Secular and Nonsecular Behavior for the Cold Plasma Equations

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The origin of "secular" behavior for the nonlinear cold electron plasma equations is studied. The equations involved are closely related to the Klein-Gordon equation with a small nonlinear term. A method is developed for arriving at perturbation theoretic solutions of this equation, and the method is then applied to the case of the higher-order effects of an electromagnetic wave propagating in the cold electron plasma. An explicit expression for the second order frequency shift is calculated.

I. INTRODUCTION

WHEN one attempts to calculate the behavior of nonlinear mechanical systems in perturbation theory which goes beyond the linearized approximation, often (though by no means always) the difficulty of "secular" behavior appears. The higher orders contain, in addition to trigonometric terms, time-proportional, or "secular" terms. The unbounded character of these terms soon invalidates the perturbation theoretical assumptions of smallness on which they were derived.

A systematic program for doing a type of perturbation theory which is free of secular terms in the case of the harmonic oscillator equation with a "small" nonlinear term was given some time ago by Krylov and Bogoliubov,1 and later refined and mathematically justified by Bogoliubov and Mitropolski.2 Recently, considerable interest has arisen in modifying these techniques to deal with partial differential and differential-integral equations, particularly in connection with the work of Frieman3 and Sandri.4 Earlier calculations, which are more closely related to this work, were made by Jackson5 and Sturrock.6

The system treated below is a partial differential equation, considerably simpler than those described in Refs. 3 and 4, but which we nonetheless believe to be of value in illuminating, in a relatively uncluttered way, some of the new features which emerge as a consequence of the fact that we are studying a partial differential, rather than an ordinary differential system.

The equation treated by Bogoliubov and Mitropolski is

$$\left( \frac{d^2}{dt^2} + \omega^2 \right) x(t) = \epsilon F\left( x, \frac{dx}{dt} \right),$$

(1)

where \( \epsilon \) is a formal expansion parameter (eventually to be set equal to one) used to indicate the relative "smallness" of the right-hand side. \( F(x, dx/dt) \) is a known nonlinear functional of \( x \) and \( dx/dt \); \( \omega^2 \) is a constant.

We examine the equation

$$\left( \frac{\partial^2}{\partial t^2} - \epsilon^2 \frac{\partial^2}{\partial x^2} + \lambda^2 \right) f(x, t) = \epsilon F\left( f, \frac{\partial f}{\partial t}, \frac{\partial f}{\partial x} \right),$$

(2)

where \( \epsilon \) is again the formal expansion parameter, \( F \) is again a known functional (which depends on the problem under consideration), and \( \epsilon^2 \) and \( \lambda^2 \) are non-negative real constants.

The immediate motivation for studying (2) was that the equations of motion for a "cold" electron plasma in a uniform positive background may be put in this form, component by component. However, the cold plasma equations are superficially less simple than Eq. (2), so we shall consider it first.

We shall first give an example of both secular and nonsecular behavior for Eq. (2), and show how the difficulties may be remedied in the secular case. Then in Sec. III, we shall give a simple physical example, that of the second order behavior of a nonlinear electromagnetic wave in a "cold" electron plasma.
II. THE KLEIN-GORDON EQUATION WITH A SMALL NONLINEAR TERM

A. Single Monochromatic Wave

We first consider perturbations about the following $\epsilon = 0$ solution to (2)

$$f = a \cos (K_0 x - \omega_0 t + \phi), \quad (3)$$

where

$$\omega_0^2 = c^2 K_0^2 + \lambda^2 \quad (4)$$

and for $\epsilon = 0$, the quantities $K_0$, $\omega_0$, $a$, and $\phi$ are constants.

This happens to be a situation for which secularity arises, so we shall set up the necessary formalism for handling it from the beginning. Following Ref. 2, we seek a solution to Eq. (2) of the form

$$f = a \cos \psi + \epsilon u_1(a, \psi) + \epsilon^2 u_2(a, \psi) + \cdots, \quad (5)$$

where the amplitude $a$ is now determined as a "slowly varying" function of $x$ and $t$ by the relations

$$\frac{\partial a}{\partial t} = \epsilon A_1(a) + \epsilon^2 A_2(a) + \cdots, \quad (6a)$$

$$\frac{\partial a}{\partial x} = \epsilon D_1(a) + \epsilon^2 D_2(a) + \cdots. \quad (6b)$$

$\psi$ is a new "phase" variable, to be chosen to coincide with the phase of (3) for $\epsilon = 0$,

$$\frac{\partial \psi}{\partial t} = -\omega_0 + \epsilon B_1(a) + \epsilon^2 B_2(a) + \cdots, \quad (7a)$$

$$\frac{\partial \psi}{\partial x} = K_0 + \epsilon C_1(a) + \epsilon^2 C_2(a) + \cdots. \quad (7b)$$

The (as yet undetermined) functions $A_1$, $A_2$, ..., $B_1$, $B_2$, ..., $C_1$, $C_2$, ..., $D_1$, $D_2$, ..., are to be chosen so as to render the solution (5) free from secular terms. Here, "secular" must be interpreted to mean $\psi$-proportional, i.e., solutions can break down due to linear growth in $x$ as well as $t$. The functions $u_1$, $u_2$, ..., are to be periodic in $\psi$.

The program is to express the various terms in (2), by means of the relations (5), (6), and (7), in terms of the $u$'s, $A$'s, $B$'s, $C$'s, and $D$'s, as functions of $a$ and $\psi$. For brevity's sake, we shall first only go to $O(\epsilon)$, though in principle the method may be carried to any order in $\epsilon$. Generally, the algebra becomes prohibitive beyond $O(\epsilon^2)$. Hereafter, the notation "$+ \cdots$" will mean "of higher order in $\epsilon$.

The result of differentiating $f$ with respect to $x$ and $t$ and using (6) and (7) is

$$\frac{\partial^2 f}{\partial t^2} - c^2 \frac{\partial^2 f}{\partial x^2} + \lambda^2 f = a \cos \psi (-\omega_0^2 + c^2 K_0^2 + \lambda^2)$$

$$+ \left\{ \lambda^2 \left( \frac{\partial u_1}{\partial \psi} + u_1 \right) + 2(\omega_0 A_1 + c^2 K_0 D_1) \sin \psi \right\} + O(\epsilon^2). \quad (8)$$

The zeroth order part of (8) vanishes identically, by virtue of (4). The coefficient of $\epsilon$ is to be equated to $F(f, \partial f/\partial t, \partial f/\partial x)$, with the arguments replaced by their zeroth order values:

$$f \to a \cos \psi, \quad \frac{\partial f}{\partial t} \to \omega_0 a \sin \psi, \quad \frac{\partial f}{\partial x} \to -K_0 a \sin \psi.$$

For the purpose of finding $u_1$, it is most convenient to write $F$ as a Fourier series in $\psi$

$$F(a \cos \psi, \omega_0 a \sin \psi, -K_0 a \sin \psi)$$

$$= g_0(a) + \sum_{n=1}^{\infty} [g_n(a) \cos n \psi + f_n(a) \sin n \psi]. \quad (9)$$

The $f_n(a)$ and $g_n(a)$ are known functions of the amplitude $a$, determined by the functional form of $F$. We shall make the assumption, always satisfied in practice, that the $f_n(a)$ and $g_n(a)$ go to zero, as $a$ goes to zero, at least as fast as $O(a^2)$.

The equation for $u_1$ may be written as

$$\lambda^2 (\partial^2 \psi^2 + 1)u_1 = g_0(a)$$

$$+ \sum_{n=1}^{\infty} \left[ g_n(a) \cos n \psi + f_n(a) \sin n \psi \right]. \quad (10)$$

This is, effectively, an ordinary differential equation in $\psi$ of a standard type. Its solution contains terms proportional to $\psi$ and $\sin \psi$ and $\cos \psi$, unless the coefficients of $\sin \psi$ and $\cos \psi$ on the right hand side of (10) vanish. If they vanish, then one may find the general solution in the form

$$u_1 = v_0(a) + \sum_{n=1}^{\infty} v_n(a) \cos n \psi + w_n(a) \sin n \psi. \quad (11)$$

Our objective is to find a solution free from $\psi$-proportional terms. This impels us to choose $A_1$, $B_1$, $C_1$, and $D_1$, so that

$$2(\omega_0 A_1 + c^2 K_0 D_1) = f_1(a), \quad (12a)$$

$$2a(\omega_0 B_1 + c^2 K_0 C_1) = g_1(a). \quad (12b)$$

Two more relations may be deduced from the conditions

$$\phi^2 \psi/\partial t \partial x = \phi^2 \psi/\partial x \partial t, \quad \phi^2 \psi/\partial t \partial x = \phi^2 \psi/\partial x \partial t.$$

Reference to Eqs. (6) and (7) shows that, to lowest significant order, these become

$$A_1 dC_1/da = D_1 dB_1/da, \quad (13a)$$

$$A_1 dD_1/da = D_1 dB_1/da. \quad (13b)$$

Equations (12) and (13) are four relations for the four unknowns $A_1$, $B_1$, $C_1$, and $D_1$, and once they have been obtained, (10) can be solved in a
straightforward way by means of (11), and \( u_t \) will contain no \( \psi \)-proportional terms. \( B_t \) can be interpreted as a "frequency shift" and \( C_t \) as a "wave number shift."

However, it is impossible to solve (12) and (13) completely without specifying the physical problem in more detail. For instance, if we wish to work a boundary value problem in which \( f \) is required to oscillate sinusoidally at a given \( x \) for all \( t \), we do not expect \( a \) or \( \psi \) to vary with \( t \), and \( B_t \) and \( A_t \) may be set equal to zero. This leaves

\[
C_t = g_1(a)/2e^2K_0a \tag{14a}
\]

as the "wave number shift" and

\[
D_t = f_1(a)/2e^2K_0. \tag{14b}
\]

If we are interested in an initial value problem in which a pure sine wave is given for all \( x \) at \( t = 0 \), we may correspondingly set \( C_t \) and \( D_t \) equal to zero and solve for \( A_t \) and \( B_t \). Other choices are necessitated by still other problems. (Note in passing that in both cases, the corrections \( \to 0 \) as \( a \to 0 \), as they must in order to make physical sense.)

All this has been for a completely general \( F \), imagined to contain all harmonics in \( \psi \). In practice, it often happens that \( F \) contains only a few harmonics. Observe that secularity may not arise for some forms of the nonlinear term. Thus, if \( F \) is proportional to \( f^2 \) or \( f \; \partial f/\partial x \), say, \( f_t(a) \) and \( g_t(a) \) will both vanish, and in this order, we may set \( A_t, B_t, C_t, \) and \( D_t \), all identically zero, which is equivalent to a completely straightforward kind of perturbation theory. On the other hand, if we were to have \( F \) proportional to \((f)^3\), \( f_t(a) \) is still zero, but \( g_t(a) = 3a^2/4 \) and a straightforward perturbation theory no longer works. The occurrence of secularity thus depends, in a given order, on the particular form of the nonlinear term.

To close this discussion, we give the full solution for \( u_t(a, \psi) \) of Eq. (11). If we make the (arbitrary) choice that all the first harmonic shall be collected in lowest order, we may choose \( \nu_t(a) = \nu_0(a) = 0 \). Then for \( n \neq 1 \),

\[
\nu_t(a) = g_0(a)/n^2(1 - n^2 + 1), \tag{15a}
\]

\[
\omega_0(a) = f_0(a)/n^2(-n^2 + 1),
\]

\[
\left[ f_0(a) \right] = 1/\pi \int_0^{2\pi} d\psi \left[ \sin n\psi \cos n\psi \right] 
\]

\[
\cdot F(a \cos \psi, \omega_0 \sin \psi, -K_0 \sin \psi). \tag{15b}
\]

When the phenomenon of secularity occurs, one must go to one higher order in \( \epsilon \) in order to get corrections to \( f \) which are uniformly valid through terms of \( O(\epsilon) \) for changes in \( x \) and \( t \) of \( O(1/\epsilon) \), and so the solution (15) is not of great interest by itself.\(^5\) However, the expressions such as (14) for frequency and wave number shifts are accurate \( O(\epsilon) \), and are usually more accessible to measurement than the expression for \( u_t \), in any case.

**B. Two Coupled Monochromatic Waves of Differing Frequencies**

We now proceed to the case in which (2) has the \( \epsilon = 0 \) solution

\[
f = a[\cos (K_1x - \omega_1 t + \psi_1) + \cos (K_2x - \omega_2 t + \psi_2)], \tag{16}
\]

where

\[
\omega_1^2 - c^2K_1^2 - \pi^2 = \omega_2^2 - c^2K_2^2 - \pi^2 = 0, \tag{17}
\]

but where there is not necessarily any other relationship between \( \omega_1, K_1, \) and \( \omega_2, K_2 \). If we were to do a straightforward perturbation theory, assuming

\[
f = f^{(0)} + \epsilon f^{(1)} + \epsilon^2 f^{(2)} + \cdots, \tag{18}
\]

with \( f^{(0)} \) given by (16), it is clear that we can always find a secularity-free perturbation theoretical solution to (2) of the form

\[
f^{(0)} = \sum \alpha^{(a)}(\psi) \exp \left\{ i\theta(K_1x - \omega_1 t + \psi_1) \right\}
\]

\[
+ s(K_2x - \omega_2 t + \psi_2) \right\} \tag{19}
\]

except when there happen to exist two integers \( \tilde{r}, \tilde{s} \) such that

\[
\lambda^2(-\tilde{r}^2 - \tilde{s}^2 + 1) - 2\beta(\omega_2 - c^2 K_1 K_2) = 0. \tag{20}
\]

When (20) is fulfilled, \( f^{(1)} \), like the \( u_t \) of Eq. (10), will in general contain terms proportional to the phases, as well as trigonometric terms. For a randomly selected pair of values \( \omega_1, K_1, \) and \( \omega_2, K_2, \) there will exist no integers \( \tilde{r}, \tilde{s} \) which fulfill (20), and there is no necessity for the Bogoliubov techniques. When, on the other hand, it happens that two such integers \( \tilde{r}, \tilde{s} \), do exist [the relation (20) will also hold for \( -\tilde{r}, -\tilde{s}, \) of course], then the Bogoliubov techniques are called for. One method of introducing them is the following, though it is not the only method. [Note added in proof. For a completely general \( F(f, \partial f/\partial t, \partial f/\partial x) \), there will exist nonvanishing Fourier coefficients on the right-hand side of the equation for \( f^{(n)} \) corresponding to
terms in which one of the integers \( r, s \) is zero and the other is plus or minus one. For this case, Eq. (20) is, of course, trivially satisfied, and the phenomenon is only a slight generalization of that of the previous subsection. For the point made here, we shall assume that \( F \) is of such a nature that, to the order we are going, there exists only one pair of integers \( r, s \) satisfying (20), neither of which is zero, and for which the Fourier coefficient corresponding to these integers on the right-hand side is nonzero. We are indebted to Dr. J. I<. Hale for a conversation which led to the discovery of an error on this point in our original manuscript.

Seek a solution of the form

\[
f = a(\cos \psi_1 + \cos \psi_2) + u_i(a, \psi_1, \psi_2) + \ldots,
\]

where now the amplitude \( a \) and the phases \( \psi_1 \) and \( \psi_2 \) develop according to

\[
\frac{\partial a}{\partial t} = \epsilon A(a, \psi_1, \psi_2) + \ldots,
\]

\[
\frac{\partial a}{\partial x} = \epsilon B(a, \psi_1, \psi_2) + \ldots,
\]

\[
\frac{\partial \psi_{1,2}}{\partial t} = -\omega_{1,2} + \epsilon C_{1,2}(a, \psi_1, \psi_2) + \ldots,
\]

\[
\frac{\partial \psi_{1,2}}{\partial x} = K_{1,2} + \epsilon D_{1,2}(a, \psi_1, \psi_2) + \ldots. \tag{21}
\]

The left-hand side of (2) may again be computed, using the relations (21), to give

\[
\frac{\partial^2 f}{\partial x^2} - c^2 \frac{\partial^2 f}{\partial \psi_1^2} + \lambda^2 f = a\left(-\omega_1^2 + c^2 K_1^2 + \lambda^2\right) \cos \psi_1
\]

\[
+ (-\omega_2^2 + c^2 K_2^2 + \lambda^2) \cos \psi_2
\]

\[
+ \epsilon \left(\sin \psi_1 \left[2\omega_1 A + \omega_1 \frac{\partial C_1}{\partial \psi_1} + \omega_2 \frac{\partial C_1}{\partial \psi_2}\right]
\]

\[
+ 2c^2 K_1 B + ac^2 K_1 \frac{\partial D_1}{\partial \psi_1} + ac^2 K_2 \frac{\partial D_1}{\partial \psi_2}\right]
\]

\[
+ \sin \psi_2 \left[2\omega_2 A + \omega_1 \frac{\partial C_2}{\partial \psi_1} + \omega_2 \frac{\partial C_2}{\partial \psi_2}\right]
\]

\[
+ 2c^2 K_2 B + ac^2 K_1 \frac{\partial D_2}{\partial \psi_1} + ac^2 K_2 \frac{\partial D_2}{\partial \psi_2}\right]
\]

\[
+ \cos \psi_1 \left[2\omega_1 C_1 - \omega_1 \frac{\partial A}{\partial \psi_1} + \omega_2 \frac{\partial A}{\partial \psi_2}\right]
\]

\[
+ 2ac^2 K_1 \frac{\partial B}{\partial \psi_1} + \epsilon c^2 K_1 \frac{\partial B}{\partial \psi_2}\right]
\]

\[
+ \cos \psi_2 \left[2\omega_2 C_2 - \omega_1 \frac{\partial A}{\partial \psi_1} + \omega_2 \frac{\partial A}{\partial \psi_2}\right]
\]

\[
+ 2ac^2 K_2 \frac{\partial B}{\partial \psi_1} + \epsilon c^2 K_2 \frac{\partial B}{\partial \psi_2}\right]
\]

\[
+ \cos \psi_1 \left[2\omega_1 C_1 - \omega_1 \frac{\partial A}{\partial \psi_1} + \omega_2 \frac{\partial A}{\partial \psi_2}\right]
\]

\[
+ 2ac^2 K_1 \frac{\partial B}{\partial \psi_1} + \epsilon c^2 K_1 \frac{\partial B}{\partial \psi_2}\right]
\]

\[
+ \cos \psi_2 \left[2\omega_2 C_2 - \omega_1 \frac{\partial A}{\partial \psi_1} + \omega_2 \frac{\partial A}{\partial \psi_2}\right]
\]

\[
+ 2ac^2 K_2 \frac{\partial B}{\partial \psi_1} + \epsilon c^2 K_2 \frac{\partial B}{\partial \psi_2}\right]
\]

\[
+ \left\{ e \left( \frac{\partial u_1}{\partial \psi_1} + \frac{\partial u_1}{\partial \psi_2} + u_1 \right) \right.
\]

\[
+ 2(\omega_1 \omega_2 - c^2 K_1 K_2) \frac{\partial^2 u_1}{\partial \psi_1 \partial \psi_2}\right\} + O(\epsilon^3). \tag{22}
\]

The zeroth order part of (22) vanishes identically. The first order part now must be equated to \( \partial^2 F(a(\cos \psi_1 + \cos \psi_2), a(\omega_1 \sin \psi_1 + \omega_2 \sin \psi_2), -a(K_1 \sin \psi_1 + K_2 \sin \psi_2)) \). It is most convenient to write this as a complex Fourier series,

\[
F = \sum_{m,n=-\infty}^{\infty} F_{mn}(a)e^{i(m\psi_1 + n\psi_2)}, \tag{23}
\]

where \( F_{mn}^* = F_{-m,-n} \), since we deal only with real quantities. Calling everything in the first curly bracket in Eq. (22), \( G(a; A; B; C_{1,2}; D_{1,2}) \), we may finally write the equation for \( u_i \) as

\[
\lambda^2 \left( \frac{\partial^2 u_1}{\partial \psi_1^2} + \frac{\partial^2 u_1}{\partial \psi_2^2} + u_1 \right) + 2(\omega_1 \omega_2 - c^2 K_1 K_2) \frac{\partial^2 u_1}{\partial \psi_1 \partial \psi_2} =
\]

\[
- G(a; A; B; C_{1,2}; D_{1,2})
\]

\[
+ \sum_{m,n=-\infty}^{\infty} F_{mn}(a)e^{i(m\psi_1 + n\psi_2)}. \tag{24}
\]

Equation (24), analogously to Eq. (10), has a readily obtainable, secularity free solution of the form

\[
u_i(\psi_1, \psi_2, a) = \sum_{m,n=-\infty}^{\infty} u_{mn}(a)e^{i(m\psi_1 + n\psi_2)} \tag{25}\]

if and only if the functions \( A, B, C_{1,2}, D_{1,2} \) are chosen to depend on \( \psi_1 \) and \( \psi_2 \) in such a manner that:

\[
\oint d\psi_1 \oint d\psi_2 \exp \left[-i(f\psi_1 + \bar{s}\psi_2)\right]
\]

\[
\cdot G(a; A; B; C_{1,2}; D_{1,2}) - F(a) = 0. \tag{26}\]

Since (26) involves complex numbers, it really amounts to two real equations upon \( A, B, C_{1,2}, D_{1,2} \), which are linear partial differential equations with periodic coefficients. (We may consider \( A, B, \ldots \) as being expressed as Fourier series in \( \psi_1, \psi_2 \); in this case, (26) leads to algebraic equations for the Fourier coefficients.) We need not write them down in full detail, since they are most cumbersome, and not of interest for our purposes here.

Three additional conditions may be added from the requirements that \( \partial^2 a/\partial x \partial t = \partial^2 a/\partial t \partial x \), \( \partial^2 \psi_{1,2}/\partial x \partial t = \partial^2 \psi_{1,2}/\partial t \partial x \); they are

\[
K_1 \frac{\partial A}{\partial \psi_1} + K_2 \frac{\partial A}{\partial \psi_2} = -\left(\omega_1 \frac{\partial B}{\partial \psi_1} + \omega_2 \frac{\partial B}{\partial \psi_2}\right), \tag{27a}
\]

\[
K_1 \frac{\partial D_1}{\partial \psi_1} + K_2 \frac{\partial D_1}{\partial \psi_2} = -\left(\omega_1 \frac{\partial C_1}{\partial \psi_1} + \omega_2 \frac{\partial C_1}{\partial \psi_2}\right), \tag{27b}
\]
Satisfaction of these five relations on the six quantities $A, B, C_{1,2}, D,,,,$ then, will guarantee that the solution shall be secularly free. It is clear that considerable latitude is left to choose the functions conveniently for whatever problem is under consideration. Nor does secularity necessarily have a simple physical interpretation as it does for Eq. (1), where the presence of secular terms is simply interpretable as a resonance between one of the frequencies present in the nonlinear coupling term and one of the natural frequencies of the system. It would take considerably more insight to apprehend the physical meaning of Eq. (20) for the case in which neither $F$ nor $s$ is zero, or any connection it might have with the secularity condition for Eq. (10).

III. THE COLD ELECTRON PLASMA

Suppose we consider a cold (no thermal motions) electron plasma of equilibrium number density $n_0$, moving in a uniform positive background, assumed immobile. If we make a perturbation about a uniform, field-free equilibrium, the appropriate variables for describing the system are $v$, the electron velocity, $-e[n_0 + n(x, t)]$, the electron charge density, $+en_0$, the positive background charge density, and the electric and magnetic fields $E, B$. All these will be treated as perturbations—i.e., first order in the amplitude—except for $n_0$.

The dynamical equations are well known:

$$
\frac{\partial n}{\partial t} + \frac{\partial}{\partial x}(nv) + \frac{\partial}{\partial x}(n v) = 0, \quad (28a)
$$

$$
\frac{\partial v}{\partial t} + v \frac{\partial}{\partial x} v = -\frac{e}{m} \left( E + \frac{1}{c} v \times B \right), \quad (28b)
$$

$$
\frac{1}{c} \frac{\partial B}{\partial t} = -\frac{\partial}{\partial x} \times E, \quad (28c)
$$

$$
\frac{1}{c} \frac{\partial E}{\partial t} = \frac{\partial}{\partial x} B + \frac{4\pi e}{c} (n_0 + n) v; \quad (28d)
$$

$$
\frac{\partial}{\partial x} E = -4\pi en, \quad (29a)
$$

$$
\frac{\partial}{\partial x} \cdot B = 0. \quad (29b)
$$

Equations (29) can be regarded as initial conditions; once fulfilled, they are preserved by Eqs. (28).

A small amount of algebraic juggling shows that Eqs. (28) lead to

$$
\frac{\partial^2}{\partial t^2} E + c^2 \frac{\partial}{\partial x} \times \left( \frac{\partial}{\partial x} \times E \right) + \omega_p^2 E = -\epsilon 4\pi e
$$

$$
\left[ n_0 \left( v \cdot \frac{\partial}{\partial x} v + \frac{e}{mc} v \times B \right) - \frac{\partial}{\partial t} (nv) \right], \quad (30a)
$$

$$
\frac{\partial^2}{\partial t^2} B + c^2 \frac{\partial}{\partial x} \times \left( \frac{\partial}{\partial x} \times B \right) + \omega_p^2 B = \epsilon \left[ 4\pi en_0 \frac{\partial}{\partial x} \times \left( \frac{\partial}{\partial x} v + \frac{e}{mc} v \times B \right) \right.
$$

$$
\left. - 4\pi e c \frac{\partial}{\partial x} \frac{\partial}{\partial t} (nv) \right], \quad (30b)
$$

where $\omega_p^2 = 4\pi n_0 e^2 / m$ is the plasma frequency. The formal expansion parameter $\epsilon$ has been written on the right-hand side of Eqs. (30) only to remind us that these terms are “small” in the sense of being second order in the amplitude.

If we now restrict ourselves to disturbances which are functions of only one spatial dimension ($z$, say), we may write the expression (30a) in the form

$$
\left( \frac{\partial^2}{\partial z^2} - c^2 \frac{\partial^2}{\partial x^2} + \omega_p^2 \right) E_i = \epsilon \delta_i, \quad (31)
$$

where $\delta_i = 0$ unless $i = 1$ and $\delta_{1,1} = 1$. $\delta_i$ represents the $i$th component of the right-hand side of (30a). Since Eq. (30b) is just the curl of (30a), it is not necessary to write it in the form of (31). It matters not a bit that we have not bothered to express the right-hand side of (30a) as a function of $E$ alone, since we always need the values of the quantities which enter into $\delta_i$ to one lower order than those under consideration on the left-hand side of (30a), and these will be shown to be obtainable in each order directly from (28).

The $\epsilon = 0$ solution to (31) is just the standard cold-plasma set of field-free normal modes, which is well understood. It will be apparent that each component of (31) has the form of Eq. (2) and is therefore immediately susceptible to the methods of Sec. II. Equation (31) represents a generalization of the equation treated by Jackson, which, however, does not exhibit all the features of (31), due to the absence of $x$-derivatives on the left-hand side in the purely electrostatic case, and the fact that the $x$ and $t$ dependences separate.

The number of possibilities from (31) is very large, due to the wide range of choices for the $\epsilon = 0$ solution. The simplest case (beyond that of a pure electrostatic oscillation) is that of a pure transverse, linearly polarized, electromagnetic wave. We shall
extend this solution to the next two orders above the linear approximation in the following subsection.

A. Plane Electromagnetic Wave

Calling our basis vectors \( \hat{e}_x, \hat{e}_y, \hat{e}_z \), the appropriate zeroth order solution is

\[
E = E_0 \hat{e}_x \cos(k_0 x - \omega_0 t),
\]

\[
B = \frac{c k_0}{\omega_0} E_0 \hat{e}_z \cos(k_0 x - \omega_0 t),
\]

\[
v = \frac{c E_0}{m \omega_0} \hat{e}_z \sin(k_0 x - \omega_0 t), \quad n = 0,
\]

where

\[
\omega_0^2 = \omega_{se}^2 + c^2 k_0^2.
\]

This zeroth order solution for \( E \) is conveniently written:

\[
E_i = \phi \hat{e}_x e^{i(k_0 x - \omega_0 t)} + \bar{\phi} \hat{e}_x e^{-i(k_0 x - \omega_0 t)}
\]

where \( \phi \) is a column vector

\[
\phi = \begin{bmatrix} 0 \\ 1 \\ 0 \end{bmatrix},
\]

and \( a = E_0 \), the amplitude.

The results of Sec. II suggest that we seek a solution of the form

\[
E_i = a_{\phi}(\phi, e^{i(k_0 x - \omega_0 t)} + \bar{\phi} \hat{e}_x e^{-i(k_0 x - \omega_0 t)}) + \varepsilon u_i(1)(a, \psi)
\]

+ \varepsilon^2 u_i(2)(a, \psi) + \cdots,

(35)

where now the amplitude \( a \) and phase \( \psi \) are to vary in \( x \) and \( t \) according to

\[
\frac{\partial a}{\partial t} = \epsilon A_1(a) + \epsilon^2 A_2(a) + \cdots,
\]

\[
\frac{\partial a}{\partial x} = \epsilon D_1(a) + \epsilon^2 D_2(a) + \cdots,
\]

\[
\frac{\partial \psi}{\partial t} = -\omega_0 + \epsilon B_1(a) + \epsilon^2 B_2(a) + \cdots,
\]

\[
\frac{\partial \psi}{\partial x} = k_0 + \epsilon C_1(a) + \epsilon^2 C_2(a) + \cdots,
\]

with the functions \( A_1, B_1, \cdots \), as yet undetermined.

The left-hand side of (31), expressed in terms of these functions, becomes:

for \( i = 1 \),

\[
\left( \frac{\partial^2}{\partial t^2} + \omega_{se}^2 \right) E_1 = \epsilon \left( \omega_{se}^2 \frac{\partial^2 u_i(1)}{\partial \psi^2} + \omega_{se}^2 u_i(1) + 2(-i \omega_0 A_1 + a_0 B_1) \phi_0 e^{i\psi} + 2(i \omega_0 A_1 + a_0 B_1) \bar{\phi} e^{-i\psi} \right) + O(\varepsilon^2);
\]

(37)

for \( i = 2 \) or \( 3 \),

\[
\left( \frac{\partial^2}{\partial t^2} + \omega_{se}^2 \right) E_i = \epsilon \left( \omega_{se}^2 \frac{\partial^2 u_i(1)}{\partial \psi^2} + \omega_{se}^2 u_i(1) + 2(-i \omega_0 A_1 + a_0 B_1) \phi_0 e^{i\psi} + 2(i \omega_0 A_1 + a_0 B_1) \bar{\phi} e^{-i\psi} \right) + O(\varepsilon^2);
\]

(38)

These three expressions are to be equated to the three components of the right hand sides of Eq. (31), with the quantities \( v, E, B, n \) replaced by their values (32), and with the substitution \( E_0 \rightarrow a \). The result can be conveniently written as a Fourier series

\[
S_i = \sum_{m=-\infty}^{\infty} \Phi_i(m, a) e^{im\psi},
\]

(39)

where the \( \Phi_i(m, a) \) are known functionals of \( a \).

Suppose we now consider whether secular terms can arise. We seek the solution for \( u_i(1) \) in the form

\[
u_i(1) = \sum_{m=1}^{\infty} \nu_i(1)(m, a) e^{im\psi}.
\]

(40)

The equation for \( u_i(1) \) then becomes

\[
\sum_{m=-\infty}^{\infty} \left[ (-m^2 + 1) \omega_{se}^2 + m^2 \delta_1 \left( \omega_{se}^2 - \omega_0^2 \right) \right] \nu_i(1)(m, a) e^{im\psi} = -2(-i \omega_0 A_1 + a_0 B_1 - i e^2 k_0 D_1 + a e^2 C_1 k_0) \phi_0 e^{i\psi}
\]

- \( 2(i \omega_0 A_1 + a_0 B_1 + i e^2 k_0 D_1 + a e^2 C_1 k_0) \bar{\phi} e^{-i\psi} \)

+ \sum_{m=-\infty}^{\infty} \Phi_i(m, a) e^{im\psi},
\]

(41)

(we have made use of \( \phi_0 = 0 \)).

We may always solve (41) for \( u_i(1)(m, a) \) with the absence of \( \psi \)-proportional terms, except when a coefficient of \( u_i(1)(m, a) \) on the left hand side of (41) vanishes for some \( m \). Since \( \omega_{se}^2 > \omega_0^2 \) it is clear that this can only occur when \( m = \pm 1 \), and only in the components \( i = 2 \) or \( 3 \). However, a term-by-term inspection of \( \Phi_i(m, a) \) reveals that for \( i = 2 \) or \( 3 \), the quantities \( \Phi_i(m, a) \) vanish identically for the choice (33) of the zeroth order values. Thus a result emerges which would have been hard to guess from the original set of equations (28): secular behavior cannot arise in second order for this case, and we may set \( A_1 = B_1 = C_1 = D_1 = 0 \), and solve in a completely straightforward way for \( u_i(1)(m, a) \)

\[
u_i(1)(m, a) = \Phi_i(m, a) \left[ (-m^2 + 1) \omega_{se}^2 + m^2 \delta_1 \left( \omega_{se}^2 - \omega_0^2 \right) \right]^{-1}, \quad m \neq \pm 1,
\]

\[
u_i(1)(\pm 1, a) = 0.
\]

(42)
The calculation of the coefficients of the Fourier coefficients is a matter of simple algebra, and we may write down in full the first correction to the linear solution (32):

\[ u^{(1)} = E^{(1)} = \frac{e_k \omega_{pe}}{2m \omega_0^2} \sin(2k_0x - \omega_0 t) \]

\[ B^{(1)} = 0, \]

\[ n^{(1)} = -\frac{E_0^2 e^{i \theta} k_0^2}{m^2 \omega_0^2} \cos(2k_0x - \omega_0 t) \]

\[ v^{(1)} = -\frac{E_0^2 e^{i \theta} k_0^2}{m^2 \omega_0^2} \cos(2k_0x - \omega_0 t) \]

The only qualitatively new feature which shows up in \(O(\varepsilon)\), then, is a longitudinal electric field. It is only in \(O(\varepsilon^2)\) that secularity manifests itself and it becomes necessary to use the Bogoliubov methods. Making use of the \(O(\varepsilon)\) solution we have just computed, we may write down the \(O(\varepsilon^2)\) part of the left-hand side of (31) in terms of the \(A's, B's, etc.\)

\[ \frac{\partial^2 E_i}{\partial t^2} - c^2 \frac{\partial^2 E_i}{\partial x^2} + \omega_i^2 E_i \]

\[ = \varepsilon^2 \left( \frac{\omega_i^2 - c^2 k_0^2}{\partial \psi^2} \right) + \omega_i^2 u_i^{(2)} \]

\[ + (-2i\omega_0 A_2 + 2a_0 B_2 - 2ik_0 \omega^2 D_2) \]

\[ + 2ac^2 k_0 C_2 \phi e^{i\psi} + (2i\omega_0 A_2 + 2a_0 B_2) \]

\[ + 2ik_0 \omega^2 D_2 + 2ac^2 k_0 C_2 \phi e^{-i\psi} \].

We may find a solution to (46) of the form

\[ u_i^{(2)}(a, \psi) = \sum_{n=-\infty}^{\infty} u_i^{(2)}(a, m) e^{in\psi}, \]

which is free of \(\psi\)-proportional terms, if and only if the coefficients of \(e^{i\psi}\) and \(e^{-i\psi}\) on the right-hand side of (46) vanish. This gives us two conditions on \(A_2, B_2, C_2, D_2\), which are algebraic relations. Two more, analogous to Eqs. (13), are given by the requirements that \(\partial^2 a/\partial t \partial x = \partial^2 a/\partial t \partial t\) and \(\partial^2 \psi/\partial t \partial x = \partial^2 \psi/\partial t \partial t\); they are

\[ D_2(dA_2/da) = A_2(dD_2/da), \]

\[ D_2(dB_2/da) = A_2(dC_2/da). \]

Let us now specialize the problem to one in which the spatial periodicity and amplitude are given, and we are to calculate the "frequency shift." This means setting \(C_2 = D_2 = 0\). The condition that the coefficients of \(e^{i\psi}\), \(e^{-i\psi}\) on the right-hand side of (46) vanish reduces to the equation

\[ (i\omega_0 A_2 - a_0 B_2) = \frac{\omega_i^2 e^{2i\psi} \phi}{4m^2 \omega_0^2 (3a_0^2 + 4c^2 k_0^2)} \]

and its complex conjugate relation. We may therefore find \(A_2 = 0\), and the frequency shift \(\Delta \omega\) becomes (substituting \(E_i\) for the amplitude \(a\))

\[ \Delta \omega = -B_2 = \frac{\varepsilon^2 k_0 \omega_{pe}^2}{4m \omega_0^2 (3a_0^2 + 4c^2 k_0^2)} \cdot E_i^2. \]

It does not seem worthwhile to write down the \(O(\varepsilon^2)\) corrections to \(E, B, v, \) and \(n\), though it would be easy to do so. In any experiment one might imagine, Eq. (50) would probably be the easiest quantity to measure.

### B. Standing Wave*

It is possible to calculate the frequency shift of a standing electromagnetic wave of given periodicity as well. One assumes a linearly polarized standing wave solution for (31) in lowest order and writes

\[ E_i = a_i \sin k_0 x \cos \psi \]

\[ + eu_i^{(1)}(a, \psi, x) + \varepsilon^2 u_i^{(2)}(a, \psi, x) + \cdots \]

Then, only the \(t\)-dependence of the phase variable and amplitude are expanded:

\[ da_i/da = \epsilon A_1(a) + \epsilon^2 A_2(a) + \cdots, \]

\[ d\psi/da = \omega_0 + \epsilon B_1(a) + \epsilon^2 B_2(a) + \cdots. \]

In lowest significant order, the frequency shift turns

* This section added in proof.
out to be exactly one eighth of the result given in Equation (50).

IV. DISCUSSION

We have given a technique for obtaining uniformly valid, perturbation theoretic solutions to the Klein–Gordon equation with a small nonlinear term. The method has been applied to calculate the second order frequency shift of an electromagnetic wave in a cold electron plasma. The smallness of the expression (50) for attainable parameters is an indication of just how good an approximation the linear theory is at these frequencies.

It is not to be inferred, however, that the method adapts itself readily to all partial differential equations. For instance, the reader can easily convince himself that it fails for a nonlinear sound wave described by the Euler equations. [It appears to fail in all situations for which the \( \epsilon = 0 \) equation is a wave equation, \( (\partial^2/\partial t^2 - \epsilon^2 \partial^2/\partial x^2)f = 0 \).] The physical reason is that, due to a steepening of the exact nonlinear wave front—obtainable from the Riemann invariants—a vertical tangent develops after a time of \( O(1/\text{amplitude}) \). This destroys any regularity properties which may have existed in the original wave profile. Since the exact solution does not remain “close” to the \( \epsilon = 0 \) solution in any sense, after a time of order \( 1/\epsilon \), it is not surprising that perturbation theory is of little use beyond this time.