

1 **Describing seasonal mixtures of cloud regimes via “regimes of regimes”**

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ABSTRACT

We propose a new type of cloud classification, relevant to monthly or longer time scales, but which inherently still encompasses cloud subgrid variability information at ~100 km scales. Our proposed classification partitions frequencies of occurrence over these scales of previously defined cloud regimes (CRs). We call the resulting distinct cloud entities regimes of regimes (RORs). While the CRs have been previously shown to successfully classify daily mesoscale subgrid variability via distributions of cloud fraction within distinct combinations of cloud top pressure and cloud optical thickness, the RORs essentially represent the prevalent seasonal mixtures of these CRs. RORs thus embody the seasonal cloudiness of a mesoscale region. We show that each ROR can still be associated with more traditional cloud classifications via composites of coincident active (lidar and cloud radar) cloud views. In a first application that gauges the potential utility of RORs, we pair them with CERES EBAF radiative fluxes to gain insight into recent trends of the cloud radiative effect. The ROR corresponding to an environment of shallow convection stands out in this analysis largely because of its declining population. Our study demonstrates the potential of RORs to categorize globally mesoscale cloudiness at monthly/seasonal scales and to serve as proxies of different atmospheric states at these scales.

27 **1. Introduction**

28 While the general features of the geographical distribution of cloudiness can be
29 consistently described in broad terms, i.e., “low” clouds occurring in subtropical marine
30 regions of atmospheric subsidence, deep convection occurring in the narrow geographical
31 zone with the strongest upward vertical motion known as the ITCZ, etc. (e.g., Fig. 7.4 in
32 Boucher et al. 2013), such vague, non-rigorous descriptions often lack consistency, which
33 limits their value. Defining cloud types from either passive or active satellite observations
34 (Welch et al. 1992; Bankert 1994; Rossow and Schiffer 1999; Wang and Sassen 2001;
35 Ceccaldi et al. 2013) is also somewhat arbitrary, having only an approximate correspondence
36 to the standard WMO cloud classification (WMO, 1956) traditionally used by surface
37 observers. Furthermore, it is often unclear what spatiotemporal range of scales are
38 appropriate for classifications based on satellite-derived cloud properties (Koren et al. 2008,
39 Wood and Field 2011). Are they meaningful only for instantaneous observations within the
40 instrument’s footprint, larger spatial scales that encompass multiple instantaneous footprints,
41 or multiple footprints across larger areas and over a certain time span? It is not uncommon for
42 the spatial persistence (coherence) of a particular cloud type to be weak, since a mesoscale
43 area of ~100 km comprising many instrument footprints can conceivably contain at any
44 instance multiple cloud types defined at the footprint level. Consider then the cloudiness in
45 this mesoscale area over a month, and even more cloud type diversity will be encountered.
46 How to characterize cloudiness at monthly or longer scales at such mesoscale regions
47 becomes then a conundrum. Distributions obtained from grid-level spatial averages of
48 instantaneous cloud properties are narrower and conceal the true nature of cloudiness we seek
49 to classify. Preservation of cloud variability at the original spatial scale of the instantaneous
50 observation should perhaps then become a consideration in cloud classification schemes.

51 Cloud regimes (CRs), aka Weather States (WSs) (Jakob and Tselioudis 2003; Rossow et
52 al. 2005; Tselioudis et al. 2013; Oreopoulos et al. 2014, Oreopoulos et al. 2016; Cho et al.
53 2021), have been previously introduced as a convenient and arguably physically meaningful
54 approach to achieve cloud classification of mixtures of different cloud types that mitigates the
55 loss of information resulting from homogenization due to spatial averaging. Based on daily or
56 three-hourly joint histograms describing the frequency of occurrence of discrete pixel-level
57 cloud top pressure (CTP) and cloud optical thickness (COT) combinations during sunlit
58 hours, where each combination can be viewed as a “cloud type”, cloud regimes provide a

59 compact way to describe semi-instantaneous cloud subgrid variability at mesoscales. Using
60 however distributions of cloud properties retrieved at the pixel-level scale for cloud
61 classifications presents its own challenges. Instead of the narrowing effect on distributions of
62 grid-level averaging, the pixel-level cloudy property distributions become broader as
63 temporal scales expand. If one were to construct, for example pixel-level CTP-COT joint
64 histograms over increasingly longer time scales and/or larger spatial scales, they would
65 become progressively more featureless (more likely to be devoid of distinct frequency peaks)
66 the longer the sampling period and the larger the domain, rendering themselves once again
67 unsuitable for cloud grouping and classification. In short, when using cloud property
68 distributions for cloud classification, the specific spatiotemporal scales chosen to build them
69 also matter.

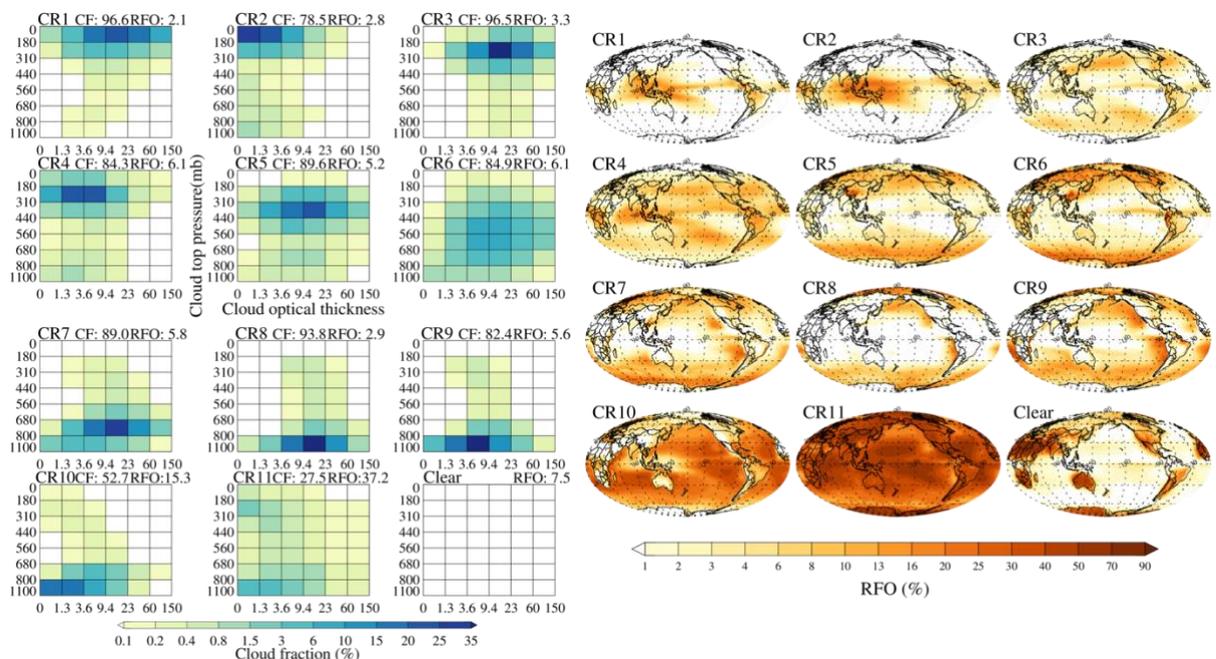
70 Motivated by this state of affairs, in this work we pay attention to the temporal aspect of
71 cloud classification and investigate whether we can overcome the inherent obstacle of spatial
72 distributions becoming progressively more featureless the longer the sampling time scale. We
73 specifically investigate whether a meaningful cloud classification at monthly or longer time
74 periods that preserves information about instantaneous subgrid variability at mesoscales is
75 possible. In essence, we seek an effective way to bypass the undesired smoothing and loss of
76 information that characterizes temporal distributions of grid-average values, while also
77 avoiding the loss of distinct features that accompanies the expansion of the reference period
78 over which pixel-level distributions are built. We make the case that a new concept of cloud
79 classification we call “regimes of regimes” (RORs) avoids both pitfalls and can serve as a
80 useful descriptor of seasonal cloudiness at mesoscales. We accordingly introduce new cloud
81 classes representing CR mixtures that tend to re-occur over staggered 3-month periods. The
82 paper’s main goal is then to examine whether RORs are indeed a physically appropriate and
83 viable concept that can serve as a novel climatological description of the planet’s cloudiness.

84 In the remainder of the paper, we detail the derivation of RORs, provide their physical
85 interpretation, and explore their potential utility for understanding not only the mean state of
86 cloud radiative effects, but also their recent trends and implied short-term feedbacks.

87 **2. ROR Derivation**

88 The RORs of this study are derived from the 1° daily MODIS CR dataset of Cho et al.
89 (2021). The salient characteristics of these CRs are expressed via: (1) their centroids, namely

90 mean joint histograms of the clusters representing the CRs obtained from applying a
 91 clustering algorithm on MODIS daytime 1° CTP-COT joint histograms, described in Cho et
 92 al (2021); and (2) the multi-year mean maps of CR relative frequency of occurrence (RFO)
 93 over a 20-year period obtained from assigning a CR daily to each grid cell. These CR
 94 centroids and RFO maps are shown in Fig. 1. To obtain RORs, we built histograms of CR
 95 normalized frequency of occurrence (FOC, per the terminology of Tselioudis and Rossow
 96 2011), over “rolling” 3-month periods for all one-degree grid cells containing CRs (Fig. 2),
 97 essentially a 12-element vector of the grid cell’s seasonal RFO. For example, a CR FOC
 98 histogram for January-February-March is followed by a February-March-April histogram,
 99 which is then followed by one for March-April-May, and so forth. The FOC histogram is
 100 nominally assigned to the middle month of the 3-month period. Such overlapping 3-month
 101 windows are employed to achieve big enough sample sizes that suppress the noise of 1°
 102 histograms. CRs derived from both morning (Terra satellite) and afternoon (Aqua satellite)
 103 MODIS cloud data enter the construction of these histograms, so each of them is built from 3
 104 (month) × ~30 (days per month) × 2 (Terra and Aqua) ≈ 180 values. An example of such a
 105 FOC histogram is shown in Fig. 2.

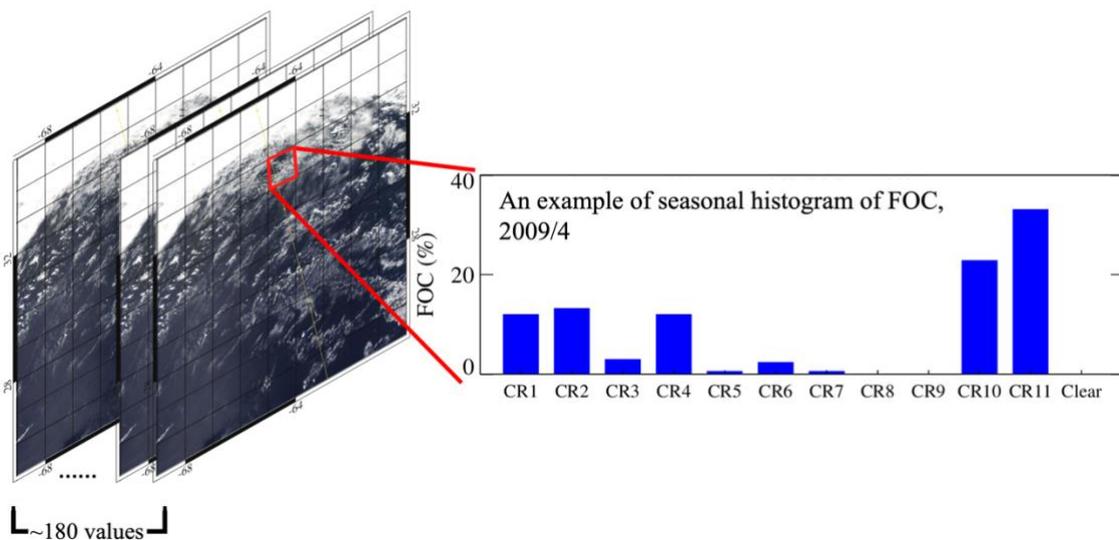


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107 Fig. 1. Centroids (cluster mean joint histograms from *k*-means clustering, left) and maps
 108 of mean relative frequency of occurrence (RFO) of the 1° daily MODIS CRs of Cho et al.
 109 (2021).

110 A large number of CR FOC histograms (~15M) such as the one in Fig. 2 were furnished
 111 to a *k*-means clustering algorithm. While objective criteria exist to determine the number of

112 clusters that optimize centroid distinctiveness (McDonald et al. 2016; Schuddeboom et al.
 113 2018; Jin et al. 2021), our experience from previous implementations of k-means in cloud
 114 classification exercises suggests that such algorithmically derived “optimal” numbers yield
 115 clusters that are too correlated geographically to be considered physically distinct. We thus
 116 follow again our previous practice of settling on the desired number clusters by trial and error
 117 based on visual inspection and judicial assessment of the output. Using such an empirical
 118 approach, we settled on K=8 as the “best” number of clusters to compactly describe the
 119 planet’s cloudiness in terms of seasonal one-degree RORs. While a larger number would
 120 theoretically provide a better separability of centroids, in practice we noticed that the
 121 algorithm does not actually split the clusters that would apparently benefit the most, i.e.,
 122 those with somewhat excessive CR mixing. Another possible avenue to improve cluster
 123 separation, namely performing additional clustering only on the histograms of the clusters
 124 deemed too mixed, was not attempted at this time because more (albeit presumably better
 125 separated) clusters undermine the practical utility of RORs when employed in subsequent
 126 analyses. Staying with K=8 has also the additional benefit of no extreme disparities in ROR
 127 global occurrences: global ROR RFOs range between ~7% and ~20% indicating that no ROR
 128 is exceedingly rare or overly frequent. While RORs can arguably be refined in future efforts,
 129 the present set satisfies our first order goal to demonstrate the physical foundation and
 130 viability of the concept.



131

132 Fig. 2. Illustration of a seasonal (3-month) histogram of CR normalized FOC (Terra and
 133 Aqua) in a 1° grid cell. ~15M of such histograms are clustered to produce RORs, with each
 134 histogram constructed from ~180 values.

135 It should be noted that the seeds of an ROR-like concept were previously planted by
136 Tselioudis and Rossow (2011). In that work their entities represented nearly instantaneous
137 tropical WS mixtures based on FOC histograms built daily at large spatial scales of $30^{\circ}\times 30^{\circ}$.
138 Hence, fundamental differences with the current work are that the WS FOC histograms of
139 Tselioudis and Rossow (2011) reflect variations in much larger domains than mesoscale, and
140 at semi-instantaneous rather than seasonal temporal scales. Hence the physical meaning of
141 their cloud mixtures is quite different than ours, as they capture the fast propagation of
142 expansive features of the tropical circulation through the prism of WS snapshots covering
143 very large areas. Our RORs on the other hand showcase the seasonal variability of persistent
144 cloud mixtures at mesoscales.

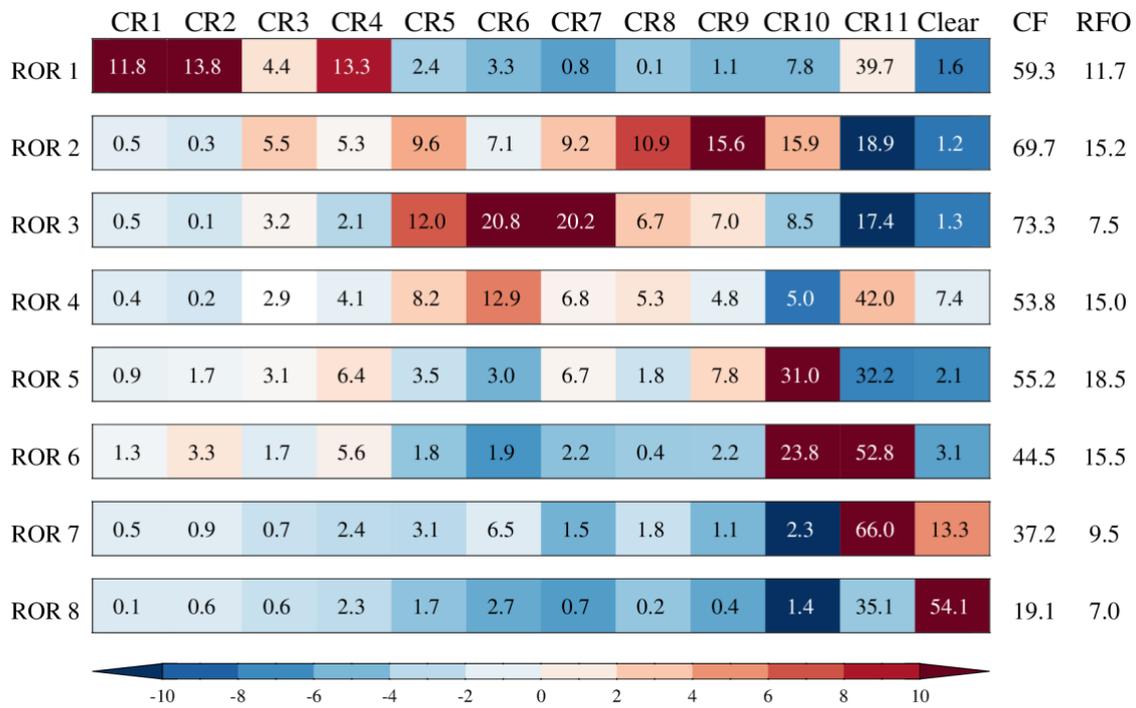
145 **3. ROR definitions and characteristics**

146 In order to come up with a physical interpretation of the RORs we examined their
147 centroids as obtained from the k -means clustering procedure, monthly “snapshot” maps of
148 their dominant geographical location, the statistics of geographical occurrences over the $20 \times$
149 $12 = 240$ months of our dataset, composites of cloud features derived from a fundamentally
150 distinct type of (coincident) space-based observations, namely active lidar and radar
151 observations, and the prevailing dynamical environment, as expressed by large-scale mid-
152 tropospheric vertical velocity, in which RORs are embedded. All these aspects of RORs are
153 presented in this section.

154 *a. ROR centroids*

155 Centroids, i.e., mean CR FOC histograms assigned to each cluster by the clustering
156 process provide the first clues on the nature of the RORs (Fig. 3). Our visualization of the
157 ROR centroids as horizontal “mosaic” bars shows both the normalized FOC of each CR
158 within an ROR (via the numbers displayed), but perhaps most appropriately, how the FOC
159 within an ROR differs from the CR’s global RFO (via the colors, where blue hues correspond
160 to values below and red hues to values above the CR’s global RFO; the deeper the colors the
161 larger the deviation from global RFOs). Next to the horizontal bars we provide: (1) the
162 ROR’s global cloud fraction (CF) obtained from globally averaging (using cosine latitude as
163 weight) the CF of successful MODIS daytime cloud retrievals of each grid cell whose CR
164 belongs to the ROR occurrence; and (2) the ROR’s global RFO obtained by normalizing the

165 number of specific ROR occurrences by the total number of all RORs encountered in our
 166 dataset, and then weighting by cosine of latitude.



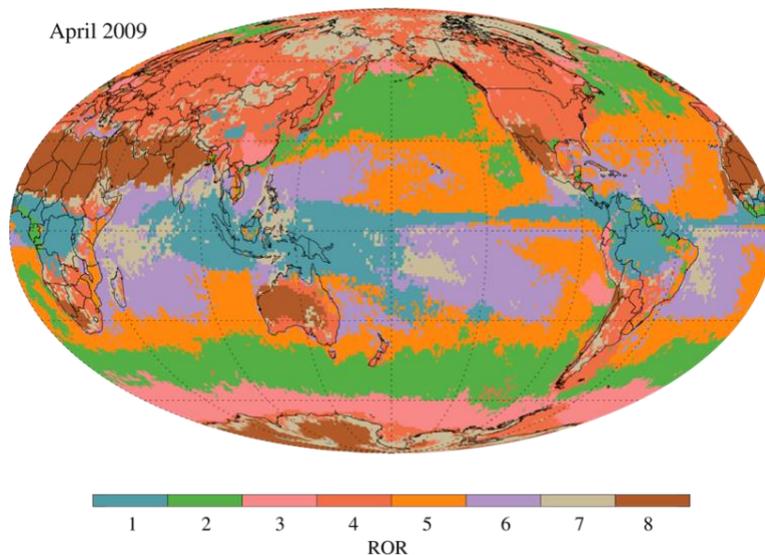
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168 Fig. 3. Centroids (mean CR FOC histograms from *k*-means clustering) corresponding to
 169 the RORs, shown as “mosaic” bars where each mosaic “piece” corresponds to a CR. The
 170 centroids are represented both by numbers showing the normalized FOC (adding to 100%) of
 171 each CR in the ROR’s CR mixture, and colors capturing the FOC deviation within the ROR
 172 from the CR’s global RFO, as explained in the text.

173 *b. Snapshots of ROR geographical occurrence*

174 For each 1° grid cell, the 3-month CR FOC histogram (which, as mentioned previously,
 175 can be thought as a 12-element array), nominally attributed to the middle month, is assigned
 176 to the ROR whose centroid produces the minimum Euclidean distance from the grid cell’s
 177 array. Ultimately, an ROR is assigned to each grid cell for every month, with the end result of
 178 this action being an “instantaneous” monthly snapshot map, as in Fig. 4. Our 20-year dataset
 179 yields 240 such maps available for download at <https://doi.org/10.5281/zenodo.11061757>,
 180 which can also be viewed in succession as an animation (available in the Supplemental
 181 Material SM) to gauge ROR seasonal and year-to-year movements. The example map of Fig.
 182 4, for April 2009 (from assignment of March-April-May 2009 CR FOC histograms) as well
 183 as the SM’s animation loop of the entire dataset supports the notion that the ROR concept has
 184 an underlying physical foundation of : RORs are generally organized geographically into
 185 large contiguous areas that reveal the climatological distribution of cloud systems with

186 discerning characteristics on a planetary scale. For this specific month, we note the following:
 187 The most convective cloud systems (ROR1) are limited to the tropics and occupy large parts
 188 of the ITCZ and SPCZ, and are surrounded by the low broken cloud tropical and subtropical
 189 shallow convection of ROR5 and ROR6. The NH land masses are blanketed by ROR4 which
 190 contains many weather-producing mid-level clouds, while the midlatitude oceans are
 191 dominated by ROR2 which contains stratiform clouds, including the marine stratocumulus
 192 not belonging to ROR5. ROR3 is encountered mostly at high latitudes, being quite prominent
 193 in the far southern oceans surrounding Antarctica, but also in the northern Atlantic south of
 194 Greenland. The low cloud fraction ROR7 does not appear to be too frequent in this season,
 195 and besides polar and regions of high relief, has also presence in oceanic tropical areas of
 196 sparse cloudiness. Finally, the semi-clear ROR8 is associated with deserts and other areas
 197 where low humidity may exhibit seasonal variations.



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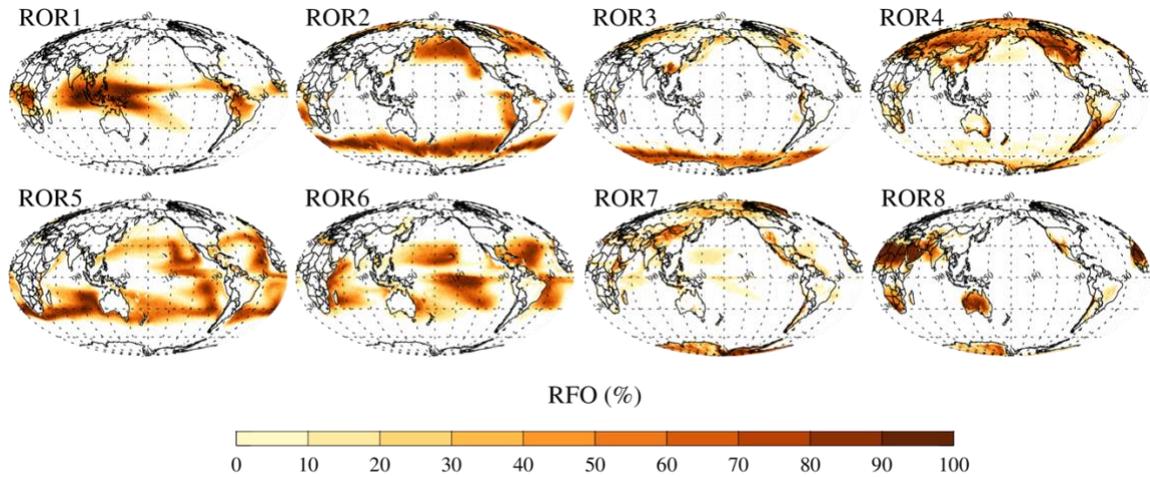
199 Fig. 4. Snapshot map of ROR assignment of 3-month CR FOC histograms for March-
 200 April-May 2009 (nominally assigned to April 2009).

201 *c. Maps of ROR RFOs*

202 In addition to global RFO and monthly snapshot maps of ROR occurrence, we can also
 203 examine the multi-annual average geographical prevalence of RORs. Geographical
 204 distributions of RFO can be created by counting for each grid cell the fraction of the time an
 205 ROR occurs over the 240-month period. The resulting ROR RFO maps (normalized counts of
 206 ROR occurrence) are shown in Fig. 5. These maps clearly show that the RORs generally

207 exhibit coherent and rather contiguous geographical distributions, broadly consistent with the
208 descriptions we provided above for April 2009.

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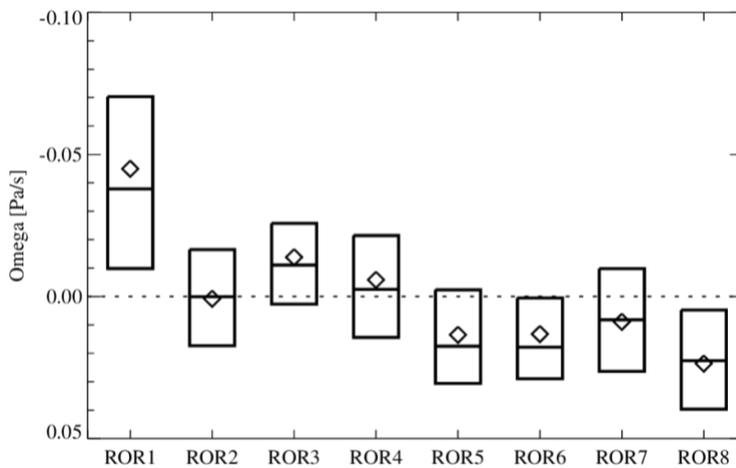


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211 Fig. 5. Multi-annual (20 years) maps of ROR RFO obtained by counting monthly ROR
212 occurrences over the analysis period in each 1° grid cell and normalizing by the total number
213 of ROR occurrences (nominally 240) in the grid cell.

214 *d. ROR interpretation*

215 Additional insight on RORs can be gained by constructing distributions of mid-
216 tropospheric 500 hPa large-scale vertical velocity (“omega”) from coincident MERRA-2
217 monthly reanalysis values. This reanalysis assimilates modern satellite observations not
218 available to its predecessor, MERRA, and includes updates to the Goddard Earth Observing
219 System (GEOS) model and analysis scheme (Gelaro et al. 2017). The omega distributions
220 corresponding to each ROR are visualized via their quartile values in the form of a boxplot in
221 Fig 6. The dynamical environment revealed by this figure, along with the previously shown
222 CR FOCs within ROR centroids, global CFs and RFOs, snapshot maps, and geographical
223 distribution of ROR RFOs provide quite a comprehensive picture on the nature of the RORs
224 which we can now summarize as follows:



225

226 Fig. 6. Mean (diamonds), median (horizontal lines), and interquartile range of large-scale
 227 500 hPa vertical velocity distributions (boxes) constructed by matching monthly MERRA-2
 228 values with ROR occurrences for our entire 20-year dataset. Negative values indicate large
 229 scale ascent while positive value large scale descent.

230 • Made of anomalously high amounts of CR1 (tropical deep convection) and CR2 (tropical
 231 anvils and cirrus), but also CR3 (tropical and midlatitude convection) and CR4
 232 (midlatitude and subtropical high clouds), ROR1 contains by far the highest proportion of
 233 high and deep clouds and is the rarest of the seven RORs with substantial average cloud
 234 fractions. Given the prevalence of CR1 and CR2, ROR1 unsurprisingly occurs in an
 235 environment of strong ascent (Fig. 6). It is the most tropical of all RORs with large
 236 presence primarily in the tropical western Pacific and secondarily over tropical Africa and
 237 South America, reflecting the mean annual position of the ITCZ. Monthly ROR snapshot
 238 maps examined either individually or as an animation (available in the SM or at
 239 <https://doi.org/10.5281/zenodo.11061757>) show that the buildup and decay of the Indian
 240 monsoon can be tracked via ROR1. Moreover, ENSO signals can also be gleaned through
 241 ROR1: for example, comparing the July ROR occurrence map of 2015 (strong El Niño
 242 initiated earlier that year) with the July map of 2010 (strong La Niña initiated earlier that
 243 year), clearly shows the anomalous 2015 expansion of deep convection represented by
 244 ROR1 into the eastern tropical Pacific.

245 • ROR2 has large mean CF and encompasses CR8 (thick low stratus) and CR9 (oceanic
 246 stratocumulus) at rates substantially above climatology, with the former representing
 247 thick low stratus and the latter mostly marine stratocumulus. ROR2's RFO map
 248 consequently shows extensive coverage at both the subtropics and mid-latitudes, with the
 249 latter regions likely contributing to the above climatology occurrence of storminess via

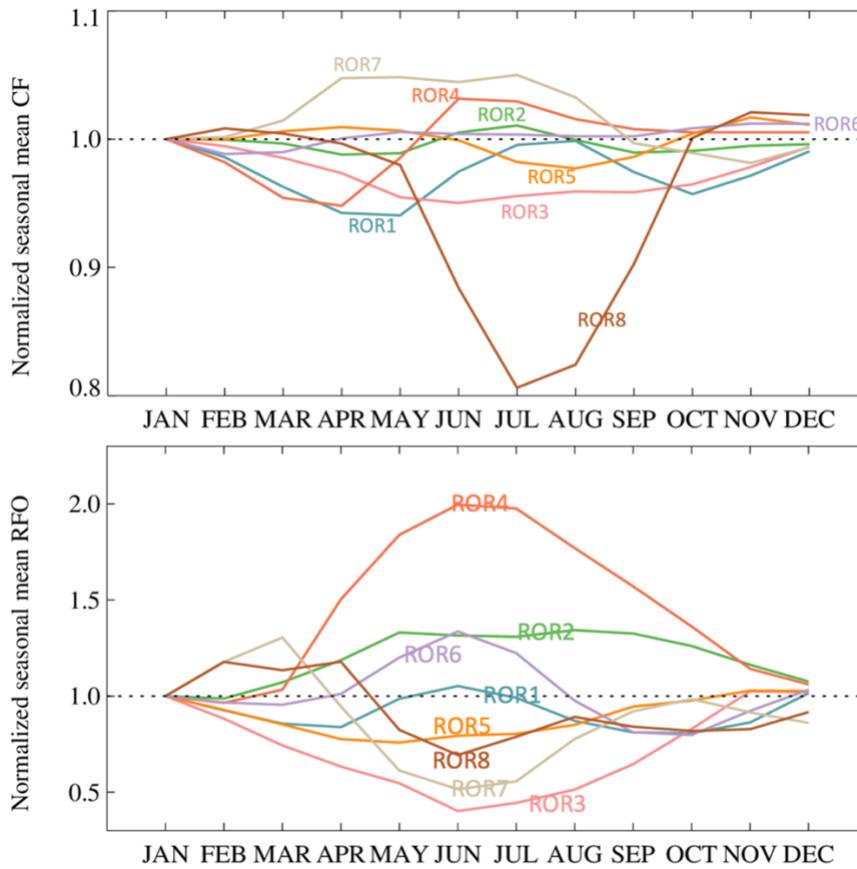
250 CR3 and CR5 (high-latitude storms). The extratropical branch of ROR2 is likely where
251 ascending motion primarily occurs, while the subtropical sector contributes to the
252 downward motions seen in Fig. 6's boxplot.

- 253 • Occurring mostly at high latitudes, ROR3 is notable for its large mean CF and high
254 relative contributions from CR6 (high-latitude midlevel clouds) and CR7 (thick high
255 stratus) which represent the bulk of mid-level clouds and the highest and thickest of
256 stratus, respectively. Fig. 6 indicates that this ROR occurs in regions of persistent large-
257 scale ascent.
- 258 • ROR4 is the most continental of the RORs, but also has substantial polar presence. It is
259 characterized by an excess of mid-level clouds (CR5 and CR6) that can occur in the
260 presence of either upward or downward large-scale motion.
- 261 • Being the most frequent of all RORs, ROR5 also stands out as the ROR with
262 anomalously high contributions from CR10 (shallow oceanic convection) and
263 anomalously low contributions from CR11 (low cloud fraction regime). While these two
264 CRs have the lowest cloud fraction, they nevertheless differ substantially with CR10
265 having nearly double the CF (~53% vs ~27%), allowing ROR5 to still have a
266 considerably large mean CF representing mostly shallow oceanic convection. This
267 interpretation is consistent with prominent presence further away from the west coast of
268 the major continents, with the exception of Australia where marine stratocumulus are of
269 more broken nature. The shallow nature of ROR5 clouds is consistent with subsidence
270 prevailing in their subtropical environment.
- 271 • In contrast to ROR5, ROR6 receives anomalously high contributions from both CR10 and
272 CR11 indicating dominant presence of broken clouds that are mixtures of both active
273 shallow convection and also suppressed conditions of fair weather (predominantly
274 downward large-scale motion in Fig. 6) in tropical and subtropical latitudes.
- 275 • ROR7 is made up chiefly from CR11 which is the CR of lowest CF associated with dry
276 conditions in the poles and mountainous areas. The upward motions seen in its omega
277 distribution probably come from the low latitudes where this ROR occasionally occurs.
- 278 • ROR8 is a semi-clear ROR containing large amounts of completely clear grid cells
279 (CR12) and thus has by far the lowest cloud fraction of all RORs. It is almost always

280 associated with large-scale descent occurring over deserts, polar regions, and (presumably
281 the leeward side of) mountains.

282 *e. Seasonal Variability*

283 Although no immediate association with the nature of cloudiness represented by each
284 ROR can be made by itself, an examination of the mean annual cycle of ROR RFO and CF
285 provides a compact view of the characteristics of their seasonal variability in terms of the
286 timing of maxima/minima and the magnitude of within-year variations. Figure 7 shows the
287 20-year-mean annual cycle of ROR CF, when the ROR occurs (top), and of ROR RFO
288 (bottom). One can immediately see that the annual cycles of CF and RFO are generally
289 distinct from each other. The convention of plotting the annual cycle as normalized values
290 with respect to the January multi-annual mean provides the cleanest visualization and
291 conveniently reveals whether January is a (near) minimum value month (i.e., mostly positive
292 deviations occurring for the other months) or a (near) maximum value month (mostly
293 negative deviations occurring for the other months). The biggest CF annual cycle is exhibited
294 by ROR8, the ROR with the smallest cloud amount (~19% annual mean value), with the
295 minimum dropping to ~16 % in the boreal summer months. This is the time of the year when
296 ROR4 RFO nearly doubles, which seems to be associated (according to the time lapse of
297 ROR occurrence in the SM) with a northward expansion of cloudiness over the NH
298 continents and polar regions, as also seen in the seasonal cycles of CR5 and CR6 in Fig. 5 of
299 Cho et al. (2021). The seasonal variabilities of the other RORs are less pronounced, with
300 perhaps ROR3 deserving some attention given its January peak of both RFO and CF. The
301 seasonal cycle of ROR3 seems to be dominated by its Southern Ocean branch, and reflects
302 the peak in thick stratiform cloudiness in the summer months, consistent with the peak in
303 CR7 DJF RFO in Fig. 5 of Cho et al. (2021).

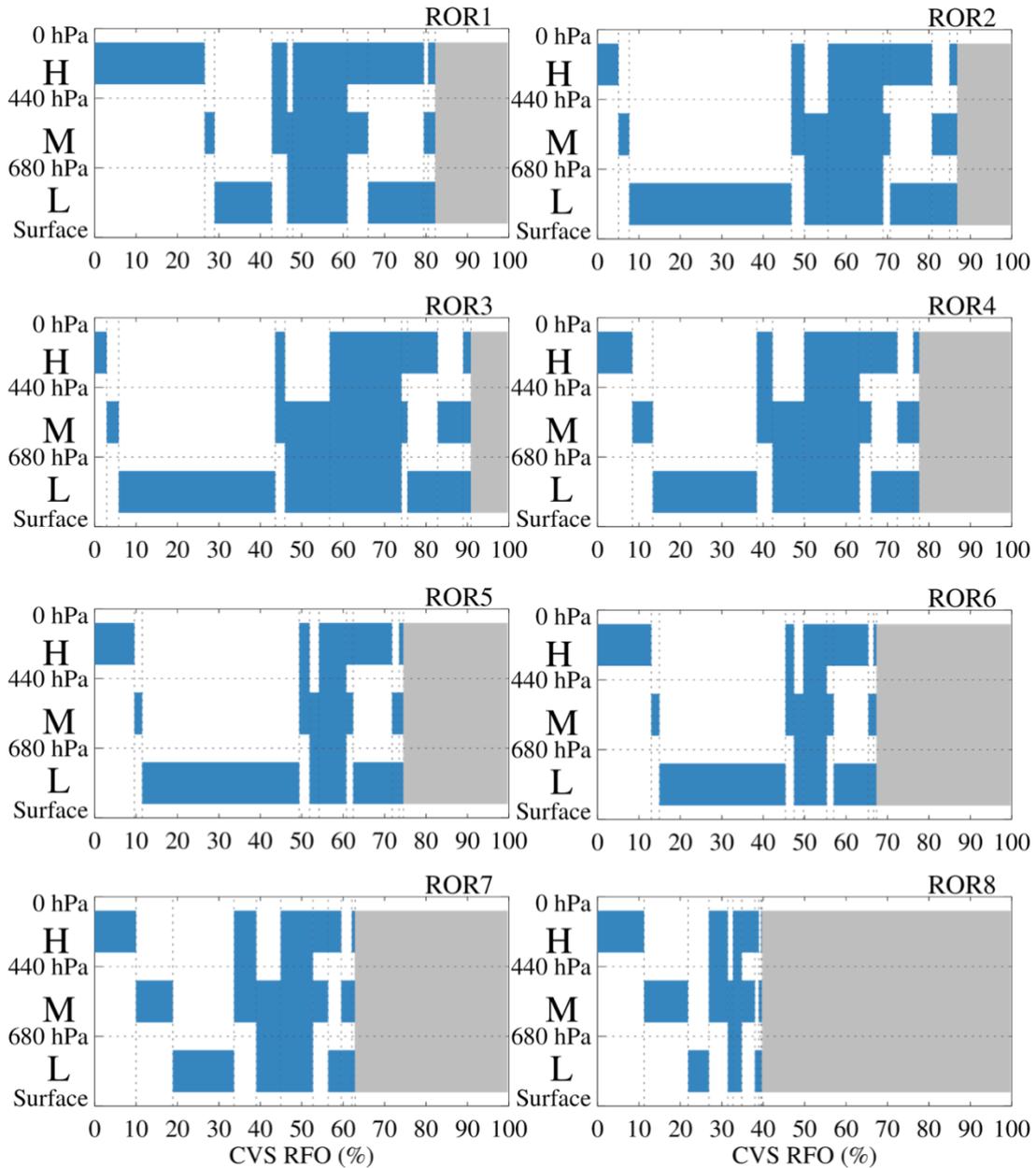


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305 Fig. 7. The mean seasonal cycle of ROR CF (top panel) and RFO (bottom panel) cycle
 306 normalized to multi-annual January values.

307 *f. Active views*

308 It has been previously shown (e.g., Cho et al. 2021) that cloud products from active
 309 observations taken contemporaneously within grid cells occupied by CRs can be used to
 310 examine their vertical structure as well as their constitution in terms of cloud types
 311 originating from an independent classification that exploits the unique capabilities of active
 312 observations. The same probing can also be done on RORs, the only material difference
 313 being that the compositing of active cloud products is performed over a three-month period.



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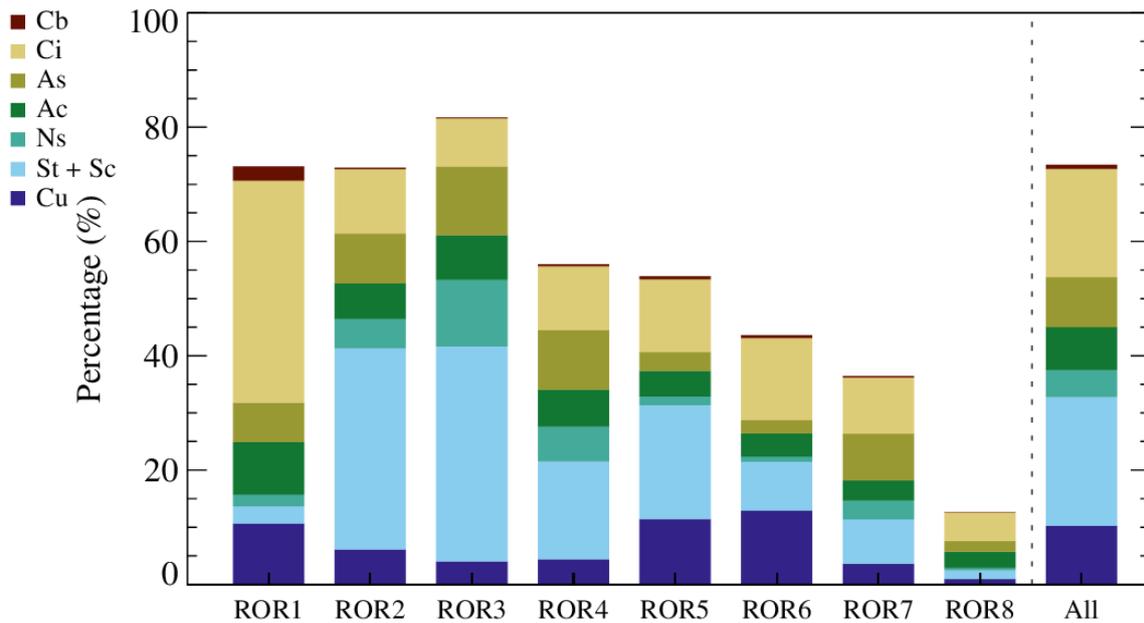
315 Fig. 8. Cloud Vertical Structure (CVS) composites for each ROR for the period 2007-
 316 2020. All panels have the same CVS ordering, with the bar width indicating the relative
 317 contribution of each CVS to the ROR. The CVSs were derived by applying the algorithm of
 318 Oreopoulos et al. (2017) on the 2BCL R05 Cloudsat dataset.

319 Figure 8 shows the makeup of each of each ROR in terms of the 10 Cloud Vertical
 320 Structures (CVSs, 11 with clear skies) introduced by Oreopoulos et al. (2017). The CVS
 321 dataset used here is similarly based to the 2B-CLDCLASS-LIDAR (2BCL) dataset (Sassen et
 322 al. 2008), but updated to reflect the most recent Release 5 (R05) version of the dataset, and
 323 was expanded to cover more years (14 years, 2007-2020). 2BCL combines CloudSat Cloud
 324 Precipitation Radar (CPR) and CALIPSO (CALIOP) lidar measurements and takes advantage
 325 of the different sensitivity to hydrometeor type of the two instruments to detect cloud phase

326 and classify clouds at various layers into different types. The fact that the CloudSat sampling
327 strategy varies throughout this period and becomes daytime-only after 2010 is immaterial for
328 our purposes since CRs and RORs are also daytime-only datasets. In the Fig. 8 visualization
329 the 10 CVS classes are represented by blocks in three broad tropospheric vertical layers
330 (segments) where low (*L*), middle (*M*) and high (*H*) clouds reside, delineated per ISCCP
331 convention by 680 hPa and 440 hPa atmospheric pressure boundaries (Rossow and Schiffer
332 1999). Blocks can appear in either one, two, or all three of these segments. When present in
333 more than one segment, clouds can occupy the segments either contiguously (indicated as
334 connected) or non-contiguously (separated blocks). The horizontal width of the block
335 indicates the percent contribution of each CVS class to the ROR.

336 Figure 8 provides the following insights: ROR2, ROR3, and ROR5 all have large
337 amounts of isolated low clouds, indicative of either isolated cumulus clouds or stratocumulus
338 fields. But ROR1, ROR2, and ROR3 also contain high proportions of vertically developed
339 clouds connected across all three segments suggestive of a deep convective nature. ROR4
340 also comprises appreciable amounts of vertically developed clouds given the second largest
341 (after ROR3) combined proportion of vertically connected clouds spanning either all three or
342 the low and middle segments (the latter suggestive of congestus and cyclonic clouds). The
343 largest by far proportion of stand-alone high clouds is found in tropical ROR1, many of
344 which are likely cirrus outflows associated with deep convection. ROR8 has the lowest cloud
345 fractions of all RORs, but even so it has clouds ~40% of the time on average according to
346 active observations which (on the lidar side) are more sensitive to thin clouds, the largest
347 proportion of which is high clouds, consistent with its prevalence in arid environments.

348 Comparing the CVSs of these RORs to the CVSs of the CRs (Cho et al. 2021), there is
349 less CVS distinction between the RORs than between the CRs, which is simply a reflection of
350 the fact that the RORs are seasonal aggregations and thus contain a more diverse mixture of
351 cloud types. Nevertheless, the fact that the CVS composition of the RORs exhibits distinct
352 patterns and contrasts (e.g., the amount of isolated high clouds in ROR1 and ROR3) is a
353 testament of the inherent ability of RORs to identify different seasonal cloud conditions.

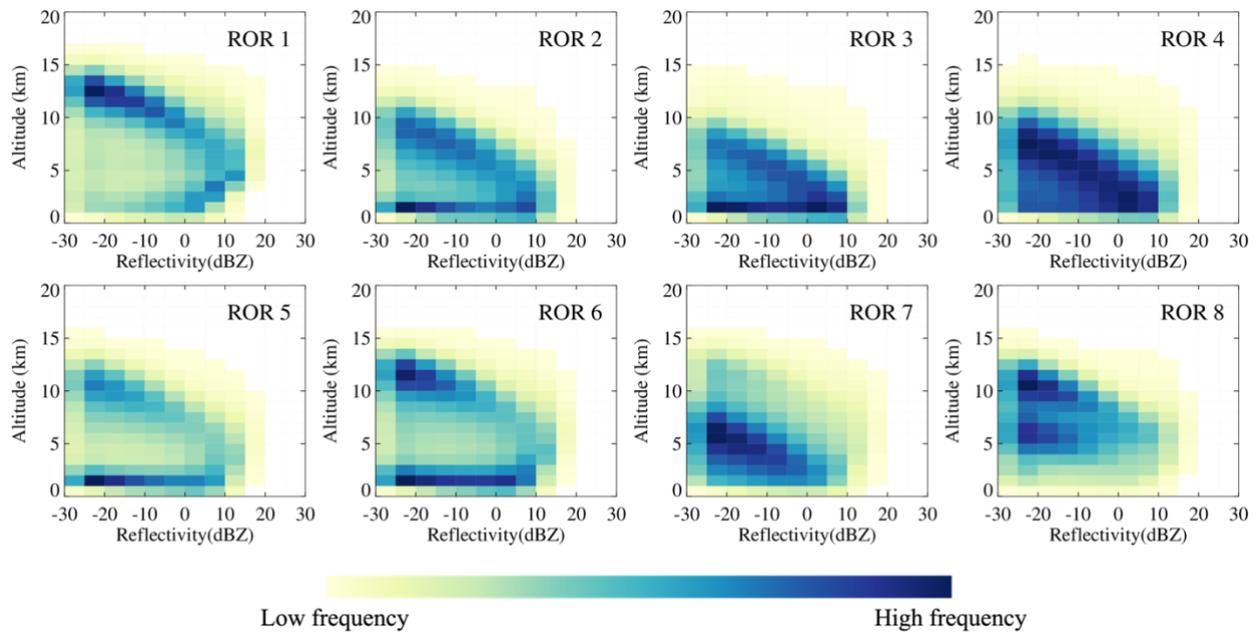


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355 Fig. 9. CloudSat-CALIPSO cloud type occurrence for each ROR created by accumulating
 356 2BCL cloud type occurrences for the 3-month periods used for ROR assignment in one-
 357 degree grid cells. The cloud type occurrences were normalized to add to the total cloud
 358 fraction for each ROR according to the coincident active observations. Cb = cumulonimbus
 359 (deep convection), Ci = cirrus (may also contain cirrocumulus and cirrostratus), As =
 360 altostratus, Ac = altocumulus, Ns = nimbostratus, St = stratus, Sc = stratocumulus, and Cu =
 361 cumulus (represents cumulus congestus and fair weather cumulus). The last bar shows the
 362 global cloud type distribution per 2BCL regardless of ROR.

363 The 2BCL dataset also contains one or more cloud “types” identified within a “ray” of
 364 coincident CALIOP and CPR observations (Sassen et al. 2008). We can thus create statistics
 365 of cloud type occurrence by accumulating the instances at which certain cloud types from
 366 active observations coincide with an ROR occurrence (i.e., all cloud type values within the 1°
 367 degree grid cell over three months). We can then compare to the contribution of each of
 368 2BCL’s seven cloud types (Sc were combined with the rare St) within each ROR to the
 369 global fractional occurrence in the last bar of Fig. 9. For example, even if rare in absolute
 370 terms, ROR1 proportionally contains ~4 times as many Cb’s as the globe as a whole; it also
 371 contains ~2.5 times more cirrus and minute amounts of St+Sc compared to the global
 372 average. The high proportion of Ns in ROR3 is consistent with its occurrence over the
 373 Southern Ocean, where persistent drizzle is common. Anomalously many Cu relative to Sc
 374 and their global populations are found in ROR5 corresponding to the aforementioned marine
 375 shallow convection mixture. About twice as many St+Sc compared to climatology are
 376 encountered in ROR2, while other types occur at rates near or below their global climatology.

377 Cloud types encountered within RORs at such rates confirm our previous interpretations of
378 their nature.



380 Fig. 10. Composite CFADs for each ROR created by accumulating 2B-GEOPROF R05
381 reflectivities for the 3-month periods used for ROR assignment in 1° degree grid cells. The
382 frequency normalization (the details of which are irrelevant) is performed separately for each
383 ROR.

384 To gauge the type and vertical locations of peak occurrence of hydrometeors within
385 RORs, composite CFADs (Contour Frequency by Altitude Diagrams) are constructed from
386 CPR reflectivities (Fig. 10). Our CFADs were created using CloudSat's 2B-GEOPROF.P1
387 R05 product (Marchand et al. 2008) for the period 2007-2020. The 2B-GEOPROF product
388 identifies those levels in the vertical column sampled by the CPR that contain significant
389 radar echo from hydrometeors (rather than noise or radar clutter) and provides an estimate of
390 the radar reflectivity factor for each of these volumes. CFADs are essentially joint histograms
391 showing the frequency of occurrence of CPR reflectivities within predetermined dBZ bins as
392 a function of height. Each CFAD was normalized separately to the number of observations
393 available for the specific ROR. Frequency peaks at low reflectivity values generally indicate
394 the presence of non-precipitating clouds, while high values above 0 dBZ signify the presence
395 of larger hydrometeors and thus substantial likelihood of precipitation. The atmospheric
396 depth throughout which large reflectivities occur indicates how vertically extensive
397 precipitation is. We see that CFADs provide a perspective consistent with the mixture of CRs
398 RORs are composed of, and their CVS and cloud type makeup previously seen in Figs. 8 and
399 9. Distinct (separated) high and low cloud layers can be seen in ROR5 and ROR6, while

400 ROR4 and to a lesser degree ROR3 show vertical development also comprising precipitating
401 parts. The strongest precipitation in the 10-15 dBZ range is unsurprisingly most frequently
402 encountered in the tropical convective ROR1, followed by ROR4 and ROR6.

403 **4. ROR Radiative Impact**

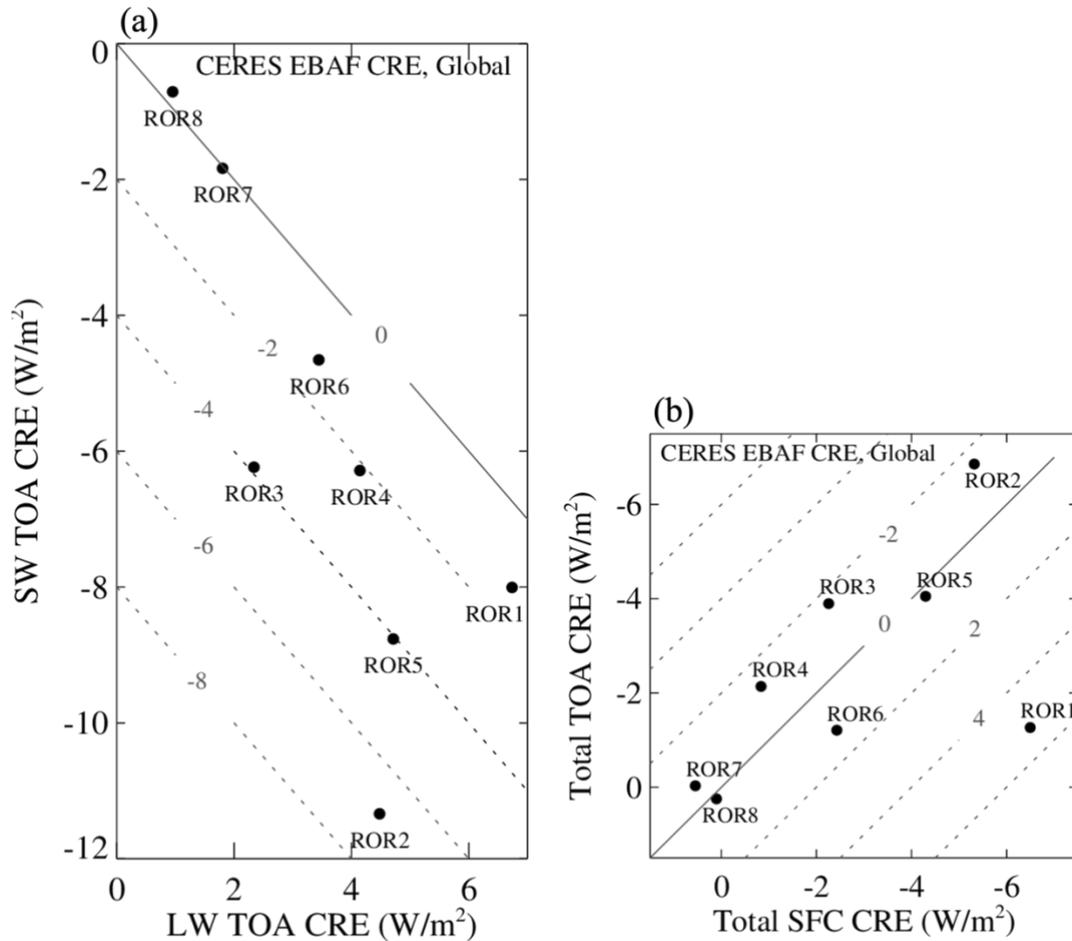
404 We express the radiative impacts associated with RORs in terms of the shortwave (SW)
405 and longwave (LW) cloud radiative effects (CRE) coming from the CERES EBAF dataset
406 (Loeb et al. 2018; Kato et al. 2018). The CERES EBAF product provides a highly stable
407 record of 1° regional, zonal and global monthly mean Top-of-Atmosphere (TOA) and surface
408 (SFC) longwave, and shortwave radiative fluxes, under clear and all-sky conditions. CRE is
409 defined as the difference between all-sky and clear-sky net (down minus up) radiative fluxes
410 at the atmospheric column boundaries (top-of-the-atmosphere TOA and surface SFC).

411 The CERES EBAF Ed4.2 Level 3b Monthly means dataset we use here contains two
412 versions of gridded clear-sky flux, one coming from cloudless portions of CERES footprints
413 and one meant to resemble the diagnostic clear-sky flux of global models where the
414 properties of the cloud-free version of the grid cell are derived from averaging the non-cloud
415 properties of both the cloud-covered and cloud-free portions of the original (cloudy) grid cell.
416 Since comparison with models does not concern us here, we elect to show results using the
417 more physically defined CREs where the clear-sky flux comes from the former method.

418 *a. Mean ROR radiative effects*

419 Fig. 11a contrasts the (RFO-weighted) SW and LW CRE at the TOA of the various
420 RORs. The CREs of each ROR were calculated by averaging the monthly CERES EBAF
421 CREs of all one-degree grid cells where an ROR is encountered. The comparison aims to
422 distinguish RORs in terms of their contribution to planetary radiative cooling. Indeed, none
423 of the RORs with substantial cloudiness contributes to planetary radiative warming, with
424 ROR7 being the sole ROR where radiative warming (from the LW) and cooling (from the
425 SW) cancel each other nearly perfectly. All other RORs contribute to various degrees to
426 cloud-induced planetary radiative cooling, with the strongest coming from ROR2, the 2nd
427 most frequent ROR which has anomalously high proportions of low clouds and large
428 presence at high latitudes. In this cloudiness mixture, the LW CRE contribution is about 2.5
429 times smaller than the SW CRE contribution, resulting in a total (=SW+LW) CRE
430 contribution approaching -7 Wm^{-2} which is about a third of the global enhancement of

431 radiative cooling by clouds (about -20 Wm^{-2}). The next two strongest radiative coolers
 432 (ROR3 and ROR5) each contribute about half of the cooling of ROR2, and their almost equal
 433 values of total CRE contribution come from very different pairings of RFO-weighted SW and
 434 LW CREs.



435

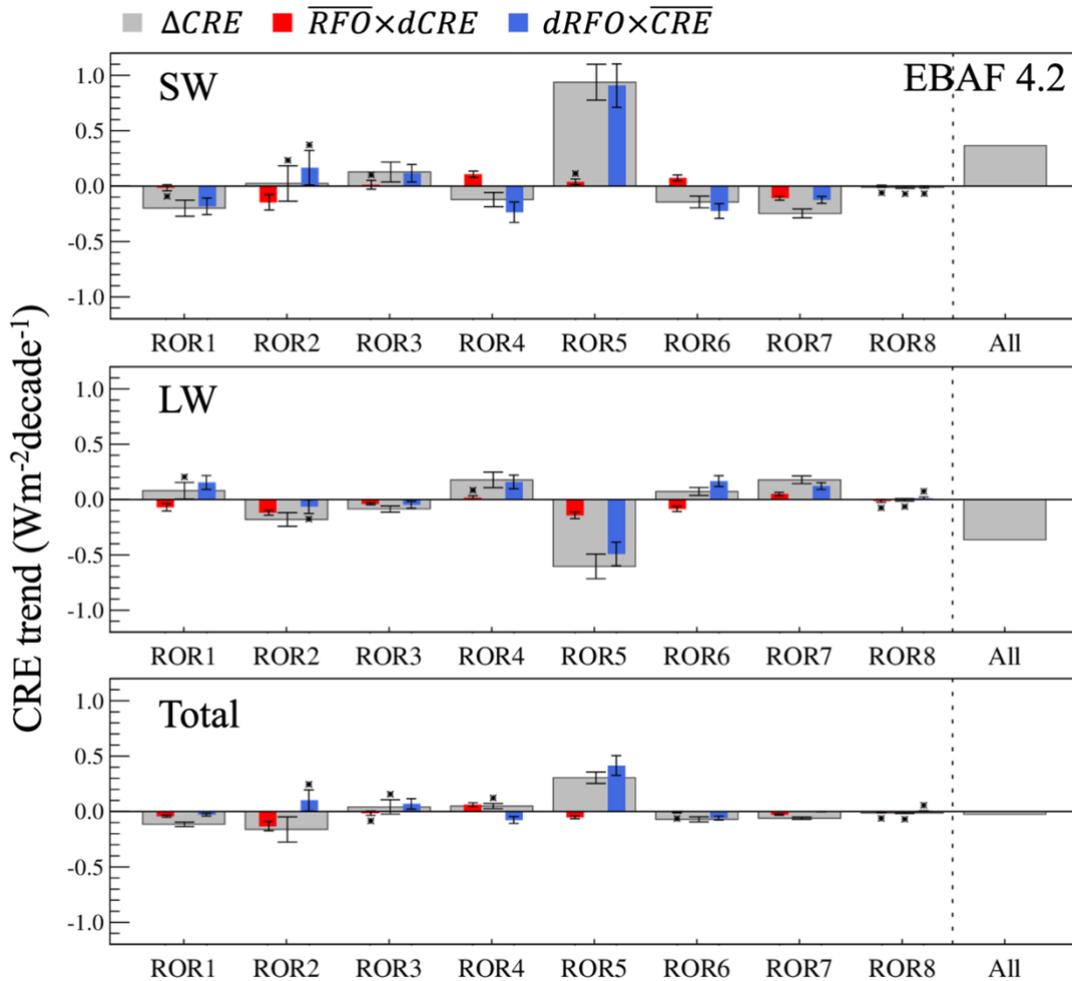
436 Fig. 11. (a) SW TOA CRE vs LW TOA CRE composited for each ROR from CERES-
 437 EBAF; the dashed diagonal lines are isolines of total=SW+LW CRE. (b) Total TOA CRE vs
 438 total SFC CRE for each ROR; the dashed diagonal lines are isolines of total atmospheric
 439 CRE where positive values indicate that clouds contribute to radiative warming of the
 440 atmosphere (vice versa for negative values). All CRE values are weighted by ROR global
 441 RFO.

442 Another cloud radiative impact worth studying is the contribution of RORs to the
 443 reduction or enhancement of atmospheric radiative cooling. This is shown in Fig. 11b where
 444 the total CRE at TOA and SFC is being contrasted. Isolines of constant atmospheric CRE, the
 445 difference between TOA and SFC CRE, are also depicted in the plot, with negative values
 446 indicating cloud radiative cooling contributions to the atmosphere, i.e., enhancements of the

447 clear-sky atmospheric radiative cooling, and isolines of positive values indicating moderating
448 contributions by clouds to the clear-sky radiative cooling. The former effect happens with
449 RORs containing CRs with plentiful low clouds; the atmospheric radiative cooling of (prior
450 versions of) such CRs has indeed been previously shown in Oreopoulos et al. (2016) and
451 Oreopoulos (2022). Three RORs, ROR2, ROR3, and ROR4, contribute nearly equal amounts
452 of about -2 Wm^{-2} additional radiative cooling to the atmosphere. The RORs containing
453 anomalously high proportions of CRs with plentiful high clouds are found to the right of the
454 neutral diagonal; the atmospheric radiative warming effects of (prior versions of) such CRs
455 were also previously recognized in Oreopoulos et al. (2016) and Oreopoulos (2022). The
456 reduction of clear-sky atmospheric radiative cooling, i.e., the warming contribution, is
457 dominated by the predominantly tropical ROR1 encompassing the largest combined amount
458 of stand-alone high and deep convection clouds.

459 *b. CRE Trends*

460 With 20 years of ROR and associated CRE data, we can examine whether any CRE
461 trends (anomalies) can be traced to changes in the properties of specific RORs. Changes in
462 RORs can be viewed as the contribution of two main types of changes: anomalies in ROR
463 populations, namely anomalies relative to the 20-year climatological mean of the ROR RFO,
464 and “internal” anomalies in ROR properties that affect their inherent CRE. We can broadly
465 refer to these as CRE anomalies as “across” and “within” ROR changes, respectively. For the
466 latter category, the major changes in clouds that constitute an ROR are changes in cloud
467 optical thickness, cloud top height, cloud amount (fraction), and changes in cloudless skies
468 within RORs.



469

470 Fig. 12. Trend values (slopes of lines obtained by linearly regressing CRE anomalies
 471 against time) expressed as $Wm^{-2}decade^{-1}$ for each ROR's CRE anomaly component, namely,
 472 CRE anomaly due to internal changes ("within" RORs, red) CRE anomaly, and CRE
 473 anomaly due to RFO changes ("across" RORs, blue); a smaller third term, the covariance
 474 term, is not shown. The broad gray bars are the sum of the three components, i.e., the overall
 475 CRE trend for each ROR. Top panel is for SW CRE, middle panel is for LW CRE, and
 476 bottom panel is for total=SW+LW CRE. The asterisk indicates statistical insignificance per
 477 an F-test. The error bars represent 1.96 times the standard error (95% confidence level) of the
 478 slope of the linear fit calculated from eq. (1).

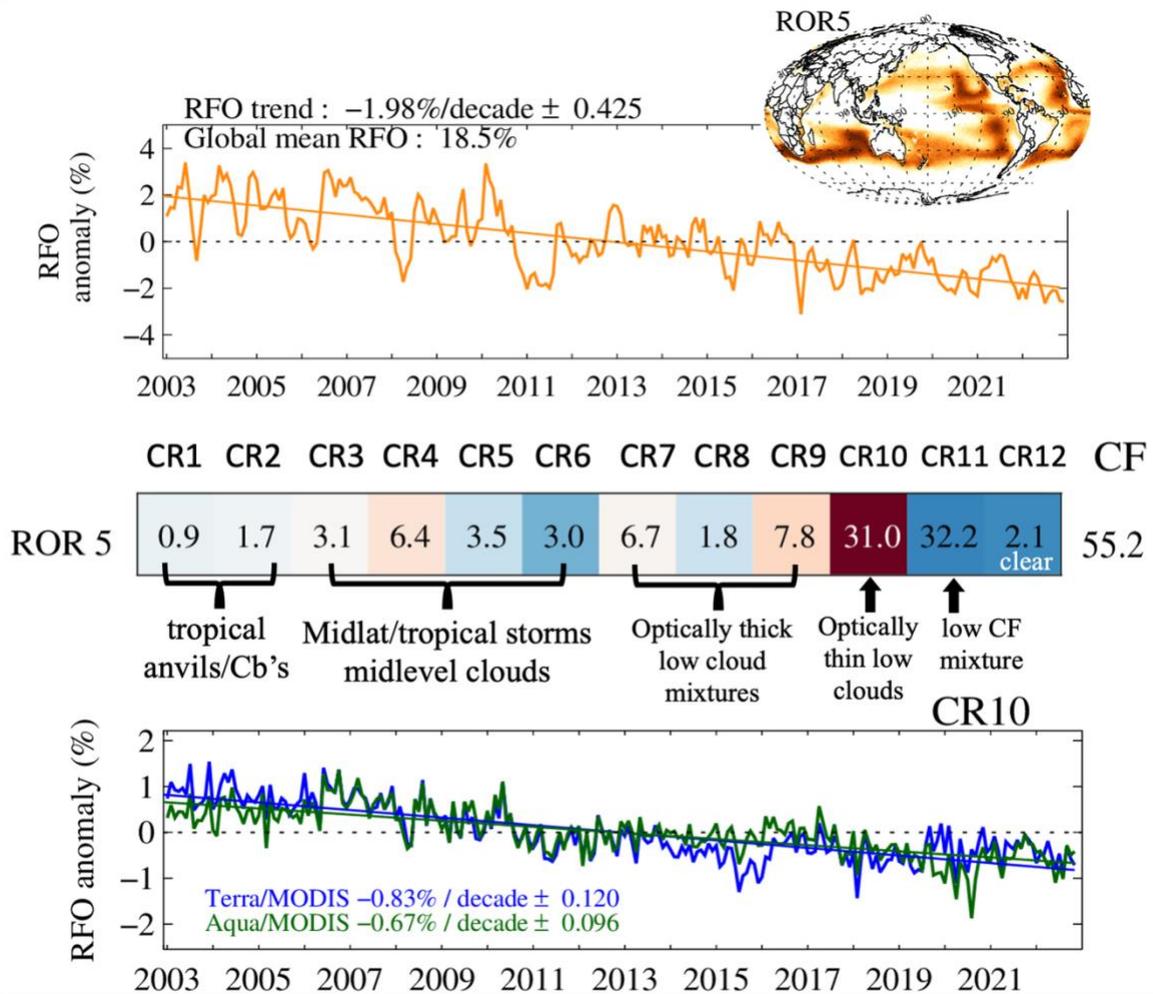
479 Following previous decompositions of observables assigned to CRs (e.g., Tan et al. 2015;
 480 Zelinka et al. 2023; Jin et al. 2023), the CRE anomaly/trend of each ROR is decomposed into
 481 a within, an across, and a covariance term. Fig. 12 shows the 20-year trends of the first two
 482 (largest) terms, expressed as $Wm^{-2}decade^{-1}$, of each ROR's SW, LW and total TOA CRE.
 483 The within-ROR term (red bars) is calculated by multiplying the global multi-annual mean
 484 RFO of the ROR with the trend, from linear regression versus time, of monthly CRE
 485 anomalies with respect to the multi-annual CRE monthly means for that ROR. Similarly, the
 486 across-ROR term (blue bars) is calculated by multiplying the global mean ROR CRE with the

487 ROR's RFO trend calculated from regressing monthly RFO anomalies versus time. The gray
 488 bars show the CRE trends of each ROR calculated directly by regressing the monthly CRE
 489 anomalies of that ROR versus time; we confirmed that the sum of the three trends from
 490 decomposition (including the covariance term not shown) match the directly calculated CRE
 491 trend. We have also verified that adding the individual ROR total CRE trends (rightmost gray
 492 bar) yields global LW and SW CRE trends that largely agree with Raghuraman et al. (2023),
 493 namely opposing $0.36 \text{ Wm}^{-2}\text{decade}^{-1}$ trends (warming for SW and cooling for LW).
 494 Overlaid on each bar of Fig. 12 is the $1.96 \times$ standard error (S.E) (representing the 95%
 495 confidence level) of the linear regression's slope, calculated from Eq. (1), as a measure of its
 496 uncertainty.

$$S.E = \frac{\sqrt{\frac{\sum_{i=1}^N (y_i - \hat{y}_i)^2}{N'}}}{\sqrt{\sum_{i=1}^N (x_i - \bar{x})^2}}$$

497 (1)

499 where N' represents the ROR-specific degrees of freedom, \hat{y}_i represents the estimated y value
 500 from the linear fit, and \bar{x} is the mean value of the x variable. Here, CRE anomalies
 501 correspond to the y variable, while for x , we use normalized time (fraction of decade) ranging
 502 from 0 to 2, to represent 20 years in monthly increments (~ 0.004 in the normalized scale).
 503 The asterisk above or below the bars mark the trends for which we cannot reject at the 95%
 504 confidence level the null hypothesis of zero correlation between CRE anomalies and time; in
 505 other words, these CRE trends are statistically insignificant at the 95% level.



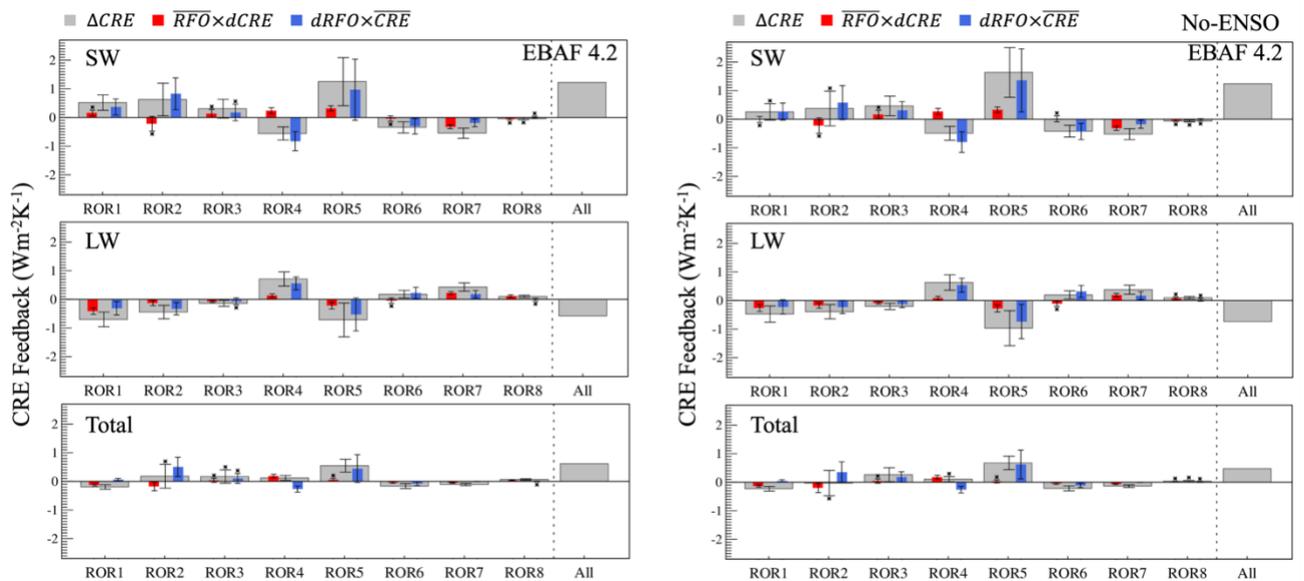
506

507 Fig. 13. Interpretation of ROR5 RFO trend in terms of CR10 RFO trend. The top panel
 508 shows the ROR5 RFO anomaly time series and trend line fit, while the bottom panel shows
 509 CR10 RFO anomaly time series (and trend line fits) separately for Terra and Aqua CR10s.
 510 The middle panel shows the ROR5 centroid.

511 The biggest CRE anomaly is associated with ROR5 and is due to RFO anomalies (large
 512 blue bar in Fig. 12). The time series of ROR5 RFO anomaly is shown in the top panel of Fig.
 513 13. A negative trend in ROR5 RFO is consistent with a weakening of both the SW CRE (less
 514 negative values) and the LW CRE (smaller positive values). ROR5 contains large amounts of
 515 CR10, as shown in the middle panel of Fig. 13, which was previously found in Cho et al.
 516 (2021) to exhibit a decreasing populations over approximately the same period examined here
 517 (an updated CR10 RFO anomaly time series is shown in the bottom panel of Fig. 13). We
 518 thus attribute ROR5's RFO decrease to CR10's RFO decrease, i.e., a decrease in the
 519 population of cloud mixtures dominated by shallow convection (Cho et al. 2021).

520 *c. CRE feedback*

521 In addition to ROR CRE trends, we also calculate ROR “CRE feedback” (Jin et al. 2023)
 522 (strictly not the “cloud feedback” which accounts for only cloud changes while everything
 523 else remains constant), using the same framework as above, namely decomposing each
 524 ROR’s CRE feedback in terms of within-ROR and across-ROR components, similarly to Jin
 525 et al. (2023). Instead of linearly regressing CRE and RFO monthly anomalies against time as
 526 in the CRE trend analysis of the previous subsection, we now regress these anomalies against
 527 the global monthly surface skin temperature (T_s) anomalies of MERRA-2. We present two
 528 versions of CRE feedback, one where we ignore ENSO effects on the T_s anomaly (left
 529 panels) and one where the ENSO effects on T_s are removed (right panels) by using the T_s
 530 anomaly residuals from the linear regression of T_s anomalies against the Niño-3.4 index, as in
 531 Jin et al. (2023). Results for both versions of the CRE feedback are shown in Fig. 14 where
 532 gray bars, blue bars, red bars, and asterisks have the same meaning as in Fig. 12 (but now for
 533 CRE feedback). The error bars once again come from Eq. (1) with x representing the T_s
 534 anomaly for the left panels, and the T_s anomaly adjusted for ENSO effects in the right panels.



535

536 Fig. 14. CRE feedback values (slopes of lines when linearly regressing CRE anomalies
 537 against T_s anomalies) expressed as $\text{Wm}^{-2}\text{K}^{-1}$ for the two major ROR CRE anomaly
 538 components (red and blue bars), and for the overall CRE anomaly (gray bars), as in Figure
 539 12. The CRE feedback on the right panel has ENSO effects on the T_s anomalies removed.

540 The CRE feedback results mirror the CRE trend results indicating positive monthly T_s
 541 anomalies for the period of analysis, i.e., a positive trend of T_s with time. Removing the
 542 ENSO effect has a notable impact on the CRE feedback of ROR5, which becomes stronger
 543 by ~23% overall, indicating a considerable masking effect of ENSO on the feedback of
 544 ROR5 clouds. This SW masking effect was previously noted by Jin et al. (2023) during the

545 DJF season for the SW CRE feedback of their optically thin low cloud regime which roughly
546 corresponds to ROR5 containing plentiful Cho et al. (2021) CR10, as shown earlier. Jin et al.
547 (2023) also noticed that for the same season ENSO amplifies “true” SW CRE feedback for
548 optically thick low clouds. In accordance, we see a reduction in the SW CRE feedback of
549 ROR2, containing high amounts of Cho et al, (2021) CR8 and CR9, when effects of ENSO
550 on T_s are removed.

551 The overall global CRE feedback over our 20-year period when adding ROR feedbacks
552 (sum of grey bars) is $1.2 \text{ Wm}^{-2}\text{K}^{-1}$ for SW and $-0.6 \text{ Wm}^{-2}\text{K}^{-1}$ for LW as depicted by the
553 rightmost gray bar of each panel in Fig. 14. The global CRE feedback values are comparable
554 to several studies using observed clouds and radiation; Raghuraman et al. (2023) found 0.80
555 $\pm 0.52 \text{ Wm}^{-2}\text{K}^{-1}$ for SW and $-0.57 \pm 0.34 \text{ Wm}^{-2}\text{K}^{-1}$ for LW tropical ($30^\circ\text{S} - 30^\circ\text{N}$) cloud
556 feedback, while Chao et al. (2024) found $0.57 \pm 0.27 \text{ Wm}^{-2}\text{K}^{-1}$ for total (LW+SW) cloud
557 feedback. Note that both Raghuraman et al. (2023) and Chao et al. (2024) calculate the cloud
558 feedback adjusted for non-cloud effects, rather than the CRE (difference between all-sky and
559 clear-sky radiative fluxes) feedback as we do. It is quite possible that changes in cloudless
560 skies have a non-negligible contribution to CRE feedback.

561 **5. Discussion and conclusions**

562 This work introduces the concept of Regimes of Regimes (RORs) as a viable option for
563 effectively capturing the main modes of global cloudiness on monthly/seasonal scales. RORs
564 are higher-order cloud classifications describing how semi-instantaneous mesoscale cloud
565 variability is organized at these temporal scales, while also preserving information on cloud
566 opacity and height co-variability at the instrument footprint level as encoded in joint
567 histograms of cloud populations within certain combinations of CTP and COT. RORs thus
568 provide an objective way for identifying areas and times of interest for specific types of
569 cloudiness without relying on arbitrary criteria for region selection. The physical
570 meaningfulness of RORs is supported by the cohesive spatial patterns in which they are
571 organized as well as the emergence of prevalent areas of occurrence over the full 20-year
572 period under examination. Their organization corresponds quite well to distinct dynamical
573 regimes and atmospheric circulation features. Further confirmation of the systematic makeup
574 of the RORs is obtained by examining composites of cloud products from active observations
575 containing information about hydrometeor vertical profiles. Additionally, RORs can be more
576 easily combined with other high-level datasets (Level-3b or Level-4), such as CERES EBAF,

577 monthly precipitation data (e.g., GPCP or IMERG) and climate indices (e.g., Niño-3.4 index),
578 which are incompatible with daily CRs.

579 When paired with radiation data from the CERES EBAF dataset, RORs reveal which
580 cloud systems are most responsible for planetary radiative cooling (ROR2, the 2nd most
581 frequent ROR with high proportions of high latitude low clouds) and counteracting (tropical
582 ROR1 with plentiful high clouds) or enhancing (three low cloud RORs) atmospheric clear-
583 sky radiative cooling. They also reveal what kinds of cloud mixtures contribute the most to
584 LW and SW CRE trends and feedbacks. We find that the ROR containing frequent
585 occurrences of shallow convection (ROR5) has been becoming progressively rarer because
586 the CR that captures the shallow convection has also experienced a shrinking population. In
587 other words, the fact that shallow convection has been decreasing in the last 20 years can now
588 be seen not only in cloud regimes dominated by such clouds, but also in even broader cloud
589 mixtures with other cloud regimes. It appears that contributions to CRE trends from internal
590 changes in the properties of RORs (i.e., “within” ROR changes due to variations in their
591 extent, optical depth, altitude, and the environment they are embedded in) are substantially
592 weaker. Larger internal changes in the CRs themselves may still be taking place, but they can
593 be conceivably compensated by changes within the other CRs that the ROR encompasses,
594 i.e., across CR changes within the ROR (although this remains to be shown).

595 We have also examined the feasibility of producing RORs in an Earth System Model,
596 specifically NASA’s GEOS model (Rienecker et al. 2008). We have confirmed that it is
597 indeed possible and actually quite trivial to generate RORs, but leave this analysis for a
598 separate study. Here we only briefly describe how this can be accomplished. What makes
599 derivation of RORs in climate models possible is the capability of classifying model clouds
600 according to observed CRs (Williams and Webb 2009) thanks to subgrid cloudiness
601 expressed in terms of the aforementioned CTP-COT joint histograms (JHs). Such JHs can be
602 produced by COSP’s ISCCP and MODIS simulators (Bodas-Salcedo et al. 2011). The
603 model’s CR occurrences are generated by assigning the model JH to the closest observed CR
604 (Williams and Webb 2009), rather than performing new clustering on the model’s JH as has
605 been suggested earlier by Williams and Tselioudis (2007). Once a model’s cloudiness is
606 expressed in terms of CRs, RORs can also be produced. There are two avenues to accomplish
607 this. One is to perform clustering of the CR frequency of occurrence (FOC) histograms, as in
608 observations. Another is to compare the model’s CR FOC histograms to the centroids of the

609 observed RORs, and assign the model histogram to the ROR with the closest matching
610 centroid.

611 The overall philosophy of clustering occurrences of distinct cloud classes over a period of
612 time to create a higher-level classification that conveys how they mix can be broadened to
613 combine datasets of different provenance. For example, one could attempt to cluster FOC
614 histograms of spatiotemporally matched classes of passive and active origin, or of combined
615 FOC histograms of MODIS CRs (Cho et al. 2021) and ISCCP WSs (Tselioudis et al. 2021).
616 We propose that the utility of the broad concept of clustering FOC histograms of different
617 (co-varying) cloud properties, cloud mixtures, or other cloud entities continues to be explored
618 in future work. Novel ways of organizing (and effectively compressing) cloud information
619 can be potentially as useful for understanding climatic trends as the fundamental
620 understanding of cloud processes and patterns at high resolutions.

621

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627

628 *Data Availability Statement.*

629 MODIS equal-angle daily CRs are available at
630 <https://doi.org/10.5067/MEASURES/MODISCR/EQANGD/DATA301>, MODIS RORs are
631 available at <https://doi.org/10.5281/zenodo.11099765>. The Clustering code used to derive the
632 ROR can be found at <https://doi.org/10.5281/zenodo.13351479>. The 2B-CLDCLASS-LIDAR
633 and 2B-GEOPROF datasets are available at [http:// www.cloudsat.cira.colostate.edu/](http://www.cloudsat.cira.colostate.edu/) and
634 CERES EBAF Level 3b dataset is available at <https://ceres.larc.nasa.gov/data/>. The Surface
635 Skin Temperature (T_s) and omega at 500 hPa from MERRA-2 are available at
636 <https://goldsmr4.gesdisc.eosdis.nasa.gov/data/MERRA2/M2T1NXSLV.5.12.4/>

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