover sources of small debris not identified by the SSN
- **The number of debris fragments increases dramatically with decreasing size**
- **Sensors such as HUSIR are vital tools for creating and validating the NASA models used to make accurate risk assessment for crewed/uncrewed missions**
- **Such sensors can be used to measure “fresh” breakups and characterize their risk to operational spacecraft**
 - Data can be used to validate our breakup models
- **The environment is dynamic, and needs to be continually monitored for dramatic changes**