



THE HABITABLE WORLDS OBSERVATORY TECHNOLOGY DEVELOPMENT PLAN

2026 July 8

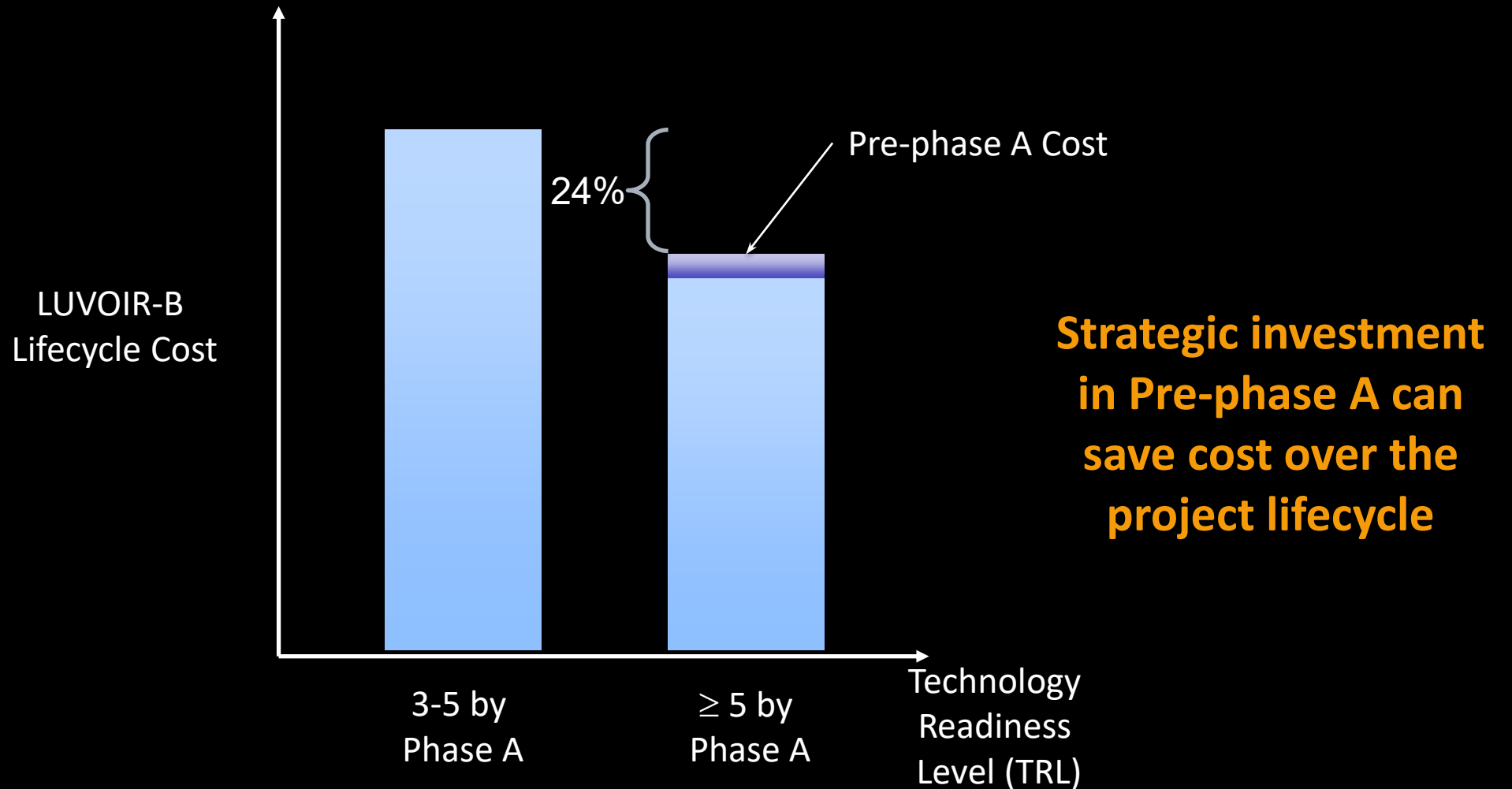
SPIE Astronomical Telescopes & Instrumentation

14145-59

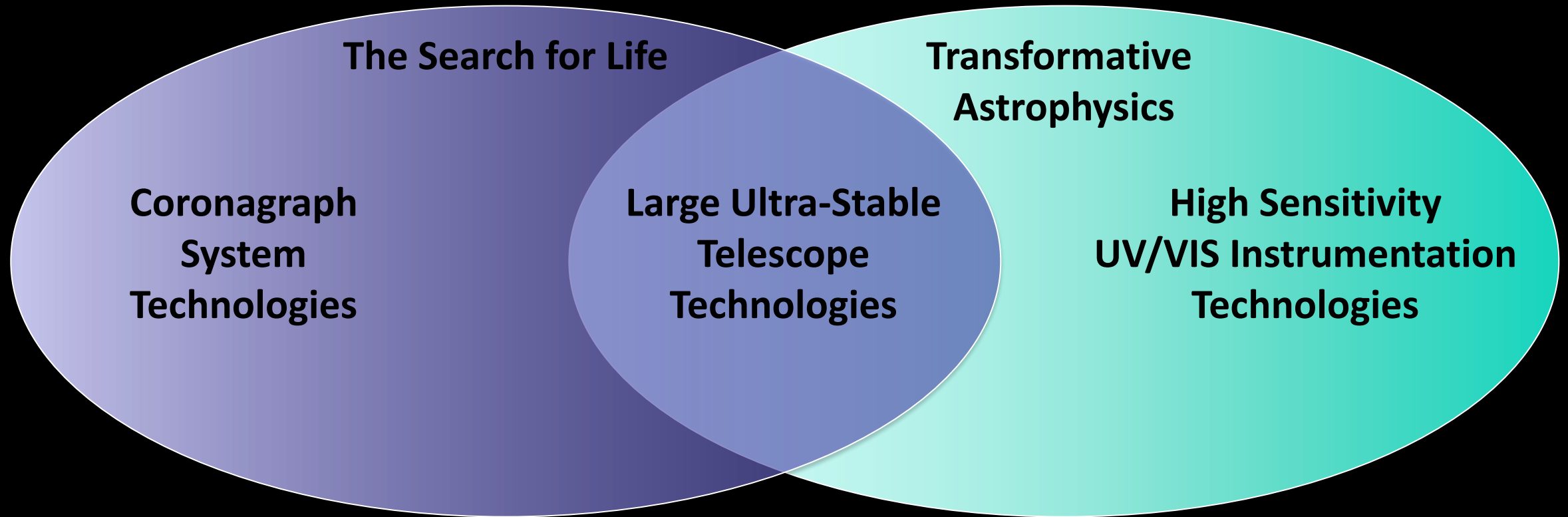
Matthew R. Bolcar (GSFC), Feng Zhao (JPL), Paul Scowen (GSFC), Breann Sitarski (GSFC), Lee Feinberg (ETH Zurich), J. Scott Smith (GSFC), Josh Abel (GSFC), Dave Muliere (JPL), Phil Stahl (MSFC), Tyler Groff (GSFC), Kevin Sims (ARC), Tomasz Lis (MSFC), Stephen Cheney (MSFC)

H A B I T A B L E
W O R L D S
O B S E R V A T O R Y

EARLY TECHNOLOGY DEVELOPMENT SAVES MONEY



TECHNOLOGY NEEDS ARE INFORMED BY SCIENCE AND ARCHITECTURE



Three Technology Tracks aligned to the system architecture.

HWO's SIX TALL POLES



By MCR, we need to show acceptable risk with respect to:

Holding a **segmented mirror** to the required stability

- Thermal stability of the optical surface
- Active sensing & control of rigid body position
- Backplane dynamic stiffness / stability
- UV-quality surface roughness, wavefront, edge roll-off

HWO's SIX TALL POLES



By MCR, we need to show acceptable risk with respect to:

Holding a **segmented mirror** to the required stability

Achieving end-to-end **system-level alignment & stability**

- Active sensing & control of system-level alignments
- Large-scale dynamic isolation & stability
- Large-scale thermal stability

HWO's SIX TALL POLES



By MCR, we need to show acceptable risk with respect to:

Holding a **segmented mirror** to the required stability

Achieving end-to-end **system-level alignment & stability**

Reaching & maintaining the necessary **contrast & contrast stability**

- Design & fabricate masks & components with necessary accuracy
- Control wavefront & amplitude errors with DMs
- Actively sense necessary modes during science operations
- Post-process images to required contrast floor

HWO's SIX TALL POLES



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Detection & characterization of exoEarths with **photon-counting detector systems**

- Required QE and total noise rate with appropriate ConOps
- Feasible operating temperatures
- Radiation hard & robust to environments

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Detection & characterization of exoEarths with **photon-counting detector systems**

A **feasible deployment scheme**

- Accommodate aperture, micrometeoroid, stray light, and thermal constraints
- Meets mass and volume constraints of available fairings

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Detection & characterization of exoEarths with **photon-counting detector systems**

A **feasible deployment scheme**

Far-UV coatings

- Uniform, repeatable, high-reflectance
- Environmentally robust
- Characterization of polarization and scatter impacts on coronagraphy

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Detection & characterization of exoEarths with **photon-counting detector systems**

A **feasible deployment scheme**

Far-UV coatings

HWO TMPO is investing strategically to advance each technology.

PARTNERING WITH INDUSTRY (1/3)



Seven contracts awarded, comprising 13 Technology Demonstrations & 10 Mission Architecture Studies



PARTNERING WITH INDUSTRY (2/3) – SYSTEM DEMONSTRATIONS



BAE SYSTEMS

Developing sub-scale segment phasing demo incorporating edge sensors, ultra-fine actuators, and laser metrology



Demonstrate sub-milli-Kelvin thermal control and picometer wavefront stability; also developing 2-m coating capabilities



Develop active, non-contact disturbance isolation system and perform end-to-end system performance modeling



Investigate ultra-stable structures, latches, and materials; investigate micrometeoroid protection approaches

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All four teams also performing mission architecture studies and parametric modeling in support of HWO systems design team.

PARTNERING WITH INDUSTRY (3/3) – TARGETED EFFORTS



Investigate innovative, cost-effective solutions to serviceability and servicing architectures



Investigate microthruster lifetime and develop capability for HWO applications



Demonstrate Al+LiF mirror coating uniformity and repeatability on 1.8-meter class optics

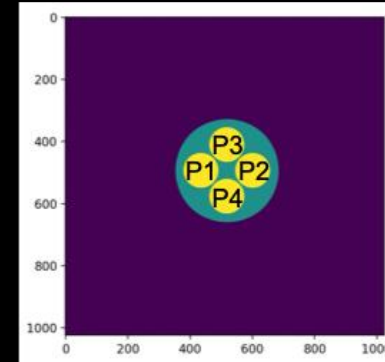
DEMONSTRATING ULTRA-STABILITY WITH NASA EXPERTISE (1/3)



Mini-MUST (Metrology and Ultra-Stable Testbed)
Chamber capable of validating picometer-level
performance of meter-class test articles

- Thermal stability $< \pm 1$ mK over hours
- Vacuum levels $\leq 10^{-5}$ Torr
- Test articles ≤ 1 -m in diameter
- Partnering with D.19 teams and academia to test mirrors, structures, and metrology critical for HWO

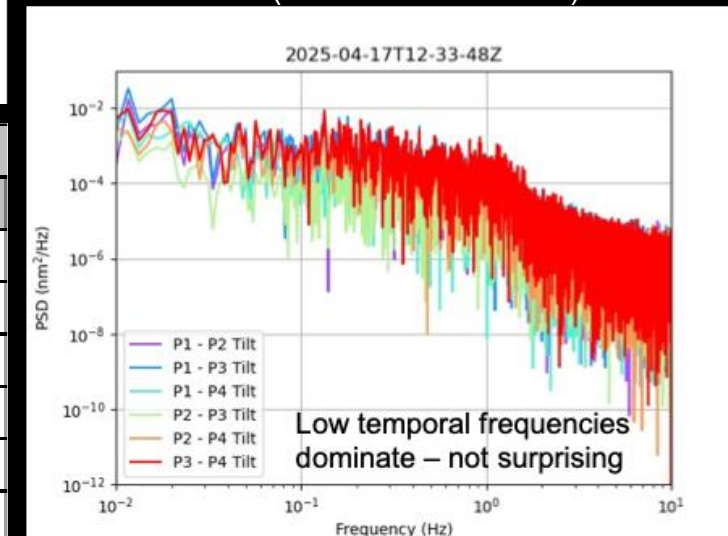
Ultra-stable performance leverages
expertise gained in the pico-chamber,
achieving picometer-level stability
with 5" optics



	Weighted σ (pm)		
0.01 - 10 Hz	Piston	Tip	Tilt
P1 - P2	2.07	3.99	8.81
P1 - P3	2.62	1.60	3.92
P1 - P4	2.85	6.76	5.09
P2 - P3	4.61	4.16	4.40
P2 - P4	6.38	4.23	8.08
P3 - P4	4.63	4.96	5.41



T. Hadjimichael and G. Hilton outside the
mini-MUST chamber with test article
(Credit: NASA/GSFC)



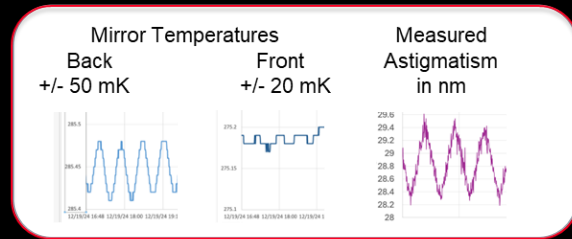
DEMONSTRATING ULTRA-STABILITY WITH NASA EXPERTISE (2/3)



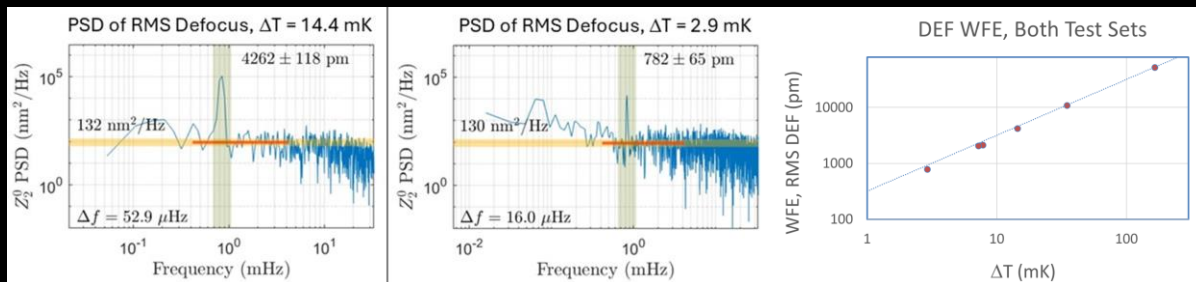
Marshall Space Flight Center developing thermal control algorithms and micrometeoroid damage models:

Thermal Modulation Testing (TMT)

Measure mirror's optical response to thermal stimuli of different amplitude & period
→ Apply frequency-domain analysis



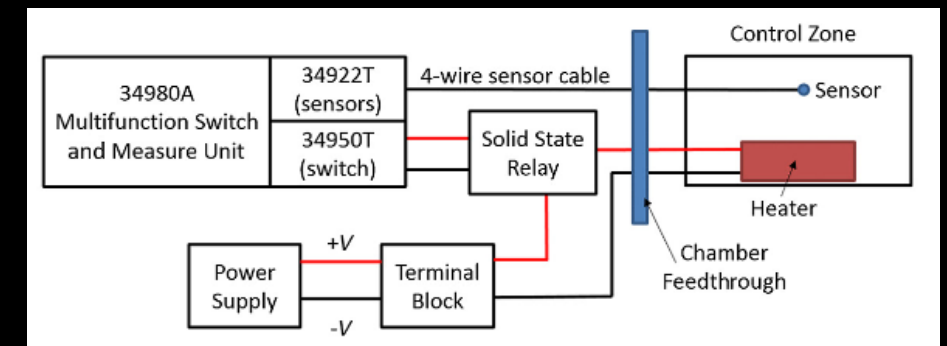
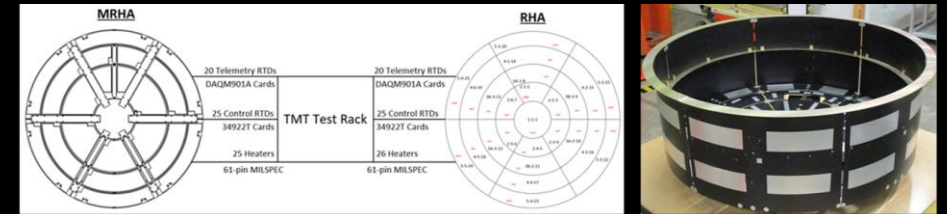
Credit: NASA MSFC



TMT of closed-back BK-7 mirror shows linear thermal response at the pm/K level

Precision Thermal Control (PTC)

Demonstrates 25 independent thermal control zones with < 2 mK RMS stability



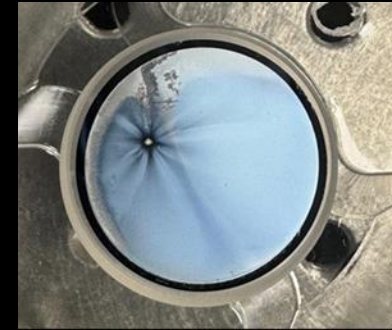
Credit: NASA MSFC

DEMONSTRATING ULTRA-STABILITY WITH NASA EXPERTISE (3/3)

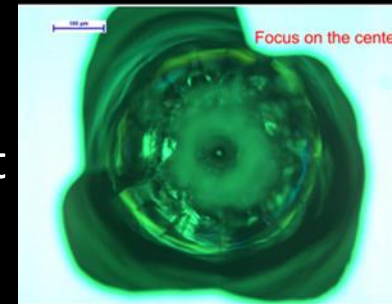
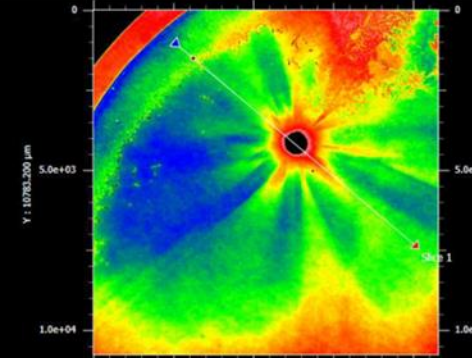


Marshall Space Flight Center developing thermal control algorithms and micrometeoroid damage models:

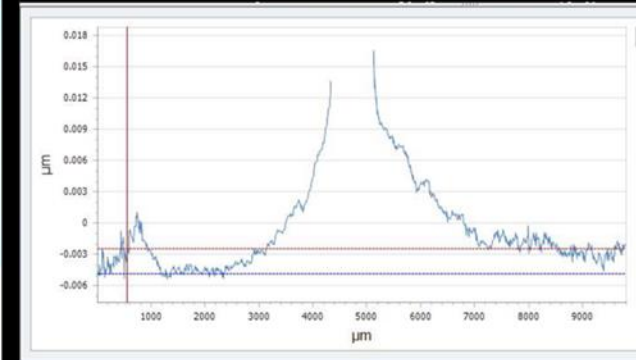
- Micrometeoroids may damage HWO optics and produce scattered light that impacts coronagraphy
- Intent of the study is to:
 - Quantify scattered light from impacts
 - Create predictive model of science longevity
 - Determine best architecture considerations with respect to material & facesheet thickness
- Test plan duplicates physical environment at Univ. of Colorado IMPACT & White Sands test facilities
 - Energies between $2.5e-11$ and 0.0015 J



Sa 4.7 nm
Sq 5.7 nm
Sz 46.2 nm



- Focused through the crater.
- Impact site ~10μm
- Crater is ~500μm
- Deeper than the full focal length of the microscope
- Different structure as focus through



Credit: NASA MSFC

BUILDING ON ROMAN CGI @ JPL...

HCIT - High Contrast Imaging Testbeds
DST2 - Decadal Survey Testbed 2
OMC - Occulting Mask Coronagraph
EPIC5 - ExoPlanet Imaging Coronagraph for TRL-5
LOWFS - Low Order WaveFront Sensing

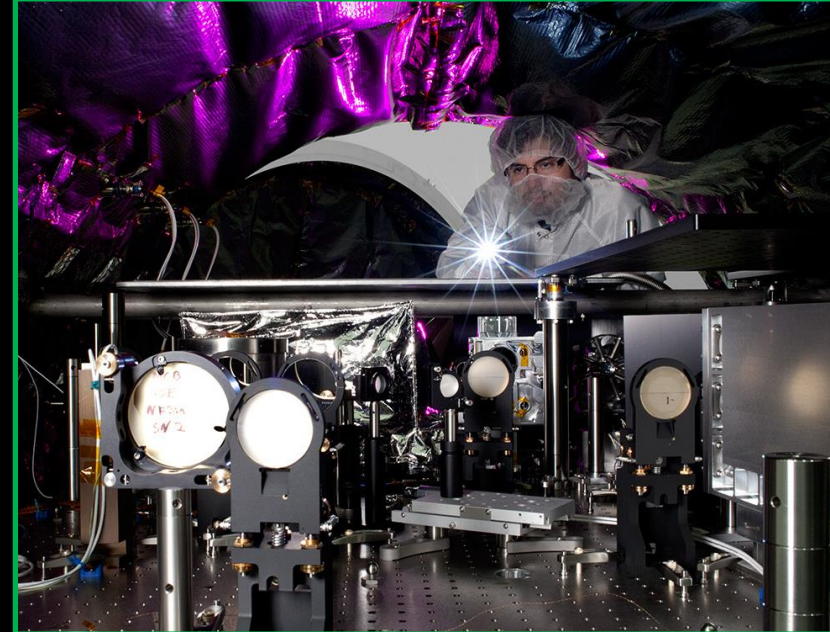


HCIT-DST2: Quasi-static Contrast



Credit: NASA JPL/Caltech

OMC-EPIC5: Contrast Stability



Credit: NASA JPL/Caltech

Design nearly
complete;
Commissioning
early next year

- Dual-testbeds for TRL-5 follows CGI success:

- CGI Key milestone #2: Shaped Pupil Coronagraph @ 1E-8 (narrowband)
- CGI Key milestone #3: Hybrid Lyot Coronagraph @ 1E-8 (narrowband)
- CGI Key milestone #4: Coronagraph demo @ 1E-8 and 10% bandwidth
- CGI Key milestone #5: LOWFS system meets pointing and WFE drift suppression req's

Efforts undertaken
on separate
testbeds

...AND GOING BEYOND (1/3) – MASKS & SPECIALTY OPTICS

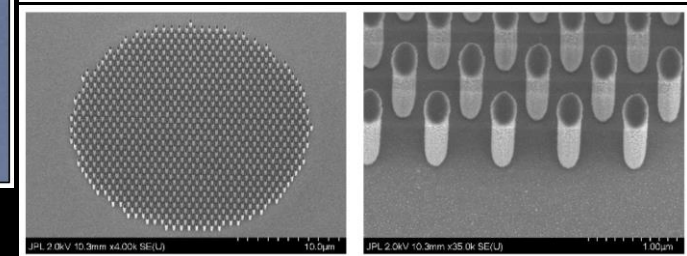
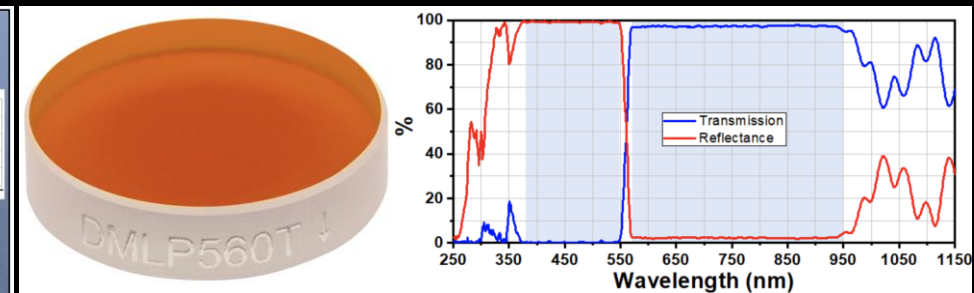
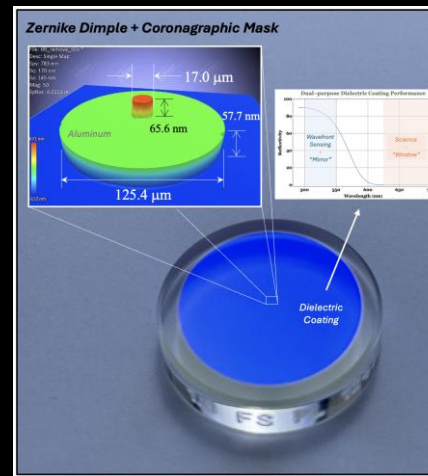


JPL ramping up effort to design, simulate, fabricate, test, and evaluate coronagraph masks for HWO

- Trade studies to include scientific yield vs. mask type, fabrication errors, alignment & calibration, etc.
- Fabricate masks and layouts for demonstration on DST2 & EPIC5 testbeds

Also performing analyses in support of specialty optic design & fabrication:

- Broadband dichroics
- Polarization optics
- Dual-purpose focal plane masks



Credit: NASA
JPL/Caltech

...AND GOING BEYOND (2/3) - DMS



Phased, parallel investments in deformable mirror (DM) development with Boston Micromachines & AO Xinetics

Phase 1

- Vendors are executing process improvements to address key technology risks
- ~64x64 TRL 5 device each delivered at end of decade

Phase 2

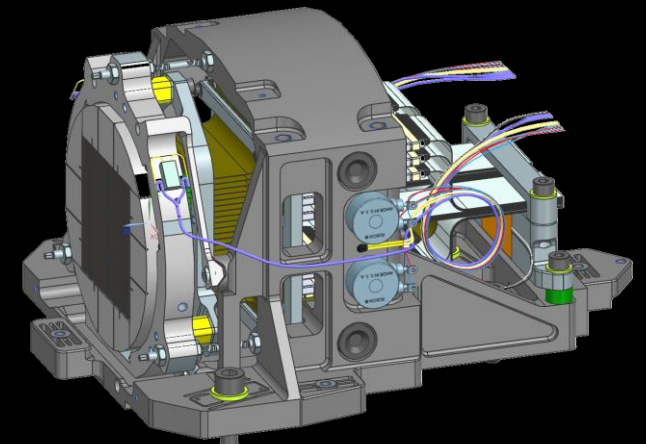
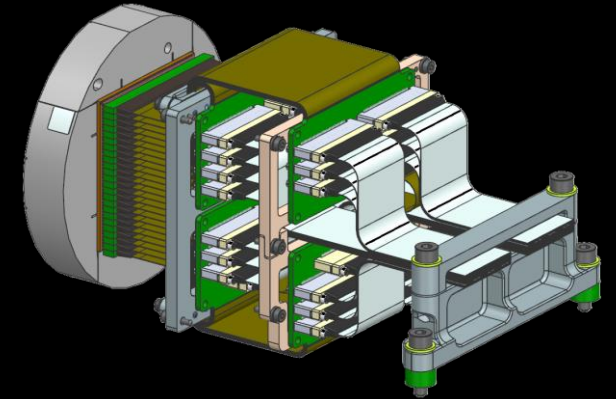
- Chance for second attempt at a full TRL 5 verification
- Two 64x64 TRL 5 devices in 2030

Phase 3

- Two 96x96 TRL 5 devices (delivery TBD, largely based on success of phase 1)

TRL 5 verifications performed on vacuum surface gauge at JPL with complementary verifications done on testbeds in development at Goddard

JPL also developing interconnect concepts to mitigate challenges encountered on Roman CGI



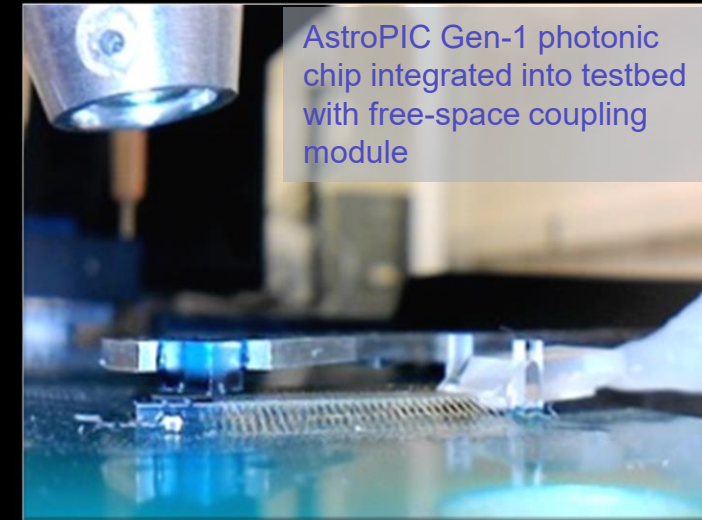
Credit: NASA JPL/Caltech

...AND GOING BEYOND (3/3) - PHOTONICS



Investigating NIR photonic-based coronagraph solutions at Ames Research Center (in-air testbed)

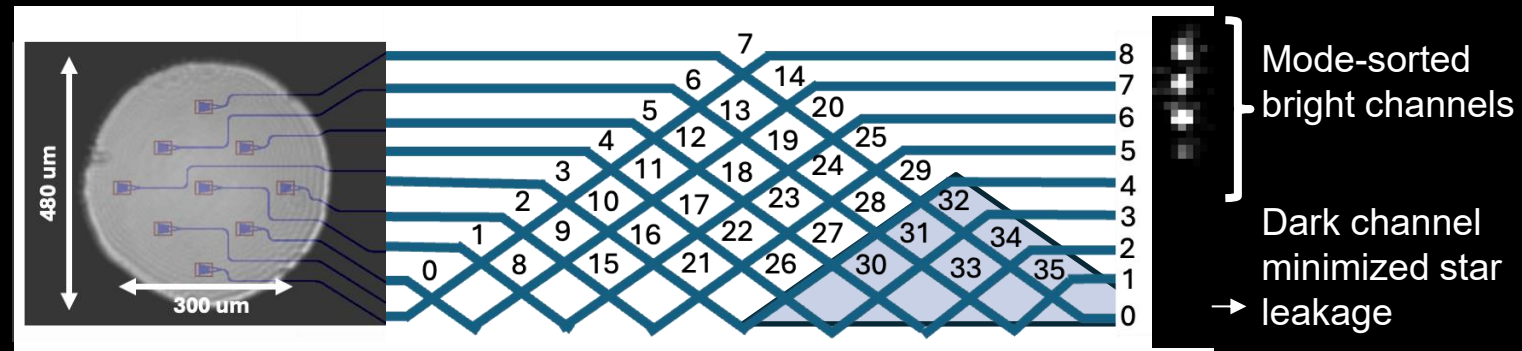
- **$2.0e-9$ suppression** on-chip (first light demonstration, Sept 2025)
- **$3.5e-10$ suppression** on-chip (component demonstration, Jan 2026)
- **$8.7e-8$ contrast** free-space coupling (TRL-3 system demo , Mar 2026)
- **AstroPIC Gen-2 chip submitted to foundry** (May 2026, spatial input channels 9 $\rightarrow$ 16, high-efficiency grating coupling, mono $\rightarrow$ 1% band)



Credit: NASA ARC

Also developing NIR vacuum coronagraph testbed, Multi-star wavefront control techniques, and phase-induced amplitude apodization (PIAA)-vortex coronagraph

See talks: 14145-280, -279, -282, -98



Credit: NASA ARC

DETECTOR SYSTEM DEVELOPMENT STUDY



Partnering with Columbus Technologies on 2-phase effort to mature detector systems

Addresses needs for **visible & NIR** coronagraph detectors, **Far-UV** detectors, and **Near-UV/visible** large-format detectors

Phase 1: Define complete technology development plan for candidate technologies to TRL 5

Phase 2: Execute technology development plan

Call closed July 7, 2026

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HWO Detector Study Request for Proposals (RFP)

HWO Detector Study Request for Proposals

Overview

Columbus Technologies and Services, Inc. is issuing this Request for Proposals (RFP) in support of NASA's Habitable Worlds Observatory (HWO) Technology Maturation Project Office (TMPO).

This RFP seeks information from qualified organizations capable of supporting detector technology maturation studies to advance candidate technologies toward Technology Readiness Level (TRL) 5.

Scope of Interest

Responses may address one or more of the following detector categories:

- Visible coronagraph detectors
- Near-infrared coronagraph detectors
- Far-ultraviolet detectors
- UV-visible large format detectors

Vendors are encouraged to provide information on technical approach, performance capabilities, environmental considerations (thermal and radiation), and maturation plans.

Response Requirement Overview

- Maximum length: 5 pages
- Format: PDF
- Include:
 - Organization overview
 - Technical approach
 - Relevant experience
 - TRL advancement strategy
 - Points of contact
- Refer to the RFP documents for complete response requirements

Submission Instructions

Interested parties shall submit responses via the submission form below. All responses must include company name, primary point of contact, and contact information.

Questions

Please submit questions to HWO-Detector-RFP@columbususa.com. We continue to receive inquiries and are updating the [Question and Answer \(Q&A\) section](#) on a daily basis. All prospective proposers are strongly encouraged to review the Q&A section regularly to ensure access to the most current and authoritative information.

Questions and Answers (Q&A) are posted [below](#).

Submission Deadline

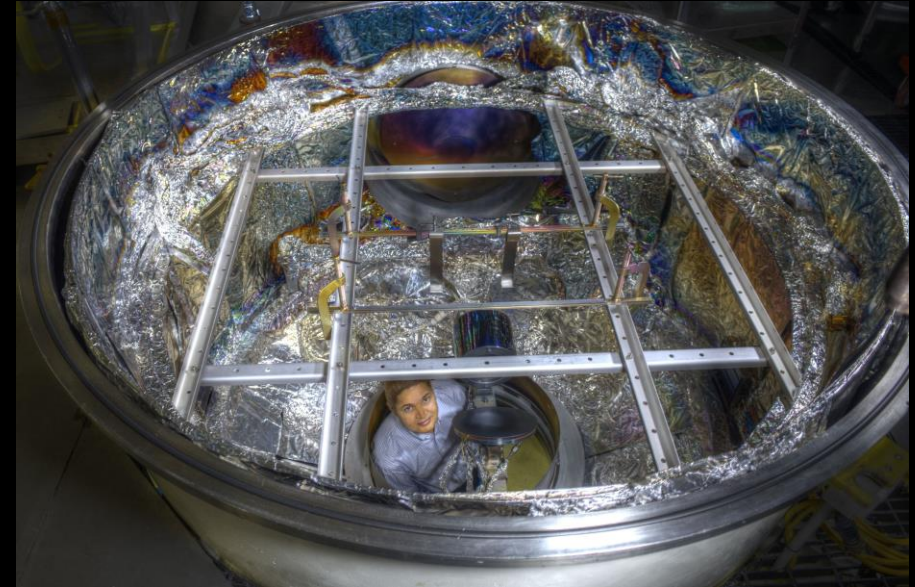
The submission deadline for proposals is July 7, 2026.

PAVING THE WAY FOR FAR-UV SENSITIVITY



GSFC completing development of 2-m coating chamber measure repeatability, uniformity, polarization, and scatter properties of Al+XeLiF coating on 1.5-m-class optics

- Working closely with partners at Zecoat and L3Harris to transfer the capability to industry



Credit: NASA GSFC

Both GSFC & JPL developing Atomic Layer Deposition (ALD) techniques for enhanced Far-UV filters and gratings

CONCLUSION



The HWO Technology Maturation Project Office has invested in a comprehensive portfolio of development efforts with our industry, academic, and NASA Center partners

Our technology development plan continues to evolve with ongoing science, architecture, and design concept investigations