

1 Roberto Aguilar Ayala¹², Morgan A. Crabtree², Matthew L. Hancock², Alexander Jarnot³, Janine E.
2 Captain², Julie Kleinhenz⁴, Erin Rezich⁵, Kris Zacny⁶, Vincent Vendiola⁶, Zack Mank, Kenneth C. Wright⁷,
3 Jaime L. Winfield⁷, Thomas Orlando⁸, Brant Jones⁸, Phil Metzger⁹, Jaqueline W. Quinn²

4

5 **Mass Spectrometry Monitoring of Drilling Operations in Water-Doped** 6 **Lunar Simulant under Cryogenic Conditions**

7

8 **Abstract:**

9 The utilization of analytical instruments for the detection and quantification of volatiles on the Moon is
10 crucial for the development of in situ resource utilization technology. A test campaign under cryogenic
11 conditions aimed to simulate the cold lunar environment of permanently shadowed regions (PSRs). The
12 test campaign was performed in a Thermal Vacuum Chamber (TVAC) equipped with a bin of regolith
13 simulant doped at three increasing water concentrations delineated at specific depths (Dry, 2.5, and 5
14 wt%, respectively). The average operating pressure of the chamber was 1.6×10^{-6} Torr with an average
15 regolith temperature of -178°C . Four holes were drilled into the regolith simulant bin using The Regolith
16 and Ice Drill for Exploring New Terrain (TRIDENT) drill. The resulting cuttings-piles and their volatile activity
17 were monitored via the Mass Spectrometer observing lunar operations (MSolo) instrument to determine
18 the extent to which a mass spectrometer can monitor the release of volatiles during drilling activities in
19 lunar-like environments. The results of this campaign helped to plan for the operation and subsequent
20 data interpretation of the Polar Resources Ice Mining Experiment-1 (PRIME-1) mission on the Moon.

21

¹ Corresponding author roberto.aguilar-ayala@nasa.gov

² NASA Kennedy Space Center, FL 32899

³ NASA Langley Research Center, Hampton, VA 23666

⁴ NASA Johnson Space Center, Houston, TX 77058

⁵ NASA Glenn Research Center, Cleveland, OH 44135

⁶ Honeybee Robotics, Altadena, CA 91001

⁷ IMI Adaptas, Palmer, MA 01069

⁸ Georgia Institute of Technology, Atlanta, GA 30332

⁹ University of Central Florida, Orlando, FL 32816

22 **Table of Figures**

23

24

25 Figure 1: Cross Beam ion source configuration of MSolo (A), showing the two linear filaments (B) in
26 correlation to the anode cage and the correlation to its field-of-view (C) used to determine optimized
27 detection of regolith cuttings pile volatile release. 5

28 Figure 2: TRIDENT Drill 6

29 Figure 3: Experimental set up in VF13 chamber: closed VF13 chamber (A), CAD drawing of experimental
30 set up showing MSolo’s location relative to the drill and regolith bin (B), cuttings cone (C). 8

31 Figure 4: Regolith bin testing conditions: diagram depicting water-doping concentration and bite
32 number by depth (A), diagram showing hole drilling locations accommodating 4 drill holes in regolith bin
33 (B). 9

34 Figure 5: MSolo detection response during drill operations (top), zoomed sections for each region
35 (bottom). 11

36 Figure 6: Individual mass spectra (top) during a peak signal (orange square) of water from drilling
37 operations (mass trace - bottom). 12

38 Figure 7: Evolution of water signal detected by MSolo throughout drilling operations. 14

39 Figure 8: Video captures of cuttings pile progression for hole 1, each different bite and its resulting pile is
40 shown. 15

41 Figure 9: MSolo peak raw intensity correlation to wt% values. 17

42 Figure 10: Water fluence per cm² (bar graph) and the water signal (*m/z* 18) intensity (line graph) for
43 holes 1-4. 19

44 Figure 11: Water fluence per cm² versus *m/z* 18 intensity shows a linear relationship between the two
45 parameters. The solid line is the linear best-fit line. 20

46 Figure 12: Linear regression fitting an approximation for water fluence from the regolith piles to the
47 partial pressures of water observed during GRC testing. 21

48 Figure 13: Predictions for the wt% water for each bite generated from hole 1 signal inputs and the
49 MATLAB algorithm 23

50 Figure 14: Predictions for the wt% water for each bite generated from sample signal inputs, lunar
51 temperature, and the MATLAB algorithm. 24

52

54 **Introduction:**

55 Missions to the moon including Clementine and Lunar Prospector indicate the presence of hydrogen-
56 bearing volatiles locked within the lunar surface (Nozette *et al.*, 1996; Feldman *et al.*, 1998). However, it
57 was not until the Lunar Crater Observation and Sensing Satellite (LCROSS) mission that the identification
58 and measurement of several volatile species was verified in a permanently shadowed region (PSR) of the
59 Cabeus crater: water ice was confirmed within its impact zone at approximately $5.6 \pm 2.9\%$ by mass
60 (Colaprete *et al.*, 2010). Knowing now that a significant concentration of volatiles is trapped within the
61 cold environments of the lunar south pole, key information such as localization of resources, the actual
62 concentrations, and the volatile distribution based on depth are all still to be determined. A fundamental
63 component for long-term space exploration requires the extraction and utilization of resources within the
64 sites of exploration (Sanders and Larson, 2013). In-Situ Resource Utilization (ISRU), as the name suggests
65 aims to develop technology to extract and use the available resources located within the surface of the
66 Moon, thus providing a foundation for human habitation and exploration.

67 A clear aspiration for future ISRU missions is to fully understand the available resources on the Moon.
68 Prospecting missions on the lunar surface that involve drilling operations along with analytical instruments
69 for the detection and quantification of volatiles on the Moon will be crucial to understanding the volatile
70 distribution on and below its surface. During a Thermal Vacuum Chamber (TVAC) test campaign in 2019,
71 a drill known as TRIDENT (The Regolith and Ice Drill for Exploration of New Terrains) was coupled to a test
72 unit of MSolo (Mass Spectrometer observing lunar operations) to monitor water ice release from regolith
73 simulant during drilling operations in simulated lunar environments. The results of this test campaign
74 discussed in this report show a relationship between water ice concentration and the signal detected by
75 MSolo during the drilling operations that could be leveraged for the PRIME-1 mission in addition to future
76 lunar missions.

77

78 **1. Experimental Set Up**

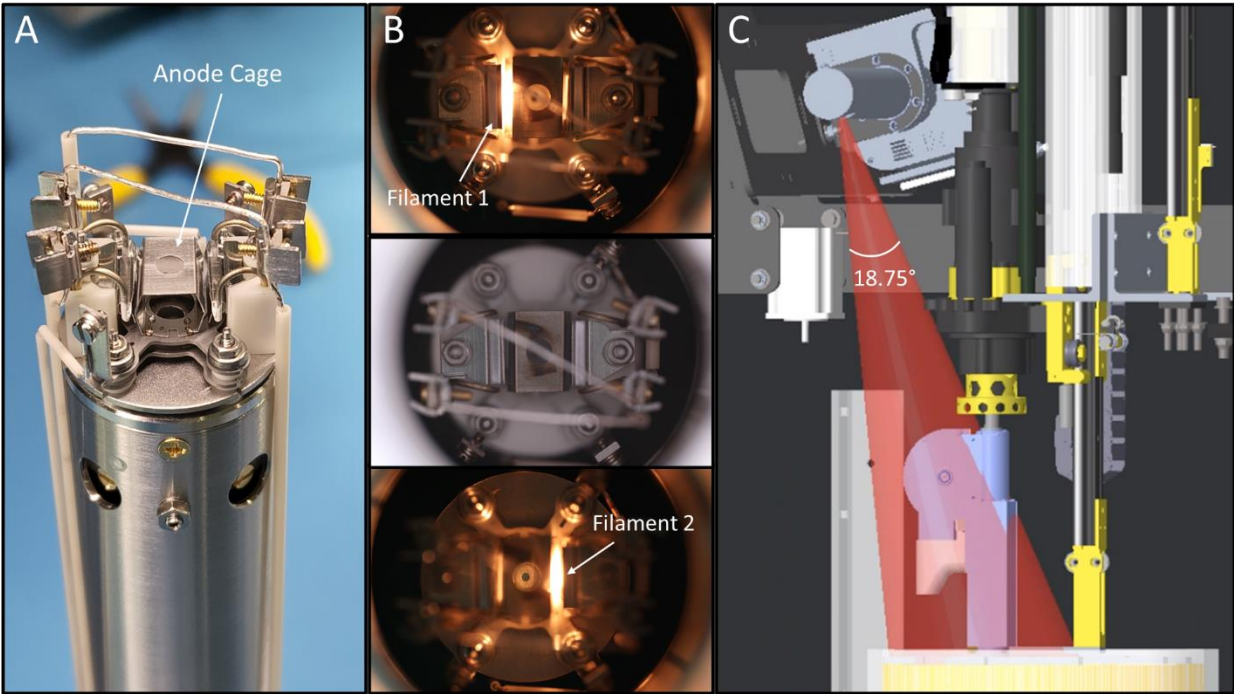
79

80 **1.1. MSolo Mass Spectrometer**

81 The MSolo instrument is a modified commercial off-the-shelf (COTS) mass spectrometer based on
82 INFICON's Transpector® MPH residual gas analyzer. This mass spectrometer was ruggedized for space
83 applications by replacing electrical components with improved, radiation-resistant components and

84 incorporating structural and thermal enhancements necessary for withstanding vibrations experienced
85 during launch and in the harshness of the space environment. All modifications were implemented
86 without changing the COTS-like operation. MSolo is capable of an operational scanning range of mass-to-
87 charge ratios (m/z) 1-100. The operating mass range allows for the detection of volatile species (for this
88 test, most notably water, hydrogen, and nitrogen), their relative intensities, and isotopic ratios. MSolo
89 utilizes a crossbeam ion source (XB) configuration, which consists of a redundant set of linear filaments
90 that ionize gas-phase molecules traveling perpendicular to the direction of the electron beam and
91 generates a small current signal as an output. The geometric configuration of the crossbeam ion source
92 design allows for a higher density of electrons to be introduced into the ionization region, thus ionizing
93 molecules traveling axially through the anode cage without interacting with its surfaces. This configuration
94 is exceptionally useful for the detection of released volatiles from a point source whilst minimizing
95 background effects. Figure 1 illustrates the crossbeam configuration in relation to the two available
96 filaments. MSolo works by introducing the generated ions into the quadrupole, where they are filtered
97 based on their m/z and subsequently detected by a Faraday cup, or an electron multiplier, if higher
98 sensitivity is required. For these tests, the instrument operated across the full scanning range, and while
99 the dwell time for all masses was set to 32 ms, the dwell time for the water signal (m/z 18) was increased
100 to 256 ms to target the release of water from drilling activities. The electron multiplier was also calibrated
101 so that the nitrogen signal (m/z 28) peak (the most abundant background peak in the chamber) reached
102 a target signal of approximately 6×10^{-8} A, thus maximizing the detectable signal of the entire mass range
103 without saturating the detector with the background signal. The baseline signal of the instrument for this
104 test was 3.71×10^{-13} A.

105



106

107 Figure 1: Cross Beam ion source configuration of MSolo (A), showing the two linear filaments (B) in
 108 correlation to the anode cage and the correlation to its field-of-view (C) used to determine optimized
 109 detection of regolith cuttings pile volatile release.

110

111 MSolo was mounted on a bracket attached to a translator frame connected to TRIDENT, allowing
 112 it to maintain a constant distance from the drill hole/pile even as the drill moved between locations.
 113 MSolo's location relative to the bracket and drill was determined based on the alignment necessary for
 114 the drill cuttings pile to be within the instrument's field of view (FOV), an 18.75-degree cone. The
 115 molecules within this FOV have a direct path from the pile surface to the detector. A visual representation
 116 of MSolo's location and FOV is shown in Figure 1C.

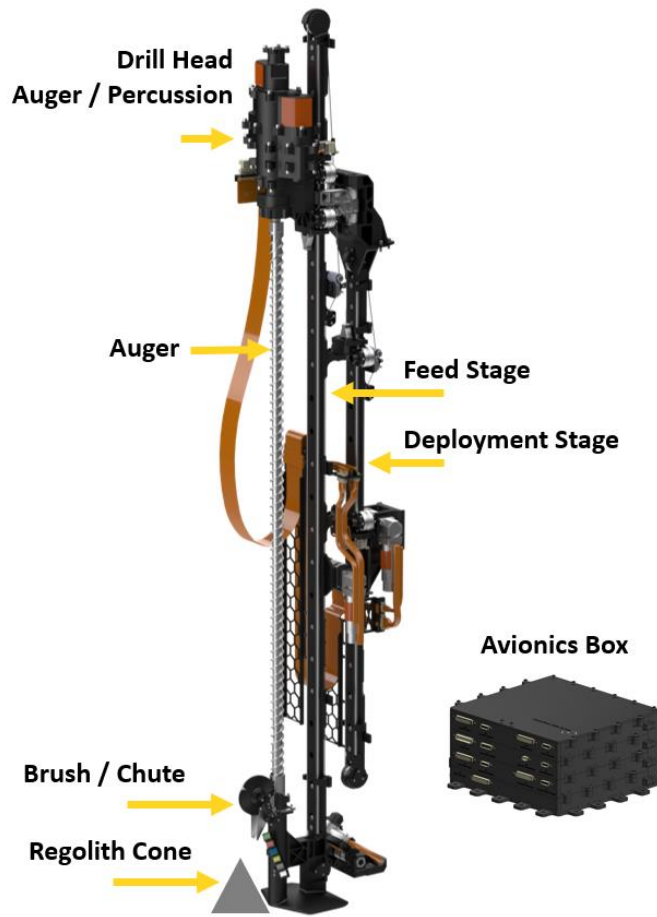
117

118 **1.2. TRIDENT Drill**

119 TRIDENT is a drilling system designed and built by Honeybee Robotics for the acquisition of
 120 volatile-rich samples from planetary subsurfaces. The drill was incrementally developed for over two
 121 decades and went through extensive technological development with validation tests performed in
 122 planetary analog locations such as Antarctica, Arctic, and Atacama, as well as in TVACs. TRIDENT's main
 123 features were developed from the Icebreaker-3 mission configuration and the drill was mass-optimized

124 during the Life in the Atacama (LITA) project (Paulsen *et al.*, 2018; Zacny *et al.*, 2023). The TRIDENT drill is
125 shown in Figure 2.

126



127

128

Figure 2: TRIDENT Drill

129

130 TRIDENT features a rotary-percussive drill head driven by two separate actuators. Splitting
131 rotation and percussion allows the drill to use three modes of operation: rotation-only, percussion-only,
132 and rotation-percussion. This flexibility helps preserve drilling energy as not all drilling operations require
133 use of percussion. It also allows use of percussion-only operation in certain circumstances including
134 clearing cuttings from auger flutes and seismic sensing. The drill has an 1.2 m auger with a full-faced
135 carbide bit at its end, allowing for a full one-meter drill depth. The “bite” section is the lower 15 cm of the
136 auger. This region has deep, low-pitched flutes, making it well-suited to cuttings capture. The upper
137 section of the auger has shallower, high-pitched flutes that are better for cuttings conveyance. The auger

138 drills and captures cuttings in 10 cm bites – after 10 cm of regolith is penetrated, the auger is lifted from
139 the resulting hole, and the captured cuttings are brushed off the auger onto the surface forming a pile via
140 a brushing subsystem. The auger is then lowered back into the hole and the process is repeated, resulting
141 in a conical cuttings pile that grows with each bite.

142 The bite sampling approach has numerous advantages. Subsurface stratigraphy is preserved in at
143 least 10 cm sections, potentially on the order of millimeters under ideal conditions. Lifting the auger every
144 10 cm lowers the energy required for drilling, as at a depth of 1 m, the power required to move the total
145 mass of cuttings from the entire hole to the surface could be significantly greater than power required to
146 drill. Finally, pauses in drilling allow the drill bit and auger to cool off between bites, thus minimizing
147 volatile gas impacts.

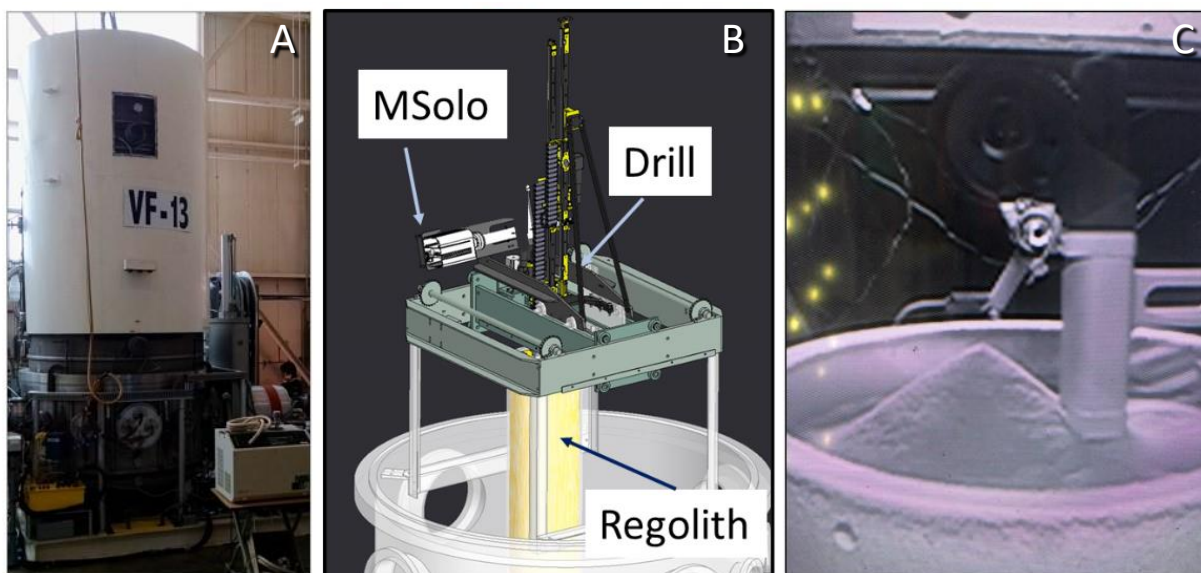
148

149 **1.3. Glenn Research Center VF-13 Chamber**

150 The VF-13 is a cylindrical chamber with an internal volume of 6.35 m³ (Figure 3A) located at NASA's
151 Glenn Research Center (GRC). The chamber includes a removable lid measuring 2.52 m in height and 1.5
152 m in diameter and a fixed base with a depth of 1.08 meters that accommodates all the electrical,
153 mechanical, and gas feedthroughs. A removable cold wall is placed inside the lid and a cold shell clamped
154 around the regolith bin sits in the fixed base of the chamber. The cold wall and regolith bin can be
155 independently temperature controlled from room temperature down to approximately 90 K using liquid
156 nitrogen, and the average operating chamber pressure for this test campaign was approximately 1.6x10⁻⁶
157 Torr.

158 All test article hardware (e.g. TRIDENT and MSolo) was mounted on a trolley rig that can be moved
159 in two dimensions using remotely-actuated stepper motors and chain drives, allowing access to the entire
160 surface of the regolith bin (Figure 3B). As such, multiple drill holes were drilled without needing to break
161 vacuum to replace the regolith sample or change location. The MSolo instrument was integrated into the
162 trolley system using a custom bracket, permitting the instrument and drill to move simultaneously and to
163 maintain a constant distance (19.32 cm) from the plane of the regolith with a high degree of rotational
164 freedom.

165



166

167 Figure 3: Experimental set up in VF13 chamber: closed VF13 chamber (A), CAD drawing of experimental
 168 set up showing MSolo's location relative to the drill and regolith bin (B), cuttings cone (C).

169

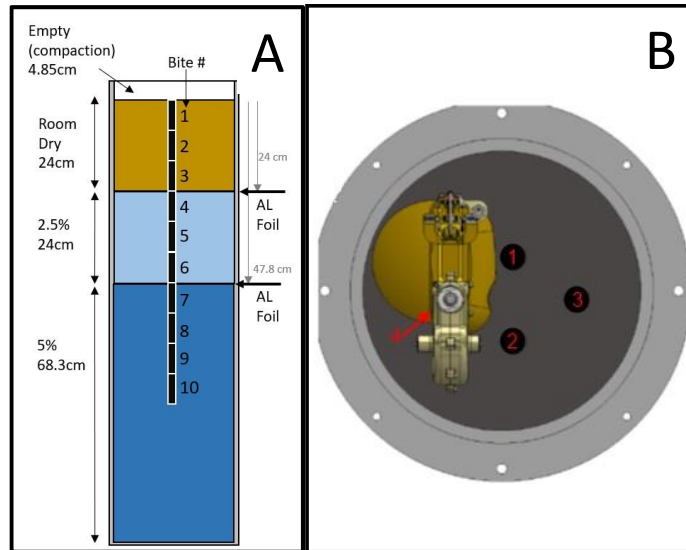
170 Regolith simulant sample was contained within a 1.2 m tall cylindrical aluminum bin with an inner
 171 diameter of 0.278 m. The bin accommodates up to four drill holes (4B). Soil temperatures were recorded
 172 via three ports at various heights along the regolith bin equipped with three to five type-T thermocouples
 173 embedded in the regolith at different radial positions.

174

175 **1.4. Test Conditions**

176 After the vacuum chamber was pumped down and cooled, the TRIDENT drill took multiple bites
 177 from a single drill hole while the MSolo instrument scanned continuously to detect volatile species being
 178 emitted from the regolith pile. This bite approach was repeated in four trolley positions so that four drill
 179 holes were analyzed within the regolith bin (Figure 4B).

180



181

182 Figure 4: Regolith bin testing conditions: diagram depicting water-doping concentration and bite
 183 number by depth (A), diagram showing hole drilling locations accommodating 4 drill holes in regolith bin
 184 (B).

185

186 The regolith simulant bin used in this test was prepared using stratified moisture layers in which
 187 different weight percent (wt%) water-doped regolith simulants are layered one over the other, with a thin
 188 sheet of aluminum foil separating the layers to retain water content during atmospheric (unfrozen)
 189 conditions (Figure 4A). The regolith bin consisted of approximately 100 kg of water-doped lunar highlands
 190 simulant (NU-LHT-3M). The layers were vibrationally compacted in 20 cm increments to a density of
 191 approximately 1.6 g/cm^3 . The water concentration increased with depth to simulate a stratification in
 192 water concentration versus drilled depth. The three moisture levels within the bin were “dry,” 2.5 wt%,
 193 and 5 wt%. The dry moisture level was measured to contain approximately 0.2 wt% water due to
 194 adsorption from the atmosphere (Kleinhenz and Linne, 2013; Kleinhenz, 2014). These moisture levels
 195 were based on the expected water concentration as detected by LCROSS (Colaprete *et al.*, 2010). A brief
 196 description of all test conditions can be reviewed in Table 1.

197 Table 1: Summary of test conditions for drill sampling of regolith simulant as discussed above and shown
198 in Figure 4.

Test Conditions	
Average Regolith Temperature	-178°C
Average Chamber Pressure	1.6×10^{-6} Torr
Bin Preparation	Top (0-24 cm): Dry Middle (24-48 cm): 2.5 wt% Bottom (48-117 cm): 5 wt%
Bite Depth	8 cm
Hole 1 Approach	All bites: 1 h hold
Hole 2 Approach	Bite 1: 5 min hold Bites 2-10: 10 min hold
Hole 3 Approach	Bite 1: 5 min hold Bites 4-6: 10 min hold Bites 7-10: 15 min hold
Hole 4 Approach	Bite 1: 5 min hold Bites 4-6: 10 min hold Bites 7-10: 15 min hold

199

200

201

202 2. Results and Discussion:

203

204 2.1. MSolo Responses to Drilling Operations

205 After each drill bite, all drill activities were paused while MSolo measured the evolution of
206 volatiles from the pile. After the water signal (m/z 18) returned to baseline or a near-baseline value, the
207 drill would then proceed to the next bite. While the focus of the test was to observe the water signal (m/z
208 18), other volatiles were also used to guide the MSolo team during drilling events. One notable example
209 is the observed changes in the hydrogen signal (m/z 2), which was found to correlate with drilling. The
210 hydrogen signal increased as the drill began to rotate, decayed as the drill retracted, and returned to
211 baseline when the drill stopped. The hydrogen signal dynamics were unique from the evolved water signal
212 dynamics, visible in Figure 5Error! Reference source not found.. It was later determined that the
213 hydrogen signal activity was caused by the release of hydrogen from metal-on-metal interactions within
214 the drill, as the hydrogen signal (m/z 2) only increased in activity when the drill was in motion.

215
216
217
218
219

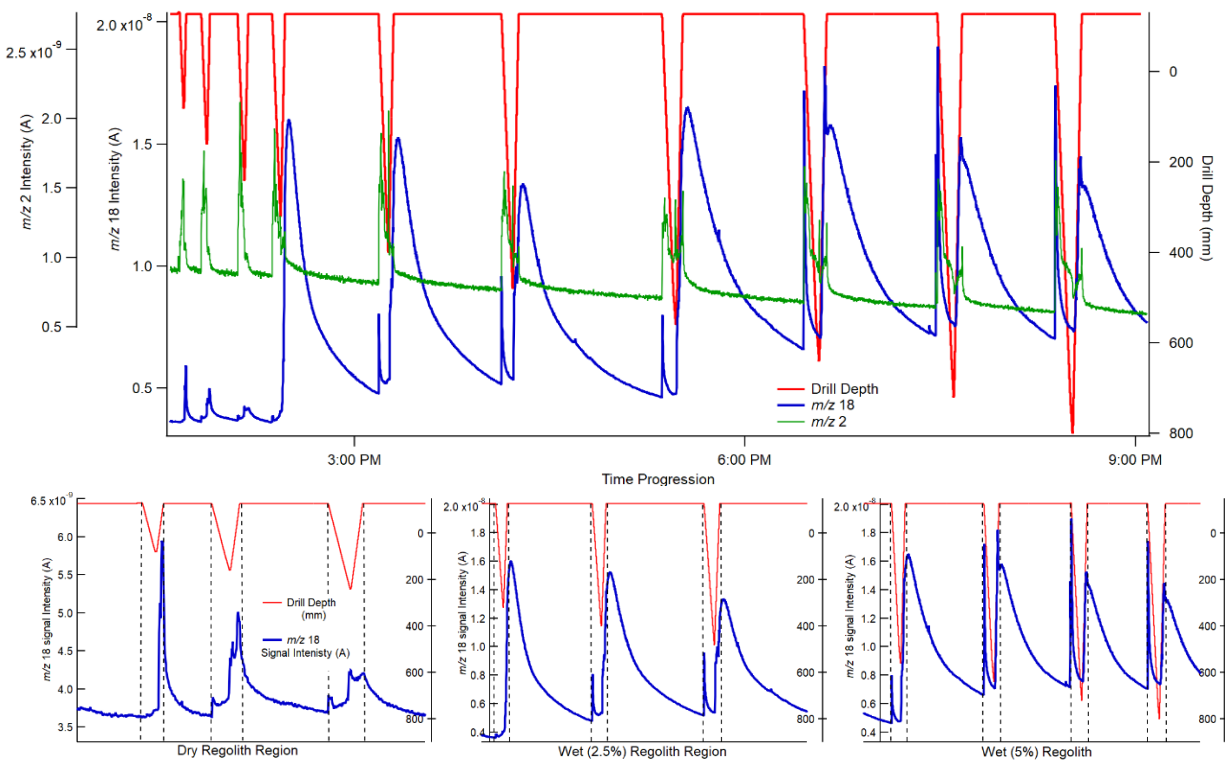
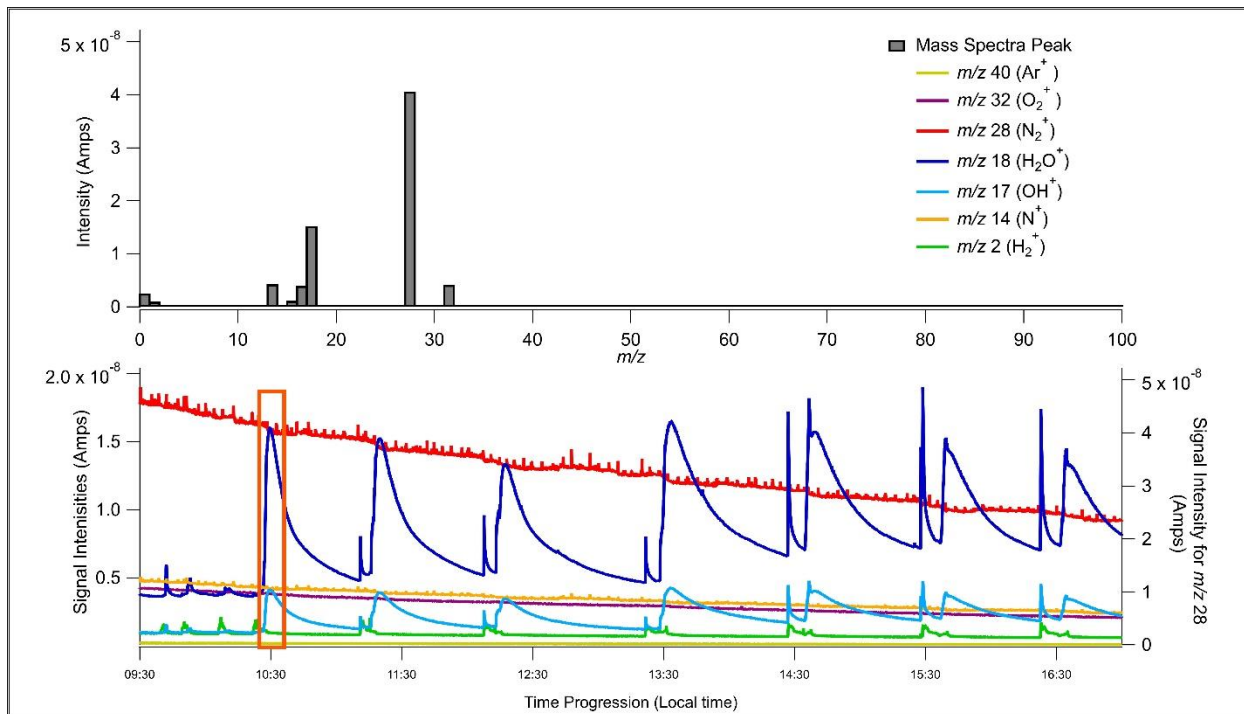


Figure 5: MSolo detection response during drill operations (top), zoomed sections for each region (bottom).

220
221
222
223
224
225
226
227

The signals measured by MSolo as a result of drilling the first hole are shown in Figure 5, where a spike in the water signal (m/z 18) was observed when regolith simulant material was extracted from the hole and deposited onto the surface, followed by a signal decay as the volatile activity reduced and therefore signal returns to a near-baseline value. This signal profile was even observed in the “dry” regolith region, demonstrating that the very low water concentration of 0.2 wt% could still be detected. There were also additional water signal spikes during drill movement activities, such as when the drill was lowered back into the drill hole or retracted out of the hole. This is visible for bites 2-10, but not for bite

228 1. These spikes in the water signal (m/z 18) likely come from the agitation of a previous drilling sample
229 retained on the auger and/or from agitation of the cuttings pile. This theory is further supported by a
230 similar water spike observed when the drill was lowered and retracted in a fashion so as to remove any
231 remaining material from the drill brush. Figure 6 shows that even during dynamic environmental
232 circumstances including changing chamber pressure and a dominant, progressively decreasing nitrogen
233 signal (m/z 28), MSolo has the capacity to observe the evolution of lower-signal volatiles.



234
235 Figure 6: Individual mass spectra (top) during a peak signal (orange square) of water from drilling
236 operations (mass trace - bottom).

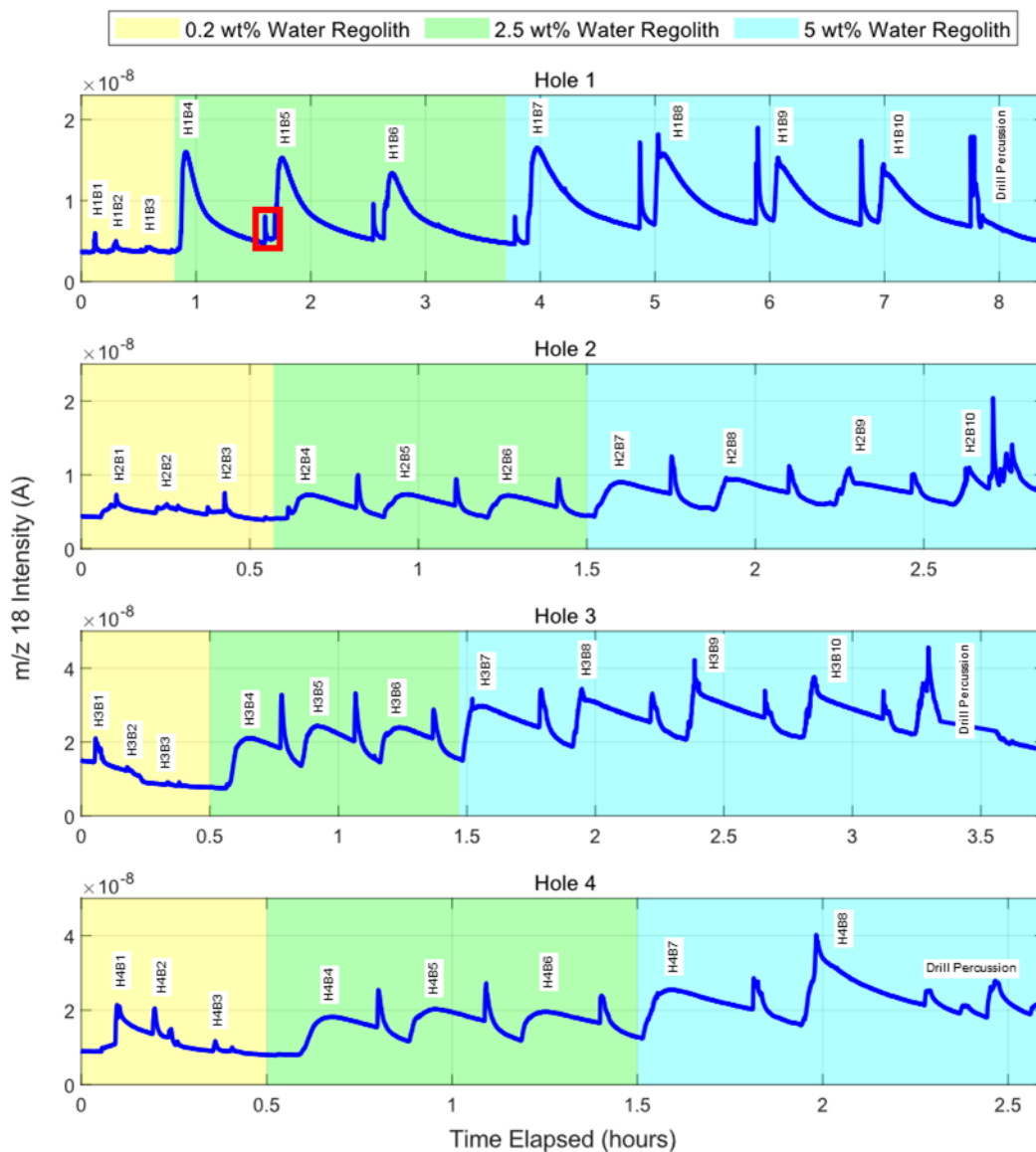
237
238
239 The original plan for the test campaign was to permit the water signal to decay back to its baseline
240 value prior to restarting drilling rotations, however, it was determined post-hole 1 that this would not be
241 feasible, as the time it would take to return to baseline exceeded the total test duration. Table 1 describes
242 that the drill position was maintained and drill activities paused for one hour between bites for hole 1,
243 however, the pause was decreased to no more than 15 minutes for holes 2-4 in order to focus on capturing
244 the shape of the initial evolution of the water peak observed during hole 1 drilling.

245

246 **2.2. MSolo Measurements**

247 As stated previously, three moisture levels were selected for the test: dry (< 0.2 wt%), 2.5 wt%,
248 and 5 wt%. Figure 7 shows the evolution of the water signal (m/z 18) observed by MSolo throughout the
249 drilling process. The water concentrations correlating with the drilling depth are depicted with different
250 colored shading. Each bite is labeled on the graphs with the hole and bite number – HXBY – where X is
251 the hole number and Y is the bite number. The label location is close to the estimated maximum result
252 from each bite, and it should be noted that although some peaks are of higher intensities, these are due
253 to secondary events such as regolith coming off the drill brush. Other events such as drill percussion and
254 drill rotations are annotated for holes 1, 3, and 4, where a corresponding water signal increase was
255 subsequently detected. Agitation of regolith material remaining within the drill, on the brush, and/or the
256 regolith pile was determined to have caused these volatile releases. Furthermore, spikes in the water
257 signal for all holes can be seen just prior to the secondary/decay spike, highlighted by a red square in
258 Figure 7 for H1B5. This phenomenon is attributed to the release of residual regolith sample remaining on
259 the drill bit upon drill rotation start up. While the deposit of this residual material resulted in the observe

260 sharp signal increases, the signal decayed quickly and did not disturb the signal trace from the new
 261 material brought to the surface during the subsequent bite's drilling operation.

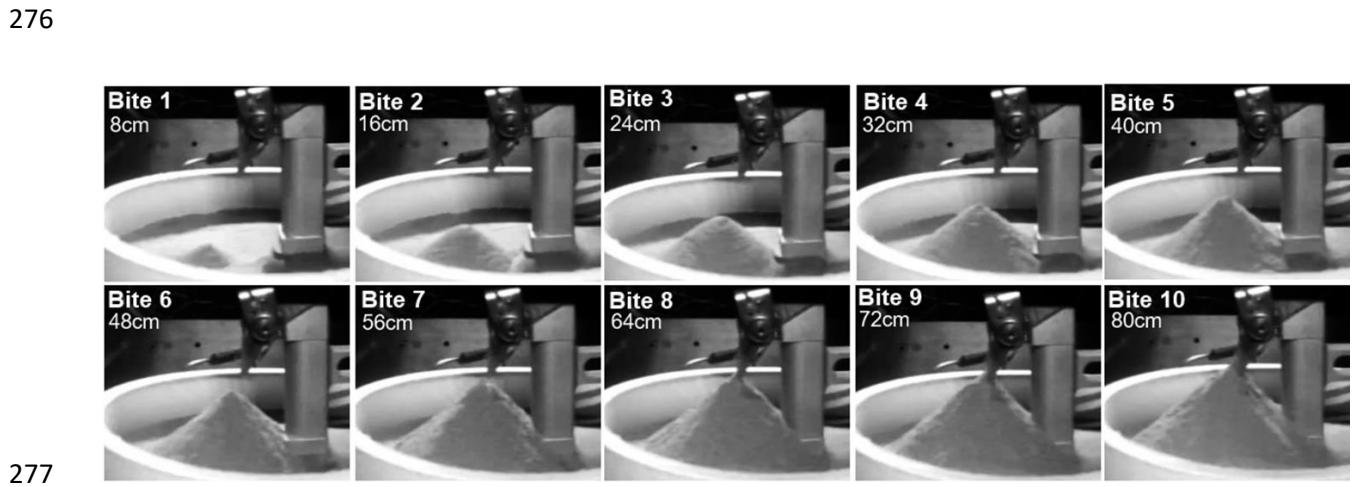


262 Figure 7: Evolution of water signal detected by MSolo throughout drilling operations.

263

264 Hole 1, bites 1-3 exhibit the lowest maximum signal intensities for the entire testing series. The
 265 max peak signal intensities for H1B1-H1B3 correspond to 5.95×10^{-9} A, 5.0×10^{-9} A, and 4.2×10^{-9} A
 266 respectively, and while low in intensity when compared to the rest of the evolved peaks, they are well
 267 above the baseline signal of the instrument (3.71×10^{-13} A) even while the pressure of the chamber is in

268 the low 10^{-6} Torr range. In contrast, H1B4-H1B10 show an increase in the water signal (m/z 18) during the
269 drill activities, corresponding to the higher concentrations of water present in the regolith. There is also
270 a distinctly higher water intensity for H1B7 than those of H1B8-H1B10, despite these bites all having the
271 same water content. This signal response was attributed to the cuttings pile having reached the same
272 height as the end of the drill chute and thus blocking the pile itself from uniformly distributing new
273 material from the drill. The pile evolution is shown in Figure 8. The pile created by H1B7 would negatively
274 affect the flow of extracted material by the remaining bites, and subsequently the signal intensities
275 corresponding to those bites.



278 Figure 8: Video captures of cuttings pile progression for hole 1, each different bite and its resulting pile is
279 shown.

280

281 Hole 2 water peak intensity values were overall lower than the rest of the drill holes, where H2B7
282 corresponding to the first bite in the 5% concentration resulted in a signal intensity of 8.9×10^{-9} A,
283 compared to values of 1.65×10^{-8} A for H1B7, 2.94×10^{-8} A for H3B7, and 2.51×10^{-8} A for H4B7, respectively.
284 This trend of lower signal intensities for hole 2 was attributed to a blockage of MSolo's FOV (calculated to
285 be approximately 12%) by the multi-layer insulation (MLI) used to cover the trolley's gearbox system. In
286 general, a negative signal intensity difference was observed when comparing intensity signals of each bite
287 for hole 2 to hole 1.

288 For hole 3, bites H3B1 to H3B6 produced smooth water signal profiles as seen in Figure 7.
289 However, from H3B7 to H3B10 a sharp spike-like signal was detected at the end of the drilling operations
290 into the 5% doped regolith, most noticeably in H3B9 and H3B10. Closer inspection via the chamber-
291 mounted cameras showed atypical regolith behavior correlating to the observed signal spikes. During the

292 later drilling activities of hole 3, sample pile disturbances were observed. Video showed that loose
293 regolith surrounding the drill hole by the end of the drill chute began to fall back into the hole as the drill
294 finished its retraction motion. It was this disruption of regolith that resulted in an unplanned release of
295 volatile material during the drilling of this hole.

296 In the case of hole 4, a build-up of the pile cuttings similar to that of hole 1 resulted in an
297 disproportionate flow of material coming off the drill chute. By the start of H4B8, a large overflow of
298 cutting material began to fall from the edge of the regolith bin, causing a large water signal spike. Since
299 any more material would only cause more disruptions, it was therefore determined that no more cuttings
300 could be obtained for this hole, and the test was stopped after H4B8.

301 For the hole 1 bites, a stepwise increase in peak signal intensities per concentration depth was
302 expected because the regolith was undisturbed prior to testing. Instead, a strong initial water signal spike
303 occurred when drilling into a new concentration depth, followed by lower peak intensities despite the
304 extracted material having the same concentration as the prior bites. Although there were some
305 experimental setbacks during hole 2 operations, the same pattern for signal intensities can be seen from
306 bite to bite, as was observed in hole 1. Holes 3 and 4 did not exhibit decreasing peak intensities within the
307 same concentration of water in the regolith for bites 4-10 and 4-8, respectively, but did for the first three
308 bites. Ignoring spikes attributed to avalanching regolith, for H3B4-10, the signals maintained consistent
309 peak intensities within the same concentration layer. For H4B4-6, the signals were also consistent.

310 It is worth noting that the signals observed during the drilling of hole 1 are unaffected by potential
311 contributions from regolith from previous holes, whereas signals for subsequent holes 2-4 may have been
312 impacted. After hole 1, the drill was fully retracted, followed by a drill percussion to remove material from
313 the drill. The pile was then flattened by the drill footpad to get the sampling surface as even as possible
314 to prepare for the next drill hole. These activities caused the extracted material from deeper layers to be
315 exposed and spread over the top (dry) layers, changing the initial conditions for the bites of the remaining
316 holes. Though this surface contamination appeared to affect the signals associated with lower water
317 concentration by creating a progressive elevated water baseline value visible in the yellow regions of
318 Figure 7, it does not appear to have significantly disrupted the higher concentrations.

319

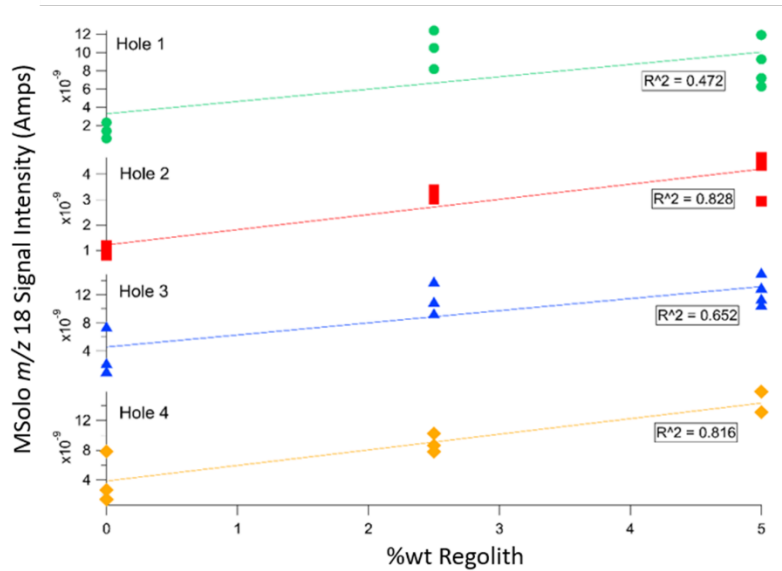


Figure 9: MSolo peak raw intensity correlation to wt% values.

320

321 Figure 9 shows the statistical correlation between the signal intensities detected for the water signal
 322 (m/z 18) and the corresponding wt% of the water in the simulant. Despite the effects of pile dynamics
 323 mentioned previously, an increase of signal intensity as a function of depth could be determined. hole 1
 324 however, displayed a greater average peak signal for the 2.5 wt% water bites than the 5 wt% water bites,
 325 which is counter to what was expected.

326

327 3. Empirical Correlation

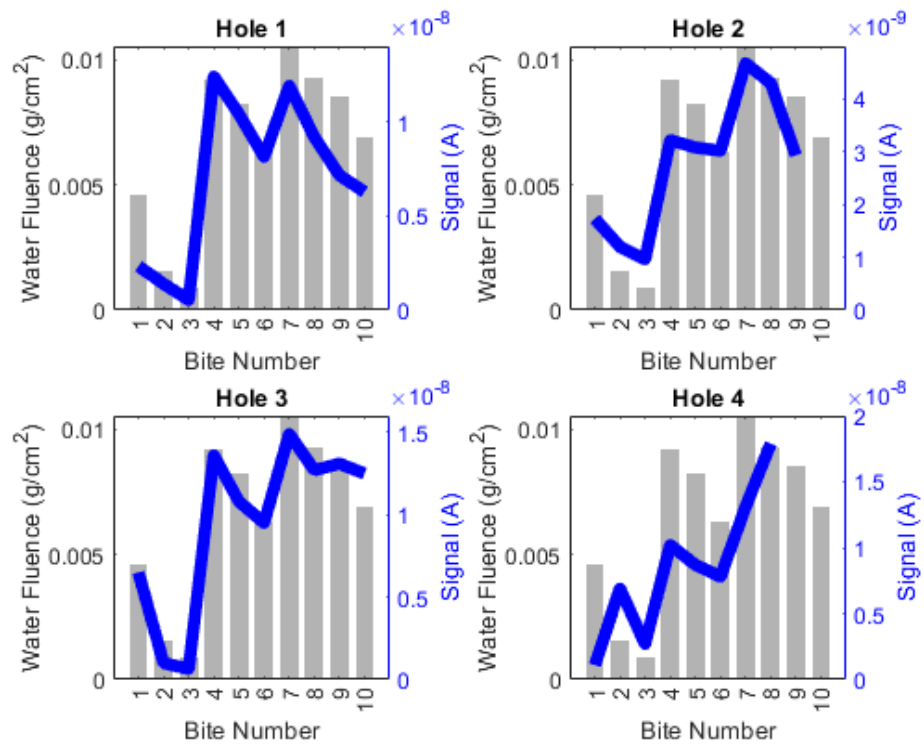
328 A quantitative relationship between the signals observed by MSolo during the test and the “visible”
 329 geometry of the regolith piles was assessed. This relationship provides a foundation for correlating the
 330 physical behavior of the water molecules as they evolve from the drill piles with the signals observed by
 331 MSolo. This relationship can be considered a pseudo-calibration curve with the potential to be used in
 332 real time during lunar missions, however, with notable limitations, to be discussed later.

333 An explanation for the decreasing signal pattern observed for bites within the same concentration
 334 layer in the first hole took into consideration the surface area of the pile. It is assumed that the same
 335 number of water molecules are contained within each bite for a particular concentration layer, and only
 336 the molecules in the most recent bite can travel to the instrument. It is known that the surface area of the
 337 pile increases with each bite. Therefore, the water molecules are less dense for repeated bites within a

338 layer of regolith at the same concentration. The signal observed by MSolo is dependent on the number of
339 molecules capable of reaching the detection zone of the instrument. With a constant instrument position,
340 and with each bite decreasing the density of water molecules, fewer molecules can enter the instrument,
341 and thus signal decreases for each bite within the same concentration.

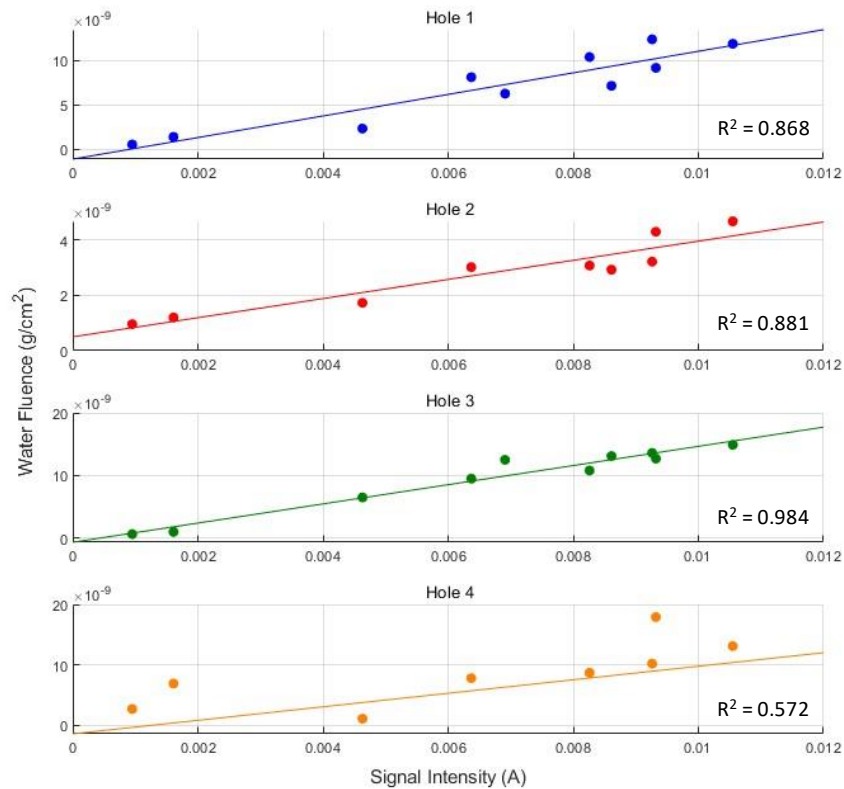
342 Utilizing cameras located adjacent to and above the bin, the diameter and height of the cuttings
343 pile from each bite was measured and used to calculate the conical surface area of the pile. The pile size
344 increases with each bite, and as such, the conical surface area increases with each bite. The initial water
345 concentration at each depth was known (< 0.2 wt%, 2.5 wt%, and 5.0 wt% water) as was the volume of
346 each bite, thus, the total amount of water in each bite could be approximated. Assuming that water could
347 only sublime from the newest bite volume, the water sublimation fluence is the approximated total
348 amount of water per bite divided by the surface area of the pile. It is assumed that this water fluence is
349 representative of the flux of molecules that are detected by MSolo. It is important to note that the
350 projected aperture area onto the pile surface is not explicitly accounted for during this approximation.
351 Bites 1-3, 4-6, and 7-10 each had the same concentration of water and thus the same number of water
352 molecules, respectively, but their total conical surface areas increase as the pile size increased. Therefore,
353 there is a decrease in water density and decrease in fluence following the initial bite at the new
354 concentration, similar to the decrease in the water signal (m/z 18) intensity from MSolo (Figure 10Error!
355 **Reference source not found.**). This phenomenon was evident in holes 1, 2, and 3. However, the pile falling
356 over the edge of the bin in hole 4 bite 8 resulted in a water intensity that was greater than what was
357 expected. Figure 10 overlays the peak signals observed for each bite for each hole (line plot) with the
358 water fluences calculated from the dimensions of the piles generated from the drilling of hole 1 (bar
359 chart), as only measurements of pile radius and height for hole 1 were available.

360



361

362 Figure 10: Water fluence per cm² (bar graph) and the water signal (*m/z* 18) intensity (line graph) for holes
 363 1-4.



365

366 Figure 11: Water fluence per cm² versus m/z 18 intensity shows a linear relationship between the two
 367 parameters. The solid line is the linear best-fit line.

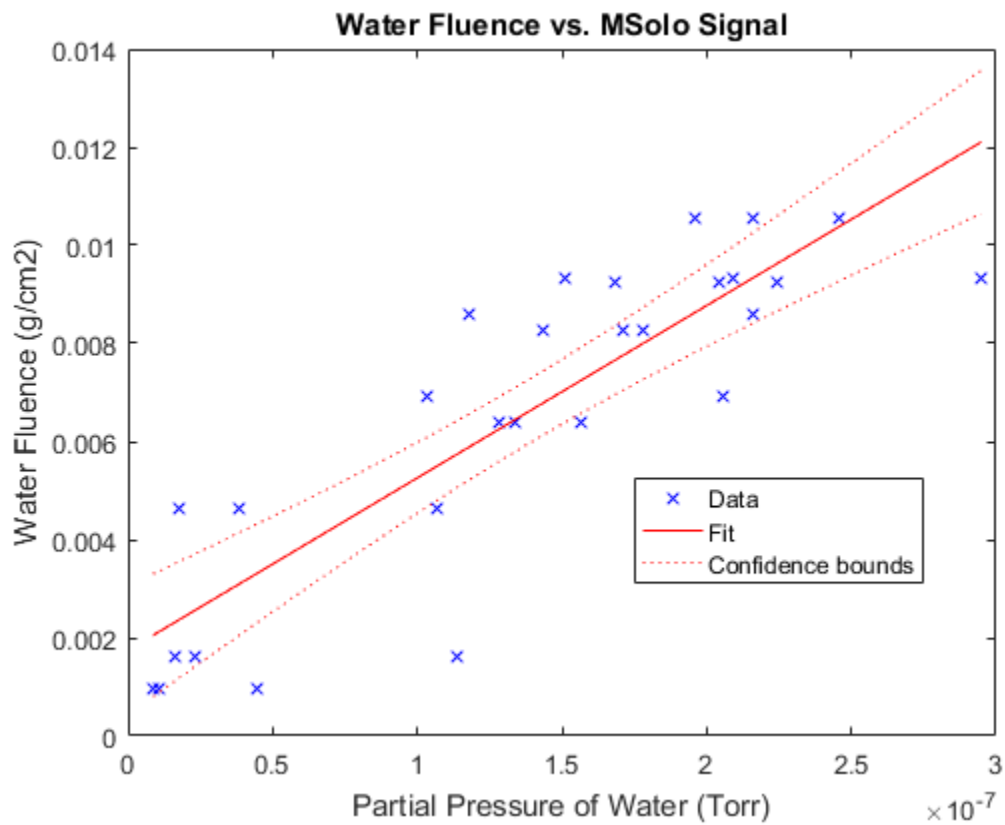
368 The relationship between the water fluence and the water signal (m/z 18) was plotted for each
 369 hole and shown to have a roughly linear relationship (see Figure 11). This water fluence is not a measured
 370 quantity, but an approximation based on the volume of the bite and the approximate density of packed
 371 regolith (Connolly and Carrier, 2023), and this value is divided by the conical surface area calculated from
 372 pile measurements. As only hole 1 pile dimensions were measured, these dimensions were used for the
 373 water fluence for holes 3 and 4, as well. The R-squared values for holes 1-3 were all above 0.85, with hole
 374 3 reporting 0.98.

375 This linear relationship between the water fluence and observed signal was used to generate a
 376 MATLAB code that takes in inputs of measured pile radius, height, bite depth, partial pressure of water
 377 (a value derived from the measured signal, the instrument sensitivity, the instrument gain, and several
 378 other constants) and outputs an approximation for the wt% of water in the regolith. To compare the

379 signals from the different depths and drill holes, it is assumed that the drill cuttings are deposited from
380 the drill at the same temperature for each bite.

381 The water fluence per bite was fit to the peak signal intensities measured for holes 1, 3 and 4 as
382 shown in the plot below. Hole 2 was not included because MSolo's FOV was blocked during drilling of
383 this hole as mentioned previously.

384



385

386 Figure 12: Linear regression fitting an approximation for water fluence from the regolith piles to the
387 partial pressures of water observed during GRC testing.

388

389

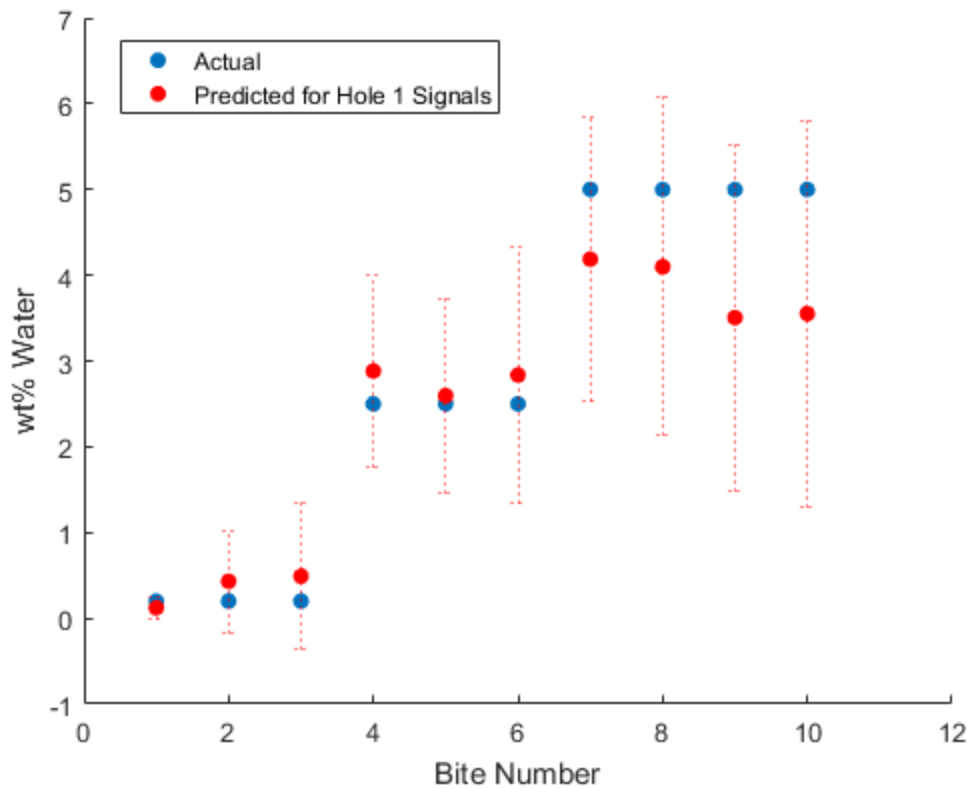
390 The algorithm used with this fit within the MATLAB code is as follows:

- 391 1. The total mass per bite is calculated by approximating the bite volume as a cylinder using
392 the radius of the drill and the depth of the bite as the height. This volume is then multiplied
393 by the density of packed regolith determined during the regolith preparation.

- 394 2. The surface area of the pile is calculated from the measured radius and height of the pile.
395 3. The measured peak signal from the MSolo instrument is converted to a partial pressure
396 using the instrument gain and sensitivity, in addition to other constants.
397 4. The fit is used to convert the partial pressure to the water fluence.
398 5. This water fluence is multiplied by the surface area calculated in step 2 and divided by the
399 total bite mass calculated in Step 1 to produce the wt% per bite.

400 The MSolo signals, pile heights and radii from hole 1 were used as inputs to this algorithm and
401 generated the plot in Figure 13. The output wt% values fall within a 95% prediction interval of the
402 predicted wt%. A prediction interval was used in this analysis as the goal is to predict the result of a
403 single future observation rather than an average of several observations (Faraway, 2009). The prediction
404 intervals increase with the bite number as the prediction interval is calculated from the water fluence vs.
405 partial pressure relationship and then converted from representing water fluence to representing mass
406 fraction using the conversion explained above in step 5. The conversion value increases with bite
407 number because the surface area of the pile increases with additional bites while the total mass of the
408 bite stays approximately the same, causing the calculated prediction interval to increase as well (see
409 “Supplementary Section” for explicit calculations).

410



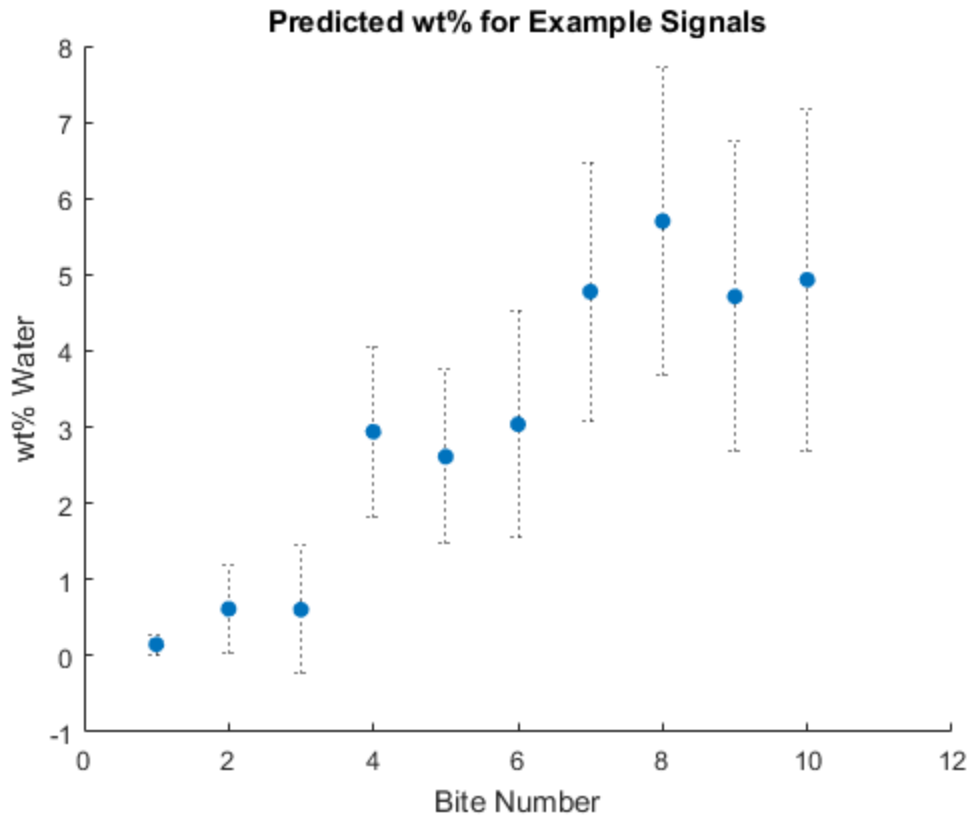
411

412 Figure 13: Predictions for the wt% water for each bite generated from hole 1 signals and the
 413 MATLAB algorithm

414

415 MSolo detects gases whose behavior is dependent on temperature, therefore, a scale factor to
 416 adjust the signal-wt% relationship will likely be necessary in order to effectively use the empirical
 417 correlation, as the lunar surface will have different environmental conditions than in the test case. An
 418 initial scale factor based on the equation for sublimation per surface area (Andreas, 2007) was
 419 calculated. The scale factor is a ratio of the sublimation per surface area at the GRC conditions over the
 420 sublimation per surface area at an input temperature condition. Based on the temperature dependence
 421 of this equation, it can be estimated that the water fluence from the piles in permanently shadowed
 422 lunar crater conditions (50 K plus 60 K added for regolith heating by the drill) would be 25 million times
 423 less than the fluence measured during the terrestrial test (100 K plus 60 K). As such, we took an average
 424 of the signals from the three holes for each bite and scaled these average signals down by an arbitrary
 425 1×10^{-8} to get reasonable wt% values. These signals were put into the MATLAB code along with an
 426 assumed lunar surface plus heated regolith temperature of 110 K. Results are shown in Figure 14.

427



428

429 Figure 14: Predictions for the wt% water for each bite generated from sample signal inputs, lunar
 430 temperature, and the MATLAB algorithm.

431

432 It is obvious that this relationship is also imperfect. Scaling the average signals by 1×10^{-8} results
 433 in signals on the order of 1×10^{-17} , which are below the limit of detection of MSolo. However, signals of
 434 this magnitude are necessary for the temperature scale factor at 110 K and the code to generate
 435 reasonable wt%, as shown in Figure 14. Typically, a calibration curve is only valid under the conditions in
 436 which it was developed because the behavior of instrument and the volatilized water could be affected
 437 by temperature and pressure. Though the GRC tests approached lunar surface conditions, there is
 438 divergence. The temperature scale factor was included to compensate for this, but it is not sufficient to
 439 account for other differences including pressure and physical characteristics of the regolith. The data
 440 used to generate the calibration curve was also highly variable and resulted in a curve fit with an R^2
 441 value of only 0.75. Gathering additional data in the future will improve the robustness of this calibration
 442 curve in addition to making progress with understanding the phenomena observed during the GRC tests
 443 and thus, future development of theoretical models.

444 **4. Conclusion**

445 Despite the limitations of this test campaign, performing relative water quantifications based on
446 MSolo’s observed signal intensity during lunar drilling operations is possible. Relative quantification of the
447 data is already possible via inspection of the raw data from the test—drilling regolith with the lowest
448 concentration of water corresponded to the lowest signal obtained by MSolo, and drilling regolith with
449 the highest concentration of water corresponded to the highest signals observed by MSolo. The empirical
450 correlation shows promise as a starting point towards estimating absolute regolith water concentrations
451 from MSolo signals, however, additional data is necessary to provide accurate and repeatable results.

452

453 **Acknowledgements:**

454 The authors would like to extend their gratitude to the entire PRIME-1 team and to the many
455 contributions from members of the KSC Prototype Design Laboratory, the KSC Vibration Laboratory, and
456 all other members who worked in the development of the MSolo and TRIDENT instruments. We would
457 also like to thank the INFICON personnel for their assistance during the test campaign. We recognize
458 NASA’s Science Technology Mission Directorate (STMD) for their financial support of the testing
459 campaign.

460

461 **References**

- 462 Andreas, E.L. (2007) 'New estimates for the sublimation rate for ice on the Moon', *Icarus*, 186(1), pp.
463 24–30. Available at: <https://doi.org/10.1016/j.icarus.2006.08.024>.
- 464 Colaprete, A. *et al.* (2010) 'Detection of Water in the LCROSS Ejecta Plume', *Science*, 330(6003), pp. 463–
465 468. Available at: <https://doi.org/10.1126/science.1186986>.
- 466 Connolly, J. and Carrier, W.D. (2023) 'An Engineering Guide to Lunar Geotechnical Properties', in *2023*
467 *IEEE Aerospace Conference. 2023 IEEE Aerospace Conference*, Big Sky, MT, USA: IEEE, pp. 1–9. Available
468 at: <https://doi.org/10.1109/AERO55745.2023.10115961>.
- 469 Cui, J. *et al.* (2023) 'Simulation of Drilling Temperature Rise in Frozen Soil of Lunar Polar Region Based on
470 Discrete Element Theory', *Aerospace*, 10(4), p. 368. Available at:
471 <https://doi.org/10.3390/aerospace10040368>.
- 472 Faraway, J.J. (2009) *Linear models with R*. London: Chapman & Hall/CRC.
- 473 Feldman, W.C. *et al.* (1998) 'Fluxes of Fast and Epithermal Neutrons from Lunar Prospector: Evidence for
474 Water Ice at the Lunar Poles', *Science*, 281(5382), pp. 1496–1500. Available at:
475 <https://doi.org/10.1126/science.281.5382.1496>.
- 476 Fraser, H.J. *et al.* (2001) 'Thermal desorption of water ice in the interstellar medium', *Monthly Notices of*
477 *the Royal Astronomical Society*, 327(4), pp. 1165–1172. Available at: [https://doi.org/10.1046/j.1365-
478 *8711.2001.04835.x*.](https://doi.org/10.1046/j.1365-8711.2001.04835.x)
- 479 Kleinhenz, J. and Linne, D. (2013) 'Preparation of a Frozen Regolith Simulant Bed for ISRU Component
480 Testing in a Vacuum Chamber', in *51st AIAA Aerospace Sciences Meeting including the New Horizons*
481 *Forum and Aerospace Exposition. 51st AIAA Aerospace Sciences Meeting including the New Horizons*
482 *Forum and Aerospace Exposition*, Grapevine (Dallas/Ft. Worth Region), Texas: American Institute of
483 Aeronautics and Astronautics. Available at: <https://doi.org/10.2514/6.2013-732>.
- 484 Kleinhenz, J.E. (2014) 'Lunar Polar Environmental Testing: Regolith Simulant Conditioning', in *7th*
485 *Symposium on Space Resource Utilization. 7th Symposium on Space Resource Utilization*, National
486 Harbor, Maryland: American Institute of Aeronautics and Astronautics. Available at:
487 <https://doi.org/10.2514/6.2014-0689>.
- 488 Li, Y. *et al.* (2023) 'The Mechanism for the Barrier of Lunar Regolith on the Migration of Water
489 Molecules', *Journal of Geophysical Research: Planets*, 128(3), p. e2022JE007254. Available at:
490 <https://doi.org/10.1029/2022JE007254>.
- 491 Metzger, P.T., Zacny, K. and Morrison, P. (2020) 'Thermal Extraction of Volatiles from Lunar and Asteroid
492 Regolith in Axisymmetric Crank–Nicolson Modeling', *Journal of Aerospace Engineering*, 33(6), p.
493 04020075. Available at: [https://doi.org/10.1061/\(ASCE\)AS.1943-5525.0001165](https://doi.org/10.1061/(ASCE)AS.1943-5525.0001165).
- 494 Metzger, P.T., Zacny, K. and Morrison, P. (2021) 'Erratum for "Thermal Extraction of Volatiles from Lunar
495 and Asteroid Regolith in Axisymmetric Crank-Nicholson Modeling" by Philip T. Metzger, Kris Zacny, and
496 Phillip Morrison', *Journal of Aerospace Engineering*, 34(5), p. 08221002. Available at:
497 [https://doi.org/10.1061/\(ASCE\)AS.1943-5525.0001312](https://doi.org/10.1061/(ASCE)AS.1943-5525.0001312).

498 Murphy, D.M. and Koop, T. (2005) 'Review of the vapour pressures of ice and supercooled water for
499 atmospheric applications', *Quarterly Journal of the Royal Meteorological Society*, 131(608), pp. 1539–
500 1565. Available at: <https://doi.org/10.1256/qj.04.94>.

501 Nozette, S. *et al.* (1996) 'The Clementine Bistatic Radar Experiment', *Science*, 274(5292), pp. 1495–1498.
502 Available at: <https://doi.org/10.1126/science.274.5292.1495>.

503 Paulsen, G. *et al.* (2018) 'The Regolith and Ice Drill for Exploration of New Terrains (TRIDENT); a One-
504 Meter Drill for the Lunar Resource Prospector Mission'. Available at:
505 <https://es mats.eu/amspapers/pastpapers/pdfs/2018/paulsen.pdf>.

506 Roozbahani, M., Borela, R. and Frost, J. (2017) 'Pore Size Distribution in Granular Material
507 Microstructure', *Materials*, 10(11), p. 1237. Available at: <https://doi.org/10.3390/ma10111237>.

508 Sanders, G.B. and Larson, W.E. (2013) 'Progress Made in Lunar In Situ Resource Utilization under NASA's
509 Exploration Technology and Development Program', *Journal of Aerospace Engineering*, 26(1), pp. 5–17.
510 Available at: [https://doi.org/10.1061/\(ASCE\)AS.1943-5525.0000208](https://doi.org/10.1061/(ASCE)AS.1943-5525.0000208).

511 Simon, L. *et al.* (no date) *STAT 501: Regression Methods, Penn State Eberly College of Science*. Available
512 at: <https://online.stat.psu.edu/stat501/> (Accessed: 24 April 2025).

513 Wilkerson, R.P. *et al.* (2023) 'Outgassing behavior and heat treatment optimization of JSC-1A Lunar
514 regolith simulant', *Icarus*, 400, p. 115577. Available at: <https://doi.org/10.1016/j.icarus.2023.115577>.

515 Zacny, K. *et al.* (2023) 'TRIDENT Drill for VIPER and PRIME-1 Missions to the Moon', in *Earth and Space
516 2022. 18th Biennial International Conference on Engineering, Science, Construction, and Operations in
517 Challenging Environments*, Denver, Colorado: American Society of Civil Engineers, pp. 465–474. Available
518 at: <https://doi.org/10.1061/9780784484470.043>.

519

520 Supplementary Section

521 Confidence Interval Calculation

522 The Linear Regression fit of the “water fluence” (y data) to the partial pressure of water (x data) is used
523 to calculate a 95% prediction interval using the equation for standard error of the prediction:

524
$$PI = \hat{Y}_h \pm t \times \sqrt{MSE \times \left(1 + \frac{1}{n} + \frac{(X_h - \bar{X})^2}{\sum(X_i - \bar{X})^2}\right)}$$

525 Where PI is the 95% confidence interval, \hat{Y}_h is the predicted value generated from a predictor value, X_h ,
526 t is the t-value which depends on the confidence level and degrees of freedom n , MSE is the mean
527 square error calculated using the equation below, \bar{X} is the mean of the known x-data used to determine
528 the fit, and X_i is the i^{th} known x value (one of the dataset used to calculate the fit)(Simon *et al.*, no date).
529

530 The mean square error is computed as follows:

531
$$MSE = \frac{1}{n} \sum_{i=1}^n (Y_i - \hat{Y}_i)^2$$

532 Where Y_i is the known y-data, \hat{Y}_i is the y-data generated from the known x-data by the fit, and n is the
533 total number of points of known data.
534

535 t for the 95% confidence interval is computed using MATLAB's built-in `tin` function.
536

537 \hat{Y}_h is calculated using the linear fit.
538

539 PI can then be calculated according to the equation above.
540

541 To convert PI to a representation of our quantity of interest – the mass percent water – PI is then
542 multiplied by the surface area of the pile at a particular bite over the total mass per bite.
543

544 The error bars in Figure 17 are the result of multiplying the second term in the PI equation by the
545 surface area of the pile at a particular bite over the total mass per bite, as mentioned in previous
546 sections. The surface area of the pile increases with the number of bites, while the total mass per bite
547 stays the same, resulting in a larger interval with larger bites.
548

549

550