Orbital Electron Capture by the Nucleus*

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The theory of nuclear electron capture is reviewed in the light of current understanding of weak interactions. Experimental methods and results regarding capture probabilities, capture ratios, and EC/\beta^+ ratios are summarized. Radiative electron capture is discussed, including both theory and experiment. Atomic wave-function overlap and electron exchange effects are covered, as are atomic transitions that accompany nuclear electron capture. Tables are provided to assist the reader in determining quantities of interest for specific cases.
CONTENTS

1. Introduction .................................................. 7
   1.1 History ....................................................... 7
   1.2 Energetics .................................................. 8
   1.3 Atomic Effects ............................................. 10
   1.4 Radiative Electron Capture ...................................... 11
   1.5 Significance .................................................. 12
   1.6 Scope of Review ............................................. 12

2. Electron-Capture Theory ........................................ 14
   2.1 The β-Decay and Electron-Capture Hamiltonian and Transition Rates ............................................. 14
   2.2 Electron-Capture Transition Rates .................................. 30
      2.2.1 General Relations for the Transition Probabilities .................................. 30
      2.2.2 Bound-State Electron Radial Wave Functions ....................................... 33
      2.2.3 Nuclear Form Factors and Nuclear Matrix Elements .................................. 44
      2.2.4 Explicit Expressions for the Quantities $m_K(k_x, k_y)$ and $m_K(k_x', k_y')$ .................................. 66
   2.3 Formulae for Allowed and Forbidden Transitions .................................. 69
      2.3.1 Allowed Transitions ........................................ 69
      2.3.2 First-Forbidden Non-Unique Transitions ................................... 70
      2.3.3 First-Forbidden Unique Transitions ....................................... 71
      2.3.4 $(L-1)$-Forbidden Unique Transitions ....................................... 72
      2.3.5 Some General Remarks on Higher-Forbidden Non-Unique Transitions ............... 73
2.4 Electron-Capture to Positron-Decay Ratios: 74
   2.4.1 General Expressions: 74
   2.4.2 Allowed Transitions: 77
   2.4.3 Non-Unique Forbidden Transitions: 81
   2.4.4 Unique Forbidden Transitions: 81

2.5 Atomic Matrix Elements: Exchange and Overlap
   Corrections: 83
   2.5.1 Introduction: 83
   2.5.2 Effect of Atomic Overlap and Exchange on
      Total Capture Rates: 84
   2.5.3 Overlap and Exchange Corrections on Capture
      Ratios: 86
   2.5.4 Evaluation of Atomic Matrix Elements: 93
   2.5.5 Comparison Among Theoretical Exchange
      Corrections to Capture Ratios: 95
   2.5.6 Correlation Effects in Electron-Capture Ratios: 96
   2.5.7 Conclusion: 96

3. Experimental Methods and Results: 98
   3.1 Determination of Capture Ratios: 101
      3.1.1 Spectrometry with Internal Gas Sources: 102
      3.1.2 Spectrometry with Internal Solid Sources: 109
      3.1.3 Spectrometry of K and L x Rays with External
         Sources: 112
   3.2 Determination of the Relative K-Capture Probability \( P_K \): 115
      3.2.1 Measurement of K x Rays or Auger Electrons
         and \( \gamma \) Rays of Conversion Electrons: 115
      3.2.2 Coincidence Measurements: 119
<table>
<thead>
<tr>
<th>Section</th>
<th>Title</th>
<th>Page</th>
</tr>
</thead>
<tbody>
<tr>
<td>3.3</td>
<td>Experimental Capture Probabilities $P_K$, $P_L$, and $P_M$</td>
<td>127</td>
</tr>
<tr>
<td>3.3.1</td>
<td>Experimental Results</td>
<td>127</td>
</tr>
<tr>
<td>3.3.2</td>
<td>Theoretical Predictions</td>
<td>129</td>
</tr>
<tr>
<td>3.3.3</td>
<td>Comparison of Experimental and Theoretical Electron-Capture Ratios</td>
<td>131</td>
</tr>
<tr>
<td>3.3.4</td>
<td>Conclusions and Recommendations</td>
<td>135</td>
</tr>
<tr>
<td>3.4</td>
<td>Determination of $K/\beta^+$ and $EC/\beta^+$ Ratios</td>
<td>136</td>
</tr>
<tr>
<td>3.4.1</td>
<td>Measurements of $K/\beta^+$ Ratios with Internal Sources</td>
<td>137</td>
</tr>
<tr>
<td>3.4.2</td>
<td>Measurements of $K/\beta^+$ Ratios with External Sources</td>
<td>140</td>
</tr>
<tr>
<td>3.4.3</td>
<td>Measurements of $EC/\beta^+$ Ratios</td>
<td>144</td>
</tr>
<tr>
<td>3.5</td>
<td>Experimental Results and Comparison with Theory for $K/\beta^+$ and $EC/\beta^+$ Ratios</td>
<td>150</td>
</tr>
<tr>
<td>3.5.1</td>
<td>Results</td>
<td>150</td>
</tr>
<tr>
<td>3.5.2</td>
<td>Theoretical Predictions</td>
<td>151</td>
</tr>
<tr>
<td>3.5.3</td>
<td>Comparison of Experiment and Theory</td>
<td>153</td>
</tr>
<tr>
<td>3.5.4</td>
<td>Conclusions and Recommendations</td>
<td>155</td>
</tr>
<tr>
<td>4.</td>
<td>Radiative Electron Capture</td>
<td>157</td>
</tr>
<tr>
<td>4.1</td>
<td>Theory</td>
<td>157</td>
</tr>
<tr>
<td>4.1.1</td>
<td>Matrix Elements and Transition Rates</td>
<td>159</td>
</tr>
<tr>
<td>4.1.2</td>
<td>IB Spectra from Allowed Transitions</td>
<td>164</td>
</tr>
<tr>
<td>4.1.3</td>
<td>IB Correlation Effects in Allowed Transitions</td>
<td>190</td>
</tr>
<tr>
<td>4.1.4</td>
<td>IB Spectra and Correlation Effects in Forbidden Transitions</td>
<td>201</td>
</tr>
<tr>
<td>4.2</td>
<td>Experiments</td>
<td>209</td>
</tr>
<tr>
<td>4.2.1</td>
<td>Experiments on Total IB Spectra</td>
<td>212</td>
</tr>
<tr>
<td>4.2.2</td>
<td>Experiments on Partial IB Spectra</td>
<td>220</td>
</tr>
</tbody>
</table>
1. INTRODUCTION

1.1 History

In $\beta$ decay, a nucleus can capture an electron (or a positron) instead of emitting one. This possibility, inherent in the Fermi (1934) theory of $\beta$ emission, was first suggested by Yukawa and Sakata (1935, 1936, 1937). The density of atomic bound electrons at the nucleus makes orbital electron capture significant, particularly for $s$ electrons in heavy atoms. Detection of the emitted neutrino is a major experimental undertaking that has not yet met with success (Davis, et al., 1968; see also Physics Today 25, August 1972, p. 17; Bahcall, 1972). Even the nuclear recoil from neutrino emission is very difficult to detect (Crane, 1948), unless extraordinary ingenuity is brought to bear (Goldhaber et al., 1958). X rays and Auger electrons emitted in the deexcitation of the ionized daughter provide more readily detectable, albeit indirect, signals of the capture process.

Alvarez (1937) first gained experimental evidence for the existence of nuclear electron capture by detecting Ti K x-rays emitted in the decay of $^{48}$V. A Geiger counter was employed; positrons were bent away by a magnetic field, and the x-ray energy was established approximately from an Al absorption curve. Gamma-ray internal conversion could not be excluded as a possible origin of the Ti K x-rays. A completely conclusive demonstration was brought about the following year, when Alvarez (1938a, b) used differential absorption to identify
Zn K x rays from the decay of $^{67}$Ga. Related cloud-chamber experiments were performed by Oldenburg (1938) and by Williams and Pickup (1938), after an unsuccessful attempt by Jacobsen (1937). The capture of L electrons was first observed by Kirkwood et al. (1948) and Pontecorvo et al. (1949), who mixed radioactive $^{37}$Ar with the gas in a proportional counter and found a peak due to Cl L x rays in the spectrum. Dougan (1961) first measured M-electron capture in $^{71}$Ge.

Following the work of Fermi (1934) and Yukawa and Sakata (1935, 1936, 1937), the theory of allowed electron capture was developed by Bethe and Bacher (1936) and Møller (1937a, b). Generalizations including forbidden transitions were carried out by Marshak (1942), Bouchez et al. (1950, Bouchez, 1952), Eyrsk and Rose (1958), Hülberd (1965), Robinson (1965), Zwiefel (1954, 1957, 1958), Konopinski (1966), and Behrens and Jänecke (1969), among others. The subject has been reviewed by Robinson and Fink (1955, 1960), Bouchez (1963a, 1968a), and Depommier (1960), and Berényi (1966). Introductions to the theory are contained in the books by Schopper (1966) and Wu and Moszkowski (1966).

1.2. Energetics

We denote by $W_0 + 1$ the energy (mass) difference between parent and daughter neutral atoms:

$$W_0 = \Delta W_{\text{nucl}} - \Delta |E_X|,$$

in units such that $m_e = c = 1$. Here, $\Delta W_{\text{nucl}}$ is the energy difference between the parent nucleus $(A,Z)$ and the daughter nucleus $(A,Z-1)$. The quantity $\Delta |E_X|$ is the total change in electron binding energy.
between parent and daughter atoms, which arises because all electron energy levels move up in the potential well as the nuclear charge decreases by one unit (the electron cloud "expands"). The binding-energy charge $\Delta |\Sigma E_x|$ is not negligible; it amounts to $\approx 20$ keV for $Z=85$, for example.

Let $E_x'$ be the binding energy of the captured electron in the daughter atom. We neglect the energy of atomic recoil from neutrino emission; its largest value, in $^7$Be decay, is only 57 eV. Because of imperfect atomic wave-function overlap, the daughter atom's electronic excitation energy will exceed $|E_x'|$ by an amount that we denote by $E_R$. The average of this rearrangement energy $E_R$, taken over many atoms, is small (of the order of a few eV), but in those individual transitions in which substantial shakeup or shakeoff (internal ionization) occurs, $E_R$ can be quite significant (Sec. 5). The neutrino energy is

$$q = W_o + 1 - |E_x'| - E_R$$

or

$$q = \Delta W_{\text{nucl}} - \Delta |\Sigma E_x'| + 1 - |E_x'| - E_R.$$  \hfill (1-3)

The atomic excitation energy $|E_x'| + E_R$ is released after the capture event in a cascade of Auger and radiative transitions, except for energy carried into the continuum in shakeoff. The energy threshold for electron capture from orbital $x$ is

$$\Delta W_{\text{nucl}} \geq 1 + \Delta |\Sigma E_x'| + |E_x'| + E_R.$$  \hfill (1-4)

Positron emission is energetically possible, and competes with
orbital electron capture (Secs. 3.4, 3.5) if $W_{\text{nucl}}$, or

$$\Delta W_{\text{nucl}} \geq 1 + \Delta |\Sigma E| + E_R.$$  \hfill (1-5)

1.3. Atomic Effects.

Nuclear electron capture by its very nature stands at the interface between nuclear and atomic physics. Only in the crudest of approximations can the atomic electron cloud be treated as merely the donor of the electron that is captured. Nevertheless, the importance of treating $\beta$ decay in general, and electron capture in particular, as transformations of the whole atom was not quantitatively taken into account until Benoist-Gueutal (1950, 1953a, b) wrote her thesis. The idea of including atomic variables in the description of initial and final states was pursued by Odint and Daudel (1956) and formulated elegantly by Bahcall (1962a, 1963a, b).

The fact that the entire atom is transformed in electron-capture decay is reflected in the energetics (Sec. 1.2) and in the effect of imperfect atomic wave-function overlap on the transition rate (Sec. 2.5). Furthermore, atomic transitions such as shakeup and shakeoff (internal ionization) can take place as an integral part of the radioactive decay (Sec. 5), quite distinct from the Auger and x-ray cascade through which the daughter atom is subsequently deexcited. Atomic effects in nuclear decay have recently been reviewed by Emery (1972), Crasemann (1973), Freedman (1974), and Walen and Briançon (1975).
1.4. Radiative Electron Capture

The existence of a low-intensity continuous photon spectrum accompanying $\beta^\pm$ decay was first observed by Aston (1927) and Bramson (1930). The basic theory of radiative $\beta$ decay was developed independently by Knipp and Uhlenbeck (1936), who were seeking an explanation for the observed photon continuum, and by Bloch (1936), who was unaware of the experimental work and was motivated by purely theoretical considerations based on Fermi's theory of $\beta$ decay and Dirac's theory of the positron. Möller (1937a, b) and Morrison and Schiff (1940) pointed out that internal bremsstrahlung (IB) should be emitted in the course of nuclear electron capture as well as in $\beta$ decay, and independently worked out the theory. Möller (1937a, b) in particular, was interested in differentiating between the Fermi and Konopinski-Uhlenbeck couplings. Internal bremsstrahlung from electron capture was first detected by Bradt et al. (1946). A number of reports followed, describing the observation of IB at high energies; all of these data were consistent with the Morrison-Schiff theory. A study of the $^{55}$Fe IB spectrum by Madansky and Rasetti (1954), however, showed an unexpected steep rise of the IB intensity at low photon energies. These data were only explained after Glauber and Martin (1956; Martin and Glauber, 1958) developed an elaborate and much more accurate theory of IB in electron capture, in which Coulomb and screening effects are taken into account and capture from L and M shells is included. Although originally restricted to allowed transitions, this theory was later generalized to electron-capture transitions
of arbitrary degree of forbiddenness by Zon and Rapoport (1968; Zon, 1971).

1.5. Significance

Research on electron-capture probabilities and ratios is being pursued as a facet of basic science and because of the importance of applications. Electron capture plays a part in the decay schemes of some 500 radionuclides, ~60 of which are commercially available. Nuclear decay by electron capture is not only relevant to nuclear science but also to geochemistry, cosmology and astrophysics (Trimble and Reines, 1973), nuclear medicine (Dillman, 1968, 1970), and technology. The measurement of $\lambda/\beta^+$ ratios is one of the more sensitive ways of determining upper limit on the Fierz interference term (Schopper, 1968). Ratios of allowed electron capture from various shells are independent of nuclear factors and reflect purely atomic properties; these ratios are sensitive to bound-electron wave functions at the nuclear surface and to electron exchange and imperfect atomic wave-function overlap (Bahcall, 1962a, 1963a, b).

1.6. Scope of Review

In Sec. 2 of this article, we discuss the theory of allowed and forbidden nuclear electron-capture. Formulae and tables are provided that enable the reader to calculate transition rates and ratios of interest. Special attention is paid, in Sec. 2.5, to electron-exchange and atomic wave-function overlap effects on the transition probability. Experimental methods for the measurement of electron-
capture probabilities and ratios of $EC/\alpha$ ratios are described and compared in Sec. 3. Published data are listed, critically evaluated, and compared with theory. In Sec. 4, the theory of radiative electron capture and experimental work on internal bremsstrahlung are thoroughly reviewed and tables for the calculation of IB spectra are provided. Section 5 is devoted to a discussion of atomic transitions that accompany nuclear electron capture.

We have made an effort at completeness in covering the subject. Some information has been included that is now of merely historical interest, but we have attempted to be adequately critical in the final evaluation and comparison of results. Meson capture, though interesting and closely related to our subject, has not been included.

We hope that this article may prove useful for both theoretical and experimental researchers in need of a complete survey of what is known about nuclear electron capture, and that it will be of help to nuclear physicists and chemists and to workers in radionuclide metrology, nuclear medicine, and in related areas.
2. ELECTRON-CAPTURE THEORY

2.1. The β Decay and Electron-Capture Hamiltonian and Transition Rates

It is usually assumed that all the weak interaction processes can be described by a universal fundamental Hamiltonian density (current-current interaction) (Marshak et al., 1969; Schopper, 1966; Blin-Stoyle, 1973). A general discussion of such phenomenological interaction currents in nuclear systems is given by Lock et al. (1974). For the special case of nuclear β decay, this Hamiltonian density has the form

\[ H_\beta(x) = -G_\beta x^{1/2} \left[ J_\mu(x)L_\mu(x) + h.c. \right] \] (2-1)

where \( J_\mu \) and \( L_\mu \) denote the hadron and the lepton current, respectively. The β-decay coupling constant \( G_\beta \) is related to the universal weak coupling constant \( G \) by

\[ G_\beta = G \cos \theta, \] (2-2)

where \( \theta \) is the Cabbibo angle.

Although Eq. (2-1) well describes such processes as β and μ decay, it represents an incomplete theory because it is not renormalizable. Thus, higher-order corrections cannot be calculated. In the last few years, however, renormalizable models (first proposed by Weinberg and Salam, 1968) have been developed.
These models are based on gauge theories unifying the weak and electromagnetic interactions (Abers and Lee, 1973; Lee, 1973; Bernstein, 1974; Weinberg, 1974; Beg and Sirlin, 1974). These gauge theories imply that the weak interaction operates through a neutral current in addition to the previously known charged current. Phenomena induced by neutral currents occur mostly in high-energy physics, but they can be found in atomic physics as well (Bouchiat and Bouchiat, 1974). Nevertheless, for the purposes of the present paper, the Hamiltonian of Eq. (2-1) is sufficient and we shall deal only with this form of the weak-interaction theory.

In nuclear $\beta$ decay, we must consider the three processes

\[
(Z,A) \rightarrow (Z+1,A) + e^- + \bar{\nu}_e \quad (\beta^- \text{ decay})
\]

\[
(Z,A) \rightarrow (Z-1,A) + e^+ + \nu_e \quad (\beta^+ \text{ decay})
\]

\[
(Z,A) + e^- \rightarrow (Z-1,A) + \nu_e \quad \text{(electron capture)}.
\]

Here, $(Z,A)$ signifies an atomic nucleus of mass number $A$ and atomic number $Z$, $e^-$ denotes an electron and $e^+$, a positron, $\nu_e$ is the neutrino, and $\bar{\nu}_e$, the antineutrino.

In order to discuss the general features of these weak-interaction processes and their interrelations, we first consider the decay of a single neutron or proton, assuming that the individual nucleons in the nucleus are independent of one another and behave like free particles.

In the case of nuclear $\beta$ decay, we need only the electron part of the lepton current, which can be expressed as

\[
L_\mu(x) = i \bar{\psi}_e(x) \gamma_\mu (1+i\gamma_5) \psi_e(x),
\]

where $\psi_e$ and $\bar{\psi}_e$ are the field operators and $\gamma_\lambda$ the Dirac matrices.
The nucleons, unlike the leptons, interact strongly as well. This leads to complications, and consequently it is not possible to express the hadron current so simply in terms of field operators (Marshak et al., 1969; Blin-Stoyle, 1973). If, however, we approximately treat the nucleons as point particles, neglecting the influence of the strong interaction, then the hadron current is

\[ J_\mu(x) = i\bar{\psi}_p \gamma_\mu (1 + \lambda \gamma_5) \psi_n \]  

where \( \lambda = -C_A/C_V = 1.251 \pm 0.009 \) (Kropf and Paul, 1974). The Hamiltonian density then has the form

\[ H_\beta(x) = G_\beta^{2-1/2} \{ i\bar{\psi}_p \gamma_\mu (1 + \lambda \gamma_5) \psi_n (x) \bar{\psi}_e(x) \gamma_\mu (1 + \gamma_5) \psi_e(x) + \text{h.c.} \} \]  

The corresponding transition matrix elements for the three basic processes in nuclear \( \beta \) decay are

\[ n \rightarrow p + e^- + \bar{\nu}_e \quad M_{n-} = \langle p \bar{\nu}_e | \int H_\beta(x) d^4x | n \rangle \]  

\[ p \rightarrow n + e^+ + \nu_e \quad M_{p+} = \langle n e^+ \nu_e | \int H_\beta(x) d^4x | p \rangle \]  

\[ p + e^- \rightarrow n + \nu_e \quad M_{EC} = \langle n \nu_e | \int H_\beta(x) d^4x | p e^- \rangle \]  

With \( H_\beta(x) \) according to Eq. (2-5), the transition matrix elements become

\[ M_{n-} = G_\beta^{2-1/2} (2\pi)^4 \delta(q_p + q_e - q_n) \bar{\psi}_p \gamma_\mu (1 + \lambda \gamma_5) u_n \bar{\psi}_e \gamma_\mu (1 + \gamma_5) \psi_e \]  

(2-7a)
\[
M_{\text{p}^+} = G_p^2 \frac{1}{2(2\pi)^4} \delta(q_n - q_e + q_v - q_e) \left[ \overline{u}_n \gamma_\mu (1 + \gamma_5) u_p \right] \left[ \overline{u}_p \gamma_\mu (1 + \gamma_5) v_e \right]
\]

(2-7b)

\[
M_{\text{p}^+} = G_p^2 \frac{1}{2(2\pi)^4} \delta(q_n + q_v - q_e) \left[ \overline{u}_n \gamma_\mu (1 + \gamma_5) u_p \right] \left[ \overline{u}_p \gamma_\mu (1 + \gamma_5) v_e \right].
\]

(2-7c)

The q's are the four-momenta of the particles indicated by the subscripts, and \( \delta(q) \) is the Dirac delta function.

Equations (2-7) have been derived for the decay of a single, point-like nucleon. To consider the decay of a nucleon in a complex nucleus, we transform the wave function used in Eqs. (2-7) from momentum space to configuration space. For this purpose, the 3-dimensional momentum-dependent part of the delta function is replaced by

\[
\delta(p) = (2\pi)^{-3} \int \frac{d^3r}{d^3p} \frac{i\mathbf{p} \cdot \mathbf{r}}{(2\pi)^3}.
\]

(Blin-Stoyle, 1973). We introduce the plane-wave solutions of the Dirac equation for the particles,

\[
\phi_a(\mathbf{p}_a, \mathbf{r}) = u_a e^{i\mathbf{p}_a \cdot \mathbf{r}},
\]

(2-9a)

and for the antiparticles,

\[
\phi_b(-\mathbf{p}_b, \mathbf{r}) = C \phi_a(\mathbf{p}_a, \mathbf{r}) = \gamma_2 \phi_a(\mathbf{p}_a, \mathbf{r}) = v_b e^{-i\mathbf{p}_b \cdot \mathbf{r}}.
\]

(2-9b)

Here, a and b denote particles and antiparticles, respectively, and C is the charge conjugation operator. We find

\[
M_{\text{p}^-} = G_p^2 \frac{1}{2(2\pi)^4} \delta(E_p + E_e + E_v - E_n) \left[ \overline{v}_n \gamma_\mu (1 + \gamma_5) \right] \left[ \overline{v}_p \gamma_\mu (1 + \gamma_5) \right] \left[ \overline{v}_e \gamma_\mu (1 + \gamma_5) \right] \left[ \overline{v}_v \gamma_\mu (1 + \gamma_5) \right]
\]

\[
\times \phi_n(\mathbf{p}_n, \mathbf{r}) \left[ \overline{v}_e(-\mathbf{p}_e, \mathbf{r}) \gamma_\mu (1 + \gamma_5) \phi_v(-\mathbf{p}_v, \mathbf{r}) \right].
\]

(2-10a)
Inside the nucleus, we replace the nucleon plane waves by bound spinor wave functions, and represent electrons or positrons by wave functions which are solutions of the Dirac equation for an extended charged nucleus surrounded by atomic electrons. Furthermore, it is convenient to split off the delta function and the factor $2\pi$ by writing

$$M = 2\pi \delta(E_i - E_f) \langle f | H_\beta | i \rangle.$$ 

The hadron parts of Eqs. (2-10a-c) can now be expanded into multipoles (Schopper, 1966; Konopinski, 1966; Bouchez and Depommier, 1960; Weidenmüller, 1961):

$$
\left[ \gamma \lambda_5 \gamma_\mu \right]_i^f = \sum_{KLM} (-1)^{K+L+M} a_{KL\alpha} (r) b_{L\alpha M} (r). 
$$

Here, $i$ and $f$ denote initial and final states, and

$$T_{LL0} = i L M 
$$

and

$$T_{KL1} = (-1)^{L-K+L} i L M \alpha 
$$

are the multipole operators.\textsuperscript{5}
The expansion coefficients $a_{KLS}^M(r)$ can be derived from the relation

$$a_{KLS}^M(r) = \int \phi_p^+(1+\lambda \gamma_5)T^M_{KLS} \phi_\nu^\dagger(r) d\Omega_{\text{nucl}}$$  \hspace{1cm} (2-13)

Inserting Eqs. (2-11) and (2-13) in Eqs. (2-10a-c), we find for the matrix elements

$$\langle f|H_p^-|1 \rangle = g_p \sum_{K\in N} (-1)^{K+M} \int \phi_p^+(1+\lambda \gamma_5)T^M_{KLS} \phi_\nu^\dagger(-q) d\Omega_{\text{lept}} [r^2 dr];$$  \hspace{1cm} (2-14a)

$$\langle f|H_p^+|1 \rangle = g_p \sum_{K\in N} (-1)^{K+M} \int \phi_p^+(1+\lambda \gamma_5)T^M_{KLS} \phi_\nu^\dagger(q) d\Omega_{\text{lept}} [r^2 dr];$$  \hspace{1cm} (2-14b)

$$\langle f|H_{\nu e}|1 \rangle = g_p \sum_{K\in N} (-1)^{K+M} \int \phi_p^+(1+\lambda \gamma_5)T^M_{KLS} \phi_\nu^\dagger(q) d\Omega_{\text{lept}} [r^2 dr].$$  \hspace{1cm} (2-14c)

Here, $q$ denotes the momentum for neutrino or antineutrino, and $\phi_{\nu e}(\vec{r})$ is the electron or positron wave function in the Coulomb field of a nucleus of atomic number $Z$.

We expand the electron (positron) and neutrino (antineutrino) continuum wave functions in partial spherical waves $\phi^\mu_e$, $\phi^\mu_{\bar{e}}$, $\phi^\mu_\nu$, $\phi^\mu_{\bar{\nu}}$ (Konopinski, 1966; Schülke, 1964; Weidenmüller, 1961):

$$\phi_{\nu e}(z) = \sum_{\kappa, \mu} a_{\kappa e}^{\mu} \phi^\mu_{\nu e}(z),$$  \hspace{1cm} (2-15a)
\[ \phi_{\nu e}(q) = \sum_{\kappa \lambda} b_{\kappa \lambda} c_{\nu \lambda} \phi_{-\nu}^\mu(q), \quad (2-15b) \]

\[ \phi_{e^+(-Z)} = -\gamma_e \phi_{e^-(Z)}^* = \sum_{\kappa \lambda} (-1)^{j} \kappa \lambda \phi_{-\nu}^\mu(-Z), \quad (2-15c) \]

\[ \phi_{-\nu e}(q) = -\gamma_e \phi_{e^+(Z)}^* = \sum_{\kappa \lambda} (-1)^{j} \kappa \lambda \phi_{-\nu}^\mu(-q). \quad (2-15d) \]

The spherical waves \( \phi_{k\nu}^\mu \) here have the form

\[ \phi_{k\nu}^\mu(z) = \left\{ \begin{array}{ll}
\left( \text{sign } \nu \right) f_k(z, r) \chi_{k\mu} \n \left[ 1, j = \ell - \frac{1}{2} \right] \n \left[ 2 + (i+1), j = \ell + \frac{1}{2} \right] \n \end{array} \right. \quad (2-16) \]

where we have

\[ \chi_{k\mu} = i \sum_{m} C_{1/2,j;m;\mu} \chi_{k,m}, \quad (2-17) \]

the \( \chi_{k,m} \) are two-component Pauli spinors, the \( C_{1/2,j;m;\mu} \) are Clebsch-Gordan coefficients, and the \( \chi_{k,m}^\mu \) are spherical harmonics. The index \( k \) is

\[ k = \begin{cases}
\ell, & j = \ell - \frac{1}{2} \\
-\ell - (i+1), & j = \ell + \frac{1}{2} 
\end{cases} \quad (2-18) \]

and \( g_k(Z, r) \) and \( f_k(Z, r) \) are the large and small radial wave functions, respectively.

The antiparticle (positron) wave function is (Rose, 1961)

\[ \phi_{k\nu}^{\mu}(-Z) = \left\{ \begin{array}{ll}
g_k(-Z, r) \chi_{k,\mu} \n \left[ 1, j = \ell - \frac{1}{2} \right] \n \left[ 2 + (i+1), j = \ell + \frac{1}{2} \right] \n \end{array} \right. \quad (2-19) \]
The neutrino radial wave functions can be written explicitly:

\[ \phi^\nu_{\kappa}(q) = \begin{bmatrix} j_{\frac{1}{2}}(qr)\chi_{-\kappa}^\mu \\ j_{\frac{3}{2}}(qr)\chi_{-\kappa}^\mu \end{bmatrix} \]  

(2-20)

where \( \kappa \) = \( -\kappa \) = \( \text{sign} \kappa \), and \( j_{\frac{1}{2}}(qr) \) is the spherical Bessel function.

For the antineutrino we have

\[ \phi^{\bar{\nu}}_{\kappa}(q) = \begin{bmatrix} j_{\frac{1}{2}}(qr)\chi_{-\kappa}^{\bar{\mu}} \\ j_{\frac{3}{2}}(qr)\chi_{-\kappa}^{\bar{\mu}} \end{bmatrix} \]  

(2-21)

The expansion coefficients \( a_{\kappa \mu} \) and \( b_{\kappa \mu} \) in Eqs. (2-15) are determined by the condition that the continuum wave functions \( \phi_e^{-}(q) \) and \( \phi_\nu^{-}(q) \) become asymptotically equal to a plane wave plus incoming (or outgoing) spherical waves (Schülke, 1964; Weidenmüller, 1961):

\[ a_{\kappa \mu}^{-}(q, s_e) = \frac{4\pi}{\sqrt{2}} \left[ \frac{1}{2} \chi_{\frac{1}{2}}(q, e) j_{\mu}^{-}\chi_{-\kappa}^{\nu} \right] \]

\[ -i\Lambda_{\kappa} (\pi/2)(\kappa + 1) \]

\[ x_y \]  

(2-22)

\[ b_{\kappa \mu}^{-}(q, s_\nu) = \frac{4\pi}{\sqrt{2}} \left[ \frac{1}{2} \chi_{\frac{1}{2}}(q, \nu) j_{\mu}^{-}\chi_{-\kappa}^{\nu} \right] \]

(2-23)

Here, \( \Lambda_{\kappa} \) is the Coulomb phase (Bühring, 1965, 1967).

It is useful, furthermore, to introduce reduced hadron and lepton matrix elements by applying the Wigner-Eckart theorem. From Eqs. (2-14a-c) and (2-15a-c) we find
\begin{align}
\langle f \mid \mathcal{H}_- \mid i \rangle &= g_{\pi} 2^{-1/2} \sum_{KLSM} \sum_{\kappa_{\mu e}} (-1) \beta_{f-M_f+\mu} \kappa_{\nu \mu} \\
&\times (-1)^{K+M+\mu} \frac{\beta_{j_f} \beta_{j_1}}{\beta_{j_\nu} \beta_{j_e}} \frac{\beta_{-\mu_e} \beta_{-\mu_\nu}}{\beta_{-\mu_\nu} \beta_{-\mu_\nu}} \\
&\times \epsilon^{a_{\kappa_{\mu e}} b_{\kappa_{\nu \mu}}^*} \left\langle \phi_p \left\| \left(1 + \gamma_5 \right) T_{KLS} \right\| \phi_n \rightangle \\
&\times \epsilon^{(\phi_{\kappa_{\mu e}} \left( -2 \right) \frac{1}{2}) \left( -\gamma_5 \right) T_{KLS} \phi_{\kappa_{\nu \mu}} \left( -2 \right) \frac{1}{2}} r^2 dr; \quad (2-24a)
\end{align}

\begin{align}
\langle f \mid \mathcal{H}_+ \mid i \rangle &= g_{\pi} 2^{-1/2} \sum_{KLSM} \sum_{\kappa_{\mu e}} (-1) \beta_{f-M_f+\mu} \kappa_{\nu \mu} \\
&\times (-1)^{K+M+\mu} \frac{\beta_{j_f} \beta_{j_1}}{\beta_{j_\nu} \beta_{j_e}} \frac{\beta_{-\mu_e} \beta_{-\mu_\nu}}{\beta_{-\mu_\nu} \beta_{-\mu_\nu}} \\
&\times \epsilon^{a_{\kappa_{\mu e}} b_{\kappa_{\nu \mu}}^*} \left\langle \phi_p \left\| \left(1 + \gamma_5 \right) T_{KLS} \right\| \phi_n \rightangle \\
&\times \epsilon^{(\phi_{\kappa_{\mu e}} \left( -2 \right) \frac{1}{2}) \left( -\gamma_5 \right) T_{KLS} \phi_{\kappa_{\nu \mu}} \left( -2 \right) \frac{1}{2}} r^2 dr; \quad (2-24b)
\end{align}

\begin{align}
\langle f \mid \mathcal{H}_{EC} \mid i \rangle &= g_{\pi} 2^{-1/2} \sum_{KLSM} \sum_{\kappa_{\nu \mu}} (-1) \beta_{f-M_f+\mu} \kappa_{\nu \mu} \\
&\times (-1)^{K+M} \epsilon^{j_f \beta_{j_1}} \epsilon^{j_\nu \beta_{j_e}} \epsilon^{\beta_{-\mu_e} \beta_{-\mu_\nu}} \\
&\times \epsilon^{a_{\kappa_{\nu \mu}} b_{\kappa_{\nu \mu}}^*} \left\langle \phi_p \left\| \left(1 + \gamma_5 \right) T_{KLS} \right\| \phi_n \rightangle \\
&\times \epsilon^{(\phi_{\kappa_{\nu \mu}} \left( -2 \right) \frac{1}{2}) \left( -\gamma_5 \right) T_{KLS} \phi_{\kappa_{\nu \mu}} \left( -2 \right) \frac{1}{2}} r^2 dr. \quad (2-24c)
\end{align}
Here, $x (= K, L_1, L_2, L_3, M_1, \ldots)$ denotes the different shells and subshells of the atomic cloud from which the electron can be captured. The states of the initial neutron or proton are specified by $|J_M\rangle$, and those of the final nucleon, by $|J'_M\rangle$.

The similarity of Eqs. (2-24a), (2-24b), and (2-24c) suggests that we need to derive the final formulae of the observables for only one type of decay ($\beta^-, \beta^+$, or EC) and can hence obtain results for the other decay modes. For this purpose, we transform Eqs. (2-24b,c) into a form that is similar to that of Eq. (2-24a), by interchanging initial and final states in the reduced lepton matrix elements. Taking into account the relation (Weidenmüller, 1961)

$$<f \left| \left( 1 + \gamma_5 \right) T_{KLS} \right| i>^* = (-1)^{K-S-J_J} \left| \left( 1 + \gamma_5 \right) T_{KLS} \right| f >$$

and the fact that here the reduced matrix elements are defined as real quantities, we obtain

$$<f \left| H_{\beta^+} \right| i> = G_F \frac{1}{2} \sum KLSM \sum (-1)^{J'_f - J_i} \kappa_e \mu_e$$

$$\times (-1)^{K-S+M+J_u-\mu_v+1} \left\{ \begin{array}{ccc} J_f & K & J_i \\ -M_f & M & J_i \end{array} \right\}^{\text{e}} \left\{ \begin{array}{ccc} j_e & K & j_u \\ -\mu_e & -M & -\mu_u \end{array} \right\}^{\text{e}}$$

$$\times a_{K}^{\beta^+} \kappa_e \mu_e \kappa_v \mu_v \int \langle \phi_n \left| \left( 1 + \lambda \gamma_5 \right) T_{KLS} \right| \phi_f \rangle^*$$

$$\times \left( \phi_e \right) \left( -Z \right) \left\{ \left( 1 + \gamma_5 \right) T_{KLS} \right\} \phi_v (q) r^2 dr.$$
The transition probability per unit time can now be found from standard quantum-mechanical formulae. By applying first-order time-dependent perturbation theory (the "Golden Rule"), the decay constant $\lambda$ and the half life $t$ are given by

$$\lambda = (2n^2)(t^-)^{-1} = 2\pi(2J_i+1)^{-1}\sum_{M_i} \sum_{s_e, s_v}$$

$$\times \iint \int |\langle f | H_{EC} | i \rangle|^2 \frac{p^2}{2E} dp d\Omega_e d\Omega_v (2\pi)^{-6}$$

(2-27)

for $\beta^\pm$-decay, and by

$$\lambda = (2n^2)(t^-)^{-1} = 2\pi(2J_i+1)^{-1}\sum_{M_i} \sum_{s_e, s_v}$$

$$\times \int |\langle f | H_{EC} | i \rangle|^2 q_x^2 d\Omega_v (2\pi)^{-3}$$

(2-28)

for electron capture from the atomic $x$-shell. By inserting the matrix element given by Eq. (2-24a) in Eq. (2-27) and making use of the orthogonality relations for Clebsch-Gordan coefficients and $3j$-symbols, we find

$$\langle f | H_{EC} | i \rangle = g_{\beta^\pm}^{-1/2} \sum_{KLM} \sum_{\mu_x \nu_x} (-1)^{K-s+M_f} J_f K J_i J_x K M_f M_i M_x M_y$$

$$\times (-1)^{J_x+H_x} \iint \int \phi_n \langle (1+\gamma_5)\Gamma_{KLM} \phi_p \rangle$$

$$\times \phi_{x\mu_x}^*(z) \langle (1+\gamma_5)\Gamma_{KLM} \phi_{x\nu_x}^*(q) r^2 dr. $$

(2-26b)
\[ \lambda_{p^-} = g_p^2(2\pi)^{-3}\left(2J_{1/2}+1\right)^{-1}\int_0^\infty \sum_K \sum_{\kappa \in K} \sum_{l_s} \left\langle \sum_{2K+1}^{1/2}(4\pi)^{1/2} \right. \]
\[ \times \left( \phi_p \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_n \right) \]
\[ \times (4\pi)^{1/2} \left( \phi_{\kappa_{e}}(Z) \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_{\kappa_{\nu}}(-q) \right) r^2 dr \right\rangle q^2 dp. \]  
(2-29)

Similarly, by combining Eqs. (2-26a) and (2-27), we have
\[ \lambda_{p^+} = g_p^2(2\pi)^{-3}\left(2J_{1/2}+1\right)^{-1}\int_0^\infty \sum_K \sum_{\kappa \in K} \sum_{l_s} (-1)^{l} \left(2K+1\right)^{-1/2}(4\pi)^{1/2} \]
\[ \times \left( \phi_n \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_p \right) \]
\[ \times (4\pi)^{1/2} \left( \phi_{\kappa_{e}}(Z) \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_{\kappa_{\nu}}(q) \right) r^2 dr \right\rangle q^2 dp. \]  
(2-30)

The electron-capture decay constant is found by inserting Eq. (2-25b) in Eq. (2-28):
\[ \lambda_{e} = g_p^2(2\pi)^{-3}(2J_{1/2}+1)^{-1}\int_0^\infty \sum_K \sum_{\kappa \in K} \sum_{l_s} (-1)^{l} \left(2K+1\right)^{-1/2} \]
\[ \times (4\pi)^{1/2} \left( \phi_n \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_p \right) \]
\[ \times (4\pi)^{1/2} \left( \phi_{\kappa_{e}}(Z) \left| (1+\gamma_Y)_{T_{KLS}} \right| \phi_{\kappa_{\nu}}(q) \right) r^2 dr \right\rangle q^2. \]  
(2-31)

The neutrino momentum \( q_x \) is given by
\[ q_x = W_o + W'_x, \]  
(2-32)

where \( W_o \) is the total transition energy between initial and final states (the difference between the atomic masses, minus \( m_0 c^2 \), see Sec. 1.2), and \( W'_x \) denotes the energy of the bound electron (in the daughter atom). This is \( W'_x = \left| E'_x \right| \), where \( E'_x \) is the binding energy of the electron. Because the electron and neutrino wave functions
are well-known [Eqs. (2-16)-(2-21)], we can evaluate the reduced lepton matrix elements explicitly (Sch"ulke, 1964; Weidenm"uller, 1961). For the three kinds of reduced lepton matrix elements appearing in Eqs. (2-29), (2-30), and (2-31), we have

\[
\left(4\pi\right)^{1/2}\frac{\langle \bar{\psi}_{e} \psi_{\nu} \rangle (Z)}{\langle 1+\gamma_{5} \rangle_{KLs}} \langle \bar{\psi}_{e} \mid q_{KLs} \mid \psi_{\nu} \rangle = g_{e}^{\nu}_{\nu} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, \kappa_{\nu}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\nu}) \right] + (\text{sign} \ k_{e}) f_{\bar{\nu}_{s}} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\nu}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, \kappa_{\nu}) \right] \quad (2-33)
\]

for the electron-neutrino matrix element,

\[
\left(4\pi\right)^{1/2}\frac{\langle \bar{\psi}_{e} \psi_{\bar{\nu}} \rangle (Z)}{\langle 1+\gamma_{5} \rangle_{KLs}} \langle \bar{\psi}_{e} \mid q_{KLs} \mid \psi_{\bar{\nu}} \rangle = g_{e}^{\nu}_{\nu} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, \kappa_{\bar{\nu}}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\bar{\nu}}) \right] + (\text{sign} \ k_{e}) f_{\bar{\nu}_{s}} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\bar{\nu}}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, \kappa_{\bar{\nu}}) \right] \quad (2-34)
\]

for the electron-antineutrino matrix element, and

\[
\left(4\pi\right)^{1/2}\frac{\langle \bar{\psi}_{e} \psi_{\nu} \rangle (-Z)}{\langle 1+\gamma_{5} \rangle_{KLs}} \langle \bar{\psi}_{e} \mid q_{KLs} \mid \psi_{\nu} \rangle = g_{e}^{\nu}_{\nu} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, \kappa_{\nu}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\nu}) \right] + (\text{sign} \ k_{e}) f_{\bar{\nu}_{s}} \left[ j_{\bar{\nu}_{s}} g_{KLs}(\kappa_{e}, -\kappa_{\nu}) + j_{\nu_{s}} g_{KLs}(\kappa_{e}, \kappa_{\nu}) \right] \quad (2-35)
\]

for the positron-neutrino matrix element. The quantity $g_{KLs}(\mathbf{n}_{1}, \mathbf{n}_{2})$, introduced by Weidenmüller (1961), represents the spin-singular part of these reduced lepton matrix elements.
\[ G_{\text{KLS}}(n_f, n_i) = \left\{ \frac{(2s+1)(2K+1)[2\ell_f(n_f)+1][2\ell_i(n_i)+1](2j_f+1)(2j_i+1)}{4\pi^2} \right\}^{1/2} \]
\[
\times \frac{G(n_f)+L(n_f)+L}{(-1)^{j_f-j_i}}
\times C\left(\ell(n_f), \ell(n_i), L; 00\right) \left(\begin{array}{ccc}
K & S & L \\
\ell_f & \frac{1}{2} & \ell(n_f) \\
\ell_i & \frac{1}{2} & \ell(n_i) \end{array}\right).
\]

Here, we have \( \ell(n) = n \) if \( n > 0 \) and \( \ell(n) = |n| - 1 \) if \( n < 0 \), where \( n \) stands for \( +K \) and \( -K \); \( C \) is a Clebsch-Gordan coefficient, and the braces denote a Wigner 9j-symbol.

We now consider the relation between \( \beta^- \) and \( \beta^+ \) decay. It is easily shown that the following relations hold:
\[
\begin{align*}
(-1)^{-s}(4\pi)^{1/2} & \langle \phi_{\kappa_e}(z) | \phi_{\kappa_v}(-q) \rangle \\
& = \delta_{\kappa_e}(z) \left[ j_{\kappa_v} G_{\text{KLS}}(\kappa_e, \kappa_v) + j_{\kappa_v} G_{\text{KLS}}(\kappa_v, \kappa_e) \right] \\
& \quad \times \text{(sign } \kappa_e) \kappa_e^{(\kappa_v) \left[ j_{\kappa_v} G_{\text{KLS}}(-\kappa_e, \kappa_v) + j_{\kappa_v} G_{\text{KLS}}(-\kappa_v, \kappa_e) \right].
\end{align*}
\]

Thus, we see from Eq. (2-35) that the product of the two reduced matrix elements in Eq. (2-30) can be replaced as follows:
\[
\begin{align*}
\left\langle \left\langle (1+\gamma_5)T_{\text{KLS}} \right\rangle \left\langle (1-\gamma_5)T_{\text{KLS}} \right\rangle \right\rangle \\
= \left\langle \left\langle T_{\text{KLS}} \right\rangle \left\langle (1-\gamma_5)T_{\text{KLS}} \right\rangle \right\rangle - \left\langle \left\langle (1-\gamma_5)T_{\text{KLS}} \right\rangle \left\langle (1+\gamma_5)T_{\text{KLS}} \right\rangle \right\rangle.
\end{align*}
\]

Consequently, we can derive the formulae for \( \beta^+ \) decay from those for \( \beta^- \) decay by making the following substitutions:
\[
\begin{align*}
\beta^- \text{ decay} & \quad \beta^+ \text{ decay} \\
\lambda & \rightarrow -\lambda \\
\not{Z} & \rightarrow -\not{Z} \\
\mathcal{G} & \rightarrow -\mathcal{G}
\end{align*}
\]

(2.39)

Here, \( \mathcal{G} \) represents the terms which are due to parity non-conservation (e.g., electron polarization or \( \beta^-\gamma \) circular polarization correlation). The relation between \( \beta^+ \) decay and electron capture is established by the substitutions [cf. Eqs. (2-30), (2-31), (2-33), (2-35)]

\[
\begin{align*}
\beta^+ \text{ decay} & \quad \text{electron capture} \\
\tilde{f}^{*}_{\kappa_e}(-z) & \rightarrow \tilde{f}_{\kappa^*_e}(z) \\
\tilde{g}^{*}_{\kappa^*_e}(-z) & \rightarrow \tilde{g}_{\kappa_{e}}(z)
\end{align*}
\]

(2.40)

where \( \tilde{g}_{\kappa_{e}}(z) \) and \( \tilde{f}^{*}_{\kappa^*_e}(z) \) are the large and small components of the bound-state electron radial wave functions, respectively.

Alternatively, we can start from \( \beta^- \) decay [cf. Eqs. (2-29) and (2-34) vs. Eqs. (2-31) and (2-33)]:

\[
\begin{align*}
\beta^- \text{ decay} & \quad \text{électron capture} \\
\tilde{j}^{\pm}_{\kappa^*_\nu}(q'r) & \rightarrow \tilde{j}^{-\pm}_{\kappa_{\nu}}(q'r) \\
\tilde{j}^{-\pm}_{\kappa^*_\nu}(q'r) & \rightarrow \tilde{j}^{\pm}_{\kappa_{\nu}}(q'r)
\end{align*}
\]

\[
\langle f \left| \left(1+\lambda \gamma_5 \right)_{\kappa_{ls}} \right| d \rangle \rightarrow (-1)^{B}\langle f \left| \left(1+\lambda \gamma_3 \right)_{\kappa_{ls}} \right| d \rangle \quad (\text{2-41a})
\]

Continuum-electron \quad \text{bound-electron}

wave function \quad \text{wave function}
For the decay probabilities as given in Eqs. (2-29) and (2-31), this
prescription can be replaced by one mentioned by Behrens and
Jünecke (1969):

$$\beta^- \text{ decay} \quad \rightarrow \quad \text{electron capture}$$

$$q \quad \rightarrow \quad -q$$

$$\langle f \parallel (1+\gamma_5)^{\pi} T_{KL} \parallel i \rangle \rightarrow (-1)^{K-S} \langle f \parallel (1+\gamma_5)^{\pi} T_{KL} \parallel i \rangle.$$ (2-41b)

In this description of the electron-capture and β-decay
processes, three important points have not been considered:

1. The hadron current in the form of Eq. (2-4) is an approx­
imation which is only valid for bare nucleons. The exact form
of this current is discussed under the heading Induced Interac­
tions in Sec. 2.2.3.

2. The Hamiltonian and transition matrix elements used
here refer to a single-particle process. The description must
be generalized for the case of many nucleons in the initial and
final states. This point is discussed in Sec. 2.2.3.

3. A complete description of the initial and final states
must include the electrons of the atomic cloud. Since the nu­
clear charge and the number of electrons are different in the
initial and final states, the atomic-electron wave functions of
these two states are not orthogonal, and the overlap between
them is not perfect. This leads to some modifications of the
transition rate (exchange and overlap corrections) and to higher-
order processes (e.g., autoionization). These points are discussed in Secs. 2.5 and 5.

2.2. Electron-Capture Transition Rates

2.2.1. General Relations for the Transition Probabilities

In discussing transition matrix elements and transition rates for the three weak nuclear decay modes, we have pointed out how these decay types are related. From here on, we consider electron capture only. We simplify Eq. (2-31), discuss the electron and neutrino radial wave functions in the lepton part, and generalize the hadron part through methods of elementary-particle physics.

We first note that Eq. (2-33) is invariant under the substitution and set . We also introduce the abbreviation (Behrens and Büning, 1971)

\[
\beta_x \sum \frac{1}{I_d} \frac{1}{2} \left[ (2I_d + 1)(2\ell + 1) \right]^{-1/2} \phi_{\ell}(k) = \sum \frac{1}{k^2} \left[ \phi_{\ell}(k) \phi_{\ell}(k) \right] \frac{1}{I_d} \frac{1}{2} \left[ (2I_d + 1)(2\ell + 1) \right]^{-1/2} \phi_{\ell}(k)
\]

\[
\times \left[ \int \phi_{\ell}(k) \phi_{\ell}(k) \right] \left[ \phi_{\ell}(k) \phi_{\ell}(k) \right] \frac{1}{I_d} \frac{1}{2} \left[ (2I_d + 1)(2\ell + 1) \right]^{-1/2} \phi_{\ell}(k)
\]

\[
+ \int \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \frac{1}{I_d} \frac{1}{2} \left[ (2I_d + 1)(2\ell + 1) \right]^{-1/2} \phi_{\ell}(k)
\]

\[
\times \left[ \int \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \phi_{\ell}(k) \right] \frac{1}{I_d} \frac{1}{2} \left[ (2I_d + 1)(2\ell + 1) \right]^{-1/2} \phi_{\ell}(k)
\]

where and \( \beta_x \) is the Coulomb amplitude (Behrens and Negele, 1969) of the bound-state electron radial wave function.
(ERWF), discussed in Sec. 2.2.2. For \( K_x = -1 \) we have \( \beta_x = g_{-1}(0) \) equal to the value of the wave function \( g_{-1}(r) \) at \( r = 0 \).

For the total capture probability from all atomic shells we then have

\[
\lambda_c = (\ln 2) (t_c)_{-1} = \alpha^2 (2\pi^3)^{\frac{1}{2}} \sum_x x C_x f_x
\]

\[ (2-43) \]

(Behrens and Jönnecke, 1969; Bouchez and Depommier, 1960; Brysk and Rose, 1958). The sum in Eq. (2-43) extends over all atomic subshells from which an electron can be captured. For closed shells, \( n_x \) equals 1. For partially filled shells, \( n_x \) is equal to the relative occupation number of electrons in the shell.

The quantity \( C_x \) corresponds to the shape factor of \( \beta \) decay.

Taking into account only the lowest-order terms in the summation over \( \kappa \) and \( \kappa_v \), \( C_x \) has the form

\[
C_x = \left[ M_L(k_x, k_v(1)) + (k_x/k_v) M_L(k_x, k_v(1)) \right]^2
\]

\[ + \left[ M_L(k_x, k_v(2)) + (k_x/k_v) M_L(k_x, k_v(2)) \right]^2 \]

\[ + \left[ M_{L+1}(k_x, k_v(2)) + (k_x/k_v) M_{L+1}(k_x, k_v(2)) \right]^2 \]

\[ + \delta_{\Delta J, 0} \left[ M_o(1,1) + (k_x/k_v) M_o(1,1) \right]^2 \]

\[ (2-44) \]

The classification of allowed and forbidden electron-capture transitions is similar to that in \( \beta \) decay (Schopper, 1966; Konopinski, 1966; Behrens and Jönnecke, 1969):
\[ \Delta J = 0, 1 \quad \pi_i \pi_f = +1 \quad \text{allowed} \]
\[ \Delta J = 0, 1 \quad \pi_i \pi_f = -1 \quad \text{first non-unique forbidden} \]
\[ \Delta J = n > 1 \quad \pi_i \pi_f = (-1)^n \quad n^{\text{th}} \text{ non-unique forbidden} \]
\[ \Delta J = n > 1 \quad \pi_i \pi_f = (-1)^{n-1} \quad (n-1)^{\text{th}} \text{ unique forbidden} \]

(2-45)

Here, \((J_i, \pi_i)\) and \((J_f, \pi_f)\) denote spins and parities of the initial and final nuclear states, and we have \(\Delta J = |J_i - J_f|\). Hence, we can write in Eq. (2-44)

\[
L = \Delta J \quad \text{for } \Delta J > 0 \\
L = 1 \quad \text{for } \Delta J = 0 \\
\begin{align*}
\kappa_y^{(1)} &= L - k_x + 1 \\
\kappa_y^{(2)} &= L - k_x + 2.
\end{align*}
\text{(2-46)}

The quantities \(\kappa_x\) and \(\kappa_x\) are related by Eq. (2-18) to the total angular momentum \(j_x\) and the orbital angular momentum \(l_x\) of the bound electron. Similarly, \(\kappa_y\) and \(\kappa_y\) determine \(j_y\) and \(l_y\) of the continuum wave function of the emitted neutrino.

The values of \(\kappa_x\) for bound electrons are as follows:

<table>
<thead>
<tr>
<th>(K)</th>
<th>(\kappa_x)</th>
<th>(M)</th>
<th>(\kappa_x)</th>
</tr>
</thead>
<tbody>
<tr>
<td>((1s))</td>
<td>(-1)</td>
<td>(1)</td>
<td>(-1)</td>
</tr>
<tr>
<td>(L_1)</td>
<td>(-1)</td>
<td>(1)</td>
<td>(+1)</td>
</tr>
<tr>
<td>(L_2)</td>
<td>(+1)</td>
<td>(3)</td>
<td>(-2)</td>
</tr>
<tr>
<td>(L_3)</td>
<td>(-2)</td>
<td>(5)</td>
<td>(+3)</td>
</tr>
</tbody>
</table>

(2-47)
The function \( f_x \) in Eq. (2-43), which corresponds to the integrated Fermi function of \( \beta \) decay, has the form

\[
f_x = (\pi/2)q_x^2 p_x^2 B_x.
\]  

(2-48)

The factor \( B_x \) takes account of the effects of electron exchange and overlap; it is discussed in Sec. 2.5.

2.2.2. Bound-State Electron Radial Wave Functions

The electron radial wave functions \( f_x(r) \) and \( g_x(r) \) are a solution of the Dirac radial differential equation or of the equivalent integral equation (Rose, 1961; Behrens and Bühring, 1971). It is convenient to consider instead the functions \( H_k(r) \), \( h_k(r) \), \( D_k(r) \), and \( d_k(r) \) introduced by Bühring (1963a):

\[
f_{k_x} (r) = \beta_x (p_x r)^{k_x-1} [(2k_x+1)!!]^{-1/2} [H_{k_x} (r) + h_{k_x} (r)] \text{ for } k_x > 0
\]  

(2-49a)

\[
g_{k_x} (r) = \beta_x (p_x r)^{k_x-1} [(2k_x+1)!!]^{-1/2} [D_{k_x} (r) + d_{k_x} (r)] \text{ for } k_x < 0
\]  

(2-49b)

\[
f_{-k_x} (r) = -\beta_x (p_x r)^{k_x-1} [(2k_x+1)!!]^{-1/2} [D_{k_x} (r) - d_{k_x} (r)] \text{ for } k_x < 0
\]  

(2-49c)

\[
g_{-k_x} (r) = \beta_x (p_x r)^{k_x-1} [(2k_x+1)!!]^{-1/2} [H_{k_x} (r) - h_{k_x} (r)] \text{ for } k_x < 0
\]  

(2-49d)

Here, \( R \) is the nuclear radius, or equivalent radius of a uniformly charged nucleus.

The first of two aspects of the electron radial wave functions that require more detailed consideration is the behavior
of these functions inside the nucleus: the dependence of the electron and neutrino radial wave functions on the distance \( r \) from the center of the nucleus must be subsumed into the nuclear matrix elements [cf. Eq. (2-42)]. The \( r \)-dependence of the electron radial wave functions inside the nucleus depends essentially on the form of the nuclear charge distribution.

Secondly, the Coulomb amplitudes \( \beta_x \) must be considered; they can only be calculated numerically by solving the Dirac equation for an extended nucleus and for a self-consistent atomic potential. The value of \( \beta_x \) is essentially determined by the shape of the charge distribution of the surrounding atomic electrons.

In many of the earlier papers on \( \beta \) decay and electron capture, the expansion of the functions \( H_k(r) \), \( D_k(r) \), \( h_k(r) \), and \( d_k(r) \) in powers of \( r \) is carried out (Behrens and Bühring, 1970):

\[
H_k(r) = \sum_{n=0}^{\infty} H_n(k)(r/R)^n;
\]

(2-50)

the different powers of \( r \) are then incorporated into the definition of the nuclear matrix elements (Behrens and Jüttner, 1969; Bühring, 1963a,b). The nuclear charge distribution has been approximated throughout by a uniformly charged sphere of radius \( R \), equal to the nuclear radius. Because this charge distribution is discontinuous at \( R \), the power-series expansion of the electron radial wave functions is only valid inside the nucleus. Usable \( \beta \)-decay and electron-capture formulae have been derived by truncating this series and extrapolating the resulting polynomials.
outside the nuclear radius (Behrens and Jänecke, 1969; Bühring, 1963a). However, this approach is unsatisfactory: contrary to general belief, a significant contribution (or even the main contribution) to the nuclear matrix elements originates from the region \( r > R \), particularly if the initial and final nucleons are in different shells (Behrens and Bühring, 1971; de Raedt, 1968).

It might be expected that this difficulty could be avoided by using a more realistic, smooth nuclear charge distribution, such as a modified Gaussian or a Fermi distribution. However, for such distributions the \( \beta \)-decay and electron-capture formulae do not converge at all (Behrens and Bühring, 1970, 1971), because the nuclear matrix elements are introduced by integrating a power series term by term, a dubious procedure if the upper limit of the integral is infinity. Only if the potential \( V(r) \) vanishes identically, as for the neutrino, is this procedure justified.

Thus, the neutrino radial wave functions (spherical Bessel functions) can be expanded in powers of \( r \). The electron radial wave functions, on the other hand, can be expanded in powers of the mass and energy parameters of the electron and the nuclear charge. The coefficients in this expansion still are functions of \( r \) and depend on the shape of the charge distribution. We find (Behrens and Bühring, 1971)

\[
H_{k_x}^\pm (r) = \sum_{\mu=0}^{\infty} \sum_{\nu=0}^{\infty} \sum_{p=0}^{2\nu} \frac{(2k_x-1)!!}{(2p)!!(2\nu+2k_x-1)!!(-1)^\nu(\nu)!^2\nu!^2\nu!(\rho/r)^{2\nu}}
\]

\[
\times \mathcal{I}_k(2\mu, 2\nu, \rho; r)(w_e^2 R_x)^{2\mu-2\nu}(w_e R_x)^{2\nu-\rho}(aZ)^\rho,
\]

(2-51a)
The symbol $m_e$ has been retained in these equations, even though we use natural units $\hbar=c=m_e=1$, because $m_e$ will be used as an expansion parameter. The expansion coefficients $I(k_x,m,n,\rho;r)$ depend on the form of the nuclear charge distribution and on the parameters $k_x$, $m$, $n$, and $\rho$. The order $m$ is the sum of the exponents of $(m_eR)$, $(WR)$, and $aZ$; the number $n$ is the sum of the exponents of $(WR)$ and $(aZ)$. The functions $I$ with $\rho=0$ are trivial:

$$I(k_x,m,n,0;r) \equiv 1.$$  \hfill (2-52)

The functions $I$ with $\rho>0$, up to order $m=3$, are listed in Appendix A2.2 (Behrens and Bühring, 1971). For $aZ=0$, Eqs. (2-51) reduce to the usual expansion in powers of $r$ (Bühring, 1963a).
As usual, \( \alpha \) is the fine-structure constant, and \( Z \) is the atomic number of the parent nucleus.

The important function \( I(k_x, l, l, l; r) \), which gives the large Coulomb terms in non-unique forbidden transitions, takes the following forms for the three most widely used charge distributions (Behrens and Bühning, 1970, 1971):

(i) For a uniform charge distribution

\[
\rho(r) = \begin{cases} 
-3\alpha z/R, & 0 < r < R \\
0, & R < r < \infty 
\end{cases}
\]  

we have (Behrens and Bühning, 1971)

\[
I(k_x, l, l, l; r) = \begin{cases} 
\frac{3}{2} - \frac{2k + 1}{2(2k + 3)} (\frac{r}{R})^2, & 0 < r < R \\
\frac{2k + 1}{2k} \frac{R}{r} - \frac{3}{2k (2k + 3)} (\frac{r}{R})^{2k + 1}, & R < r \leq \infty 
\end{cases}
\]  

(ii) In a shell-model or modified Gaussian distribution (Behrens and Bühning, 1970)

\[
\rho(r) = N \left[ 1 + \alpha(k_x)^2 \right] e^{-\frac{(r/r_0)^2}{2}},
\]  

\[ (2-56a) \]
where

\[ N = \frac{8\alpha^2}{(2+3A)a^2 \sqrt{\pi}} \quad (2-56b) \]

the equivalent uniform radius \( R \) is related to the parameters \( a \) and \( A \) as

\[ R = a\left[5(2+5A)/2(2+3A)\right]^{1/2}; \quad (2-57) \]

for this distribution, we have

\[ I(k_x,1,1;\alpha) = \frac{2k_x+1}{2k_x} \frac{R_x}{r} \left[ \text{erf}(y) - (1+\frac{2k_x}{2+2A}) \frac{(2k_x-1)!!}{k_x 2k_x} \right] \]

\[ \times \left[ \text{erf}(y) + \frac{\beta R}{\sqrt{\pi}} \sum_{m=0}^{\infty} \frac{2^{m-1} \Gamma(2m+1)}{(2m+1)!} \right] \quad (2-58) \]

where \( \text{erf}(y) \) is the error function,

\[ \text{erf}(y) = \frac{2}{\sqrt{\pi}} \int_0^y e^{-t^2} dt, \quad y = \frac{\beta R}{\pi} \]

and

\[ \beta = \left[5(2+5A)/2(2+3A)\right]^{1/2}. \]

(iii) For a Fermi distribution (Behrens and Bühring, 1970)

\[ \rho(r) = -3\alpha^2e^{-3N[1+e^{(r-c)/b}-1]}, \quad (2-59) \]

with

\[ N = (1+2\alpha c^2)^2 - 6w_2(b,c;0) \]

the equivalent uniform radius \( R \) is

\[ R = c \left[ \frac{3c^5 + 10\pi^2 c^2 + 27 \pi^4 c^4 b - 360 \beta \left( \frac{w_5(b,c;0)}{w_2(b,c;0)} \right) \right]^{1/2} \quad (2-60) \]
The function $I(k_x, l, l, l; r)$ takes the form

$$
I(k_x, l, l, l; r) = \frac{2}{2} \frac{B}{c} N \left\{ -\frac{1}{2} \frac{1}{r} \frac{(b, c; r)}{2^{k_x + 1}} \right\}
$$

$$
- 2(2k_x + 1) \frac{S}{r} \left[ w_1(b, c; r) - \frac{1}{k_x} w_2(b, c; r) \right]
$$

$$
+ 2(2k_x + 1)(2k_x + 2) \frac{S}{r} \sum_{m=0}^{2k_x - 1} \left( -1 \right)^m
$$

$$
\times \frac{(2k_x - 1)!}{(2k_x - 1 - m)!} \left( \frac{S}{r} \right)^m w_4(b, c; r),
$$

(2-61)

where the functions $w_n$ are defined as

$$
w_n(b, c; r)
$$

$$
= \left\{ \begin{array}{lr}
\left( \frac{b}{c} \right) \sum_{m=1}^{\infty} (-1)^m n^{-n} \left( e^{-b} \right)^m, & r \leq c \\
\left( \frac{b}{c} \right) \sum_{m=0}^{\infty} a_m \left( \frac{r-c}{b} \right)^{m-2} \left( -1 \right)^m \sum_{m=1}^{\infty} (-1)^m n^{-n} \left( e^{-b} \right)^m, & r \geq c
\end{array} \right.
$$

(2-62)

(Schucan, 1965). Here, $a_m(n)$ stands for

$$
a_m(n) = (-1)^m \frac{(2m-2)!}{(2m)!} \frac{(2m-2)}{(2m)!} B_m,
$$

where the $B_{2m}$ are Bernoulli numbers:

$$
B_0 = 1, B_2 = \frac{1}{6}, B_4 = -\frac{1}{30}, B_6 = \frac{1}{42}, \ldots
$$

At $r = c$, $w_n$ is given by

$$
w_n(b, c; c) = \left( \frac{b}{c} \right) \sum_{m=1}^{\infty} (-1)^m n^{-n} = \left( \frac{b}{c} \right) n (2^{1-n} - 1) \zeta(n),
$$

(2-63)

where $\zeta(n)$ is the Riemann zeta function.
The functions $I(k_x, l, l, l)$ [Eqs. (2-55), (2-58), and (2-61)] have been derived neglecting the small influence of the atomic electron cloud on the $r$-dependence of the electron radial wave functions inside the nucleus. These functions are illustrated in Fig. 2-1.

We consider next the Coulomb amplitudes $\beta_x$ of the bound atomic electrons. These quantities can be calculated by integrating the Dirac equation in the potential of the nuclear and atomic charge distributions.\(^6\) The value of $\beta_x$ is essentially determined by the potential outside the nucleus, i.e., by the electronic screening of the nuclear electrostatic field. The finite nuclear size and the shape of the nuclear charge distribution have less influence on $\beta_x$. The potential produced by the nuclear charge and the atomic electron cloud can be derived approximately from statistical models (Gombas, 1956, 1967), by solving the Thomas-Fermi or the Thomas-Fermi-Dirac equations. A more exact form of the potential can be derived through self-consistent Hartree-Fock methods (Hartree, 1957; Slater, 1960; Myers, 1972; Burke and Grant, 1967; Grant, 1970; Lindgren and Rosen, 1974).

Both methods of finding the extranuclear potential can only be carried out numerically and have been pursued by many investigators. The Thomas-Fermi and Thomas-Fermi-Dirac equations have been solved for potentials and eigenvalues, for example, by Gombas (1956), Thomas (1954), Latter (1955), Shalitin (1965, 1967), and Yonsi (1966, 1967). The self-consistent field methods offer the best possibility for obtaining good atomic electron wave
functions, but require extensive numerical calculations (Hartree, 1957; Slater, 1960; Mayers, 1972; Burke and Grant, 1967; Grant, Lindgren and Rosen, 1974).

In a simplification first introduced by Slater (1960, Vol. 2), the exact exchange potential is approximated by the exchange potential of an electron gas with local electron density ρ(r), i.e.,

\[ V_{\text{ex}}(r) = \alpha \frac{3g(r)}{2} \frac{1}{r^{1/3}} \]  

(2-64)

This Slater exchange potential tends to zero as the radius becomes large, while the exact potential tends to \( \alpha/r \). To correct this discrepancy, Latter (1955) has suggested replacing the Slater term in the region of large radius by \( \alpha/r \). Statistical exchange potentials have been discussed extensively by Gombas (1967).

Herman and Skillman (1963) have tabulated nonrelativistic Hartree-Fock-Slater potentials and wave functions for elements with Z=2 to 103, including the Latter tail correction. Extensive nonrelativistic calculations with the exact Hartree-Fock form of the exchange potential have been performed by Froese-Fischer (1972b) and Mann (1967, 1968). Approximate analytic nonrelativistic Hartree-Fock wave functions have been derived by Watson and Freeman (1961a,b), Malli (1966), and Roetti and Clementi (1974).

Because relativistic effects in atomic structure are remarkably important, even for light elements, a number of relativistic self-consistent-field calculations (mostly Hartree-Fock-Slater) have been carried out (Liberman et al., 1965;
Nestor et al., 1966; Tucker et al., 1969). Most comprehensive is the work of Lu et al. (1971), who have published tables of energies and of expectation values of $r$, $r^{-1}$, $r^{-3}$, $r^2$, and $r^4$ for each orbital, of the total energy, and of the potential function. They have included the effect of finite nuclear size, using a Fermi charge distribution.

The possibility of making better approximations than Slater's Kohn and Sham (1965), for the exchange potential has been discussed by Rosen and Lindgren (1968) and Lindgren and Schwarz (1971).

The most sophisticated method of calculating atomic wave functions involves the use of relativistic Hartree-Fock codes; here the exchange term is included without approximation (see e.g. Mann and Weber, 1973; Desclaux, 1973).

Unfortunately, in most published atomic-structure calculations no explicit values are given for the Coulomb amplitudes or electron wave functions at the nuclear radius. For applications to electron capture, special calculations have therefore been carried out; these are listed in Table 2.1. For comparison among the various calculations, the most important electron radial wave-function ratios are listed in Tables 2.2-2.8. For $s$ electrons, the nonrelativistic ratios for a point nucleus are included in the comparison. For $p_{1/2}$ electrons, on the other hand, it is meaningless to compare nonrelativistic wave functions in the field of a point nucleus (proportional to $ar$ at small $r$) with relativistic electron wave functions in the field of a finite nucleus [proportional to $b(1+cr^2+...)$]. We also do not compare
absolute values of electron wave functions, nor do we list other ratios than those contained in Tables 2.2-2.8, because the magnitudes of the nuclear radius $R$ chosen by different authors are not the same, and moreover, some authors report $g_k(R)$ and $f_k(R)$, while others instead report the amplitudes $a_x$.

We can draw the following conclusions from Tables 2.2-2.8:

(1) For the $s$-electron ratios (Tables 2.2-2.5), there is excellent agreement between the nonrelativistic Hartree-Fock calculations of Froese-Fischer (1972b) and Winter (1968) and the relativistic Hartree-Fock calculations of Mann and Waber (1973). An exception is the $g_0^2/g_{n_1}^2$ ratio. However, here relativistic effects might play some role because of the high atomic numbers ($Z>70$).

(2) Relativistic effects become notable in $g_{l_2}^2/g_k^2$ for $Z>15$, in $g_{l_1}^2/g_{l_1}^2$ for $Z>30$, and in $g_{n_1}^2/g_{n_1}^2$ and $g_{o_1}^2/g_{n_1}^2$ for $Z>60$. For $g_{l_1}^2/g_k^2$, relativistic and nonrelativistic ratios differ by ~50% for very heavy nuclei. For all other ratios, relativistic effects are small (<2% for the $g_{m_1}^2/g_{l_1}^2$ and $g_{n_1}^2/g_{m_1}^2$, <8% for the $g_{o_1}^2/g_{n_1}^2$).

(3) The electron radial wave-function ratios from Hartree-Fock calculations lie systematically below those from Hartree-Fock-Slater and Thomas-Fermi-Dirac calculations, especially for low atomic numbers.

(4) For the $K$, $L$, and $M$ ratios, the Hartree-Fock-Slater calculations agree with the Thomas-Fermi-Dirac calculations to within 2.5% for $Z>40$. 
The $g_{1}^{2}/g_{K}^{2}$ ratios of Brysk and Rose (1958) deviate systematically from all other calculations in the range $20<Z<80$ (Table 2.2). Therefore, these values should be discarded.

Of the various methods discussed above, the self-consistent relativistic Dirac-Hartree-Fock calculations of atomic structure are based on the soundest theoretical grounds (Mayer, 1972; Burke and Grant, 1967). It might consequently be expected that the wave functions of Mann and Weber (1973) would be most accurate, and should preferably be used for analyzing electron-capture experiments. Table 2.9 contains a compilation of electron radial wave-function amplitudes according to Mann and Weber (1973).

For practical reason, we have listed the products $\beta_{x}^{k-1}p_{x}$ instead of the amplitudes $\beta_{x}^{k-1}$. It is always this product which appears in the formulation for the decay constant [Eq. (2-49)].

Because the electron-capture rate is essentially proportional to the electron density at the nucleus, different chemical environments or other macroscopic perturbations (pressure, temperature, etc.) can affect the decay constant. Such effects are most noticeable in capture from outer electron shells (Emery, 1972; Crasemann, 1973).

2.2.3. Nuclear Form Factors and Nuclear Matrix Elements

Form factors and form factor coefficients. The electron-capture transition matrix elements [Eq. (2-4)] were formulated in Sec. 2.1 under the assumption that the vector and axial vector interactions govern the process. However, the hadron part of this...
transition matrix element is only an approximation. In the most
general case, we must make the substitution
\[
\langle \bar{u} n^{-1} \gamma_{\mu} (1 + \lambda_5) u \rangle \rightarrow \langle \bar{f} | V_{\mu} - A_{\mu} | i \rangle
\]
in Eq. (2-7c), where \( f \) and \( i \) represent the final and initial
nuclear states, respectively. The vector and axial vector had­
ron weak current are denoted by \( V_{\mu} \) and \( A_{\mu} \). According to Stech
and Schülke (1964) and Schülke (1964), we decompose this \( V-A \)
nuclear current into form factors depending on the square of
the momentum transfer (Armstrong and Kim, 1972; Bottino and
Holstein, 1974). Ciochetetti, 1973; Donnelly and Walecka, 1972, 1973; we use a
covariant decomposition, which is strictly valid in the Breit
system. A transformation in the frame in which the initial nu­
cleus is at rest is easily performed because the decay energies
are low compared with the nuclear rest masses. The correction
due to this transformation is of the order \( |\vec{k}|/M \), where \( \vec{k} \) is
the momentum of the nucleus and \( M \) is its mass. In this approxi­
mation, the hadron matrix element depends only on the momentum
transfer \( q = k_f - k_i \). It can be expanded as
\[
\langle \bar{f} | V_{\mu} - A_{\mu} | i \rangle \gamma_{\mu} = \sum_{KLM} (-1)^{J_f - M_f + M_i} (-1)^L (4 \pi)^{1/2}
\times (2J_f + 1)^{1/2} \left[ \begin{array}{ccc}
J_f & K & J_i \\
- & M_f & M_i
\end{array} \right] T_{KLM}^{M_f}(q) \frac{(qR)^L}{(2L+1)\Gamma} F_{KLM}(q^2).
\] (2-66)

Here, \( T_{KLM}^{M_f}(q) \) is the irreducible tensor defined by Eqs. (2-12);
\( J_i, J_f \) and \( M_i, M_f \) denote the spins and magnetic quantum numbers
of the initial and final nuclear states, respectively, and $R$ is
the nuclear radius.

This treatment of the nuclear current, similar to methods
used in elementary-particle physics, has the advantage of being
completely independent of any assumption about the detailed form
of the $\beta$-decay operators. All information about the nuclear cur-
rent and all effects due to the strong interaction (induced terms,
exchange currents, relativistic nucleon motion inside the nucleus,
etc.) are contained in the form factors $F_{K\ell s}(q^2)$; they determine
the outcome of $\beta$-decay and electron capture experiments and are
the only quantities, as far as nuclear structure is concerned,
that can be extracted from experimental data.

We neglect, for the moment, the initial electromagnetic
interaction between electron and nucleus, i.e. the fact that
there is a bound electron in the initial state. Then the form
factors $F_{K\ell s}(q^2)$ can be expanded in powers of $q^2$ (in analogy
with the expansion of spherical Bessel functions (Stech and
Schülke, 1964)):

$$F_{K\ell s}(q^2) = F^{0}_{K\ell s} - \frac{(qR)^2}{2(2\ell+3)} F^{1}_{K\ell s} + \cdots$$ (2-67)

The form-factor coefficients are then

$$F^{N}_{K\ell s} = \frac{(-1)^{N}(2N+2\ell+1)!}{R^{2N}(2L+1)!N!} \left( \frac{\partial}{\partial q^2} \right)^{N} F_{K\ell s}(q^2) \bigg|_{q^2=0}$$ (2-68)

These form-factor coefficients contain all the information about
the initial and final nuclear states and the $V-A$ operator. Since
$q$ equals $W_0$ if the initial nucleus is at rest, we have $qR<0.1$,
whence the form factors are slowly varying functions of $q^2$. Therefore, only the first one or two terms will be significant.

In reality, however, we must take into account the fact that there is a bound-state electron wave function in the initial state. Hence, the momentum transfer $\hat{q}_N^+$ to the nucleus is $\hat{q}_N^+ = \hat{p}_e - \hat{k}_x$ if the center of mass of the initial nucleus and electron is at rest. The Fourier transform of the lepton part of Eq. (2-10c) is

$$L(\hat{q}_N^+) = \int e^{-i\hat{q}_N^+ \cdot \hat{r}} \bar{\nu}_\gamma(1+\gamma_5)\phi_{-\rho}(\hat{r}) d^3r$$

or

$$L(\hat{q}_N^+) = \bar{\nu}_\gamma(1+\gamma_5)\phi_{-\rho}(\hat{q}_N^+ \hat{q}_x^+)$$

(Schopper, 1966; Stech and Schülke, 1964). Hence, Eq. (2-7c) becomes

$$M_{EC} \approx \int <f| e^{-i\hat{q}_1^+ \cdot \hat{k}_1^-} \bar{\nu}_\gamma(1+\gamma_5)\phi_{-\rho}(\hat{q}_N^+ \hat{q}_x^+) dq_N^+ dq_1^+ dq_x^+$$

The hadron matrix element corresponds to a transition from the initial state $i'$ to the final state $f$, whereas the Fourier transform $\phi_{-\rho}(\hat{q}_N^+ \hat{q}_x^+)$ induces an electromagnetic transition from $i$ to $i'$. The integral over $q_N^+$ corresponds to an integration over all momenta of the intermediate initial states, because we have $\hat{q}_N^+ = \hat{k}_1^- - \hat{k}_x^-$. The Coulomb interaction in the initial state therefore entails that terms of the form

$$\int I(q_N^+) \bar{F}_{KLs}(q_N^2) q_N^2 dq_N^+$$

(2-72)
appear in electron-capture formulae (Schulke, 1964), where \( I(q_N^1) \)
has four different possible forms [Eq. (2-42)]:

\[
I(q_N^1) = \frac{(q_N^1)^L}{(2L+1)^{1/2}} \int \left\{ g_k(r) j_L(q_N^1 r) j_L(q_N^1 r) \right\} r^2 dr
\]

\[
\int f_k(r) j_L(q_N^1 r) j_L(q_N^1 r) \right\} r^2 dr
\]

\[
\int f_k(r) j_L(q_N^1 r) j_L(q_N^1 r) \right\} r^2 dr.
\] (2-73)

By expanding the spherical Bessel functions in powers of \( r \) and
the electron radial wave functions \( g_k(r) \) and \( f_k(r) \) as discussed
in Sec. 2.2.2. [Eqs. (2-49) and (2-51)], we obtain new form-fac­
tor coefficients (Behrens and Bühring, 1971)

\[
F_{KLS}^N (k, m, n, \rho) = \int_0^\infty J(q) F_{KLS}^N (q^2) q^2 dq,
\] (2-74)

where

\[
J(q) = \frac{2}{\pi} \frac{(qR)^L}{(2L+1)^{1/2}} \int_0^{kR} (k, m, n, \rho; r) j_L(qr) r^2 dr.
\] (2-75)

Terms in which these new form factors occur always contain powers
of \( aZ \). Terms that are independent of \( aZ \) contain the simpler form­
factor coefficients \( F_{KLS}^N \) [Eqs. (2-67) and (2-68)].

**Relation between form-factor coefficients and nuclear matrix elements.** The form factors or form-factor coefficients can only
be expressed in terms of nuclear matrix elements, in general, if
some approximations are made. First, it is assumed that the
nucleons inside the nucleus interact with leptons in the same way as free nucleons do (impulse-approximation treatment). Meson exchange (Blin-Stoyle, 1973; Lock et al., 1974) and other many-body effects are hence neglected.

The β-decay Hamiltonian must be used with various many-body nuclear wave functions that can only be calculated in the framework of specific nuclear models. Thus, the uncertainties of nuclear-structure theory are carried over into the nuclear matrix elements or form-factor coefficients.

Finally, the axial-vector constant $\lambda$ for nucleons embedded in a complex nucleus is renormalized in a different way from that for free nucleons, because the mesonic currents behave differently for free and bound nucleons, and new mesonic currents appear that are absent for free nucleons. Thus, $\lambda$ is in principle not a constant over the whole range of nuclei. For light nuclei, a deviation of $\lambda$ from the free-nucleon value by $\sim 7\%$ has been found (Szybisz, 1975; Wilkinson, 1973a, 1973b, 1974a; Ericson et al., 1973; Ohta and Wakamatsu, 1974).

Under these assumptions, we develop the relation between form-factor coefficients and nuclear matrix elements for a pure $V$-$A$ nucleon current of the form given in Eq. (2-4). Induced terms will be discussed later. We have (Stech and Schullke, 1964; Behrens and Bühning, 1971)

\[
V_{\text{KLS}}^{F_{R}}(k_x, m, n, \rho) = (-1)^{K-L} V_{\text{KLS}}^{N}(k_x, m, n, \rho)
\]

\[
A_{\text{KLS}}^{F_{R}}(k_x, m, n, \rho) = (-1)^{K-L} A_{\text{KLS}}^{N}(k_x, m, n, \rho),
\]
where the nuclear matrix elements are denoted by $V_{KLs}^{N}(k_{x},m,n,\rho)$ and $A_{KLs}^{N}(k_{x},m,n,\rho)$. The meaning of the indices has been explained in connection with the form-factor coefficients.

The nuclear matrix elements are [Eq. (2-42)]

$$\begin{align*}
(-1)^{J_{f}-M_{f}} & \left[\begin{array}{ccc}
J_{f} & K_{f} & J_{i} \\
-M_{f} & M & M_{i}
\end{array}\right] \left\{ V_{KLs}^{N}(k_{x},m,n,\rho)+\lambda A_{KLs}^{N}(k_{x},m,n,\rho) \right\} \\
& \times \sum_{j=1}^{A} \left[ I(k_{x},m,n,\rho;1+\lambda)T_{KLs}^{M} \right]^{j} \\
& \times \psi_{f}(1,2,\ldots;A;J_{f},M_{f},\pi_{f})d\tau_{1}d\tau_{2}\ldots d\tau_{A} \tag{2-77}
\end{align*}$$

Here, $\psi_{f}$ and $\psi_{i}$ are the nuclear many-particle wave functions of the final and initial state, respectively, which depend on all the coordinates of the A nucleons. The sum over $j$ runs over the A nucleons, and all the operators are single-particle operators operating on the $j$th nucleon only. The $t^{+}$ is the isospin operator changing a proton into a neutron. The term with 1 gives the $V$ matrix element while the term with $\lambda\gamma_{5}$ leads to the $A$ matrix element. The multipole operators $T_{KLs}^{M}$ have been defined in Eqs. (2-45).

The nuclear matrix elements of Eq. (2-77) must be calculated on the basis of appropriate nuclear models. This is a complicated problem which requires special considerations for each particular $\beta$ transition. One-body operators $O_{r}^{M}$ must be used in Eq. (2-77), which can be expanded (Bohr and Mottelson, 1969) as

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\[ 0 = \sum_{\alpha, \beta} \langle \alpha | 0 | \beta \rangle a^\dagger_\beta a_\alpha, \quad (2-78) \]

where \( a^\dagger_\beta \) is the annihilation operator for a proton in the single-particle state \( \beta \) and \( c^\dagger_\alpha \) is the creation operator for a neutron in the single-particle state \( \alpha \). Here, \( \alpha \) and \( \beta \) represent a complete set of single-particle quantum numbers. We can, therefore, write

\[ \sum_{\alpha, \beta} (-1)^{J_\alpha + J_\beta - K} (2K + 1)^{1/2} \langle \alpha | c^\dagger_\alpha | 0 \rangle \langle 0 | a_\beta | \beta \rangle, \quad (2-79) \]

with \( i = V, A \) and

\[ V_{i,k} = \left( \frac{J}{K} \right)^{L_{x+y} + 2N} I(k, n; \rho; r) Y_{KL}^M, \]

\[ A_{i,k} = \left( \frac{J}{K} \right)^{L_{x+y} + 2N} I(k, n; \rho; r) Y_{KL}^M. \]

The expansion coefficients \( C_{\alpha \beta} \) are

\[ C_{\alpha \beta} = \langle \psi_f | \left[ \left( \frac{J}{K} \right)^{J_{F} + K} \right] \left[ \frac{-1}{2} \langle 0 | a^\dagger_\beta a_\alpha | 0 \rangle \rangle, \quad (2-81) \]

where

\[ a_{i=0} = (-1)^{J_{F} + 1} a_{J_{F} - 1}. \]

It follows that, however complicated the nuclear states may be, the exact nuclear matrix elements between many-body states can be expanded in a linear combination of single-particle matrix elements (Donnelly and Walecka, 1972, 1973). For example, methods of calculating the coefficients \( C_{\alpha \beta} \) in the framework of the shell model are discussed by de Shalit and Talmi (1963). Formulae for nuclear matrix elements within the isospin formalism are also given by de Shalit and Talmi (1963).
Once the set of numerical coefficients \( C_{\alpha\beta} \) has been determined, the nuclear matrix elements can be computed if we are able to deduce reliable values for the single-particle matrix elements. In Eqs. (2-84), we therefore list the single-particle expressions for all the nuclear matrix elements in terms of radial-integral and angular-momentum quantum numbers (Brysk, 1952; Talmi, 1953; Rose and Osborn, 1954; Berthier and Lipnik, 1966; Lipnik and Sunier, 1966; Delabaye and Lipnik, 1966; Strubbe and Callebaut, 1970).

The compact form of Eqs. (2-84) is that given by Behrens and Bühring (1971). The orbits of the nucleons are assumed to have definite angular momentum, as in the \( jj \)-coupling shell model. In the same notation as used for the electron wave functions [Eqs. (2-16) and (2-17)], the nuclear wave functions can be written

\[
\phi^\mu_K(r) = \begin{cases} 
(\text{sign } \kappa) f^\mu_K(r) \chi^\mu_{\kappa} \\
g^\mu_K(r) \chi^\mu_{\kappa} 
\end{cases} \tag{2-82}
\]

The orbit of a nucleon is identified by the number \( \kappa \), defined as for leptons:

\[
|\kappa| = j+ (1/2); \begin{cases} 
k > 0 \text{ if } \ell = j+(1/2) \\
k < 0 \text{ if } \ell = j-(1/2)
\end{cases} \tag{2-83}
\]

The large component of the nuclear radial wave functions is denoted by \( g_K \) and the small component, by \( f_K \). The single-particle values of the nuclear matrix elements then are
\[ V_{VX}^N(k_x, m, n, \rho) = \sqrt{2J_{\frac{1}{2}} + 1}^{-1/2} \left\{ \frac{g_{\text{KKO}}(\kappa_f, \kappa_i)}{g_{\text{KLO}}(\kappa_f, \kappa_i)} \right\} \]

\[ \times \int_0^\infty g_f(r, \kappa_f) \left( \frac{r}{R} \right)^{K+2N} I(k_x, m, n, \rho; r) g_1(r, \kappa_i) r^2 dr \]

\[ + \text{sign} (\kappa_f) \text{sign} (\kappa_i) g_{\text{KJO}}(-\kappa_f, -\kappa_i) \]

\[ \times \int_0^\infty f_f(r, \kappa_f) \left( \frac{r}{R} \right)^{K+2N} I(k_x, m, n, \rho; r) f_1(r, \kappa_i) r^2 dr \}, \quad (2-84a) \]

\[ A_{VX}^N(k_x, m, n, \rho) = \sqrt{2J_{\frac{1}{2}} + 1}^{-1/2} \left\{ \frac{g_{\text{LLO}}(\kappa_f, \kappa_i)}{g_{\text{LLO}}(\kappa_f, \kappa_i)} \right\} \]

\[ \times \int_0^\infty g_f(r, \kappa_f) \left( \frac{r}{R} \right)^{L+2N} I(k_x, m, n, \rho; r) g_1(r, \kappa_i) r^2 dr \]

\[ + \text{sign} (\kappa_f) \text{sign} (\kappa_i) g_{\text{LLO}}(-\kappa_f, -\kappa_i) \]

\[ \times \int_0^\infty f_f(r, \kappa_f) \left( \frac{r}{R} \right)^{L+2N} I(k_x, m, n, \rho; r) f_1(r, \kappa_i) r^2 dr \}, \quad (2-84b) \]

\[ A_{\text{KKO}}^N(k_x, m, n, \rho) = \sqrt{2J_{\frac{1}{2}} + 1}^{-1/2} \left\{ \frac{\text{sign} (\kappa_i)}{g_{\text{KJO}}(\kappa_f, \kappa_i)} \right\} \]

\[ \times \int_0^\infty g_f(r, \kappa_f) \left( \frac{r}{R} \right)^{K+2N} I(k_x, m, n, \rho; r) f_1(r, \kappa_i) r^2 dr \]

\[ + \text{sign} (\kappa_f) g_{\text{KJO}}(-\kappa_f, \kappa_i) \]

\[ \times \int_0^\infty f_f(r, \kappa_f) \left( \frac{r}{R} \right)^{K+2N} I(k_x, m, n, \rho; r) g_1(r, \kappa_i) r^2 dr \}, \quad (2-84c) \]
The indices $i$ and $f$ refer to the initial and final states of the nucleon undergoing decay. The radial quantum numbers of the orbits are not explicitly indicated. The quantity $G_{KIS}(n_f,n_i)$ is defined through Eq. (2-36).

If relativistic nuclear wave functions are used (Miller and Krutov et al., 1974; Green, 1972; Miller, 1972; Krutov and Savashkin, 1973), the nuclear radial wave functions must be normalized to satisfy the condition

$$\int_0^{\infty} g^2(r,\kappa) r^2 dr + \int_0^{\infty} f^2(r,\kappa) r^2 dr = 1. \quad (2-85)$$

In most cases, relativistic nuclear wave functions are not known, whence actual calculations must be performed in the context of nonrelativistic nuclear models. It is then necessary to find the small components $f(r,\kappa)$ of the nuclear radial wave functions. It is possible to express $f(r,\kappa)$ in terms of $g(r,\kappa)$ by using the Dirac equation in the nonrelativistic limit, if the spin angular and the radial parts of the wave functions are considered separately (Behrens and Bühning, 1971). In the nonrelativistic limit one then finds
where $M$ is the nucleon mass, and $g(r, \kappa)$ is the solution of the single-particle Schrödinger equation. In this case, the radial wave functions $g(r, \kappa)$ must be normalized according to

$$\int_0^\infty g^2(r, \kappa) r^2 dr = 1. \quad (2-87)$$

The matrix elements of Eqs. (2-84a) and (2-84b) are usually called nonrelativistic because their radial parts depend only on the radial functions $g(r, \kappa)$. The terms containing both $f_f(r, \kappa_f)$ and $f_i(r, \kappa_i)$ constitute small relativistic corrections that can usually be omitted. On the other hand, the matrix elements of Eqs. (2-84c) and (2-84d), which contain $f(r, \kappa)$, are called relativistic matrix elements.

The radial momentum operator $p_r$ is

$$p_r = \frac{1}{i} \left( \frac{d}{dr} + \frac{1}{r} \right); \quad (2-88)$$

hence we have

$$f(r, \kappa) = \frac{1}{2M} \left[ p_r + \frac{\kappa}{r} \right] g(r, \kappa). \quad (2-89)$$

For a bound nucleon state in a spherical potential, on the other hand, the relation

$$\frac{1}{2M} \left( p_r^2 + \frac{\kappa(\kappa+1)}{r^2} \right) g(r, \kappa) = E_{\text{kin}} g(r, \kappa) \quad (2-90)$$

holds, where $E_{\text{kin}}$ is the kinetic energy of the nucleon. The ratio of relativistic to nonrelativistic single-particle matrix elements
can, therefore, be estimated as

\[ \frac{M_R}{M_{NR}} \approx \left( \frac{E_{kin}}{2m} \right)^{1/2} \approx 0.1. \]  

(2-91)

It has been shown that some approximations must be made in going from relativistic nuclear wave functions to the nonrelativistic limit. Some of the relativistic form-factor coefficients, however, can be related to nonrelativistic coefficients on the basis of CVC theory (Stech and Schülke, 1964; Fujita, 1962; Eich-Schopper, 1966; Damgaard and Winther, 1965; Blin-Stoyle and Nair, 1966; Blin-Stoyle, 1973). The most important such relations are (Behrens and Bühring, 1971)

\[ -V_{C1}^{N-1}(k_x, m, n, \rho) = \left( \frac{W_0 + 2.5}{R} \right) \int_0^R \left( \frac{k_x}{R} \right)^{2N-1} I(k_x, m, n, \rho; x) dx T_{000i} T_{000j}^{-1} \]

\[ -\alpha Z \left( \frac{W_0 + 2.5}{R} \right) \int_0^R \left( \frac{k_x}{R} \right)^{2N-1} I(k_x, m, n, \rho; x) dx U(r) T_{000i} T_{000j}^{-1} \]  

(2-92)

and

\[ -(2K+1+2N) \left[ K/(2K+1) \right]^{1/2} V_{F_{KK-11}}^{N-1} - 2N \left[ (K+1)/(2K+1) \right]^{1/2} V_{F_{KK+11}}^{N-1} \]

\[ = \left( \frac{W_0 + 2.5}{R} \right) \left[ \left( \frac{k_x}{R} \right)^{K+2N} U(r) T_{KK00}^{-1} \right]. \]  

(2-93)

Additional relations are given by Behrens and Bühring (1971).

Because the old Cartesian notation for nuclear matrix elements is used in many papers, the connection between form-factor coefficients and nuclear matrix elements is listed in Cartesian notation in Table 2.10.
**Induced interactions.** As indicated in Sec. 2.1, the hadron current is influenced by the presence of the strong interactions. It can be shown (Delorme and Rho, 1971), that hence the simple nuclear current of Eq. (2-4) must be replaced by the most general current

\[
J_\mu = i\bar{\psi} (p') \sum_{i=A, P, T} \left[ g_1 O^{\delta}_\mu (i\gamma_\mu + M) \psi (p) \right] + \sum_{i=A, P, T} \left[ g_1 O^{\delta}_\mu (i\gamma_\mu + M) \psi (p) \right]
\]

where we have

\[
O^{V}_\mu = \gamma_\mu, O^{S}_\mu = i q_\mu, O^{M}_\mu = q_\mu, O^{A}_\mu = \gamma_\mu \gamma_5, O^{P}_\mu = i q_\mu \gamma_5
\]

and

\[
O^{\gamma}_\mu = \gamma_\mu q_\lambda \gamma_5
\]

Because the binding energy \( B \) of the nucleons inside the nucleus is always small compared with their mass \( M \), the off-mass-shell effects are expected to be negligible (of order \( B/M \)). In the standard impulse-approximation treatment, the nucleons are therefore taken on their mass shell, i.e., \( (i\gamma_\mu + M) u(p) = 0 \) is assumed.

Then the terms associated with the coupling constants \( F_i, G_i, \) and \( H_i \) vanish. On replacing \( q_\mu \) by the corresponding differential operator (Behrens and Bühring, 1971; 1974) we obtain

\[
J_\mu = i\bar{\psi} (p') \left[ \gamma_\mu + \frac{ieA_\mu}{\partial x_{\mu}} + \frac{iA_\mu}{\partial x_{\mu}} \right] \psi (p) \]

\[
+ \lambda \gamma_\mu \gamma_5 + \frac{ieA_\mu}{\partial x_{\mu}} + \frac{iA_\mu}{\partial x_{\mu}} \right] \psi (p) \]

\[
= i\bar{\psi} (p') \left[ \gamma_\mu + \frac{ieA_\mu}{\partial x_{\mu}} + \frac{iA_\mu}{\partial x_{\mu}} \right] \psi (p)
\]

\[
+ \lambda \gamma_\mu \gamma_5 + \frac{ieA_\mu}{\partial x_{\mu}} + \frac{iA_\mu}{\partial x_{\mu}} \right] \psi (p)
\]
for the case of $\beta^-$ decay. In $\beta^+$ decay and electron capture, the Hermitian conjugate current is

$$
\jmath_\mu^+ = i \bar{\psi}_n \left[ \gamma_\mu \left( \frac{\partial}{\partial x_\nu} - i e A_\nu - f S \left( \frac{\partial}{\partial x_\mu} - i e A_\mu \right) \right) \right] \psi_p.
$$

(2-96)

By comparing Eqs. (2-95) and (2-96), the formal substitutions can be determined that must be made for the induced coupling constants in going from $\beta^-$ to $\beta^+$ decay, from $\beta^+$ decay to electron capture, or from $\beta^-$ decay to electron capture (Behrens and Bühring, 1974).

Between $\beta^-$ and $\beta^+$ decay, the following correspondences hold [in addition to those indicated in Eq. (2-39)]:

$$
\begin{align*}
\beta^- \text{ decay} & \quad \beta^+ \text{ decay} \\
 f_M & \rightarrow f_M \\
f_S & \rightarrow -f_S \\
f_T/\lambda & \rightarrow -f_T/\lambda \\
f_F/\lambda & \rightarrow f_F/\lambda \\
-i e A_\mu & \rightarrow -i e A_\mu
\end{align*}
$$

(2-97)

For $\beta^+$ decay and electron capture, the hadron current (and therefore also the hadron part of the transition matrix element) has the same form. Thus, beyond the substitutions indicated in Eq. (2-40), it is only necessary to replace $\mathcal{W}$ by $\mathcal{W} + \mathcal{W}_x$ to go from $\beta^+$ decay to electron capture.
Starting from $\beta^-$ decay, on the other hand, the following substitutions apply [in addition to those indicated in Eqs. (2-41) and (2-42)]:

\[
\begin{align*}
\beta^- \text{ decay} & \quad \text{electron capture} \\
\begin{array}{c|c|c}
\hline
f_M & f_M \\
\hline
f_S & -f_S \\
\hline
f_p/\lambda & -f_p/\lambda \\
\hline
f_p/\lambda & f_p/\lambda \\
\hline
ieA_\mu & -ieA_\mu \\
\hline
\end{array}
\end{align*}
\]

The quantities $f_M$, $f_S$, $f_T$, and $f_p$ are the coupling constants for the weak magnetic, induced scalar, induced tensor, and induced pseudoscalar interactions, respectively (Marsbak et al., 1969; Schopper, 1966; Blin-Stoyle, 1973; Blin-Stoyle and Nair, 1966; Kim, 1974).

The conserved-vector-current theory predicts the values (Blin-Stoyle and Nair, 1966)

\[
\begin{align*}
\frac{f_M}{2M} = \frac{\mu_p - \mu_n}{2M} \approx 0.0010, \\
\frac{f_S}{2M} = 0
\end{align*}
\]

for $f_M$ and $f_S$; here, $\mu_p$ and $\mu_n$ are the anomalous magnetic moments of the proton and neutron, and $M$ is the nucleon mass.

The quantity $A_\mu = (A, i\phi)$ in Eq. (2-98) is the potential of the external electromagnetic field, which in this case is the static electric field of the nuclear charge, for which we have $A_\mu = 0$. 

The terms containing $A$ must be included to assure gauge invariance of the Hamiltonian.

By applying the Dirac equation, the operator $\frac{\partial}{\partial \xi_4}$ in Eq. (2-96) can be replaced by the transition energy $\omega = \omega_0 + i \omega_x$.

Like the simple current of Eq. (2-4), the general current given by Eq. (2-96) consists of two parts, one of which Lorentz-transforms like a four-vector, the other like an axial vector. We make use of this property. In the nuclear matrix elements without induced interactions [Eq. (2-77)], the spherical tensor operators $T_{KLs}$ and $\gamma_5 T_{KLs}$ occur:

$$
\begin{align*}
1 \cdot T^{M}_{LLLO} &= 1 \cdot L^{M}_{L} \\
\lambda \gamma_5 T^{M}_{LLLO} &= \lambda \gamma_5 \cdot \frac{1}{i} Y^{M}_{L} \\
1 \cdot T^{M}_{KL} &= \lambda \cdot (-1)^{L-K+1} L^{M}_{L} \\
\lambda \gamma_5 T^{M}_{KL} &= \lambda \gamma_5 \cdot (-1)^{L-K+1} L^{M}_{L} 
\end{align*}
$$

(2-100)

The nuclear operators $1$, $\lambda \gamma_5$, $\lambda$ and $\lambda \gamma_5$ behave under rotation like scalars, pseudoscalars, vectors, and axial vectors, respectively. Introduction of the general current of Eq. (2-96) makes it necessary to replace these operators by more complicated operators which have the same transformation properties (Behrens and Bühring, 1971):
These substitutions, in Eq. (2-76) via Eq. (2-77), lead to the form-factor coefficients that correspond to the general nuclear current. The expressions for the observables in terms of form-factor coefficients remain unchanged (Sec. 2.2.3; Stech and Schulke, 1964; Bühring and Schulke, 1965). Only the definition of the form-factor coefficients in terms of nuclear matrix elements and coupling constants is changed.

The form-factor coefficients in terms of nuclear matrix elements appropriate for electron capture are as follows:

\[
V_F^N(k, m, n, \rho) = V_M^N(k, m, n, \rho)
\]

\[
+ f_M^{-1}\left\{ \frac{K}{(2K+1)} \right\}^{1/2}\left[ \left( \frac{r}{R} \right)^{K+2N-1} \left( 2(2K+2N)I(r) + rI'(r) \right) \beta_{KK-11}^T \right] + \left( \frac{K+1}{2K+1} \right)^{1/2} \left( \frac{2N+1}{2N+2} \right) \left( \frac{r}{R} \right)^{K+2N-1} \left( 2N+1 \right) \beta_{KK+11}^T \right) \right] \]

\[
+ f_S^{-1}\left[ \left( \frac{r}{R} \right)^{K+2N} \left( w_{1R} - \alpha Z U(r) \right) \beta_{KK}^T \right] \right), \quad (2-102a)
\]

\[
A_F^N(k, m, n, \rho) = \lambda A_M^N(k, m, n, \rho)
\]

\[
- f_M^{-1}\left\{ \frac{K}{(2K+1)} \right\}^{1/2}\left[ \left( \frac{r}{R} \right)^{K+2N-1} \left( 2(2K+2N)I(r) + rI'(r) \right) \beta_{KK-11}^T \right] + \left( \frac{K+1}{2K+1} \right)^{1/2} \left( \frac{2N+1}{2N+2} \right) \left( \frac{r}{R} \right)^{K+2N-1} \left( 2N+1 \right) \beta_{KK+11}^T \right) \right] \]

\[
- f_P^{-1}\left[ \left( \frac{r}{R} \right)^{K+2N} \left( w_{1R} - \alpha Z U(r) \right) \beta_{KK}^T \right] \right), \quad (2-102b)
\]
\[
V_{FKKL}(k, m, n, \rho) = V_{KKL}(k, m, n, \rho)
\]

\[
\begin{aligned}
& V_{FKKL}(k, m, n, \rho) = V_{KKL}(k, m, n, \rho) \\
& \left. \right\} \text{(2-102c)}
\end{aligned}
\]

\[
A_{FKKL}(k, m, n, \rho) = \lambda_{KKL}(k, m, n, \rho)
\]

\[
\begin{aligned}
& A_{FKKL}(k, m, n, \rho) = \lambda_{KKL}(k, m, n, \rho) \\
& \left. \right\} \text{(2-102d)}
\end{aligned}
\]

\[
V_{FKKO}(k, m, n, \rho) = V_{KKO}(k, m, n, \rho)
\]

\[
\begin{aligned}
& V_{FKKO}(k, m, n, \rho) = V_{KKO}(k, m, n, \rho) \\
& \left. \right\} \text{(2-102e)}
\end{aligned}
\]

\[
A_{FKKO}(k, m, n, \rho) = \lambda_{KKO}(k, m, n, \rho)
\]

\[
\begin{aligned}
& A_{FKKO}(k, m, n, \rho) = \lambda_{KKO}(k, m, n, \rho) \\
& \left. \right\} \text{(2-102f)}
\end{aligned}
\]
$$-V_{F}^{N\kappa+1}(k, m, n, \rho) = V_{m_{N}}^{K+1}(k, m, n, \rho)$$

$$+ f_{N}^\alpha \left\{ \frac{[K/(2K+1)]^{1/2}}{2} \left( \int \frac{(r/R)^{K+2N} I(r + rI(r)) \beta_{S}^{N_{K+1}}} {W_{R}^{\alpha}(r) \beta_{S}^{N_{K+1}}} \right) \right\}$$

$$- f_{N}^\alpha \left\{ \frac{[(K+1)/(2K+1)]^{1/2}}{2} \left( \int (r/R)^{K+2N} I(r + rI(r)) \beta_{S}^{N_{K+1}} \right) \right\}$$

For brevity, we have written $I(r)$ instead of $I(k_{x}, m, n, p; r)$. We have $I'(r) \equiv dI/dr$.

In addition to the single-particle matrix elements of Eqs. (2-84), the following are required:

$$\left( \int \frac{(r/R)^{K+2N} \phi(r) \beta_{S}^{N_{K+1}}} {W_{R}^{\alpha}(r) \beta_{S}^{N_{K+1}}} \right) = \frac{2}{2J_{l}+1} \left\{ - C_{K+1}(k_{\|}, k_{\perp}) \right\}$$

$$\times \int_{0}^{\infty} \bar{g}_{l}(r, k_{\|}) \left( \frac{r}{r/R} \right)^{K+2N} \phi(r) \bar{g}_{l}(r, k_{\|}) r^{2} dr$$

$$+ \text{sign} \left( k_{\|} \right) \text{sign} \left( k_{\perp} \right) G_{K+1}(k_{\|}, k_{\perp}) \int_{0}^{\infty} \phi(r) \bar{g}_{l}(r, k_{\|}) r^{2} dr$$

(2-103a)
\[
\left\{ \int (r/R)^{L+2N} \phi(r) \beta_{Y^{K\text{KL}}_L} \right\} = \left[ 2/(2J_{i+1}) \right]^{1/2} \left\{ -g_{KL} (\kappa_1, \kappa_1) \right. \\
\times \int_0^\infty g_{\zeta}(r, \kappa_1) (r/R)^{L+2N} \psi(r) g_{\zeta_1}(r, \kappa_1) r^2 dr \\
+ \text{sign} (\kappa_1) \text{sign} (\kappa_1) G_{KL} (\kappa_1, -\kappa_1) \\
\left. \times \int_0^\infty f_{\xi}(r, \kappa_1) (r/R)^{L+2N} \psi(r) f_{\xi_1}(r, \kappa_1) r^2 dr \right\}. \tag{2-103b}
\]

\[
\left\{ \int (r/R)^{K+2N} \phi(r) \beta_{Y^{T\text{KKO}}_L} \right\} = \left[ 2/(2J_{i+1}) \right]^{1/2} \left\{ -\text{sign} (\kappa_1) G_{KKO} (\kappa_1, -\kappa_1) \right. \\
\times \int_0^\infty g_{\zeta}(r, \kappa_1) (r/R)^{K+2N} \psi(r) f_{\zeta_1}(r, \kappa_1) r^2 dr + \text{sign} (\kappa_1) G_{KKO} (-\kappa_1, \kappa_1) \\
\left. \times \int_0^\infty f_{\xi}(r, \kappa_1) (r/R)^{K+2N} \psi(r) g_{\xi_1}(r, \kappa_1) r^2 dr \right\}. \tag{2-103c}
\]

\[
\left\{ \int (r/R)^{L+2N} \phi(r) \beta_{Y^{T\text{KL}}_L} \right\} = \left[ 2/(2J_{i+1}) \right]^{1/2} \left\{ -\text{sign} (\kappa_1) G_{KL} (\kappa_1, -\kappa_1) \right. \\
\times \int_0^\infty g_{\zeta}(r, \kappa_1) (r/R)^{L+2N} \psi(r) f_{\zeta_1}(r, \kappa_1) r^2 dr + \text{sign} (\kappa_1) G_{KL} (-\kappa_1, \kappa_1) \\
\left. \times \int_0^\infty f_{\xi}(r, \kappa_1) (r/R)^{L+2N} \psi(r) g_{\xi_1}(r, \kappa_1) r^2 dr \right\}. \tag{2-103d}
\]

Here, \( \phi(r) \) stands for \( I(k, m, n, \rho; r) \) or \( \rho I'(k, m, n, \rho; r) \) or a linear combination of these integrals. The question whether a finite coupling constant \( f_{T} \) exists for the induced tensor interaction has aroused great interest of late. Second-class currents (Weinberg, 1958) manifest themselves in principle only through the
coupling constants $f_S$ and $f_T$, and $f_S$ vanishes in accord with the conserved vector current theory. Hence, the determination of $f_T$ is connected with the very question of the existence of second-class currents in $\beta$ decay and electron capture. Although this problem has been discussed extensively in the literature (Wilkinson, 1970a, 1971, 1972a, 1971/72, 1974a; Alburger and Wilkinson, 1970; Kim, 1971; Holstein and Treiman, 1971; Vatai, 1971, 1972a; Wilkinson and Alburger, 1971; Blomquist, 1971; Wolfenstein and Henley, 1971; Lipkin, 1970, 1971; Kim and Fulton, 1971; Blin-Stoyle et al., 1971; Laverne and Dang, 1971; Alburger, 1972; Tribble and Garvey, 1974; Towner, 1973; Greenland, 1975) an unambiguous answer concerning the existence of second-class currents has not yet been obtained. An excellent review of this matter has been written by Wilkinson (1971/72).

In view of the uncertainty about second-class currents, Kubodera et al. (1973) have recently pointed out that one cannot neglect the nucleon binding effects, i.e., off-mass-shell phenomena and exchange currents. Thus, at least as far as the axial-vector part is concerned, one should start with the most general current [Eq. (2-94)]. But then the large number of coupling constants complicates the problem to such an extent that it can be dealt with only under some simplifying assumptions, i.e., minimal coupling. Furthermore, special models for the meson exchange current must be used. Following this line of attack, Kubodera et al. (1973) were able to calculate explicitly off-mass-shell and meson-exchange effects for some special cases, and to demonstrate their importance (Eman et al., 1973).
2.2.4. Explicit Expression for the Quantities $M_K(k_x,k_y)$ and $m_K(k_x,k_y)$

By expanding electron and neutrino radial wave functions as outlined in Sec. 2.2.2 and introducing the form-factor coefficients defined in Sec. 2.2.3, we can derive from Eq. (2-42) explicit expansions of the quantities $M_K(k_x,k_y)$ and $m_K(k_x,k_y)$. If we take into account only dominant terms (of lowest order in the expansion of electron and neutrino radial wave functions), we arrive at the following simple forms for $M_K(k_x,k_y)$ and $m_K(k_x,k_y)$ (Behrens and J"{u}nck, 1969; Behrens and B"{u}hring, 1971):

For allowed transitions,

\[
M_0(1,1) = V_{F000},
\]

\[
M_1(1,1) = -A_{F101};
\]  

(2-104)

for first-forbidden transitions,

\[
M_0(1,1) = A_{F000} + (1/3)aZ A_{F011}(1,1,1,1) - (1/3)\tilde{W}_0 R A_{F011},
\]

\[
m_0(1,1) = (1/3)R A_{F000},
\]

\[
M_1(1,1) = -V_{F101} + (1/3)aZ(1/3)^{1/2} \tilde{V}_{F110}(1,1,1,1)
\]

\[
- (1/3)\tilde{W}_0 R(1/3)^{1/2} \tilde{V}_{F110} - (1/3)aZ(2/3)^{1/2}
\]

\[
\times A_{F111}(1,1,1,1) - (1/3)(\tilde{W}_x + \tilde{W}_y) R(2/3)^{1/2} A_{F111},
\]

\[
m_1(1,1) = (1/3)R \left[ (1/3)^{1/2} \tilde{V}_{F110} - (2/3)^{1/2} A_{F111} \right].
\]
\[ M_1(1,2) = -(1/3)q_x R \left( (2/3)^{1/2} V_{FLO}^{0} (k_1,2) \right)^{1/2} \]
\[ M_1(2,1) = -(1/3)p_x R \left( (2/3)^{1/2} V_{FLO}^{0} (k_2,1) \right)^{1/2} \]
\[ M_2(1,2) = -(1/3)q_x R A_{FLO}^{0} \]
\[ M_2(2,1) = -(1/3)p_x R A_{FLO}^{0} \]  \hspace{1cm} (2-105)

For higher forbidden transitions, we have

\[ m_L(k_x, k^{(1)}_y) = K_L(p_x R) \left[ k^{(1)}_x \right]^{-1} (q_x R) \left[ k^{(1)}_y \right]^{-1} \left\{ \left[ (2L+1)/L \right]^{1/2} V_{FLO}^{0} \right\} \]
\[ + (2k_x+1)^{-1/2} \zeta \left[ V_{FLO}^{0} (k_x,1,1,1) + \left[ (2k_x+1)^{-1} \right] \left[ (2k_y+1)^{-1} q_x R \right] \right] \]
\[ \times \left[ V_{FLO}^{0} \left[ (2L+1)/L \right]^{1/2} A_{FLO}^{0} \right] \]  \hspace{1cm} (2-106a)

\[ m_L(k_x, k^{(2)}_y) = -K_L(p_x R) \left[ k^{(2)}_x \right]^{-1} (q_x R) \left[ k^{(2)}_y \right]^{-1} \left\{ \left[ (2L+1)/L \right]^{1/2} V_{FLO}^{0} \right\} \]
\[ \times \left[ V_{FLO}^{0} + (k_x k^{(2)}_y) (L+1)^{-1/2} \left[ (2k_x-1) (2k_y-1) \right]^{1/2} \right] \]  \hspace{1cm} (2-106b)

\[ m_L+1(k_x, k^{(2)}_y) = -K_L(p_x R) \left[ k^{(2)}_x \right]^{-1} (q_x R) \left[ k^{(2)}_y \right]^{-1} A_{FLO}^{0} \]  \hspace{1cm} (2-106c)
Here we have introduced the abbreviations

\[ K_L = (1/2)^{1/2} \left[ (2L)!! / (2L+1)!! \right]^{1/2} \left[ (2k_x - 1)!! (2k_v - 1)!! \right]^{-1/2} \]  

(2-107a)

\[ \tilde{K}_L = \left[ (2L)!! / (2L+1)!! \right]^{1/2} \left[ (2k_x - 1)!! (2k_v - 1)!! \right]^{-1/2}. \]  

(2-107b)

The two quantities \( K_L \) and \( \tilde{K}_L \) are related by

\[ \tilde{K}_L = [(2L+1)/L]^{1/2} K_L. \]  

(2-108)

The energy of the bound electron in the parent atom is defined as \( \psi_x = |E_x| \), where \( E_x \) is the binding energy in the parent atom. The electron momentum \( p_x \) is given by

\[ p_x = (1 - \kappa_x^2)^{1/2}. \]  

(2-109)

The form-factor coefficients are \( V F_{KLS}^{MN}, A F_{KLS}^{MN}, V F_{KLS}^{MN}(k_x, m, n, \rho) \) and \( A F_{KLS}^{MN}(k_x, m, n, \rho) \); they are related to the nuclear matrix elements as indicated previously. The symbols \( V \) and \( A \) refer to vector and axial vector; \( K \) specifies the rank, \( L \) the multipolarity, and \( s \) the spin of the spherical tensor operators that are involved. The radial dependence of this operator is \( r^{L+2N} \) or \( r^{L+2N} I(k_x, m, n, \rho; r) \).

These form-factor coefficients occur in accordance with the expansion of the electron radial wave functions discussed in Sec. 2.2.2.

In Eqs. (2-47) through (2-49) we have only presented the dominant terms of the multipole expansion and the expansion of the electron radial wave functions for linear combinations of form-factor coefficients. Complete expressions are listed in Appendix A2.1 (Behrens and Bühring, 1971). Unless there are strong cancellations between different terms connected with the form-factor coefficients, the higher-order terms can be neglected.
2.3. Formulae for Allowed and Forbidden Transitions

2.3.1. Allowed Transitions

In allowed transitions, electrons can only be captured from orbits with \( \kappa = \pm 1 \), i.e., from the K, L_1, L_2, M_1, M_2, \ldots \) shells [cf. Eqs. (2-44)-(2-47)]. This result is based on the approximate neglect of contributions from higher-order (so-called second-forbidden) terms (see Appendix A2.2). Capture from orbits with \( \kappa = \pm 2 \), for example, would be governed by matrix-element combinations \( M_1(2,1), M_2(2,1) \), etc., which are smaller than \( M_0(1,1) \) and \( M_1(1,1) \) by at least a factor \( p_{RS} \approx 0.02 \). Consequently, we have

\[
C_x(\kappa=\pm2) \leq 4 \times 10^{-4} C_x(\kappa=\pm1)
\]

[Eq. (2-44)], and capture from orbits with \( \kappa = \pm 2 \) can be expected to be difficult to observe. However, capture from such states in principle offers a possibility of determining the higher-forbidden contributions separately from the leading terms.

For the quantity \( C_x \) we find

\[
C_x = (V_0^0)^2 + (A_{101}/A_{101})^2
\]

[ Eqs. (2-44) and (2-104) ]. Inserting this result in Eq. (2-43) leads to

\[
\lambda_0 = (g^2/4\pi^2) \left[ (V_0^0)^2 + (A_{101}/A_{101})^2 \right] n_{1} q_{1}^{2} b_{1}^{2} + \ldots
\]

\[
+ n_{1} q_{1}^{2} b_{1}^{2} + \ldots + n_{1} q_{1}^{2} b_{1}^{2} + \ldots
\]

(2-111)
for the decay constant. Hence it is easy to derive the ratios of the capture probabilities from different subshells. The \( \frac{\lambda_{L_1}}{\lambda_K} \) ratio, for example, is

\[
\frac{\lambda_{L_1}}{\lambda_K} = \left( \frac{n_{L_1} q_{L_1}^2 B_{L_1}}{n_K q_K^2 B_K} \right). \tag{2-112}
\]

2.3.2. *First-Forbidden Non-Unique Transitions*

Considering, as before, only the dominant terms in non-unique, first-forbidden transitions, we find that electrons with the quantum numbers \( k_x = \pm 1, \pm 2 \) are captured. \( K, L_1, L_2, M_1, M_2, \ldots \) capture, we have

\[
C_x = \left[ M_0(1,1) F_{000}^0(1,1) \right]^2 + \left[ k_x(1,1) F_{011}^0(1,1) \right]^2 + M_2^2(1,2) + M_2^2(1,2) \tag{2-113}
\]

[Eqs. (2-44)-(2-46)]. The upper sign holds for \( K, L_1, M_1, \ldots \) capture and the lower, for \( L_2, M_2, \ldots \) capture. The quantities \( M_L(k_x,k_y) \) in Eq. (2-113) are defined through Eqs. (2-105). If there is no cancellation between the different terms in Eqs. (2-105), we can simplify Eq. (2-113). Because we have \( W_x = 1 - |E_K| \), with \( |E_K| < 0.2 \) and \( R = 0.0031 A^{1/3} < 0.02 \), we can usually neglect terms multiplied by \( R \) and \( W_x R \). Then we find (Vatai, 1973)

\[
C_x = \left[ A_{F011}^0(1/3) a_z A_{F000}^0(1/3) - (1/3) a_z A_{F011}^0(1/3) \right]^2 + \left[ V_{F101}^0 - (a_z/3)(1/\sqrt{3}) \left\{ V_{F110}^0(1,1,1,1) - (a_z^2 A_{F111}^0(1,1,1,1) \right\} \right]^2
+ (W_x R/3)(1/\sqrt{3}) \left\{ V_{F110}^0 + (a_z A_{F111}^0) \right\}^2 + (W_x R)^2/9 \left\{ (2/3)^{1/2} V_{F110}^0(1/\sqrt{3}) A_{F111}^0 \right\}^2 + (A_{F211}^0)^2. \tag{2-114}
\]
This result shows that, even in the case of first-forbidden non-unique transitions, the quantity \( C_x \) to a very good approximation does not depend on the particular subshell from which the electron is captured. As for allowed electron capture, the ratios of the capture probabilities from different subshells are therefore independent of the form-factor coefficients. Thus, these ratios have the same form as given in Eq. (2-81).

In many cases, especially for the heavier nuclei, we have \( aZ \gg W_0 R \). Then Eq. (2-114) can be simplified further:

\[
C_x = \left[ A_{F 00}^2 (1/3) aZ A_{F 011}^2 (1,1,1,1) \right]^2
+ \left[ V_{F 10}^O (aZ/3) (1/\sqrt{3}) \left[ V_{F 110}^O (1,1,1,1) - \sqrt{2} A_{F 111}^O (1,1,1,1) \right] \right]^2. \tag{2-115}
\]

For capture from \( \kappa = \pm 2 \) \((L_2, M_2, M_4, \ldots)\) states, we have

\[
C_x = \{ M_1 (2,1) \}^2 + \{ M_2 (2,1) \}^2, \tag{2-116}
\]

or explicitly [cf. Eqs. (2-105)],

\[
C_x = \left[ (p_x R)^2 / 2 \right] \left[ (2/3)^{1/2} V_{F 110}^O (1/\sqrt{3}) A_{F 111}^O \right]^2 \left[ A_{F 211}^O \right]^2. \tag{2-117}
\]

Comparison of Eq. (2-117) with Eq. (2-114) suggests that \( \kappa = \pm 2 \) capture is negligibly small as against capture from \( \kappa = \pm 1 \) orbits.

### 2.3.3. First-Forbidden Unique Transitions

Considering dominant terms in Eqs. (2-105) for unique first-forbidden transitions, we find that subshells with \( \kappa = \pm 1, \pm 2 \) can contribute (Behrens and Jönecke, 1969). For capture from \( \kappa_x = \pm 1 \) \((K, L_1, L_2, L_1, M_1, M_2, \ldots)\) states, we have
and for capture from $x_1 = \pm 2$ \( L_3, M_3, M_4, \ldots \) orbits, we find

\[ C_x = \left( A_{F,0}^{211} \right)^2 (R^2/9) q_x^2. \]  

(2-119)

It follows from Eqs. (2-43) and (2-118) that the \( L_1/K \) capture ratio is

\[ \lambda_{L_1} / \lambda_K = \frac{n_{L_1} \beta_{L_1}^2 \beta_{L_1}^2 \beta_{L_1}^2}{n_K \beta_{K}^2 \beta_{K}^2 \beta_{K}^2}. \]  

(2-120)

Expressions for the \( L_2/K, M_1/K, L_2/L_1, \) \( M_1/L_1 \) capture ratios are entirely analogous. For the \( L_3/L_1 \) ratio, on the other hand, we have

\[ \lambda_{L_3} / \lambda_{L_1} = \frac{n_{L_3} \beta_{L_3}^2 \beta_{L_3}^2 \beta_{L_3}^2}{n_{L_1} \beta_{L_1}^2 \beta_{L_1}^2 \beta_{L_1}^2}. \]  

(2-121)

Other \( k_x = 2 \) to \( k_x = 1 \) capture ratios are analogous to Eq. (2-121).

2.3.4. \( (L-1) \)-Forbidden Unique Transitions

Taking only dominant terms in Eq. (2-106d) into account, we have for \( L_x k_x \)

\[ C_x = \left( \frac{(2L-2)!!}{(2L-1)!!} \right) \left( A_{F,0}^{L1-1} \right)^2 R_{2(L-1)} \frac{P_x 2(L-1) q_x}{(2k_x-1)! (2(L-k_x)+1)!}. \]  

(2-123)

For \( K, L_1, L_2, M_1, M_2, \ldots \) capture, for example, we obtain

\[ C_x = \left( A_{F,0}^{L1-1} \right)^2 \left\{ \frac{(2L-1)!!}{(2L-1)!!} \right\}^{-2} (q_x R)^2 (2L-1). \]  

(2-124)
2.3.5. Some General Remarks on Higher-Forbidden Non-Unique Transitions

Special formulae for the higher-forbidden non-unique capture rates can easily be derived from Eqs. (2-106) in analogy with the first-forbidden non-unique transition rate [Eqs. (2-113) to (2-115)]. The following general statements can be made regarding such higher-forbidden capture transitions:

(i) As for $\Delta J = 1$ first-forbidden non-unique transitions, these capture rates depend only on six different form-factor coefficients, viz., $V_F(0)$, $V_F(0)$, $V_F(0)$, $V_F(L_0, L_0^0, L_0, L_0^0, L_0, L_0^0)$, $A_F(0)$, $A_F(L_0, L_0, L_0)$, $A_F(L_0, L_0, L_0)$. Expressions for these rates are therefore no more complicated than those for first-forbidden transitions.

(ii) If we neglect terms multiplied by $R$ and $W_x$, as in Eq. (2-114), the capture ratios from shells with the same $k_x$ value do not depend on the nuclear form-factor coefficients. Form-factor coefficients can therefore be determined by investigating capture ratios only if ratios of capture from states with different $k_x$ are measured (e.g. $L_2/K$, $M_2/K$) (Vatai, 1973).

(iii) Non-unique $L^{th}$-forbidden capture rates are always proportional to a factor

$$\frac{1}{(2L+1)!} \frac{\hbar (q_x R)^{2L}}{(p_x/q_x)^{2k_x}}$$

[Eqs. (2-106)]. Consequently, such capture probabilities decrease very rapidly with increasing degree of forbiddenness.
2.4. Electron-Capture to Positron-Decay Ratios

2.4.1. General Expressions

For allowed as well as forbidden transitions, the following general result for $\lambda_{EC}/\lambda_{\beta^+}$ ratios holds [Eqs. (2-2), (2-7), (2-10)]:

$$\lambda_{EC}/\lambda_{\beta^+} = \left( \sum_\chi n_x C_x \right) \left( \frac{f_{\beta^+} \overline{C(W)}}{f_{\gamma}} \right). \quad (2-125)$$

Here, $f_{\beta^+}$ is the integrated Fermi function (Behrens and Jönecke, 1969):

$$f_{\beta^+} = \int_0^{p_0} p^2 (W_0 - W)^2 F(Z', W) dp, \quad (2-126)$$

where $p$ is the positron momentum (in units of $m_0c$), the maximum momentum is $p_0 = (W_0^2 - 1)^{1/2}$, $W$ is the positron energy (in units of $m_0c^2$), $Z'$ is the atomic number of the daughter nucleus, $F(Z', W)$ is the Fermi function, and $\overline{C(W)}$ is the spectrum shape factor, averaged over the $\beta^+$ spectrum. The form of the shape factor for different types of $\beta^+$ decay has been discussed, for example, by Schopper (1966), Behrens and Jönecke (1969), and Behrens and Bühring (1971).

To calculate the integrated Fermi function $f$ we need the continuum-electron radial wave functions $g_{-1}(r)$ and $f_{+1}(r)$. Conventionally, these functions (and hence the Fermi function) are evaluated at the nuclear radius ($r=R$). However, recent discussions indicate that a less ambiguous result is achieved if the Fermi function is evaluated at the center of the nucleus ($r=0$) (Schopper, 1966; Behrens and Bühring, 1968, 1972; Blin-Stoyle, 1969). This
latter definition of the Fermi function is appropriate for the
electron-capture formalism in the present paper (Sec. 2.2.2).
A number of detailed calculations and tabulations of the Fermi
function $F(Z, W)$ and of the integrated Fermi function $f(Z, \omega_0)$
exist. However, in many instances finite nuclear size and
screening by orbital electrons has not been taken fully into
account. The Fermi function for a point nucleus without screen­
ing is listed in the National Bureau of Standards tables (1952)
and in a paper by Rose and Perry (1953). Dzhelepov and Zyryanova
(1956) have calculated the Fermi function and the integrated
Fermi function (at $r=R$) by adding corrections for screening and
finite size to the functions for a point nucleus. Several
authors (Matese and Johnson, 1966; Durand, 1964; Brown, 1964),
however, have noted that the screening corrections of Reitz
(1950) used by Dzhelepov and Zyryanova are incorrect for higher
electron momenta.

Fermi functions evaluated numerically (at $r=R$) from an exact
solution of the Dirac equation for a nucleus with finite size,
but without screening, have been tabulated by Bhalla and Rose
(1960, 1961, 1962, 1964). It was later shown, however, that
these tables are not entirely correct for positrons of higher
momenta (Bühring, 1967; Huffacker and Laird, 1967; Behrens and
Asai and Ogata, 1974). Bühring, 1968; Blin-Stoyle, 1973, p. 38; For a few elements,
Bühring (1965) has carried out an exact numerical integration
of the Dirac equation, taking into consideration finite nuclear
size and screening. By employing a method similar to that of Bühring, extensive tables of the Fermi function (at r=0) and graphs of the integrated Fermi function have been published by Behrens and Jäncke (1969); this calculation takes exact account of both finite nuclear size and electron screening. Numerical integration of the Dirac equation, including finite size and screening, has also been carried out by Suslov (1966, 1967, 1968a). Theoretical K/β⁺ ratios have also been listed by Suslov (1970b). The extensive tabulations of the Fermi function (at r=R) and of the integrated Fermi function by Dzhelepov, Zyryanova, and Suslov (1972) are based on these calculations. Suslov, however, included in the electrostatic potential caused by the atomic electrons a Slater exchange term. While the exchange term is applicable to the bound orbital electrons, it is not appropriate for the continuum states; this is self-evident for positrons and has also been shown for emitted β⁻ particles (Matese and Johnson, 1966; Behrens and Jäncke, 1969, p. 25).

It may be for this reason that Suslov's calculations do not agree at low β⁺ energies with his Thomas-Fermi-Dirac calculations and with results of other authors (Behrens and Jäncke, 1969; Bhalla and Rose, 1960, 1961, 1962, 1964).

An extensive tabulation of log f (at r=R) and of capture-to-positron ratios, with an accuracy of two to three digits, has been compiled by Gove and Martin (1971). These values were obtained
by correcting point-nucleus continuum radial wave functions for
finite nuclear size and screening.

In all calculations discussed so far, the finite size of
the nucleus was represented by the simplest model, viz., a uni-
formly charged sphere of radius R, equal to the nuclear radius.
A more realistic charge distribution has been employed by Behrens
and Bühring (1970), who have shown that the influence of the
shape of the charge distribution on the Fermi function can be
(see also Asai and Ogata, 1974).

An analytical parametrization of the
Fermi function and of the integrated Fermi function (for a point-
like nucleus), of the screening corrections, the finite nuclear-
size effects, and of the dependence of allowed β decay on the
nuclear radius has been derived by Wilkinson (1970b, 1970c, 1970d;

2.4.2. Allowed Transitions

For allowed transitions, for which we have C(W)=C_x=(V\_00)^2
+(A\_01)^2, the EC/β^+ ratio has a very simple form:

$$\lambda_K/\lambda_{\beta^+} = f_K/f_{\beta^+}$$

(2-127)

This ratio consequently does not depend on the form-factor coef-
cients, just like the capture ratios. However, for the EC/β^+ ratio there are two effects that can lead to small deviations
from the result predicted by Eq. (2-127):

(i) If higher-order terms (Appendix A2.1) contribute sig-
nificantly, the differences between C_x and C(W) must be taken
into account [Sec. 2.1; Eq. (2-40)]. For allowed transitions, the cor-
rection factor of Eq. (2-127) can be given explicitly. Neglecting terms
in $F_{121}^{(N)}(l,m,n,p)$ and form-factor coefficients of rank two, we find
(Appendix A2.1; Behrens and Bühning, 1971)

$$\lambda_{K}^\ast = \left( \frac{E_{K}}{E_{\beta^+}} \right)_{\ast} \frac{\lambda_{\beta^+}}{[1+(A_{1}y^{2}A_{2})/(1+y^{2})]} \quad (2-128)$$

where

$$A_{1} = (2/3) \frac{3/2}{2(\overline{W}+\overline{W})} \left[ 1+\left( \frac{1}{\overline{W}} \right) \right] R \left( \frac{V_{F(0)}^{(0)}}{A_{F(0)}^{(0)}} \right)_{111}^{101}$$

$$+ 2(3)^{-3/2} \left[ 1+\left( \frac{1}{\overline{W}} \right) \right] R \left( \frac{A_{F(0)}^{(0)}}{A_{F(0)}^{(0)}} \right)_{110}^{101} - (2/3) \left( \frac{W_{K}+W}{W} \right) \right] R \left( \frac{1}{\overline{W}} \right)$$

$$+ \left[ (1/9) \left( \frac{A_{F(1)}}{A_{F(0)}} \right) \right]_{101}^{101} \left( \frac{A_{F(1)}}{A_{F(0)}} \right)_{101}^{101}$$

$$- (1/2) \left[ 2(\overline{W}+\overline{W}) \right] \left( \frac{A_{F(0)}}{A_{F(0)}} \right)_{101}^{101}$$

$$A_{2} = - (2/3) \left[ 1+\left( \frac{1}{\overline{W}} \right) \right] R \left( \frac{V_{F(0)}^{(0)}}{V_{F(0)}^{(0)}} \right)_{011}^{000} - (2/3) \left( \frac{W_{K}+W}{W} \right) \right] R \left( \frac{1}{\overline{W}} \right)$$

$$\times \left[ \left( \frac{V_{F(1)}^{(1)} \cdot V_{F(0)}^{(0)}}{V_{F(0)}^{(0)} \cdot V_{F(0)}^{(0)}} \right) \right]_{000}^{000}$$

$$+ (2/9) \left[ 2(\overline{W}+\overline{W}) \right] \left[ 1+\left( \frac{1}{\overline{W}} \right) \right] R \left( \frac{V_{F(0)}^{(0)}}{V_{F(0)}^{(0)}} \right)_{101}^{101}$$

$$y = \frac{V_{F(0)}^{(0)}}{A_{F(0)}^{(0)}}. \text{ Here, the energy } \overline{W} \text{ and the Coulomb function } \overline{W} \text{ are averaged over the } \beta^+ \text{ spectrum (Behrens and Jänecke, 1969); } \gamma_1 \text{ stands for } [1-(\alpha Z)^{2}]^{1/2}.$$
and hence very small. Thus, we generally have $y \ll 1$ (Blin-Stoyle, 1973; Bertsch and Mekjian, 1972). Hence $A_1$ is the important correction term. The form-factor coefficient ratio $V_{\text{rel}}(0)/A_{\text{rel}}(0)$, relativistic over nonrelativistic, depends sensitively on the nuclear structure and is difficult to calculate. This ratio is of the order $\sim 0.1$. The ratios $A_{\text{rel}}(0)/A_{\text{rel}}(0)$, $A_{\text{rel}}(1)/A_{\text{rel}}(1)$, and $A_{\text{rel}}(1)/A_{\text{rel}}(1)$ can however be estimated more easily. They generally lie in the range $0.5-2.0$. Taking into account only the latter form-factor coefficient ratios leads to the estimate $A_1 = 0.03$ for $Z=80$.

(ii) A second cause for deviations of the $E0/\beta^+$ ratio from the prediction of Eq. (2-127) lies in electromagnetic radiative corrections to the electron-capture and $\beta^+$ decay rates, for example for the emission of internal bremsstrahlung. Radiative corrections for allowed $\beta$ transitions, especially for the superallowed $0^+ \to 0^+$ transitions, have been discussed extensively (Marshak et al., 1969; Sirlin, 1967; Källen, 1967; Dicus and Norton, 1970; Beg et al., 1972; Jaus and Rasche, 1970; Jaus, 1972; Sirlin, 1974; Roos, 1974; Suzuki and Yokoo, 1975).

For allowed $\beta$ transitions, the effect of radiative corrections can be described, first, by a renormalization of the vector and axial-vector coupling constants,

\begin{align*}
C_V &= C_V(1 + \frac{\alpha_C}{\hbar c}), \\
C_A &= C_A(1 + \frac{\alpha_D}{\hbar c}).
\end{align*}

(2-131)
(Blin-Stoyle, 1973), and second, by a known modification of the β spectrum. This second point affects the integrated Fermi function:

\[ f_{\beta^+} + f_{\beta^+}(1 + \delta_R(W,Z)). \tag{2-133} \]

In Eqs. (2-131) and (2-132), C and D are the so-called model-dependent radiative corrections; they depend on details of the weak and strong interaction theories (Sirlin, 1967; Källen, 1967; Dicus and Norton, 1970; Beg et al., 1972; Sirlin, 1974; Roos, 1974; Wilkinson, 1975). These model-dependent radiative corrections cannot as yet be calculated without ambiguity, but they cancel in EC/β⁺ ratios. The model-independent radiative correction factor \( [1 + \delta_R(W,Z)] \) is well-known to order \( a \) (Sirlin, 1967; Källen, 1967; Dicus and Norton, 1970). This correction factor can be found, for example, in the work of Wilkinson and Macefield (1970), where semianalytical formulae and nomograms are given. The terms of order \( Z a^2 \) and \( Z^2 a^3 \) have also been calculated (Jaus and Rasche, 1970; Jaus, 1972).

For electron capture this model-independent part of the radiative correction differs, however, from that discussed for β⁺ decay. Unfortunately, no explicit calculation has been carried out as yet.

Some contrary statements notwithstanding (Vatai, 1971, 1972b; Eman et al., 1973), Behrens and Bühring (1974) have pointed out that the existence of second-class currents, i.e., of a finite value of \( f_T \), does not significantly affect EC/β⁺ ratios. This fact follows in principle from the equality \( 14 \)
of the hadron parts, or of the form-factor coefficients, for electron capture and $\beta^+$ decay (Sec. 2.2.3).

2.4.3. **Non-Unique Forbidden Transitions**

The $EC/\beta^+$ ratios for non-unique forbidden transitions are proportional to an additional factor $C_x/C(W)$. The quantity $C_x$ is given by Eqs. (2-44), (2-105), and (2-106). The corresponding formulae for the shape factor $C(W)$ can, for example, be found in the papers by Behrens and Jänecke (1969) and in Behrens and Bähring (1971). These formulae show that the $EC/\beta^+$ ratios for non-unique forbidden transitions generally depend on the relative values of the nuclear form-factor coefficients, i.e., on the details of the nuclear structure.

There is one exception from this rule, however, in the case of non-unique first-forbidden transitions. When the $\xi$-approximation [Eq. (2-115)] is applicable, the $EC/\beta^+$ ratios from $k_x = 1$ states are independent of the nuclear matrix elements, and have the same values as for allowed transitions. The applicability of the $\xi$-approximation can be tested experimentally by investigating the shape factor of the $\beta^+$ spectrum.

2.4.4. **Unique Forbidden Transitions**

For the $(I-1)^{st}$ unique forbidden transitions, explicit expressions for the ratios $C_x/C(W)$ can be given. The formulae for $C_x$ can be taken from Eq. (2-123), and for $C(W)$, for example, from the work of Behrens and Jänecke (1969). We find
Here, $X$ is a special Coulomb function defined, for example, by Behrens and Jönncke (1969). As before, barred symbols denote quantities averaged over the $\beta^+$ spectrum.

For $K, L_1, L_2, M_1, M_2, \ldots$ capture, Eq. (2-134) takes the simpler form

$$C_x/C(y) = \left( \frac{2(L-1)}{2(L+1)+1} \right)^{-1} \left( \frac{2(k_x-1)}{q_x} \right)^{2(L-k_x)}$$

$$\times \left\{ \sum_{n=1}^{L} \lambda_n \frac{2(n-1)}{q_x} \left( \frac{2n}{2n+1} \right) \right\}^{-1} \left( \frac{2(2n-1)}{2(2n+1)+1} \right)^{-1}. \quad (2-135)$$
2.5. Atomic Matrix Elements: Exchange and Overlap Corrections

2.5.1. Introduction

According to the usual theory of allowed orbital electron capture (Sec. 2.3), the probability that a K electron is captured by the nucleus is

\[
\lambda_K = G^2 q \xi |\psi_K(0)|^2 ,
\]  

(2-136)

where \( G \) is the \( B \)-decay coupling constant, \( q \) is the energy of the neutrino that is emitted, \( \xi \) is the appropriate combination of nuclear matrix elements, and \( |\psi_K(0)|^2 \) is the square of the parent atom's ls electron wave function at the nucleus. In Eq. (2-136), no atomic matrix elements are included.

Benoist-Gueutal (1950, 1953b) first suggested that atomic electrons must be included in a complete description of the nuclear electron-capture process. She estimated the effect of imperfect atomic overlap on the total electron capture rate of \(^7\)Be by calculating the electron-capture probability for various final atomic states. Due to the lack of accurately known wave functions for excited Li atoms, Benoist-Gueutal only concluded that the decrease in the total decay rate was less than 30%. Odiot and Daudel (1956) made a quantitative calculation of the \(^{37}\)Ar L-to-K capture ratio, using wave functions for the entire atom. Odiot and Daudel's prediction of 0.10 for the \(^{37}\)Ar L-to-K capture ratio has subsequently been verified by experiment.

The discrepancy between the traditional theory of electron capture (Brysk and Rose, 1958) and experiments on L-to-K electron-capture ratios indicated that a critical examination of the theory was needed. Bahcall
(1962a, 1963a, b, 1965a) made a comprehensive study of the role of atomic electrons in the nuclear electron-capture process, emphasizing the importance of the indistinguishability of electrons and of the change in nuclear charge by one unit from initial to final atomic states, aspects which were neglected in the usual theory. In Bahcall's work, ground-state wave functions were used for the initial and final atoms. The importance of the presence of an inner-shell hole in the daughter atom was pointed out by Vatai (1968b).

In this section, we consider the effect of atomic overlap and exchange corrections on the total electron-capture rate and on various subshell capture ratios. We also discuss the calculation of atomic matrix elements. This subject has recently been reviewed by Genz (1973a) and Vatai (1973c). The calculations of electron density at the nuclear surface are discussed in Sec. 2.2.

2.5.2. Effect of Atomic Overlap and Exchange on Total Capture Rates

Bahcall (1963a, b) used second quantization to formulate the nuclear electron-capture process. For allowed transitions, the probability per unit time that a nucleus will capture any of its atomic electrons and leave the daughter atom in the final state $|A'\rangle$ is

$$
\lambda(A') = g^2 \xi (2\pi)^{-1} q^2 (A') M^\dagger (A') (1+\gamma) M(A'),
$$

(2-137)

where

$$
M(A') \equiv <A'|\phi_e(0)|G>
$$

(2-138)

and

$$
q(A') = W_0 + 1 + [E(G) - E(A') - 1].
$$

(2-139)
Here, $W_0$ is the difference between initial and final nuclear masses; $E(G)$ and $E(A')$ are the total energies of the initial and final atomic electrons, including their rest masses.

If one uses a single-particle representation of $|G>$, the total electron-capture rate can be written

$$\lambda = \lambda_0 G_1 \frac{\Delta G}{\lambda_0 E_0},$$  \hspace{1cm} (2-140)$$

where

$$\lambda_0 = G_2 \xi (2\pi)^{-1} \sum_b q_b |\psi_b(0)|^2$$  \hspace{1cm} (2-141)$$
is the usual total electron capture rate. We have

$$\lambda' = G_2 \xi^{-1} q_b |\psi_b(0)|^2 \left[ -\sum_{b',b''} \frac{q_{b''} - q_{b'}}{2} \right]$$

$$+ \sum_{b',b''} q_{b''} |\psi_{b''}(0)|^2 \left[ -\sum_{b',b''} \frac{q_{b''} - q_{b'}}{2} \right].$$  \hspace{1cm} (2-142)$$

and

$$\Delta \lambda = q(ls') G_2 \xi^{-1} \sum_{b_1,b_2} q_{b_1} |\psi_{b_1}(0)|^2 \left[ -\sum_{b_1,b_2} \frac{q_{b_2} - q_{b_1}}{2} \right].$$  \hspace{1cm} (2-143)$$

and

$$q(ls') = W_0 + E(G) - E(G') - \epsilon (ls'),$$  \hspace{1cm} (2-144)$$

and

$$\Delta q(A') = E(G') - E(A') + \epsilon (ls'),$$  \hspace{1cm} (2-145)$$

where $\epsilon (ls')$ is the $K$ binding energy in the final atom.

The second and third terms in Eq. (2-140) are the contributions due to imperfect atomic overlap and-exchange capture, respectively. By applying closure to sum the electron-capture probability over all possible final atomic states, Bahcall found
The contributions of overlap and exchange are of the opposite sign. They partially cancel each other in the total capture rate. The net effect on the total capture probability does not exceed a few percent if $q(1s')$ is greater than, or of the order of, 50 keV.

2.5.3. Overlap and Exchange Corrections on Capture Ratios

The electron-capture rate, including the atomic matrix element in the theory, can be written

$$\lambda_i = \lambda_i^0 B_i, \quad i = K,L,M,N,...,$$  

(2-148)

where $\lambda_i$ is the transition rate from the usual theory and $B_i$ is the exchange-correction factor introduced by Bahcall to take account of the exchange and overlap contribution.

For allowed transitions, the L/K capture ratio can then be written

$$\frac{\lambda_L}{\lambda_K} = \frac{\lambda_L^0 B_L}{\lambda_K^0 B_K} \left[ 1 + \frac{B_L}{B_K} \left( \frac{\lambda_L^0}{\lambda_K^0} \right) \right].$$  

(2-149)

For unique forbidden transitions, the L/K ratio becomes

$$\frac{\lambda_L}{\lambda_K} = \frac{\lambda_L^0 B_L}{\lambda_K^0 B_K} \left[ 1 + \frac{B_L}{B_K} \left( \frac{\lambda_L^0}{\lambda_K^0} \right) \right] \left[ \frac{3(\Delta J-1)(2\Delta J-1)}{(g_{L2} R_{L2})^2 \frac{B_{L2}}{B_{L1}}} \right].$$  

(2-150)
where

\[
\left( \frac{\lambda_i}{\lambda_j} \right)^0 = \frac{\frac{2}{q_i} - \frac{2}{q_j}}{\frac{2}{q_j}}
\]

(2-151)

\[
\left( \frac{\lambda_{L1}}{\lambda_{K}} \right)^0 = -\frac{2}{q_{L1}} \left( \frac{\Delta J}{q_{K}} \right)
\]

(2-152)

The \( q \)'s are neutrino energies and the \( g \)'s, charge densities, at the nuclear surface.

In Eqs. (2-149) and (2-150), the difference in binding energy among the \( L \) subshells has been neglected.

A similar expression applies for \( M/L \) capture ratios:

\[
\frac{\lambda_M}{\lambda_L} = \left( \frac{\lambda_{M1}}{\lambda_{L1}} \right) \frac{B_M}{B_L} \left[ 1 + \left( \frac{\lambda_{M2}}{\lambda_{M1}} \right) \frac{B_M}{B_L} - \left( \frac{\lambda_{L2}}{\lambda_{L1}} \right) \frac{B_L}{B_L} \right]
\]

(2-153)

Most theoretical and experimental work has been done on \( K, L \) and \( M \) capture for allowed transitions. Little research has been performed on \( N \) capture. We proceed to review various theoretical calculations dealing with the overlap and exchange corrections.

Bahcall's approach. In order to overcome the difficulty of calculating and summing an infinite number of separate contributions from the final atomic states, Bahcall (1962a, 1963a, b, 1965a) used the following approximations: (1) The innermost electrons are almost inert. (2) The outer-electron states (outside the 3s shell) form a practically complete set. (3) The energy available for a given nuclear transition is nearly independent of the particular states occupied by the outer electrons in the final atom.
Bahcall separated the atomic state vectors into two independent parts,

$$|\text{atomic}\rangle = |\text{inner}\rangle \times |\text{outer}\rangle.$$  \hspace{1cm} (2-154)

He then invoked closure to perform the sum over the infinite number of final atomic states, obtaining

$$\lambda_1^0 = \lambda_1 B_1,$$  \hspace{1cm} (2-155)

where $\lambda_1^0$ is the usual electron-capture rate, and we have

$$B_1 = \left| \frac{\psi_1(0)}{\psi_1(0)} \right|^2.$$  \hspace{1cm} (2-156)

The capture amplitudes are

$$f(3s') = \langle 1s' | 1s \rangle \langle 2s | 3s \rangle - \langle 1s' | 3s \rangle \langle 2s | 1s \rangle (0),$$

$$f(2s') = \langle 1s' | 1s \rangle \langle 3s | 3s \rangle - \langle 1s' | 3s \rangle \langle 3s | 1s \rangle (0),$$

$$f(1s') = \langle 2s' | 2s \rangle \langle 3s | 1s \rangle - \langle 2s' | 1s \rangle \langle 3s | 2s \rangle (0).$$

(2-157, 2-158, 2-159)

The primed orbitals pertain to the daughter atom.

The $L_1$-to-$K$ and $M_1$-to-$L_1$ capture ratios can be written

$$\frac{\lambda_{L_1}}{\lambda_K} = \left( \frac{\lambda_{L_1}}{\lambda_K} \right)^0 B_{L_1} \left( \frac{\lambda_{L_1}}{\lambda_K} \right)^0 \frac{L_1/K}{X},$$

$$\frac{\lambda_{M_1}}{\lambda_{L_1}} = \left( \frac{\lambda_{M_1}}{\lambda_{L_1}} \right)^0 B_{M_1} \left( \frac{\lambda_{M_1}}{\lambda_{L_1}} \right)^0 \frac{M_1/L_1}{X},$$

(2-160, 2-161)
where the exchange correction factors are

\[
\frac{L_1/K}{X} = \frac{B_{L_1}}{B_{K}} = \left| \frac{f(2s')\psi_{1s}(0)}{f(1s')\psi_{2s}(0)} \right|^2
\]

(2-162)

\[
\frac{M_1/L_1}{X} = \frac{B_{M_1}}{B_{L_1}} = \left| \frac{f(3s')\psi_{2s}(0)}{f(2s')\psi_{3s}(0)} \right|^2
\]

(2-163)

To compare these calculated capture ratios with measurements, correction must be made for capture from \( n_{3/2} \) states.

To calculate the atomic matrix elements \( \langle ns'|ns \rangle \), Bahcall used nonrelativistic Hartree-Fock ground-state wave functions for parent and daughter atoms (Watson, 1960; Watson and Freeman, 1961b).

The following comments can be made on Bahcall's theory:

1. The assumption that the neutrino energy is independent of final states of the atom, and the use of the closure approximation without correction for occupied states, tend to lead toward underestimation of the overlap correction.

2. The overlap correction is small for \( K \) and \( L \) capture, but is much larger for \( M \) capture. Therefore, Bahcall's approach will overestimate the \( M \)-to-\( L \) capture ratio correction factor \( X \).

3. Multiple exchange processes and the exchange between inner and outer electrons are neglected.

4. The effect of the inner-shell vacancy in the daughter atom is neglected.

\[ \text{Vatai's ansatz.} \]

\[ \text{Vatai (1968b, 1970a) calculated the capture transition to the most prominent state } |A> \text{ of the final atom. In state } |A>, \]
except for the captured electron, all the other electrons retain their quantum numbers. Vatai obtained the exchange and overlap correction coefficients as

\[ B_i = \left( \frac{f_i}{\psi_i(0)} \right)^2 \]  

(2-164)

and

\[ f_K = \psi_{1s}(0)|2s\rangle\langle 2s'|2p\rangle\langle 2p'|3s\rangle\cdots \]

\[ - \psi_{2s}(0)|2s\rangle\langle 1s'|2p\rangle\langle 2p'|3s\rangle\cdots \]

\[ - \psi_{3s}(0)|3s\rangle\langle 1s'|2s\rangle\langle 2p'|3p\rangle\cdots \]  

(2-165)

Similar expressions for \( f_L \) and \( f_M \) are obtained by exchange of \( 1s \) with \( 2s \) and \( 1s \) with \( 3s \), respectively, in the \( f_K \) expression. If overlap corrections for \( p \) and \( d \) electrons are neglected, one obtains the same \( f_i \) expressions as those of Bahcall [Eqs. (2-157) to (2-159)].

In Vatai's calculation, the effect of the inner hole in the daughter atom on the exchange integral is estimated by perturbation theory.

Vatai used the analytic Hartree-Fock wave functions of Watson and Freeman (1961b) for the initial state and as unperturbed wave functions for the final-state calculation. He estimated the overlap correction for the inner \( p \) and \( d \) electrons including the multiplicity by calculating the overlap integral with the wave functions of Watson and Freeman for both parent and daughter atoms. The overlap integrals of outer electrons are set equal to 1 in Vatai's calculation.

With regard to Vatai's approach, we note the following points:
(1) Some contributions due to processes involving shakeup or
shakeoff are neglected.

(2) The use of perturbation theory to calculate the exchange
integrals introduces a discrepancy of 10-40% in the value of these
integrals compared with Froese's HF calculations (Faessler et al.,
1970).

(3) The overlap corrections are only rough estimates.

(4) Vatai, like Bahcall, neglects multiple exchange processes.

**Faessler's calculation.** Faessler et al. (1970) recalculated the
Bahcall exchange corrections, taking into account the inner-shell vacancy
that after electron capture exists in the daughter atom. Faessler et al.
used the Herman-Skillman (1963) Hartree-Fock-Slater and Froese-Fischer
(1965, 1969) Hartree-Fock programs to calculate hole-state wave func-
tions and exchange and overlap integrals. Although some of the exchange
integrals calculated with the two programs differ by as much as 50%,
the exchange correction factors agree to within 3%. This indicates
that the exchange correction, being a ratio, is insensitive to the model
wave functions, due to cancellation of errors. Faessler et al. con-
cluded that the influence of rearrangement effects on the L/K and M/L
capture ratios is far too small to account for the discrepancy between
theory and experiment, although it does affect the theoretical capture
ratios in the right direction.

**Relativistic calculations.** Suslov (1970a) followed Bahcall's
approach and used relativistic Hartree-Fock-Slater wave functions to
calculate the exchange and overlap corrections for $^{14}\text{ZS}98$. The wave functions were obtained by numerical integration of Dirac's equation, using a nonrelativistic potential (Herman and Skillman, 1963) for $^{14}\text{ZS}73$, and an analogous relativistic potential (Liberman et al., 1965) for $^{15}\text{ZS}74$. Finite nuclear size was included through the uniformly-charged-sphere model. For $^{15}\text{ZS}37$, the new relativistic values of $B_K$, $B_{L_1}$, $B_{M_1}$, $X^{K}$, and $X^{L_1/L_1}$ are quite close to Bahcall's (1963a, b) results; the differences do not exceed 5%. For $^{22}\text{ZS}38$, the exchange correction decreases as Z increases, and for large Z it is nearly constant. The relativistic exchange-corrected capture ratios do not narrow the gap between theory and experiment.

Martin and Blichert-Toft (1970) performed another relativistic calculation of electron-capture ratios for $^{6}\text{ZS}98$ using the same approach as Vatai's. The required wave functions and electron radial densities were calculated with a relativistic Hartree-Slater program with finite nuclear size. The K and $L_1$ electron radial density at the nuclear surface, calculated by Martin and Blichert-Toft (1970), agrees with other calculations (Zyryanova and Suslov, 1968; Behrens and Jänecke, 1969; Winter, 1969; Suslov, 1970b) within 1%, and the exchange-overlap factors agree very well with the present results based on Vatai's approach.
2.5.4. Evaluation of Atomic Matrix Elements

Atomic matrix elements $\langle ms' | ns \rangle$ are not only required for the calculation of exchange and overlap corrections, but also for determining autoionization rates in $\beta$-decay and electron-capture transitions, and for shake-up calculations (Sec. 5). The degree of orthogonality of the wave functions is the important point in the evaluation of the overlap integrals $\langle ms' | ns \rangle$. Overlap integrals that involve ground-state wave functions from parent to daughter atoms are not very sensitive to the choice of the atomic potential, because the inner shells are closed shells. Overlap integrals calculated with the analytic Hartree-Fock wave functions of Watson and Freeman, with Herman-Skillman Hartree-Fock-Slater wave functions, or with Froese-Fischer Hartree-Fock wave functions, all agree to better than 5% (Faessler et al., 1970). However, for calculations of inner-shell vacancy states (e.g., 1s and 2s hole states), the atomic model is important, as the hole-state wave functions are sensitive to the potential. In the Herman-and-Skillman (1963) code, single electronic configurations having open shells are treated on the same basis as configurations having only closed shells. Consequently, the wave function of an electron in an open shell is not necessarily orthogonal to a single-electron wave function that describes an electron of the same symmetry species and in the same configuration, but from a closed shell. For example, the 1s electron wave function for an atom
with a K vacancy may not be orthogonal to the 2s wave function of the atom, if it has a full $l_1$ subshell. The overlap integrals between open-shell and closed-shell single-electron wave functions, involving the ground state of the parent atom and a deep hole state of the daughter, can therefore contain a sizable error if it is computed with Herman-Skillman wave functions (Faessler et al., 1970).

In Froese-Fischer's (1965, 1969) and Bagus' (1964, 1965) approaches, the orthogonality between self-consistent field orbital wave functions with the same symmetry is taken into account by introducing off-diagonal Lagrangian multipliers into the Hartree-Fock equations. For closed shells, a unitary transformation can be found between the occupied orbitals, such that the Lagrangian multipliers are in diagonal form. The additional requirement that the off-diagonal Lagrangian multipliers be zero serves as a unique definition of the self-consistent field orbitals. For open-shell systems, it is not possible to reduce the Lagrangian multipliers that couple open and closed shells of the same symmetry to zero (Roothaan, 1960; Roothaan and Bagus, 1963).

The Ne-like and Ar-like ns hole states have been calculated by Bagus (1964, 1965). The off-diagonal Lagrangian multipliers between open and closed shells $\Theta_{ns,ms}$ are large for ls-hole states and become smaller for 3s hole states. The effect of including the off-diagonal Lagrangian multipliers for Ar-like ions is that the ls orbitals of the 1s-hole states have a node; an extended tail appears in the 1s wave functions (Bagus, 1964). For large $r$, $P_{ls}(r)$ becomes

$$P_{ls}(r) \approx -\frac{\Theta_{2s,ls}}{\varepsilon_{ls}} 2s(r) - \frac{\Theta_{3s,ls}}{\varepsilon_{ls}} 3s(r).$$  (2-166)
The features introduced by the off-diagonal Lagrangian multipliers in the Froese-Fischer Hartree-Fock hole-state wave functions explain the differences between overlap integrals obtained by using Herman-Skillman and Froese-Fischer wave functions in the work of Faessler et al. (1970).

To resolve the discrepancy between the overlap matrix elements \( \langle n' l | n l \rangle \) of Faessler et al. and of Vatai, we have recalculated these matrix elements for Ar K-, L-, and M-capture with Bagus' accurate analytic Hartree-Fock Ar ground-state and Cl- ns hole-state wave functions (Bagus, 1964). Our results from Bagus' wave functions agree with the overlap matrix elements calculated by Faessler et al. (1970) with the Hartree-Fock program of Froese-Fischer to better than 1%.

2.5.5. Comparison Among Theoretical Exchange Corrections to Capture Ratios

In Sec. 2.5.4., we have described evidence that the Hartree-Fock program of Froese-Fischer is best suited for the evaluation of the exchange and overlap integrals. We have therefore recalculated the exchange correction factors using the Froese-Fischer program (Froese-Fischer, 1972a) and have included the effect of the ns hole present in the daughter atom. Two sets of values were computed, one based on Bahcall's approach, the other following Vatai's ansatz that includes the overlap correction for both inner and outer electrons (Table 2.11). The results computed by various workers according to Bahcall's approach (Faessler et al., 1970; Suslov, 1970a; Bahcall, 1963a,b, 1965a, and our present calculations) agree very well (within 5%). The results of
Martin and Blichert-Toft (1970) coincide with our present calculations based on Vatapi's approach. In Figs. 2-2 and 2-3, the exchange correction factors $X_{1}/K$ and $X_{2}/L_{1}$ are shown, as recalculated by us with the Froese-Fischer (1972a) code. For comparison, the results from the two relativistic calculations (Suslov, 1970a; Martin and Blichert-Toft, 1970) are also included. In general, the results from Vatapi's approach are smaller than those following Bahcall's theory.

2.5.6: Correlation Effects in Electron-Capture Ratios

All theoretical work reviewed in Sec. 2.5.3 contains the independent-particle approximation. Effects due to electron correlations are neglected.

Goverse and Blok (1974c) have observed that the experimental L/K capture ratios seem to oscillate about the theoretical curve, and suggested that correlation effects between the orbital electrons might cause this discrepancy. This assertion remains to be proven.

2.5.7: Conclusion

The exchange and overlap correction factors are not very sensitive to the choice of the atomic potential, due to compensation between the electron density at the nucleus and the atomic matrix element $<\text{ns}|\text{ms}>$. The importance of including an appropriate inner-shell hole in the daughter atom after electron capture, stressed by Vatapi (1968b, 1970a), is not borne out by the work of Faessler et al. (1970) nor by our present calculations, if Bahcall's approach is followed. On the other hand, the presence of the inner hole has a significant effect on these correction factors if they are calculated with Vatapi's formulae.
The effect of exchange on electron-capture ratios has been treated in a similar way in the two existing theories, those of Bahcall (1963a,b, 1965a) and Vatai (1970), while the overlap corrections are treated differently. Because the overlap corrections are important for low-Z elements, the difference in exchange and overlap correction factors between Bahcall's and Vatai's approaches shows up clearly in light atoms.

Our recalculated correction factors permit a direct comparison of results based on Bahcall's and Vatai's approaches. Vatai's formulation causes an underestimation of L/K capture ratios at low-Z, but leads to M/L capture ratios in fair agreement with experiment. On the other hand, Bahcall's approach yields better agreement to L/K ratios with experiment, but overestimates the M/L capture ratios.

To solve this problem, a new calculation is needed in which overlap corrections are treated more carefully. Electron correlation must be included, at least by means of configuration interactions. More accurate experimental capture ratios in the low-Z region are needed to provide a better test of theory.
3. EXPERIMENTAL METHODS AND RESULTS

The experimental determination of nuclear electron-capture ratios from various atomic shells and of K-capture to positron-emission (K/β+) ratios has been the subject of considerable effort because of the importance of these quantities in various contexts. Aspects of orbital electron capture have been reviewed by Robinson and Fink (1955, 1960), Boucher and Depommier (1960, 1965), Depommier (1968), Fink (1965, 1966, 1968a, 1968b), Benseny (1963a, 1963b), Genz (1971b, 1973a), and Fitzpatrick (1973). In recent years, several new measurements of L/K, M/L and K/β+ ratios have been performed and much effort has been devoted to reducing experimental uncertainties, so that comparisons can be made with different theoretical calculations of atomic wave functions and of electron exchange and imperfect atomic wave-function overlap effects.

In this section we classify the methods employed to determine capture ratios and compare their potential reliability. From the vast body of experimental data reported in the literature, we select a limited list of capture and K/β+ ratios that can be considered highly reliable and use these values for comparison with theory.

Relative transition probabilities are commonly used in experimental work; these are related as follows to the transition probabilities per unit time as defined in Eqs. (2-27), (2-28) and (2-43):

\[ P_{EC} = \frac{\lambda_c}{\lambda_{tot}}, \quad P_{\beta^+} = \frac{\lambda_{\beta^+}}{\lambda_{tot}}, \quad P_{\beta^-} = \frac{\lambda_{\beta^-}}{\lambda_{tot}} \]  (3-1)
where

\[ P_{EC} + P_{\beta^+} + P_{\beta^-} = 1, \]  

(3-2)

and

\[ P_K = \frac{\lambda_K}{\lambda_C}, \quad P_L = \frac{\lambda_L}{\lambda_C}, \quad P_M = \frac{\lambda_M}{\lambda_C}, \ldots \]  

(3-3)

where

\[ P_K + P_L + P_M + \ldots = 1. \]  

(3-4)

Corresponding relations hold for capture from subshells.

The probability of orbital electron capture from the K shell or from any of the L or M subshells depends upon the nature and energy of the transition. The capture process cannot be detected directly because of the extremely low interaction probability of the emitted neutrino. The capture rate can therefore only be determined from the intensity of subsequently emitted radiation, such as x rays or Auger electrons given off during reorganization of the electronic cloud after capture and \( \gamma \) rays or conversion electrons from the daughter nucleus. In principle, the recoil of the final nucleus can also be measured, but the recoil kinetic energy is always very small. The largest recoil (57 eV) occurs in the transition \( ^7Be \rightarrow ^7Li \).

Methods for measuring capture probabilities vary according to the decay scheme of the radionuclide, the energy and relative intensity of the emitted radiation, available detectors, and requirements for necessary corrections. The methods can be classified according to the information they provide.

One group of methods yields ratios of capture probabilities from
From these ratios, a consistent set of capture probabilities can be deduced with the aid of Eq. (3-4):

\[
P_L = \left( \frac{P_{LM}^2}{P_L^2} \right) \left( \frac{1 + \frac{P_L}{P_M} \left( 1 + \frac{P_M}{P_N} \right)}{1 + \frac{P_L}{P_M}} \right)^{-1},
\]

\[
P_M = \frac{P_L}{P_L^2},
\]

Equations (3-6) can also be used with reliable theoretical capture ratios.

Some methods pertain to situations in which the L and M x-ray or Auger-electron peaks cannot be resolved. Such methods lead to the determination of a capture ratio \( P_{LM} / P_K \) from which the relative K-capture probability can be obtained directly:

\[
\frac{P_{LM}}{P_K} = \frac{1}{P_K} - 1.
\]

In several other methods, \( P_{MK} \) is determined, where \( \omega_K \) is the K-shell fluorescence yield. With the appropriate value for \( \omega_K \) (Bambynek et al., 1972), the relative K-capture probability can be calculated.

If the transition energy exceeds twice the electron rest energy \( (2mc^2) \), then positron emission is possible as an alternative nuclear
decay process. In such cases, it is of interest to measure ratios of K-capture to positron-emission probability or of the total electron-capture to positron-emission rate,

\[
\frac{P_K}{P_{\beta^+}} = \frac{\lambda_K}{\lambda_{\beta^+}}, \quad \frac{P_{EC}}{P_{\beta^+}} = \frac{\lambda_c}{\lambda_{\beta^+}}.
\]

Table 3.1 contains a compilation of methods reported in the literature; these are discussed in Secs. 3.1, 3.2, and 3.4. The usual corrections for background, dead time, detector efficiency, etc., are taken for granted.

3.1. Determination of Capture Ratios

Capture ratios have been determined both with external and internal sources. In general, it is difficult to measure capture ratios with external sources, because large corrections are required for source self-absorption, air scattering, window absorption, and fluorescence yields. During the last few years, capture ratios have therefore more frequently been measured by internal-source techniques in which these difficulties are avoided, provided the radioactive atoms can be dispersed throughout the sensitive volume of the counter. Internal-source methods fall into two major classes: at low atomic numbers, gaseous compounds are mixed with the counting gas of a proportional counter, while at high Z crystal scintillators are preferred that have the radioactive atoms built into the lattice, thus minimizing distortions due to escape of x rays from the sensitive counter volume.
1.1. Spectrometry with Internal Gas Sources

A radioactive gas or the vapor of a radioactive metal-organic compound is added to the counting gas of a proportional counter. The prompt cascade of x rays and Auger electrons, which follows the capture event, is integrated by the detector to produce a single K peak at the K-electron binding energy of the daughter atom. Similarly, L, M, ... peaks are produced by events from higher shells. It is usually assumed that all L and M x rays and Auger electrons are completely absorbed inside the counter. However, as Vatai (1968a, 1970b) has pointed out, the escape of L x rays is not always negligible a priori, and becomes especially important if the L x-ray energy lies just below the K-shell binding energy of the counter gas. The L peak contains a contribution from K-capture events which arises from K x rays that escape from the sensitive volume of the counter.

Typical K, L, and M peaks from an internal $^{71}$Ge source are shown in Fig. 3-1. From the measured intensities $I_K$, $I_L$, $I_M$ of these peaks, the ratio of capture probabilities can be deduced:

$$\frac{P_L}{P_K} = \frac{I_L}{I_K} [1 - \omega_\alpha (1 + K_\alpha K_\alpha + \frac{P_{K_\alpha K_\alpha} P_{K_\beta}}{P_{K_\alpha K_\alpha} P_{K_\beta}})] - \omega_\alpha P_{K_\alpha K_\alpha}$$  \hspace{1cm} (3-9)

$$\frac{P_M}{P_L} = \frac{I_M}{I_L} \left[ 1 + \frac{P_{K_\alpha K_\alpha} P_{K_\beta}}{P_{K_\alpha K_\alpha} P_{K_\beta}} \right] - \frac{P_{K_\alpha K_\alpha} P_{K_\beta}}{P_{K_\alpha K_\alpha} P_{K_\beta}}$$

$$- \omega_\alpha \left\{ 1 + \frac{P_{K_\alpha K_\alpha} P_{K_\beta}}{P_{K_\alpha K_\alpha} P_{K_\beta}} \right\}.$$  \hspace{1cm} (3-10)
Here, \( \omega_K \) and \( \omega_L \) are the K- and L-shell fluorescence yields of the daughter atom; and \( k_{\alpha} \), \( k_{\beta} \), and \( k_{\gamma} \) the fractions of \( K\alpha \), \( K\beta \), and \( L\alpha \) x rays in the K and L series. The K and L x-ray escape probabilities from the detector sensitive volume are denoted by \( P_K \), \( P_{K\beta} \), and \( P_{L\alpha} \).

There are two limiting cases. The first of these is Method 1 of Table 3.1, in which escape of x rays from the counter volume is avoided. Then Eqs. (3-9) and (3-10) have the simple form

\[
\frac{P_L}{P_K} = \frac{I_L}{I_K}, \quad \frac{P_M}{P_L} = \frac{I_M}{I_L}.
\] (3-11)

Absence of x-ray escape can be realized approximately when the counter is operated at high pressure. Gas fillings of argon-propane and argon-methane mixtures at up to 22 atm have been used. Since the development of the wall-less multiwire proportional counter (Drever et al., 1957a, 1957b), this type of detector has been employed successfully by various groups. The principal advantage of such a multiwire counter is that escape can be made very small. A central counter is surrounded by a ring of additional counters (Fig. 3-2). An inner circle of wires serves as the cathode for the central counter. Alternate wires in an outer circle serve as anodes and cathodes of a set of ring counters. The sensitive volume of the detector is then separated into two parts. The main central counter and the ring counters are operated in anticoincidence.

A block diagram of electronics for the operation of a multiwire proportional counter is shown in Fig. 3-3. Negative high voltage is often applied to the outer case of the counter and to the field tubes.
This approach is superior to grounding the cathode and using positive high voltage on the center wire, with a large potential difference across the coupling capacitor between center wire and the first preamplifier stage, leading to problems of leakage and spurious discharge.

For the determination of L/K ratios at Z<20 and M/L ratios at Z<40 it is necessary to detect Auger electrons and soft x rays below 500 eV, down to a few eV. Most recent advances in low-energy proportional-counter technique are related to the electronic system (Dougan et al., 1962a; Renier et al., 1968; Genz et al., 1971a). Proportional-counter spectrometry of radiation below ~500 eV is affected by certain problems that are less important or negligible at higher energies: (1) After-pulses from primary ionizing events can occur (Dougan et al., 1962a; Renier et al., 1968; Genz et al., 1971a; Campion, 1968, 1973); (2) degradation tails from peaks of higher energy can appear (Renier et al., 1968; Genz et al., 1971a; Heuer, 1966; Vaninbroukx and Spernol, 1965; Spernol, 1967); (3) small pulses can be mutually induced between ring and center counters in multiwire detectors (Genz et al., 1971a; Drever et al., 1957); (4) the anticoincidence gate may cause front- and back-edge clipping of large pulses, producing smaller pulses (Dougan et al., 1962a; Renier et al., 1968; Genz et al., 1971a); (5) large deadtime may arise when radiation of higher energy is present in high intensity (Dougan et al., 1962a; Renier et al., 1968; Genz et al., 1971a). The electronic system shown in Fig. 3-3 is designed to overcome these problems, except for long deadtime and degradation tails.

The shape of the spectrum produced by events between a few and 500 eV
depends on the initial number of ion pairs. The energy required to produce an ion pair in an argon-propane mixture is \textasciitilde 27 eV. Peaks produced by several ion pairs can be satisfactorily fitted with a Poisson distribution (Campbell and Ledingham, 1966), while the spectrum due to single-electron events cannot be represented accurately by an exponential or quasi-exponential function, as it varies with gas multiplication (Gold and Bennet, 1966; Genz, 1968, 1973b).

Corrections for several effects must be applied. (1) Escape probabilities \( P_{K\alpha} \) and \( P_{K\beta} \) of \( K\alpha \) and \( K\beta \) x rays from the sensitive volume of the counter must be accounted for. These escape probabilities can be separated into the additive probabilities \( P_1 \), that a x ray escapes from the central counter through the ends, \( P_2 \), that a x ray escapes from the central counter and hits a cathode wire, and \( P_3 \), that an x ray escapes from the central counter and passes through a ring counter without being detected. All these corrections can be kept below 1%. A careful study of the escape probability in multiwire counters has been made by Vatai (1970b). (2) An important correction must be made for degraded L and K events in the energy region below the peaks. The total contribution from such events can be determined by extrapolation parallel to the energy axis to low energy, as has been demonstrated down to 80 eV (Genz et al., 1971a). The degradation correction can amount to several percent but has not been taken into account in many investigations. This leads to appreciable differences in results (Heuer, 1966; Totzek and Hoffmann, 1967; Genz et al., 1971a; Pengra et al., 1972). (3) Condensation of radioactive metal-organic
vapor on the counter wall can lead to an increase in background. (4) Values of the fluorescence yield \( \omega_K \) and of the Ka/K\( \beta' \) x-ray intensity ratios can usually be taken from literature. The largest source of error in this method arises from the uncertainty in the \( I_L/I_K \) or \( I_L/I_K \) intensity ratio. In the determination of M/L capture ratios, errors in \( P_{L}/P_K \) largely cancel [see Eq. (3-10)]. Uncertainties in \( k_\alpha \) and \( k_\beta \) have been greatly reduced since the new calculations of Scofield (1974) became available, which agree very well with experiment (Scofield, 1975).

If transitions take place to several levels in the daughter nucleus, then only mean capture ratios are measured. Several of the most reliable mean ratios have been measured by internal gas-source spectrometry. In the use of nuclides that decay by electron capture to a level that is deexcited by a \( \gamma \) transition, coincidences can be measured between K and L events detected in a multiwire counter and \( \gamma \) rays detected with NaI(Tl) scintillators surrounding the proportional counter. The capture ratio for transitions to the excited state can be deduced from the measured intensities \( I_{L-\gamma} \) and \( I_{K-\gamma} \) of L and K events gated by the \( \gamma \) rays. Equation (3-9) applies, with \( I_L \) and \( I_K \) replaced by \( I_{L-\gamma} \) and \( I_{K-\gamma} \). An analogous procedure can be employed in M/L-ratio measurements. In addition to the corrections already mentioned, accidental and sum coincidences must be taken into account.

In the second limiting case of internal gas spectrometry (Method 2 of Table 3.1), all K x rays are allowed to escape from the sensitive volume of the counter. Then we have \( P_{K\alpha} = 1 \) and \( P_{K\beta} = 1 \), Eq. (3-9) yields

\[
\frac{P_L}{P_K} = \frac{I_L}{I_K} \left( 1 - \omega_K \right) - \omega_K \alpha' \tag{3-12}
\]
and Eq. (3-10) becomes

\[
\frac{P_M}{P_L} = \frac{T_M}{T_L} \left[ 1 + \frac{P_K}{P_L} \left( 1 + K \right) \right] - \frac{P_K}{P_L} \left( 1 + K \right).
\] (3-13)

Here, L x-ray escape is considered negligible. Experimentally, total K x-ray escape has been approximated with single-wire proportional counters filled with a low-Z gas at low pressure (Pontecorvo et al., 1949; Langevin, 1954, 1955, 1956; Langevin and Radvanyi, 1954, 1955; Radvanyi, 1955a; Scobie, 1957a; Kiser and Johnston, 1959). Corrections are needed to account for (1) non-escape of K x rays, (2) escape of L x rays, (3) wall and end effects, (4) the fluorescence yield \( \omega_K \), and (5) the fraction \( K_\alpha \) of Ka x rays in the total x-ray group. Additional uncertainties may arise from separation of the K and L peaks and from their degradation tails.

With single-wire proportional counters containing a gaseous radioactive source mixed with the counter gas, reliable measurements are no longer limited to events with energies above \( \sim 200 \text{ eV} \). Recent advances in single-wire proportional-counter techniques (No. 3 in Table 3.1) have extended the sensitivity of precision measurements to make possible the detection of single- and few-electron events down to essentially zero energy, even in the presence of intense more energetic radiation (Fink, 1968; Genz, 1968, 1973a). These improvements were attained with more sophisticated low-noise electronics and through an understanding of the degradation spectrum (Genz et al., 1971a) and of after-pulses (Genz et al., 1968). Single- and few-electron peaks have been resolved on the basis of their spectral shape (Renier et al., 1968) or by fitting
a Poisson distribution (Genz et al., 1971a, 1972; Pengra et al., 1972).

The techniques of single-electron spectrometry have been applied by Renier et al. (1968) in a precision measurement of the M/L capture ratio of $^{37}\text{Ar}$. In this case, the peak due to capture of L-shell electrons has a mean energy of 280 eV, and the M spectrum is a single-electron peak because the energy released in a capture event ($\approx 5$ eV) is lower than that required to produce an ion pair ($\approx 26.5$ eV in argon-propane). The spectrum due to single electrons was determined experimentally by introducing ultraviolet photons into the counter to produce photoelectrons of only a few eV. This experimentally determined single-electron spectrum was fitted in the M region (Fig. 3-4) of the composite M and L spectrum (Fig. 3-5) and extrapolated to zero energy. The small afterpulses, which may follow a primary event in the counter gas were kept from entering the analyzer by introducing an electronic paralysis time of up to 3.8 ms following each primary pulse. A block diagram of the electronic circuit is shown in Fig. 3-6.

The principal errors in this method arise from fitting the single-electron spectrum to the M-peak shape and from establishing the zero-energy calibration of the analyzer. The spectrum must be corrected for background and degradation tails. The ratio $P_M/P_L$ is a very sensitive function of $k_a$, but it is rather insensitive to $w_K$ [Eq. (3-13)].

Internal gas spectrometry for the precision determination of electron-capture ratios is limited to sources with atomic numbers below ~50, because with heavier atoms too many x rays escape from the sensitive counter volume, even at high counting-gas pressures. Although
this escape probability can be calculated in principle, the accuracy of the measurements is severely affected.

In earlier days, some L/K capture ratios were determined by measuring trajectories produced in a cloud chamber by K and L events from a radioactive gas (Radvanyi, 1952a, 1955a). This approach is included in Table 3.1 for historical reasons as Method 4.

3.1.2. Spectrometry with Internal Solid Sources

The internal gas spectrometry technique fails at high Z because too many K x rays escape. To circumvent the problem, the proportional counter can be replaced by scintillation crystals if the radioactive atoms can be built into the crystal lattice (Der Mateosian, 1953). From the measured intensities of K, L, and M events the capture ratios can then be deduced. The advantage of the method (No. 5 in Table 3.1) is that self-absorption of the emitted radiation can be neglected.

It is required, however, that the scintillation behavior not be disturbed by addition of the source material. Clustering must be avoided.

The source crystal can be placed directly on the photocathode of the multiplier tube. Groups at Heidelberg have used NaI(Tl) and CsI(Na) crystals doped with appropriate isotopes for the determination of electron-capture ratios by the internal-source technique. Leutz et al. (1966) grew NaI(Tl) crystals containing $^{202}$Tl and $^{204}$Tl as a constituent of the crystal lattice, and Schulz (1967a) doped the scintillator with $^{83}$Rb and $^{185}$Os. Furthermore, $^{131}$Cs has been built into the lattice of CsI(Na) scintillation crystals. To use doped crystals for spectrometry it is necessary that the radioactive nuclei
be uniformly distributed in the scintillator. To avoid absorption
effects caused by possible surface concentration and precipitation of
activity at grain boundaries, Ravn and Bøgeholt (1971) used \( \text{Cs}_2 \text{Pt(CN)}_4 \cdot \text{H}_2\text{O} \) doped with \( \text{Pt} \) for the determination of the M/L capture ratio in
the decay of \( \text{Pt} \). This scintillator material has several advantages.
Platinum being one of the main constituents of the crystal, \( \text{Pt} \) is
for chemical reasons ensured a completely uniform distribution. The
crystal exhibits light yields and relaxation times comparable to those
of NaI(Tl).

Two principal sources of error must be overcome in this method.
The radioactive source must form a true solution; if the radioactivity
lodges non-uniformly at dislocations or grain boundaries, absorption
effects occur. Schulz (1967b) has investigated the problem and has
developed a chemical and a physical criterion to decide which radio-
active isotopes form true mixed crystals with NaI(Tl). She finds that:
Rb, Cs, Ba, Os, Tl, and Pb do form such mixed crystals, whereas P, Ca,
Mn, Zn, As, Y, Sn, Ce, and Bi do not. Joshi et al. (1963) have studied
the effects of non-uniformity of mixing and the phenomena of over-
activation and poisoning. The second main source of error arises from
K x-ray escape from regions near the surface, which results in the
recording of K-capture events as L- or M-capture events.

To correct for x-ray escape, basically two methods have been used.
A well-type NaI(Tl) hollow crystal can be employed to surround the
NaI(Tl) crystal that contains the internal radioactive source (Fig.
3-7). Escaping K x rays from electron capture and iodine K x rays
associated with the detection process are absorbed in the outer crystal and are recorded as simultaneous events, so that no x-ray escape corrections are required. The method has been used by Joshi and Lewis (1960), Joshi (1961), Smith and Lewis (1966), Goedbloed (1970a) and by Goedbloed et al. (1968, 1970b) who have discussed it in detail.

An alternative approach to correct for x-ray escape involves measurement of the ratios of the areas A, B, and C of the K, L, and M peaks for several source crystals of different sizes (Figs. 3-8 and 3-9). Leutz et al. (1966) have shown that correction for escape can be most accurately performed by plotting the ratios A/(A+B) and C/B against the surface-to-volume ratio of the doped crystal and extrapolating linearly to a surface-to-volume ratio of zero. Thus, values of \( P_K/(P_K + P_L) \) and \( P_M/P_L \) are found that correspond to a measurement with an infinitely large crystal.

Corrections must be applied for (1) sum effects, (2) self-absorption, if clustering occurs, (3) possible influence of internal conversion or \( \beta^- \) background (as in the case of \( ^{204}_{\text{Tl}} \)). X-ray escape is accounted for if one of the above-described techniques is used.

The method of internal solid source spectrometry can be made very accurate.

A reduction in the noise level was attained by Ravn and Bøgelund (1971) by means of a coincidence system in which two low-noise photomultiplier tubes were coupled to a \( ^{193}_{\text{Pt}} \)-doped crystal. Crystal and photomultiplier assembly were cooled to \(-35^\circ\text{C}\) to reduce dark current (Fig. 3-10).
In the case of nuclides that undergo electron capture to an excited state, internal solid-source spectrometry with coincident γ rays is possible. The intensities of L and K events are measured in the source crystal in coincidence with ensuing γ rays (Fig. 3-11). Accidental coincidences must be taken into account. In favorable cases this method can be made quite accurate.

3.1.3. Spectrometry of K and L x Rays with External Sources

This method (No. 6 in Table 3.1) is based on the determination of the intensities $I_{LX}$ of L x rays and $I_{KX}$ of K x rays from singles spectra as measured with proportional counters or NaI(Tl) detectors. The sources, placed outside the sensitive volume of the detector, are usually prepared by drop deposition, but metal grains (Johns et al., 1957), sources prepared by painting (Fujiwara et al., 1964), and vacuum-evaporated sources (Venugopala Rao and Crasemann, 1965) have been used. The L/K ratio is deduced from the relation

$$\frac{P_L}{P_K} = \frac{I_{LX} \omega_K}{I_{KX} \omega_{LL}} - \frac{\omega_{LK}}{\omega_{LL}} \eta_{KL},$$

where $\omega_K$ is the K-shell fluorescence yield, $\omega_{LL}$ is the partial L-shell fluorescence yield following L capture, $\omega_{LK}$ is the partial L fluorescence yield following Kα x-ray emission, and $\eta_{KL}$ is the number of L-shell vacancies produced on the average when a K-shell vacancy is filled.

Corrections must be made to account for (1) self-absorption, (2) absorption between source and detector, (3) solid angle, if different
detectors are used for the measurement of L and K x rays, (4) efficiency of the detectors, (5) interfering effects due to γ rays and internal-conversion electrons. There is some uncertainty in $n_{KL}$ and in the fluorescence yields $ω_K$, $ω_{LK}$, and $ω_{LL}$, which can usually be found in the literature (Bambynek et al., 1972). An additional uncertainty can arise from degraded L x rays at the low-energy side of the L peak.

Capture ratios can be determined by this method in the case of nuclides that decay from ground state to ground state or to an excited metastable state. For nuclei that decay by a prominent transition, among others, to the ground state of the daughter, mean L/K ratios can be obtained. Though often used, the method is not very accurate, because $P_L/P_K$ is a small difference between two large quantities, and the partial L-shell fluorescence yields greatly affects the result.

Venugopala Rao and Crasemann (1965), and Venugopala Rao et al. (1966a) have measured the L and K x-ray intensities relative to the K x-ray intensity of a $^{109}$Cd reference source and thus deduced $P_L/P_K$ of $^{181}$W and $^{204}$Tl. Kramer et al. (1956) have determined $P_L/P_K$ of $^{202}$Tl by comparing the intensity ratio $I_{LX}/I_{KX}$ with that of a $^{203}$Hg reference source. In addition to the need for corrections indicated earlier, the quantities $n_{KL}$, $ω_{LL}$, $ω_{LK}$ and the internal conversion coefficients $a_L$ and $a_K$ of the reference source must be known. With an appropriately chosen reference nuclide these corrections can partly cancel.

For nuclides decaying to an excited state that is followed by γ-ray emission, coincidences can be determined between K x-rays and γ-rays and between L x rays and γ rays. From the measured coincidence
counting rates \( I_{KX-Y} \) and \( I_{LX-Y} \) and from the singles rate \( I_\gamma \), the L/K-capture ratio can be found:

\[
\frac{P_L}{P_K} = \left( \frac{I_{LX-Y}}{I_{KX-Y}} \frac{\omega_K}{\omega_L} \right) \frac{\lambda_{KL}}{\lambda_{LL}}.
\]

The L x rays have usually been measured with proportional counters, and the K x rays and \( \gamma \) rays, with NaI(Tl) detectors. This method is an extension of that based on Eq. (3-14). It requires the same principal corrections and suffers from the same uncertainties; accidental and sum coincidences must also be taken into account.

A special technique was employed by McCann and Smith (1968) in their work on \( ^{133}\)Ba. These authors used a NaI(Tl) detector to measure the L and K x-ray spectra gated by the sum coincidence peak of the 356-keV and 81-keV \( \gamma \) rays, which were absorbed in another NaI(Tl) detector.

Measurement of (L-event)-(K-x-ray) Coincidences. This method (No. 7 in Table 3.1) has been employed by Christmas (1964) to determine the L/K-capture ratio of \( ^{204}\)Tl. Coincidences between L x rays and K x rays were measured by means of two NaI(Tl) detectors, and \( \frac{P_L}{P_K} \) was deduced.

In a similar approach, Konstantinov and Perepelkin (1961) used a \( 4\pi \) proportional counter filled with a Xe-CH\(_4\) mixture. Coincidences between L events (L x rays and L Auger electrons) in the top part and K x rays in the bottom part of the counter were detected. A sufficiently thick backing material permitted only K x rays to penetrate to the bottom counter.

The method requires corrections for (1) self-absorption of L x rays.
and Auger electrons, (2) absorption of K x rays in the backing foil, (3) escape of K x rays from the detectors, (4) detector efficiencies, including solid angle, (5) accidental and sum coincidences, and (6) influence of possible γ rays. Values of n_\text{KI} and K-shell and L-shell fluorescence yields can usually be found in the literature (Bambynek et al., 1972); they contribute to the overall uncertainty. The method yields mean P_L/P_K values if the nuclide decays by more than one electron-capture branch.

3.2. Determination of the Relative K-Capture Probability P_K

In addition to the determination of capture ratios, there are various other methods from which the relative capture probability P_K can be deduced. Some of these constitute a direct measurement of P_K. In various others the product P_Kω_K is determined. All measurements described in this section employ external sources, placed outside the sensitive volume of the detector.

3.2.1. Measurement of K x Rays or Auger Electrons and γ Rays or Conversion Electrons

Spectrometry of K x rays and γ rays. The principle of this method (No. 8 in Table 3.1) is to measure the intensities I_{KX} of the emitted K x rays and I_γ of the γ rays and hence to deduce the K-capture probability:

\[
\frac{I_{KX}}{I_\gamma} = \alpha_K \omega_K \left[ 1 + P_K (1+\alpha) / \alpha_K \right].
\]  (3-16)
Here, \( \omega_K \) is the K-shell fluorescence yield, while \( \alpha_K \) and \( \alpha \) are the K-shell and total conversion coefficients. Sources have been prepared by simple drop deposition. Proportional counters as well as NaI(Tl) and Ge(Li) detectors have been used.

Principal corrections are required for (1) self-absorption of the K x rays, (2) absorption between source and sensitive volume of the detectors, (3) efficiencies of the detectors for K x rays and \( \gamma \) rays, and (4) solid angles. Values for the fluorescence yield \( \omega_K \) and the conversion coefficients are required. If internal conversion can be neglected, Eq. (3-16) becomes simply \( I_{Kx}/I_\gamma = P_K \omega_K \).

Bayer et al. (1972) used this method to measure the K x-ray intensities in the \( ^{140}\text{Nd} \rightarrow ^{140}\text{Pr} \rightarrow ^{140}\text{Ce} \) decay chain and to deduce \( P_K \) of \( ^{140}\text{Nd} \). Wapstra et al. (1954, 1957) and Friedlander and Orr (1951b) employed two nuclides that decay to the same excited level in the daughter nucleus, one by electron capture and the other by \( \beta^- \) emission. The intensity ratio of the K x rays and \( \gamma \) rays from the two nuclides was determined and hence the K-capture probability:

\[
(I_{Kx}/I_\gamma)^{EC} (I_{Kx}/I_\gamma)^{-1} = 1 + P_K (1 + \alpha)/\alpha_K.
\]  

(3-17)

Corrections are required mainly for (1) sum effects, and (2) contributions of radiation from higher levels. K-shell and total conversion coefficients are usually taken from the literature.

Spectrometry of K x rays or Auger electrons and K conversion electrons. The principle of this method (No. 9) is to measure the intensity \( I_{Kx} \) of K x rays and \( I_{eK} \) of K conversion electrons (Avignon
et al., 1955). The K-capture probability is found from the equation

\[ \frac{I_{KX}}{I_{eK}} = \omega_K \frac{1 + \rho_K (1 + \alpha)/\alpha_K}{1 + \rho_K (1 + \alpha)/\alpha_K}. \] (3-18)

Moussa and Juillard (1956) have measured the intensities \( I_{KA} \) of K Auger electrons and \( I_{eK} \) of K conversion electrons and used a relation similar to Eq. (3-18) with \( I_{KX} \) and \( \omega_K \) replaced by \( I_{KA} \) and \( (1 - \omega_K) \), respectively. Magnetic \( \beta^- \) spectrometers were used to detect the electrons and a NaI(Tl) scintillation counter for the x rays.

Corrections must be made for (1) self-absorption of the K x rays or Auger electrons, (2) absorption between source and detector; (3) efficiencies of the detectors including solid angles; and (4) radiation from higher levels, if present. Fluorescence yields and internal conversion coefficients are usually taken from the literature.

**Determination of K x-ray emission rate and disintegration rate.**

This method (No. 10) requires determination of the K x-ray emission rate \( I_{KX} \), preferably with a large proportional counter filled to a sufficient pressure to absorb all K x rays. In addition, the disintegration rate \( I_0 \) must be determined, preferably by means of a coincidence technique as used in absolute standardization of radioactive sources. The value \( \rho_K \omega_K \) is found from the relationship

\[ \rho_K \omega_K = \frac{I_{KX}}{I_0}, \] (3-19)

where \( \omega_K \) is the K-shell fluorescence yield.

The method is described in detail by Taylor and Merritt (1965). To check the K x-ray emission rate, a second fairly independent approach
can be used (Bambynek, 1967a) utilizing a medium-solid-angle arrangement with a proportional counter or a thin NaI(Tl) crystal as detector (Bambynek et al., 1966; Bambynek, 1967b). The detection system for determining the disintegration rate has been described by Campion (1959). It consists of a 4π flow-type pillbox proportional counter placed between two NaI(Tl) detectors. A calibrated γ spectrometer also (Vaninbroukx and Grosse, 1966) has been used to determine the disintegration rate.

Radioactive sources have been prepared for experiments of this type by drop deposition, electrodeposition, and evaporation in vacuum. Sources have been mounted on thin metallized plastic foils for the determination of the disintegration rates, then they were sandwiched between absorber foils to stop all Auger electrons, so that K x-ray emission rates could be measured in a high-pressure proportional counter.

The principal corrections that must be applied in the K x-ray measurements are for (1) self-absorption, (2) foil absorption, (3) x-ray counter efficiency (normally near unity), and (4) the effect of γ rays and β⁺ particles, if present. The corrections in the determination of the disintegration rate by the coincidence method are small and well-understood, and involve only parameters that can be determined experimentally as an integral part of the measurement. The fluorescence yield ωₖ is usually taken from the literature (Bambynek et al., 1972).

This method has been applied in laboratories specializing in the standardization of radionuclides, and has yielded several of the most reliable Pₖωₖ values.
3.2.2. Coincidence Measurements

With nuclides that decay by electron capture to an excited level in the daughter nucleus, coincidences can be measured between x rays or Auger electrons (from the capture process) and y rays or conversion electrons (from the deexcitation of the daughter state). Such measurements can serve to determine capture probabilities or their ratios.

**Measurement of K x-ray and γ-ray coincidences.** In this method (No. 11), coincidences are measured between K x rays in one detector and γ rays in another detector. One finds

\[ P_{Kx} = \frac{I_{Kx-\gamma}}{I_\gamma}, \tag{3-20} \]

where \( I_{Kx-\gamma} \) is the (K x-ray)-(γ-ray) coincidence counting rate, \( I_\gamma \) is the singles γ rate, and \( \omega_k \) is the K-shell fluorescence yield of the daughter atom. Sources for such experiments have mostly been prepared by drop evaporation; however, plated (Grotheer et al., 1969), electroplated (Thomas et al., 1963), gaseous external sources (Bresesti et al., 1964; Winter et al., 1965b) and metal powders (Ferrin, 1960; Millar et al., 1959) have also been used.

Different combinations of detectors have been employed; in most cases proportional counters served for the K x rays and NaI(Tl) detectors for the γ rays or for both radiations. Solid state detectors have also been used recently: NaI(Tl)-Ge(Li) (Raeside et al., 1969; Myslek et al., 1971); Ge(Li)-Ge(Li) (Schmidt-Ott and Pink, 1972), and Si(Li)-Ge(Li) (Genz et al., 1973c).

Corrections must be applied principally for the following effects:
(1) self-absorption and absorption of K x-rays between source and sensitive volume of the detector, (2) efficiency of the K x-ray detector, including solid angle, (3) detection of γ rays or conversion electrons in the x-ray detector, (4) contributions from positrons, if present, and (5) sum and accidental coincidences. Values of the fluorescence yield $\omega_K$ can usually be taken from the literature. In order to avoid uncertainties due to the insufficiently known fluorescence yields, De Wit and Wapstra (1965) in their measurements on $^{195}$Au and $^{197}$Hg compared the intensity ratios $I_{K\gamma\gamma}/I_{\gamma\gamma}$ with that of a $^{202}$Hg reference source. With an appropriately chosen reference nuclide, the fluorescence yields practically cancel. On the other hand, knowledge of $P_K$ of the reference nuclide is required.

With nuclides decaying to an excited level that is followed by a γ-γ cascade to the ground state, triple coincidences have been measured. The K-capture probability can then be found from the relation

$$P_{K\omega K} = \frac{I_{K\gamma_1\gamma_2}}{I_{\gamma_1\gamma_2}}, \quad (3-21)$$

where $I_{K\gamma_1\gamma_2}$ is the rate of the (K x-ray)-(γ₁)-(γ₂) triple coincidences, and $I_{\gamma_1\gamma_2}$ is the (γ₁)-(γ₂) coincidence rate. In addition to the corrections mentioned previously, directional correlations must be taken into account.

The coincidence method permits determination of the K-capture probability for transitions to an excited level in the daughter nucleus. By appropriate choice of γ-ray window settings one can select a particular electron-capture transition among several in the same decay.
This technique (No. 12) has been employed to determine the ratio of K-capture probabilities to different levels (denoted here by 1 and 2):

\[ \frac{P_{K1}}{P_{K2}} = \left( \frac{I_{K\alpha - \gamma 1}}{I_{\gamma 1}} \right) / \left( \frac{I_{K\alpha - \gamma 2}}{I_{\gamma 2}} \right). \] (3-22)

The result does not depend upon the fluorescence yield and the efficiency of the K x-ray detector. In most cases, NaI(Tl) detectors have been used for the K x rays and γ rays (Lewin et al., 1965), but NaI(Tl)-Ge(Li) (Schmidt-Ott et al., 1968; Schmidt-Ott, 1970; Cook and Johns, 1969; Lourens et al., 1970) and Si(Li)-Ge(Li) combinations (Lourens, et al., 1970) have also been employed. The method has been used mostly to determine the energies of electron capture transitions.

**Measurement of (K x-ray and Auger-electron)-(γ-ray) coincidences.**

If coincidences between K x rays or Auger electrons and γ rays are measured (Method 13), the K-capture probability \( P_K \) can be directly deduced:

\[ P_K = \frac{I_{K-\gamma}}{I_{\gamma}}. \] (3-23)

Very thin (e.g., vacuum-evaporated) sources of large area are required to keep self-absorption down. Kramer et al. (1962a) employed this method with a double proportional counter operated at sufficiently high pressure to detect all K x rays and Auger electrons. The source was placed so as to attain a ≈4π solid angle. Gamma rays were detected with a NaI(Tl) scintillation counter. Vatai and Hohmuth (1968) employed a 4π CsI(Tl) detector system to register K events and a CsI(Tl) detector for the γ rays.
Corrections are required for (1) self-absorption of K x rays and Auger electrons, (2) absorption of x rays and electrons in the backing foil of the source, (3) incomplete realization of the 4π solid angle, (4) accidental coincidences, (5) detection of γ rays in the K-event detector, and (6) influence of positrons, if present.

Measurement of (K x-ray)-(γ-ray) sum coincidences. In this method (No. 14), which was first used by Gupta and Iha (1956), the pulse-height spectrum of K x rays and γ rays is measured in one single detector. The spectrum (Fig. 3-12) contains a K x-ray peak, a γ-ray peak, and a sum peak arising from (K x-ray)-(γ-ray) coincidences in the detector. From the measured areas $A_\gamma$ and $A_{XY}$ of these peaks, the capture probability can be deduced:

$$P_{K\gamma} = \frac{I_{KX-\gamma}}{I_\gamma} = \frac{A_{XY}}{A_\gamma + A_{XY}}.$$  (3-24)

In most cases, a NaI(Tl) detector has been employed for measurements of this type. Das Mahapatra and Mukherjee (1974) used a Ge(Li) detector, and Campbell and McNelles (1972) employed a sandwich detector consisting of two CsI(Tl) crystals with the source in between.

Corrections must be made for (1) self-absorption and absorption of K x rays between source and sensitive volume of the detector, (2) efficiency of the K x-ray counter, including solid angle, (3) accidental coincidences, and (4) separation of overlapping parts of the γ-ray and sum peaks.

Gupta (1958) has used this method with triple sum coincidences. He observed the pulse-height spectrum in a single NaI(Tl) detector and
determined the areas $A_{12}$ and $A_{12}$ of $Kx-\gamma_1-\gamma_2$ and $\gamma_1-\gamma_2$ sum coincidence peaks. The K-capture probability is

$$P_{K} = \frac{I_{Kx+\gamma_1+\gamma_2}}{I_{\gamma_1+\gamma_2}} = \frac{A_{12}}{A_{12}^{+}}.$$  \hspace{1cm} (3-25)

Instead of employing a single detector, it is possible to measure coincidences between $Kx$ rays in one NaI(Tl) detector and sum coincidences of $\gamma_1$ and $\gamma_2$ in a second NaI(Tl) detector. The $Kx$ rays are then gated by the $\gamma_1+\gamma_2$ sum coincidences. The ratio of the corresponding intensities is equal to $P_{K}$. In a few cases, in which K capture is forbidden due to energetics, the I-capture fraction can be measured directly (Supstra et al., 1972; see Reer et al., 1974; Penzias, 1976).

Measurement of (K x-ray)-(y-ray) and (K x-ray)-(K x-ray) or (K x-ray)-(K conversion electron) coincidences. This method (No. 15) can be applied to nuclides that decay to an excited level in the daughter nucleus that is deexcited by a converted $\gamma$ transition. The approach was developed by Prueett and Wilkinson (1954); it is based on measuring coincidences between $Kx$ rays from the electron-capture process and $\gamma$ rays from the daughter nucleus, and additionally, coincidences between $Kx$ rays from the electron capture process and $Kx$ rays from internal conversion. The K-capture probability can be deduced from the relation

$$2 \left( \frac{I_{Kx-\gamma}}{I_{\gamma}} \right) / \left( \frac{I_{Kx-Kx}}{I_{Kx}} \right) = 1 + P_{K} (1+\alpha) / \alpha_{K},$$  \hspace{1cm} (3-26)

where $I_{Kx-\gamma}$ and $I_{Kx-Kx}$ are the coincidence counting rates, and $I_{\gamma}$ and $I_{Kx}$ the corresponding singles rates. Drop-deposited sources and NaI(Tl) detectors were used in these experiments. Results are independent of the fluorescence yields, but the K-shell and total con-
Hansen (1975) has determined $P_K$ of $^{139}$Ce by measuring coincidences between K x rays and γ rays and K x rays and K conversion electrons. The photons were measured by Si(Li) and NaI(Tl) detectors, and the electrons, by means of a magnetic β spectrometer. $P_K$ can be deduced from the relation

$$\frac{1+P_K}{P_K} = \left( \frac{I_{Kx-eK}}{I_{eK}} \right) / \left( \frac{I_{Kx-γ}}{I_γ} \right).$$

(3-27)

In addition to the usual corrections, sum and accidental coincidences must be considered. Fluorescence yield and conversion coefficients need not be known. The method is only applicable to nuclides with a simple decay scheme lacking a γ cascade in the daughter.

Measurements of coincidences between K x rays or Auger electrons and conversion electrons. Coincidence measurements of this type (Method 16) for the determination of $P_K$ were first made by Brosi et al. (1959), who observed the K x-ray spectrum gated by K- and L-conversion electrons (Fig. 3-13) and determined coincidence and singles intensities. The K-capture probability can be deduced from the relation

$$\frac{1+P_K}{P_K} = \left( \frac{I_{Kx-eK}}{I_{eK}} \right) / \left( \frac{I_{Kx-eL}}{I_{eL}} \right),$$

(3-28)

where $I_{Kx-eK}$ and $I_{Kx-eL}$ are the (K x-ray)-(K-conversion electron) and (K x-ray)-(L-conversion electron) intensities, respectively, and $I_{eK}$ and $I_{eL}$ are the corresponding singles rates. The K x rays have been measured with NaI(Tl) detectors, and the conversion electrons, with magnetic β spectrometers. Knowledge of the K-shell fluorescence yield
and the x-ray and electron detector efficiencies is not required.
Corrections must be made to account for (1) accidental coincidences,
(2) sum effects due to K x rays from electron capture and internal con-
version, (3) possible effects of other converted γ transitions in
cascade, and (4) possible effects of electron capture to higher levels.

Instead of utilizing coincidences between x rays and conversion
electrons, it is possible to determine $P_K$ from coincidences between K
Auger electrons and K or L conversion electrons. From the measured
intensities, $P_K$ is found:

\[
\frac{1 + P_K}{P_K} = \left( \frac{I_{KA-eK}}{I_{eK}} \right) / \left( \frac{I_{KA-eL}}{I_{eL}} \right) \tag{3-29}
\]

Here, $I_{KA-eK}$ and $I_{KA-eL}$ are the coincidence rates between K Auger
electrons and K and L conversion electrons, respectively. This method
(No. 17) has been used by Marelius et al. (1967), who employed two
magnetic spectrometers. The necessary corrections are essentially the
same as those in Method 16.

A slight variation of this approach has been used by Sparrman et al.
(1966), who measured the K Auger-electron spectrum in coincidence with
K and L conversion electrons by means of two long-lens spectrometers.
The value for $P_K$ was found from

\[
\frac{1 + P_K}{P_K} = \left( \frac{I_{KA-eK}}{I_{KA-eL}} \right) \frac{\alpha_L}{\alpha_K} \tag{3-30}
\]

The K and L conversion coefficients must be known. In addition to the
corrections mentioned above, efficiencies for detecting K and L con-
version electrons and the absorption of these electrons between source
and detector must be taken into account.

Ploch et al. (1971) measured the K x-ray spectrum in a Ge(Li) detector gated by K conversion electrons which were detected in a proportional counter. By this method (No. 18), they determined $P_K$ from the ratio of the (K x-ray)-(K-conversion electron) coincidence rate $I_{KX-eK}$ and the K conversion-electron singles intensity $I_{eK}$:

$$\omega_K = \frac{I_{KX-eK}}{I_{eK}}.$$  \hspace{1cm} (3-31)

Corrections are needed for (1) accidental and sum coincidences, (2) self-absorption and absorption of K x rays between source and detector, and (3) efficiency of the K x-ray detector.

With nuclides decaying to a metastable level of the daughter, Durosinmi-Etti et al. (1966) have measured K x rays by means of a NaI(Tl) detector in coincidence with K conversion electrons detected with a surface barrier detector. The K-capture probability was deduced from the equation

$$P_K = \frac{I_{KX}}{I_{KX-eK}} \frac{1}{1+\alpha} - \frac{\alphaK}{1+\alpha}.$$  \hspace{1cm} (3-32)

Here, $I_{KX}$, $I_{eK}$, and $I_\gamma$ are the measured intensities of K x rays, K conversion electrons, and $\gamma$ rays, respectively; $I_{KX-eK}$ is the (K x-ray)-(K-conversion electron) coincidence rate, $\alpha_K$ is the K conversion coefficient, and $\alpha$, the total conversion coefficient. These conversion coefficients must be known. Corrections are needed for (1) X and $\gamma$ detector efficiencies, including solid angle; (2) absorption between source and detectors, and (3) overlap of spectrum peaks.
Measurement of triple coincidences between K x rays, \( \gamma \) rays, and internal-conversion electrons. This method (No. 19) was used by Thun et al. (1966), who determined the triple coincidence rate \( I_{KX-\gamma-eL} \), measuring K x rays with a NaI(Tl) crystal, \( \gamma \) rays with a Ge(Li) detector, and L conversion electrons with a magnetic spectrometer; the coincidence rate \( I_{\gamma-eL} \) between \( \gamma \) rays and L conversion electrons was simultaneously determined. Then we have

\[
P_{K}^{K} = \frac{I_{KX-\gamma-eL}}{I_{\gamma-eL}} \tag{3-33}
\]

A different approach was taken by Törnkvist and Ström (1968) in their measurements on \(^{133}\)Ba decay. These workers determined \( P_{K} \) directly from triple coincidences between K x rays, \( \gamma \) rays, and K or L conversion electrons detected with a lens spectrometer. The K-capture probability was deduced from

\[
\frac{1+P_{K}}{P_{K}} = \left( \frac{I_{KX-\gamma-eK}}{I_{\gamma-eK}} \right) \left( \frac{I_{KX-\gamma-eL}}{I_{\gamma-eL}} \right) \tag{3-34}
\]

Sources were prepared by evaporation in vacuum. Corrections must account for (1) accidental and sum coincidences, (2) directional correlations (which can be minimized by proper choice of the angle between detectors), and (3) escape of iodine K x rays from the NaI(Tl) detector.

3.3. Experimental Capture Probabilities \( P_{K}, P_{L}, \text{ and } P_{M} \)

Comparison with Theory

3.3.1. Experimental Results

All experimentally determined, published values of \( P_{L}/P_{K}, P_{M}/P_{L}, \)

\( P_{LM}/P_{K}, P_{K}^{K}, \text{ and } P_{K} \) are listed in Table 3.2. In the many cases in
which authors quote $P_K$ while they actually have measured $P_K \omega_K$, we list the latter product, recalculated from the authors' $P_K$ and $\omega_K$. In some cases, authors do not specify the value of $\omega_K$ which they used; these are indicated by "+." Some entries in Table 3.2 have been revised from the original publication. For example, the $P_{LM}/P_K$ ratio for $^{109}$Cd (Moler and Fink, 1965) was revised by the authors, who communicated this to Durosimmi-Etti (1966). Vatai (1968b, 1970b) has noted that the $^{109}$Cd $P_M/P_L$ value of Moler and Fink (1965) was not corrected for escape of Ag L x rays. Applying a corresponding correction and making use of newly reported values for $k_\alpha$ and $k_\beta$ (Salem et al., 1974) and $\omega_K$ (Bambynek et al., 1972) and a theoretical $P_L/P_K$ ratio yields $P_M/P_L = 0.205 \pm 0.020$. Similar corrections have been made to the $^{113}$Sn $P_M/P_L$ ratio of Manduchi et al. (1964b).

From among the entries in Table 3.2, we have selected those results that can with certainty be judged as reliable, because they were derived from measurements with pure, carefully prepared sources, all necessary corrections being determined and clearly described. (The importance of pure sources has been emphasized, for example, by Raman et al. (1973), who suggest that discrepancies in measured $P_L/P_K$ ratios of $^{113}$Sn may be due to variable amounts of $^{250}$d $^{119}$Sn present in the $^{115}$d $^{113}$Sn.) We have omitted results published without indication of error limits, or with errors in excess of 15%. The information provided in most publications is unfortunately less than complete. It is therefore probable that we have omitted some "good" results from the list of selected values. The selected $P_L/P_K$ measurements are
listed in Table 3.3, the $P_M/P_L$ ratios in Table 3.4, and selected values of $P_{LM}/P_K$, $P_{KL}$, and $P_{L}$ in Table 3.5. The K-shell fluorescence yields used to deduce the capture ratios $P_K$ in Table 3.5 were calculated from the equation

$$\left(\frac{\omega_K}{1-\omega_K}\right)^{1/4} = A + BZ + CZ^3.$$  \hspace{1cm} (3-35)

The constants $A$, $B$, $C$ were determined by fitting this expression to the selected "most reliable" experimental fluorescence yields listed by Bambeynek et al. (1972), with exception of those deduced from $P_K$ measurements. The fluorescence yields calculated in this manner are practically the same as those recommended by Bambeynek et al. (1972); slight changes in the last digit are within the stated error limits.

We use the transition energies $Q_{EC}$ evaluated by Wapstra and Gove (1971), except in cases where these were deduced from measurements of electron capture ratios. In those cases, we have used $Q_{EC}$ determined from measurements of internal-bremsstrahlung spectra or $(p,n)$ reaction thresholds. For a few transitions, no independent $Q_{EC}$ energies were available; these are indicated by an asterisk in Tables 3.3 and 3.5.

### 3.3.2. Theoretical Predictions

The last three columns of Tables 3.3 and 3.5 contain theoretical $L/K$ and $M/L$ ratios. These were calculated (see Sec. 2) from the relations

$$\frac{P_L/P_K}{(q_{L1}/q_K)^2} = \left(\frac{q_{L1}}{q_K}\right)^2 \left[1 + \frac{\epsilon_{L2}}{q_{T1}}\right] X_{L/K}$$  \hspace{1cm} (3-36)

and
for allowed transitions, and

$$\frac{\mathcal{P}_M/\mathcal{P}_L}{(q_M/q_L)^2} = \left(\frac{g_{M_1}}{g_{L_1}}\right)^2 \left[1 + \left(\frac{f_{M_2}/q_{M_1}}{1 + (f_{L_2}/q_{L_1})^2}\right)^2\right] L^M/L^L \quad (3-37)$$

for unique first-forbidden transitions. The electron radial wave

function amplitudes $g_K, g_{L_1}, f_{L_2}, g_{M_2}, f_{M_2}$, as well as $p_{L_3} g_{L_3}$ were taken from the relativistic Hartree-Fock calculations of Mann and Waber (1973) as listed in Table 2.9. The exchange and overlap correction factors $X_{L_1} = \frac{B_{L_1}}{A_{L_1}}$ and $X_{M_1} = \frac{B_{M_1}}{A_{M_1}}$ were recalculated in the present work according to the ansatz of Bahcall (1963a, c, 1965a) and that of Vatai (1968b, 1970a) as described in Sec. 2.5. For $Z > 32$, the correction factors of Suslov (1970) are used in continuation of the Bahcall factors, and those of Martin and Blichert-Toft (1970) in extension of the recalculated Vatai factors. Assumptions and approximations underlying the calculation of these correction factors are discussed in Sec. 2.5. Equations (3-36)-(3-38) contain the simplifications

$$\left(\frac{q_{L_2}/q_{L_1}}{q_{L_3}/q_{L_1}}\right)^2 = \left(\frac{q_{M_2}/q_{M_1}}{q_{M_3}/q_{M_1}}\right)^2 = 1 \quad (3-39)$$

and

$$\frac{L_2/L_1}{X} = \frac{L_3/L_1}{X} = \frac{M_2/M_1}{X} = 1. \quad (3-40)$$

These approximations affect the capture ratios by less than 0.04% for $Z = 20$ and less than 0.3% for $Z = 75$. 
The theoretical K-capture probabilities $P_K$ listed in the last column of Table 3.5 were calculated from theoretical capture ratios $P_L/P_K$, $P_M/P_L$, $P_N/P_M$ for $Z > 37$ and also $P_O/P_N$ for $Z > 67$, according to Eq. (3-6). Exchange and overlap corrections $X_{L/K}$ and $X_{M/L}$ were applied as discussed above, using our recalculated factors for $Z \leq 32$ and those of Suslov or Martin and Blichert-Toft for heavier atoms. For the outer shells no exchange correction was made, none being available.

The theoretical capture ratios and probabilities listed in Tables 3.3-3.5 for first-forbidden non-unique transitions are calculated for allowed transitions. This approximation is justified because for such transitions the ratios of capture probabilities from the $ns_{1/2}$ and $np_{1/2}$ subshells are independent of the form-factor coefficients (Sec. 2.3.2).

### 3.3.3. Comparison of Experimental and Theoretical Electron Capture Ratios

For comparison with theory, the selected experimental $L/K$ and $M/L$ ratios for allowed and non-unique first-forbidden transitions (Tables 3.3 and 3.4) were divided by the energy-dependent factors

$$\left(\frac{q_{L_{1/2}}}{q_{K}}\right)^2 = \left(\frac{E_{EC} - E_{L_{1/2}}}{E_{EC} - E_{K}}\right)^2$$ (3-41)

and

$$\left(\frac{q_{M_{1/2}}}{q_{L_{1/2}}}\right)^2 = \left(\frac{E_{EC} - E_{M_{1/2}}}{E_{EC} - E_{L_{1/2}}}\right)^2$$ (3-42)

respectively, where the capture transition energy is

$$E_{EC} = \frac{Q_{EC} - E}{\gamma}$$ (3-43)
and $E_K$, $E_{L_1}$, and $E_{M_1}$ are electron binding energies taken from Bearden and Burr (1967). In the case of measurements pertaining to transitions to several levels, we divided the measured mean L/K capture ratios by the factor

$$\left(\frac{q_{L_1}}{q_K}\right)^2 = \sum_v a_v \left(\frac{q_{L_1}}{q_K}\right)_v^2 .$$

(3-44)

The index $v$ labels the final-state levels; the $a_v$ are branching ratios subject to $\sum a_v = 1$. A corresponding procedure was used for mean M/L ratios. The branching ratios were taken from the Nuclear Data Sheets edited by the Nuclear Data Group, Oak Ridge National Laboratory.

The reduced experimental capture ratios $(P_L/P_K)/(q_{L_1}/q_K)^2$ and $(P_M/P_L)/(q_{M_1}/q_{L_1})^2$ are compared with theoretical ratios (Tables 3.3 and 3.4) in Figs. 3-12 and 3-13. For clarity, we have combined results for each atomic number and plotted weighted mean values and their uncertainties.

$P_L/P_K$ capture ratios. Figure 3-12 shows that agreement between experimentally determined L/K capture ratios and exchange-corrected theoretical predictions is fairly good for all atomic numbers, both for allowed and for non-unique first-forbidden transitions. The difference between theoretical ratios, due to different exchange and overlap corrections, is largest for light atoms (Sec. 2.5).

In cases in which the electron-capture transition energy is not much larger than the K-shell binding energy, the $(q_{L_1}/q_K)^2$ ratio is very sensitive to $Q_{EC}$. Errors in $Q_{EC}$ can then lead to erroneous conclusions in the comparison with theory. Such is the case for $^{206}$Bi, and probably also for $^{109}$Cd, $^{133}$Ba, $^{159}$Dy, $^{195}$Au, and $^{202}$Tl. More accurate
measurements on $^{79}$Kr and $^{159}$Dy should be performed. For $^{126}$I a mean L/K ratio has been measured, due to 60% non-unique and 40% unique first-forbidden transitions. The experimental result agrees well with predictions for either type of transition. The few available measurements pertaining to pure unique first-forbidden transition also agree well with theory. Table 3.3 includes the 4 measured L/K ratios for non-unique second-forbidden transitions, but these are not compared with theoretical ratios.

Vatai (1973a, 1974) has suggested that the ratio of non-relativistic to relativistic nuclear matrix elements could be estimated from $L_3/K$ ratios, and attempted to do this by evaluating the $L_3/K$ fraction of the measured L/K ratios of $^{93}$Mo (Hohmuth et al., 1964) and $^{97}$Tc (Kattoff, 1958), and the $LM.../K$ ratio of $^{139}$La (Turchinetz and Pringle, 1956). The fact that the $(L_1+L_2)/K$ ratio is independent of nuclear matrix elements made the separation possible. The experimental ratios unfortunately are not very accurate; improved measurements on these cases and on additional second and higher forbidden non-unique transitions would be useful. Vatai (1973a, 1974) has further pointed out that in the presence of K capture determinations of M/K ratios would be more useful than of M/L ratios, because the former are more sensitive to nuclear matrix elements. Chew et al. (1974a) have followed Vatai's suggestion and calculated the ratio of nuclear matrix elements

$$ R = \left( \frac{V_{220}^0}{V_{221}^0} - \sqrt{3/2} \cdot \frac{A_{221}^0}{V_{221}^0} \right) \frac{V_{221}^0}{V_{221}^0} \text{ in the decay of } ^{59}$Mn \text{ from } L_3/K \text{ deduced from the total measured L/K ratio. Daniel (1969) has noticed that for allowed transitions the reduced capture ratios } \frac{P_L/P_K}{(q_L/q_K)^2} \text{ are in surprisingly good agreement with the ratios of the MI internal conversion coefficients } \alpha_L/\alpha_K, $$

$$ P_L/P_K \text{ capture ratios. From Fig. 3-13 it is seen that experimental } \alpha_L/\alpha_K.
calculations for all Z. Precision measurements of additional M/L ratios of light atoms would be most useful to test exchange and overlap corrections.

A new more precise measurement on $^{65}$Zn is needed. Further experimental evidence is also required in the medium-Z region; the M/L capture ratios in the decay of $^{81}$Kr, $^{109}$Cd and $^{127}$Xe should be determined.

$P_N/P_L$ capture ratios. The only measurement of an $N/M$ capture ratio performed to date is that of Fengra (1974) on $^{205}$Pb. With a gaseous source of $^{205}$Pb tetraethyl, Fengra determined $P_N/P_L = 0.524 \pm 0.010$ and $P_N/P_K = 0.286 \pm 0.026$. Comparison with theory is impeded by lack of reliable information on the transition energy. An indirect determination of the $(N+...) / M$ ratio of $^{202}$Tl has been made from measurements of $(K+\bar{M}+...) / L$ and M/L ratios (Leutz et al., 1966), but the accuracy of this result is insufficient for meaningful comparison with theory.

Capture probability $P_K$. Selected K-capture probabilities for allowed and first-forbidden transitions are compared in Table 3.5 with theoretical predictions for allowed transitions. Two selected measurements on $^{40}$K are compared with theoretical capture probabilities for unique first-forbidden transitions. The K-capture probability, unless the reduced capture ratio, depends on the transition energy as well as on the atomic number. In Fig. 3-14 we have plotted the ratio of experimental to theoretical $P_K$ vs. Z. The recalculated exchange and overlap corrections according to Bahcall (1963a,c, 1965a) (Sec. 2.5) were used in the theoretical calculations. For several nuclides (e.g. $^{133}$Ba, $^{145}$Pm, $^{151}$Gd, $^{195}$Au), the energy $\omega_{EC}$ is not known with sufficient accuracy.
New, more accurate measurements for $P_K$ are desirable for some nuclides, e.g. for $^{73}$As, $^{75}$Se, $^{83}$Rb, $^{84}$Rb, $^{156}$Yb, and $^{195}$Au. The spin of the 307-keV level of $^{151}$Eu is not exactly known, it is quoted as $(3/2)^+$ or $(7/2)^+$. The transition from the $(7/2)^-\ ^{151}$Gd ground state to this level can therefore be non-unique or unique first-forbidden. Comparison of the measured $P_K = 0.811\pm 0.021$ with the theoretical $P_K = 0.740$ for a non-unique and $P_K = 0.428$ for a unique transition supports the $(7/2)^+$ assignment.

Experimental and theoretical K-capture ratios are seen from Fig. 3-14 to agree within $\pm 5\%$; there is no systematic difference between allowed and first-forbidden non-unique transitions.

3.3.4. Conclusions and Recommendations

From Tables 3.3-3.5 and from Figs. 3-12, 3-13, and 3-14 we find that experimental and theoretical electron-capture data agree rather well, viz., on the average to $\pm 3\%$ in the case of $L/K$ ratios, $\pm 9\%$ for $M/L$ ratios, and $5\%$ for $P_K$ values. The experimental accuracy is insufficient to distinguish between the theoretical correction factors for exchange and overlap effects. These effects are expected to be largest in the decay of $^7$Be (Odilet and Daudel, 1956; Bahcall, 1963). Experiments to measure the $P_L/P_K$ ratio of $^7$Be have been unsuccessful due to experimental limitations (Renier et al., 1968).

New, more accurate measurements of capture ratios and $P_K$ should be performed. More accurate results for second- and higher-order forbidden transitions would be useful to deduce nuclear matrix elements. Furthermore, more accurate $Q_{\alpha\alpha}$ energies are very much needed.
3.4 Determination of $K/\beta^+$ and $EC/\beta^+$ Ratios

In Secs. 3.4 and 3.5 we list all available experimental $K/\beta^+$ and $EC/\beta^+$ ratios and describe the experimental techniques involved in these measurements. We compare experimental ratios for allowed, unique first-forbidden and non-unique first-forbidden transitions with the appropriate theoretical values.

Source preparation is an important aspect of these measurements. Allowed $\beta^+$ emitters are generally short-lived, many of them having half-lives of the order of seconds, minutes or hours (Fig. 3-15). In order to study $\beta^+$ emitters with comparative ease a continuous supply of the source is therefore often necessary. Positron emitting nuclei are normally deficient in neutrons, hence one cannot prepare them by slow-neutron bombardment of stable isotopes in reactors. Instead, the stable isotopes are usually converted to radioactive isotopes by such reactions as $(\gamma,n)$, using machines like synchrotrons or electron linear accelerators, or by $(n,2n)$ reactions with fast neutrons from such devices as Cockroft-Walton generators or high-current electrostatic accelerators. Cyclotron irradiation with protons, deuterons or alpha particles to produce proton-rich (neutron-deficient) nuclei is another useful method of preparing positron emitters.

The radioactive source must be transported to the detector in a time that is short compared with the half-life. This problem has been solved, for example, by fast pneumatic transfer systems in which solid sources can be conveyed from the irradiation site to the detector in a fraction of a second. Continuous gas flow systems (Fig. 3-16) have also been
used extensively (Ledingham et al., 1965); if narrow-bore tubing is used in conjunction with a gas pressure of several atmospheres, the radioactive source (in gaseous form) can be conveyed to the detector in a very short time. Where the sources cannot be obtained in suitable solid or gaseous forms, the problem can often be solved by using liquids under pressure with the radioactive source dissolved in the medium or in suspension.

The main types of measurement used to determine K/β⁺ and EC/β⁺ ratios are summarized in Table 3.1. These various techniques and the sources of error involved in them are described in Secs. 3.4.1-3.

3.4.1. Measurements of K/β⁺ Ratios with Internal Sources

Internal-source proportional counter. In this method (No. 20), the radioactive source in gaseous form is mixed with the normal proportional-counter gas. If the half-life of the source is sufficiently long, the gases may be static, but for short-lived nuclei continuous production of the source and gas flow through the counter is employed. The electron-capture events are detected as discrete peaks superimposed on the positron continuum. A major part of the error in these measurements comes from the procedure adopted in separating the K-capture peak from the continuum.

Measurements of K/β⁺ ratios by this technique have generally been made under conditions where K x-ray escape from the counter is very small. For high-Z nuclei, the proportional counter must therefore be operated at high pressure. For low Z nuclei, counters can be operated at normal pressure, but for such nuclei the K/β⁺ ratio is usually
extremely small, whence it is often difficult to resolve the K peak from the positron spectrum.

We assume that the radioactive source can be produced with negligible competing activities, a situation which is usually attainable in practice. The positrons and K-capture events are detected with practically 100% efficiency. Then we have

\[
\frac{I_K}{I_{\beta^+}} = \frac{I_{K}}{I_{\beta^+}},
\]

where \(I_K\) and \(I_{\beta^+}\) are the measured intensities of the K peak and the \(\beta^+\) spectrum, respectively. Corrections have to be applied to \(I_{\beta^+}\) to account for the number of positrons which, unlike the K x-rays and Auger electrons, may enter the sensitive volume from the ends of the proportional counter. This correction was calculated to be 4.6% in the case of \(^{18}F\) (Drever et al., 1956).

Solid internal sources may also be employed (e.g., Avignon, 1956) but corrections for the absorption of the x-rays, Auger electrons and positrons in the source itself must then be taken into account.

In cases where the decay leads to an excited state of the daughter nucleus it is sometimes possible to measure coincidences between the spectrum in the proportional counter and the de-excitation \(\gamma\) ray, thus reducing the background. This technique was applied by Kramer et al. (1962b) to the decay of \(^{58}Co\).

**Internal-source proportional counter with anticoincidence.** This technique (No. 21) is similar to Method 20 and is particularly suitable
generally very much less intense than the positrons. In order to resolve weak K-capture peaks from the positron continuum, an anticoincidence counter is employed. One such counter with a plastic scintillator as anticoincidence detector is shown in Fig. 3-17. Both the positron and electron-capture events are detected in the central proportional counter; only the positrons can reach the surrounding counter. Thus, if signals from the central counter are taken in anticoincidence with those from the surrounding plastic scintillator, a well-resolved K peak is obtained. Figure 3-18 shows a typical K peak from $^{30}$P, measured with the counter shown in Fig. 3-17.

From the total spectrum in the central counter and the K peak in the anticoincidence spectrum, $I_K$ and $I_{\beta^+}$ are obtained and Eq. (3-45) applies as in Method 20.

Unless high-pressure counters are employed, this method becomes complicated for nuclei with $Z \geq 18$ because corrections for x-ray escape must be made. The method then becomes intrinsically less accurate, and hence, has so far been employed only in the low-Z region.

Internal-source scintillation counter. In this technique (Method 22), the radioactive source is distributed in a scintillating crystal (usually NaI) by introducing it into the melt from which the crystal is grown. The capture and positron events are detected in the scintillator, with the K x rays and K Auger electrons producing a well-defined peak so that the $K/\beta^+$ ratio can be determined. The interpolation of the continuum under the peak is a major source of error. Examples of this technique are the measurements of the $K/\beta^+$ ratios for $^{22}$Na with an error
of 9% (McCann and Smith, 1969) and for $^{60}$Co with an error of 2% (Joshi and Lewis, 1961). In both of these isotopes the decay leads to an excited state of the daughter nucleus which then de-excites by $\gamma$-ray emission. To reduce background, the positron and electron-capture events were measured in coincidence with the de-excitation gamma rays, detected in a second scintillation counter.

 Corrections must be applied for the escape of positrons from the source crystal before they have deposited sufficient energy to be detected. If coincidences are taken with a de-excitation gamma ray, allowance should furthermore be made for the loss of positron counts due to the summing of the gamma ray with a 511-keV positron-annihilation photon. A K peak from $^{22}$Na (McCann and Smith, 1969) is shown in Fig. 3-19. The difficulty of obtaining peaks at these very low energies with a scintillation counter is considerable. Specially selected low-noise photomultiplier tubes must be used in conjunction with an electronic system that is capable of eliminating afterpulses from long-lived phosphorescence associated with large energy deposition by positrons in the radioactive scintillator.

 Because the positrons and the K-capture events are detected with approximately 100% efficiency, Eq. (3-45) again applies, allowing for the corrections described above.

3.4.2. Measurements of K/$\beta^+$ Ratios with External Sources

Spectroscopy of positrons and K Auger electrons. In this type of measurement (No. 23), the areas under the Auger lines and the positron spectrum are measured. Since the Auger electrons and the positrons are
oppositely charged, a magnetic spectrometer with a Geiger, proportional, or scintillation counter is often used to analyze the radiations. The difficulty of subtracting a $\beta^+$ spectrum from a K peak is thus avoided.

In order to determine a $K/\beta^+$ ratio from such measurements, the value of the K-shell fluorescence yield $\omega_K^+$ must be known. There were often fairly large errors in the values of $\omega_K^+$ employed in the early experiments. However, Bambynek et al. (1972) have selected reliable measurements of $\omega_K^+$ and carried out a semi-empirical fit to these values. Thus, for many cases, uncertainty in $\omega_K^+$ need no longer seriously limit the accuracy of this method.

The relation

$$\frac{P_K}{P_{\beta^+}} = \frac{I_{KA}}{(1-\omega_K^+) I_{\beta^+}}$$

(3-46)

applies, where $I_{KA}$ is the total intensity of the K Auger lines. Corrections for absorption of low-energy Auger electrons and $\beta^+$ in the source are very important and contribute significantly to the errors involved in this technique.

Spectroscopy of K x rays and positrons. In this method (No. 24), a solid source is placed outside of semiconductor or scintillation counters. The K x rays and positron continuum are detected either in the same or separate counters, the Auger electrons generally being absorbed before reaching the detectors. A major uncertainty again arises from the subtracting of the $\beta^+$ spectrum from the K x-ray peak. As with Method 23, this technique requires knowledge of the fluorescence yield. Assuming that there are no competing activities, and correcting for
absorption, the equation applicable to this method is
\[ \frac{P_K}{P_{\beta^+}} = \frac{I_{KX}}{I_{\beta^{uma}K}}. \] (3-47)

Account must be taken of any differences in solid angle for the detection of K x rays and positrons. Self-absorption of x rays in the source is an important factor in this technique and makes the use of thin sources desirable.

Figure 3-20 shows how clearly the K x rays may be separated from the continuum in the decay of $^{91}$Mo (Fitzpatrick et al., 1975). This spectrum was obtained from a 5-mg/cm$^2$ thick, activated molybdenum foil placed 2 cm from a Si(Li) detector (area 30 mm$^2$, thickness 5 mm). The Ka and K\$ x rays of Nb are well resolved and the fluorescent K x rays of Mo caused by positron excitation of the foil can also be seen. Although the intensity of the K-capture branch in the decay of $^{91}$Mo is small (\(\sim 5\%\)), the error in estimating the areas of the K x-ray peaks can easily be kept as low as 1\%. There is, however, a difficulty in ensuring that the solid angles for the x rays and the positrons are the same, even when a single detector is employed. This difficulty can be reduced by using a detector with a large surface area. The K x-ray spectrum of $^{91}$Mo measured with a 5.1-cm \(\times\) 0.63-cm NaI(Tl) detector is shown in Fig. 3-21. The fine structure in the spectrum of Fig. 3-20 is unfortunately lost due to the intrinsically inferior resolution of NaI(Tl). Corrections are required for absorption of the K x rays and positrons and for the scattering of positrons out of the detector before they have deposited sufficient energy to be detected.
An interesting development of this technique is shown in Fig. 3-22 (Campbell et al., 1975). Here, the radioactive sample is placed between two CaF₂(Eu) scintillators in a 4π arrangement. This arrangement overcomes the problem caused by positrons being scattered out of the detector before depositing sufficient energy to be detected, or being scattered into the detector from surrounding material.

Many of the early K/β⁺ measurements in this category (Method 24) employed absorption techniques, typically with a Geiger counter and different absorbers to determine the relative intensities of the K x rays and the positrons. The accuracy of these measurements is very poor.

**Spectroscopy of K x rays and β⁺ annihilation photons.** This technique (No. 25) is similar to the previous method, but instead of detecting the β⁺ continuum, the positrons are stopped in an absorber and the 511-keV annihilation photons are detected. The source must be surrounded by sufficient material to ensure that all positrons are stopped at a well-defined position, as close to the source as possible. The K x rays and the annihilation photons may be counted simultaneously, with corrections applied to both intensities to allow for the presence of the β⁺ absorber. Alternatively, when the half-life of the source is sufficiently long, spectra taken with and without the absorber may be used to determine I_511 and I_KX⁺ respectively.

The K/β⁺ ratio is deduced from the relation

\[
\frac{P_K}{P_{\beta^+}} = \frac{2I_{KX}}{I_{511}}. \tag{3-48}
\]
A correction must be applied to $I_{511}$ for the loss of 511-keV $\gamma$-rays due to the summing of two such $\gamma$ rays; the size of this correction depends on details of geometry and the type of detector. The effect on $I_{511}$ of $\beta^+$ annihilation in flight (e.g., Kantele and Valkonen, 1973) must also be considered, although in many cases this has been assumed to be negligibly small.

### 3.4.3. Measurements of EC/$\beta^+$ Ratios

EC/$\beta^+$ ratios are determined by measuring the number of positrons emitted by the parent leading to an excited state of the daughter nucleus, and the number of $\gamma$ rays or conversion electrons from that level in a given time interval. Since the total number of $\gamma$ rays plus conversion electrons is equal to the total number of positrons and electron capture events--corrected with reference to the decay scheme where necessary--the ratio EC/$\beta^+$ of total electron capture to $\beta^+$ emission can be determined. Errors in these measurements can be kept very small, especially if the decay scheme is well-known. For example, the EC/$\beta^+$ ratios for $^{22}$Na and $^{58}$Co have been determined to $\approx 0.3\%$ and $\approx 0.7\%$, respectively. Errors in the decay scheme can, however, be large, and have led to large systematic errors in many of these measurements.

**Spectroscopy of $\gamma$ rays or conversion electrons and $\beta^+$ annihilation photons.** One of the simplest forms of EC/$\beta^+$ measurements consists of a comparison of the relative photopeak intensities of the de-excitation $\gamma$ rays and the $\beta^+$ annihilation photons in, for example, a scintillation or semiconductor detector (Method 26). As for Method 25, the source
must be surrounded by sufficient material to annihilate the positrons near the source to ensure that the solid angle is the same for both the nuclear and the annihilation photons. Corrections are required for absorption in the source and detector window, for decays to other levels in the daughter nucleus, for summing, and for annihilation of positrons in flight. In cases where the energy of the de-excitation γ rays is high it may be necessary to correct for a contribution to the annihilation photons due to internal and external pair production (e.g., Rupnik, 1972).

The total capture to β⁺ emission ratio is given by

$$\frac{P_{EC}}{P_{β⁺}} = \frac{2I_γ (1+\alpha)}{I_{511}} - 1,$$

(3-49)

where $I_γ$ and $I_{511}$ are the photopeak areas of the de-excitation γ ray and β⁺ annihilation photons, respectively, and $\alpha$ is the internal conversion coefficient.

A variation of this technique which has often been employed, particularly in the early measurements, is the comparison of the photopeak areas of the 511-keV and de-excitation γ rays for the source being investigated with similar areas for a source with a known EC/β⁺ ratio. Thus, if the subscripts $a$ and $b$ refer to the source with known and unknown EC/β⁺ ratio, respectively, we have

$$\left(\frac{P_{EC}}{P_{β⁺}}\right)_b = \left(\frac{I_γ}{I_{511}}\right) \left(\frac{I_{511}}{I_γ}\right) \frac{\varepsilon_{γa}}{\varepsilon_{γb}} \left[\frac{P_{EC}}{P_{β⁺}}\right]_a + \frac{1-\varepsilon_{γa}}{\varepsilon_{γa}} \left(\frac{I_{511}}{I_γ}\right) \left(\frac{I_γ}{I_{511}}\right) - 1.$$  

(3-50)

This method is suitable when the de-excitation γ rays for the two sources are of similar energy, since the ratio of efficiencies $\varepsilon_{γa}/\varepsilon_{γb}$
is then approximately unity. Hence the EC/$\beta^+$ ratio is then independent of detector efficiency. The accuracy of this method is obviously limited by the error in the EC/$\beta^+$ ratio of the standard source. Often $^{22}$Na was used for this comparison but the range of reported EC/$\beta^+$ values for this isotope is large (Table 3.6). Some authors did not even state which comparison value they employed.

A less common variation of this technique consists of measuring the intensities of the $\beta^+$ annihilation photons and the conversion electrons, rather than the de-excitation $\gamma$ rays. This method is only feasible in special cases where the internal conversion coefficient is high.

Several measurements have been carried out employing a similar technique in which the positron activity was determined from the area under the $\beta^+$ spectrum rather than from the intensity of the annihilation photons. As above, comparison with an isotope with a well-known EC/$\beta^+$ ratio was often employed. The results reported from this technique, however, have very large errors (>20%).

Measurement of $\beta^+ - \gamma$-ray coincidences. The principle of this method (No. 27) is to determine the number of $\gamma$ rays, $I_\gamma$, and of positron-$\gamma$-ray coincidences, $I_{\beta^+ - \gamma}$. Various combinations of detectors may be employed. Typically, scintillation or semiconductor detectors have been used for the $\gamma$ rays while the positrons were detected in proportional or scintillation counters. A $4\pi$ proportional counter or an internal-source scintillation counter (Leutz and Wenninger, 1967) have also been employed to detect the positrons.

The EC/$\beta^+$ ratio is given by
Comparison of $I_\gamma$ and $I_{\beta^+\gamma}$ with measurements for a source of known EC/$\beta^+$ ratio has often been employed.

**Sum-coincidence technique.** In this more sophisticated coincidence technique (No. 28), the quantities measured are the positron intensity $I_{\beta^+}$, the $\gamma$-ray intensities $I_{\gamma N}$ and $I_{\gamma S}$, and the positron-$\gamma$-ray coincidence intensities $I_{\beta^+-\gamma N}$ and $I_{\beta^+-\gamma S}$, where $\gamma_N$ refers to the normal de-excitation $\gamma$ ray and $\gamma_S$ is the sum of a $\beta^+$ annihilation photon and $\gamma_N$. It can be shown (Williams, 1964) that the relation

$$\frac{I_{\beta^+}I_{\gamma N}}{I_{\beta^+\gamma N}} = I_0$$

holds, where $I_0$ represents the total number of disintegrations.

Furthermore, we have

$$\frac{I_{\beta^+}I_{\gamma S}}{I_{\beta^+\gamma S}} = I_0 P_{\beta^+},$$

whence

$$\frac{P_{EC}}{P_{\beta^+}} = \frac{I_{\beta^+\gamma S}I_{\gamma N}}{I_{\beta^+\gamma N}I_{\gamma S}} - 1.\tag{3-54}$$

In a measurement of the EC/$\beta^+$ ratio for $^{22}$Na (Williams, 1964), the $\beta^+$ activity was determined with a $4\pi$ proportional counter. For the detection of $\gamma_S$, two large NaI(Tl) crystals were used to obtain a high efficiency for the summation events. For $\gamma_N$, one smaller NaI(Tl) crystal was used to minimize the efficiency to summation events. The simplifying assumptions involved in Eq. (3-54) and the corrections which must be applied are discussed in detail by Williams (1964).
Measurement of triple coincidences. The EC/β⁺ ratio can be obtained by taking the γ-ray spectrum in triple coincidence with two β⁺ annihilation photons (Method No. 29). The two counters for annihilation photons are placed opposite each other with analyzer channels set to record the 511-keV photopeaks only. Due to the nature of the annihilation process, the efficiency for the detection of coincidences of two 511-keV γ rays at 180° is sufficiently increased over other coincidences that even very weak positron emission can be detected. A typical electronic arrangement for this type of measurement is shown in Fig. 3-23. The γ-ray singles intensity Iγ and of the triple coincidence intensity IC are measured. If similar measurements are made for a source a whose EC/β⁺ ratio is known, then the unknown EC/β⁺ ratio source b is

\[
\left( \frac{P_{EC}}{P_{β⁺}} \right)_b = \left( \frac{I_{γ}}{I_C} \right)_b \left( \frac{I_C}{I_{γ}} \right)_a \left( \frac{P_{EC}}{P_{β⁺}} \right)_a \left( \frac{I_{EC} + 1}{1 + \alpha_a} \right) - 1.
\]

(3-55)

Corrections are required for such effects as summing, β⁺ annihilation in flight, differences in the detection of annihilation radiation for the two sources due to possible differences in solid angle and in summing of the γ rays and the annihilation radiation, and the possibility of coincidences due to Compton events from high-energy γ rays being registered in the analyzer window of the annihilation detectors.

Measurement of (γ-ray)-(β⁺-annihilation-photon) coincidences. The various coincidence techniques are very similar in principle and this method (No. 30) is essentially a variation of Method 27. The quantities measured are the number of nuclear γ rays Iγ and the number of coincidences of nuclear and β⁺ annihilation photons I511-γ. The usual corrections
for absorption, summing, and \( \beta^+ \) annihilation in flight are required.

The EC/\( \beta^+ \) ratio is given by

\[
\frac{P_{EC}}{P_{\beta^+}} = \frac{2I_4}{I_{511-\gamma}} - 1.
\]

\( (3-56) \)

**Miscellaneous.** The experiments in this group (No. 31) do not fall readily into any of the other categories. Many of the experiments were carried out by employing various combinations of methods 20-30. No loss of accuracy need be implied. This category also includes methods which have been employed in only very few, exceptional cases and because of their limited application do not warrant description as a separate technique. Also included in the miscellaneous category are a few experimental results whose methods are in doubt due to incomplete details provided in the published papers.

One different approach to EC/\( \beta^+ \) measurements is the technique employed by Allen et al. (1955) for the determination of the EC/\( \beta^+ \) ratio for \( ^{22}\text{Na} \). This involves a comparison of the number of positrons emitted from the source with the number of daughter atoms produced (Alvarez, 1937). The positron activity was determined using a 4\( \pi \) Geiger counter and the rate of evolution of the daughter (Ne) was determined by gas analysis.

Another interesting technique has been applied by Gleason (1959) to \( ^{65}\text{Zn} \) which decays by electron capture and \( \beta^+ \) emission to the ground state and by electron capture to the first excited state of \( ^{65}\text{Cu} \). Using a measured value for the efficiency of detection of the de-excitation \( \gamma \) ray, the total electron capture decay rate and the electron capture
branching ratio were determined from measurements of the K x-ray counting rate, the γ-ray singles rate and (K-x-ray)-(γ-ray) coincidence counting rate. The assumption was made that the ratio of K-electron capture to total electron capture was the same for both branches. The $\beta^+$ emission rate was determined by counting coincidences of annihilation photons in two detectors at 180° and thus the EC/$\beta^+$ ratio for the ground state transition was found. The important feature of this technique is that although K x rays were used to indicate the occurrence of electron capture, the deduced value of the EC/$\beta^+$ ratio is independent of the fluorescence yield.

3.5. Experimental Results and Comparison with Theory for K/$\beta^+$ and EC/$\beta^+$ Ratios

3.5.1. Results

All published experimental K/$\beta^+$ and EC/$\beta^+$ ratios are listed in Table 3.6. Table 3.7 contains selected experimental K/$\beta^+$ and EC/$\beta^+$ ratios for allowed transitions. Only ratios for transitions to a single final state in the daughter nucleus are included. Unfortunately, information provided on some measurements was not complete and these results had to be rejected. Where the $\omega_K$ values were stated, results were recalculated using the latest reliable fluorescence yields, derived with the aid of Eq. (3-35).
The remaining $K/\beta^+$ and $EC/\beta^+$ ratios were found to lie in two distinct groups, one with errors ranging up to 12.5% and the other, consisting mainly of the earlier measurements, with considerably larger errors. Since the two groups are well separated only the results from the former are considered further.

Tables 3.8 and 3.9 contain selected results for first-forbidden unique and first-forbidden non-unique transitions. Results with errors greater than 25% or without quoted errors were excluded.

### 3.5.2. Theoretical Predictions

**Allowed transitions.** The theoretical $K/\beta^+$ ratios for allowed transitions in Table 3.7 have been calculated according to the relation

$$\frac{P_K}{P_{\beta^+}} = \frac{\pi \rho^2 (W_0 + W_n)^2 \rho K}{2 \rho_0^2 \rho_0^2 (W_0 - n)^2 F(Z', W) dp}$$

(3-57)

[See Eqs. (2-111), (2-125), and (2-126)]. Small corrections [Eq. (2-128)] were neglected. The values of $\rho K$ were taken from Mann and Waber (1973) (Sec. 2.2.2) and the intensity of the $\beta^+$ spectrum was computed with the tables of Fermi functions of Behrens and Jünecke (1969). The energies $W_0$ were taken from the atomic mass tables of Wapstra and Gove (1971). Errors in the theoretical $K/\beta^+$ ratios in Table 3.7 reflect only the uncertainty in $W_0$ obtained from Wapstra and Gove. The value of $\rho K$ used in these calculations is discussed in Sec. 3.5.3.
The theoretical EC/\beta^+ ratio for allowed transitions is

\[
\frac{P_{EC}}{P_{\beta^+}} = \frac{P_{KC}}{P_{\beta^+}} \frac{P_{EC}}{P_{KC}} = \frac{P_{KC}}{P_{\beta^+}} \sum_x \beta_x^2 \beta_x^2,
\]

where \( x \) stands for \( K, L_1, L_2, M_1, \) or \( M_2 \), and \( P_{KC}/P_{\beta^+} \) is the theoretical \( K/\beta^+ \) ratio for an allowed transition [Eq. (3-58)].

**Unique forbidden transitions.** In general, the \( K/\beta^+ \) ratio for forbidden transitions is

\[
\frac{P_{KC}}{P_{\beta^+}} = \frac{\pi \beta^2}{2} \frac{C_K}{C_{\beta^+}} \frac{C_{K^*}}{C(W)}
\]

where \( C_K \) and \( C(W) \) are shape factors and the bar represents averaging over the \( \beta^+ \) spectrum [Eq. (2-134)]. The shape factors contain matrix elements and are functions of \( W \) and \( W_0 \). For unique forbidden transitions it is possible to separate the matrix-element and the energy dependence of \( C_K \) and \( C(W) \) to give explicit expressions for the ratio \( C_K/C(W) \) (Sec. 2.4.4).

The first-forbidden unique transitions Eq. (2-135) is simplified to

\[
\frac{C_K}{C(W)} = \frac{(W + W_0)^2}{q^2 + \lambda^2 P^2},
\]

where \( q \) is the neutrino momentum, \( p \) is the positron momentum, and the bar represents averaging over the \( \beta^+ \) spectrum. The theoretical first-forbidden unique \( K/\beta^+ \) ratios shown in Table 3.8 have been calculated using these expressions, with values of \( \lambda_2 \) from the tables of Behrens and Jänecke (1969). For comparison of theory and experiments, one can
use the approximations \( W_K \approx 1 \) and \( q^2 + \lambda^2 \psi^2 \psi(W - 1) \), whence

\[
\frac{C_K}{C(W)} \approx \frac{2(W + 1)}{W - 1}.
\]  

Equation (3-61) has an accuracy of a few percent.

Non-unique forbidden transitions. For non-unique forbidden transitions, \( K/\beta^+ \) ratios cannot, in general, be calculated explicitly (Sec. 2.4.3). For the special case of non-unique first forbidden transitions, however, which exhibit a \( \beta^+ \) spectrum with an allowed shape, the \( K/\beta^+ \) ratio is expected to be the same as for allowed transitions. Information about the shapes of some \( \beta \) spectra is given by Paul (1966) and Daniel (1968). For many of the non-unique first-forbidden decays listed in Table 3.9, however, details of the spectrum shape are not available. Nevertheless, to provide a general comparison, allowed theoretical \( K/\beta^+ \) and \( EC/\beta^+ \) ratios are indicated for all cases.

3.5.3. Comparison of Experiment and Theory

Allowed transitions. Theoretical and selected experimental values for \( K/\beta^+ \) and \( EC/\beta^+ \) ratios are listed in Table 3.7. Exchange and overlap corrections have been neglected in the theoretical ratios; they affect the total capture probability and \( \beta^+ \) emission rate only slightly (Bahcall, 1963a). The EC probability for \( ^7 \)Be, e.g., is affected by \(<0.1\% \), and that of \(^{37}\)Ar, by \(<0.7\% \) through exchange and overlap; the \(^{65}\)Zn \( \beta^+ \) decay rate is affected by \(<0.1\% \), and that of \(^{14}\)C, by \(<0.1\% \). The theoretical \( K/\beta^+ \) ratios in Table 3.7 include a correction factor according to Bahcall (Table 2.11); from \( Z>32 \), the factors of
Suslov (1970) were used. At present, EC/β⁺-ratio measurements (Table 3.7) are not nearly accurate enough to help decide between the two sets of exchange and overlap correction factors listed in Table 2.11.

Figure 3-24 shows the ratio of experimental to theoretical values for all the results in Table 3.7. The interesting and very accurate point for ²²Na is plotted in the inset. For most of the decays, the experiment/theory ratio is less than unity; exceptions are ¹¹C, ¹⁵O, ¹⁹Ne, ⁸⁹Zr, ⁸⁹ᵐZr, and ¹¹¹Sn. The disagreement between experiment and theory apparently increases with Z.

In the theory of allowed transitions, only s-wave leptons are considered and the EC/β⁺ and K/β⁺ ratios are independent of nuclear matrix elements. In the general case, leptons do not leave the nucleus only radially, and small contributions from p and d waves must be considered. This gives rise to higher-order matrix elements that do not cancel in the ratios (Sec. 2.4.2). A correction factor has been determined [Eq. (2-128)] that slightly reduces the theoretical ratios, by as much as 3% at Z=80.

The possible existence of second-class currents does not significantly affect electron-capture to positron-emission ratios (Behrens and Bühring, 1974).

First-forbidden unique transitions. For these transitions the experimental K/β⁺ ratios are compared in Table 3.8 with first-forbidden unique theoretical ratios. There is agreement within the errors between experiment and theory, but the experimental accuracy is fairly poor.
First-forbidden non-unique transitions. The experimental $K/\beta^+$ and $EC/\beta^+$ ratios for these transitions are compared in Table 3.9 with the corresponding theoretical ratios for allowed transitions. The comparison is made for interest only; a complete theoretical treatment requires knowledge of the nuclear matrix elements which for these transitions do not cancel.

3.5.4. Conclusions and Recommendations

It can be seen from Fig. 3-24 that theoretical allowed $K/\beta^+$ and $EC/\beta^+$ ratios are systematically larger than experimental ratios; the discrepancy apparently increases with Z. Higher-order effects, such as second-class currents, corrections of the type described by Eq. (2-128), and radiative corrections are insufficient to resolve the difficulty. The question of radiative corrections is still unsettled; it has been shown (Sec. 2.4.2) that these corrections partially cancel out. There remains a model-independent part of the radiative corrections, however, which differs in the case of electron capture from that in positron emission. This model-independent correction includes the well-known emission of real photons (internal bremsstrahlung). Calculations for $\beta^+$ emission have been carried out to order $\alpha$, e.g. by Tilkinson and Macefield (1970); an increase in the probability of $\beta^+$ emission is found which thus reduces the theoretical capture-to-positron ratios. The correction factor increases as $W_0$ decreases and as Z increases and amounts to 1.5% for $^{58}\text{Co}$ (Williams, 1970) if it is assumed that the correction is multiplicative and not additive. Radiative corrections for electron capture
have not yet been calculated.

It would be of interest to establish with greater accuracy the Z dependence of the trend shown in Fig. 3-24, if indeed such a simple functional dependence on Z exists. Remeasurements, preferably using different techniques, for any of the decays in Table 3.7 would be useful. The decays of $^{65}\text{Zn}$, $^{111}\text{Sn}$ and any high-Z isotope are possibly the most interesting for study. The question of whether there is real agreement between theory and experiment in the case of first-forbidden unique transitions is still open; measurements on $^{84}\text{Rb}$, $^{122}\text{Sb}$ and $^{126}\text{I}$ should be repeated with greater accuracy.

The theory of atomic exchange and imperfect wave-function overlap effects needs to be refined and calculations must be extended to low Z. Critical experiments on capture/$\beta^+$ ratios which would differentiate between theoretical approaches have yet to be carried out. The problem of establishing the overlap and exchange correction for the K shell cannot be resolved by measuring K/$\beta^+$ ratios alone. The most sensitive isotope for study is $^7\text{Be}$ which decays solely by electron capture; a measurement of $P_K$ for this isotope is very desirable (Sec. 3.3.4).

Some new and interesting EC/$\beta^+$ ratios have recently been reported by Firestone et al. (1974, 1975a). Anomalously high ratios are found for hindered allowed transitions in $^{145}\text{Gd}$ and $^{143}\text{Sm}$; these are attributed to the interference of higher-order nuclear matrix elements. It would be of great value to verify this experimental finding.

Theoretically K/$\beta^+$ ratios for allowed transitions are plotted in Figs. 3-25 and 3-26 as functions of Z and of the $\beta^+$ end-point kinetic energy. These ratios were calculated according to Eq. (3-58) with $B_K = 1$; the graphs may be used where an accuracy of $\pm 10\%$ is sufficient.
4. RADIATIVE ELECTRON CAPTURE

4.1. Theory

Radiative electron capture consists of processes which lead to the production of a continuous spectrum of electromagnetic radiation during electron capture (EC) decays. Such processes involve the emission of one or more photons during a single EC event. The energy released in the decay is shared statistically among these photons and the neutrino, thus accounting for the continuous nature of the resulting spectra. The most probable radiative-electron capture events are those in which a single photon accompanies the neutrino. The radiation emitted even in this mode is quite weak, the total probability for the emission of a single photon being of the order of $10^{-4}$ per EC event. Radiative electron capture processes in which more than one photon is emitted occur with far smaller probabilities. Their contributions to the radiation spectra are completely insignificant and will not be considered further.

From the point of view of perturbation theory, radiative electron capture is a second-order process involving both beta and electromagnetic radiative transitions. The two transitions connect the initial and final states of the system through a set of virtual intermediate states. In general, there are two fundamentally different types of intermediate states through which the process can proceed. They are represented pictorially by the
Feynman diagrams shown in Fig. 4-1. The first type [Fig. 4-1(a)] involves only excited electronic states, and the radiation is produced by the sudden acceleration of charge and magnetic moment associated with the orbital electron's capture. This radiation is commonly referred to as internal bremsstrahlung (IB). The second type [Figs. 4-1(b) and (c)] involves excited nuclear states and the radiation arises from a nuclear transition which may either precede or follow the virtual EC decay. These two decay modes are variously denoted as electronic beta-gamma and nuclear beta-gamma transitions or, more simply, direct and detour transitions. In allowed decays, detour transitions are expected to occur at a \( \sim 10^6 \) times smaller rate than direct transitions. In forbidden decays, this difference can be less pronounced (Longmire, 1949; Horowitz, 1952).

Extensive calculations on detour transitions were carried out by Rose et al. (1962) and Lessile (1963) for the especially interesting situation in which the initial and intermediate nuclear states, connected by a virtual EC transition, are almost degenerate. It was shown that the spectrum of the radiation arising from detour transitions is sharply peaked near the endpoint under these circumstances, in contrast to the usual IB spectrum. It was hoped that this deviation of the photon spectrum from its IB form might be observable, revealing the presence of detour transitions, even though their contribution was still expected to be quite small. An experiment designed to test these ideas was reported shortly thereafter by Schmorak (1963), who
studied $^{59}$Ni, a nucleus possessing a decay scheme with the required characteristics, and found that the observed spectrum did indeed show a very small distortion from the predicted IB form near the end point. Attributing this distortion to the presence of detour transitions, Schmorak (1963) concluded that such transitions account for no more than $0.6\%$ of the total radiative K-capture transition rate.

While the contribution of detour transitions is of great interest for the study of nuclear structure, such transitions usually do not significantly affect the shape or intensity of radiative electron-capture spectra. For this reason, and because available theoretical results on detour transitions are very limited, such transitions will be disregarded here and all calculations will be confined to the determination of the direct-transition amplitude shown diagrammatically in Fig. 4-1(a). Clearly, a highly accurate theory of the direct-transition process will be necessary to permit the identification of any detour-transition contributions in observed spectra.

4.1.1. Matrix Elements and Transition Rates

Radiative electron capture is expected to occur with significant probability only for the innermost electrons of the atom. Since the available energy is usually greater than the K-shell binding energy, the K electrons, which spend the most time in the neighborhood of the nucleus, are expected to provide the dominant contribution to the IB spectra (except at very low photon
energies where \(2p\)-state capture provides the dominant contribution. In all but the very lightest atoms, the potential in which the innermost electrons move is primarily the Coulomb potential of the nucleus. For this reason, all electron-electron interactions and the screening and correlation effects for which they are responsible are neglected in current theories, and it is assumed that each orbital electron is initially moving under the influence of only the nuclear Coulomb field.

Accordingly, the unperturbed electron-field operator \(\psi_e(x)\) is chosen to satisfy a Dirac equation containing the nuclear Coulomb field,

\[
(\gamma_\mu \partial_\mu + \gamma_4 Z e/r)\psi_e(x) = 0. \quad (4-1)
\]

In this representation, the interaction Hamiltonian density is

\[
H_I(x) = H_{\gamma}(x) + H_{EC}(x), \quad (4-2)
\]

where \(H_{\gamma}\) represents the interaction of the electron field with the Maxwell field and \(H_{EC}\) represents the EC interaction, assumed to be of the standard \(V-\lambda A\) form. The matrix element associated with diagram (a) of Fig. 4-1 is derivable by standard quantum-field theoretic methods. As Glauber and Martin (1956) have shown, it can be written

\[
M_\alpha = i e C_\gamma (2\pi/k)^{1/2} \int d\mathbf{r}_N \int d\mathbf{r}_N' \int d\mathbf{r}_N'' \int d\mathbf{r}_N''' \frac{\bar{\Phi}^{(\gamma)}(\mathbf{r}_N') \Gamma_\mu \Phi^{(\gamma)}(\mathbf{r}_N''')}{i k \mathbf{r}_N - \mathbf{r}_N'} \times \gamma_\mu e^{-ik \cdot \mathbf{r}_N} \psi_\alpha(\mathbf{r}), \quad (4-3)
\]
with the matrix element of the nuclear EQ current density defined by

\[ j^N_{\mu}(r_N) = \langle N | \overline{\psi}_n(r_N, 0) \gamma_\mu \psi_p(r_N, 0) | N \rangle. \tag{4-4} \]

In these equations, \( C_\nu \) is the weak-interaction vector coupling constant, and we have \( \lambda = \left| C_A / C_\nu \right| \), \( \lambda_\mu = \gamma_\mu (1 + \lambda \lambda_5) \), and \( \lambda_\mu = \gamma_\mu (1 + \lambda_5) \).

The \( \psi_n \) and \( \psi_p \) are the nucleon field operators, and \( \psi_\nu \) and \( \psi_\alpha \) are the Dirac spinor wave functions for a neutrino of momentum \( P_\nu \) and an initial electron in state \( \alpha \), respectively. The one-photon state, characterized by momentum \( k \) and polarization \( \epsilon \), has been normalized to a unit volume. The intermediate-state sum which appears in Eq. (4-3) has been identified as the eigenfunction expansion for the Dirac-Coulomb Green's function,

\[ G_{E}(x_N') = \sum_\beta \frac{\phi_\beta(x_N') \overline{\phi}_\beta(x)}{(E_\beta - E)}, \tag{4-5} \]

with \( E = E_\alpha - k \), where \( E_\alpha \) is the relativistic energy of the orbital electron undergoing capture.

Two comments on the structure of the matrix element are in order. First, it should be noted that the role played by positrons in the radiative capture process is included implicitly in the structure of \( M_\alpha \). One type of path through which the radiative capture process can proceed is the emission of a virtual positron by the nucleus followed by its single-quantum annihilation with an orbital electron. Such paths are accounted for by the presence of the various negative-energy eigenstates in the expansion of the Green's function. Thus, the structure of the
Green's function is such that complete account is taken of the role of positrons in the radiative capture process.

Since the theory developed so far assumes the presence of any number of orbital electrons moving independently in the Coulomb field of the nucleus, the Pauli exclusion principle forbids virtual radiative transitions to intermediate states which are already occupied. Presumably such occupied intermediate states should then be excluded from the eigenfunction expansion. However, as was first pointed out quite generally by Feynman (1949) and emphasized by Glauber and Martin (1956) in reference to radiative electron capture, the presence of an obstructing electron makes another path possible for the radiative capture process, which is not otherwise available. This path consists of virtual transitions of the obstructing electron followed by a radiative transition.

Feynman has shown that, for a noninteracting system, the total amplitude for such a new path exactly compensates for that of the forbidden intermediate states; thus one may perform the calculation as if all the other states were unoccupied.

Feynman's result is easily generalized to include the presence of a static external field, such as the field of the nucleus, and consequently it has been assumed valid in all theoretical studies on radiative electron capture. However, as pointed out by Persson and Koonin (1972), radiation before capture takes place in the Coulomb field of element Z, while radiation following capture takes place in the field of element Z-1. Consequently, those terms in the eigenfunction expansion for the Green's function which
correspond to occupied atomic states should really be represented by Coulomb eigenfunctions for element Z-1 rather than element Z. Undoubtedly, for \( Z \gg 1 \) the corrections resulting from such a modification of the eigenfunction expansion are entirely negligible. However, for very low-Z elements, especially at the lower photon energies (i.e., very low values of \( \hbar \omega \)) where the poles corresponding to the bound states contribute strongly to the transition amplitude, such a modification of the Green's function may prove to be of importance.

The Green's function introduced in Eq. (4-3) and defined by the eigenfunction expansion [Eq. (4-5)] is seen to satisfy the inhomogeneous differential equation

\[
G_{\alpha}(r_N, r) \gamma_\mu[H_\alpha(r) - \mathcal{E}] = \delta(r_N - r), \quad (4-6)
\]

where \( H_\alpha \) is the Dirac-Coulomb Hamiltonian. As Glauber and Martin (1956) have shown, the evaluation of \( M_\alpha \) is facilitated by the introduction of the second-order Dirac-Coulomb Green's function \( G_E(r_N, r) \), defined by

\[
G_E(r_N, r) \equiv G_E(r_N, r)[\gamma_\mu \gamma_\nu \gamma_\lambda \gamma_\sigma (E + \hbar \omega/r) + 1] \quad (4-7)
\]

and satisfying the inhomogeneous second-order equation

\[
g_{\alpha}(r_N, r) [\mathcal{V}^2 + (\mathcal{E} + \hbar \omega/r)^2 - 1 - i \mathcal{E} \alpha^\ast (1/\alpha)] = -\delta(r_N - r). \quad (4-8)
\]

With the introduction of Eq. (4-7), the matrix element of Eq. (4-3) lends itself to considerable simplification and can be written

\[
M_\alpha = i e C_v (2\pi/\hbar)^{1/2} \int d^3r_N \psi^\ast_N \sum_{\mu} \int d^3r \, E_E(r_N, r) e^{-ik \cdot r} \chi[-2\mathcal{E} \gamma_\nu \gamma_\lambda \gamma_\sigma \sum_{\lambda_0} \hbar \omega \gamma_\mu \gamma_\lambda \gamma_\sigma \Phi_\alpha(r), \quad (4-9)
\]
In the:Sécs. 4.1.2: 4.1.4, the evaluation of $\mathcal{M}$ and related quantities is described and final results are presented for allowed and first-forbidden transitions. We note that the differential transition rate is determined by the usual formula of time-dependent perturbation theory (Fermi's "Golden Rule") and is given by

$$
\frac{d\omega}{dx} = (2\pi)^{-\frac{5}{2}} |M|^2 \delta(E_{\nu} + k - q) \frac{dP_{\nu}}{d\omega} \frac{dk}{d\omega}.
$$

where $q^2 = q_{\nu}^2 - E^2_{\nu}$ has been introduced to represent the total available energy, shared between the photon and the neutrino.

4.1.2. 1B Spectra: from Allowed Transitions

For allowed transitions, the lepton functions of $r_N$ are usually replaced by their values at $r_N = 0$. However, one must exercise some caution in doing this since the Coulomb Green's function $g(r)$ is known to be weakly singular at $r = 0$. To get around this difficulty it is necessary to take account of the fact that the EC interaction actually takes place over a finite nuclear volume, by averaging the Green's function over this volume.

This averaged Green's function will be denoted by $\langle g_E(r) \rangle_0$. Thus, for allowed transitions the matrix element of Eq. (4-9) is simplified to

$$
\mathcal{M}_{\alpha \alpha} = ie_N \left( \frac{2\pi}{k} \right)^{\frac{1}{2}} J_{\frac{3}{2}}(0) \Gamma \int d\mathbf{r} \langle g_E(r) \rangle_0 e^{-ik \cdot r} \times \left[ -2\mathbf{e}_\lambda \cdot \mathbf{r} + \mathbf{e}_\lambda \cdot k - \mathbf{e}_\lambda \cdot r \right] \tilde{\phi}_{\alpha}(r).
$$

where the nuclear EC transition current has been introduced,
defined by \( J^N = \int \sum_\mu^N \mathcal{E}^N_{\mu}(r_N) \). A nonrelativistic approximation for the nuclear motion leads to \( J^N = \langle \psi(\nu), \langle \psi \rangle \rangle \), where \( \langle \psi \rangle \) and \( \langle \psi \rangle \) are the familiar matrix elements associated with Fermi and Gamow-Teller transitions.

**Coulomb-Free Theory**

The earliest theory of IB spectra in allowed transitions was developed independently by Møller (1937a) and by Morrison and Schiff (1940). This theory is presently of interest because of its simplicity and because it yields IB spectra that are accurate at high photon energies. The more sophisticated theory developed later by Glauber and Martin (1956) may be viewed as providing correction factors for the basic results.

Morrison and Schiff (1940) simplified the problem by neglecting the momentum (and binding energy) of the initial electron and by neglecting the influence of the Coulomb field on the intermediate electron states. The first of these assumptions is only valid when the recoil momentum of the electron after photon emission greatly exceeds its initial momentum (of average value \( Z_\alpha \)). The second approximation consists of assuming a Born-approximation treatment of the intermediate states. For its validity, this approximation requires that \( Z_\alpha/v << 1 \), where \( v \) is the velocity of the electron after photon emission. It is evident that both approximations restrict the results to photons in the high-energy region where \( k \) is much larger than \( Z_\alpha \).

Ignoring the Coulomb field in the intermediate states amounts to using the free-particle relativistic Green's function found by
solving Eq. (4-8) with $E=0$. The initial momentum of the electron is neglected by approximating its wave function by a constant, equal to the value of the wave function at the origin. Under these approximations, the calculation of the matrix element of Eq. (4-11) is greatly simplified and leads to the result:

$$M_{n^s} = -\frac{ie\hbar k}{v} \left( \frac{\hbar}{2m} \right)^{1/2} J_{n^s} \left( \frac{\hbar}{m} \right)^{1/2} \phi_{n^s}^* \left( \phi_{n^s} \right) \cdot \phi_{n^s}^* \cdot \phi_{n^s}. \quad (4-12)$$

It is important to note that as a consequence of neglecting the momentum of the initial electron, radiative electron capture from an initial electron state of nonvanishing angular momentum is forbidden. Both the electric and magnetic contributions to the IR radiation vanish under these circumstances; this is immediately evident from the structure of the matrix element of Eq. (4-12).

When Eq. (4-12) is substituted into Eq. (4-10) and the appropriate momentum and spin summations are completed, the IR spectra associated with ns-state capture are found to be:

$$d\omega_{n^s} = \frac{2\pi k}{n^2} C_{s^2} N_{s^2} \left| \phi_{n^s}^* \left( \phi_{n^s} \right) \cdot \phi_{n^s}^* \cdot \phi_{n^s} \right|^2 \left( k(q_{n^s} - k) \right)^2 dk. \quad (4-13)$$

The ratio of the radiative capture rate to that for ordinary K capture is:

$$\frac{d\omega_{n^s}}{\omega_K} = \frac{a}{n} \left| \phi_{n^s}^* \left( \phi_{n^s} \right) \cdot \phi_{n^s}^* \cdot \phi_{n^s} \right|^2 \frac{k(q_{n^s} - k)^2}{a_{1s}^2} dk. \quad (4-14)$$

Hence the total radiative capture rate per K-capture event is
More generally, if only photons with \( k_0 \) are detected, the integrated radiative capture rate per \( K \)-capture event is given by

\[
\frac{\nu_{ns}}{\nu_K} = \frac{1}{\nu_K} \int_{k_0}^{\infty} \frac{d\nu_{ns}}{dk} dk = \frac{a q_{ns}}{\pi} \left( \frac{q_{ns}}{q_{1s}} \right)^2 \left( \frac{q_{ns}}{q_{1s}} \right)^2,
\]

Equations (4-17) and (4-18) were first derived by Møller (1937a) and by Morrison and Schiff (1940). The more general results for arbitrary \( s \)-state capture [Eqs. (4-14) and (4-15)] were first reported by Glauber and Martin (1956).
which account is taken of both relativistic and Coulomb effects, have been reported by Glauber and Martin in two well-known papers.

In their first paper on the subject, Glauber and Martin (1956) developed the general formalism for allowed transitions (Sec. 4.1.1) and evaluated it to a relative accuracy of order \( Z_a \) for both s- and p-state capture. Certain relativistic corrections that are important for s-state capture at low energies were also calculated. In their second paper, Martin and Glauber (1958) developed more elaborate methods which make detailed calculations possible in which relativistic and Coulomb effects are included to all orders in \( Z_a \). These results lead to certain integrals which cannot be evaluated exactly in closed form or tabulated easily. To obtain numerical results, Martin and Glauber developed \( Z_a \) expansions for these integrals and carried out their evaluation to a relative accuracy of order \((Z_a)^2\). This limitation on their otherwise exact results for radiative K capture has been removed recently however, by Intemann (1971), who has shown how to evaluate the integrals exactly using partly numerical methods. We briefly outline this theory and summarize its final results.

Nonrelativistic calculations. For moderately light nuclei it is expected that the initial electronic states can be described adequately by nonrelativistic Coulomb wave functions, especially for capture from the higher shells. In view of the greater complexity attendant to the use of Dirac-Coulomb wave functions, it is natural that nonrelativistic calculations be considered first. In general, these are expected to yield results with a relative
accuracy of order $Z\alpha$. In order to preserve this level of accuracy at all photon energies, however, it is necessary to employ somewhat more accurate wave functions, in order to correct for certain relativistic effects which have a pronounced influence on the low-energy portions of the s-state spectra (Glauber and Martin, 1956).

A particular advantage of introducing the second-order Green's function is that, consistent with the use of nonrelativistic wave functions, an approximate Green's function $g_2^i(r, r')$ can be employed, which has a particularly simple structure. This Green's function, obtained by neglecting the $(Z\alpha/r)^2$ and $Z\alpha a^r(\mu l/r)$ terms in Eq. (4-8) and solving the resulting equation, has been studied in considerable detail by Glauber and Martin (1956). In particular, $g_2^i(0, r)$ has been shown to possess the integral representation

$$g_2^i(0, r) = (\mu/2\pi)e^{-\mu r} \int_0^\infty ds s^\eta (1 + 5s^2)^\eta e^{-2\mu rs},$$  \hspace{1cm} (4-19)

where $\mu = (1-E^2)^{1/2}$ and $\eta = Z\alpha E/\mu$.

(i) $s$-state radiative capture. For radiative capture from an $s$ state, the contribution to $M_\alpha$ from the $e^{-V}$ term vanishes from symmetry considerations; when terms of order $Z\alpha$ are neglected, the remaining contribution can be evaluated using only very general properties of the Green's function. Final results for the transition rates are identical with those of the Coulomb-free theory [Eq. (4-14) et seq.]. The calculations of Glauber and Martin (1956) reveal, however, that the range of validity of the Coulomb-
free theory is much greater than could have been anticipated on
the basis of the calculations of Möller (1937a) or Morrison and
Schiff (1940). Indeed, it was established by Glauber and Martin
(1956) that the Coulomb-free theory yields results for the 1B
spectra associated with s-state capture which are formally cor-
rect to order \( Z \alpha \) for all photon energies. It is also true how­
ever, that for the low-energy portion of s-state spectra, the
factor of \( Z \alpha \) is partially compensated by an increased probability
of radiation. Consequently, in order to obtain results for which
the actual error is not greater than order \( Z \alpha \), it is necessary to
carry the calculations to the next order in \( Z \alpha \) and omit only
those terms which are actually of order \( Z \alpha \) or less. Glauber and
Martin (1956) accomplished this by means of a Foldy-Wouthysen
transformation applied to the Dirac-Coulomb wave functions and
Green's function. The result, valid for \( n<2 \) and \( k \leq Z \alpha \), is

\[
\begin{align*}
W_{ns} & = \frac{1}{2} \mathcal{E}_o \left( \frac{n^2}{2k^2} \right)^{1/2} \frac{1}{\mu} \frac{\partial}{\partial \mu} \left[ \Sigma^{+} e^{i \mu \cdot k} & - e^{i \mu \cdot \mathbf{k}} \right] \hat{B}_{ns} \phi \left( \frac{r}{n} \right). \quad (4-20)
\end{align*}
\]

The function \( B_{ns}(k) \) is defined by

\[
B_{ns}(k) = 1 + \frac{2}{3 \phi_{ns}(0)} \int dr \mathcal{E}_E^{(1)}(0, r) \frac{d}{dr} \phi_{ns}(r), \quad (4-21)
\]

where \( \mathcal{E}_E^{(1)}(0, r) \) is the p-wave contribution to the partial-wave
expansion of the approximate Green's function \( g_i^{(0, r)} \), \( g_i^{(0, r)} \)
= \( g_i^{(0, r)} + g_i^{(1, r)} \). The transition rate is calculated
as before, with the result

\[
\frac{dw_{ns}}{w_K} = \left( \frac{dw_{ns}}{w_K} \right)_{CF} \frac{\phi_{ns}(0)}{\phi_{is}(0)} \frac{k(\mathbf{q}_{ns} - \mathbf{k})^2}{q_{is}^2} R_{ns} dk. \quad (4-22)
\]

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The correction factor $R_{ns}(k)$, which describes the modification of the Coulomb-free result brought about by inclusion of the most important relativistic and Coulomb effects, is defined by

$$R_{ns}(k) = (1/2)(1 + B_{ns}^2).$$  \hspace{1cm} (4-23)

The evaluation of the functions $B_{ns}(k)$ has been described in great detail by Glauber and Martin (1956). Here we quote only the final results,

$$B_{1s}(k) = 1 - \frac{4}{3} \left( \frac{\eta_1^2}{(1 + \eta_1^2)} \right) \left\{ 1 + \frac{\eta_1^2}{(1 - \eta_1^2)} \left[ 2K(\lambda_1) - 1 \right] \right\},$$  \hspace{1cm} (4-24)

with $\lambda_1 = (1 - \eta_1)/(1 + \eta_1)$, and

$$B_{2s}(k) = 1 - \frac{\eta_2^2}{(1 + \eta_2^2/4)} \left\{ \frac{4}{3} + \frac{5}{6} \eta_2^2 \right\} - \frac{\eta_2^2}{(1 - \eta_2^2/4)^2} \times \left[ \frac{8}{3} (1 - \eta_2^2) K(\lambda_2) - 3 - \eta_2^2 + \frac{5}{4} \eta_2^2 \right],$$  \hspace{1cm} (4-25)

with $\lambda_2 = (2 - \eta_2)/(2 + \eta_2)$. The function $K(\lambda)$ is

$$K(\lambda) = \int_0^1 \frac{\lambda x^{\lambda-1}}{(1 + \lambda x)} \, dx.$$  \hspace{1cm} (4-26)

For the purpose of evaluation, $K(\lambda)$ can be represented conveniently by the rapidly converging series expansion

$$K(\lambda) = \ln(1 + \lambda) - \sum_{j=1}^{\infty} \frac{(\lambda)^j}{j(j - 1)}.\hspace{1cm} (4-27)$$

In arriving at these final results, advantage has been taken of the fact that $E$ may be set equal to one in the correction term,
so that $\eta = 2\alpha/\mu$. Consequently, the two parameters $\lambda$ and $k$, upon which the functions $B_{1s}$ depend, enter only in the single combination $\eta$ which, in the present approximation, is given by

$$
\eta = (1+k/B_{1s})^{-1/2}
$$

for $1s$-state capture and by

$$
\eta = (1/k + B_{1s})^{1/2}
$$

for $2s$-state capture. Here, $B_{1s}$ is the $1s$-state binding energy. This simplification greatly facilitates tabulation of final results.

With the aid of Eq. (4-27), we have evaluated Eq. (4-24) and Eq. (4-25) numerically (Tables 4.1 and 4.2). Although for energies not greatly exceeding the binding energy, $B_{1s}(k)$ increases quite rapidly from its value of zero at $k=0$, the function approaches its asymptotic value of unity quite slowly. The correction factor $R_{1s}(k)$ therefore remains substantially less than unity, even at energies very much larger than the binding energy. Like $R_{1s}(k)$, $R_{2s}(k)$ also slowly approaches unity for large $k$.

Unlike $B_{1s}(k)$, however, $B_{2s}(k)$ does not go to zero as $k$ approaches zero; rather, as may be shown analytically, $B_{2s}(0) = 3/2$.

The functions $B_{n\lambda}(k)$ for $n \geq 3$ can be evaluated similarly. However, the contributions to radiative electron capture from $ns$ states with $n \geq 3$ can usually be neglected entirely, compared with contributions from $1s$ and $2s$ states. For example, according to the above results, the $3s$-state intensity is only $1/4$ of the $1s$-state intensity; when screening-effects are taken into account, its contribution is reduced even more.

(ii) $p$-state radiative capture: From the calculations of Morrison and Schiff (1940) it can be concluded that the $p$-state
capture contribution to the IB spectrum is negligibly small for $k \gg Z \alpha$. As the calculations of Glauber and Martin (1956) bear out, however, the p-state intensity becomes quite appreciable for $k \leq Z \alpha$ and indeed exceeds the s-state spectrum over a large part of this range. Discussion of p-state radiative capture can therefore be restricted to photon energies $k \leq Z \alpha$. In this energy region, the transition matrix element can be reduced to

$$M_{np} = -2i \alpha \gamma C (2n/k)^{1/2} N_n \phi'_0(0) \int dr \mu \Phi_{np}(r) e^{-\mu} \phi_{np}(r)$$

when terms of order $Z \alpha$ are neglected. It is clear that the IB radiation associated with p-state capture is distributed isotropically.

Since the three np-state wave functions transform like the components of a vector under rotations, one of them can conveniently be chosen to be the component in the direction of $e$. The remaining two-component states then do not contribute to the matrix element, and a single calculation takes into account the contributions from all three magnetic substates. On this basis, the matrix element can be written

$$M_{np} = -2i \alpha \gamma C (Z \alpha/k)^{1/2} N_n \phi'_0(0) \chi_{np} Q_{np}$$

where $\chi_{np}$ is the spin part of the np-state wave function, and the integral $Q_{np}(k)$ is

$$Q_{np}(k) = (2n/Z \alpha)^{1/2} \int dr \mu \Phi_{np}(r) e^{-\mu} \phi_{np}(r).$$
The transition rate is calculated as before, with the result

$$\frac{d\omega_{np}}{w_k} = \frac{k}{\omega_{ns}^2} \left[ Q_{np}(k) \right]^2 \frac{(q_{np} - k)^2}{q_{1s}^2} \text{d}k. \quad (4.31)$$

Evaluation of the integrals $Q_{np}(k)$ is similar to that of $E_{ns}(k)$ and has also been described in detail by Glauber and Martin (1956). The final results are

$$Q_{2p}(k) = \frac{\eta_2^2}{4(1-\eta_2/2)^2} \left[ 1 + (2/3) \eta_2 - (7/12) \eta_2^2 + (4/3) \eta_2^2 \right] \quad (4.32)$$

and

$$Q_{3p}(k) = \frac{\eta_3^2}{27} \left[ \frac{1}{(1-\eta_3/3)^2} \right] \left\{ \left[ (1-\eta_3/3) \right]^2 - 2(\eta_3/3)^2 + 2(\eta_3/3)^3 \right\} \quad (4.33)$$

where \( \eta_3 = (1/9 + k/\lambda_{1s})^{-1/2} \), all other quantities having been defined previously. Evaluation of Eqs. (4.32) and (4.33) yields the results shown in Table 4.3.

(iii) Results. To illustrate the results of the theory of Glauber and Martin (1956), the predicted spectra associated with $1s^{-}$, $2s^{-}$, $2p^{-}$, and $3p$-state radiative capture in $^{55}$Fe have been plotted in Fig. 4-2. As stated, terms of order $2\alpha$ were neglected, introducing an error of ~20% for $^{55}$Fe. It is evident from Fig. 4-2 that the s-state spectra do not differ greatly in form from the simple $k(q_{ns} - k)^2$ shape predicted by the Coulomb-free theory. Figure 4-2 also shows the existence of very intense p-state spectra at low photon energies. Indeed, p-state contributions to
the IB spectrum become more dominant with increasing charge and decreasing available energy.

For states of still higher orbital angular momentum, the radiative capture probability is expected to be much smaller than for capture from s or p states, because the probability of finding the electron in the neighborhood of the nucleus is smaller and the radiation emitted is of a higher multipole order than the predominantly M1 and E1 radiation associated with s- and p-state radiative capture, respectively. Indeed, within the framework defined by the approximations used in treating p-state capture, the transition amplitude for radiative capture from a state of orbital angular momentum \( \ell > 1 \) vanishes.

Relativistic calculations. The preceding calculations were intended to provide results with a relative accuracy of order 2%. To achieve even this level of accuracy requires that some consideration be given to relativistic effects when treating radiative capture from s states. The importance of relativistic effects in s-state capture, even for moderately light nuclei, is primarily due to the fact that such transitions involve a spin flip, a process which results in large photon energies, and hence, in a relativistic recoil by the electron. Furthermore, a nonrelativistic calculation does not take account of paths that involve virtual positron emission and neglects electron capture through intermediate p states, a path made possible by spin-orbit coupling.

The results described above are usually adequate to determine the IB spectra of moderately light nuclei for photon energies that
are small compared with the electron rest energy. For heavy nuclei or large photon energies, these results are wholly inadequate.

Martin and Glauber (1958) therefore developed a more general theory, taking full account of relativistic and Coulomb effects. The nonrelativistic results indicate that relativistic and Coulomb effects to all orders in $Z_e$ are most important in radiative capture from ls states, hence Martin and Glauber (1958) applied their full theory to this specific calculation.

It should be noted that Yukawa (1956) has also attempted a fully relativistic calculation of the $K$-capture $1B$ spectrum. Yukawa found it necessary, however, to introduce an approximation in constructing a usable form for the relativistic Coulomb Green’s function; it is not entirely clear how reliable this approximation is. The results of Yukawa (1956) are at least as complicated as those of Martin and Glauber (1958) and have the serious drawback of being inapplicable to heavy nuclei. For these reasons, we do not discuss Yukawa’s calculations further.

(i) 1s-state radiative capture. The starting point for the fully relativistic calculations of Martin and Glauber (1958) is the general expression (4.11) for the allowed-transition matrix element. To evaluate this matrix element exactly within the one-electron Coulomb approximation, appropriate forms for $\hat{a}_{ls}$ and $\hat{g}_E(N,N', \tau)$ must first be introduced. For $\hat{a}_{ls}$, the usual ground-state solution of the Dirac equation for an electron moving in the Coulomb field of a nuclear charge $Z_e$ is chosen. For the exact second-order Green’s function $\hat{g}_E(N,N', \tau)$, Martin and Glauber (1958)
constructed an eigenfunction expansion from the solutions of Eq.
(4-8). The smallness of the nuclear radius \((2\rho r_N \approx 10^{-3})\) allows
some simplification. The region occupied by the nucleus may be
safely neglected in integrating over \(r \), and those functions in
the Green's-function expansion which depend on \(r_N \) can be replaced
by the first term in their power-series expansion. The errors
associated with the use of this simplified form of the exact
Green's function are expected to be no greater than \(\approx 10^{-3}\).

Using the above representations, Martin and Glauber (1958)
calculated the transition matrix element for allowed radiative
\(K\) capture without further approximations,

\[
M_{_1s} = \frac{i e C \lambda}{k} (\pi/k)^{1/2} J_{\lambda} N_{\lambda} \int_0^\infty \Phi^*(0) \Gamma_{\lambda} [A_{_1s} \Sigma_{\lambda} e^* x_{k+i k B} B_{_1s} e^*] \chi_{_1s}^* \tag{4-34}
\]

The particular angular-momentum substate of the initial \(K\) electron
is represented through the spin function \(\chi_{_1s}^* = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \), where \(\chi_{_1s}^*\) are the
usual two-component Pauli spinors, and the integrals \(A_{_1s}(k)\) and
\(B_{_1s}(k)\) are defined by

\[
A_{_1s}(k) = \frac{\lambda_{_1s}(1+ \lambda_{_1s})^2}{\Gamma(2\lambda_{_1s}+1)\mu} \int_0^\infty dr \int_0^\infty ds 
\]

\[
\left\{ j_0(kr) \left[ 1+ \frac{a^2}{3(\lambda_{_1s}+1)^2} \right] - \frac{j_1(kr)a^2}{kr(\lambda_{_1s}+1)^2} - \frac{j_2(kr)a^2}{3(\lambda_{_1s}+1)^2} \right\} 
\times \begin{pmatrix} \eta_{_1s+1} \mu \\ -\eta_{_1s+1} \end{pmatrix} \begin{pmatrix} \eta_{_1s}+1 \\ -(2s+1)\mu \end{pmatrix} e^{i r} e^{i \lambda_{_1s} \rho} , \tag{4-35a}
\]
In Eqs. (4-35), the previous definitions of \( \mu \) and \( \eta \) have been retained, and we have \( \alpha = \frac{Z}{2} \) and \( \lambda_1 = \frac{1}{2} \frac{1}{2} = \frac{1}{2} \).

The energy spectrum of the IB, calculated as before, is

\[
\frac{\text{d}w_{1s}}{\text{d}w_k} = \frac{\alpha}{n} \frac{k(q_{1s} - k)^2}{q_{1s}^2} R_{1s} \text{ dk}.
\]

This expression is the same as Eq. (4-22) for \( 1s \) capture, except that \( R_{1s}(k) \) is defined by

\[
R_{1s}(k) = \frac{1}{2} \left( A_{1s} + B_{1s} \right),
\]

with \( A_{1s}(k) \) and \( B_{1s}(k) \) given by Eqs. (4-35). Unfortunately, the integrals appearing in Eqs. (4-35) cannot be evaluated exactly analytically in closed form and depend separately on \( Z \) and \( k \), rather than on the single combined parameter \( k/(2\alpha)^2 \) as do the integrals \( B_{ns} \) and \( Q_{np} \) discussed earlier.

A number of limited and, in most cases, approximate analytic results for \( A_{1s} \) and \( B_{1s} \) are reported by Martin and Glauber (1958).
For example, Eqs. (4-35) can be simplified, transformed, and expanded if one neglects terms of order \( (Z\alpha)^2 \) or smaller and the remaining contribution to \( B_{ls}(k) \) from the \( j_2(kr) \) term. The results are

\[
A_{ls}(k) = \sum_{n=0}^{\infty} \frac{1}{n(n-\eta)} \left( \frac{a+ik}{a+ik+\mu} \right)^n, \tag{4-38a}
\]

\[
B_{ls}(k) = A_{ls}(k)(1+ \frac{2}{3} \frac{a^2}{k}) - \frac{2}{3} \left( \frac{a}{k} \right)^2 \zeta, \tag{4-38b}
\]

where

\[
\zeta = 2 \left[ \ln \left( \frac{1+ik}{2\mu} \right) + \sum_{n=1}^{\infty} \frac{1}{n(n-\eta)} \left( \frac{a+ik}{a+ik+\mu} \right)^n \right]. \tag{4-39}
\]

Because of the underlying approximations, these expressions are expected to hold well only at low photon energies and for elements which are not too heavy.

For \( k > 1 \), it is feasible to expand the Green's function and the initial-state wave function in powers of \( Z\alpha \). Carried to first order in \( Z\alpha \), such expansions yield

\[
A_{ls}(k) = 1 - Z\alpha \left\{ \mu/k + 2(1 - \frac{1}{k})\tan^{-1}(k/\mu) \right\}, \tag{4-40a}
\]

\[
B_{ls}(k) = 1 - Z\alpha \left\{ \mu/k \left( 1 + \frac{1}{k} \right) + 2(1 - \frac{1}{k^2})\tan^{-1}(k/\mu) \right\}. \tag{4-40b}
\]

For three particular photon energies, more accurate results can easily be obtained because of special circumstances which simplify the calculation in each case. In the neighborhood of \( k = 0 \), \( A_{ls} \) and \( B_{ls} \) are given, exact to all orders in \( Z\alpha \), by
Als(k) = \frac{(2\lambda_1 + 1)}{3} (1-k), \hspace{1cm} (4-41a)

Bls(k) = 0 + \Phi(k/2\alpha). \hspace{1cm} (4-41b)

The integrals can be evaluated conveniently to second order in Z\alpha for k=\lambda_1 (m=0),

Als(k) = 1 - 2\alpha + \pi(2\alpha)^2/4, \hspace{1cm} (4-42a)

Bls(k) = 1 - 2\alpha + (4-\pi/2)(2\alpha)^2; \hspace{1cm} (4-42b)

and for k=1+\lambda_1 (m=0),

Als(k) = 1 - \pi\alpha/2 + 2(2\alpha)^2, \hspace{1cm} (4-43a)

Bls(k) = 1 - 3\alpha\pi/4 + 9(2\alpha)^2/2. \hspace{1cm} (4-43b)

These approximate results are expected to be fairly reliable for the lighter elements. In general, however, it is necessary to resort to numerical procedures. The above results are still of interest, though, since they provide a valuable check on the accuracy of numerical computations.

A relatively simple procedure for obtaining exact numerical results for Als and Bls for arbitrary k and Z has been reported by Intemann (1971). The integration over r in Eqs. (4-35) is performed first, then the change of variable x=s/(1+s) is made in the remaining integrals. After algebraic reduction, one finds
\[ A_{1s}(k) = C \int_{0}^{1} dx \, x^{-\eta+\lambda_{1}-1} f_A(x), \]  

\[ B_{1s}(k) = \frac{C}{2k(1-\lambda_{1})} \int_{0}^{1} dx \, x^{-\eta+\lambda_{1}-1} f_B(x), \]

where \( C = (2\mu)^{2\lambda_{1}-1} / [\lambda_{1}(2\lambda_{1}-1)k^{2}] \). To define \( f_A \) and \( f_B \), it is convenient to introduce the definitions

\[ \Sigma = k^{2} + (\mu a)^{2}, \quad s = \Sigma + \delta x^{2}, \]

\[ \epsilon = 2(\mu^{2} - a^{2} - k^{2}), \quad \sigma = a^{2}(1-x)/(1-x), \]

\[ \delta = k^{2} + (\mu - a)^{2}, \quad \theta = \tan^{-1}(k/c), \]

whence \( f_A \) and \( f_B \) can be written

\[ f_A(x) = \frac{[2k\lambda_{1}\cos(2\lambda_{1}\theta) - \sigma^{2}\sin(2\lambda_{1}\theta)]}{s^{\lambda_{1}}}, \]  

\[ f_B(x) = (2k\lambda_{1}\sigma[\sigma - 2a^{2} + (1-\lambda_{1})k] \cos(2\lambda_{1}\theta)) \]

\[ + \left\{ k^{2}(2\lambda_{1}-1)[k(\lambda_{1}-1)-2a^{2}+\sigma] + \sigma^{2}[2a^{2}-k(1-\lambda_{1})-\sigma] \right\} \sin(2\lambda_{1}\theta) / s^{\lambda_{1}}. \]

Now \( f_A \) and \( f_B \) are very slowly varying functions of \( x \) over the entire range of integration, for all physical values of \( k \) and \( \lambda_{1} \) of interest. After an integration by parts to remove the weak singularity in each of the integrands at \( x=0 \), the remaining integrals which appear in \( A_{1s} \) and \( B_{1s} \) thus can easily be evaluated.
in this manner for several nuclides of interest, are displayed in Fig. 4-3.

It is considerably easier to evaluate $A_{1s}$ and $B_{1s}$ by means of the low-k approximation [Eqs. (4-38) and (4-39)] or the high-k approximation [Eqs. (4-40)], rather than by employing the exact results [Eqs. (4-44)]. Therefore it is of interest to compare the functions $R_{1s}(k)$ obtained in these three ways, in order to assess the circumstances under which either of the approximate results can be employed without significant error. We have evaluated $R_{1s}(k)$ exactly and in the high- and low-k approximations for three very different values of $Z$. The results, shown in Fig. 4-4, are indistinguishable for very small $Z$ over almost the entire energy range. For intermediate $Z$, the low-k approximation fits the exact curve quite well, even in the high-energy region where it does better than the high-k approximation. For large $Z$, neither approximation fits the exact result very well, and both approximations are totally wrong in their description of the low-energy behavior of $R_{1s}(k)$.

To compare the various calculations and indicate the importance of relativistic and Coulomb effects, we have plotted in Fig. 4-5 the $1s$-state radiative capture spectra predicted for the moderately light nucleus $^{55}$Fe. It is evident from Fig. 4-5 that the shape of the $1s$ spectrum is not substantially altered by the inclusion of relativistic and Coulomb effects, but that the overall intensity experiences a very significant reduction. As is to
exact calculation (ME) and that in which terms of order \( \alpha \) are neglected (GM) agree fairly well. There will be no such agreement for heavy nuclei or for photon energies \( kT \).

(ii) \( L \) and \( M \)-shell radiative capture. Although Martin and Glauber (1958) limited their fully relativistic calculations to \( ls \)-state capture, their relativistic theory provides an equally valid basis for describing radiative capture from an arbitrary atomic shell. The general results of such a calculation are given by Zon (1971). The complexity of the expressions has precluded the derivation of analytical results; not even approximate results have been derived in which only terms through first order in \( \alpha \) are retained. Zon (1971) does, however, report the construction of a computer program which permits numerical evaluation of the amplitude for radiative capture from the \( L \) and \( M \) shells, although few details are given and the only spectra reported in Zon's paper are those for \( ^{165}\text{Er} \) (Fig. 4-6).

Two general features of the \( ^{165}\text{Er} \) spectra are worth noting since they undoubtedly will be exhibited by the radiative capture spectra of other nuclei as well. A resonance in the \( 2s \)-state capture spectrum appears which is associated with a forbidden \( 2s-1s \) atomic transition. This resonance is quite sharp and therefore modifies the result of Glauber and Martin (1956) only in the binding-energy region. Elsewhere, the results of Zon (1971) and of Glauber and Martin (1956) are indistinguishable. Also to be noted are the modifications of the \( p \)-state spectra brought about by the inclusion of all relativistic and Coulomb effects. While
these modifications appear to be only slight for capture from 3p states, they are quite important for the 2p-state spectrum (at least for heavy nuclei). In the case of \(^{165}\)Er, they cause a reduction by a factor of \(\approx 2\) in the overall intensity of the 2p-state spectrum. There is, however, no appreciable change in the form of the p-state capture energy distributions.

Some years ago it was suggested by Koh et al. (1962), and again by Koh (1965), that the IB spectrum possesses a cusp-shaped irregularity in the neighborhood of the positron threshold. To confirm this idea, Zon and Rapoport (1968) carried out extensive calculations. Their results, accurate to order \((Za)^2\), show however that the form factor for radiative K capture varies continuously in this region, and thus, there is no such anomaly in the predicted spectrum at this level of accuracy.

Influence of Uncaptured Atomic Electrons

In all of the foregoing calculations, only the electron which undergoes radiative capture is considered, and the presence of all other atomic electrons has been ignored. We now consider how, and to what extent, the one-electron results are modified when the presence of the remaining atomic electrons is taken into account.

Screening corrections. Screening by the remaining electrons affects the amplitude for radiative capture both by altering the initial configuration of the electron to be captured and by altering the probability amplitude for an electron to reach the nucleus.
after the virtual emission of a photon. To analyze these effects most simply, Martin and Glauber (1958) employed an independent-particle model in which the stationary states of the individual electrons are determined as the self-consistent-field solutions for the full many-body atomic Hamiltonian. In this approximation, no further account of the remaining electrons needs to be taken when the radiative transition probability for a single electron is calculated.

By far the more important effect of screening is the modification of the wave function that describes the initial electronic state. This modification is quite similar to that which occurs in ordinary electron capture, except that the effective size of the region from which capture can occur is somewhat larger. In ordinary electron capture, this region is determined by the nuclear radius, while in radiative electron capture it is determined by the range of the Green's function. For photon energies of greatest practical interest, above the binding energy of the initial electron and below the threshold for positron production, the range of the Green's function is of the order of the electron's Compton wavelength. While it is much larger than the nuclear radius, this range is still very small on an atomic scale. Thus, it is argued by Martin and Glauber (1958) that a simple and seemingly reasonable procedure for taking into account screening effects on the initial state of an electron undergoing radiative capture is to multiply the unscreened results for the radiative capture probability amplitude by the ratio of the screened to
unscreened initial-state wave functions, evaluated in the neighborhood of the origin.

The second effect of screening, the alteration in the structure of the Green's function, is expected to be quite small; this can be understood qualitatively from the following considerations (Martin and Glauber, 1958). Over the relatively small region defined by the range of the Green's function, the electron field is well approximated by the nuclear Coulomb field. Indeed, if the electronic charge cloud associated with the remaining atomic electrons did not penetrate this region at all, its external presence would simply result in a shift in the zero of energy and thus produce no physical effects. For all but the lowest-energy photons, the range of the Green's function is so small that penetration of the electronic charge cloud into the region defined by this range is not expected to be appreciable, and therefore no significant modification in the Green's function is expected. It should be emphasized, however, that this reasoning is not valid for photon energies near the binding energy, where the range of the Green's function becomes quite large and a more elaborate treatment of screening is required.

To establish in quantitative terms the accuracy of the simple approximation scheme of Martin and Glauber (1958), these authors carried out more extensive calculations in which the screened Coulomb potential was approximated by a Hulthen potential. The results of these calculations indicate that the above conclusions are quite well-founded. In particular, Martin and Glauber (1958).
calculated the screening corrections for the 2p state of Fe to lowest order in the Hulthen parameter. The results were compared with unscreened results multiplied by the ratio of the screened to unscreened probability densities at the origin. At a photon energy equal to the K-shell binding energy in Fe, the difference between these two results was found to be \( \approx 20\% \) (i.e., of order \( Z_\alpha \)), while at a photon energy three times as large the difference is only \( \approx 2\% \). Thus it appears that, except at very low photon energies (in the immediate neighborhood of the K-shell binding energy), screening effects can be taken into account satisfactorily by simply multiplying the unscreened rate for radiative capture from the state \( \alpha \) by the screening factor

\[
S_\alpha = \frac{\phi^{SC}_\alpha(R_N)^2}{|\phi_\alpha(R_N)|^2},
\]  

(4-46)

where \( R_N \) is the nuclear radius.

From results of Brysk and Rose (1958) and available Hartree calculations, Martin and Glauber (1958) have constructed a graph of \( S_\alpha \) vs. \( Z \) for initial states of interest (see Fig. 4-7). It appears that the intensities of the IB spectra for radiative capture from the L shell are considerably reduced by screening effects and those for radiative capture from higher shells become insignificant.

If the intensities of the various IB spectra are normalized to a single K-capture event, or to a single EC event, then only the ratios \( S_\alpha/S_1 \) appear in the final formulas. To evaluate these ratios for the most important case, the L shell, results of
Sec. 2.2.2 can be used when a high degree of accuracy is desired. The ratios are

\[
\frac{S_{2s}}{S_{1s}} = \left| \frac{G_K}{G_L} \frac{g_{L_1}}{g_K} \right|^2
\]  

(4-47a)

and

\[
\frac{S_{2p}}{S_{1s}} = \left| \frac{G_K}{F_{L_2}} \frac{f_{L_2}}{g_K} \right|^2
\]  

(4-47b)

Here, \(G_K\), \(G_{L_1}\), and \(F_{L_2}\) are the large components of the unscreened Dirac wave function for the 1s, 2s, and 2p states, respectively, evaluated at the nuclear radius of a hydrogenic atom. The large components are denoted by \(g_K\), \(g_{L_1}\), and \(f_{L_2}\), respectively, when the effects of screening are included.

Plots of \(G_K\), \(G_{L_1}\), and \(F_{L_2}\) for a point nucleus and corrections for finite nuclear size are given by Brysk and Rose (1958) (finite-nuclear-size corrections to the L-shell screening ratios are always <1%). As discussed in Sec. 2.2.2, the ratios \((g_{L_1}/g_K)^2\) and \((f_{L_2}/g_{L_1})^2\) have been calculated by several authors; the most reliable results being those displayed in Table 2.9. These ratios were computed with a relativistic Hartree-Fock self-consistent potential with allowance for finite nuclear size.

The procedure described above is but one possible way in which screening effects can be treated. Alternatively, Zon (1971) has included screening effects by employing relativistic initial-state Coulomb wave functions with effective charges. These effective charges, as suggested by work on internal conversion, were taken to be \(Z_{eff} = Z - \sigma\), with \(\sigma_K = 0.3\), \(\sigma_L = 3.5\), and \(\sigma_M = 5\). Zon (197
has carried out several numerical calculations but does not compare his results with those obtained by the simpler procedure of Martin and Glauber (1958).

**Exchange and overlap corrections.** All results described so far, including screening corrections, are based on independent-particle approximations and take no account of exchange and overlap effects which result from the many-particle nature of the atom (see also Sec. 2.5). Corrections for such effects have been applied to the Martin-Glauber theory by Persson and Koonin (1972), using a procedure analogous to that applied by Bahcall (1962) to L/K electron-capture ratios. The calculations of Persson and Koonin (1972) deal specifically with the electron-capturing nucleus $^7\text{Be}$, but are easily generalized.

It is found that, for EC-decays of $^7\text{Be}$ to the 477-keV state of $^7\text{Li}$, the predominant effect of exchange and overlap corrections is to increase the ratio of 2s-state radiative capture to 1s-state radiative capture $\omega_{2s}/\omega_{1s}$ by a factor of 2.9. The ratio of the 1s-state radiative capture rate to the total (K+L) nonradiative capture rate $\omega_{1s}/(\omega_{K}+\omega_{L})$ is decreased by 7%. However, the net effect on the ratio $(\omega_{1s}+\omega_{2s})/(\omega_{K}+\omega_{L})$ is found to be negligibly small (<1%). Changes in the shape of the FB spectrum at energies above 50 keV are found to be negligible.

Calculations of overlap and exchange effects in radiative EC of $^{51}\text{Cr}$ and $^{54}\text{Mn}$ are reported by Koonin and Persson (1972), who find that $\omega_{2s}/\omega_{1s}$ is increased by 15% over the Martin-Glauber predictions. This increase is cancelled, however, by a similar
increase in the corresponding ratio for nonradiative capture, so
that the correction to the ratio \((w_{1s} + w_{2s})/(w_K + w_L)\) is again found
to be insignificant (<0.5%).

4.1.3. IB Correlation Effects in Allowed Transitions

With the discovery of parity nonconservation in weak inter-
actions, interest in radiative-electron capture shifted to studies
of those correlation effects whose existence requires a parity-
violating interaction. Calculations on such phenomena were re-
ported by Cutkosky (1957), Koh et al. (1957, 1962),
Berestetskii (1958), Martin and Glauber (1958), Gandelsman (1959),
Bloom and Uretsky (1960), and Tihomirov and Kamin (1960).

Cutkosky (1957) first showed that a two-component neutrino-
theory predicts that IB radiation will be circularly polarized.
Terms of order \(Z^a\) were neglected in Cutkosky's calculations, but
a determination of the polarization of the IB associated with K
capture, valid to all orders in \(Z^a\), was reported shortly there-
after by Martin and Glauber (1958). Only the polarization of the
1s-state contribution to the IB spectrum is considered in these
papers, yet it is evident from the results of Sec. 4.1.2. that
at low photon energies the contributions from L- and M-shell
radiative capture must also be taken into account. For allowed
transitions, this is easily accomplished using the theory of
Glauber and Martin (1958). More elaborate calculations, based
on a generalization of the Martin-Glauber theory, are reported
by Zon (1971), who lists numerical results for \(^{37}\)Ar.
The parity-nonconserving character of the weak interaction is also responsible for the existence of an anisotropy in the angular distribution of the EB radiation emitted from oriented nuclei, as may be inferred from the work of Cutkosky (1957). This makes EB angular-distribution studies of interest as a potential source of information on nuclear spin changes and the relative magnitudes of the EC nuclear matrix elements. The angular distribution of the EB emitted from oriented nuclei during K capture was first calculated by Timashev and Kaminskii (1960) and by Koh et al. (1962), assuming a nonrelativistic description of the electronic motion and neglecting all Coulomb effects on the intermediate states of the electron. The results of these calculations are quite simple, but they have proved inadequate to explain the experimental data at low photon energies, where both intermediate-state Coulomb effects and the contributions from L- and M-shell radiative capture become important. More exact and extensive calculations, based on the work of Glauber and Martin, have been reported by Intemann (1971) and by Zen (1971).

While the existence of the EB correlation effects described above depends on the parity-nonconserving property of the weak interaction, a variety of other correlation phenomena exist which could arise even if parity were conserved. (From the point of view of testing weak-interaction theory, these phenomena are of little interest, but they can provide information on nuclear structure.) In particular, Koh et al. (1957, 1962) have studied
the correlations between the direction of nuclear spin, the momentum of the IB photon, and the momentum of a subsequent nuclear \(\gamma\) ray, and have reported detailed results on the correlation between the directions of the IB photon and the nuclear \(\gamma\) ray. These calculations were, however, carried out for allowed and first-forbidden transitions and neglect Coulomb effects on the intermediate electron states; thus they are limited to high photon energies. More extensive calculations of this correlation function, based on a generalization of the work of Martin and Glauber (1958), have been reported by Zon (1971). This latter work includes a determination of the correlation between the directions of the IB photon and a subsequently emitted atomic x ray.

**IB Circular Polarization**

The polarization \(P(k)\) of the IB accompanying electron capture from the state \(\alpha\) is defined as the difference in the intensities of the right- and left-circularly polarized radiation, divided by their sum:

\[
P_{\alpha}(k) = \frac{dw_{\alpha}^{+1} - dw_{\alpha}^{-1}}{dw_{\alpha}^{+1} + dw_{\alpha}^{-1}}.
\]  

For \(ls\)-state radiative capture, the required expressions for the intensity of the polarized radiation are obtained from Eq. (4-34) by squaring and summing over all final states of the unobserved neutrino and over the spin states of the initial electron. The
result for randomly oriented nuclei is

\[ n_{ls}^s(k) \propto |A_{ls}(k) + sB_{ls}(k)|^2, \quad (s = \pm 1). \]

The polarization of the IE accompanying 1s-state capture is found to be

\[ P_{ls}(k) = \frac{(A_{ls} + B_{ls})^2 - (A_{ls} - B_{ls})^2}{(A_{ls} + B_{ls})^2 + (A_{ls} - B_{ls})^2} \frac{A_{ls} B_{ls}}{R_{ls}}. \quad (4-49) \]

At low photon energies, the 1s-state radiation is almost completely unpolarized since \( B_{ls}(k) \rightarrow 0 \) as \( k \rightarrow 0 \). At high energies, we have \( A_{ls} = B_{ls} = 1 \), neglecting terms of order \( Z \alpha \), and due to cancellation, \( P_{ls}(k) = 1 \) neglecting terms of order \( (Z \alpha)^2 \). More precisely, the high-energy form of \( P_{ls}(k) \) is

\[ P_{ls}(k) = 1 - (Z \alpha)^2 \left\{ \left( \frac{\mu}{k} \right)^2 + 2 \left( 1 - \frac{1}{k} \right) \right\} / 2k^2, \quad (4-50) \]

which follows from Eq. (4-49).

The polarization of the 2s-state radiation can be analyzed similarly, starting with Eq. (4-20). The final result has the same structure as Eq. (4-49) except that, in the approximation which underlies Eq. (4-20), \( A_{2s}(k) = 1 \). Thus, we have

\[ P_{2s}(k) = B_{2s}/R_{2s}. \quad (4-51) \]

While it is expected in the high-energy limit that \( P_{2s}(k) = 1 - (Z \alpha)^2 \), the results which follow from Eq. (4-20) are not sufficiently accurate to allow the determination of the coefficient of the \( (Z \alpha)^2 \) term. The low-energy limit of \( P_{2s}(k) \) is easily obtained, however,
by using the fact that $B_{2s}(0) = -3/2$. From this result it follows that $P_{2s}(0) = -12/13$.

To illustrate the above results, the functions $A_{ls}$, $B_{ls}$, $B_{2s}$ and the resulting polarization functions have been evaluated for two nuclei of interest, viz. $^{37}$Ar and $^{119}$Nb (Figs. 4-8 and 4-9).

It is evident from Eqs. (4-29) and (4-30) that, when terms of order $Z\alpha$ are neglected, p-state radiation should be completely unpolarized. At low photon energies, where the p-state spectra dominate, one therefore expects an even greater reduction of the $\Pi_B$ polarization than predicted by the function $P_{ls}(k)$. The overall polarization of the total $\Pi_B$ radiation accompanying electron capture is

$$P(k) = \frac{2}{\sum_{n=1}^{\infty} \sum_{\alpha} P_{ns}(k)dw_{\alpha}/\sum_{\alpha} \alpha}$$

where the sum on $\alpha$ extends over $1s$, $2s$, $2p$, and $3p$ states.

**Angular Distribution of $\Pi_B$ from Oriented Nuclei**

When the initial nuclei are aligned, it is convenient to represent each by its polarization vector $P_N = \langle J_1 | M | J_1 \rangle / J_4$, where $J$ is the angular-momentum operator and $J_1$, $M$ are the angular-momentum eigenvalues which label the initial nuclear state.

In this case, squaring Eq. (4-34) and summing over all final states of the neutrino, the spin states of the initial electron, and the final magnetic substates of the nucleus leads to the following result for $1s$-state radiative capture:
\[ \frac{d\omega}{d\Omega} \propto (A_{ls} + sB_{ls})^2 [1 + s a_k P_M \cos \theta]. \] (4-53)

Here, \( \theta \) is the angle between the vectors \( P_M \) and \( k \). The factor \( a_k \) vanishes for a pure Fermi transition, while for a pure Gamow-Teller transition we have

\[
a_k = \begin{cases} 
-\frac{J_f}{(J_f+1)} & \text{if } J_f = J_f + 1 \\
\frac{1}{(J_f+1)} & \text{if } J_f = J_f - 1 \\
1 & \text{if } J_f = J_f - 1,
\end{cases} \] (4-54)

where \( J_f \) is the angular-momentum quantum number of the final nuclear state. For transitions in which both allowed EC matrix elements are operative, \( a_k \) is given by

\[
a_k = \left[ \frac{\lambda^2 |R|}{(J_f+1)^2} + \frac{\lambda J_2(R^*\sigma R)}{[J_J(J_J+1)]^{1/2}} \right] (1 + \lambda^2 |R|^2)^{-1} \] (4-55)

with

\[ R = \langle f | \sigma | i \rangle / \langle f | 1 | i \rangle. \]

If the circular polarization \( s \) of the IB is measured, then the angular-distribution function has the simple form

\[
W_{ls}(\theta, s) = 1 + s a_k P_M \cos \theta, \] (4-56)

whence the shape of the angular distribution is seen to be independent of the energy of the IB photon.

If the photon polarization is not measured, Eq. (4-53) must be summed over \( s = \pm 1 \). This leads to an angular-distribution function of the form

\[
W_{ls}(\theta) = 1 + a_k (k) a_k P_M \cos \theta. \] (4-57)
The function \( a_{ls}(k) \) is defined by

\[
a_{ls}(k) = \frac{A_{ls} B_{ls}}{R_{ls}}
\]  

(4.53)

and is seen to be identical with the polarization function \( P_{ls}(k) \) discussed earlier. Indeed, this one function is sufficient to account for the IB energy dependence of all 1s-state capture correlations considered here.

The angular distribution of the IB radiation accompanying 2s-state capture can be determined in a similar manner, starting with Eq. (4-20). It is found that the distribution function \( W_{2s}(\theta, s) \) is identical with \( W_{ls}(\theta, s) \). The function \( W_{2s}(\theta) \) has the same general form as that for 1s-state capture, viz.,

\[
W_{2s}(\theta) = 1 + c_{2s}(k) a_{kp} \frac{P_{2s}}{2s} \cos \theta,
\]  

(4.59)

but \( c_{2s}(k) \) is defined as

\[
c_{2s}(k) = B_{2s}/2s,
\]  

(4.60)

describing the dependence of the angular-distribution function on the energy of the IB photon. Again we have \( c_{2s}(k) = P_{2s}(k) \).

With regard to \( np \)-state radiation, it has already been noted that the structure of Eq. (4-29) implies an isotropic distribution, i.e., \( a_{np}(k) = 0 \). This result is expected to be valid only to a relative accuracy of order \( 2\alpha \). Indeed, Zon (1971) reports that exact computer calculations for \( a_{np} \) show \( a_{2p} \) and \( a_{3p} \) to be small negative quantities.
The overall angular-distribution function is given by

\[ W(\theta) = \sum_{\alpha} s_{\alpha} dw_{\alpha} W(\theta) / \sum_{\alpha} s_{\alpha} dw_{\alpha} = 1 + A(k) s_p P_k \cos \theta, \quad (4-61) \]

with the overall asymmetry function \( A(k) \) defined by

\[ A(k) = \sum_{\beta} s_{\beta} dw_{\beta} / \sum_{\beta} s_{\beta} dw_{\beta} \quad (4-62) \]

In view of the equality of the asymmetry function and the polarization function for both s and p states, it follows that the overall asymmetry function \( A(k) \) is identical with the overall polarization function \( P(k) \) [Eq. (4-52)]

**Correlation of IB and Subsequent Nuclear \( \gamma \) Rays**

The simplest type of decay scheme for which the directional correlation between an IB photon and a subsequent nuclear \( \gamma \) ray can be studied is one in which the radiative capture transition leads to an excited nuclear state \( |N_f> \) from which there is a single \( \gamma \)-ray mode for deexcitation, leading to the final nuclear state \( |N_f> \). To determine the correlation between the directions of emission of the IB and \( \gamma \)-ray photons, a knowledge of the radiative capture matrix element must be combined with results from the theory of nuclear angular correlations (Frauenfelder and Steffen, 1966). The required calculation is straightforward but employs much mathematical machinery from the theory of angular momentum (Edmonds, 1960) and will not be described here. Such calculations were first reported by Gandelman (1959) for allowed transitions, and by Koh et al. (1962) for allowed and
first-forbidden transitions. Although Coulomb effects on the intermediate electron states are neglected in these calculations, Zon (1971) has reported results of much more extensive calculations based on a generalization of the Martin-Glauber theory to radiative capture from arbitrary shells and for any order of forbiddenness of the EC transition. Only for the case of allowed radiative K capture, however, has Zon's theory been worked out in complete detail, and we shall restrict our discussion to this particular case.

For allowed K-capture transitions, the radiative capture matrix element of Zon (1971) reduces to that of Martin and Glauber (1958). For this particular case, Zon's final results can be summarized as follows. For an \( I \) quantum of circular polarization \( s \) and a nuclear \( \gamma \) quantum of circular polarization \( t \), the directional correlation function is of the form:

\[
\Psi_\gamma(\theta, s, t) \propto \left( A_{1s} + sB_{1s} \right)^2 + \frac{1}{4} A_1(\mu_1 J_f J_f) b_k S(A_{1s} + sB_{1s})^2 \cos \theta, \quad (4-63)
\]

where the quantum numbers \( J_f \) and \( J_f \) refer to the angular momentum of the nuclear states \( \left| N_f^s \right\rangle \) and \( \left| N_f^\gamma \right\rangle \) respectively, and \( \theta \) is the angle between the directions of the two photons. The factor \( b_k \) vanishes for a pure Fermi transition, while for a pure Gamow-Teller transition it is

\[
b_k = \begin{cases} 
\left( \frac{J_f+1}{J_f+1} \right)^{1/2} & \text{if } J_f = J_f + 1 \\
\frac{1}{\left[ J_f(J_f+1) \right]^{1/2}} & \text{if } J_f = J_f \\
\frac{-1}{\left[ J_f(J_f+1) \right]^{1/2}} & \text{if } J_f = J_f - 1.
\end{cases} \quad (4-64)
\]
For transitions in which both EC matrix elements are operative, we have

\[ b_k = \left[ \frac{\lambda_2 |b_{jk}|^2}{[\ell_j (\ell_j + 1)]^{1/2}} + \lambda (\ell + R') \right] (1 + \lambda^2 |R|^2)^{-1}. \] (4-65)

The coefficient \( A_{\alpha} \), familiar from the theory of angular correlations, is defined by

\[ A_{\alpha}(LL'J_f f f) = \frac{F_1(\ell L, \ell L', J_f f f)}{\Gamma(2L + 1)(2L' + 1)(2J_f + 1)}, \] (4-66)

where the angular-momentum and parity quantum numbers \( L, \ell, L', \ell' \) characterize the multipolarities of the \( \gamma \) transition, and the ratio of the corresponding reduced matrix elements is

\[ \delta = \langle J_f f f \mid [L' \ell' \ell \mathcal{J}_f f f] \mid J_f f f \rangle. \] For pure multipole radiation, we have \( L' = L \) and \( \ell' = \ell \). The \( F \) coefficients are defined by

\[ F_1(\ell L, \ell L', J_f f f) = (-1)^{J_f + J_f - 1} \frac{\Gamma[(2L + 1)(2L' + 1)(2J_f + 1)]^{1/2}}{(2L + 1)(2L' + 1)(2J_f + 1)}, \] (4-67)

where the standard designations \( \{ \} \) and \( \{ \} \) indicate Wigner 3j and 6j symbols.

It is immediately apparent from the form of Eq. (4-63) that the circular polarization of the \( \gamma \)-ray photon must be measured if one is to observe any correlation between the directions of the two photons. If the circular polarization of the IB photon is also measured, then the directional correlation function is
\( \nu_{\gamma}(\theta, s, t) = 1 + \frac{t_{\text{INS}}}{\sqrt{2}} A_{\gamma}(\pi J, J) b_{\gamma} \cos \theta, \)  

independent of the IB-photon energy. If the polarization of the IB photon is not measured, Eq. (4-53) must be summed over \( s=\pm 1. \)

In this case, the directional correlation function is given by

\( \nu_{\gamma}(\theta, t) = 1 + \frac{t_{\text{INS}}}{\sqrt{2}} A_{\gamma}(\pi J, J) b_{\gamma} a_{\text{INS}}(k) \cos \theta, \)

and shows a dependence on the energy of the IB photon characterized by the asymmetry function \( a_{\text{INS}}(k) \) previously discussed.

The above results are exact, but in the derivation of the IB-\( \gamma \) directional-correlation functions it is assumed that no forces act on the nucleus while it is in the intermediate state \( |N_f^-\rangle. \) Generally, this assumption is not well satisfied, because the hole in the atomic shell produces strong magnetic and inhomogeneous electric fields at the nucleus, leading to a perturbation of the directional correlations.

**Correlations of IB and Succeeding Atomic X Rays**

The determination of the directional correlation function for an IB photon and a succeeding x-ray quantum requires a calculation which is essentially analogous to that of IB-photon-\( \gamma \)-ray directional correlations. Zon (1971) has carried out such a calculation and reported final formulas for the case of radiative \( \gamma \) capture. For allowed transitions, these results can be summarized as follows.

For an IB quantum of circular polarization \( s, \) and an atomic x-ray quantum of circular polarization \( t, \) the directional-
correlation function is of the form

\[ x(\theta, s, t) \propto (2J+1)(d_+ + s_+ s_0)^2 (-1)^J s_+ \frac{1}{2} \int \frac{d^3 k}{(2\pi)^3} \left( \frac{1}{\xi_{1s}} \right) \xi_{1s}^2 \cos \theta, \quad (4-70) \]

where \( \theta \) is the angle between the directions of the two photons and \( J=1/2, 3/2 \) is the angular momentum of the atomic electron which fills the hole in the K shell.

Further results are parallel to those for the nuclear \( \gamma \)-ray case. For example, it is evident from Eq. (4-70) that, in order to observe a directional correlation between the two photons, the circular polarization of the x-ray photon must be measured. If the circular polarization of the IB photon is also measured, we have

\[ x(\theta, s, t) = 1 + (\frac{-1}{2})(\frac{J+1/2}{2J+1}) \xi_{1s} t \cos \theta, \quad (4-71) \]

and the directional correlation shows no dependence on the energy of the IB photon. If the polarization of the IB photon is not measured, we have

\[ x(\theta, t) = 1 + (\frac{-1}{2})(\frac{J+1/2}{2J+1}) t \xi_{1s} \cos \theta, \quad (4-72) \]

and the correlation function again displays a dependence on the energy of the IB photon characterized by the asymmetry function \( \xi_{1s}(k) \).

4.1.4. IB Spectra and Correlation Effects in Forbidden Transitions

Early attempts to formulate a theory of IB for forbidden transitions were made by Cutkosky (1954), Turovtsev and Shapiro (1954),
Yukawa (1955), and Koh et al. (1957, 1962). Turovtshev and Shapiro calculated the radiative $\beta^-$-capture spectrum for first-forbidden transitions, assuming vector and tensor couplings, while Cutkosky derived the matrix element for radiative $\beta^-$ capture for arbitrary coupling, neglecting terms strictly of order $Z^2$ or smaller and terms contributing only to third- or higher-order transitions. Cutkosky's principal result was a theorem, often referred to as the "Cutkosky rule," which relates the spectra and angular correlations of the K-capture $\beta^-$ to the spectra and angular correlations of positrons. Basically, these calculations are extensions of the work of Morrison and Schiff (1940) to forbidden transitions. Yukawa (1956) made an attempt to include relativistic and Coulomb effects in the calculation of allowed and first-forbidden K-capture $\beta^-$ spectra. The formulas he obtained proved to be so complicated that this work has never led to useful results. Koh et al. (1957, 1962) first reported correlation studies for first-forbidden transitions; Coulomb effects were neglected in these calculations.

The modern theory of radiative electron capture in forbidden transitions is due to Zon and Rapoport (1968), who developed a generalization of the theory of Martin and Glauber (1958) to transitions of arbitrary order of forbiddenness. They also derived general formulas for the $\beta^-$ energy spectra. For K capture, detailed results were obtained. Zon (1971) developed this theory further for radiative capture from an arbitrary atomic shell, derived general formulas for various correlation and polarization
effects, and obtained detailed results for the case of capture.

The theory of Zon and Rapoport (1968) starts from Eqs. (4-9) and (4-10) for the radiative capture matrix element and transition rate. In order to evaluate the matrix element (4-9) exactly, including relativistic and Coulomb effects to all orders in $\hbar$, Zon and Rapoport first decompose and simplify it by introducing the irreducible tensor operators and the second-order Dirac-Coulomb Green's function of Martín and Clauser. This decomposition makes the angular-momentum dependence of the transition amplitude explicit. Integration over the angular coordinates is then completed through extensive use of the methods of the theory of angular momentum.

In evaluating the transition amplitude, Zon and Rapoport introduce the Konopinski-Uhlenbeck approximation,

$$\left\{ \begin{array}{c}
R_N \\
\lambda_L \\
\cdots \\
R_N \\
\end{array} \right\} \sim \left\{ \begin{array}{c}
R_N \\
\lambda_{L-N} \\
\cdots \\
R_N \\
\end{array} \right\} \left\{ \begin{array}{c}
\cdots \\
\cdots \\
\cdots \\
\cdots \\
\end{array} \right\},$$

where $\lambda_L = (l^2 - a^2)^{1/2}$, and the $\xi$ approximation which is based on the assumption $(\xi \omega)(L-N) < \omega$, a condition that is always well-satisfied when competing positron emission is not energetically possible. Under only these approximations, Zon and Rapoport obtain a general expression for the transition rate for radiative electron capture from an arbitrary shell. The form of the result reveals that for radiative electron capture in the $\xi$ approximation, just as in $\beta$ decay, nonunique forbidden spectra have
the same shape as the unique spectra of the next-lower order of forbidderness.

Only for $K$ capture do Zon and Rapoport carry their calculations to completion. For capture from higher shells, the theory is developed further by Zon (1971), but the resulting expressions prove to be too complicated to permit exact analytic evaluation or even the development of expressions that are correct to first order in $Z\alpha$. Indeed, the only detailed results which have so far been reported are those contained in the table of Zon (1973) for the $L$- and $K$-shell IB spectra associated with the first-forbidden unique transition in $^{41}$Ca; these results were derived through completely numerical procedures.

Zon and Rapoport's transition rate for radiative $K$ capture can be summarized as follows. Assuming the polarization of the IB radiation is not observed, the transition rate can be written

$$\omega_{1s} = \frac{8\pi}{\hbar} \left| \langle \varphi_{1s} | \varphi_{0s} \rangle \right|^2 \frac{k(q_{1s} - k)^2}{4\pi} F_{1s}(k) \frac{dk}{k}.$$  \hspace{1cm} (4-73)

The form factor $F_{1s}(k)$ is defined in terms of two corrections factors, $R_{1s}^{(1)}(k)$ and $R_{1s}^{(2)}(k)$, and the appropriate combination of nuclear matrix elements $N_{ij}^A$:

$$F_{1s}(k) = \sum_{Nj\lambda} (2\mu R_H) 2^{(\lambda - \lambda_1 - 2(N-1))} \left[ \frac{2^{N-1}(N-1)!}{(2j_\nu)!((2N-2)!)} \right] 2 N_{ij}^A \sigma_j^2$$

$$\times k_{2j_\nu}^{2j_\nu-1} k^{2N-2}, \hspace{1cm} (4-74)$$
Here, $R_N$ is the nuclear radius, and $l^*_N$ equals $(N^2-a^2)^{1/2}$. The quantities $E_{\ell}^{l^*_N}$, for all contributing values of $A$, $N$, and $j_N$, have been tabulated by Zon and Rapoport for up to third-forbidden transitions (Table 4.4). The correction factors $\tilde{\alpha}_{\ell_1}^{(n)}$ are defined in terms of the more fundamental quantities $\alpha_{\ell_1}^{(n)}$ and $\beta_{\ell_1}^{(n)}$, which are generalizations of the functions $\alpha_{\ell_1}$ and $\beta_{\ell_1}$ of Martin and Glauber (1958):

$$R_{\ell_1}^{(n)} = \frac{1}{4\pi} \left[ (\pi-1) \left( \frac{\alpha_{\ell_1}^{(n)}}{\beta_{\ell_1}^{(n)}} \right)^2 + \frac{\beta_{\ell_1}^{(n)} + \alpha_{\ell_1}^{(n)}}{2} \right] \left[ \frac{\alpha_{\ell_1}^{(n)}}{\beta_{\ell_1}^{(n)}} \right]^{(\pi-1)} \left( \frac{\alpha_{\ell_1}^{(n)}}{\beta_{\ell_1}^{(n)}} \right)^2 \right].$$

(4-75)

To specify the transition rate, formulas for $A$ and $B$ are required. Zon and Rapoport (1968) have developed exact general expressions for these functions, but these formulae contain a large number of integrals involving Whittaker functions, none of which can be evaluated exactly analytically.

For moderately light nuclei, it may be sufficient to expand the above-mentioned integrals in powers of $Z_N$ and thereby evaluate $A$ and $B$ to first order in $Z_N$. Such calculations are reported by Zon and Rapoport (1968) with the result

$$A_{\ell_1}^{(n)} B_{\ell_1}^{(n)} = 1 - Z_N \sum_{n=0}^{N} \left[ a_n \left( \frac{\mu}{k} \right) + b_n \tan^{-1} \left( \frac{\mu}{k} \right) \right] (1/k)^n.$$

(4-76)

The coefficients $a_n$ and $b_n$ are listed in Table 4.5. At the present time, Eq. (4-76) is the only formula available for the determination of $A$ and $B$. Unfortunately, even for light elements these formulas are not valid for low $k$. The nature of the expansion underlying Eq. (4-76) is such that these results are expected to break down for $k < Z_N$. For the special case $k = 0$, however, $A$ and $B$
can be evaluated exactly to all order in $Z$. The $k=0$ results, listed in the last column of Table 4.5, are valuable for estimating the low-energy behavior of $A$ and $B$ and to test numerical procedures for the exact evaluation of $A$ and $B$ for arbitrary $k$.

In examining the predicted $I_B$ spectra for $K$ capture in further detail, we restrict our discussion to first-forbidden transitions.

**Nonunique First-Forbidden Transitions**

In the $\xi$ approximation, the $K$-capture $I_B$ spectrum of a nonunique first-forbidden transition is predicted to have the allowed shape. Indeed, when the above results are evaluated for this case and normalized by the corresponding nonradiative $K$-capture rate, exactly the same result is obtained as for the allowed case,

$$\frac{\omega_{ls}}{I_B} = \frac{k(a_{ls} - k)^2}{q_{ls}^2} R_{ls}(1) \, dk,$$

where $R_{ls}(1)$ is easily identified as the function $R_{ls}$ of Martin and Glauber (1958), defined by Eq. (4-37) and displayed in Fig. 4-3.

**Unique First-Forbidden Transitions**

In unique forbidden transitions, only one nuclear matrix element contributes and the $\xi$ approximation becomes irrelevant. For unique first-forbidden electron capture, the radiative transition rate normalized by the corresponding nonradiative $K$-capture rate is
The factor in the brackets replaces the function $n_{ls}^{(1)}$ which appears in the allowed result.

While the factor $R_{ls}^{(1)}$ has been evaluated exactly numerically and in several analytic approximations, this has not been done for $R_{ls}^{(2)}$. The only available basis for the evaluation of $R_{ls}^{(2)}$ is Eq. (4-76), from which $R_{ls}^{(2)}$ can be calculated to first order in $Z\alpha$, with results only valid for $kZ\alpha$. For $k=0$, it follows from Table 4.5 that $R_{ls}^{(2)}<1/k^2$. Thus, it may be expected that $R_{ls}^{(2)}$ will contribute substantially to the determination of the IB spectrum at all photon energies. Little is therefore gained by evaluating $R_{ls}^{(1)}$ to any greater accuracy than $R_{ls}^{(2)}$. To illustrate the behavior of the correction factors, we have evaluated the functions $R_{ls}^{(1)}$ and $R_{ls}^{(2)}$ to first order in $Z\alpha$, using Eqs. (4-75) and (4-76), for two atomic numbers (Fig. 4-10).

It is of interest to consider the limit $Z=0$, corresponding to the neglect of Coulomb effects on the intermediate electron states and the momentum of the initial electron. In this limit, $N=N^\lambda$, $R_{ls}^{(1)}=R_{ls}^{(2)}=1$, and Eq. (4-78) is simplified to

$$
\left( \frac{dw_{ls}}{\omega_K} \right)_{CF} = \frac{a}{\pi} \frac{k(q_{ls}-k)^2}{q_{ls}^2} \left[ (1-k/q_{ls})^2+q_{ls}^2 \right] dk.
$$

(4-79)
This result can also be obtained by extending the calculations of Morrison and Schiff (1940) to unique first-forbidden transitions. Equation (4-79) is interesting because of its simplicity but is not expected to be very accurate, although it does describe the general shape of the IB spectrum. The expression is useful for estimating the integrated intensity over any given portion of the spectrum, providing an upper bound.

In order to assess the importance of Coulomb effects in unique first-forbidden transitions and to illustrate the difference between allowed and forbidden shapes, we have plotted several different predictions for the K-capture IB spectrum of $^{41}$Ca in Fig. 4-11.

The two Morrison-and-Schiff curves, labeled MS-A and MS-F, illustrate the basic differences in spectral shape between allowed and first-forbidden unique transitions. The behavior of the Zon-Rapoport (ZR) result at $k>2c$ suggests that, for unique first-forbidden transitions, the main effect resulting from the inclusion of Coulomb effects is an overall reduction in the intensity of the IB spectrum, similar to that found in the allowed case.
Experimental studies of the radiative capture process are valuable for providing information on electron-capture decay, analogous to the information on \( \beta \) decay derived from the study of \( \beta \) spectra. The energy spectrum and the intensity of internal-bremsstrahlung (IB) photons provide a measure of the total energy release and the change of spin and parity in the decay. Experiments on the circular polarization and on various angular correlations provide basic information on weak interaction and nuclear structure. Furthermore, bremsstrahlung experiments may yield supplementary data for the characterization of nuclear decay schemes and for the determination of capture ratios from various subshells.

Precise experimental investigations of radiative electron capture do, however, require rather complicated techniques for experiment and analysis, due to the very low intensity (\( \approx 10^{-4} \) photons per capture event) and the continuous nature of the IB spectra. The interpretation of experimental results is made difficult by the fact that electrons captured from different atomic subshells contribute to the emitted radiation.

Much effort has been devoted to IB experiments during the last thirty years. Critical reviews were compiled by Zylicz (1968) and Kádár (1972), and to a lesser extent by Bouchez and Depommier (1965), Petterson (1965), Schopper (1966), and Berényi (1968). Considerable progress has been achieved since, especially in the development of experimental techniques.
The low probability of radiative capture makes its observation sensitive to interference from other electromagnetic radiation. Especially nuclear γ rays, annihilation radiation and x rays emitted in the course of radioactive decay can considerably limit the energy range of an IB measurement and distort the measured IB pulse-height spectrum through pileup and summing effects. The measurement of coincidences between such primary radiation and the rare IB photons requires sophisticated techniques. In decays with competing β⁺ or β⁻ branches or with highly converted γ transitions, corrections may be required for other types of electromagnetic radiation, comparable in intensity with IB: (i) internal bremsstrahlung accompanying β⁺ or β⁻ decays, (ii) external bremsstrahlung emitted during absorption of β particles or conversion electrons in the source or surrounding materials, and (iii) continuously distributed annihilation radiation for positron annihilation in flight. In view of the large number of possible interfering effects, it is not surprising that IB measurements performed up to now have been restricted to electron-capture transitions in simple decay schemes. In most of the many nuclei decaying by electron capture, radiative capture has not yet been investigated.

It is also evident that IB experiments are very sensitive to small amounts of γ-ray emitting impurities in the sources. Experimental results therefore are only reliable if the source material is carefully checked and purified if necessary to remove spurious contaminants. Impurity checks of the required sensitivity were hardly possible before the advent of high-resolution Ge(Li) spectrometers,
whence older experimental results must be regarded with reservations.

On reviewing the experimental literature, it appears that most measurements of IB spectra have been performed only to derive electron-capture transition energies from the IB end-point energies. This procedure was initiated by the early theory of Morrison and Schiff (1940) and Jauch's proposition to linearize IB spectra in a way that resembles the construction of Kurie plots for β spectra (Jauch, 1951; Bell et al., 1952). For this purpose, most IB spectra were measured without normalization to the electron-capture rate. These shape measurements qualitatively confirmed the spectral shapes predicted by theory for s- and p-type radiation; in the case of forbidden decays, they yielded an estimate of the relative abundance of detour transitions. Measurements of spectral shapes alone, however, are not adequate for a detailed test of modern IB theory (Martin and Glauber, 1958; Intemann, 1971): as shown in Sec. 4.1, relativistic and Coulomb effects, screening, exchange and overlap influence the absolute IB yield, while affecting spectral shapes only slightly. Absolute IB measurements are, however, scarce. Some early results exist, of poor accuracy, pertaining to ground-state transitions; a few results on decays that include γ transitions were obtained recently.

In Secs. 4.2.1-4.2.2, we have compiled the available experimental material and classified the techniques employed in the measurement of normalized IB spectra associated with different decay schemes. We do, however, frequently refer to incomplete studies and list all experiments known to us, to provide a guide for accurate future investigations.
4.2.1. Experiments on Total IB Spectra

An IB spectrum that is not measured in coincidence with x rays or Auger electrons constitutes the total spectrum $\Delta w_{IB}$, which is a superposition of partial spectra $\Delta w_{nl}$ due to electron capture from different atomic states $nl$. This spectrum is mainly determined by $s$ radiation for energies above $\nu_{2n}$ (in units of $mc^2$) and by $p$ radiation at lower energies; contributions from the innermost $1s$ and $2p$ shells dominate.

Experimental techniques applicable to the determination of total spectra can be divided broadly into two categories: single-spectrum methods and coincidence methods. In single-spectrum methods, IB spectra are measured relative to other emitted radiation that can be normalized to the ordinary capture rate. Measurements in coincidence with $\gamma$ rays or conversion electrons permit separation of the IB spectra associated with individual electron-capture branches in a given decay.

In Table 4.6, we list published experiments on total IB spectra and indicate what methods and spectrometers were used and what quantities were deduced. A somewhat more detailed description of experimental methods follows.

Spectrometry of IB and of x rays and Auger electrons. Total IB spectra can most advantageously be observed in pure ground-state transitions and in decays that feed only low-energy transitions. Table 4.6 shows that numerous total IB measurements have been performed on such simple decays, viz., on $^{37}$Ar, $^{41}$Ca, $^{49}$V, $^{55}$Fe, $^{71}$Ge, $^{119}$Sb, $^{125}$I, $^{131}$Cs,
\(^{145}\)Sm, \(^{159}\)Dy, \(^{165}\)Er, \(^{181}\)W and \(^{193}\)Pt. NaI(Tl) and Ge(Li) spectrometers have been used. In cases in which the electron-capture transition energy is high compared with the K x-ray or \(\gamma\)-ray energy, a large fraction of the IB spectrum can be measured. Counting problems produced by the much higher x-ray or \(\gamma\)-ray rates can be avoided by placing suitable absorbers between source and spectrometer. For example, Fig. 4-12 shows IB pulse-height spectra of \(^{131}\)Cs recorded with a NaI(Tl) spectrometer, a variety of Cu absorbers having been interposed (Saraf, 1954a). The procedure fails for transition energies not far above the K x-ray energy; in such cases, pileup from the K x-ray pulses strongly affects the IB spectrum. Methods for pileup reduction and correction are described below.

In most cases listed in Table 4.6, only IB spectral shapes were measured, and the accuracy is generally poor. Precise shape determinations with different types of NaI(Tl) spectrometers have been performed on \(^{55}\)Fe (Berényi et al., 1965b), and on the forbidden spectra from \(^{36}\)Cl (Berényi et al., 1965a, b; Smirnov and Batkin, 1973) and \(^{59}\)Ni (Schmorak, 1963). Only recently were Ge(Li) spectrometers used, resulting in accurate shape measurements on \(^{41}\)Ca (Myszek et al., 1973) and \(^{59}\)Ni (Berényi et al., 1976) and on the IB spectrum from higher shells only in \(^{193}\)Pt (Hopke and Naumann, 1969).

To obtain normalized IB spectra, the ordinary K-capture rate \(W_K\) must be determined from separate measurements of the K x-ray or K Auger-electron emission rates. Normalized IB spectra have been determined in only a few cases: for \(^{37}\)Ar (Saraf, 1956), \(^{55}\)Fe (Michalowicz, 1953; Saraf, 1956), \(^{71}\)Ge (Bisi et al., 1955a), \(^{119}\)Sb
214

\[ \text{Olsen et al., 1957, } ^{131}\text{Cs (Michalowicz, 1956), } ^{145}\text{Sm (Sujkowski et al., 1968), } ^{159}\text{Dy (Sujkowski et al., 1965) and } ^{165}\text{Er (Zylicz et al., 1963; Sujkowski et al., 1965). All these workers used NaI(Tl) crystals to detect the K x rays, with the exception of Saraf (1956), who applied a low-geometry proportional counter for the K x rays in } ^{55}\text{Fe and used internal gas counting to determine the K Auger-electron rate from } ^{37}\text{Ar. In the cases of } ^{145}\text{Sm and } ^{159}\text{Dy, the K x rays could not be resolved from low-energy } \gamma \text{ rays, and decay-scheme corrections were applied. The accuracy of these early normalized IB spectra is generally poor (rarely better than 50%); considerable improvements would be possible today (see Sec. 3). New measurements of total IB spectra would be of great value, especially for pure ground-state transitions which are listed in Table 4.7.}

\text{IB and } \gamma \text{-ray spectrometry. For decays that involve emission of energetic } \gamma \text{ rays, the measurement of total IB spectra is much more complicated. On the other hand, the } \gamma \text{ rays make it possible to normalize the IB spectra, independently of fluorescence yields. If no IB- } \gamma \text{ coincidences are measured, the available energy range is generally limited to energies above the highest } \gamma \text{ energy. These measurements depend strongly on the details of the decay scheme, such as } \gamma \text{ and electron-capture energies and branching ratios, and internal-conversion coefficients.}

\text{To date, IB and } \gamma \text{ spectroscopy has only been applied to relatively simple decays, such as that of } ^{7}\text{Be (Mutterer, 1973b), } ^{51}\text{Cr (Bisi et al., 1955b; Cohen and Ofer, 1955; Van der Kooi and Van der Bold, 1956;}}
The isotopes $^7\text{Be}$ and $^{51}\text{Cr}$ have favorable decay schemes for this type of measurement. Both nuclides decay by two electron-capture branches, $\sim 90\%$ to the ground state and $\sim 10\%$ to an excited state with an energy of $\sim Q_{EC}/2$. Thus, a large fraction of the $\text{IB}$ spectrum associated with the ground-state branch can be measured without interference from the second electron-capture branch. A single $\gamma$ spectrometer can be used to determine $d\omega_{\text{IB}}$ relative to the $\gamma$ emission rate. In order to normalize $d\omega_{\text{IB}}$ to the ground-state electron-capture rate, the $\gamma$ branching ratio $P_{\gamma} = N_{\gamma}/N_0$ is found precisely from measurements of the disintegration rate $N_0$ through $4\pi$ (x-ray, Auger)$-\gamma$ coincidence counting, and of the $\gamma$ rate $N_\gamma$ by integral $\gamma$ counting (Mutterer, 1971; De Roost and Lagoutine, 1973).

Above the $\gamma$-ray energy, the $\text{IB}$ spectrum must be corrected for $\gamma$-ray pileup. In early measurements on $^{51}\text{Cr}$ and $^{113}\text{Sn}$, NaI(Tl) spectrometers were used. With these, poor resolution and long pulse rise times cause the pileup spectra to be smeared out (Waibel, 1969, 1970) and it is not clear whether the measured $\text{IB}$ spectra are free of pileup distortions. These measurements were considerably refined by Ribordy and Huber (1970) and Mutterer (1973a, b) who used Ge(Li) spectrometers with electronic pileup-rejection systems. Such systems prevent pileup of pulses spaced by $2100\ \text{ns}$ and can reduce total pileup by an order of magnitude. Furthermore, the residual pileup spectra show sharp sum peaks that can be distinguished from the smooth $\text{IB}$ spectra (Fig. 4-13). A complete separation of the $\text{IB}$ spectrum from
the residual pileup spectrum cannot, however, be achieved in a single measurement, even with a weak $^{51}$Cr source (Ribordy and Huber, 1970); an extrapolation from measurements with sources of different strengths is required. The extrapolation procedure used by Mutterer (1973a) is illustrated in Fig. 4-14. This technique has yielded normalized IB spectra of $^7$Be and $^{51}$Cr (Fig. 4-15) of good accuracy.

The spectrometry of IB in the presence of $\gamma$ rays could be further improved by using large Ge(Li) detectors and suitable absorbers, in order to optimize the ratio of IB to $\gamma$-ray counting rates, and by using low-background arrangements. The reduction of background, either by applying optimal shielding or by using anticoincidence devices (Persson and Koonin, 1972), allows the use of weak sources and reduces the pileup correction accordingly. It would also be interesting to apply Ge(Li) anti-Compton spectrometers operated with pileup rejectors, because here pileup is confined to the region of the coincidence sum peaks. It can be expected that with improved techniques the accuracy with which total IB spectra of $^{51}$Cr and $^7$Be are now measured can also be attained in cases of decay schemes with higher $P_\gamma$, larger ratios of $\gamma$-ray energy to $Q_{BC}$, or with several $\gamma$ branches. Spectrometry of the ground-state bremsstrahlung offers the possibility of determining ground-state branching ratios that in complex decays can otherwise only be obtained (often with very poor accuracy) from total $\gamma$ and x-ray intensities.

**IB spectrometry in coincidence with $\gamma$ rays.** Spectrometry of internal bremsstrahlung in coincidence with $\gamma$ rays or conversion
electrons permit one to separate the IB spectra accompanying decay to different excited states. Spectra can be measured over their entire energy range, above the K x-ray region, for EC transitions that feed states which decay by prompt \( \gamma \)-ray emission to the ground state or to a lower-lying metastable state of the daughter nucleus. Normalization is easily accomplished by dividing a coincidence IB spectrum by the singles \( \gamma \) counting rate.

IB-\( \gamma \) coincidence experiments have been performed on EC transitions to excited states in the decays of \( ^7 \text{Be} \) (Lancman and Lebowitz, 1971a; Persson and Koonin, 1972), \( ^{51} \text{Cr} \) (Koonin and Persson, 1972), \( ^{54} \text{Mn} \) (Lancman and Lebowitz, 1969; Kádár et al., 1970; Koonin and Persson, 1972), \( ^{57} \text{Co} \) (Lancman and Lebowitz, 1971b), and \( ^{113} \text{Sn} \) (Bosch et al., 1967). The main difficulty in IB-\( \gamma \) coincidence spectrometry arises from the large difference in intensity (\( \times 10^{-4} \)) between IB and \( \gamma \) radiation, because the \( \gamma \)-ray spectra usually cover the same energy range as the weak IB spectra. Very short coincidence resolving times and high-efficiency detectors are therefore necessary to attain good true-to-chance coincidence ratios within reasonable counting times.

Furthermore, scattering between the IB and \( \gamma \) detectors must be avoided to prevent false prompt coincidences and counting losses produced by sum effects in both channels. To meet these conditions, NaI(Tl) scintillators have been used as IB and \( \gamma \) detectors, arranged in close face-to-face geometry. Scattering has been reduced with suitable absorbers (Lancman and Lebowitz, 1969, 1971a, b), sometimes combined with lead collimators (Persson and Koonin, 1972; Koonin and Persson, 1972) (Fig. 4-16). Kádár et al. (1970) employed a 90\(^\circ\) crystal
arrangement of lower geometry with lead collimators. Timing was accomplished by Bosch et al. (1967), Lancman and Lebowitz (1969, 1971a, b), and Kádár et al. (1970) with conventional fast-slow coincidence circuits of 20-35 nsec resolving time. Even so, random coincidences between $\gamma$ events in both detectors made the main contribution to the measured coincidence spectra. Bremsstrahlung spectra were found by subtracting singles $\gamma$ spectra, recorded with the IB detector, from the measured coincidence spectra; both sets of spectra had been normalized to equal photopeak areas (Fig. 4-17).

A considerable improvement in technique was achieved by Persson and Koonin (1972) by using a fast time-to-pulse-height converter and applying two-parameter analysis: the IB pulse-height spectrum and the energy-dependent delay between IB and $\gamma$ pulses were recorded simultaneously. A block diagram of the electronic circuit is shown in Fig. 4-18. This technique has led to effective coincidence resolving times of ~4 nsec over the entire IB-spectrum range. Persson and Koonin (1972) have furthermore reduced the background rate by surrounding both crystals with a plastic-scintillator anticoincidence shield, allowing the use of weak sources. With this technique, random coincidences could be reduced to a much lower level, and accurate IB spectra could be measured for $^7$Be, $^{51}$Cr, and $^{54}$Mn. The result for $^7$Be is shown in Fig. 4-19.

It would undoubtedly be of interest to apply this IB-$\gamma$ coincidence technique to additional cases. Large Ge(Li) detectors or plastic scintillators might be used. In cases that involve low-energy $\gamma$ transitions, measurements might be performed in coincidence with conversion electrons; this has been done only with $^{145}$Sm (Sujkowski et al., 1965).
Spectrometry of IB and of positrons or annihilation radiation.

For high-energy transitions in which electron capture competes with positron emission, total IB spectra can be measured relative to the $\beta^+$ decay rate or to the annihilation radiation ($\gamma_A$) produced by the positrons in suitable source encapsulations. Methods for measuring EC/$\beta^+$ ratios, with which the IB spectra can be normalized, are discussed in Sec. 3.4. Interference of annihilation radiation with the IB spectrum can be reduced by measuring $\gamma_A - \gamma_A$ anticoincidences with a detector placed opposite the IB detector. Alternatively, very thin sources and backings can be used and the $\beta^+$ particles can be magnetically bent away from the IB detector. This technique has been applied by Berényi and Varga (1969) to measure internal bremsstrahlung from $\beta^-$ emission with minimal contribution from external bremsstrahlung. In isotopes that decay by electron capture and $\beta^+$ emission, the positrons give rise to such other continuously distributed radiation as internal and external bremsstrahlung and photons from positron annihilation in flight (Kantele and Valkonen, 1973). The EC bremsstrahlung spectra will be affected by these effects at energies below the $\beta^+$ end point and in the neighborhood of 511 keV.

The only reported IB measurements on a nuclide decaying by EC and $\beta^+$ emission are on $^{36}$Cl, which has a very weak (0.001%) $\beta^+$ branch and decays 98.1% of the time by $\beta^-$ emission. The bremsstrahlung accompanying the 1.9% EC branch has been studied by Dougan et al. (1962), Berényi (1962, 1963b, 1965a), Lipnik et al. (1964), Berényi et al. (1965b), and Smirnov and Batkin (1973) with various types of NaI(Tl) spectrometers. The latter two experiments yielded very
accurate results on the IB spectrum shape at energies above the $\beta^-$ end point at 712 keV. No attempt was made, however, to normalize the spectra to the electron-capture rate.

Bremsstrahlung studies on isotopes that decay by electron capture and $\beta^+$ emission do not contribute decay-scheme information that could not be derived more readily from $\beta^+$ spectrometry. Measurements of normalized IB spectra at high energies would, however, be extremely useful to check the treatment of relativistic and Coulomb effects in the theory, which predicts that these effects reduce the IB yield with increasing energy (Fig. 4-3).

4.2.2. Experiments on Partial IB Spectra

Supplementary to the experiments described in Sec. 4.2.1, considerable effort has been expended to measure partial IB spectra associated with electron capture from specific shells, mainly the $1s$ IB spectrum associated with K capture. Such spectra can be observed by IB spectrometry in coincidence with x rays or Auger electrons. Higher-shell spectra can be determined by subtracting accurately measured $1s$ IB spectra from total IB spectra.

The $1s$ IB spectrum. Spectrometry of internal bremsstrahlung in coincidence with K x rays or K Auger electrons singles out the $1s$ IB spectrum $d\omega_{1s}$. The spectrum can be normalized to the corresponding K-capture rate by dividing the coincidence IB spectrum by the singles K x-ray (K Auger-electron) counting rate.

Only IB-K-x-ray coincidence experiments have been reported
Most of these experiments have yielded only spectral shapes. Normalized spectra have been determined only for some simple decays, viz., for $^{55}$Fe (Biavati et al., 1962), $^{131}$Cs (Michalowicz, 1966; Biavati et al., 1962) (Fig. 4-20), $^{145}$Sm (Sujkowski et al., 1969), and $^{165}$Er (Zylicz et al., 1963; Sujkowski et al., 1965). NaI(Tl) detectors were used in these experiments for both the IB and K x-ray photons; interference of K x-rays in the IB spectrometer was avoided with absorbers. For all isotopes but $^{145}$Sm and $^{165}$Er, only poor accuracy was achieved in these early experiments.

Measurements of bremsstrahlung in coincidence with K x rays can also be performed in the presence of higher-energy $\gamma$ rays, with the restriction that prompt $\gamma$ rays limit the observable IS radiation to energies above the $\gamma$ energy. Spectra accompanying EC decays that feed a state deexcited by prompt $\gamma$ rays of energies in excess of the EC transition energy cannot be obtained by IB-K x-ray coincidences with any degree of accuracy. One IB result on such a cascade in $^{54}$Mn, reported by Jung and Pool (1956), should be disregarded. Delayed $\gamma$ rays, such as arise if electron capture feeds isomeric states, have no direct influence but may contribute considerably to the random-coincidence rate below the $\gamma$ energy. This was the case in the older coincidence experiments on $^{85}$Sr by McDonnell and Ramaswamy (1969), $^{109}$Cd by Gopinathan and Rubinson (1968), and $^{113}$Sn by Jung and Pool (1956). Modern coincidence techniques, as used by Persson and Koonin (1972) in IB-$\gamma$ spectrometry, would permit measurements of entire IS IB spectra. Some results on IS spectra have been reported for EC transitions to isomeric states with mean lives of the order of the...
coincidence resolving time, viz., on $^{125}$I (Gopinathan and Rubinson, 1968), $^{145}$Sm (Sukowski et al., 1968), and $^{197}$Hg (Jasinski et al., 1965). In such cases, only the spectrum above the $\gamma$-ray energy is usually observable, and normalization is complicated.

In all measurements of coincidences between bremsstrahlung and $K$ x rays on radioisotopes that emit prompt or delayed $\gamma$ rays, a correction must be applied for the $\gamma$ contribution in the $K$ x-ray channel. This correction is determined through a second measurement with a discriminator window setting above the $K$ x-ray line. Corrections for $K$ x rays from internal conversion must also be considered.

Bremsstrahlung from $1s$ capture can be measured in coincidence with $K$ x rays even in cases where $\beta^+$ or $\beta^-$ decay competes with electron capture because the only $K$ vacancies created in $\beta$ decay are the few produced by $K$-shell internal ionization or shakeup (Sec. 5). Thallium-204 has often been investigated; this isotope decays by 97.9% $\beta^-$ emission and 2.1% electron capture. Landmann and Bond (1973) have pointed out that double internal bremsstrahlung associated with the $\beta^-$ branch may have to be considered.

Most measurements of $1s$ IB spectra could be considerably improved today. Careful new measurements on pure ground-state decays and EC decays to isomeric states would be especially useful.

Higher-shell IB spectra. The possibilities for measuring the bremsstrahlung that accompanies electron capture from higher shells are more limited. Radiation from $ns$ capture, $n>1$, has very similar shape to $1s$ radiation and constitutes only ~10% of the total
bremsstrahlung in the energy range above $\gamma \alpha$. The radiation accompanying capture of p electrons dominates only at low energies, $k<\gamma \alpha$; for low Z, this is difficult to separate from the characteristic x-ray lines.

The IB spectra associated with capture from higher shells are quite easily observed in the few low-energy transitions in which K capture is energetically forbidden, e.g., $^{193}$Pt and $^{163}$Ho. An accurate IB shape measurement on $^{193}$Pt was performed by Hopke and Naumann (1969) with a Ge(Li) spectrometer (Fig. 4-21). In more energetic transitions, however, such spectra are very difficult to measure with good accuracy.

The 2s IB spectrum, associated with radiative capture of L$_1$ electrons, can only be singled out in coincidence with L x rays if it is possible to gate on the L$_2^3$ and L$_2^4$ lines. Other L x rays can also arise from L$_2^3$ capture or follow K x-ray emission after K capture. The method is thus restricted to high-Z atoms for which the L x-ray components can be resolved and the K fluorescence yield is large. For other nuclides, the 2s spectrum (including s spectra from higher shells) can only be obtained indirectly by comparing accurately measured ls and total IB spectra. No experimental results on separated 2s IB spectra have been reported.

The 2p IB spectrum associated with radiative capture of electrons from the L$_2$ subshell (plus the small amount captured from the L$_3$ subshell) can be singled out by coincidence IB spectrometry in cases in which K capture is forbidden, such as $^{193}$Pt. Here, an IB measurement in coincidence with those L x rays that fill L$_2$ and L$_3$ vacancies (all
but L₃ and L₄ can be performed. The total p radiation, however, that differs strongly in spectral shape from s radiation, can be determined by subtracting from the total IB spectrum measured at energies below Za the s IB spectrum that is measured at higher energies and extrapolated to the p IB region. Alternatively, one can subtract the l's IB spectrum measured by IB-X-ray coincidences and corrected for the ~10% contribution from higher s states (Biavati et al., 1962).

Measurements of total IB spectra at low energies where p IB dominates, have been performed on several nuclides, viz., on ⁵⁵Fe by Biavati et al. (1962), ¹³¹Cs by Michalowicz (1956), Hoppes and Haywards (1966), and Biavati et al. (1962) (Fig. 4-20), ¹⁴⁵Sm by Sujkowski et al. (1968), and ¹⁵⁹Dy and ¹⁶⁵Er by Sujkowski et al. (1965). Relative intensities of p radiation and s radiation were determined for ¹⁴⁵Sm, ¹⁵⁹Dy and ¹⁶⁵Er. In all these experiments, NaI(Tl) detectors were used. With scintillation detectors, however, distortions of the IB spectrum due to pileup contributions from K x rays and K-L x-ray sum effects are difficult to control in the vicinity of the K x-ray energy. For the measurement of p radiation, Ge(Li) spectrometers should be used, preferably with pileup rejectors, and corrections for residual pileup should be considered. Platinum-193 would be a good case for study.

The measurement of the relative intensity N_{IB} from s and p capture represents an independent method to determine the capture ratios; this may supplement corresponding x-ray and Auger-electron experiments.
4.2.3. Analysis of IB Pulse-Height Spectra

In this Section, we consider methods for deriving IB energy spectra \( d_{IB}(k) \) or \( d_{IB,n}(k) \) from measured pulse-height distributions. For continuous spectra, spectrometer calibration is more complicated and analysis more laborious than for line spectra. The calibration procedure must include determination of total detector response over the entire range of energy \( k \) and pulse height \( E \) that is covered by the continuous spectrum. The pulse-height spectrum \( d_n(E) \) and the corresponding photon energy spectrum \( dw(k) \) are, in general, related as follows:

\[
d_n(E) = \int_0^{k_{max}} R(E,k) \, dw(k).
\]  \hspace{1cm} (4-80)

The response function \( R(E,k) \) defines the probability that a photon emitted with energy \( k \) produces a pulse of height between \( E \) and \( E+dE \) when detected. The accuracy to which a measured spectrum \( d_n(E) \) can be compared with a predicted IB spectrum \( d_{IB}(k) \) depends both on the accuracy of \( R \) and the method used to solve Eq. (4-80).

In analogy to extensive work on \( \beta \) spectra, various methods for making response corrections on continuous \( \gamma \) spectra have been worked out that are applicable to measurements with NaI(Tl) and Ge(Li) spectrometers. In the present paper, we can only make a few remarks on essential features. Electron-capture bremsstrahlung spectra have been subjected successively to procedures designed to correct for resolution, Compton distribution, total efficiency, iodine K x-ray escape, etc. (Liden and Starfelt, 1954; Lindqvist and Wa, 1955; Persson and Koonin, 1972). As an example, Fig. 4-22 shows the various
corrections applied by Lindqvist and Wu (1955) to the 19 spectrum of
37Ar. These procedures depend very much on the peculiarities of the
detector arrangement and can differ considerably in accuracy. We
discuss, instead, more generally applicable methods based on the
application of complete response functions.

Determination of response functions for NaI(Tl) and Ge(Li)
spectrometers. The response function R [Eq. (4-80)], which varies
strongly with the type of spectrometer and the measured energy range
(see e.g. Heath, 1963), can in principle be calculated in terms of
the different fundamental absorption processes in the detector. Monte-
Carlo calculations have been performed, e.g., by Beattie and Byrne
(1972) for scintillators and by Meixner (1974) for Ge(Li) specatro-
meters. These calculations have reached a high level of accuracy;
their application, however, is limited by the fact that the true di-
mensions of the detector’s sensitive volume and the thickness of dead
zones and encapsulations are often not accurately known. In fact,
calculations deviate from measured response functions, especially at
low energies.

All pertinent effects are correctly taken into account if the
response function is determined empirically by interpolation, starting
from pulse-height spectra produced by monoenergetic γ rays of known
energies and intensities.

For NaI(Tl) spectrometers, numerous peak-fitting procedures have
been developed (e.g., Prescott, 1963); these allow one to derive the
energy dependence of the fitting parameters. For the interpolation
of Compton distributions, Chester et al. (1963) have fitted parametrized analytical curves to measured spectra and determined the sets of parameters as functions of the energy $k$. Wapstra and Oberski (1963) and others have interpolated between calibration spectra that were transformed so as to bring all Compton edges to a common value of the transformed pulse height. Under special conditions, e.g. with large crystals in close geometry yielding a small Compton-to-peak ratio, it may suffice to approximate the Compton distributions by simple rectangular or trapezoidal shapes. The possibility of such simplification has been demonstrated by Persson and Koonin (1972) for IB spectra measured with a 3x3-in. NaI(Tl) crystal (Fig. 4-16).

With Ge(Li) spectrometers, the correct determination of peak areas is important, whereas the peak shapes can be approximated because the continuous spectra vary little over an energy interval corresponding to a peak width. On the other hand, correct fitting of the Compton distribution is of the utmost importance, especially for small detectors, because the Compton-to-peak ratio is large. Methods for interpolating Compton distributions by fitting parametrized curves (Ribordy and Huber, 1970) and by interpolating transformed calibration spectra (Mutterer, 1973a, c) were reported. Both procedures have yielded accurate Ge(Li)-response functions (Fig. 4-23).

**Correction methods.** With a known response function, a measured IB pulse-height spectrum $d_n^\text{IB}$ can be compared in either of two ways with a theoretical spectrum $d_w^\text{IB}$: (i) the theoretical spectrum can be converted according to Eq. (4-80) into a "predicted" pulse-height
spectrum that is compared with the measured spectrum (folding method), or the measured pulse-height spectrum can be converted into an experimental energy spectrum by solving Eq. (4-80) (unfolding method).

The folding method has been applied most often in the evaluation of IB results, e.g. for $^7$Be by Lancman and Lebowitz (1971b); and Mutterer (1973b, c), $^{37}$Ar by Anderson and Wheeler (1953), Lindqvist and Wu (1955), and Saraf (1956), $^{49}$V by Hayward and Hoppes (1956), $^{51}$Cr by Mutterer (1973a, c), $^{54}$Mn by Lancman and Lebowitz (1969), $^{55}$Fe by Mäder and Freiswerk (1951), Michalowicz (1953), and Biavati, et al. (1962), $^{57}$Co by Lancman and Lebowitz (1971a), $^{113}$Sn by Phillips and Hopkins (1962), $^{165}$Er by Ryde et al. (1963a), and $^{204}$Ti by Lancman and Bond (1973). The folding method is simplest but has the great disadvantage that no direct experimental energy spectrum $d\omega_{IB}$ is obtained. It is thus less valuable for a detailed comparison of IB experiments with theory. Furthermore, the important method for determining the transition energy $E_{EC}$ by constructing a Jauch plot of $d\omega_{IB}$ (Sec. 4.2.4) cannot be applied. Instead, a variational procedure has often been used to determine $E_{EC}$: $dn_{IB}$ is calculated from IB theory and the known detector response as a function of the end-point energy $q$, and $q$ is varied to give the best fit to the measured spectrum (Fig. 4-24). To obtain experimental results for the IB yield as well as $E_{EC}$, both a constant factor and $q$ have often been varied in fitting calculated to experimental IB pulse-height spectra (e.g., Lancman and Lebowitz, 1969, 1971a, b). Experimental results on the IB yield obtained by this method evidently imply theoretical assumptions on the spectral shape.
The unfolding method is consequently to be preferred. Various procedures have been reported in literature. The solution of Eq. (4-80) by matrix inversion usually has to be limited to small matrices. This difficulty can be overcome by iterative methods, such as the correction-factor method of Scofield (1963) and the Gauss-Seidel method (e.g., Zurmühl, 1965). Both of these methods, which also have often been used for unfolding measured β spectra, normally lead to quite rapid convergence if the diagonal elements (E=k) in the response function dominate. Ribordy and Huber (1970) have compared different iterative methods for unfolding the IB spectrum of $^{51}$Cr and find comparable results. The Gauss-Seidel method was applied to the $^{51}$Cr IB spectra by Mutterer (1973a, b), who found rapid convergence of the iteration, provided that the response function was renormalized to unit peak areas. These unfolding techniques performed with the aid of modern computers have yielded accurate response corrections for bremsstrahlung spectra. It should be noted, however, that some problems remain concerning the propagation of statistical experimental errors (Weise, 1968).

4.2.4. Determination of Electron-Capture Transition Energies from Measured IB Spectra

The determination of IB endpoint energies is of particular interest because it provides a direct method for measuring EC transition energies $E_{EC}$ and the corresponding isobaric atomic mass differences $Q_{EC}$. The endpoint of an IB spectrum is equal to the energy $q$ of the neutrino emitted during ordinary (nonradiative) electron
capture and, consequently, the transition energy is obtained by adding to the endpoint energy the atomic binding energy (in the daughter atom) of the shell from which capture has occurred (Rubinson, 1971).

Transition energies have been determined in most IB experiments. In Tables 4.6 and 4.8, $E_{\text{EC}}$ results are listed which were obtained from measurements of total and ls IB spectra. Tables 4.6 and 4.8 also contain $E_{\text{EC}}$ values deduced from the atomic-mass compilation of Wapstra and Gove (1971). With few exceptions ($^{113}\text{Sn}$, $^{125}\text{I}$, $^{197}\text{Hg}$ and $^{204}\text{Tl}$), the IB data are in fair agreement with the atomic-mass differences. It should, however, be noted that the two sets of data are not independent. Wapstra and Gove (1971) have considered part of the listed IB data in assigning the isobaric mass differences, supplementing data from nuclear reaction thresholds and electron-capture ratios. Especially in the medium and high-$Z$ region (e.g., $^{103}\text{Pd}$, $^{109}\text{Cd}$, $^{119}\text{Sb}$, $^{131}\text{Cs}$, $^{145}\text{Sm}$ and $^{181}\text{W}$), the listed $E_{\text{EC}}$ values from IB experiments appear, with slight changes, also in the atomic mass tables.

Because of the great importance of accurate mass differences, some comments on the determination of IB endpoint energies are in order. Many electron-capture transition energies listed in Tables 4.6 and 4.8 originate from early experiments and are of low accuracy. These measurements could be much improved with modern techniques. The overall accuracy of $E_{\text{EC}}$, however, depends also upon the theoretical model which is used to extract the IB endpoint energy $q$ from a measured IB spectrum. This dependence on theory is most obvious in $E_{\text{EC}}$ determinations based on the fitting of calculated spectra to measured
ones, with $q_{1s}$ as the fitting parameter (Sec. 4.2.3). That different theoretical assumptions in this procedure can yield quite different values of $q_{1s}$ was demonstrated by Lancman and Bond (1973) in the case of the first forbidden unique EC decay of $^{204}$Tl: fitting procedures with different allowed IB shapes yielded $E_{\text{EC}}$ values that differ by 25 keV. Shape functions from theories for forbidden transitions were, however, not considered. Berényi et al. (1976) studied the variation of $E_{\text{EC}}$ of $^{59}$Ni, obtained from an accurately measured $^{59}$Ni IB spectrum, by fitting spectra calculated from different theoretical approaches to forbidden radiative capture; they found differences of a few keV.

Most experimental transition-energy determinations from (unfolded) IB spectra $\delta w_{\text{IB}}$ (or $\delta w_{1s}$) have been made by linearizing the spectra in a way that resembles the construction of Kurie plots for $\beta$ spectra. The procedure for constructing such a Jauch plot (Jauch, 1951; Bell et al., 1952) is based on the elementary shape of the $1s$ IB spectrum (Eq. 4-17) as predicted by the early Coulomb-free theory of Morrison and Schiff (1940). A linear plot is obtained by converting a measured spectrum $\delta w_{1s}$ into Jauch coordinates by plotting $(\delta w_{1s}/k)^{1/2}$ vs. $k$. Because of the predicted proportionality

$$(\delta w_{1s}/k)^{1/2} = k - q_{1s}$$

the intercept with the $k$-axis occurs at $q_{1s}$. The accuracy of this procedure evidently depends on how closely the investigated spectrum is approximated by the Morrison-Schiff theory. For a strictly correct linearization, various corrections to the spectrum must be considered which appear in the modern theory for allowed decays (Sec. 4.1.2) and for forbidden decays (Sec. 4.1.4).
The $1s$ IB spectrum from allowed and first-forbidden nonunique decays can be linearized more strictly on the basis of the relation

$$\left( \frac{d\omega_{1s}}{kR_{1s}(k)} \right) = k - q_{1s}.$$  

(4-82)

The $1s$ IB shape function $R_{1s}(k)$ corrects for relativistic and Coulomb effects; it can be calculated exactly from Eq. (4-44). This shape function is displayed in Fig. 4-3 for various atomic numbers. The influence of $R_{1s}$ on the determination of $q_{1s}$ has been studied by Zylicz et al. (1963) in the case of the $^{165}$Er IB spectrum (Fig. 4-15).

It was found that a Jauch plot according to the relation (4-82) yields an endpoint energy that differs by 3 keV from that obtained with a simple plot based on the proportionality (4-81). In this analysis, however, an approximate result for the relativistic shape factor $R_{1s}$ was used, as derived by Martin and Glauber (1958). For $^{165}$Er ($Z=68$) and in the measured energy range, $150$ keV $\leq k \leq 300$ keV, the approximate function deviates considerably from the exact function $R_{1s}$ (Fig. 4-4), so that a greater effect of $R_{1s}$ on $q_{1s}$ is expected. Larger differences are also expected in the case of IB spectra that cover wider energy intervals and are not measured as close to the endpoint.

In order to determine $1s$ IB endpoint energies from measured total IB spectra $d\omega_{IB}$, a correction must be applied for the higher-shell components which have endpoints $q_{nl}$ larger than $q_{1s}$. In the energy range $k < q_{1s}$, this correction can be written

$$d\omega_{IB} = d\omega_{1s} \left[ 1 + \left( \sum_{nl} (\omega_{nl})/d\omega_{1s} \right) \right].$$  

(4-83)
this leads to an additional energy-dependent correction $f(k)$ in the Jauch coordinate:

$$[d\omega_{IB}/kR_{1s}(k)f(k)]_s = k - q_{1s}.$$  \hspace{1cm} (4-84)

The $k$-dependence of $f(k)$ is complicated because it generally contains the higher-shell shape functions $R_{ns}$ and $Q_{np}^2$ (Tables 4.2 and 4.3) as well as corrections for the different endpoint energies $q_n$. In most practical cases, however, only higher-shell $s$ radiation is important at higher energies. The correction for the dominant $2s$ radiation is adequately taken into account by

$$f(k) = 1 + \frac{P_L}{P_K} \frac{R_{2s}(k)}{R_{1s}(k)} \left( 1 + \frac{k_{Kx}}{q_{1s} - k} \right)^2.$$  \hspace{1cm} (4-85)

Here, $P_L/P_K$ is the $L/K$ electron-capture ratio, and the $K$ x-ray energy $k_{Kx}$ has been written for the difference $q_{2s} - q_{1s}$. The term containing $k_{Kx}$ constitutes an important correction for experimental data that are close to $q_{1s}$, within a few times $k_{Kx}$. Because this term implicitly also contains $q_{1s}$, the correction (4-85) can only be calculated iteratively.

Equation (4-84) has been used by Mutterer (1973a, b) in determining the IB endpoint energies of $^7$Be and $^{51}$Cr, with $R_{ns}$ functions calculated from Martin and Glauber's theory, setting $R_{1s} = R_{2s} = 1$.

In total IB spectra that accompany low-energy transitions between high-Z nuclei, $p$ radiation dominates; a correction function $f(k)$ can be calculated from theory, using shape functions $Q_{np}^2$ and the corresponding subshell capture ratios. Because $p$-type spectra deviate considerably from the Morrison-Schiff spectrum, it is expected
that a simple Jauch plot according to Eq. (4-81) may yield quite incorrect results. Consequently, the result for $E_{EC}$ of $^{193}$Pt derived by Hopke and Naumann (1969) from both the L- and M-capture bremsstrahlung (Fig. 4-21) by using Eq. (4-81) should be regarded with reservations.

It is clear that reliable theoretical calculations are necessary for obtaining accurate $E_{EC}$ values from measured IB spectra. A strong argument for the performance of accurate new IB experiments is implied. Measurements on those decays for which accurate $E_{EC}$ values are available from independent experiments are most valuable for testing IB theories.

4.2.5. Experimental Results and Comparison with Theory:

Allowed and First-Forbidden Nonunique Transitions

Most experiments described so far deal with allowed EC decays. They are to be compared with the theory of Martin and Glauber (1958) which, in the $\xi$ approximation, is expected to apply also for first-forbidden nonunique decays.

In most experiments, only spectral shapes have been determined. The results, of varying accuracy, generally agree with theory. This agreement is found both for total IB spectra (experiments listed in Table 4.6) with dominating s- and p-type radiation and for 1s IB spectra singled out by IB-K-x-ray coincidences (Table 4.8). The situation is illustrated in Figs. 4-15, 4-26, and 4-27 for the s IB spectra of $^{51}$Cr, $^{49}$V, and $^{55}$Fe, which cover different energy ranges. Figure 4-20 contains a comparison with theory of the 1s IB spectrum of $^{131}$Cs and of the total IB spectrum which in this case covers an
energy range below $aZ$, so that $p$ radiation dominates. Measured $s$ IB spectra (including $ls$ IB spectra) however, are generally not sufficiently accurate to reveal the weak energy dependence of the predicted IB shape factors (mainly $R_{ls}$). The measured IB shapes can therefore not be used to distinguish between the theory of Martin and Glauber and the pioneering work of Möller (1937a) and Morrison and Schiff (1940). It is seen from Fig. 4-3 that $R_{ls}$ depends on the energy $k$ quite differently for different atomic numbers and in different energy regions. The most precise measurement of an allowed IB shape was performed by Berényi et al. (1965b) on $^{55}$Fe. In Fig. 4-27, their result is displayed in terms of the effective shape factor $R_{eff}$, defined by

$$R_{eff}(k) = \frac{\Delta w_{IB}(k)}{k(q_{ls}^2 - k^2)}. \quad (4-86)$$

The function $R_{eff}$ is equal to $R_{ls}$, multiplied by the correction function $f(k)$ for higher-shell contributions [Eq. (4-85)]. The accuracy of the experiment of Berényi et al. (1965b) is comparable with the accuracy attainable in determinations of $\beta$ shape factors. The constancy of $R_{eff}$ within 1%, found in this measurement, can also be compared in this special case with the Martin-Glauber theory. It is seen from Fig. 4-3 that, for $Z=26$, the theoretical $ls$ shape factor has a flat maximum between 100 and 218 keV, a range covered by the $^{55}$Fe experiment. To reveal the dependence of $R_{ls}$ on $k$, accurate shape measurements below $\sim 100$ keV and on transitions of high energy (e.g. $^{37}$Ar and $^{49}$V) should be performed.

Only in a few experiments has the IB intensity been measured in
addition to the shape. Some of the data on normalized IB spectra have been compiled and compared with theory by Bouchez and Depommier (1965), Kádár et al. (1970), Lancman and Lebowitz (1971a), Vanderleeden et al. (1971), Kádár (1971), Koonin and Persson (1972), and Mutterer (1973c). Conclusions from these summaries were partly inconsistent, depending on whether or not theoretical values were recalculated and which values for the transition energies $E_{BC}$ were inserted. Here we therefore compare experimental results with consistently calculated theoretical values.

The selected experimental data on normalized IB spectra from allowed and first forbidden nonunique decays are compiled in Table 4.9. These data represent integral values $I_{IB}(k_1,k_2)$ of normalized spectra $d\omega_{IB}$ measured between energy limits $k_1$ and $k_2$. The upper limits $k_2$ are always equal to or slightly below the endpoint energies $q$. We did not consider data for which $k_1$ and $k_2$ were not specified, as in measurements on $^{51}$Cr by Cohen and Ofer (1955), Ofer and Wiener (1957), Murty and Jnanananda (1957), and Ribordy and Huber (1970). In these experiments, the $\gamma$ branching ratio in $^{51}$Cr was determined by comparing IB and $\gamma$ intensities. We also have omitted results on $^{51}$Cr reported by Vanderleeden et al. (1971) and Kuphal et al. (1973), which were deduced from measurements of circularly polarized bremsstrahlung, because the measured energy range could not be inferred clearly. It is to be noted that experimental $I_{IB}$ values obtained from measured IB pulse-height spectra through the folding method (Sec. 4.2.3) also do not exactly represent the IB intensity within stated limits, but
rather constitute ratios of counting rates between corresponding limits \( E_1 \) and \( E_2 \). These values are included in the comparison, but are specially identified in Table 4.9. In this Table, we distinguish between values for 1s IB intensities \( I_{1s} \) and intensities of total IB spectra, measured relative to the ordinary K or total EC rates.

The experimental values listed in Table 4.9 are compared with predictions according to different approaches to the theory of allowed radiative capture. Predictions for 1s IB intensities (which dominate as well in most of the listed total spectra) have been calculated from the Coulomb-free theory of Morrison and Schiff (MS) (1940) and from the theory of Martin and Glauber (MG) (1958). Results are listed from both the (analytical) low-energy approximation of Eq. (4-38) (MG) and the (numerically calculated) exact solution [Eq. (4-44)] as derived by Intemann (Int) (1971). The higher-shell contributions were consistently calculated from the approximate relativistic Martin-Glauber theory (Glauber and Martin, 1956) with screening corrections of Fig. 4-7. We consistently used transition energies \( E_{EC} \) derived from the atomic mass compilation of Wapstra and Gove (1971) and, for EC decays to exited states, from \( \gamma \) energies as evaluated by Meixner. Energies \( q_{ni} \) were calculated from atomic binding energies (Bearden and (1971)) The only exception is \(^{197}\text{Hg} \), where the \( E_{EC} \) value of 336±20 keV derived from the mass table (which originates from a \( P_{K} \) measurement by DeWit and Wapstra, 1965) falls completely outside of the range of the measured IB spectrum. Uncertainties in the calculated intensities which are due to the stated errors in the energy \( E_{EC} \) were estimated from the Coulomb-free theory by differentiating Eq. (4-16) with respect to \( q_{1s} \). These uncertainties were found to be generally
below 1%, except for $^{71}$Ge, $^{133}$Cs, $^{145}$Sm and $^{165}$Er, where they lie between 2% and 7%, and for $^{197}$Hg, where it is 40%.

In Figs. 4-28 to 4-30, ratios $\frac{p_{E,T}}{\rho_{E,T}}$ of experimental and theoretical bremsstrahlung intensities are plotted as a function of atomic number. We have included only data which pertain predominantly to $s$ radiation. The indicated error bars correspond to the sum of experimental and theoretical errors.

In Table 4.10, unweighted average values $<p_{E,T}>$ of independent measurements are listed.

The summary of experimental and theoretical IB intensities proves the advantage of the theory of Martin and Glauber over the Coulomb-free approach. The measured data, although widely scattered, clearly reveal the predicted reduction of the IB intensity, increasing with $Z$, that is caused by relativistic and Coulomb effects. While this lowering of the intensity is most obvious in the heavier nuclides with $51 < Z < 80$, it can also be noted, on the average, in the data on lighter nuclides with $18 < Z < 32$.

The ratios $\frac{p_{E,T}}{\rho_{E,T}}$ between measured data and theoretical intensity according to either MG or Int deviate from unity by up to 50%; the deviations in most cases are larger than the error bars. The available data do not allow one to distinguish between the approximate solution of MG and the exact solution of Int. The average values $<p_{E,T}>$ in Table 4.10 provide no evidence that experiments deviate systematically from the Martin-Glauber theory, either at low $Z$ or in general, contrary to indications in previous surveys by Lancman and Lebowitz (1971a) and Vanderleeden et al. (1971). On the other hand, the
inconsistency between experimental results makes it difficult to assign a limit within which the present theory correctly seems to describe the intensity of bremsstrahlung.

Inspection of the experimental data shows that there is a special discrepancy between some of the most recent results obtained by different groups of authors, with quoted probable errors of 5% to 20%. The IB-γ coincidence experiments of Lancman and Lebowitz (1969, 1971a, b) on 7Be, 54Mn, and 57Co yielded intensities that fall 20% to 50% below theoretical predictions, whereas similar experiments, performed with improved techniques by Persson and Koonin (1972) and Koonin and Persson (1972) on 7Be, 51Cr, and 54Mn led to results that exceed theoretical intensities by up to 30%. The spectrometry of IB and γ rays in 7Be and 51Cr by Mutterer (1973a, b) yielded IB intensities which are in agreement with theory to within ±8%. This inconsistency suggests that unknown sources of systematic errors of ±10% remain in the experimental techniques and in the procedures applied for the response correction.

**Double IB.** Experimental evidence for the simultaneous emission of two IB photons during electron capture, or double internal bremsstrahlung, has been reported by Ljubič et al. (1974). Coincidences between two IB photons from 37Ar were measured at an angle of 90° to each other. In the energy range from 210 to 810 keV, the ratio of double IB to single IB was found to be $(4.8±0.4) \times 10^{-5}$, which is comparable to the IB/EC rate, as might be expected. The only presently available theoretical results on double IB are those of
Menhardt (1957), which unfortunately are not applicable to the experimental situation realized by Ljubičić et al. (1974).

4.2.6. Experimental Results and Comparison with Theory: IB Spectra from Higher-Forbidden Decays

Experimental information on IB spectra that accompany higher-forbidden transitions is limited to a few cases of ground-state transitions by EC alone ($^{41}$Ca and $^{59}$Ni) or with competing $\beta^-$ branches ($^{36}$Cl and $^{204}$Tl). Some of these decays have been measured extensively (Tables 4.6 and 4.8).

Bremsstrahlung from first-forbidden unique transitions has been studied with $^{41}$Ca and $^{204}$Tl. The total IB spectrum from $^{41}$Ca was measured with a Ge(Li) spectrometer by Mysłek et al. (1973). The observed shape (Fig. 4-31) is not in accord with the theory of Zon and Rapoport (1968) and Zon (1971). The shape agrees with theory only at low energies, $k<250$ keV; at higher energies the spectrum has nearly allowed shape. Mysłek et al. have also derived a crude value for the IB intensity by estimating the K capture rate from the weight of the source. A value of $3.9\times10^{-4}$ IB photons between 90 and 421.5 keV per EC event was found, much in excess of theoretical predictions. For forbidden transitions, theory in the Coulomb-free approach of Eq. (4-79) leads to an upper limit of $5.7\times10^{-5}$ IB photons per decay. The low-Z expansion of Zon and Rapoport (1968) results in a value of $3.3\times10^{-5}$ photons and more detailed calculations of Zon (1973) have yielded $4.9\times10^{-5}$ photons per EC transition.

On $^{204}$Tl, several IB-K-x-ray coincidence experiments have been
performed to measure the 1s IB spectrum that accompanies the 2.1% EC branch to the ground state of $^{204}\text{Hg}$. Severe doubts exist regarding the reliability of early experiments by Der Mateosian and Smith (1952), Jung and Pool (1956), and Biavati et al. (1962), because no corrections for the bremsstrahlung from the 97.9% $\beta^-$ branch to $^{204}\text{Pb}$ were applied. It was first pointed out by Goudsmit et al. (1966) and established by a recent experiment of Lancman and Bond (1973) that scattering effects due to the $\beta^-$ bremsstrahlung and double-IB emission can cause a continuous spectrum that closely resembles the IB spectrum expected from the weak EC branch. The shape of the IB spectrum measured by Lancman and Bond, taking account of corrections for these effects, agrees well with the theory for allowed transitions. Such agreement is expected because only a small energy range is involved. The intensity per K capture was found to be $2.8 \times 10^{-6}$ IB photons above 103 keV. This result is to be compared with an upper limit of $1.08 \times 10^{-5}$, from Eq. (4-79). More accurate theoretical results are not available. Zon (1973) has reported only values for the shape factors, and the Zu expansion of Zon and Rapoport (1968) is not applicable for $^{204}\text{Tl}$, because the entire IB spectrum lies in the region of the K binding energy.

Bremsstrahlung spectra from second-forbidden nonunique transitions were studied with $^{36}\text{Cl}$ and $^{59}\text{Ni}$ by several groups. It has been well established that the total IB spectrum of $^{36}\text{Cl}$ closely follows an allowed shape at energies above 600 keV. This observation agrees with the calculation of Zon (1971), which predicts a noticeable deviation from allowed shape only at lower energies. The result of the
most recent IB measurement on $^{36}$Cl by Smirnov and Batkin (1973) is shown in Fig. 4-32. An attempt was made to look for possible contributions from detour transitions; a clear indication could, however, not be established.

In the case of $^{59}$Ni, a distinct deviation of the IB spectrum from allowed shape was observed already in an early measurements by Saraf (1956); the IB intensity above 100 keV was reported to be $1.4 \pm 0.4$ times the theoretical value calculated from the early theory of Cutkowsky (1954). The shape of the $^{59}$Ni IB spectrum was very carefully measured by Schmorak (1963), who used 3x3-in. and 5x5-in. NaI(Tl) detectors. An apparent deviation from the calculated shape, observed near the endpoint, was attributed to destructive interference with detour transitions, as predicted by Rose et al. (1962). Schmorak (1963) estimated the amount of detour transitions as between $6 \times 10^{-3}$ and $5 \times 10^{-4}$ of the total IB intensity. A careful Ge(Li) measurement of the $^{59}$Ni IB spectrum was recently performed by Berényi et al. (1976), who report that the measured shape agrees well with calculations of Zon (1971) and shows no evidence for detour transitions.

In most experiments performed hitherto, only spectral shapes have been determined albeit often with high precision. Without question, accurate measurements of normalized spectra would be of great value to improve our present knowledge of radiative capture in forbidden EC transitions. Pertinent EC nuclides are listed in Table 4.7.

4.2.7. Experiments on IB Correlation Effects

Experiments on the various IB correlation effects that are
to allowed decay's and discussed in Secs. 4.1.3 are scarce. They are confined to measurements of circular polarization and to some work on the angular distribution of IB photons emitted from oriented nuclei.

**Circular polarization of internal bremsstrahlung.** Experiments on the circular polarization of the IB accompanying electron capture are listed in Table 4.11. Polarimeters employed in these measurements are based on the effect of spin-dependent Compton scattering from electrons in magnetized iron. Usually, forward-scattering magnets have been used, but in the most recent experiment by Kuphal et al. (1974), a specially designed radial-transmission magnet was employed.

The polarization $P$, defined by Eq. (4-88), is proportional to the relative change $\Delta N/N$ of the measured intensity when the magnetic field in the scattering magnet is reversed:

$$P \propto \frac{\Delta N}{N} = \frac{2(N_+ - N_-)}{(N_+ + N_- - 2N_0)}.$$  \hspace{1cm} (4-87)

Here, $N_+$ ($N_-$) is the counting rate with the electron spins in iron parallel (antiparallel) to the incident-photon momentum. The counting rate $N_0$ is due to background, including $\gamma$ impurities in the source.

If nuclides are measured which emit also nuclear $\gamma$ rays, the denominator in Eq. (4-87) is represented by the counting rate of the unpolarized $\gamma$ rays. The measured effect is then extremely low.

The polarization of IB in pure ground-state transitions was studied in early measurements on $^{37}$Ar by Hartwig and Schopper (1958) and Mann et al. (1959), $^{55}$Fe by Paste nova (1960), and $^{71}$Ge by Bernardini et al. (1958). Only recently, IB polarization has also been measured in the presence of a background of much more intense...
γ rays. In such experiments on $^{51}$Cr, Vanderleeden et al. (1971) used a Ge(Li) detector, whereas Kuphal et al. (1974) used a ring of 8 NaI(Tl) scintillators. In both experiments, very strong $^{51}$Cr sources of up to 500 Ci were employed, and current integration was applied instead of counting techniques. The statistical errors could be kept below $10^{-6}$. Clearly, the polarimeter efficiency must be accurately known to derive the absolute polarization from the measured rate $\Delta N/N$. This efficiency has generally been calculated from basic assumptions. Kuphal et al. (1974) have tested their calculation by measuring polarized (internal and external) bremsstrahlung from several $\beta^-$-decaying nuclides.

Measurements summarized in Table 4.11 confirm within errors that $s$-type bremsstrahlung is nearly 100% right-circularly polarized, due to the parity-nonconserving character of the weak interaction. The measured polarization of $\mathrm{IB}$ from $^{37}$Ar (Hartwig and Schopper, 1958) is displayed in Fig. 4-3. Figure 4-3 shows $\Delta N/N$ values measured for $^{51}$Cr, compared with calculations from theory. The incomplete polarization observed in $^{37}$Ar at low energies and the low result for $^{71}$Ge found by Germanoli et al. (1958) can qualitatively be explained by Coulomb effects and the influence of unpolarized $p$-type bremsstrahlung. Both effects, which enter in the overall polarization function $P(k)$ according to Eq. (4-52), reduce the polarization at low energies. A noticeable reduction of $P$ is not expected, however, in the high-energy bremsstrahlung from $^{51}$Cr; the low value of 0.67±0.07 found by Vanderleeden et al. (1971) can probably be attributed to an erroneous calculation of the polarimeter efficiency, in view of the work of Kuphal et al. (1974).
Angular distribution of IB emitted from oriented nuclei. Anisotropy of IB emitted from oriented nuclei has been observed only once. Brewer and Shirley (1968) studied the forward-backward asymmetry of IB from oriented $^{119}$Sb. Carrier-free $^{119}$Sb had been implanted in an iron lattice, cooled to $\approx 0.02^\circ K$, and magnetized in a field of 2.3 kOe. The IB radiation was measured with two 3×3-in. NaI(Tl) detectors placed at 0$^\circ$ and 180$^\circ$ relative to the direction of the magnetic field. Figure 4-35 shows the measured asymmetry $W(\theta)/W(0)$ as a function of the sample temperature $T$ that defines the degree of source polarization. The measurement of this ratio for different energy intervals has revealed an unexpected energy dependence of the asymmetry.

The experimental results of Brewer and Shirley have been compared with theory by Intemann (1971) in terms of the overall asymmetry function $A(k)$ of Eq. (4-62). It was found that the measured decrease of $A(k)$ at low energies can be well explained (Fig. 4-36). This decrease is consistent with the observed decrease of the overall polarization, described above. Other nuclei that might be suitable for measuring IB angular correlations have been listed by Koh et al. (1962).

A preliminary measurement of the angular correlation between bremsstrahlung and nuclear $\gamma$ rays in the decay of $^{84}$Rb has been performed by Chasan and Chandra (1967). The result was reported to be in approximate agreement with calculations of Koh et al. (1962). The experimental error, however, is $\approx 50\%$ and details of the measurement have not been fully reported, so that a detailed comparison with theory is not feasible.
4.2.8. Concluding Remarks

The study of second-order effects, such as internal bremsstrahlung, is of particular interest in electron-capture decay because experimental information on the primary process is very limited because of the extremely low interaction probability of the emitted neutrino. The main features of the low-intensity radiative-capture process are generally understood today. There is still a great need, however, for experimental work to test the details of the theory. Open questions remain concerning the influence of screening and exchange and overlap effects (Persson and Koonin, 1972) on the shape and intensity of IB spectra. Experimental information is still very scarce on forbidden decays, where nuclear-matrix elements play an important role.

Precise measurements of normalized IB spectra are very much needed. The same holds for measurements of partial spectra that accompany the capture of electrons from specific atomic subshells. Experimental techniques have been developed to high accuracy in recent years. This applies especially to the determination of electron-capture rates, to coincidence experiments with bremsstrahlung, and to calibration procedures for γ spectrometers which yield complete response functions. It can be expected that it will be possible to measure normalized IB spectra in the near future with an overall accuracy of a few percent, at least in some favorable decays such as pure ground-state transitions. As pointed out before, precise experiments are also very valuable for providing accurate isobaric atomic-mass differences, supplementary data on subshell capture ratios, and spectroscopic information on branching ratios. In this context, refined computa-
tions of higher-shell IB spectra, based on the present theory, would be of interest.

The variety of IB correlation effects, discussed in Secs. 4.1.3 - 4.1.4 opens another wide field for future experimental work, from which valuable information on the weak interaction and nuclear structure can be expected. Bremsstrahlung measurements may also help to solve some specific problems of radionuclide metrology (Spernol et al., 1973; Matterer, 1973c), such as the (relative and absolute) determination of disintegration rates of pure EC nuclides.
5. ATOMIC TRANSITIONS ACCOMPANYING NUCLEAR ELECTRON CAPTURE

5.1. Introduction

In first approximation, the probability of allowed capture of a K electron by the nucleus is

\[ \lambda_K = G^2 q^2 |\psi^{(K,Z)}(0)|^2, \tag{5-1} \]

where \( G \) is the \( \beta \)-decay coupling constant, \( q \) is the energy of the emitted neutrino, \( \xi \) the appropriate combination of nuclear matrix elements, and \( \psi^{(K,Z)}(0) \) is the 1s-electron wave function, evaluated at the origin (Sec. 2). The only electron wave function contained in this formulation is that of the electron which is destroyed. Two significant aspects of the problem are neglected in Eq. (5-1): (1) the indistinguishability of electrons, and (2) the nuclear charge change by one unit, which entails that parent and daughter atomic wave functions are eigenfunctions of different Hamiltonians.

The importance of treating \( \beta \) decay and nuclear electron capture as transformations of the whole atom, and hence, of including atomic variables in the description of initial and final states, was first emphasized by Benoist-Gueutal (1950, 1953a, b), pursued by Odiot and Daudel (1956), and comprehensively formulated by Bahcall (1962a, 1963a, b). In fact, an infinite number of final atomic states, including continuum states, contribute to any given electron-capture probability. The effect on transition rates is discussed in Sec. 2. In the present section, we consider observable atomic effects that result during
nuclear decay by electron capture. Atomic rearrangements that take place after the decay process are not included in this discussion, even though x rays emitted in the course of such rearrangements have led to the discovery of the process (Alvarez, 1937, 1938a, b) and constitute the most readily detectable signals indicating that capture has taken place. Details of the rearrangement process have been surveyed by Rao et al. (1972) and Bambeyek et al. (1972). Here, we consider atomic transitions that take place in the course of the electron-capture decay process, due to imperfect overlap between parent and daughter atomic wave functions. This effect is variously denoted as electron shakeup and shakeoff, autoionization, or internal ionization.

5.2. Internal Ionization: Nonrelativistic Theory

Nuclear electron capture is accompanied by the emission of low-intensity, continuous photon and electron spectra. The internal-bremsstrahlung photon spectrum emitted during radiative electron capture was first calculated by M"uller (1937a) and by Morrison and Schiff (1940); this subject is discussed in Sec. 4. The process of internal ionization was first treated by Primakoff and Porter (1953), who calculated the probability of K-electron ejection during K capture and derived an expression for the ejected-electron spectrum, in analogy with work by Migdal (1941) and Feinberg (1941) on orbital-electron ejection accompanying β-particle emission.

The weak interaction which is responsible for nuclear electron capture is of very short range. On the atomic time scale, the
transformation of the parent nucleus with \( Z \) protons into the daughter nucleus with atomic number \( Z' \) can be assumed to be instantaneous. One can gain an intuitive feeling for the mechanism that causes internal ionization by considering the nucleus plus the orbital vacancy created by the capture simply as the source of a suddenly changing Coulomb potential. A 1s electron, for example, with the wave function \( \psi^{(K,Z)} \) in the parent atom does not have time to adjust its wave function adiabatically to the change in potential; the sudden approximation of time-dependent perturbation theory applies. The amplitude of the probability that the electron retains its original quantum numbers is then proportional to the overlap of its original wave function with the 1s wave function in the daughter ion with one inner vacancy:\(^{25}\)

\[
P_{\text{remain}} = |\int \psi^{(K,Z')}(r) \psi^{(K,Z)}(r) dr|^2. \tag{5-2}
\]

Similarly, the overlap of \( \psi^{(K,Z)} \) with excited- and continuum-state wave functions in the potential of the daughter ion provides an indication of the probability amplitudes of excitation or ejection of the K electron. The Pauli principle excludes excitation into occupied orbitals, and conservation of angular momentum allows only \( l=0 \) final states for s-electron shakeup or shakeoff.

It is a gross oversimplification, however, to consider the nucleus-plus-vacancy as a mere source of an abruptly changing electrostatic potential, and the internal excitation and ionization probabilities as determined by wave-function overlap alone. In particular, energy conservation and the demands of quantum
statistics are not included unless the process is treated as a transformation of the whole atom, and nuclear and lepton variables (including those characterizing the pertinent atomic electrons) are incorporated in the description of the initial and final states of the system. Especially, the available energy is shared statistically between ejected electron and neutrino, and the transition probability is weighted by the density of available final states. The energy-conserving delta function must be included in the expression for the transition probability.

Quantitatively, the transition probability can be expressed in the sudden perturbation approximation through "Fermi's Golden Rule No. 2." The general applicability of this approach to the present problem has been examined by Bahcall (1963a). The transition rate for $K$-electron ejection during $K$ capture is

$$\frac{d\omega}{d\omega} = 2\pi \left(\frac{1}{2}\right) |H|^2 \delta \left(W + W_{-} - W'_{K} + W - q \right) dq dp, \quad (5-3)$$

where $p$ and $W$ are the momentum and total relativistic energy of the ejected electron, $q$ is the neutrino momentum and $q$ its energy, $1 - |E'_{K}|$ is the total energy of a $K$ electron in the daughter atom (with binding energy $E'_{K}$), and $W_{+1}$ is the energy difference between the parent atom and the neutral daughter atom (Sec. 1.2). The matrix element $H$ is discussed below. The units used throughout this discussion are such that $\hbar = m = c = 1$, and hence, $e^2 = \alpha = 1/137$. The summation sign in Eq. (5-3) indicates summing over spin states of ejected electron and neutrino, and over spin states of the two initial $K$ electrons. One must also sum over final nuclear spin states and
average over initial nuclear spin states. Because \( \sum |M|^2 \) is independent of \( \mathbf{q} \), the integration over all possible neutrino momenta can be carried out at this stage.\(^{26}\) The result, after performing the spin summations, is

\[
dw = 16\pi^2 (W^2 + 1 - |E_{K'}| - |E_{K''}|)^2 \rho_{\mathbf{p}}^2 |M|^2
\]

(Intemann, 1972).

In a representation in which the interaction Hamiltonian consists of the \( \beta \) interaction \( H_\beta \) alone, the matrix element for K-electron ejection during K capture can be written

\[
M = (1-P_{12}) \int_0^\infty \psi_N(0) \psi_v(0) H_\beta \psi_{N1} \psi_{N2} (K,Z)(0,\mathbf{r}) d\mathbf{r}.
\]

Here, \( \psi_N \) and \( \psi_v \) are the final and initial nuclear wave functions, respectively, and \( \psi_v \) is the neutrino wave function. The wave functions of the leptons that participate in the \( \beta \) interaction have been replaced by their values at the origin. It is assumed that all but the two K electrons retain their original quantum numbers, and that their initial and final wave functions overlap perfectly. The exchange operator \( P_{12} \) exchanges the two K electrons.

The main difficulty in explicitly writing out the matrix element (5-5) resides in expressing the initial-state two-electron wave function \( \psi_{1,2}^{(K,Z)}(r_1,r_2) \), including correlation effects between the two electrons. Primakoff and Porter (1953), in their classic calculation, used an approximate wave function of the form

\[
\psi_{1,2}^{(K,Z)}(r_1,r_2) = N \psi_1^{(K,Z)}(r_1) \psi_2^{(K,Z)}(r_2) e^{\gamma_1 r_1 e^{\gamma_2 (r_1 + r_2)}}.
\]
where $\psi_1$ and $\psi_2$ are hydrogenic ls wave functions, and $N$ is chosen to assure normalization. The factor $e^{\alpha_1 x_1}$ takes account of the effect of the electron-electron Coulomb interaction on their spatial correlation, and the factor $e^{\alpha_2 (x_1 + x_2)}$ accounts for screening of the nucleus, effectively replacing $Z$ by $Z - \gamma_2$ in $\psi_1$ and $\psi_2$. The parameters $\gamma_1$ and $\gamma_2$ were chosen so that $\psi_{12}$ is a good approximation to the Hylleraas variational wave function for a two-electron atom.

With a Coulomb wave function $\psi^{(\omega, z')}$ to describe the ejected electron and a plane wave to describe the neutrino, the matrix element (5-5), and hence, the transition rate (5-3) were computed. Dividing by the transition rate $w_K$ for ordinary allowed K capture, Primakoff and Porter derived an expression for the probability, per K-capture event, for ejection of the other K electron with a momentum in the range $dp$. This result can be written

$$\frac{dW_{\text{ejec}}}{w_K} = \frac{16\alpha^2 \frac{\pi}{p}}{(\frac{\pi}{2} + \frac{1}{2})^4 (1 - \frac{p^2}{2\pi^2})} \left[ 1 - \frac{p^2}{2\pi^2} \frac{\pi}{p} \right]^2 dp. \quad (5-7)$$

Again, $W_0$ and the K-electron binding energy $E'_{K}$ are in units of $mc^2$, the ejected-electron momentum $p$ is in multiples of $mc$, and $\zeta$ stands for $\alpha Z$. We have neglected $\gamma_1 + \gamma_2 \approx 0.5$ and unity compared with $Z$ in the final result, and have set $\beta = p$, i.e., $W_{\text{ejec}}$ for the ejected electrons, whose kinetic energy is generally very much smaller than $mc^2$. As before, $\Delta W = \Delta W_{\text{nuc}} - \Delta (\Sigma E_x) + 1$ is the mass difference between parent and daughter neutral atoms; $\Delta W_{\text{nuc}}$ is the nuclear energy release, and $\Delta (\Sigma E_x)$ is the change in the total electronic binding energy between parent and daughter atoms—a positive quantity in electron capture (Sec. 1.2).
A very different method for constructing the initial two-electron wave function was devised by Intemann and Pollock (1967), who calculated it from perturbation theory. They treated the electron-electron interaction as a perturbation on the nuclear Coulomb interaction, including it in the perturbed part of the Hamiltonian, rather than in the unperturbed part as Primakoff and Porter had done. In essence, they performed a perturbation expansion on the exact two-electron wave function, with the perturbation taken to be the electron-electron interaction. With this approach, the problem of K-shell internal ionization during K capture is one in third-order perturbation theory, involving a sum over intermediate electron states. Intemann and Pollock found it possible to represent this sum in closed form by drawing upon the analogy between internal ionization and internal-bremsstrahlung emission. In fact, the electron-ejection process can be looked upon as a radiative capture process in which the emitted photon is virtual, and is absorbed by the electron that is ejected. Exploiting this aspect of the problem, Intemann and Pollock were able to take advantage of a crucial observation made by Glauber and Martin (1956; Martin and Glauber, 1958) in their development of the theory of radiative capture, viz., that the sum over intermediate electron states which appears in the calculation is the Green's function for the Dirac equation with a nuclear Coulomb potential and can be represented in closed form. This approach made a more exact analysis of the internal-ionization process possible. The result for the differential transition rate per K-capture event is
Here, we have $\mu = [2(1-\varepsilon)]^{1/2}$, where $\varepsilon$ is the intermediate-state energy (in units of $m^2c^2$) of the electron undergoing capture: $\varepsilon = \varepsilon_1 + \varepsilon_2 - \varepsilon$, where $\varepsilon_1$ and $\varepsilon_2$ are the energies of the initial K electrons, and $\varepsilon$ is the energy of the ejected electron. In the Intemann-Pollock treatment, the relation $\varepsilon = 1 - \frac{\varepsilon_1^2}{2} - \frac{\varepsilon_2^2}{2}$ holds, because $\varepsilon_1 = \varepsilon_2 = 1 - \frac{\varepsilon}{2}$ and $\varepsilon = 1 + \frac{\varepsilon_2^2}{2}$.

The integral $I$ is

$$I = \int_0^1 \frac{x^{-\eta-1}[1-(1-x)^{\eta}]^{\varepsilon}(x)}{x} dx,$$

where

$$f(x) = \frac{e^{-\frac{(2z/p)\tan^{-1}[(2z+\mu)/p]}{e^{2(2z/p)\tan^{-1}[(2z+\mu)/p]}+\mu(1+x)/p}}}{(1+\lambda x)^{\varepsilon}(1+\sigma^2 x)}$$

and the remaining symbols are defined as $\eta = \xi / \mu$, $\sigma = (\mu - 2\xi - ip) / (\mu + 2\xi + ip)$, $\lambda = (\mu - \xi) / (\mu + \xi)$. Fortunately, a rapidly converging Maclaurin series exists for $I$:

$$I = \sum_{n=0}^{\infty} \frac{\xi^{(n)}(0)}{n1} \left[ -\frac{1}{(n-\eta)} + \frac{4}{(n+1-\eta)} - \frac{6}{(n+2-\eta)} + \frac{4}{(n+3-\eta)} - \frac{1}{(n+4-\eta)} \right].$$

Intemann and Pollock find that, for $Z=26$, an error only of order $\xi^2$ results from breaking the series off after the first three terms.
The energy spectrum of electrons ejected from $^{55}\text{Fe}$, predicted by this more exact theoretical approach, does not appear to differ materially from that of Eq. (5-7) when placed on a semi-logarithmic plot (Intemann and Pollock, 1967). Some writers have consequently assumed that the results of the two theories are truly identical. This is not the case. In fact, the momentum spectra from the two approaches differ appreciably on a linear plot; they have approximately the same shape, but the Intemann-Pollock spectrum has somewhat lower intensity. Hence it yields significantly smaller values for the total ejection rate than the Primakoff-Porter theory (Intemann, private communication). That the difference is not greater appears to indicate that Primakoff and Porter's variational wave function takes unexpectedly accurate account of screening and correlations in the initial two-electron state. Improving the accuracy of the continuum wave function to take screening in the final state into consideration is only expected to affect the results of Eqs. (5-7) and (5-8) by $\lesssim 5\%$ for $Z=26$.

The neglect of relativistic effects inherently limits the accuracy of the results to a relative error of order $\alpha Z$, even at the lowest ejection energies. The nonrelativistic calculations were pushed to this limit in a refinement, due to Intemann (1972), of the Intemann-Pollock approach, which involves the use of a more elaborate Coulomb Green function. This modification has the effect of considerably reducing the calculated K ejection probabilities $w_{\text{eject}}$, particularly at high $Z$, as compared with the Intemann-Pollock results.
The reduction in the predicted intensity of the ejected-electron spectrum can be understood in the following terms (Intemann, 1975). In all calculations based on the Intemann-Pollock approach (including the one discussed in Sec. 5.3), retardation effects are neglected and the interaction between the two electrons is taken to be an instantaneous Coulomb interaction, so that the exchange of only longitudinal and scalar virtual photons can be considered. In the approximation used by Intemann and Pollock (1967), only s-wave intermediate states make a contribution to the transition amplitude, and thus, only scalar photon exchange is taken into account. In his later paper, Intemann (1972) employed the more accurate Green's function used by Glauber and Martin (1956). In this more refined calculation, p-wave intermediate states also contribute to the amplitude, and thus, longitudinal photon exchange is also being taken into account. The relative importance of longitudinal photon exchange is indicated by the extent to which the intensity of the electron spectrum is reduced (Fig. 5-1).

5.3. Relativistic Calculations of Electron Ejection

Both of the basic approaches described in Sec. 5.2 have been extended to include relativistic effects. Intemann (1969) modified the work of Intemann and Pollock (1967), using the solutions of the symmetric Hamiltonian of Biedenharn and Swamy (1964). This is a relativistic Hamiltonian with symmetry so that the radial parts of the spinor components of its solutions are formally nonrelativistic. The solutions form a complete canonical basis, and their close
correspondence to the nonrelativistic problem leads to substantial computational simplifications. The Biedenharn Hamiltonian differs from the exact Dirac-Coulomb Hamiltonian by a precisely known fine-structure term; the eigenfunctions differ from the exact Dirac-Coulomb eigenfunctions by terms of order \((aZ)^2\).

Except for the use of semi-relativistic Coulomb eigenfunctions in the overlap integral and an appropriately modified expression for the density of final states available to the ejected electron, the calculation of Intemann (1969) follows the lines of his earlier work, i.e., the interaction between the two K electrons is treated as a perturbation along with the weak interaction, leading to an exact calculation of the electron ejection probability without the need of introducing adjustable parameters such as screening constants or effective nuclear charges. Even though relativistic effects partly cancel the reduction in \(w_{\text{ejec}}\) that arises when longitudinal photon exchange is included, the ejected-electron spectrum calculated semirelativistically by Intemann (1969) is considerably less intense than that derived from the Primakoff-Porter (1953) approach (Fig. 5-2).

An independent relativistic calculation of auto-ionization in electron-capture decay was performed by Law and Campbell (1973b), in terms of second-quantization formalism and in analogy with extensive work by the same authors on internal ionization accompanying \(\beta\) decay (Campbell et al., 1971; Campbell and Law, 1972; Law and Campbell, 1972a, 1972b, 1973a). It was, however, shown by Intemann (1974) that the model of Law and Campbell (1973b) is actually identical with that of Intemann and Pollock (1967) and Intemann
(1969), and that the large difference in the results can be traced to
the fact that Law and Campbell cut off the eigenfunction expansion
for the Coulomb Green function too soon. Law and Campbell approxi-
mated the infinite series by a few terms because it appeared to con-
verge rapidly; Intemann (1974), drawing upon an analogous calculation
by Paquette (1962), pointed out that the sum over discrete eigenstates
in the Green function expansion does indeed converge rapidly, but
that continuum states make a large contribution that cannot be
neglected.

The (historically older) alternative to the Intemann approach
for the calculation of internal ionization is the "overlap" ansatz,
used in the pioneering work of Primakoff and Porter (1953). As
indicated in Sec. 5.2, in this method one attempts to take account
of all screening and correlation effects in the initial two-electron
wave function by an adjustable parameter, viz.; the effective nuclear
charge. The calculations are simplified considerably, but it is
difficult to make a choice of the key parameter, and some arbitrariness
is bound to remain. Moreover, the near-orthogonality of the wave
functions makes the overlap integral very sensitive to the exact
form of the wave functions and to the values chosen for the effective
charges. Thus, the accuracy of the results cannot be established
_a priori_, as in the Intemann approach; on the other hand, the overlap
method does not rely on the condition $Z \gg 1$, and hence, may be
superior for very light elements.

The most recent and complete calculation based on the "overlap"
method is due to Mukoyama et al. (1973). In their formulation, Mukoyama et al. draw upon the work of Stephas (1969), who had employed an atomic matrix element calculated from analytic hydrogenic relativistic wave functions for the purpose of studying internal ionization accompanying $\beta$ decay (Stephas and Crasemann, 1967, 1971; Crasemann and Stephas, 1969). However, in their evaluation of the wave-function overlap integral, Stephas and Crasemann (1967) made an approximation that causes their expression to diverge at low electron momenta, where most electrons are ejected; thus, the result cannot meaningfully be integrated to compute total electron-ejection probabilities (Isozumi and Shimizu, 1971; Kitahara et al., 1972; Nagy et al., 1972). Mord (1972, 1973) and, independently, Mukoyama et al. (1973) have calculated the atomic matrix element by alternative techniques and derived a result that is exact, within the limitations stated above; it agrees in the non-relativistic limit with the formulae of Primakoff and Porter (1953) and Stephas and Crasemann (1971).

The screening constants $\sigma$ that determine the effective nuclear charge $Z_{\text{eff}} = Z - \sigma$, to take account of electron-electron interaction, are determined by Mukoyama et al. (1973) in the following manner. In the parent atom, they take

\[ \sigma = Z(1 - \frac{\bar{r}_Z}{\bar{r}_{\text{SCF}}}), \quad (5-12) \]

where $\bar{r}_Z$ is the mean value of $r$ determined from the relativistic hydrogenic wave functions, and $\bar{r}_{\text{SCF}}$ is $\bar{r}$ from relativistic self-consistent field wave functions, as computed by Carlson et al. (1970).
For the continuum electron, Mukoyama et al. use the same screening constant as for the bound electron to be ejected. They take account of the fact that a vacancy resulting from electron capture is present in the daughter atom by reducing $\sigma$ from Eq. (5-12) by the ratio of the appropriate Slater screening constant for an atom that is ionized in an inner shell to that for a neutral atom (Slater, 1930).

The total K-electron ejection probabilities per K capture, calculated by Mukoyama et al. (1973), agree with those of Intemann (1969, 1974) as well as could be expected, given the uncertainties in the choice of screening parameters (Table 5.1).

Excitation to a bound state ("shakeup") of the second K electron, while the first one is captured, has also been computed by Mukoyama et al. (1973). Such calculations are important for comparison with experiments, in which double K x-ray emission is measured. The main difficulty here is to make adequate provision for omitting occupied final states to which shakeup is forbidden by the Pauli principle. Mukoyama et al. (1973) find that the probability for double K-vacancy production (including excitation), just as the K electron-ejection probability, is reduced when relativistic effects are included, compared with the nonrelativistic results of Primakoff and Porter (1953) (Table 5.2).

5.4. Electron Ejection from Higher Shells

It was first emphasized by Wolfsberg (1954) that a spectrum of electrons ejected during nuclear electron capture, measured in coincidence with a single K x ray, contains contributions from L
electrons shaken off during K capture and from K electrons ejected during L capture. Wolfsberg evaluated these effects in terms of the Primakoff-Porter formalism. Internal ionization of this type, resulting in K and L\textsubscript{1} vacancies, has also been discussed by Law and Campbell (1973\textsuperscript{a,b}). The energy distribution of K electrons ejected during nuclear electron capture from higher shells was considered by Ryde \textit{et al}. (1953).

The subject has been extensively treated in terms of the wavefunction overlap approach by Mukoyama and Shimizu (1974). Starting with the formalism of Stephas (1969), but using the relativistic hydrogenic atomic matrix element of Mukoyama \textit{et al}. (1973), these workers have computed the probability per K capture for L\textsubscript{1}-shell electron ejection with total energy W:

\[
\left. P_{\text{Ki}} (W) \right| \text{dW} = \frac{n_{\text{L}}^2}{2\pi} \left| M_{\text{Ki}} \right|^2 \frac{S(W - W_{\text{K}}) (W_{\text{K}} - W)^2}{S(W)} \frac{\text{p} \text{dW}}{W_{\text{K}}^2}.
\]

Here, \( W_{\text{K}} \) is the transition energy for K capture, \( W_{\text{K}} \) is the maximum total energy of the ejected electron, and \( n_{\text{L}} \) is the number of electrons in the L\textsubscript{1} shell. \( S \) is the shape factor, and the wavefunction overlap integral is

\[
M_{\text{Ki}} = \langle \psi(Z-1, W) | \cdot \psi(Z, L_{\text{1}}) \rangle.
\]

Similarly, Mukoyama and Shimizu have computed the K-shell internal ionization probability per L\textsubscript{1} capture, expressed as a ratio to the K-capture probability:
\[ p_{iK}(W) dW = \frac{n_i}{2\pi^2} \frac{e_i}{e_K} |M_{iK}|^2 \frac{S(W_2 - W)}{S(W_1 - W)} \frac{(W_2 - W)^2}{W_1^2} pW dW, \tag{5-15} \]

where \( e_i / e_K \) is the \( L_i \)-to-\( K \) capture ratio, and \( W_1' \) is the mass difference between initial and final nuclei, minus the \( L_i \) binding energy, plus one (in units of \( mc^2 \)). The atomic matrix element is

\[ M_{iK} = \langle \psi(Z-1,W) | \psi(Z,K) \rangle. \tag{5-16} \]

The authors construct a properly antisymmetrized expression for the total probability for the direct and exchange processes (5-13) and (5-15) and evaluate the result for cases of practical interest (Table 5.1). It is predicted that the \( L \)-shell internal-ionization probability accompanying \( K \) capture is of almost the same order of magnitude as the \( K \) ejection probability during \( L \) capture. The probability that the atom undergoing electron capture and internal ionization is left with holes in the \( K \) and \( L \) shells increases with \( Z \), relative to the double \( K \)-hole production probability. The \( L \)-shell ionization probability decreases more slowly with \( Z \) than the \( K \)-electron ejection probability, per \( K \) capture.

Calculated spectra of electrons ejected during \( K \) and \( L \) capture of \(^{55}\)Fe are shown in Fig. 5-2. It is predicted that electrons ejected from the \( L_1 \) shell contribute substantially over the entire spectrum.

Comparable calculations of \( L \)-shell internal ionization accompanying \( L \) capture have been carried out by Mukoyama \textit{et al.} (1974).
In this context, it should be noted that only allowed transitions have so far been treated by the Intemann-Pollock approach. By contrast, because of its simplicity, the overlap-integral approach has led to results for arbitrary beta transitions. The simplifying feature of this approach is the assumption that the initial state of the two electrons involved in the process is describable in terms of an independent-particle model, i.e., the two-electron wave function can be written as an uncorrelated product of one-electron wave functions. It is this assumption which permits the factorization of the matrix element. For forbidden transitions, however, with the entrance of higher beta moments, it is to be expected that the amplitude for internal ionization will be more sensitive to the details of the structure of the initial electronic configuration, and therefore the overlap-integral approach will be less reliable. On the other hand, relativistic effects, which are of particular importance for forbidden transitions, are much more easily included in this approach than in the Intemann-Pollock approach.

Furthermore, in connection with all wave-function overlap calculations, on which the most extensive predictions of internal-ionization probabilities are based, it must be kept in mind that near-orthogonality makes the atomic matrix element exceedingly sensitive to the accuracy of the wave functions. This point is discussed in detail in Sec. 2.5. It is likely that quantitative results derived from hydrogenic wave functions may lack in accuracy, particularly in the case of outer shells.
5.5. Measurements of Internal Ionization

Excellent critical reviews of experimental work on internal ionization and excitation accompanying electron capture have been compiled by Law and Campbell (1973), Mukoyama et al. (1973), Freedman (1974), and Walen and Briand (1975); somewhat earlier results have been discussed by Stephan (1969).

Experiments on shakeup and shakeoff during electron capture are made difficult a priori by the fact that the probability of these processes is much lower, perhaps by an order of magnitude, than in $\beta$ decay: the effect of the sudden increase in nuclear charge upon the Coulomb field seen by the atomic electrons is, to a considerable extent, compensated by the reduction in screening that ensues when one K electron is captured. Consequently, the experimental information on the subject is quite limited; it is confined to the five isotopes with simple ground-state-to-ground-state decays listed in Tables 5.1 and 5.2, and to some recent work on $^7$Be (Mutterer, 1970).

Relatively least difficult are measurements of the probability of double K-vacancy production through the detection of coincidences between two K x rays (or K'Auger electrons, or both). Two decades ago, Charpak (1953) used two 2π proportional counters for such measurements on $^{55}$Fe. Langevin (1957, 1958) measured the K Auger-electron sum peak in a single proportional counter with a gaseous internal $^{71}$Ge source. Miskel and Perlman (1954) and Kiser and Johnston (1959) measured K Auger electrons and K shakeoff electrons from $^{37}$Ar in a proportional counter.
Upon the advent of NaI(Tl) scintillation detectors, these were employed in several measurements (Daniel et al., 1960; Lark and Perlman, 1960; Ryde et al., 1963; Smith, 1964). A further advance in the technique was made possible when solid state detectors were developed with which K x rays from elements with adjoining atomic numbers can be resolved, so that one can discriminate sensitively against impurities. Nagy et al. (1972) used a Si(Li) semiconductor detector in coincidence with a scintillation counter in double K-vacancy production measurements on $^{131}\text{Cs}$ and $^{165}\text{Er}$.

The creation of double K holes can also be determined by detecting radiative transitions to the empty K shell. Such transitions produce Kα x-ray "hypersatellites" that are shifted up in energy with respect to the diagram line. A hypersatellite measurement was first used by Oërtzen (1964), who employed a bent-crystal diffraction spectrometer to determine the double K-vacancy production rate in $^{71}\text{Ge}$; the result agrees extremely well with that of Langevin (1957, 1958). Briand et al. (1971) measured the Kα hypersatellite from $^{71}\text{Ge}$ decay in coincidence with the ensuing Kα$_{3,4}$ satellite.

Results of all these measurements of double K-vacancy production probability during nuclear K capture are included in Table 5.2.

Total electron ejection probabilities are much more difficult to determine. Spectrum measurements necessarily have a low-energy threshold, determined by detector noise, electron scattering, and window transmission problems. Because most electrons are ejected with very low energies (Fig. 5-3), total ejection probabilities can only
be inferred from measured spectra, that extend over a limited range, by fitting the data to some theoretical spectral shape. The admixture of L electrons ejected during K capture, and of K electrons ejected during L capture, introduces additional uncertainties that are difficult to account for, unless the electron counts are gated by double K x-ray events. The results depend so heavily on the theoretical model in terms of which the data are interpreted and often contain such large probable errors that they have not been included in Table 5.1. Pertinent information can be found in the original literature and in the papers by Stephas (1969), Mukoyama et al. (1973), Freedman (1974), and Walen and Briançon (1975).

While ejected-electron spectrum measurements have not, in the past, led to unequivocal and precise determinations of the total electron ejection probability, they are nevertheless of value for testing theoretically predicted spectrum shapes. The \(^{7}\text{Be}\) electron spectrum has been measured by Mutterer (1970), and that of \(^{37}\text{Ar}\), by Miskel and Perlman (1954), with proportional counters. Pengra and Crasso (1963) gated on Mn K x rays, detected with a scintillation counter, to measure the \(^{55}\text{Fe}\) electron spectrum with a proportional counter, at low energies, and with an early solid state detector, at higher energies. Modern measurements of the \(^{55}\text{Fe}\) electron spectrum have been performed by Nagy (1971) with two plastic scintillators in coincidence, and by Kitahara and Shimizu (1975), who performed a triple-coincidence (x-x-\(^{5}\)) experiment with proportional counters. The \(^{71}\text{Ge}\) spectrum was determined by Langevin (1958) with a proportional counter. Daniel et al. (1960) used a
magnetic spectrometer to study the spectrum from $^{131}$Cs; this spectrum was measured more recently by Sujkowski et al. (1973) with a Si(Li) detector placed at the focus of a zero-dispersion homogeneous magnetic-field spectrometer. A magnetic $\beta$-ray spectrometer was used by Ryde et al. (1963) on $^{165}$Er.

The measured spectra appear to agree, within errors, with the general shape that all theories predict; this shape is largely determined by the statistical factor. Without question, precise absolute measurements of ejected-electron spectra, preferably in coincidence with two $K$ x rays, would be of great value as a guide for more refined computations of the atomic matrix element.

5.6. Correlation of X Rays and $\gamma$ Rays Following Electron Capture

If aligned nuclei undergo electron capture, the atomic inner-shell vacancies created thereby can be polarized, and subsequently emitted x rays can be circularly polarized (for an illustrative example, see Emery, 1975). Dolginov (1956-1957, 1958a, b) has described these circular polarization effects and pointed out that even in the decay of unaligned atoms a correlation can exist between the circular polarization of x rays and of $\gamma$ rays emitted following the nuclear decay.

An anisotropic directional correlation of the type

$$W(\theta) = 1 + \lambda_2 P_2 (\cos \theta)$$

(5-17)

can exist between x rays and $\gamma$ rays emitted after nuclear electron
capture if the intermediate atomic state is characterized by a vacancy with $j > l/2$. The theory has been developed by Dolginov (an early discussion of the problem is given by Tol'ko et al., 1958b). Somewhat simplified expressions based on Dolginov's theory are given by Rupnik and Crasemann (1972), who also worked out the directional-correlation function for x rays emitted in transitions to the $L_3$ level and $\gamma$ rays, following second-forbidden nonunique electron capture transitions.

The experimental detection of anisotropic x-$\gamma$ correlations is hampered by the condition that the intermediate atomic vacancy must have $j > l/2$, whence only $L_3$ capture is of interest. The $L_3/L_1$ capture-ratio in allowed transitions is always small ($< 10^{-7}$); one must choose a radioisotope that decays through a second or higher forbidden electron-capture transition to a short-lived excited state of the daughter. The only readily available isotope that fulfills these requirements is $^{207}$Bi, but its decay scheme is cluttered with other transitions. Efforts to detect anisotropy in the x-$\gamma$ directional correlation from $^{207}$Bi decay have been unsuccessful (Rupnik and Crasemann, 1972; Cambiaggio et al., 1975), although the results are not inconsistent with theoretical predictions.
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Appendix

A2.1. Expressions for $M_K(k_x,k_x)$ and $m_K(k_x,k_x)$

Only the dominant terms of the quantities $M_K(k_x,k_x)$ and $m_K(k_x,k_x)$ are given in Eqs. (2-106). The complete formulae for $k_x^{(1)}$ and $k_x^{(2)}$ follow (Behrens and Bühring, 1971, modified for electron capture).

1. $k_p = k_x^{(1)} = K-k_x+1$

\[
M_x(k_x,k_x) = \sqrt{\frac{K!}{(2K+1)!!(2K+1)!!}} \frac{1}{(k_x-1)!((k_x-1)!!(k_x-1))} \left( qR \right)^{k_x-1} \left( qR \right)^{k_x-1} \times \sum_{\lambda=0}^{\infty} \frac{(2\lambda-1)!!}{(2\lambda+1)!!(2\lambda+1)!!} \left( qR \right)^{\lambda} \sum_{\sigma=0}^{\infty} \left( -1 \right)^{\lambda+\sigma} \left( \mu \sigma \right) \\
\times \sum_{p=0}^{2\sigma-1} \left( m_R \right)^{2\sigma-2} \left( W_R \right)^{2\sigma-2} (2\eta\sigma+2) \left( - \sqrt{\frac{2K+1}{K}} \right) \left( \frac{2\sigma}{\rho} \right) F_{Kx-1}^{+1}(k_x,2\mu,2\sigma,\rho) \\
- \frac{1}{2K+1} \left( 2\lambda+2K+1 \right) qR \left( 2\sigma \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho) \\
+ \left[ \frac{1}{2\lambda+2K+1} W_R \left( 2\sigma+1 \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho) \right] \\
\times \left[ \frac{1}{2\lambda+2K+1} W_R \left( 2\sigma+1 \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho) \right] \times \left( - \sqrt{\frac{K+1}{2K+1}} \right) \\
+ \frac{1}{2\lambda+2K+1} qR \left( 2\sigma \rho \right) F_{Kx-1}^{+1}(k_x,2\mu,2\sigma,\rho) \times \left( - \sqrt{\frac{K+1}{2K+1}} \right) \\
\times \frac{2}{2\lambda+2K+1} qR \left( 2\sigma \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho) \\
\times \frac{1}{2\lambda+2K+1} qR \left( 2\sigma \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho) \\
\times \frac{1}{2\lambda+2K+1} qR \left( 2\sigma \rho \right) F_{Kx-1}^{+1}(k_x,2\mu+1,2\sigma+1,\rho)
\]
\[
m_{k}(k, k_\nu) = \sqrt{\frac{K!}{(2K+1)!!(2k_\nu-1)!!(2k_\nu-1)!!(k_\nu-1)!!(k_\nu-1)!}} \left(\frac{\rho R}{\rho}\right)^{k_\nu-1} (qR)^{k_\nu-1} \\
\times \sum_{j=0}^{\infty} \sum_{\mu=0}^{\infty} \frac{(2k_\nu-1)!!}{(2\mu+2k_\nu-1)!!} \frac{(2\mu-1)!!}{(2\mu+2k_\nu-1)!!} \frac{1}{(2\mu+2k_\nu-1)!!} q^j R^{j+1} \left(\frac{\mu}{\rho}\right) \\
\times \sum_{\sigma=0}^{2\sigma} (m_\nu R)^{2\sigma} (W_\nu R)^{2\sigma-\rho} (\alpha Z)^{\rho} \left(-\sqrt{\frac{K+1}{K}} \left[m_\nu R (W_\nu R)^{-1} A \left(\frac{2\sigma-1}{\rho}\right) \right] \\
\times F_{KK-11}^* (k_\nu, 2\mu, 2\sigma-1, \rho) - \frac{1}{2K+1} \frac{1}{2\mu+2k_\nu+1} \frac{1}{2\mu+2k_\nu+1} q R m_\nu R \left(\frac{2\sigma}{\rho}\right) \\
\times F_{KK-11}^* (k_\nu, 2\mu, 2\sigma, \rho) \right] + \left[ \frac{1}{2\mu+2k_\nu+1} m_\nu R \left(\frac{2\sigma}{\rho}\right) F_{KK-11}^* (k_\nu, 2\mu, 2\sigma-1, \rho) \right] - \sqrt{\frac{K+1}{K}} \\
\times \left(\frac{2\sigma-1}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu, 2\sigma-1, \rho) \right] - \sqrt{\frac{K+1}{K}} \frac{2}{2K+1} \frac{1}{2\mu+2k_\nu+1} \frac{1}{2\mu+2k_\nu+1} q R m_\nu R \\
\times \left(\frac{2\sigma}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu, 2\sigma, \rho) \right)
\]

2. \( k_\nu = k_\nu (2) = K-k_\chi+2 \)

\[
\sum_{\sigma=0}^{2\sigma} \left[ \left(\frac{2\sigma}{\rho}\right) F_{KK-10}^* (k_\nu, 2\mu, 2\sigma, \rho) \right] + \frac{1}{2\mu+2k_\nu+1} \frac{1}{2\mu+2k_\nu+1} q R W_\nu R \\
\times \left(\frac{2\sigma+1}{\rho}\right) F_{KK+10}^* (k_\nu, 2\mu, 2\sigma+1, \rho) \right] + \frac{k_\nu-k_\chi}{K+1} \sqrt{\frac{K+1}{K}} \\
\times \left[ \left(\frac{2\sigma}{\rho}\right) F_{KK-11}^* (k_\nu, 2\mu, 2\sigma, \rho) \right] + \frac{1}{2\mu+2k_\nu+1} \frac{1}{2\mu+2k_\nu+1} q R W_\nu R \\
\times \left(\frac{2\sigma+1}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu, 2\sigma+1, \rho) \right] + \sqrt{\frac{1}{K(2K+1)}} \\
\times \left[ 2(k_\nu-1) \frac{1}{2\mu+2k_\nu+1} W_\nu \left(\frac{2\sigma+1}{\rho}\right) F_{KK-11}^* (k_\nu, 2\mu+1, 2\sigma+1, \rho) \right] \\
\times \left[ (k_\nu-1) \frac{1}{2\mu+2k_\nu+1} W_\nu \left(\frac{2\sigma+1}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu+1, 2\sigma+1, \rho) \right] \\
\times \left[ (k_\nu-1) \frac{1}{2\mu+2k_\nu+1} W_\nu \left(\frac{2\sigma+1}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu+1, 2\sigma+1, \rho) \right] \\
\times \left[ (k_\nu-1) \frac{1}{2\mu+2k_\nu+1} W_\nu \left(\frac{2\sigma+1}{\rho}\right) F_{KK+11}^* (k_\nu, 2\mu+1, 2\sigma+1, \rho) \right]
\]

(2-A3)
In these expressions, we have
\[ \Delta = \begin{cases} 
0 & \text{if } \mu = 0 \\
0 & \text{if } \sigma = 0 \\
0 & \text{if } \rho = 0 \\
1 & \text{otherwise}
\end{cases} \]

For \( n \)th-forbidden unique transitions, Eqs. (2-A1) through (2-A4) apply, with \( K = n + 1, A_F^{(n+1)nl}, A_F^{(n+1)(n+1)0}, V_F^{(n+1)(n+1)l} \).
For \( n \)th forbidden non-unique transitions, Eqs. (2-A1)-(2-A4) apply with \( K=n \), \( V_F^{(n+1)(n+2)} \), \( V_{F_{n(n-1)}} \), \( V_{F_{nn0}} \), \( A_{F_{nnl}} \), \( V_{F_{n(n+1)}} \). If \( n=1 \), there is a further contribution from Eqs. (2-A3)-(2-A4) with \( K=0 \), \( A_{F_{000}} \), \( A_{F_{011}} \). Allowed transitions involve Eqs. (2-A3) and (2-A4) with \( K=0 \), \( V_{F_{000}} \), \( V_{F_{011}} \) and Eqs. (2-A1) and (2-A2) with \( K=1 \), \( A_{F_{101}} \), \( A_{F_{110}} \), \( V_{F_{111}} \), \( A_{F_{121}} \). The magnitude of the various terms in Eqs. (2-A1) through (2-A4) is determined, first, by powers of the factors \( (p_R) \), \( (q_R) \), \( (\omega_R) \), \( (m_R) \), and \( (\alpha^2) \); and second, by the difference of one order of magnitude between the relativistic and nonrelativistic form-factor coefficients. Thus, the dominant terms of Eqs. (2-A1)-(2-A4) are a subset of the terms with \( \lambda=0 \), \( \mu=0 \) [Eqs. (2-104)-(2-106)]. The correction terms of the next order are

(i) terms with \( \mu=0 \), \( \lambda=0 \) which were not included in Eqs. (2-104)-(2-106);

(ii) terms with \( \mu=0 \), \( \lambda=1 \) and \( \mu=1 \), \( \lambda=0 \) corresponding to the terms with \( \mu=0 \), \( \lambda=0 \) of Eqs. (2-104)-(2-106).

All terms, however, with powers of \( m_R \) and \( \omega_R \) can usually be omitted since \( m_R \) and \( \omega_R \) are generally much smaller than \( \alpha^2 \). It should be noted that for electron capture the correction terms are important only in cases where cancellations occur among the dominant terms.
A2.2. Expansion Coefficients $I(k,m,n,p;r)$ up to Order $m=3$

The expansion coefficients $I(k,m,n,p;r)$ of the electron radial wave functions, up to order $m=3$, are as follows (Behrens and Bühring, 1971):

\begin{align*}
I(k, 1, 1, 1; r) &= (2k+1)r^{-2k-1} \int_0^\infty x^{2k} U(x) dx,
I(k, 2, 2, 2; r) &= 2(2k+1)r^{-2k-2} U(y) y^{-2k-1} \int_0^\infty x^{2k} U(x) dx dy,
I(k, 2, 1, 1; r) &= -\frac{2k+1}{2k-1} r^{-2k-1} \int_0^\infty x^{2k} U(x) dx + \frac{4k}{2k-1} r^{-2} \int_0^\infty x U(x) dx,
I(k, 2, 1, 1; r) &= +\frac{2(2k+1)}{2k-1} r^{-2k-1} \int_0^\infty x^{2k} U(x) dx - \frac{4}{2k-1} r^{-2} \int_0^\infty x U(x) dx,
I(k, 3, 3, 3; r) &= 2(2k+1)(2k+3)r^{-2k-3} \int_0^\infty x^{2k} U(x) \int_0^\infty U(y) y^{-2k-1} \int_0^\infty x^{2k} U(x) dx dy dz,
I(k, 3, 3, 2; r) &= \frac{4(2k+3)r^{-2k-2}}{2k-1} U(y) y^{-2k-1} \int_0^\infty x^{2k} U(x) dx dy + \frac{8k(2k+3)}{3(2k-1)} r^{-2k-3} \int_0^\infty y U(y) \int_0^\infty x^{2k} U(x) dx dy,
I(k, 3, 2, 2; r) &= 2(2k+3)r^{-2k-2} U(y) y^{-2k-1} \int_0^\infty x^{2k} U(x) dx dy - \frac{4(2k+3)}{2k-1} r^{-2k-3} \int_0^\infty y U(y) \int_0^\infty x^{2k} U(x) dx dy,
I(k, 3, 3, 1; r) &= \frac{4k(2k+3)}{3(2k+1)} r^{-2k-3} \int_0^\infty x^{2k+1} U(x) dx - \frac{2(2k+1)(2k+3)}{3(2k-1)} r^{-2k-1} \int_0^\infty x^{2k} U(x) dx + \frac{8k(2k+3)}{3(2k+1)(2k-1)} r^{-2} \int_0^\infty x U(x) dx,
I(k, 3, 2, 1; r) &= \frac{2(2k+3)}{2k+1} r^{-2k-2} \int_0^\infty x U(x) dx - \frac{2(2k+3)}{2k+1} r^{-2k-3} \int_0^\infty x^{2k+2} U(x) dx,
I(k, 3, 1, 1; r) &= \frac{4(k+1)(2k+3)}{2k+1} r^{-2k-3} \int_0^\infty x^{2k+2} U(x) dx - \frac{(2k+1)(2k+3)}{2k-1} r^{-2k-1} \int_0^\infty x^{2k} U(x) dx + \frac{4(2k+3)}{(2k+1)(2k-1)} r^{-2} \int_0^\infty x U(x) dx.
\end{align*}
The function $U(x)$ in these expressions is defined by

$$V(x) = -\left(\frac{aZ}{R}\right)U(x),$$  \hspace{1cm} (2-A6)

where $V(x)$ is the potential of the nuclear and atomic charge distributions.
Footnotes

1We have $A^+ = \{A_1^+, A_2^+ , A_3^+ , -A_4^+ \}$.

2The field operators are given by

$$\psi(x) = \sqrt{2} \sum_q \sum_r \{e^{-i q x} a_{r}^\dagger(q) u_{r}(q) + b_{r}^\dagger(q) v_{r}(q) e^{-i q x}\},$$

$$\bar{\psi}(x) = \sqrt{2} \sum_q \sum_r \{e^{-i q x} a_{r}^\dagger(q) \bar{u}_{r}(q) + b_{r}^\dagger(q) \bar{v}_{r}(q) e^{i q x}\};$$

$r = 1, 2$.

The $a_{r}(q)$ and $a_{r}^\dagger(q)$ are the annihilation and creation operators for a fermion of momentum $q$ and spin $r$, respectively,

$b_{r}(q)$ and $b_{r}^\dagger(q)$ are the corresponding operators for the antiparticles. The $u_{r}(q)$ and $v_{r}(q)$ are both the free-particle Dirac spinors. We have $\bar{u}_{r}(q) = u_{r}^\dagger(q) \gamma_4$ and $\bar{v}_{r}(q) = v_{r}^\dagger(q) \gamma_4$.

3We use the Dirac equation $\left(\slashed{p} - m - i\gamma_5\right) \psi = 0$ and the notation

$$\gamma_\mu = i \sigma_\mu \gamma_4, \gamma_5 = -\gamma_1 \gamma_2 \gamma_3 \gamma_4, \sigma_\mu = \frac{1}{2} (\gamma_\mu \gamma_\nu - \gamma_\nu \gamma_\mu).$$

4In the following, we use natural units $\hbar = m = c = 1$.

5We have $r = r_0$, and $d\Omega$ is the solid angle.

6The electron radial wave functions can also be calculated approximately as hydrogenic wave functions for a point nucleus of charge reduced by the appropriate Slater screening constants (Slater, 1930).

7However, for the inner shells and for medium and high atomic numbers, there is only a negligible difference between the wave functions of Mann and Waber (1973) and those from Hartree-
Fock-Slater calculations (Suslov, 1969 and 1970; Chelepov et al., 1972; Martin and Blichert-Toft, 1970) or Thomas-Fermi-Dirac calculations (Behrens and Jeannecke, 1969; Band et al., 1956, 1958).

The Fourier transform of the electron wave function \( \phi_e(r) \) is

\[
\phi_e^+(p_e) = \int e^{-ip_e \cdot r} \phi_e(r) \, dr^3.
\]

In the formulae for \( \beta^- \) and \( \beta^+ \) decay, the axial-vector coupling constant \( \lambda \) has the opposite sign [Sec. 2.1, Eq. (2-39)]. For electron capture, there are hence two ways of defining the axial-vector form-factor coefficients in terms of matrix elements and coupling constants, i.e., by using the same sign definition for \( \lambda \) as in the \( \beta^- \)-decay formulae or the same as in the \( \beta^+ \)-decay formulae. In the present work, the definition of the form-factor coefficients corresponds, as in Behrens and Jeannecke (1969), to those in \( \beta^- \) decay. Consequently, in addition to the substitution indicated in Eq. (2-40), we must replace \( A_{FKL_q} \) by \( A_{FKL_q}^* \) when going from \( \beta^+ \) decay to electron capture.

There is another possibility of going to the nonrelativistic limit. By applying the Foldy-Wouthuysen transformation on the total (nuclear plus \( \beta \)-decay) Hamiltonian, one can construct an effective V-A transition operator that can be used with nonrelativistic single-particle wave functions (Rose and Osborn, 1954; Blin-Stoyle, 1973; Konopinski, 1966).
The operators $\hat{\alpha}$ and $\gamma_5$ which appear in the relativistic matrix elements are replaced in the nuclear space by

$$\alpha \rightarrow -\frac{\not p}{M}; \quad \gamma_5 \rightarrow \frac{\sigma \cdot \not p}{M}.$$ 

This treatment of the relativistic nuclear single-particle matrix elements is fully equivalent to that described in the text.

First- and second-class currents are defined on the basis of their behavior under a $G$ operation. If we split the hadron current into first- and second-class terms (Weinberg, 1958; Kim and Primakoff, 1969),

$$J^I_\mu = J^I_\mu + J^{II}_\mu,$$

we have

$$2J^I_\mu = J^I_\mu + GJ^I_\mu G^{-1}$$

$$2J^{II}_\mu = J^{II}_\mu - GJ^{II}_\mu G^{-1},$$

and hence,

$$GJ^I_\mu G^{-1} = -J^I_\mu$$

$$GJ^{II}_\mu G^{-1} = -J^{II}_\mu.$$ 

Here, $J^I_\mu$ is a first-class element of the hadron current, and $J^{II}_\mu$ is a second-class element. The $G$ operator is defined as

$$G = C e^{-\sqrt{2} i \mathbf{T}_2},$$

where $C$ is the charge-conjugation operator, and $\mathbf{T}_2$ is the second isospin component.
12. Also used the nonrelativistic self-consistent Hartree- 
Fock-Clater potential of Heinen and Skillman (1965).

12. Note that $\chi = \chi - \chi$. A correction formula given by 
Firestone et al. (1975b) might contain an error of sign 
and therefore should not be used.

14. Form-factor coefficients not included in Table 7.10 
are related as follows to the nuclear matrix elements 
(without induced terms):

\[ V_{\text{IL}} \equiv \int \left( \frac{r}{R} \right)^2 \; ; \]

\[ V_{\text{II}} \equiv \int \left( \frac{r}{R} \right)^4 \; ; \]

\[ V_{\text{CL}} \equiv \int \frac{r^2}{R} \; ; \]

\[ V_{\text{CLII}} \equiv \int \frac{r}{R} \; ; \]

\[ A_{\text{IL}} \equiv \lambda \int \frac{r}{R} \; ; \]

\[ A_{\text{CL}} \equiv -\lambda \int \frac{r}{R} \; . \]

\[ A_{\text{CLII}} \equiv -\lambda \int \frac{r}{R} \; . \]

15. All the form-factor coefficients for electron capture and \( \beta^+ \) decay, because 
of the different decay energies. In the former case, 
the decay energy is $W_0 = W_0 + \chi$, while in the latter case 
it is $W_0$. This energy change leads to the
following correction factor (Behrens and Bühring, 1974):

\[
\frac{\lambda_k}{\lambda_{B^+}} = \frac{f_K}{f_{B^+}} \left( 1 + 2 \frac{f_T}{\lambda} \right) \left( W - \frac{1}{3} \left( 1 + \frac{\mu_1 Y_1}{W} \right) \right)
\]

Because we can assume \(|\frac{f_T}{\lambda}| < 3 \times 10^{-3}\) (Blin-Stoyle, 1973; Wilkinson, 1970; Alburger and Wilkinson, 1970; Wilkinson and Alburger, 1971; Eman et al., 1973), we obtain a correction

\[
\left| \frac{f_T}{\lambda} \right| \left( W - \frac{1}{3} \left( 1 + \frac{\mu_1 Y_1}{W} \right) \right) \leq 2 \times 10^{-3}.
\]

This value is smaller than the contributions from higher-order terms [Eqs. (2-128)-(2-131)] and from the radiative corrections.
The total integrated intensity of a two-photon spectrum, for example, is expected to be no greater than $10^{-4}$ times that of the corresponding one-photon spectrum. Two-photon $\beta$ and the directional correlation between the photons have been studied by Henhardt [1967].

The adequacy of this procedure has been questioned by Tousin and Persson [1972], but it underlies all theoretical work reported so far.

For a possible exception to this statement, see Smirnov and Batkin [1974].

Unfortunately, Møller's work is much less well-known than that of Morrison and Schiff. Thus, the theory has come to be known by the names of the latter authors. Yet, it was Møller who first envisaged $\beta$ as arising from the emission of a virtual positron, followed by its single-quantum annihilation with one of the $K$ electrons.

Winter [1957] has shown how to construct a simple classical model for radiative $K$ capture which correctly predicts the low-energy portion of the $\beta$ spectrum [Eq. (4-17)] and, to within a factor of $\ln 2$, the total radiative capture rate [Eq. (4-18)]. Neither the high-energy portion of the $\beta$ spectrum, however, nor the $\beta$ angular distribution are correctly given by the model.
Because $g_E^{\dagger}(\mathbf{r}, \mathbf{R})$ is well behaved as $\mathbf{r} \to 0$, it is unnecessary to average it over the nuclear volume.

Of greatest interest are those EC transitions for which competing positron emission is energetically impossible. Then we have $k \leq 2 - B_\omega$ and $|E| < 1$. In this case, the Green's function cannot represent a freely propagating wave. Rather, it decreases rapidly with distance from the nucleus and has a range which depends on $k$.

In this approximation, Glauber and Martin neglect the retardation factor $e^{ik \cdot \mathbf{r}}$ for photon energies $k \leq 2\alpha$. This approximation is discussed further and a calculation of the 1s-state capture spectrum of $^{37}$Ar in which this approximation is not made is given by Paquette (1962).

As pointed out in Sec. 2.2.2, the $(g_L^2 / g_K^2)$ -ratios given in Brysk and Rose (1958) deviate systematically from all other reported calculations on screened electron wave functions. However, these deviations, and the resulting uncertainties in Fig. 4-7, appear to be never greater than about 5-6%. The errors, of order $2\alpha$, associated with the results of Glauber and Martin (1956) for the 2s, 2p, 3p -spectra are always much larger (except for the special case of $^4$Be). Thus the results displayed in Fig. 4-7 are more than adequate for present purposes and, as a convenience, will be used to determine all screening corrections in Sec. 4.2 unless otherwise noted.

An excellent summary of these results is given by Schopper (1966).
The vacancy created by nuclear electron capture tends to counteract the effect of the decrease in nuclear charge from Ze to \((Z-1)e\).

For this reason, the overlap integral of Eq. (5-2) is smaller than its analogs in \(\beta^+\) decay.

The upper limit of the neutrino energy is only approximately \(W_0 + l - |E_K| - W\) as implied by the energy-conserving delta function in Eq. (5-3). The neutrino energy is reduced by the binding energy of the second K electron in the daughter atom that already contains one K hole, and increased by the additional relaxation energy of the electron cloud.

The directional correlation function for x rays from L\(_3\)-shell internal conversion of an M\(_4\) \(\gamma\) transition and a cascade \(\gamma\) ray in \(^{207}\text{Bi}\), given by Rupnik and Crasemann (1972) [their Eqs. (36) and (37)] is in error: contrary to these authors' assumption, the radial integrals cannot be factored out of the correlation expression (J. S. Geiger, private communication, 1974). New calculations are being carried out by Geiger and Ferguson (1974) and Carvalho et al. (1975).

While nuclear electron capture as a rule occurs predominantly from s states, it is interesting to note that \(\sim 97\%\) of the primary vacancies produced in the decay of \(^{202}\text{Pb}\) and \(^{205}\text{Pb}\) are in the \(L_3\) shell (Emery, 1975; Bambahen et al., 1974).
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Grigor'ev, O.I., B.S. Kuznetsov, N.S. Shimanskaia, and I.A. Iutlandov,
Grinberg, B., J.P. Brethon, F. Lagoutine, Y. Le Gallic, J. Legrand,
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TABLE 2.1. List of calculations of electron radial wave functions inside or near the nucleus.

<table>
<thead>
<tr>
<th>Reference</th>
<th>R/NR&lt;sup&gt;a&lt;/sup&gt;</th>
<th>Potential&lt;sup&gt;b&lt;/sup&gt;</th>
<th>Distribution</th>
<th>Z</th>
<th>Shells</th>
<th>Remarks</th>
</tr>
</thead>
<tbody>
<tr>
<td>Brysk and Rose (1955)</td>
<td>R</td>
<td>TFD</td>
<td>uniform</td>
<td>10-100</td>
<td>K, L</td>
<td>Results presented graphically</td>
</tr>
<tr>
<td>Band et al. (1956, 1958)</td>
<td>R</td>
<td>TFD</td>
<td>uniform</td>
<td>18-98</td>
<td>K, L</td>
<td></td>
</tr>
<tr>
<td>Brewer et al. (1951)</td>
<td>R</td>
<td>TFD</td>
<td>uniform</td>
<td>55-90</td>
<td>M</td>
<td>Every fifth atomic number is listed</td>
</tr>
<tr>
<td>Watson and Freeman (1961)</td>
<td>NR</td>
<td>HF</td>
<td>point</td>
<td>3-42</td>
<td>all</td>
<td>Analytical wave functions are used</td>
</tr>
<tr>
<td>Herman and Skillman (1963)</td>
<td>NR</td>
<td>HFS</td>
<td>point</td>
<td>2-100</td>
<td>all</td>
<td></td>
</tr>
<tr>
<td>Winter (1968)</td>
<td>NR&lt;sup&gt;c&lt;/sup&gt;</td>
<td>HF</td>
<td>point</td>
<td>3-42</td>
<td>K, L</td>
<td>L&lt;sub&gt;1&lt;/sub&gt;/K ratios only</td>
</tr>
<tr>
<td>Behrens and JHnecke (1959)</td>
<td>R</td>
<td>HF&lt;sup&gt;(Z&lt;96)&lt;/sup&gt;</td>
<td>TFD&lt;sup&gt;(Z&gt;96)&lt;/sup&gt;</td>
<td>uniform</td>
<td>1-102</td>
<td>K, L, M</td>
</tr>
<tr>
<td>Suslov (1969, 1970b)</td>
<td>R</td>
<td>NR HFS&lt;sup&gt;(Z&lt;72)&lt;/sup&gt;</td>
<td>uniform</td>
<td>2-98</td>
<td>K, L, M</td>
<td>N&lt;sub&gt;1&lt;/sub&gt;, N&lt;sub&gt;2&lt;/sub&gt;</td>
</tr>
<tr>
<td>Martin and Blichert-Toft (1970)</td>
<td>R</td>
<td>HFS&lt;sup&gt;d&lt;/sup&gt;</td>
<td>Fermi</td>
<td>5-98</td>
<td>K, L</td>
<td></td>
</tr>
<tr>
<td>Froese-Fischer (1972b)</td>
<td>NR</td>
<td>HF</td>
<td>point</td>
<td>2-86</td>
<td>all</td>
<td></td>
</tr>
<tr>
<td>Mann and Waber (1975)</td>
<td>R</td>
<td>HF</td>
<td>Fermi</td>
<td>1-102</td>
<td>all</td>
<td></td>
</tr>
</tbody>
</table>

<sup>a</sup>NR=nonrelativistic; R=relativistic.

<sup>b</sup>TFD=Thomas-Fermi-Dirac; HF=Hartree-Fock; HFS=Hartree-Fock-Slater.

<sup>c</sup>Supplementary relativistic corrections are applied to results from NR analytic wave functions of Watson and Freeman (1961) and Malli (1966).

<sup>d</sup>Nestor <em>et al.</em> (1966); Tucker <em>et al.</em> (1969); Lu <em>et al.</em> (1971).
TABLE 2.2. Calculated electron radial wave-function ratios $\frac{g_L^2}{g_K^2}$.

<table>
<thead>
<tr>
<th>Z</th>
<th>Nonrelativistic</th>
<th>Relativistic</th>
<th>Relativistic</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>HFS</td>
<td>HF</td>
<td>HFS</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>Calculated$^a$</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>Winter (1968)</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>Martin and</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>Blichert-Toft (1970)</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>Corrected</td>
</tr>
<tr>
<td>5</td>
<td>0.049</td>
<td>0.041</td>
<td>0.049</td>
</tr>
<tr>
<td>10</td>
<td>0.058</td>
<td>0.055</td>
<td>0.059</td>
</tr>
<tr>
<td>15</td>
<td>0.075</td>
<td>0.074</td>
<td>0.076</td>
</tr>
<tr>
<td>20</td>
<td>0.085</td>
<td>0.084</td>
<td>0.086</td>
</tr>
<tr>
<td>30</td>
<td>0.095</td>
<td>0.095</td>
<td>0.100</td>
</tr>
<tr>
<td>40</td>
<td>0.101</td>
<td>0.101</td>
<td>0.109</td>
</tr>
<tr>
<td>50</td>
<td>0.104</td>
<td>0.105</td>
<td>0.119</td>
</tr>
<tr>
<td>60</td>
<td>0.107</td>
<td>0.107</td>
<td>0.128</td>
</tr>
<tr>
<td>70</td>
<td>0.109</td>
<td>0.109</td>
<td>0.139</td>
</tr>
<tr>
<td>80</td>
<td>0.110</td>
<td>0.111</td>
<td>0.152</td>
</tr>
<tr>
<td>90</td>
<td>0.111</td>
<td>0.119</td>
<td>0.167</td>
</tr>
<tr>
<td>100</td>
<td>0.112</td>
<td>0.184</td>
<td>0.187</td>
</tr>
</tbody>
</table>

$^a$The parameters in the Slater exchange term (Eq. (6) of Fricke et al. (1971)) are $C=1$, $n=1$, and $m=1$.

$^b$Nonrelativistic results multiplied by a correction factor for relativistic effects.
TABLE 2.3. Calculated electron radial wave-function ratios $g_{n,l}^2 / g_{l,l}^2$.

<table>
<thead>
<tr>
<th>Z</th>
<th>Nonrelativistic</th>
<th>Relativistic</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>HFS</td>
<td>HF</td>
</tr>
<tr>
<td>15</td>
<td>0.095</td>
<td>0.075</td>
</tr>
<tr>
<td>20</td>
<td>0.132</td>
<td>0.118</td>
</tr>
<tr>
<td>25</td>
<td>0.144</td>
<td>0.134</td>
</tr>
<tr>
<td>30</td>
<td>0.148</td>
<td>0.144</td>
</tr>
<tr>
<td>40</td>
<td>0.174</td>
<td>0.172</td>
</tr>
<tr>
<td>50</td>
<td>0.194</td>
<td>0.193</td>
</tr>
<tr>
<td>60</td>
<td>0.208</td>
<td>0.208</td>
</tr>
<tr>
<td>70</td>
<td>0.218</td>
<td>0.218</td>
</tr>
<tr>
<td>80</td>
<td>0.225</td>
<td>0.225</td>
</tr>
<tr>
<td>90</td>
<td>0.231</td>
<td>0.242</td>
</tr>
<tr>
<td>100</td>
<td>0.235</td>
<td>0.245</td>
</tr>
</tbody>
</table>

Here, $g_{n,l}^2$ has been taken from the tables of Behrens and Jünecke (1969).
TABLE 2.4. Calculated electron radial wave-function ratios $\frac{\gamma_{N_1}}{\gamma_{M_1}}$.

<table>
<thead>
<tr>
<th>Z</th>
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<th></th>
<th>Relativistic</th>
<th></th>
</tr>
</thead>
<tbody>
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<td></td>
<td>HFS</td>
<td>HF</td>
<td>HFS</td>
<td>HF</td>
</tr>
<tr>
<td>35</td>
<td>0.116</td>
<td>0.094</td>
<td>0.116</td>
<td>0.094</td>
</tr>
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<td>40</td>
<td>0.162</td>
<td>0.143</td>
<td>0.163</td>
<td>0.143</td>
</tr>
<tr>
<td>45</td>
<td>0.184</td>
<td>0.168</td>
<td>0.186</td>
<td>0.185</td>
</tr>
<tr>
<td>50</td>
<td>0.203</td>
<td>0.188</td>
<td>0.205</td>
<td>0.189</td>
</tr>
<tr>
<td>60</td>
<td>0.233</td>
<td>0.224</td>
<td>0.235</td>
<td>0.225</td>
</tr>
<tr>
<td>70</td>
<td>0.237</td>
<td>0.232</td>
<td>0.245</td>
<td>0.236</td>
</tr>
<tr>
<td>80</td>
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<td>0.248</td>
<td>0.257</td>
<td>0.251</td>
</tr>
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<td>0.271</td>
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<td>0.283</td>
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</tr>
<tr>
<td>Z</td>
<td>Nonrelativistic</td>
<td></td>
<td></td>
<td>Relativistic</td>
</tr>
<tr>
<td>----</td>
<td>----------------</td>
<td>-----</td>
<td>-----</td>
<td>--------------</td>
</tr>
<tr>
<td></td>
<td>HFS</td>
<td>HF</td>
<td></td>
<td>HFS</td>
</tr>
<tr>
<td></td>
<td>Herman and Skillman (1963)</td>
<td>Froese-Fischer (1972)</td>
<td>Calculated with the codes of Fricke et al. (1971)</td>
<td>Mann and Waber (1973)</td>
</tr>
<tr>
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<td>0.186</td>
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<td>0.232</td>
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<td></td>
<td>0.252</td>
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<td>0.263</td>
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<td></td>
<td>0.266</td>
</tr>
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</table>
### Table 2.6: Calculated relativistic electron radial wave-function ratios $f_L^2 / g_L^2$.

<table>
<thead>
<tr>
<th>Z</th>
<th>TFTD</th>
<th>HFS</th>
<th>HF</th>
</tr>
</thead>
<tbody>
<tr>
<td>10</td>
<td>Brysk and Rose (1955)</td>
<td>0.001</td>
<td>0.00052</td>
</tr>
<tr>
<td>15</td>
<td>Band et al. (1956, 58)</td>
<td>0.002</td>
<td>0.00155</td>
</tr>
<tr>
<td>20</td>
<td>Behrens and Jahnke (1969)</td>
<td>0.005</td>
<td>0.00308</td>
</tr>
<tr>
<td>25</td>
<td>Suslov (1969, 1970b); Dzhelepov et al. (1972)</td>
<td>0.00515</td>
<td>0.00512</td>
</tr>
<tr>
<td>30</td>
<td>0.005</td>
<td>0.00525</td>
<td>0.00512</td>
</tr>
<tr>
<td>40</td>
<td>0.005</td>
<td>0.006751</td>
<td>0.00774</td>
</tr>
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<td>50</td>
<td>0.007</td>
<td>0.008786</td>
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<td>0.013</td>
<td>0.01545</td>
<td>0.0147</td>
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<tr>
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<td>0.022</td>
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<td>0.0244</td>
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<td>0.034</td>
<td>0.0368</td>
<td>0.0371</td>
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<td>0.0538</td>
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<td>-----------------------------</td>
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<td>0.0865</td>
</tr>
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<td>0.1201</td>
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<td>0.1661</td>
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</tbody>
</table>
TABLE 2.8. Calculated relativistic electron radial wave-function ratios $\frac{f_2^2}{g_2^2}$.

<table>
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<tr>
<th>Z</th>
<th>HFS</th>
<th>HF</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>Suslov (1969, 1970b)</td>
<td>Dzhelepov et al. (1972)</td>
</tr>
<tr>
<td>35</td>
<td>0.0078</td>
<td>0.0076</td>
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<tr>
<td>40</td>
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TABLE 2.9. Amplitudes $\beta_{p_k}$ of the bound electron radial wave functions. (After Mann and Waber, 1973 and private communication.) Columns are headed by atomic numbers Z. (Outermost electrons have been omitted.)

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### TABLE 2.10. Relations between form-factor coefficients and nuclear matrix elements in Cartesian notation. (After Behrens and Jänecke, 1969; Bühring, 1963b; Bühring and Schülke, 1965.)

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<th>Form-factor coefficient</th>
<th>Matrix element in Cartesian notation</th>
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<td>( A_{F}(0) ) ( \frac{\tilde{A}}{101} )</td>
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<td>( A_{F}(0) ) ( \frac{\tilde{A}}{011}(1,1,1,1) )</td>
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<td>$L^{th}$ forbidden nonunique</td>
<td>$V_F^{(0)}_{L100}$</td>
<td>$\left(\frac{(2L+1)!!}{L!!}\right)^{-\frac{1}{2}} \sum_{i_1 \ldots i_L} I_{1 \ldots i_L}^R \sum_{R} I(k;1,1,1;i)$</td>
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<td>$V_F^{(0)}_{L100}(k,1,1,1)$</td>
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<td>$\left(\frac{(2L+1)!!}{L!!}\right)^{-\frac{1}{2}} \sum_{i_1 \ldots i_L} I_{1 \ldots i_L}^R \sum_{R} I(k;1,1,1;i)$</td>
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<td>$V_F^{(0)}_{L11-1,1}$</td>
<td>$\frac{1}{L!} \left(\frac{(2L+1)!!}{L!!}\right)^{-\frac{1}{2}} \sum_{i_1 \ldots i_L} I_{1 \ldots i_L}^R \sum_{R} I(k;1,1,1;i)$</td>
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<td>$A_F^{(0)}_{LLL}(k,1,1,1)$</td>
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<td>$\frac{1}{L!} \left(\frac{(2L+1)!!}{L!!}\right)^{-\frac{1}{2}} \sum_{i_1 \ldots i_L} I_{1 \ldots i_L}^R \sum_{R} I(k;1,1,1;i)$</td>
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<td>$(L-1)^{st}$ forbidden unique</td>
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<td>$-\lambda \frac{1}{L!} \left(\frac{(2L-1)!!}{(L-1)!!}\right)^{-\frac{1}{2}} \sum_{i_1 \ldots i_L} I_{1 \ldots i_L}^R \sum_{R} I(k;1,1,1;i)$</td>
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Table 2.11. Exchange and overlap correction factors \( b \) for 1-electron capture and \( \alpha / \alpha \) for 1/j capture ratios, recalculated with the Hartree-Fock code of Freeze-Pitcrew. Columns labeled "*x" are computed according to the approach of Volyi; columns labeled "*y" are calculated with Scholl's formula. In both approaches, the effect of the hole in the daughter atom has been included. Asterisks in the elements for which the calculations were performed at exciting '\*x' relating factors were determined by a 6-point Lagrange interpolation procedure. (K. A. Chen, private communication).

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* Takes into account rearrangement in final state
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* Takes into account rearrangement in final state
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<th>Recalculated in this work as described in Sec. 2.5 after Bahcall</th>
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### TABLE 3.1 Methods that have been used for the determination of electron capture probabilities.

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<th>Deduced</th>
<th>Estimated accuracy of the method (percent)</th>
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<td>Source</td>
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<td>Deduced&lt;sup&gt;b&lt;/sup&gt;</td>
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<td>Miscellaneous</td>
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<sup>a</sup> The following abbreviations are used: apc, anticoincidence proportional counter; cc, cloud chamber; gm, Geiger-Müller counter; pc, proportional counter; pl, plastic scintillator; mw, multi-wire proportional counter; sc, semiconductor; sd, double-focusing spectrometer; ls, lens spectrometer.
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<th>$P_L/P_K$</th>
<th>$P_M/P_L$</th>
<th>$P_{LM+}/P_K$</th>
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<td>$P_{LM..}/P_K$</td>
<td>$P_{KwK}$</td>
<td>$P_K$</td>
<td>Method</td>
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<td>24</td>
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<td>51 751</td>
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<td>0.10 ± 0.02</td>
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<td>$\frac{7^-}{2^+}$</td>
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<td>0.1033 ± 0.0031</td>
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<td>A</td>
<td>Q_{EC}^a (keV)</td>
<td>Final state</td>
<td>( j_i ^{\pi} \rightarrow j_f ^{\pi} )</td>
<td>P_L/P_K</td>
<td>P_M/P_L</td>
<td>P_{LM...}/P_K</td>
<td>P_{Kw}^b</td>
<td>P_K</td>
<td>Method</td>
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<td>26</td>
<td>Fe</td>
<td>55 232</td>
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<td>( \frac{3}{2} ^{-} \frac{5}{2} ^{-} )</td>
<td>0.108 +0.006</td>
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<td>27</td>
<td>Co</td>
<td>57 837</td>
<td>136</td>
<td>( \frac{7}{2} ^{-} \frac{5}{2} ^{-} )</td>
<td>0.20 +0.13</td>
<td>0.254 +0.011</td>
<td>0.262 +0.008</td>
<td>0.3044+0.0043</td>
<td>0.87 +0.02</td>
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<td>136;706 ( \frac{7}{2} ^{-} \frac{5}{2} ^{-} )</td>
<td>0.099 +0.011</td>
<td>0.15 +0.02</td>
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<td>0.92 +0.03</td>
<td>11</td>
<td>Bosch(1969)</td>
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<td>136 ( \frac{7}{2} ^{-} \frac{5}{2} ^{-} )</td>
<td>0.088+0.040</td>
<td>0.317 +0.006</td>
<td>0.317 +0.006</td>
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<td>11</td>
<td>Mukerji(1973)</td>
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<td>28</td>
<td>Ni</td>
<td>56 2133</td>
<td>1720</td>
<td>( 0^+ )</td>
<td>0.115 +0.006</td>
<td>0.115 +0.006</td>
<td>0.100 +0.006</td>
<td>0.121 +0.002</td>
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<td>28</td>
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<td>( E_{\text{EC}} ) (keV)</td>
<td>( J_i \rightarrow J_f )</td>
<td>( P_L/P_K )</td>
<td>( P_M/P_L )</td>
<td>( P_{LM}/P_K )</td>
<td>( P_K )</td>
<td>Method</td>
<td>Reference</td>
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<tr>
<td>30</td>
<td>Zn</td>
<td>65 1350</td>
<td>1115 ( \frac{3}{2} \rightarrow \frac{5}{2} )</td>
<td>0.369 ±0.023</td>
<td>1115;0</td>
<td>0.13 ±0.002</td>
<td>7 Konstantinov(1961)</td>
<td>7</td>
<td>Perrin(1960)</td>
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<td></td>
<td>1115;0 ( \frac{3}{2} \rightarrow \frac{5}{2} )</td>
<td>0.119 ±0.007</td>
<td>7</td>
<td>Santos-Ocampo(1962)</td>
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<td>1115;0 ( \frac{3}{2} \rightarrow \frac{5}{2} )</td>
<td>0.111 ±0.006</td>
<td>7</td>
<td>Taylor(1963)</td>
<td>10</td>
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<td>Totzek(1967)</td>
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<td>1115;0 ( \frac{3}{2} \rightarrow \frac{5}{2} )</td>
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<td>1115</td>
<td>0.400 ±0.006</td>
<td>10 Taylor(1963)</td>
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<td>1115</td>
<td>0.386 ±0.010</td>
<td>11 Hammer(1968)</td>
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<td>7</td>
<td>Konstantinov(1961)</td>
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<td>1115;0 ( \frac{3}{2} \rightarrow \frac{5}{2} )</td>
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<td>7</td>
<td>Manduchi(1962)</td>
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<td>32</td>
<td>Ge</td>
<td>71 235</td>
<td>0 ( \frac{1}{2} \rightarrow \frac{3}{2} )</td>
<td>0.30 ±0.02</td>
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<td>0 ( \frac{1}{2} \rightarrow \frac{3}{2} )</td>
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<td>Drevor(1959)</td>
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<td>33</td>
<td>As</td>
<td>73 340</td>
<td>67 ( \frac{3}{2} \rightarrow \frac{1}{2} )</td>
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<td>Final state (keV)</td>
<td>( P_L/P_K )</td>
<td>( P_M/P_L )</td>
<td>( P_{LM.}/P_K )</td>
<td>( P_K )</td>
<td>Method</td>
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<td>74 2564</td>
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<td>( 5^+ \rightarrow 2^- )</td>
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<td>0.085 ( \pm 0.020 )</td>
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<td>34 Se</td>
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<td>( \frac{5}{2} )</td>
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<td>0.460 ( \pm 0.004 )</td>
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<td>0.462 ( \pm 0.012 )</td>
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<td>0.516 ( \pm 0.021 )</td>
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<td>36 Kr</td>
<td>79 1631</td>
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<td>( 7^+ \rightarrow 3^- )</td>
<td>( 7^+ \rightarrow 3^- )</td>
<td>0.27 ( \pm 0.09 )</td>
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<td>0.26 ( \pm 0.03 )</td>
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<td>0.108 ( \pm 0.005 )</td>
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<td>81 290</td>
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<td>( \frac{7}{2} )</td>
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<td>0.146 ( \pm 0.005 )</td>
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<td>37 Rb</td>
<td>83 1038 571;562</td>
<td>( \frac{5}{2} )</td>
<td>( \frac{3}{2} )</td>
<td>( \frac{3}{2} )</td>
<td>0.121 ( \pm 0.002 )</td>
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<td>571</td>
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<td>0.128 ( \pm 0.002 )</td>
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<td>0.580 ( \pm 0.025 )</td>
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<td>0.5959 ( \pm 0.0035 )</td>
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</table><p>ightarrow j^T_f$ | $P_L/P_K$ | $P_M/P_L$ | $P_{LM}/P_K$ | $P_{w}^{10}$ | $P_K$ | Method | Reference        |
|----|----|---------------|-------------------|-----------------------|-----------|-----------|-------------|-------------|-------|--------|-----------------|
| 39 | Y  | 3619          | 2734              | $4^{-3}$              | 0.629 ±0.0032 |          |           |             |             |       | 11     | Groothed (1969) |
|    |    | 2734;1836     |                   | $4^{-3}$              | 0.613 ±0.004   |          |           |             |             |       | 10     | Barnbynek (1973)|
| 42 | Mo | 398           | 30;0              | $5^+;\frac{1}{2};\frac{3}{2}$ | 0.36 ±0.04   |          |           |             |             |       | 6      | Hoffman (1964)  |
| 43 | Tc | 97            | 347               | $0^+;\frac{3}{2};\frac{5}{2}$ | 0.21 ±0.14    | -0.10    |           |             |             |       | 6      | Katcoff (1958)  |
| 45 | Rh | 101           | 554;27           | $1^{-};\frac{1}{2};\frac{3}{2}$ | 0.65          |           |           |             |             |       | 11     | Perrin (1960)   |
|    |    | 101           | 87.7              | $1^{-};\frac{3}{2};\frac{5}{2}$ | 0.56 ±0.7     |           |           |             |             |       | 8      | Avignon (1953)  |
|    |    | 101           | 87.7              | $0^+;\frac{1}{2};\frac{3}{2}$ | 0.95          |           |           |             |             |       | 9      | Avignon (1955)  |
| 47 | Ag | 105           | 1341              | $\frac{1}{2};\frac{1}{2};\frac{3}{2}$ | 0.128 ±0.003  |          |           |             |             |       | 5      | Schulz (1967d)  |
|    |    | 1088          |                   | $\frac{1}{2};\frac{3}{2};\frac{5}{2}$ | 0.152 ±0.002  |          |           |             |             |       | 5      | Schulz (1967d)  |
| 48 | Cd | 109           | 182;87.7         | $\frac{1}{2};\frac{3}{2};\frac{5}{2}$ | 0.32 ±0.04    |          |           |             |             |       | 6      | Bertolini (1954) |</p>

* $P_{w}^{10}$ is a notation for $P_{w}$ with 10% uncertainty.
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<th>P_{M}/P_{L}</th>
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a Final QEC state
b NAK

REFERENCE PAGE 374
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Note: \( Q_{EC} \) values are approximate.
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$^a$ $E_{EC}$ values are given in keV.

$^b$ Final states are given in terms of the quantum numbers $J$ and $\pi$.

$^c$ Method of measurement.

References:
- Hanson(1968)
- Ravn(1971)
- Bisi(1959, 1954)
- Goebbrod(1964)
- De Wit(1965)
- Harris(1965)
- Jasin(1968)
- Govers(1973)
- Gupta(1958a)
- De Wit(1965)
- Pich(1971)
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<td>( J^m_{1} - J^m_{1} )</td>
<td>( P_L/P_K )</td>
<td>( P_M/P_L )</td>
<td>( P_{LM-}/P_K )</td>
<td>( P_{MK} )</td>
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<td>94</td>
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<td></td>
<td></td>
<td>( \frac{7}{2} ) ( - \frac{5}{2} )</td>
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<td>97</td>
<td>Bk</td>
<td>245</td>
<td>819</td>
<td>250</td>
<td>( \frac{3}{2} ) ( - \frac{3}{2} )</td>
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<td>0.74</td>
<td>±0.03</td>
<td>11</td>
<td>Magnusson(1956)</td>
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</table>

**TABLE 3.2 (continued)**

Kalkstein(1957)

Hoffman(1958)

Magnusson(1956)
\( Q_{EC} \) values are taken from Wapstra and Gove (1971). There are some values that originate from electron capture measurements. They are replaced by values obtained from other methods, except for a few cases, indicated by an asterix, where no recent other result is available.

If \( P_K \) is given, the fluorescence yield used by the authors was used to calculate the measured value \( P_K \omega_K \). There are some cases in which \( \omega_K \) is not quoted. They are indicated by the sign "+1".

Methods are identified by numbers explained in Table 3.1.

\( Q_{EC} \) value from Bertrand (1974).

Revised value using \( k_B/k_\alpha=0.212 \) (Salem et al., 1974) and \( \omega_K=0.832 \) (see Table III.V).

Value revised by the author, private communication to Durosinmi-Etti (1966).

Revised value using \( k_B/k_\alpha=0.217 \) (Salem et al., 1974), and \( \omega_K=0.852 \) (see Table III.V).

\( Q_{EC} \) value from Henry (1974).

\( Q_{EC} \) value from Berényi (1970).

\( Q_{EC} \) value from Kroger (1973).

\( Q_{EC} \) very uncertain.

\( Q_{EC} \) value from Ellis (1973).

\( Q_{EC} \) value from Jasinski (1963b).

\( Q_{EC} \) value from Auble (1971a).

\( Q_{EC} \) value from Auble (1971b).

\( P_{MN...P_L} \) value.

\( Q_{EC} \) value from Gopinathan (1968).

\( Q_{EC} \) value from Ford (1970).
<table>
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<th>Z</th>
<th>A</th>
<th>Q_{EC} (keV)</th>
<th>Final state (keV)</th>
<th>J_1^\pi - J_f^\pi</th>
<th>P_L/P_K</th>
<th>(q_{L_1}/q_K)^2</th>
<th>Me-b</th>
<th>Method</th>
<th>Reference</th>
<th>( P_L/P_K )</th>
<th>( (q_{L_1}/q_K)^2 )</th>
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<td>0.101 ±0.008</td>
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<td>0.102 ±0.003</td>
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<td>Santos-Camp (1960)</td>
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<td>0.0965±0.0005</td>
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<td>Manduchi(1961)</td>
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<td>0.102 ±0.003</td>
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<td>0.101 ±0.003</td>
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<td></td>
<td></td>
<td>0.098±0.003</td>
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<td>0.097±0.003</td>
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<td>Totzek(1967)</td>
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<td>0.097±0.002</td>
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<td>4^+ - 4^+</td>
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<td>0.103±0.004</td>
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<td>0.115±0.015</td>
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<td>0.114±0.015</td>
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<td>Bertmann(1972)</td>
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<td>1.06±0.004</td>
<td>1.015</td>
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<td>0.101</td>
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<td>Q_{EC}^a (keV)</td>
<td>Final state (keV)</td>
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<td>P_L/P_K</td>
<td>(g_{L1}/g_{K})^2_k</td>
<td>P_L/P_K</td>
<td>(g_{L1}/g_{K})^2</td>
<td>Mz-b</td>
<td>Method</td>
<td>Reference</td>
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<td>24</td>
<td>Cr 51</td>
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<td>320</td>
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<td>0.1044±0.0021</td>
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<td>0.1021±0.0021</td>
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<td></td>
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<td>7^-5^-/2^-2^-</td>
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<td>1.022</td>
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<td>1.008</td>
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<td>0.107</td>
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<td>$Q_{EC}^\alpha$ (keV)</td>
<td>$J^\Pi_{f}$</td>
<td>$P_L/P_K$</td>
<td>$(q_{L_1}/q_K)^2$</td>
<td>Method</td>
<td>Reference</td>
<td>$P_L/P_K^c$</td>
<td>$(q_{L_1}/q_K)^2$</td>
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<td>30</td>
<td>Zn</td>
<td>65 1350.7±1.1</td>
<td>$^{5/2}_{-5/2}^{+3/2}$</td>
<td>1115;0</td>
<td>0.111 ±0.006</td>
<td>0.106 ±0.006</td>
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<td>1115;0</td>
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<td>0.109 ±0.008</td>
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<td>1115;0</td>
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<td>1.043</td>
<td>0.113 ±0.003</td>
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<td>McCann (1968)</td>
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<td>0.109 ±0.008</td>
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<td>0.120 ±0.003</td>
<td>1.071</td>
<td>0.112 ±0.003</td>
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<td>Kraftt (1970)</td>
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<td>0.109 ±0.008</td>
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<td>0.112</td>
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<td>0.108 ±0.005</td>
<td>1.017</td>
<td>0.106 ±0.005</td>
<td>1</td>
<td>Drever (1959)</td>
<td>0.115</td>
<td>0.113</td>
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<td>37</td>
<td>Rb</td>
<td>83 1038 ±32</td>
<td>571;562</td>
<td>$^{5}<em>{2}^{+3}</em>{2}$</td>
<td>1.056</td>
<td>0.115 ±0.002</td>
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<td>Schulz (1967a)</td>
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<td>0.115</td>
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<td></td>
<td>571</td>
<td>0.128 ±0.002</td>
<td>1.056</td>
<td>0.121 ±0.002</td>
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<td>562</td>
<td>0.132 ±0.002</td>
<td>1.056</td>
<td>0.125 ±0.002</td>
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<td>Goedbloed (1970b)</td>
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### TABLE 3.3 (continued)

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<th>$Q_{EC}$ (keV)</th>
<th>Final state</th>
<th>$J^\pi_1 - J^\pi_T$</th>
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<th>$(q_{L_1}/q_K)^2$</th>
<th>$P_L/P_K$</th>
<th>$(q_{L_1}/q_K)^2$</th>
<th>Method</th>
<th>Reference</th>
<th>Bahcall</th>
<th>Vata</th>
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<td>Cd</td>
<td>109 182.0±3.0</td>
<td>88 $\frac{5}{2}^+ - \frac{3}{2}^+$</td>
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<td>55</td>
<td>Ca</td>
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<td>0 $\frac{5}{2}^- - \frac{3}{2}^-$</td>
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<td>0.129 ±0.007</td>
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<td>Joshi(1960)</td>
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<tr>
<td>56</td>
<td>Ba</td>
<td>131 1340 ±19</td>
<td>696;620 $\frac{1}{2}^+ - \frac{1}{2}^+$</td>
<td>0.135 ±0.009 1.091</td>
<td>0.124 ±0.008</td>
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<td>Smith(1963)</td>
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<tr>
<td>56</td>
<td>Ba</td>
<td>133 515.8±3.0</td>
<td>437 $\frac{1}{2}^- - \frac{1}{2}^-$</td>
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<td>Schulz(1967c)</td>
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</table>

Note: $P_L/P_K$ and $(q_{L_1}/q_K)^2$ are the experimental and theoretical values respectively.
### TABLE 3.3 (continued)

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<th>Z</th>
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<th>$Q_{EC}^a$ (keV)</th>
<th>Final state $(\hbar J_i J_f^{\pm})$</th>
<th>$P_L/P_K^c$</th>
<th>$P_{L_1}/P_{K_1}^c$</th>
<th>$\left(q_{L_1}/q_{K_1}\right)^2$</th>
<th>$\left(q_{L_1}/q_{K_1}\right)^2$</th>
<th>Method</th>
<th>Reference</th>
<th>Bahcall Vatai</th>
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<td>37 Rb</td>
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<td>Schulz(1967a)</td>
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<td>0.115</td>
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<td></td>
<td></td>
<td>0.119±0.002</td>
<td>0.117±0.002</td>
<td>5</td>
<td>Goedbloed(1970b)</td>
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<td>47 Ag</td>
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<td>1341</td>
<td>$1^-1^+$</td>
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<td>Schulz(1967d)</td>
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<td></td>
<td>0.152±0.002</td>
<td>0.137±0.002</td>
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<td>Scobe(1958)</td>
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<td>53 I</td>
<td>126</td>
<td>2151</td>
<td>several</td>
<td>0.142±0.005</td>
<td>0.137±0.005</td>
<td>5</td>
<td>Goverse(1974a)</td>
<td>0.133</td>
<td>0.131</td>
<td></td>
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<td></td>
<td></td>
<td></td>
<td>0.136±0.001</td>
<td>0.130±0.002</td>
<td>5</td>
<td>Goverse(1974a)</td>
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<td>Leiper(1971)</td>
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</table>

**First non-unique forbidden transitions** $\Delta J=0, 1$; $\pi \hbar J = -1$
<p>| Z  | A | Q_{EC}^a (keV) | Final state | J_f^π | J_i^π | P_L/P_K | P_L/P_K \left( \frac{q_{L_1}}{q_K} \right)^2 | Method | Reference            | Babcall | Vatai |
|----|---|----------------|-------------|--------|--------|---------|-----------------------------|--------|---------------------|
| 76Os | 185 | 1015.0±0.7 | 874;880   | \frac{1}{2}^-; \frac{1}{2}^+; \frac{1}{2}^- | 0.600±0.006 | 3.62 | 0.166±0.007 | 5 | Schulz(1967a) | 0.162 | 0.160 |
| 646 | \frac{1}{2}^-; \frac{1}{2}^+ | 0.228±0.004 | 1.438 | 0.160±0.003 | 5 | Schulz(1967a) | 0.168 | 0.165 |
| 79Au | 195 | 229.0±1.0^m | 99 | \frac{3}{2}^+; \frac{3}{2}^- | 0.873±0.044 | 5.047 | 0.173±0.009 | 5 | Goverse(1973) | 0.163 | 0.166 |
| 130 | \frac{3}{2}^+; \frac{3}{2}^- | 3.055±0.086 | 16.74 | 0.183±0.008 | 5 | Goverse(1973) | 0.165 | 0.168 |
| 0 | \frac{3}{2}^+; \frac{3}{2}^- | 0.337±0.007 | 2.040 | 0.165±0.003 | 5 | Goverse(1973) | 0.168 | 0.165 |</p>
<table>
<thead>
<tr>
<th>Z</th>
<th>A</th>
<th>$Q_{EC}$ (keV)</th>
<th>$^\text{a}$Final state (keV)</th>
<th>$J_1^\pi$</th>
<th>$J_2^\pi$</th>
<th>$P_L/P_K$</th>
<th>$(q_{L_1}/q_K)^2$</th>
<th>$P_L/P_K$</th>
<th>$(q_{L_1}/q_K)^2$</th>
<th>Method</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>81</td>
<td>Tl 202</td>
<td>1372 $^{+22}_{-22}$</td>
<td>440 $^{2-2+}_{-2-2+}$</td>
<td>0.196 $^{+0.002}_{-0.002}$</td>
<td>1.167 $^{+0.002}_{-0.002}$</td>
<td>0.168 $^{+0.002}_{-0.002}$</td>
<td>5</td>
<td>Leutz (1966)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>960 $^{2-2+}_{-2-2+}$</td>
<td>0.305 $^{+0.020}_{-0.017}$</td>
<td>1.458 $^{+0.017}_{-0.017}$</td>
<td>0.209 $^{+0.014}_{-0.014}$</td>
<td>5</td>
<td>Leutz (1966)</td>
<td></td>
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</tr>
<tr>
<td>83</td>
<td>Bi 206</td>
<td>3652 $^{+25}_{-25}$</td>
<td>3279 $^{6+5-}_{6+5-}$</td>
<td>0.264 $^{+0.010}_{-0.010}$</td>
<td></td>
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<tr>
<td></td>
<td></td>
<td></td>
<td>3403 $^{6+5-}_{6+5-}$</td>
<td>0.281 $^{+0.009}_{-0.009}$</td>
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<tr>
<td></td>
<td></td>
<td></td>
<td>3563 $^{6+5-}_{6+5-}$</td>
<td>0.509 $^{+0.015}_{-0.015}$</td>
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<td>17</td>
<td>Cl 36</td>
<td>1144.1$^{+1.7}_{-1.7}$</td>
<td>0 $^{2+0+}_{2+0+}$</td>
<td>0.112 $^{+0.008}_{-0.008}$</td>
<td>1</td>
<td>Doughan (1962b)</td>
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<tr>
<td>28</td>
<td>Ni 59</td>
<td>1073.1$^{+1.1}_{-1.1}$</td>
<td>0 $^{\frac{3}{2}-\frac{1}{2}}_{\frac{3}{2}-\frac{1}{2}}$</td>
<td>0.121 $^{+0.002}_{-0.002}$</td>
<td>1</td>
<td>Chew (1974a)</td>
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<tr>
<td>42</td>
<td>Mo 93</td>
<td>398 $^{+4}_{-4}$</td>
<td>30;0 $^{\frac{5}{2}+1-\frac{3}{2}}_{\frac{5}{2}+1-\frac{3}{2}}$</td>
<td>0.36 $^{+0.04}_{-0.04}$</td>
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<td>Hohmuth (1964)</td>
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<tr>
<td>43</td>
<td>Tc 97</td>
<td>347 $^{+9}_{-9}$</td>
<td>0 $^{\frac{5}{2}+5^+}_{\frac{5}{2}+5^+}$</td>
<td>0.21 $^{+0.14}_{-0.10}$</td>
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<td>Katcoff (1958)</td>
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</tr>
</tbody>
</table>

Second non-unique forbidden transitions $^{\Delta J=2}$ $^{\pi^\pi_{1}=1}$
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<tr>
<th>Z</th>
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<th>$Q_E$ (keV)</th>
<th>Final state</th>
<th>$J_{II}$</th>
<th>$J_{II}^T$</th>
<th>$P_L/P_K$</th>
<th>$(q_{L1}/q_K)^{4}$</th>
<th>$(q_{L1}/q_K)^{4}$</th>
<th>Method</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>36</td>
<td>Kr</td>
<td>81</td>
<td>290</td>
<td>+100</td>
<td>$0 \ - \ 1^{+}$</td>
<td>+0.005</td>
<td>1.179</td>
<td>0.124</td>
<td>+0.006</td>
<td>1 Chew(1974b)</td>
</tr>
<tr>
<td>53</td>
<td>I</td>
<td>125</td>
<td>1251</td>
<td>+5</td>
<td>several</td>
<td>+0.005</td>
<td>1.071</td>
<td>0.133</td>
<td>+0.005</td>
<td>5 Scobie(1958)</td>
</tr>
<tr>
<td>81</td>
<td>Tl</td>
<td>202</td>
<td>1372</td>
<td>+22</td>
<td>$0 \ - \ 2^{+}$</td>
<td>+0.02</td>
<td>1.230</td>
<td>0.179</td>
<td>+0.016</td>
<td>5 Leutz(1966)</td>
</tr>
<tr>
<td>81</td>
<td>Tl</td>
<td>204</td>
<td>345</td>
<td>+4</td>
<td>$0 \ - \ 2^{+}$</td>
<td>+0.05</td>
<td>2.256</td>
<td>0.17</td>
<td>+0.02</td>
<td>5 Joshi(1961)</td>
</tr>
</tbody>
</table>

*First unique forbidden transitions* \( \Delta J=2 \); \( \pi_{II}^T=-1 \)

- $P_L/P_K$ and $(q_{L1}/q_K)^4$ are theoretical values.
- $Q_E$ is the experimental value in keV.
- Reference: Bahcall and Vatai.
$Q_{EC}$ values are taken from Wapstra and Gove (1971). There are some values that originates from electron capture measurements. They are replaced by values obtained from other methods, except for a few cases, indicated by an asterisk where no recent other result is available.

Methods are identified by numbers explained in Table 3.1.

The theoretical $L/K$ ratios are derived from wave functions of Mann and Wabes (1973) and exchange and overlap corrections $X_{L/K}$ as described in Sec. 2.5. For $Z > 54$ the correction factors of Suslov (1970) are used in continuation of the Bancal factors and those of Martin and Bichert-Toft (1970), in extension of the recalculated Vateli factors.

$Q_{EC}$ value from Henry (1974).

$Q_{EC}$ value from Bertzand (1974).

Revised value; see Grinberg et al. (1973).

$Q_{EC}$ value from Aubie (1971b).

The $Q_{EC}$ value is obviously too low. No reliable comparison with theoretical values can be given.

First unique forbidden transitions; the factor $(q_{L1}/q_K)^2$ has been squared and in the theoretical values the contribution of the $L_3$ shell has been allowed for.
### TABLE 3.4 Experimental and theoretical $P_M/P_L$ ratios

<table>
<thead>
<tr>
<th>$Z$</th>
<th>$A$</th>
<th>$Q_{EC}$ (keV)</th>
<th>$Q_{EC}$ (keV)</th>
<th>$J^\pi_1,J^\pi_2$</th>
<th>$P_M/P_L$</th>
<th>$P_M/P_L$</th>
<th>$P_M/P_L$</th>
<th>$P_M/P_L$</th>
<th>Reference</th>
<th>Method</th>
<th>Technique</th>
<th>Reference</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>18</td>
<td>Ar</td>
<td>814.1±0.6</td>
<td>$\frac{3}{2}^+,-\frac{1}{2}^+$</td>
<td>0.104±0.007, -0.003</td>
<td>1.000</td>
<td>1.000</td>
<td>1.000</td>
<td>1.000</td>
<td>Renier(1968)</td>
<td>0.130</td>
<td>0.116</td>
<td></td>
<td></td>
</tr>
<tr>
<td>26</td>
<td>Fe</td>
<td>231.7±0.7</td>
<td>$\frac{3}{2}^-,-\frac{1}{2}^-$</td>
<td>0.157±0.003, 1.006</td>
<td>0.156±0.003</td>
<td>1.006</td>
<td>0.156</td>
<td>1.006</td>
<td>Pengra(1972)</td>
<td>0.163</td>
<td>0.156</td>
<td></td>
<td></td>
</tr>
<tr>
<td>30</td>
<td>Zn</td>
<td>1350.7±1.1</td>
<td>$\frac{5}{2}^-,-\frac{3}{2}^-$</td>
<td>0.153±0.020, 1.008</td>
<td>0.152±0.020</td>
<td>1.008</td>
<td>0.152</td>
<td>1.008</td>
<td>Krafft(1970)</td>
<td>0.167</td>
<td>0.160</td>
<td></td>
<td></td>
</tr>
<tr>
<td>32</td>
<td>Ge</td>
<td>235.1±1.7</td>
<td>$\frac{1}{2}^-,-\frac{1}{2}^-$</td>
<td>0.142±0.010, 1.010</td>
<td>0.141±0.010</td>
<td>1.010</td>
<td>0.141</td>
<td>1.010</td>
<td>Manduchi(1962b)</td>
<td>0.170</td>
<td>0.164</td>
<td></td>
<td></td>
</tr>
<tr>
<td>48</td>
<td>Cd</td>
<td>182.0±3.0</td>
<td>$\frac{5}{2}^-,-\frac{3}{2}^+$</td>
<td>0.205±0.020, 1.070</td>
<td>0.192±0.019</td>
<td>1.070</td>
<td>0.192</td>
<td>1.070</td>
<td>Moler(1965)</td>
<td>0.206</td>
<td>0.202</td>
<td></td>
<td></td>
</tr>
</tbody>
</table>

**Allowed transitions** $\Delta J = 0, 1$; $\pi_1\pi_2 = +1$

| 50  | Sn  | 1025 ±15       | $\frac{1}{2}^-,-\frac{1}{2}^+$ | 0.220±0.010, 1.011 | 0.218±0.010 | 1.011      | 0.218      | 1.011      | Manduchi(1964b) | 0.209  | 0.205      |            |            |
| 76  | Os  | 1015.0±0.7     | $\frac{1}{2}^-,-\frac{1}{2}^+$ | 0.254±0.005, 1.055 | 0.241±0.005 | 1.055      | 0.241      | 1.055      | Schulz(1967a) | 0.245   | 0.236      |            |            |

**First non-unique forbidden transitions** $\Delta J = 0, 1$; $\pi_1\pi_2 = -1$
### Table 3.4 (continued)

<table>
<thead>
<tr>
<th>Z</th>
<th>A</th>
<th>Q_EC (^a) (keV)</th>
<th>Final state (J^\pi_f)</th>
<th>(J^\pi_f - J^\pi_i)</th>
<th>(P_M/P_L)</th>
<th>(P_M/P_L) (\left(\frac{q_{M_1}/q_{L_{11}}}{q_{M_2}/q_{L_{21}}}\right)^2)</th>
<th>Me-(^b)</th>
<th>Method</th>
<th>Reference</th>
<th>P_M/P_L (\left(\frac{q_{M_1}/q_{L_{11}}}{q_{M_2}/q_{L_{21}}}\right)^2)</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>78</td>
<td>Pt</td>
<td>193</td>
<td>61.2(^\pm)3.0</td>
<td>0</td>
<td>(\frac{1}{2}^-)(^\pm)(\frac{3}{2}^+)</td>
<td>0.386 (+)0.014</td>
<td>1.475</td>
<td>0.262(+)0.010</td>
<td>5 Ravn(1971)</td>
<td>0.247</td>
<td>0.239</td>
</tr>
<tr>
<td>81</td>
<td>Tl</td>
<td>202</td>
<td>1372</td>
<td>0</td>
<td>(\frac{1}{2}^-)(^\pm)(\frac{3}{2}^+)</td>
<td>0.269 (+)0.007</td>
<td>1.025</td>
<td>0.262(+)0.007</td>
<td>5 Leutz(1966)</td>
<td>0.249</td>
<td>0.240</td>
</tr>
<tr>
<td>83</td>
<td>Bi</td>
<td>206</td>
<td>3652</td>
<td>0</td>
<td>(\frac{1}{2}^-)(^\pm)(\frac{5}{2}^+)</td>
<td>0.228 (+)0.007</td>
<td>0.228 (+)0.007</td>
<td>5 Goverse(1974b)</td>
<td>0.250</td>
<td>0.242</td>
<td></td>
</tr>
</tbody>
</table>

\(^a\) From T. Gockenbach, 1967. \(^b\) Average of available literature data.

Theor. values

\[
\frac{P_M}{P_L} \left(\frac{q_{M_1}/q_{L_{11}}}{q_{M_2}/q_{L_{21}}}\right)^2
\]

Reference Bahcall Vatai.
\( Q_{EC} \) values are taken from Wapstra and Gove (1971).

Methods are identified by numbers explained in Table 3.1.

The theoretical \( M/L \) ratios are determined from wave functions of Mann and Waber (1973) and exchange and overlap corrections \( X^{M/L} \) as described in Sec. 2.5. For \( Z > 54 \) the correction factors of Suslov (1970) are used in continuation of the Bahcall factors and those of Martin and Blichert-Toft (1970) in extension of the recalculated Vatái factors.

Revised value using \( k_B/k_\alpha = 0.212 \) (Salem et al. 1974) and \( w_K = 0.832 \) (see Table 3.5).

Revised value using \( k_B/k_\alpha = 0.217 \) (Salem et al. 1974) and \( w_K = 0.852 \) (see Table 3.5).

\( Q_{EC} \) value from Auble (1971b).

The \( Q_{EC} \) value is obviously too low. No reliable comparisor with theoretical values can be given.
### Table 3.5: Experimental and theoretical $P_K$ values

<table>
<thead>
<tr>
<th>Z</th>
<th>A</th>
<th>$\Omega_{EC}$ (keV)</th>
<th>Final state</th>
<th>$J_1^\Pi - J_2^\Pi$</th>
<th>$\omega_K^b$</th>
<th>$P_{PLM} / P_K$</th>
<th>$P_{K^2}$</th>
<th>Method</th>
<th>Reference</th>
<th>$P_K$</th>
<th>Theor. $P_K$</th>
</tr>
</thead>
<tbody>
<tr>
<td>23</td>
<td>V</td>
<td>48.4</td>
<td>4015.4</td>
<td>several</td>
<td>$0.2005 \pm 0.030$</td>
<td>$0.225 \pm 0.009$</td>
<td>$0.891 \pm 0.036$</td>
<td>11</td>
<td>Albrecht(1975)</td>
<td>0.892</td>
<td>0.896</td>
</tr>
<tr>
<td>24</td>
<td>Cr</td>
<td>59</td>
<td>751.4</td>
<td>$320 \ T^{-5} \ t^{-2}$</td>
<td>$0.227 \pm 0.003$</td>
<td>$0.256 \pm 0.007$</td>
<td>$0.887 \pm 0.008$</td>
<td>10</td>
<td>Taylor(1965)</td>
<td>0.890</td>
<td>0.893</td>
</tr>
<tr>
<td>25</td>
<td>Mn</td>
<td>54</td>
<td>1374.9</td>
<td>$835 \ T^{-3} \ t^{-2}$</td>
<td>$0.257 \pm 0.004$</td>
<td>$0.283 \pm 0.007$</td>
<td>$0.908 \pm 0.008$</td>
<td>10</td>
<td>Taylor(1965)</td>
<td>0.889</td>
<td>0.891</td>
</tr>
<tr>
<td>27</td>
<td>Co</td>
<td>57</td>
<td>836.9</td>
<td>$136 \ T^{-5} \ t^{-2}$</td>
<td>$0.3044 \pm 0.0043$</td>
<td>$0.344 \pm 0.008$</td>
<td>$0.885 \pm 0.009$</td>
<td>11</td>
<td>Rubinson(1968)</td>
<td>0.887</td>
<td>0.890</td>
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</table>

*Allowed transitions $\Delta J=0, \Delta I=1$; $\Pi_{IF}^*=-1$*
<table>
<thead>
<tr>
<th>Z, A</th>
<th>Q_{EC}^a (keV)</th>
<th>E_{final} (keV)</th>
<th>( J_1^\pi - J_2^\pi )</th>
<th>Experimental values</th>
<th>Theor. ( P_K ) values^c</th>
<th>Method</th>
<th>Reference</th>
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</tr>
<tr>
<td>27 Co</td>
<td>58</td>
<td>2308.0</td>
<td>1675; 2^+ - 2^+; 2^+</td>
<td>0.3050 ± 0.0022, 0.3444 ± 0.008</td>
<td>0.8873 ± 0.008</td>
<td>10 Banbynek(1968b)</td>
<td>0.887</td>
</tr>
<tr>
<td>30 Zn</td>
<td>65</td>
<td>1350.7</td>
<td>1115; 5^- - 5^-</td>
<td>0.441 ± 0.009</td>
<td>0.8783 ± 0.006</td>
<td>13 Kramer(1962a)</td>
<td>0.882</td>
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<td>33 As</td>
<td>73</td>
<td>340</td>
<td>67</td>
<td>3^- - 1^-</td>
<td>0.3927 ± 0.0026</td>
<td>0.890 ± 0.009</td>
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<td>34 Se</td>
<td>75</td>
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<td>2^+ - 5^-</td>
<td>0.386 ± 0.010</td>
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<td>11 Mukerji(1973)</td>
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<tr>
<td>37 Rb</td>
<td>83</td>
<td>1038</td>
<td>562</td>
<td>5^- - 3^-</td>
<td>0.400 ± 0.006</td>
<td>0.907 ± 0.011</td>
<td>10 Taylor(1965)</td>
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</table>

^a Q_{EC} = \text{Total Energy Content} 

^b \( \omega \) = \text{Angular Distribution Coefficient} 

^c Theor. \( P_K \) = Theoretical Probability of \( K \) Capture 

<table>
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</thead>
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<table>
<thead>
<tr>
<th>Reference</th>
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TABLE 3.5 (continued)
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<th>$Z$</th>
<th>$A$</th>
<th>$\Omega_{EC}$ (keV)</th>
<th>$\Omega_{EC}$ (keV)</th>
<th>$J_{1}^{M}$</th>
<th>$J_{2}^{M}$</th>
<th>$P_{LM} / P_K$</th>
<th>$P_{K}^{m}$</th>
<th>$\alpha_{K}$</th>
<th>$P_{K}$</th>
<th>Method</th>
<th>Reference</th>
<th>Bahcall</th>
<th>Vatai</th>
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<tr>
<td>58</td>
<td>Cd</td>
<td>275</td>
<td>$165 \frac{5}{2}^{+}$</td>
<td>$\frac{3}{2}^{+}$</td>
<td>$\frac{5}{2}^{+}$</td>
<td>0.37 $\pm$ 0.02</td>
<td>0.73 $\pm$ 0.01</td>
<td>15 Ketelle(1956)</td>
<td>0.72 4 $\pm$ 0.014</td>
<td>0.729 $\pm$ 0.014</td>
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<td>0.68 $\pm$ 0.02</td>
<td>17 Marelus(1967)</td>
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<td>0.75 $\pm$ 0.01</td>
<td>16 Adamowicz(1968)</td>
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<td>0.69 $\pm$ 0.02</td>
<td>13 Vatai(1968a)</td>
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<td></td>
<td>0.707 $\pm$ 0.018</td>
<td>0.906 $\pm$ 0.02</td>
<td>0.78 $\pm$ 0.03</td>
<td>11 Schmidt-Ott(1972)</td>
<td>0.724 $\pm$ 0.011</td>
<td>0.729 $\pm$ 0.011</td>
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<td></td>
<td></td>
<td>0.649 $\pm$ 0.017</td>
<td>0.716 $\pm$ 0.03</td>
<td>14 Campbell(1972)</td>
<td>0.645 $\pm$ 0.015</td>
<td>0.726 $\pm$ 0.011</td>
<td></td>
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<td></td>
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<td></td>
<td></td>
<td></td>
<td>0.639 $\pm$ 0.006</td>
<td>0.705 $\pm$ 0.03</td>
<td>11 Pich(1975)</td>
<td>0.645 $\pm$ 0.015</td>
<td>0.726 $\pm$ 0.011</td>
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<tr>
<td>64</td>
<td>Gd</td>
<td>151</td>
<td>352</td>
<td>$\frac{7}{2}$</td>
<td>$\frac{3}{2}$</td>
<td>0.664 $\pm$ 0.009</td>
<td>0.930 $\pm$ 0.015</td>
<td>0.74 $\pm$ 0.017</td>
<td>11 Genz(1973c)</td>
<td>0.704 $\pm$ 0.015</td>
<td>0.709 $\pm$ 0.015</td>
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<td></td>
<td></td>
<td></td>
<td>0.68 $\pm$ 0.06</td>
<td>0.946 $\pm$ 0.02</td>
<td>0.72 $\pm$ 0.06</td>
<td>14 Jasinski(1963a)</td>
<td>0.711 $\pm$ 0.011</td>
<td>0.715 $\pm$ 0.011</td>
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<tr>
<td>70</td>
<td>Yb</td>
<td>166</td>
<td>82</td>
<td>$\frac{2}{2}$</td>
<td>$\frac{3}{2}$</td>
<td>0.68 $\pm$ 0.06</td>
<td>0.946 $\pm$ 0.02</td>
<td>0.72 $\pm$ 0.03</td>
<td>14 Jasinski(1963a)</td>
<td>0.711 $\pm$ 0.011</td>
<td>0.715 $\pm$ 0.011</td>
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<td></td>
</tr>
<tr>
<td>81</td>
<td>Tl</td>
<td>201</td>
<td>167</td>
<td>$\frac{1}{2}$</td>
<td>$\frac{1}{2}$</td>
<td>0.67 $\pm$ 0.04</td>
<td>0.964 $\pm$ 0.01</td>
<td>0.70 $\pm$ 0.04</td>
<td>11 Gupta(1960)</td>
<td>0.722 $\pm$ 0.014</td>
<td>0.726 $\pm$ 0.014</td>
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</table>

*Note:* The table continues with similar entries for other elements.
### Table 3.5 Experimental and theoretical $P_K$ values

<table>
<thead>
<tr>
<th>Z</th>
<th>A</th>
<th>Q_{EC} (keV)</th>
<th>Final state (keV)</th>
<th>$J_i^-J_f^+$</th>
<th>$P_{LM}/P_K$</th>
<th>$P_{Kw}^wK$</th>
<th>$w_K^b$</th>
<th>$P_K$</th>
<th>Method</th>
<th>Reference</th>
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<tbody>
<tr>
<td>37</td>
<td>Rb</td>
<td>84 2679.8</td>
<td>$^2_2$</td>
<td>$^2_2$</td>
<td>0.580 ±0.025</td>
<td>0.653±0.030</td>
<td>0.888±0.039</td>
<td>11 Welker(1955)</td>
<td>0.876</td>
<td>0.878</td>
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<tr>
<td>61</td>
<td>Pm</td>
<td>145 170</td>
<td>$^5_2$ $^3_2$</td>
<td>$^5_2$</td>
<td>0.558 ±0.022</td>
<td>0.919±0.024</td>
<td>0.607±0.033</td>
<td>11 Tolea(1974)</td>
<td>0.676</td>
<td>0.681</td>
</tr>
<tr>
<td>62</td>
<td>Sm</td>
<td>145 647</td>
<td>$^7_2$ $^1_2$</td>
<td>$^7_2$</td>
<td>0.509 ±0.022</td>
<td>0.554±0.033</td>
<td>11 Tolea(1974)</td>
<td>0.660</td>
<td>0.665</td>
<td></td>
</tr>
<tr>
<td>64</td>
<td>Gd</td>
<td>151 484</td>
<td>$^3_2$ $^7_2$</td>
<td>$^3_2$</td>
<td>0.754 ±0.014</td>
<td>0.930±0.015</td>
<td>0.811±0.021</td>
<td>11 Genz(1973c)</td>
<td>0.754</td>
<td>0.759</td>
</tr>
<tr>
<td>66</td>
<td>Dy</td>
<td>159 365</td>
<td>$^3_2$ $^5_2$</td>
<td>$^3_2$</td>
<td>0.752 ±0.024</td>
<td>0.936±0.022</td>
<td>0.803±0.033</td>
<td>11 Genz(1973c)</td>
<td>0.793</td>
<td>0.797</td>
</tr>
<tr>
<td>72</td>
<td>Hf</td>
<td>175 607</td>
<td>$^3_2$ $^3_2$</td>
<td>$^3_2$</td>
<td>0.64 ±0.04</td>
<td>0.950±0.020</td>
<td>0.67 ±0.04</td>
<td>11 Funke(1965)</td>
<td>0.689</td>
<td>0.693</td>
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</tbody>
</table>

First non-unique forbidden transitions $\Delta J=0,1; \pi,\pi^-,-1$
<table>
<thead>
<tr>
<th>Z</th>
<th>A</th>
<th>$Q_{EC}^{a}$ (keV)</th>
<th>Final state</th>
<th>$J_l^\pi - J_f^\pi$</th>
<th>P&lt;sub&gt;LM&lt;/sub&gt; / P&lt;sub&gt;K&lt;/sub&gt;</th>
<th>P&lt;sub&gt;K&lt;/sub&gt;</th>
<th>ω&lt;sub&gt;K&lt;/sub&gt;</th>
<th>P&lt;sub&gt;K&lt;/sub&gt;</th>
<th>Method</th>
<th>Reference</th>
<th>Theor. P&lt;sub&gt;K&lt;/sub&gt;</th>
<th>Values</th>
<th>Bahcall Vatal</th>
</tr>
</thead>
<tbody>
<tr>
<td>78</td>
<td>Pt</td>
<td>188</td>
<td>540</td>
<td>0&lt;sup&gt;+&lt;/sup&gt;-1&lt;sup&gt;-&lt;/sup&gt;</td>
<td>195</td>
<td>0.744±0.020</td>
<td>16</td>
<td>0.748</td>
<td>0.752</td>
<td>Hanson(1968)</td>
<td>0.766±0.023</td>
<td>16</td>
<td>0.750</td>
</tr>
<tr>
<td>79</td>
<td>Au</td>
<td>195</td>
<td>229.0</td>
<td>0&lt;sup&gt;+&lt;/sup&gt;-1&lt;sup&gt;-&lt;/sup&gt;</td>
<td>187</td>
<td>0.766±0.023</td>
<td>16</td>
<td>0.750</td>
<td>0.754</td>
<td>Hanson(1968)</td>
<td>0.766±0.023</td>
<td>16</td>
<td>0.750</td>
</tr>
<tr>
<td>80</td>
<td>Hg</td>
<td>197</td>
<td>684&lt;sup&gt;+&lt;/sup&gt;14&lt;sup&gt;+&lt;/sup&gt;</td>
<td>268;77 1&lt;sup&gt;-&lt;/sup&gt;-2&lt;sup&gt;+&lt;/sup&gt;</td>
<td>1&lt;sup&gt;-&lt;/sup&gt;-3&lt;sup&gt;+&lt;/sup&gt;</td>
<td>0.741±0.012</td>
<td>18</td>
<td>0.754</td>
<td>0.758</td>
<td>Plch(1971)</td>
<td>0.754±0.002</td>
<td>18</td>
<td>0.754</td>
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<tr>
<td>81</td>
<td>Tl</td>
<td>202</td>
<td>1372&lt;sup&gt;+&lt;/sup&gt;22&lt;sup&gt;+&lt;/sup&gt;</td>
<td>440&lt;sup&gt;+&lt;/sup&gt;-2&lt;sup&gt;+&lt;/sup&gt;</td>
<td>0.76±0.05</td>
<td>0.764±0.017</td>
<td>11</td>
<td>0.790</td>
<td>0.793</td>
<td>Hagedorn (1958)</td>
<td>0.761±0.015</td>
<td>14</td>
<td>0.789±0.019</td>
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<td></td>
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<td>0.751±0.014</td>
<td>0.751±0.014</td>
<td>14</td>
<td>0.789±0.022</td>
<td>0.799±0.034</td>
<td>Gupta(1960)</td>
<td>0.778±0.034</td>
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<td>0.789±0.019</td>
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<td></td>
<td>0.75±0.03</td>
<td>0.75±0.03</td>
<td>14</td>
<td>0.791±0.006</td>
<td>0.791±0.006</td>
<td>Leutz(1966)</td>
<td>0.791±0.006</td>
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<td>0.791±0.019</td>
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<td>$Z$</td>
<td>$A$</td>
<td>$Q_{EC}$</td>
<td>Final state</td>
<td>$J_i^\pi - J_f^\pi$</td>
<td>Experimental values</td>
<td>Theor. $P_K$ values $^c$</td>
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<tr>
<td>82</td>
<td>Pb 203</td>
<td>982</td>
<td>0.66 ±0.04</td>
<td>0.69 ±0.04</td>
<td>11 Hagedoorn (1958): 0.709 ±0.003, 0.713 ±0.003</td>
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<td>680</td>
<td>0.755±0.014</td>
<td>0.783±0.022</td>
<td>11 Hagedoorn (1958): 0.770 0.780</td>
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<td></td>
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<td>279</td>
<td>0.750±0.019</td>
<td>0.776±0.025</td>
<td>11 Persson(1961): 0.776 0.780</td>
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<tr>
<td>19</td>
<td>K 40</td>
<td>1505:1 1460; 4$^-$2$^+; 0^+$</td>
<td>0.34 ±0.08</td>
<td>0.75 ±0.05</td>
<td>5 McCann(1967): 0.741 0.749</td>
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<td></td>
<td></td>
<td>0.44 ±0.09</td>
<td>0.69 ±0.04</td>
<td>8 Ahrman(1968): 0.741 0.749</td>
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</table>

First unique forbidden transitions $\Delta J=2$; $i_i=i_f=-1$
Values are taken from Wapstra and Gove (1971).
There are some values that originate from electron capture measurements. They are replaced by values obtained from other methods, except for a few cases, indicated by an asterix, where no recent other result is available.

Fluorescence yields were calculated from the equation \[ \frac{L}{(-\omega_K)}^{1/4} = A + B\omega + C\omega^3 \]. The constants A, B, C were determined by fitting the selected "most reliable" experimental values of Bambynek et al. (1972) to this equation. We have omitted from the list of the "most reliable" values those that were deduced from PKL measurements.

The theoretical PK values were derived from wave functions of Mann and Weber (1973) and exchange and overlap corrections as described in Sec. 2.5. For \( Z > 54 \) the correction factors of Suslov (1970) are used in continuation of the Bahcall factors and those of Martin and Blichert-Toft (1970) in extension of the recalculated Vatai factors. Uncertainties are quoted only in those cases where they are significant. They originate from the uncertainties of the QEC value.

Methods are identified by numbers explained in Table 3.1.

- QEC value from Gopinathan (1968).
- QEC value from Auble (1971a).
- QEC value from Jasinski (1963b).
- QEC value from Auble (1971b).
- Theoretical value for a unique 1st forbidden transition.
<table>
<thead>
<tr>
<th>Element</th>
<th>2(E_{\text{3z}}) (keV)</th>
<th>Final state (J^\pi_i - J^\pi_f)</th>
<th>Data</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>6 C</td>
<td>6662.2 ± 1.0</td>
<td>(3^+ \rightarrow 3^+)</td>
<td>(1.9 ± 0.3) \times 10^{-3}</td>
<td>20. Scobie (1957b)</td>
</tr>
<tr>
<td>7 N</td>
<td>2220.5 ± 0.9</td>
<td>(1^- \rightarrow 1^+)</td>
<td>(1.68 ± 0.12) \times 10^{-3}</td>
<td>21. Campbell (1967)</td>
</tr>
<tr>
<td>8 O</td>
<td>2759.2 ± 0.9</td>
<td>(1^- \rightarrow 1^+)</td>
<td>(1.07 ± 0.06) \times 10^{-3}</td>
<td>21. Leiper (1972)</td>
</tr>
<tr>
<td>9 F</td>
<td>1655.3 ± 0.9</td>
<td>(1^+ \rightarrow 0^+)</td>
<td>(3.00 ± 0.18) \times 10^{-2}</td>
<td>20. Drever (1956)</td>
</tr>
<tr>
<td>10 Ne</td>
<td>3232.2 ± 0.9</td>
<td>(1^+ \rightarrow 1^+)</td>
<td>(9.6 ± 0.3) \times 10^{-4}</td>
<td>21. Leiper (1972)</td>
</tr>
<tr>
<td>11 Na</td>
<td>2842.3 ± 0.5</td>
<td>(3^+ \rightarrow 2^+)</td>
<td>0.105 ± 0.009</td>
<td>22. McCa (1969)</td>
</tr>
</tbody>
</table>

**Notes:**
- Data collected from various sources, with references provided.
- Errors are given in parentheses.
- The table represents experimental results for various elements and their corresponding ratios.

**References:**
- Scobie (1957b)
- Campbell (1967)
- Leiper (1972)
- Drever (1956)
- McCa (1969)
- See the original page for more details.
<table>
<thead>
<tr>
<th>Z, Blcment</th>
<th>$Q_{20}$ (keV)</th>
<th>Final state $J^π$</th>
<th>$F_K/F_{θ}^+$</th>
<th>$P_{E2}/P_{θ}^+$</th>
<th>Method, Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>11 Na '22</td>
<td>*0.124 ± 0.012</td>
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<td></td>
<td>27 Hagedorn(1957)</td>
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<tr>
<td>15 Al 26</td>
<td>*0.109 ± 0.008</td>
<td></td>
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<td></td>
<td>27 Konijn(1958/59)</td>
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<tr>
<td>15 Al 26</td>
<td>*0.112 ± 0.004</td>
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<td>27 Ramaswamy(1959a)</td>
</tr>
<tr>
<td>15 Al 26</td>
<td>*0.104 ± 0.0007</td>
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<td></td>
<td>28 Williams(1964, 1968)</td>
</tr>
<tr>
<td>15 Al 26</td>
<td>*0.103 ± 0.018</td>
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<td>27 Leitz(1967)</td>
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<td>15 Al 26</td>
<td>*0.1042 ± 0.0010</td>
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<td>27 Steyn(1966)</td>
</tr>
<tr>
<td>15 Al 26</td>
<td>*0.1077 ± 0.0005</td>
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<td></td>
<td>27 Vatai(1968c)</td>
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<tr>
<td>15 Al 26</td>
<td>*0.135 ± 0.023</td>
<td></td>
<td></td>
<td></td>
<td>26 MacMahon(1970)</td>
</tr>
<tr>
<td>15 Al 26</td>
<td>*0.12</td>
<td></td>
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<td></td>
<td>31 Jastram(1961)</td>
</tr>
<tr>
<td>15 Al 26</td>
<td>*0.105 ± 0.044</td>
<td></td>
<td></td>
<td></td>
<td>28 Ledingham(1971)</td>
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<tr>
<td>17 Cl 36</td>
<td>1144.1 ± 1.7</td>
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<td></td>
<td></td>
<td>31 Dougan(1962b)</td>
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<tr>
<td>21 Sc 44</td>
<td>3549 ± 6</td>
<td>several $2^+$</td>
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<td></td>
<td>25 Langevin(1954c)</td>
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<tr>
<td>21 Sc 44</td>
<td>0 ± 0.1</td>
<td>several $2^+$</td>
<td></td>
<td></td>
<td>25 Langevin(1954c)</td>
</tr>
<tr>
<td>21 Sc 44</td>
<td>0.05 ± 0.15</td>
<td>several $2^+$</td>
<td></td>
<td></td>
<td>25 Blue(1955)</td>
</tr>
<tr>
<td>Z</td>
<td>Element</td>
<td>$Q_{30}$</td>
<td>Final State</td>
<td>$P_K/P_8$</td>
<td>$P_{EC}/P_8$</td>
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<td>-------------</td>
<td>-----------</td>
<td>-------------</td>
</tr>
<tr>
<td>21</td>
<td>Sc 44</td>
<td>0.075 \pm 0.017</td>
<td>27</td>
<td>Blue(1955)</td>
<td></td>
</tr>
<tr>
<td>23</td>
<td>V 48</td>
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References:
- Blue(1955)
- Konijn(1958/59)
- Dillman(1963)
- Good(1946)
- Casson(1953)
- Bock(1955)
- Konijn(1967b) revised by Konijn(1967b)
- Hagedoorn(1957)
- Ristinen(1963)
- Biryukov(1966)
- Konijn(1967a)
- Konijn(1967b)
- Konijn(1967b)
- Albrecht(1975)
- Good(1946)
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References:
- van Patter (1964)
- Cooper (1965)
- Fitzpatrick (1975)
- Sehr (1954)
- Levi (1954)
- Monaro (1962)
- Matuszek (1963)
- Hamilton (1964)
- Medicus (1950)
- Levi (1957)
- Unik (1959)
- Cretzu (1965)
- Levi (1959)
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2.7 ± 0.4

2.2

25' Mitchell(1959)
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|        |        |               |             |                     |                 |                 |        | Friedlander (1962) |
|        |        |               |             |                     |                 |                 |        | Jha (1961) |
|        |        |               |             |                     |                 |                 |        | Severse (1974a) |
|        |        |               |             |                     |                 |                 |        | Robinson (1962) |
|        |        |               |             |                     |                 |                 |        | Taylor (1963) |

|        |        |               |             |                     |                 |                 |        | Taylor (1963) |

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$a$ From $\beta$ decay.

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<th>( P_{X}/P_{B+} )</th>
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<td>1.07 ± 0.11</td>
<td>25</td>
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<td>53</td>
<td>Eu ^{143}</td>
<td>5000 ± 200</td>
<td>( \frac{3}{2} \rightarrow \frac{3}{2} )</td>
<td>( \frac{3}{2} \rightarrow \frac{3}{2} )</td>
<td>0.62 ± 0.06^d</td>
<td>31</td>
<td>Firestone(1974)</td>
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<td>$J^\pi_1$</td>
<td>$J^\pi_2$</td>
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<td>$P_{EC}/P_{E^+}$</td>
<td>Method</td>
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<td>2720 ± 15</td>
<td>0 $\frac{5}{2}^+ \rightarrow \frac{7}{2}^-$</td>
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<td>120</td>
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<td>195 ± 50</td>
<td>302 ± 150</td>
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<td>257 $\pm$ 100</td>
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<td>$P_K/P_{\beta^+}$</td>
<td>$P_{EC}/P_{\beta^+}$</td>
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<td>$P_K/P_\beta^-$</td>
<td>$P_{EC}/P_\beta^-$</td>
<td>Method</td>
<td>Reference</td>
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<td>25, 26</td>
<td>Wilson(1969), Natovich(1974)</td>
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<td>van Nooijen(1962)</td>
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</table>

a $Q_{EC}$ values are taken from Wapstra and Gove(1971).

b Methods are identified by numbers explained in Table 143.

c Relative measurements, normalised to the transition to the 1056.6 keV state of $^{143}$Pm.

d Relative measurements, normalised to the transition to the 1107.2 keV state of $^{143}$Sm.
<table>
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<th>Final state (keV)</th>
<th>$J_1$ - $J_f$</th>
<th>$P_K/P_{\beta^+}$</th>
<th>Method</th>
<th>Reference</th>
<th>Theoretical values</th>
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<td>(1.800±0.006) $10^{-3}$</td>
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<td>(0.911±0.002) $10^{-3}$</td>
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TABLE 3.7 (continued) (b) Results for EC/\beta^+ Ratios.

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TABLE 3.7 (b) continued

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<td>Bolotin(1961)</td>
<td>830 ±80</td>
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</tbody>
</table>

$Q_{EC}$ values are taken from Wapstra and Gove (1971).

Methods are identified by numbers explained in Table 3.1.
TABLE 3.8 First Forbidden Unique Transitions

<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
<th>Q_{EC} (keV)</th>
<th>Final state (keV)</th>
<th>( J_f^{\pi} - J_i^{\pi} )</th>
<th>( P_K/P_{\beta^+} )</th>
<th>Method</th>
<th>Reference</th>
<th>1st unique forbidden values</th>
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<tbody>
<tr>
<td>37</td>
<td>Rb 84</td>
<td>2679.8±2.9</td>
<td>0</td>
<td>2^-0^+</td>
<td>1.12 ±0.25</td>
<td>31</td>
<td>Konijn(1958/59)</td>
<td>0.94 ±0.01</td>
</tr>
<tr>
<td>51</td>
<td>Sb 122</td>
<td>1610.1±3.3</td>
<td>0</td>
<td>2^-0^+</td>
<td>300 ±50</td>
<td>31</td>
<td>Perlman(1958) and Glaubman(1955)</td>
<td>254 ±11</td>
</tr>
<tr>
<td>53</td>
<td>I 126</td>
<td>2151 ±5</td>
<td>0</td>
<td>2^-0^+</td>
<td>20.2 ±2.0</td>
<td>31</td>
<td>Koerts(1955)</td>
<td>21.1 ±0.7</td>
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</table>

\( Q_{EC} \) values are taken from Wapstra and Gove (1971).

Methods are identified by numbers explained in Table 3.1.
**TABLE 3.9. First Forbidden Non-Unique Transitions**

(a) $K/\beta^+$ Ratios

<table>
<thead>
<tr>
<th>$Z$</th>
<th>Element</th>
<th>$A$</th>
<th>$Q_{EC}^{a}$ (keV)</th>
<th>$\Delta E_C$ (keV)</th>
<th>$J_1^\Pi - J_2^\Pi$</th>
<th>$P_K/P_{\beta^+}$</th>
<th>Method</th>
<th>Reference</th>
<th>$P_K/P_{\beta^+}$</th>
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<tbody>
<tr>
<td>37</td>
<td>Rb</td>
<td>84</td>
<td>2679.8±2.9</td>
<td></td>
<td>2$^-$-2$^+$</td>
<td>5.15 ±0.38</td>
<td>31</td>
<td>Welker(1925)</td>
<td>3.51 ±0.06</td>
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<td></td>
<td></td>
<td></td>
<td></td>
<td>3.96 ±0.16</td>
<td>22</td>
<td>Goedbloed(1970c)</td>
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</tr>
<tr>
<td>53</td>
<td>I</td>
<td>126</td>
<td>2151 ±5</td>
<td></td>
<td>2$^-$-2$^+$</td>
<td>95 ±10</td>
<td>31</td>
<td>Koerts(1955)</td>
<td>138 ±7</td>
</tr>
<tr>
<td>63</td>
<td>Eu</td>
<td>145</td>
<td>2720 ±15</td>
<td></td>
<td>$\frac{5^+}{2}$ - $\frac{7^-}{2}$</td>
<td>3.0 ±0.5</td>
<td>31</td>
<td>Muziol(1966)</td>
<td>3.39 ±0.14</td>
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<tr>
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<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>100 ±20</td>
<td>31</td>
<td>Avotina(1965a)</td>
<td>43.9 ±4.2</td>
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<td></td>
<td></td>
<td>70 ±9</td>
<td>31</td>
<td>Muziol(1966)</td>
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</tr>
<tr>
<td>63</td>
<td>Eu</td>
<td>147</td>
<td>1762 ±9</td>
<td></td>
<td>$\frac{5^+}{2}$ - $\frac{3^-}{2}$</td>
<td>160 ±30</td>
<td>31</td>
<td>Avotina(1966)</td>
<td>197 ±16</td>
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<tr>
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<td></td>
<td></td>
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<td>31</td>
<td>Avotina(1966)</td>
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<td></td>
<td>165 ±30</td>
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<td>Muziol(1966)</td>
<td>119 ±8</td>
</tr>
<tr>
<td>55</td>
<td>Cs</td>
<td>132</td>
<td>2099 ±23</td>
<td></td>
<td>2$^-$-2$^+$</td>
<td>53.5 ±8.9</td>
<td>22</td>
<td>Goverse(1974)</td>
<td>264 ±71</td>
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<tr>
<td>81</td>
<td>Tl</td>
<td>200</td>
<td>2454 ±5</td>
<td></td>
<td>2$^-$-2$^+$</td>
<td>1.10 ±1.0</td>
<td>31</td>
<td>Konijn(1960)</td>
<td>65.7 ±1.4</td>
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<tr>
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<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>1.02 ±0.9</td>
<td>27</td>
<td>van Nooijen(1962)</td>
<td></td>
</tr>
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<td>$Q_{\text{EC}}$ (keV)</td>
<td>$J_i^\pi - J_f^\pi$</td>
<td>$P_{\text{EC}}/P_{\beta^+}$</td>
<td>Method</td>
<td>Reference</td>
<td></td>
<td></td>
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<tr>
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<td>------------------------</td>
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<td>-----------</td>
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</tr>
<tr>
<td>33</td>
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<td>596</td>
<td>$2^- - 2^+$</td>
<td>1.32 ±0.14</td>
<td>31</td>
<td>Grigor'ev(1958a)</td>
<td></td>
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</tr>
<tr>
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<td>Vatai(1968c)</td>
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<td>Konijn(1958a)</td>
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<td></td>
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<tr>
<td>53</td>
<td>I</td>
<td>2151 ±5</td>
<td>667</td>
<td>$2^- - 2^+$</td>
<td>165 ±5</td>
<td>29</td>
<td>Harmer(1959)</td>
<td></td>
<td></td>
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</tbody>
</table>

$Q_{\text{EC}}$ values are taken from Wapstra and Gove (1971).

Methods are identified by numbers explained in Table 3.1.
II. (a) The function $\lambda_s (E)$ given by Eq. (4-24), and the associated relativistic correction factor $\lambda_s (E)$, for various values of the photon energy $E$, measured in units of the $s$-shell binding energy $R_s = (2\pi)^2/\alpha$.

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<th>$\lambda_s (E)_{\text{rel}}$</th>
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<td>0.5463</td>
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<td>0.3175</td>
<td>0.5504</td>
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<td>0.5622</td>
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<td>0.5733</td>
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<td>0.6043</td>
</tr>
<tr>
<td>$k$</td>
<td>$\mathbb{E}_{18}(k)$</td>
<td>$\mathbb{R}_{18}(k)$</td>
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<td>-----------------</td>
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<td>3.0</td>
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<td>0.6635</td>
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<td>11.0</td>
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<td>12.0</td>
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<td>0.7083</td>
</tr>
<tr>
<td>θ</td>
<td>( Z_{13}(k) )</td>
<td>( Z_{18}(k) )</td>
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<tr>
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### Table 4.2

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<th>778 (K)</th>
<th>778 (°C)</th>
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</tr>
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<tr>
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The function $\lambda_{2s}(k)$ given by $\lambda_{n}(h-k)$, and the associated relativistic correction factor $\lambda_{2s}(k)$, for various values of the photon energy $k$, measured in units of the $K$-shell binding energy $E_{b} = (\hbar c)^{2}/2e$.

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<th>$\lambda_{2s}(k)$</th>
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<td>0.5077</td>
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<td>( x )</td>
<td>( \lambda_{PS}(k) )</td>
<td>( \lambda_{PS}(\mu) )</td>
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TABLE 4.3: The functions $\phi_0(k)$, over $q \cdot (k-\bar{k})$, and $\phi_0(k)$, given by (4.33), for various values of the photon over $k$, measured in units of the 3-shell, 3-disk, e.m.: $E = \hbar^2/2$.

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$$\frac{1}{\sqrt{A_{12}^2 + \sum}} \left( B_{12} \int_{-330}^{330} + 2 \int_{-331}^{331} \right)$$
The table below shows coefficients $a_{n}$ and $b_{n}$ from Eq. (4-21), after correcting for the proper scaling, according to corrections by Son (1971). The table includes values for $\lambda = 1$ and $k + 6$.

<table>
<thead>
<tr>
<th>$\lambda$</th>
<th>$a_{1}$</th>
<th>$a_{2}$</th>
<th>$a_{3}$</th>
<th>$b_{1}$</th>
<th>$b_{2}$</th>
<th>$b_{3}$</th>
<th>$b_{4}$</th>
<th>$k + 6$</th>
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The table is continued with similar entries for different values of $\lambda$ and $k + 6$. The entries are calculated using various mathematical expressions involving gamma functions and constants like $\lambda$, $\alpha$, and $\beta$.
<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
<th>A</th>
<th>Final state (keV)</th>
<th>J(_i) - J(_f)</th>
<th>E(_{EC}) (keV)</th>
<th>Deduced quantities</th>
<th>Spectrometer</th>
<th>Method</th>
<th>Reference</th>
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</thead>
<tbody>
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<td>4</td>
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<td>7</td>
<td>O</td>
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<td>477.6</td>
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<td>I(<em>{IB}), R(</em>{eff})</td>
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<td>Cl</td>
<td>36</td>
<td>O</td>
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<td>1144.1 ± 1.7</td>
<td>1170 ± 40</td>
<td>1158 ± 18</td>
<td>I(_{IB})</td>
<td>NaI</td>
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<tr>
<td>18</td>
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<td>37</td>
<td>O</td>
<td>3/2(^{+}) - 3/2(^{+})</td>
<td>814.1 ± 0.6</td>
<td>818 ± 15</td>
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<td>R(_{eff}), D(_T)</td>
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<td>IB/N(_{0})</td>
<td>Mysłek (1973)</td>
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<td>IB/K-Auger</td>
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<td>$E_{EC}^{a}$ (keV)</td>
<td>Deduced quantities</td>
<td>Spectrometer</td>
<td>Method $^d$</td>
<td>Reference</td>
</tr>
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<td>54</td>
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<td>786 ± 50 $\gamma$ from IIB</td>
<td>NaI</td>
<td>IB/γ</td>
<td>Cohen(1955)</td>
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<td>752 ± 22</td>
<td>NaI</td>
<td>IB/γ</td>
<td>Van der Kooi(1956)</td>
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<td>730 ± 20 $\gamma$ from IIB</td>
<td>NaI</td>
<td>IB/γ</td>
<td>Ofer(1957)</td>
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<td>794 ± 60 $\gamma$ from IIB</td>
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<td>IB/γ</td>
<td>Murty(1967)</td>
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<td>748 ± 14 $\gamma$ from IIB</td>
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<td>IB/γ</td>
<td>Ribordy(1970)</td>
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<td>760 ± 15 IIB Reff</td>
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<td>IB/γ</td>
<td>Mutterer(1973a)</td>
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<td>IB/γ</td>
<td>Koonin(1972)</td>
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<td>IB-γ-coinc.</td>
<td>Koonin(1972)</td>
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<tr>
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<td>NaI</td>
<td>IB/Kx</td>
<td>Bell(1952)</td>
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<td>222 ± 10 IIB</td>
<td>NaI</td>
<td>IB/γ</td>
<td>Michalovicz(1954)</td>
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<td>GM-count.</td>
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<td>232 ± 10</td>
<td>NaI</td>
<td>IB/γ</td>
<td>Madansky(1954)</td>
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</table>

$^a$ Deduced quantities |
$^b$ Others |
$^d$ Reference
<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
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<th>Final state (keV)</th>
<th>$J_1^\pi - J_F^\pi$</th>
<th>$E_{EC}$ (keV)</th>
<th>Deduced quantities $E_{EC}$ (keV)$^b$. others$^c$</th>
<th>Spectrometer</th>
<th>Method$^d$</th>
<th>Reference</th>
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<td>27</td>
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<td>57</td>
<td>136.3</td>
<td>$7/2^- - 5/2^-$</td>
<td>1073.1 ± 1.1</td>
<td>$1075.1^{+1.3}_{-1.3}$</td>
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<td>IB/Kx</td>
<td>Saraf(1956)</td>
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<td>Ni</td>
<td>59</td>
<td>0</td>
<td>$3/2^- - 7/2^-$</td>
<td>235.1 ± 1.7</td>
<td>236 ± 12</td>
<td>IB/Kx</td>
<td>IIB</td>
<td>Biavati(1959,62)</td>
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<td>$1/2^- - 3/2^-$</td>
<td>39.7</td>
<td>513 ± 27</td>
<td>IB/Kx</td>
<td>IB</td>
<td>Raj(1965b)</td>
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<td>517 ± 27</td>
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<td>IB</td>
<td>Rietjens(1954)</td>
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<td>$J_i^\pi - J_f^\pi$</td>
<td>$E_{EC}$ (keV)</td>
<td>$E_{EC}$ (keV)&lt;sup&gt;a&lt;/sup&gt;</td>
<td>DEDUCED QUANTITIES</td>
<td>$E_{EC}$ (keV)&lt;sup&gt;b&lt;/sup&gt;</td>
<td>OTHERS&lt;sup&gt;c&lt;/sup&gt;</td>
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<tr>
<td>50</td>
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<td>391.0 1/2&lt;sup&gt;+&lt;/sup&gt;-1/2&lt;sup&gt;-&lt;/sup&gt;</td>
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<td>930 ± 300</td>
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<tr>
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<td>646.5 1/2&lt;sup&gt;+&lt;/sup&gt;-3/2&lt;sup&gt;-&lt;/sup&gt;</td>
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<td>51</td>
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<td>119</td>
<td>23.8 5/2&lt;sup&gt;+&lt;/sup&gt;-3/2&lt;sup&gt;-&lt;/sup&gt;</td>
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<td>555 ± 20</td>
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<td>125</td>
<td>35.5 5/2&lt;sup&gt;+&lt;/sup&gt;-3/2&lt;sup&gt;-&lt;/sup&gt;</td>
<td>112.5 ± 1.0</td>
<td>141.5 ± 2.0</td>
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<tr>
<td>55</td>
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<td>131</td>
<td>0  5/2&lt;sup&gt;+&lt;/sup&gt;-3/2&lt;sup&gt;-&lt;/sup&gt;</td>
<td>355 ± 6</td>
<td>356 ± 10 I&lt;sub&gt;IB&lt;/sub&gt;</td>
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<td>145</td>
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<td>584 ± 15</td>
<td>NaI</td>
<td>IB/(X)&lt;sub&gt;y&lt;/sub&gt;+γ</td>
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<td>66</td>
<td>Dy</td>
<td>159</td>
<td>0  3/2&lt;sup&gt;-&lt;/sup&gt;-3/2&lt;sup&gt;+&lt;/sup&gt;</td>
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<tr>
<td>68</td>
<td>Er</td>
<td>165</td>
<td>0  5/2&lt;sup&gt;-&lt;/sup&gt;-7/2&lt;sup&gt;-&lt;/sup&gt;</td>
<td>371 ± 4</td>
<td>370 ± 10</td>
<td>NaI</td>
<td>IB/Kx</td>
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<td>74</td>
<td>W</td>
<td>181</td>
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<td>187 ± 10</td>
<td>190 ± 16</td>
<td>Ge(Li)</td>
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<sup>a</sup> Deduced quantities

<sup>b</sup> $E_{EC}$ (keV)<sup>a</sup> is calculated from the final state and $J_i^\pi - J_f^\pi$

<sup>c</sup> Others include possible experimental uncertainties.

<sup>d</sup> Method: IB-γ-coinc., IB/Kx, NaI, Ge(Li).
<table>
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<tr>
<th>Z</th>
<th>Element</th>
<th>A</th>
<th>Final State (keV)</th>
<th>$J_i^m - J_f^m$</th>
<th>$E_{EC}^a$ (keV)</th>
<th>$E_{EC}^b$ (keV)</th>
<th>others $^c$</th>
<th>Spectrometer</th>
<th>Method $^d$</th>
<th>Reference</th>
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<td>60.8 ± 3.0</td>
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<td>Ge(Li)</td>
<td></td>
<td>Hopke (1969)</td>
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$^a$Calculated using $E_{EC}$ values from Wapstra and Gove (1971).

$^b$Partly recalculated from measured $1s-1B$ end-point energies, using electron binding energies from Bearden and Burr (1967).

$^c$Symbols are used for the bremsstrahlung intensity ($I_{LB}$), the effective shape function ($R_{eff}$), information on the influence of detour transitions ($D_T$) and the $\gamma$-branching ratio ($P_\gamma$). Information on the different spectral shapes are not indicated.

$^d$Indicated only if normalized $1B$ spectra have been determined.

$^e$Includes bremsstrahlung of the 8%-EC branch to the ground state of $^{145}$Pm.

$^f$Includes bremsstrahlung of the 26%-EC branch to the 58.2 keV-excited state in $^{159}$Tb.
<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
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<th>Degree of forbiddenness</th>
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<td>allowed</td>
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<td>V</td>
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<td>Fe</td>
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<td>$5/2^{+}$-$3/2^{+}$</td>
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<td>421.2 ± 0.5</td>
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<td>290 ± 100</td>
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<td>$2 \times 10^{6}$ y</td>
<td>597.3 ± 1.2</td>
<td>$7/2^{-}$-$3/2^{-}$</td>
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</tr>
<tr>
<td>28</td>
<td>Ni</td>
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<td>$8 \times 10^{4}$ y</td>
<td>1073.1 ± 1.1</td>
<td>$3/2^{-}$-$7/2^{-}$</td>
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<td>$9/2^{+}$-$5/2^{+}$</td>
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<td>57</td>
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<td>500</td>
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<tr>
<td>52</td>
<td>Te</td>
<td>123</td>
<td>$1.2 \times 10^{13}$ y</td>
<td>57.2 ± 2.4</td>
<td>$1/2^{+}$-$7/2^{+}$</td>
<td>second unique</td>
</tr>
</tbody>
</table>

$^{a}$ from Wapstra and Gove (1971)
<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
<th>A</th>
<th>Final state (keV)</th>
<th>$J_i - J_f$</th>
<th>$E_{EC}^a$ (keV)</th>
<th>Deduced quantities $E_{EC}(keV)^b$ others$^c$</th>
<th>Spectrometer</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>25</td>
<td>Mn</td>
<td>54</td>
<td>835.0</td>
<td>$3^+ - 2^+$</td>
<td>540.1 ± 3.6</td>
<td>528 ± 20</td>
<td>NaI</td>
<td>Jung (1956)</td>
</tr>
<tr>
<td>26</td>
<td>Fe</td>
<td>55</td>
<td>0</td>
<td>$3/2^- - 5/2^-$</td>
<td>231.7 ± 0.7</td>
<td>$I^{1s}_{IB}$</td>
<td>NaI</td>
<td>Biavati (1959, 62)</td>
</tr>
<tr>
<td>38</td>
<td>Sr</td>
<td>85</td>
<td>514.0</td>
<td>$9/2^+ - 9/2^+$</td>
<td>550 ± 7</td>
<td>493 ± 30</td>
<td>NaI</td>
<td>McDonnell (1969)</td>
</tr>
<tr>
<td>48</td>
<td>Cd</td>
<td>109</td>
<td>87.7</td>
<td>$3/2^+ - 7/2^+$</td>
<td>94 ± 3</td>
<td>94 ± 3</td>
<td>Ge(Li)</td>
<td>Gopinathan (1968)</td>
</tr>
<tr>
<td>50</td>
<td>Sn</td>
<td>113</td>
<td>646.5</td>
<td>$1/2^+ - 3/2^-$</td>
<td>378 ± 14</td>
<td>100 ± 10</td>
<td>NaI</td>
<td>Jung (1956)</td>
</tr>
<tr>
<td>53</td>
<td>I</td>
<td>125</td>
<td>35.5</td>
<td>$5/2^+ - 3/2^+$</td>
<td>112.5 ± 1.0</td>
<td>141.5 ± 2</td>
<td>Ge(Li)</td>
<td>Gopinathan (1968)</td>
</tr>
<tr>
<td>55</td>
<td>Cs</td>
<td>131</td>
<td>0</td>
<td>$5/2^+ - 3/2^+$</td>
<td>355 ± 6</td>
<td>$I^{1s}_{IB}$</td>
<td>NaI</td>
<td>Michalowicz (1956)</td>
</tr>
<tr>
<td>62</td>
<td>Sm</td>
<td>145</td>
<td>61.2</td>
<td>$7/2^- - 7/2^+$</td>
<td>577 ± 7</td>
<td>$I^{1s}_{IB}$</td>
<td>NaI</td>
<td>Biavati (1959, 62)</td>
</tr>
<tr>
<td>68</td>
<td>Er</td>
<td>165</td>
<td>0</td>
<td>$5/2^- - 7/2^-$</td>
<td>371 ± 4</td>
<td>$I^{1s}_{IB}$</td>
<td>NaI</td>
<td>Sujkowski (1968)</td>
</tr>
<tr>
<td>74</td>
<td>W</td>
<td>181</td>
<td>0</td>
<td>$9/2^+ - 7/2^+$</td>
<td>187 ± 10</td>
<td>184 ± 12</td>
<td>Ge(Li)</td>
<td>Rao (1966a)</td>
</tr>
<tr>
<td>80</td>
<td>Hg</td>
<td>197</td>
<td>77.3</td>
<td>$1/2^- - 1/2^+$</td>
<td>338 ± 20</td>
<td>686 ± 40</td>
<td>$I^{1s}_{IB}$</td>
<td>NaI</td>
</tr>
<tr>
<td>81</td>
<td>Tl</td>
<td>204</td>
<td>0</td>
<td>$2^- - 0^+$</td>
<td>345 ± 4</td>
<td>335</td>
<td>NaI</td>
<td>Der Mateosian (1956)</td>
</tr>
<tr>
<td>Z</td>
<td>Element</td>
<td>A</td>
<td>Final state (keV)</td>
<td>$J_1^\pi - J_f^\pi$</td>
<td>$E_{EC}^a$ (keV)</td>
<td>Deduced quantities $E_{EC}^b$ (keV)</td>
<td>others$^c$</td>
<td>Spectrometer</td>
</tr>
<tr>
<td>---</td>
<td>--------</td>
<td>---</td>
<td>------------------</td>
<td>-------------------</td>
<td>-----------------</td>
<td>----------------</td>
<td>------------</td>
<td>-------------</td>
</tr>
<tr>
<td>376 ± 20</td>
<td>NaI</td>
<td>Jung (1956)</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>393 ± 10</td>
<td>NaI</td>
<td>Biavati (1959, 62)</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>385 ± 20</td>
<td>NaI</td>
<td>Goudsmit (1966)</td>
<td></td>
<td></td>
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<td></td>
<td></td>
</tr>
</tbody>
</table>

$^a$Calculated using $Q_{EC}$ values from Wapstra and Gove (1971).

$^b$Partly recalculated from measured $1s_{\text{IB}}$ end-point energies, using K-electron binding energies from Bearden and Burr (1967).

$^c$$1s_{\text{IB}} = 1s_{\text{IB}}$ intensity

$^d$Includes bremsstrahlung of the $8\%_{\text{EC}}$ branch to the groundstate of $^{145}_{\text{Pm}}$. 
<table>
<thead>
<tr>
<th>Ele-</th>
<th>Final-</th>
<th>$J_i^\pi - J_f^\pi$</th>
<th>$E_{EC}$</th>
<th>Energy</th>
<th>Inten-</th>
<th>Experiment</th>
<th>Theoretical</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>ment A</td>
<td>state (keV)</td>
<td>(keV)</td>
<td>range (keV)</td>
<td>Intensity ratio</td>
<td>value ($10^{-5}$)</td>
<td>MS</td>
<td>MG</td>
<td>Int</td>
</tr>
</tbody>
</table>

**Allowed transitions** \( \Delta J = 0,1; \; \pi_i \pi_f = +1 \)

Be 7 O 3/2$^-$-3/2$^-$ 861.75±0.09 523.7-$k_{\text{max}}$ I$_{\text{IB/WEC}}$ 9.0±0.6 9.35 8.56 8.57 Mutterer(1973b)

Be 7 O 3/2$^-$-1/2$^-$ 477.6 3/2$^-$-1/2$^-$ 384.1 $^+0.1$ 50 - 360 I$_{\text{IB/WEC}}$ 10.3±0.6 9.95 9.19 9.20 Persson(1972) \(^a\)

A 37 O 3/2$^+$-3/2$^+$ 814.1 $^+0.6$ 35 -$k_{\text{max}}$ I$_{\text{IB/WK}}$ 52 $^+13.5$ 52.1 36.9 37.6 Saraf(1956) \(^d\)

Cr 51 O 7/2$^-$-7/2$^-$ 751.4 $^+0.9$ 348.1-$k_{\text{max}}$ I$_{\text{IB/WEC}}$ 9.56±0.60 14.6 9.13 9.43 Mutterer(1973c) \(^e\)

Mn 54 835.3 3$^+$-2$^+$ 540.1 $^+3.6$ 130 - 425 I$_{\text{IB/WEC}}$ 7.2±0.4 8.34 5.41 5.58 Koonin(1972) \(^f\)

Fe 55 O 3/2$^-$-5/2$^-$ 231.7 $^+0.7$ 50 -$k_{\text{max}}$ I$_{\text{IB/WK}}$ 4.0±1.04 3.42 2.20 2.28 Michalowicz(1955) \(^{15}\)}
<table>
<thead>
<tr>
<th>Element</th>
<th>Final-state</th>
<th>$J_i - J_f$</th>
<th>$E_{EC}$ (keV)</th>
<th>Energy range (keV)</th>
<th>Intensity value ($x10^{-5}$)</th>
<th>Experiment</th>
<th>Theoretical values ($x10^{-5}$)</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>Z</td>
<td></td>
<td></td>
<td>$E_{BC}$ (keV)</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>27 Co</td>
<td>57</td>
<td>136.3</td>
<td>$7/2^- - 5/2^-$</td>
<td>700.4 ±0.7</td>
<td>100 $-k_{max}$ $I_{IB/W_K}$</td>
<td>$1.4^{+0.4}_{-0.5}$</td>
<td>1.68</td>
<td>1.07</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>0 $-k_{max}$ $I_{IB/W_K}$</td>
<td>$1.5^{+0.8}_{-0.6}$</td>
<td>3.76</td>
<td>2.15</td>
</tr>
<tr>
<td>32 Ge</td>
<td>71</td>
<td>0</td>
<td>$1/2^- - 3/2^-$</td>
<td>235.1 ±1.7</td>
<td>180 $-465$ $I_{IB/W_K}$ $I_{EC/W_K}$</td>
<td>$8.8^{+1.3}_{-0.4}$</td>
<td>21.0</td>
<td>12.8</td>
</tr>
<tr>
<td></td>
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<td></td>
<td></td>
<td>70 $-k_{max}$ $I_{IB/W_K}$</td>
<td>$2.3^{+0.5}_{-0.4}$</td>
<td>2.82</td>
<td>1.70</td>
</tr>
<tr>
<td>51 Sb</td>
<td>119</td>
<td>23.8</td>
<td>$5/2^+ - 3/2^+$</td>
<td>555 ±20</td>
<td>0 $-k_{max}$ $I_{IB/W_K}$</td>
<td>$10.6^{+1.2}_{-1.0}$</td>
<td>22.0</td>
<td>8.46</td>
</tr>
<tr>
<td>55 Cs</td>
<td>131</td>
<td>0</td>
<td>$5/2^+ - 3/2^+$</td>
<td>355 ±6</td>
<td>0 $-k_{max}$ $I_{IB/W_K}$</td>
<td>$1.4^{+1.0}_{-0.6}$</td>
<td>7.61</td>
<td>2.25</td>
</tr>
<tr>
<td>68 Er</td>
<td>165</td>
<td>0</td>
<td>$5/2^- - 7/2^-$</td>
<td>371 ±4</td>
<td>93 $-306$ $I_{IB/W_K}$ $I_{EC/W_K}$</td>
<td>$1.6^{+0.16}_{-0.12}$</td>
<td>4.87</td>
<td>1.05</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>182 $-306$ $I_{IB/W_K}$ $I_{EC/W_K}$</td>
<td>$0.53^{+0.06}_{-0.04}$</td>
<td>1.52</td>
<td>0.34</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>185 $-300$ $I_{IB/W_K}$ $I_{EC/W_K}$</td>
<td>$0.9^{+0.2}_{-0.1}$</td>
<td>2.39</td>
<td>1.28</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>185 $-300$ $I_{IB/W_K}$ $I_{EC/W_K}$</td>
<td>$0.89^{+0.12}_{-0.10}$</td>
<td>2.39</td>
<td>1.28</td>
</tr>
<tr>
<td>Element A</td>
<td>Final-state</td>
<td>$J^M_1 - J^M_f$</td>
<td>$E_{EC}$ (keV)</td>
<td>Energy range (keV)</td>
<td>Intensity ratio</td>
<td>Experiment values (x10^-5)</td>
<td>Theoretical values (x10^-5)</td>
<td>Reference</td>
</tr>
<tr>
<td>-----------</td>
<td>-------------</td>
<td>-----------------</td>
<td>----------------</td>
<td>-------------------</td>
<td>----------------</td>
<td>----------------------------</td>
<td>-----------------------------</td>
<td>-----------</td>
</tr>
<tr>
<td>62 Sm 145</td>
<td>7/2^-7/2^+</td>
<td>577 ±7</td>
<td>120 -239</td>
<td>$I_{IB}/W_{K}$</td>
<td>6.7±0.7</td>
<td>11.9 5.87 6.73</td>
<td>Sujkowski(1968)</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>120 -412</td>
<td>$I_{IB}/W_{K}$</td>
<td>10.1±1.0</td>
<td>20.7 9.18 10.6</td>
<td>Sujkowski(1968)</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>169 -412</td>
<td>$I_{IB}/W_{K}$</td>
<td>7.3±1.2</td>
<td>15.1 6.08 7.14</td>
<td>Sujkowski(1968)</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>129 -412</td>
<td>$I_{IB}/W_{K}$</td>
<td>4.6±0.8</td>
<td>14.9 3.90 5.28</td>
<td>Sujkowski(1968)</td>
<td></td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td>169 -412</td>
<td>$I_{IB}/W_{K}$</td>
<td>3.6±0.5</td>
<td>12.2 3.19 4.25</td>
<td>Sujkowski(1968)</td>
<td></td>
</tr>
<tr>
<td>66 Dy 159</td>
<td>3/2^-3/2^+</td>
<td>365.4±1.0</td>
<td>185 -300</td>
<td>$I_{IB}/W_{K}$</td>
<td>1.0±0.3</td>
<td>2.24 1.18 1.32</td>
<td>Sujkowski(1965)</td>
<td></td>
</tr>
<tr>
<td>80 Hg 197</td>
<td>1/2^-1/2^+</td>
<td>686 ±40</td>
<td>350 -550</td>
<td>$I_{IB}/W_{K}$</td>
<td>1.6±0.3</td>
<td>5.49 0.83 1.26</td>
<td>Jasinski(1965)</td>
<td></td>
</tr>
</tbody>
</table>
Table 4.10. Average experimental-to-theoretical IB yield $\langle p_{E,T} \rangle$ for various regions of the atomic number $Z$.

<table>
<thead>
<tr>
<th>Region of $Z$</th>
<th>$4 \leq Z \leq 80$</th>
<th>$Z = 4$</th>
<th>$18 \leq Z \leq 32$</th>
<th>$51 \leq Z \leq 80$</th>
</tr>
</thead>
<tbody>
<tr>
<td>Number of independent measurements</td>
<td>19</td>
<td>3</td>
<td>11</td>
<td>5</td>
</tr>
<tr>
<td>Theory of:</td>
<td>$\langle p_{E,T} \rangle$</td>
<td>$\langle p_{E,T} \rangle$</td>
<td>$\langle p_{E,T} \rangle$</td>
<td>$\langle p_{E,T} \rangle$</td>
</tr>
<tr>
<td>Morrison and Schiff</td>
<td>0.66</td>
<td>0.91</td>
<td>0.76</td>
<td>0.32</td>
</tr>
<tr>
<td>Martin and Glauber</td>
<td>1.19</td>
<td>0.98</td>
<td>1.19</td>
<td>1.31</td>
</tr>
<tr>
<td>Intemann</td>
<td>1.06</td>
<td>0.98</td>
<td>1.10</td>
<td>0.92</td>
</tr>
</tbody>
</table>
TABLE 4.11. Circular Polarization of IB in allowed EC transitions

<table>
<thead>
<tr>
<th>Z</th>
<th>Element</th>
<th>A</th>
<th>Final state (keV)</th>
<th>E$_{EC}$ a) (keV)</th>
<th>Energy range (keV)</th>
<th>Degree of Polarization</th>
<th>Polari- meter type b)</th>
<th>Reference</th>
</tr>
</thead>
<tbody>
<tr>
<td>18</td>
<td>A</td>
<td>37</td>
<td>O</td>
<td>814.1 ± 0.6</td>
<td>200 - k$_{max}$</td>
<td>1.03 ± 0.04</td>
<td>f.s.m.</td>
<td>Hartwig(1958)</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td>0.97 ± 0.15</td>
<td>f.s.m.</td>
<td>Mann(1958)</td>
</tr>
<tr>
<td>24</td>
<td>Cr</td>
<td>51</td>
<td>O(90.2%)</td>
<td>751.4 ± 0.9</td>
<td>320.1(9.8%)</td>
<td>0.67 ± 0.07</td>
<td>f.s.m.</td>
<td>Vandcrleeken(1971)</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
<td></td>
<td>431.3 ± 1.0</td>
<td></td>
<td>1.2 ± 0.1</td>
<td>r.t.m.</td>
<td>Kuphal(1974)</td>
</tr>
<tr>
<td>26</td>
<td>Fe</td>
<td>55</td>
<td>O</td>
<td>231.7 ± 0.7</td>
<td>85 - 220</td>
<td>0.98 ± 0.1</td>
<td>f.s.m.</td>
<td>Parfenova(1960)</td>
</tr>
<tr>
<td>32</td>
<td>Ge</td>
<td>71</td>
<td>O</td>
<td>235.1 ± 1.7</td>
<td>70 - 120</td>
<td>~ 0.4</td>
<td>f.s.m.</td>
<td>Bernardini(195')</td>
</tr>
</tbody>
</table>

a) from Wapstra and Gove (1971)

b) f.s.m. = forward-scattering magnet

r.t.m. = radial transmission magnet.
TABLE 51. Electron ejection probabilities per K capture (in multiples of $10^{-5}$).

<table>
<thead>
<tr>
<th>Isotope</th>
<th>Primaikoff-Porter $^a$</th>
<th>MIK$^b$</th>
<th>Intemann $^c$</th>
<th>K- or L-electron ejection $^d$</th>
</tr>
</thead>
<tbody>
<tr>
<td>$^{37}$Ar</td>
<td>27.7</td>
<td>14.2</td>
<td>21.12</td>
<td>57</td>
</tr>
<tr>
<td>$^{38}$Ar</td>
<td>11.2</td>
<td>8.81</td>
<td>8.26</td>
<td>6.4</td>
</tr>
<tr>
<td>$^{71}$Ge</td>
<td>6.68</td>
<td>4.56</td>
<td>4.72</td>
<td>3.3</td>
</tr>
<tr>
<td>$^{131}$Cs</td>
<td>1.62</td>
<td>0.709</td>
<td>0.92</td>
<td>2.6</td>
</tr>
<tr>
<td>$^{165}$Er</td>
<td>0.767</td>
<td>0.304</td>
<td>0.39</td>
<td>2.9</td>
</tr>
</tbody>
</table>

$^a$Primaikoff and Porter (1953), evaluated by Mukoyama et al. (1973).

$^b$Mukoyama et al. (1973).


TABLE 5.2. Double K-vacancy production probability (due to internal ionization and excitation), per K-capture event (in multiples of $10^{-5}$).

<table>
<thead>
<tr>
<th>Isotope</th>
<th>Theory</th>
<th>Experiments</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>Primakoff-Porter&lt;sup&gt;a&lt;/sup&gt;</td>
<td>MIKS&lt;sup&gt;b&lt;/sup&gt;</td>
</tr>
<tr>
<td>37&lt;sup&gt;18&lt;/sup&gt;Ar</td>
<td>38.6</td>
<td>23.0</td>
</tr>
<tr>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>55&lt;sup&gt;26&lt;/sup&gt;Fe</td>
<td>18.5</td>
<td>15.8</td>
</tr>
<tr>
<td>71&lt;sup&gt;32&lt;/sup&gt;Ge</td>
<td>12.2</td>
<td>8.85</td>
</tr>
<tr>
<td></td>
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</tr>
<tr>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>131&lt;sup&gt;55&lt;/sup&gt;Cs</td>
<td>4.13</td>
<td>1.79</td>
</tr>
<tr>
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<tr>
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<tr>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>165&lt;sup&gt;68&lt;/sup&gt;Er</td>
<td>2.70</td>
<td>1.09</td>
</tr>
<tr>
<td></td>
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<td></td>
</tr>
</tbody>
</table>

<sup>a</sup>Primakoff and Porter (1953), as evaluated by Mukoyama et al. (1973).

<sup>b</sup>Mukoyama et al. (1973).

<sup>c</sup>K-x-ray-K-x-ray coincidence experiments, except for K x-ray satellite measurements on $^{71}$Ge by Oertzen (1964) and Briand et al. (1971).
Figure Captions

FIG. 2-1. The function $I(l_1, l, r)$ vs. distance $r$ from the origin (in multiples of the nuclear radius $A$) for various nuclear charge distributions: a) uniform charge distribution [Eq. (2-55)]; b) Fermi distribution, with $t=0.4A$ [Eq. (2-61)]; c) Gaussian distribution, with $A=0$ [Eq. (2.58)]; d) modified Gaussian distribution, with $A=1$.

FIG. 2-2. $L_1/K$ exchange and overlap correction factors. The solid and broken curves were recalculated according to the approaches of Bahcall (1963a, b, 1965a) and Vatai (1968, 1970), respectively, with wave functions from the Hartree-Fock program of Froese-Fischer (1972a). Results of the relativistic calculation of Suslov (1970a), following Bahcall's theory, are indicated by triangles, and those of the calculation of Martin and Blichert-Toft (1970), based on the same approach as Vatai's, are indicated by crosses.

FIG. 2-3. $L_1/L_1$ exchange and overlap correction factors. See caption of Fig. 2-2 for details.

FIG. 3-1. Typical K, L, and M spectra from the decay of $^{71}$Ge measured with a multiwire counter system. In the M spectrum, background and degradation tails were subtracted and a Poisson distribution fitted to the data (after Genz, 1971a).
FIG. 3-2. Multiwire proportional counter (after Scobie et al., 1959).

FIG. 3-3. Block diagram of multiwire-proportional-counter electronic system (after Genz et al., 1971a).

FIG. 3-4. The M region of the $^{37}$Ar spectrum, with the single-electron spectrum produced by introducing ultraviolet photons from an external source, normalized to the M spectrum (after Renier et al., 1968).

FIG. 3-5. The normalized M and L spectra from $^{37}$Ar decay, corrected for dead time and background (after Renier et al., 1968).

FIG. 3-6. Block diagram of single-wire proportional-counter electronic system (after Genz et al., 1972).

FIG. 3-7. Assemblies of source and enveloping crystals (after Goedbloed et al., 1970a).

FIG. 3-8. Spectrum of $^{131}$Cs measured with a doped NaI(Tl) crystal. Elimination of escape effects by extrapolating to a zero surface-to-volume ratio (after Schulz, 1967a).

FIG. 3-9. M-electron capture decay to the 646-keV level of $^{185}$Re. (a) Spectrum of M events. (b) Extrapolation to correct for escape effects (after Schulz, 1967a).

FIG. 3-10. Block diagram of coincidence apparatus to measure $^{193}$Pt M- and L-capture peaks (after Ravn and Bøgholt, 1971).
FIG. 3-11. Block diagram for coincidence measurements with internal solid sources (after Leutz et al., 1966).

FIG. 3-12. Comparison of experimentally determined L/K capture ratios for allowed transitions (solid circles) and first-forbidden non-unique transitions (open circles) with theoretical predictions based on wave functions of Mann and Waber (1973) and exchange and overlap corrections $X^{L/K}$ according to Bahcall (1963, 1965), Vatai (1970a) and Martin and Blichert-Toft (1970).

FIG. 3-13. Comparison of experimentally determined M/L capture ratios for allowed transitions (solid circles) and first-forbidden non-unique transitions (open circles) with theoretical predictions based on wave functions of Mann and Waber and exchange and overlap corrections $X^{M/L}$ according to Bahcall (1963, 1965), Vatai (1970a), and Martin and Blichert-Toft (1970).

FIG. 3-14. Comparison of experimentally determined $P_X$ values for allowed transitions (solid circles), first-forbidden non-unique transitions (open circles), and first-forbidden unique transitions (squares) with theoretical predictions based on wave functions of Mann and Waber (1973) and exchange and overlap corrections according to Bahcall (1963, 1965).

FIG. 3-15. Number of allowed positron emitters, as a function of half-life.

FIG. 3-16. Continuous gas-flow system used for K/$\beta^+$ measurements.
with short-lived low-Z isotopes.

FIG. 3-17. Diagram of counter used to determine $K/\beta^+$ ratios of $\text{^{11}C, ^{13}N, ^{15}O, ^{19}Ne}$ and $\text{^{30}P}$. K-capture events and positrons are detected in the central counter; only positrons have sufficient energy to be detected in the plastic scintillator.

FIG. 3-18. Typical pulse-height spectrum from the central proportional counter in Fig. 3-17, in anticoincidence with the plastic scintillator. The counter gas, introduced in flow mode, was 90% Ar and 10% CH$_4$. Radioactive phosphine (PH$_3$) was introduced in trace amounts (<1% of Ar/CH$_4$) from an irradiation vessel to the main flow line carrying the counting mixture.

FIG. 3-19. The 870-eV K-capture peak of $\text{^{22}Na}$, measured with an internal-source scintillation counter in coincidence with another NaI detector, closely located to register the 1.274-MeV deexcitation γ rays of $\text{^{22}Ne}$.

FIG. 3-20. Niobium K X rays from the decay of $\text{^{91}Mo}$, measured with a Si(Li) detector with a resolution of 185 eV at 5.9 keV. The Nb Kα, and Kβ peaks are well-resolved, even in the presence of a $\beta^+$ spectrum twenty times as intense as the K-capture branch. The Mo Kα peak is caused by $\beta^+$ induced fluorescence in the source.

FIG. 3-21. Molybdenum-91 K X-ray spectrum measured with a 5.7 × 0.63 cm NaI(Tl) of 28% resolution at 22 keV. The fine structure evident in Fig. 3-20 is no longer visible.
FIG. 3-22. Thin, self-supporting evaporated sources are placed between two CaF₂(Eu) crystals. Although CaF₂ has inherently a lower light output than NaI(Tl), the crystals are nonhygroscopic and can be used without windows between source and crystal.

FIG. 3-23. Typical electronic arrangement for triple-coincidence measurements.

FIG. 3-24. Ratio of experimental to theoretical allowed K/β⁺ and EC/β⁺ ratios.

FIG. 3-25. Theoretical K-capture to positron-emission ratios for allowed transitions.

FIG. 3-26. Theoretical K/β⁺ ratios.

FIG. 4-1. Feynman Diagrams for electronic and nuclear mode contributions to radiative electron capture.

FIG. 4-2. IB spectra for radiative capture from various atomic shells of ⁵⁵Fe, according to the theory of Glauber and Martin (1956).

FIG. 4-3. Relativistic correction factor R_{ls}(k), according to the exact results of Martin and Glauber (1958) and Intemann (1971).

FIG. 4-4. Comparison of several theoretical results for the relativistic correction factor R_{ls}(k). The exact result is deduced from Eqs. (4-44) and (4-45), the low-k expansion, from Eqs. (4-38) and (4-39), and the high-k approximation, from Eq. (4-40).

FIG. 4-5. K-capture IB spectrum for ⁵⁵Fe according to the theories of Morrison and Schiff (1940) (MS) [Eq. (4-14)], Glauber and
Martin (1956) (GM) [Eq. (4-22)], and Martin and Glauber (1958) (MG) [Eq. (4-36)]. GM includes relativistic effects to lowest order in $Z_a$, while MG is fully relativistic.

FIG. 4-6. IB spectra for radiative capture from various atomic shells of $^{165}$Er. The solid curves represent the fully relativistic results of Zon (1971), while the dashed curves are deduced from the results of Glauber and Martin (1956). [After Zon (1971)].

FIG. 4-7. Screening factors $s_a$, according to Martin and Glauber (1958).

FIG. 4-8. Polarization and asymmetry functions, $P_{ls}(k)=\alpha_{ls}(k)$ and $P_{2s}(k)=\alpha_{2s}(k)$, and related functions for $Z=58$. The $ls$-state curves are deduced from the exact results of Martin and Glauber (1958) and Intemann (1971), the $2s$-state curves, from the results of Glauber and Martin (1956).

FIG. 4-9. Polarization and asymmetry functions, $P_{ls}(k)=\alpha_{ls}(k)$ and $P_{2s}(k)=\alpha_{2s}(k)$, and related functions for $Z=51$. The $ls$-state curves are deduced from the exact results of Martin and Glauber (1958) and Intemann (1971), the $2s$-state curves, from the results of Glauber and Martin (1956).

FIG. 4-10. Relativistic correction factors $R_{ls}^{(1)}(k)$ and $R_{ls}^{(2)}(k)$, according to Zon and Rapoport (1968), for several atomic numbers. The function $R_{ls}^{(1)}$ is the same as $R_{ls}$ of Martin and Glauber (1958); it has been evaluated using the high-$k$ approximation [Eqs. (4-40) or (4-76)],
\( R_{1S}^{(2)} \) has also been evaluated in the high-k approximation [Eq. (4-76)]. The three points shown on the ordinate represent the results of an exact evaluation of \( R_{1S}^{(1)}(0) \), using Eq. (4-41) or Table 4.5, for \( Z = 20, 50, 80 \) (in descending order).

**FIG. 4-11.** Comparison of theoretical results for the K-capture IB spectrum for \(^{41}\)Ca. The theories of Morrison and Schiff (1941) and Martin and Glauber (1958) for an allowed transition are represented by the curves MS-A and MG, respectively. For a unique first-forbidden transition, the corresponding curves are those labeled MS-F and ZR, deduced from Eq. (4-79) and the results of Zon and Rapoport (1968), evaluated to first order in \( Z \).

**FIG. 4-12** IB pulse-height spectra of \(^{131}\)I, measured with a 3.5 3.5-cm NaI(Tl) spectrometer. Copper absorbers were placed between source and detector, ranging from 710 mg/cm\(^2\) (A) to 2200 mg/cm\(^2\) (H). [From Saraf (1954a)].

**FIG. 4-13** Pulse-height spectrum of \(^{51}\)Cr, as recorded with a 1.2-cm\(^3\) Ge(Li) detector with a pileup rejector. The measured spectrum \( (N_{ap}) \) is shown in the energy range above the 320.1-keV \( \gamma \)-ray peak, with its individual components: internal bremsstrahlung \( (N_{IB}) \), residual pileup \( (N_{pu}) \), and background \( (N_B) \). [From Mutterer (1973a)].

**FIG. 4-14** Extrapolation plot for pileup correction of \(^{51}\)Cr spectra, recorded from sources of different strengths with a Ge(Li) spectrometer. Ratios of integral counting rates \( (N_{int}) \) for different energy ranges \( E > E_1 \) above the 320.1-keV \( \gamma \) line and total counting rates \( (N) \) are plotted against corrected total counting rates \( (N') \). The intercepts at \( N' = 0 \) give ratios of IB to \( \gamma \) counting rates above different energy thresholds \( E_1 \). [From Mutterer (1973a)].
FIG. 4-15 IB pulse-height spectrum ($n_{IB}$) of $^{51}$Cr, deduced from $^{1}$-$^{7}$$\gamma$\textsuperscript{2} a set of spectra that were recorded with a Ge(Li) spectrometer and corrected for pileup applying the extrapolation method. The corresponding energy spectrum ($E_{IB}$) is shown in the inset. Solid lines represent theoretical spectra of Martin and Glauber (1958). [From Mutterer (1973a)].

FIG. 4-16 Arrangement of two 3x3-in. NaI(Tl) detectors, used for IB spectrometry in coincidence with $\gamma$ rays. [From Persson and Koonin (1972)].

FIG. 4-17 Pulse-height spectra of $^{7}$Be photons gated by $\gamma$ rays, as obtained with two 3x3-in. NaI(Tl) spectrometers in close face-to-face geometry. The coincidence spectrum (open circles) is compared with the random coincidence spectrum (filled circles). The difference between the two spectra represents the pulse-height spectrum of internal bremsstrahlung that accompanies the EC transition to the excited state in $^{7}$Li. [From Lancman and Lebowitz (1971b)].

FIG. 4-18 Electronic circuit of IB spectrometry in coincidence with $\gamma$ rays, for a device with two NaI(Tl) detectors (Fig. 4-16). [From Persson and Koonin (1972)].

FIG. 4-19 IB pulse-height spectrum of $^{7}$Be measured in coincidence with $^{7}$Be $\gamma$-ray with the 477-keV $\gamma$ rays. The peak at 477 keV remained after correction for coincidences. The corresponding Compton distribution is shown as a dashed line. [From Persson and Koonin (1972)].

FIG. 4-20 Total IB spectrum of $^{131}$Cs, measured with a 1x1$\frac{1}{4}$-in. NaI(Tl) crystal with a 0.0005-in. thick aluminum window. The 1s IB
spectrum gated by 131I Kx rays that were recorded with a 1.5 x 0.080-in. NaI(TL) crystal, is also shown. [From Balit et al. (1962)].

FIG. 4-21 Pure higher-shell TB spectrum of 393Pt, measured with a 7-cm³ coaxial Ge(Li) detector. [From Hopke and Naumann (1969), and private communication].

FIG. 4-22 Corrections applied to the predicted 37Ar TB spectrum to convert it into a pulse-height spectrum as would be measured with a 1 x 1-in. NaI(Tl) spectrometer: (A) theoretical curve corrected for γ efficiency, and distribution curves for (B) photo electrons, (C) Compton electrons, (D) backscattered photons, (E) escaped photons, and (F) absorbed photons. [From Lindqvist and Wu (1955)].

FIG. 4-23 Compton distributions of 54Mn (a) and 85Sr (b) recorded with small Ge(Li) spectrometers. Measured spectra are compared with those calculated from constructed response matrices. [(a) from Ribordy and Huber (1970), courtesy of Birkhäuser Publishing Co.; (b) from Nutterer (1973c), unpublished].

FIG. 4-24 IB pulse-height spectrum of 54Mn, measured with a 3 x 3-in. NaI(TL) spectrometer, in coincidence with the 835-keV γ rays from 54Cr. The solid line is the best fit (corresponding to the minimum value of $\chi^2$) of the curves obtained by folding the theoretical IB spectrum with the response matrix. The endpoint energy is used as fitting parameter. [From Iancuan and Lebowitz (1969)].
FIG. 4-25 The ls IB spectrum of $^{165}$Er, as measured with an 1½×1-in. NaI(Tl) spectrometer, in coincidence with Ho K x rays. Jauch plots are shown according to Eq. (4-81) (curve a) and Eq. (4-82) (curve b). [From Zylicz et al. (1963), courtesy of North-Holland Publishing Co.].

FIG. 4-26 IB spectrum of $^{49}$V, measured with a well-type NaI(Tl) spectrometer. The full line represents the theoretical spectrum of Glauber and Martin (1956). The experimental points are normalized at 200 keV. [From Hayward and Hoppe (1956)].

FIG. 4-27 Total IB spectrum of $^{55}$Fe, measured by Berényi et al. (1965) with a 10.2×15.2-cm NaI(Tl) spectrometer. Data points represent the effective shape factor $R_{\text{eff}}$, obtained by dividing the corrected spectrum ($N_{\text{corr}}$) by the Morrison-Schiff term $k(k_o-k)^2$. [From Varga (1970), courtesy of Hungarian Academy of Science].

FIG. 4-28 Ratios $p_{\text{E,T}}$ of experimental to theoretical bremsstrahlung yields, calculated with theoretical IB intensities according to Morrison and Schiff (1940). Values are given for ls IB intensities per K capture (open circles); total IB intensities in which s contributions predominate, relative to K-capture rates (triangles); and total IB intensities relative to total EC rates (filled circles). References for experimental data are given in Table 4.9.
FIG. 4-29 Ratios of experimental to theoretical bremsstrahlung yields, calculated with theoretical IB intensities from the theory of Martin and Glauber (1958), with $R_{ls}(Z,k)$ from Eqs. (4-37) and (4-38). For notation and references see caption of Fig. 4-28.

FIG. 4-30 Ratios of experimental to theoretical bremsstrahlung yields, calculated with theoretical IB intensities according to Intemann (1971) for $R_{ls}(Z,k)$ (Eq. 4-44). For notation and references see Fig. 4-28.

FIG. 4-31 IB spectrum of $^{41}$Ca, measured with a Ge(Li) spectrometer in two different geometries. The spectra ($\frac{W_{IB}}{}$) divided by the Morrison-Schiff spectrum $k(k^2-k^2)$ are compared with predicted IB shape factors. [From Myszêk et al., (1973), courtesy of North-Holland Publishing Co.].

FIG. 4-32 Total IB spectrum accompanying the second-forbidden nonunique EC decay of $^{36}$Cl, measured with a 10 x 10-cm NaI(Tl) spectrometer. The spectrum is shown in Jauch coordinates (a) and in form of the effective shape function $R_{\text{eff}}$ (b). The latter is compared with theoretical shape functions, calculated with (1) and without (2) including detour transitions. [From Smirnov and Batkin (1973), courtesy of Nauka Press].

FIG. 4-33 Circular Polarization (P) of the IB from $^{37}$Ar as function of energy, measured with a forward-scattering Compton polarimeter provided with a NaI(Tl) detector. (Hartwig and Schopper, 1958). The solid line is the theoretical curve, cal-
culated from Eq. (4-52) with the polarization functions \( \alpha_{1s} \) and \( \alpha_{2s} \) of Fig. 4-8. The IB spectrum of \(^{37}\text{Ar}\) is also shown.

**FIG. 4-4** Relative change \( \delta(x) \) in the Compton absorption of \(^{51}\text{Cr}\) photons in iron that is magnetized parallel and anti-parallel, respectively, to the photon momentum. The Compton polarimeter has a special radial-transmission magnet. Photons are recorded with NaI(Tl) detectors applying current integration techniques. Values of \( \delta(x) \) for different Pb absorbers between source and magnet are shown. Solid lines are calculated from IB theory. [From Kuphal et al. (1974)].

**FIG. 4-5** Forward-backward asymmetry \( W(\pi)/W(0) \) in the emission of IB photons from polarized \(^{119}\text{Sb}\) nuclei, as a function of sample temperature. [From Brewer and Shirley (1968)].

**FIG. 4-6** Overall asymmetry coefficient \( A(k) \) of IB emission from oriented \(^{119}\text{Sb}\) nuclei, measured by Brewer and Shirley (1968). Data points are compared with theoretical predictions, calculated with the asymmetry functions \( \alpha_{1s} \) and \( \alpha_{2s} \) of Fig. 4-9 (full curve), and with \( \alpha_{1s} = \alpha_{2s} = 1 \) (dashed curve). [From Intemann (1971)].
FIG. 5-1. Theoretical momentum spectrum of K electrons ejected during K capture of $^{55}$Fe. The upper curve is calculated according to Intemann and Pollock (1967), taking into account only the exchange of scalar virtual photons during transitions between spherically symmetric states. The lower curve, calculated by Intemann (1972), results if p-wave intermediate states and the exchange of longitudinal virtual photons are taken into account. (After Intemann, 1972).

FIG. 5-2. Calculated momentum spectrum of K electrons ejected during K-capture decay of $^{131}$Cs. Curve A is according to the non-relativistic theory of Primakoff and Porter (1953); curve B represents the semirelativistic calculation of Intemann (1969). (From Intemann, 1969).

FIG. 5-3. Calculated energy spectra of electrons ejected in the decay of $^{55}$Fe. The dashed curve labeled "K-K" represents K electrons ejected during K capture; the curves "K-L$_1$" indicate L$_1$-electrons ejected during K capture plus the exchange effect, viz., K electrons ejected during L$_1$ capture. All rates are given per K-capture event. After Mukoyama and Shimizu (1974).
FIG. 2-2

- Bahcall
- Vatai
- Martin and Blichert-Toft
- Suslov
$$\frac{M_1}{L_1}$$

FIG. 2-3
FIG. 3-1
M Region
Ar$^{37}$

- M region
- Single-electron shape
- Background

COUNTS PER CHANNEL

CHANNEL NUMBER

FIG. 3-4
FIG. 3-5

COMPOSITE M AND L SPECTRUM

$^{37}$Ar
FIG. 3-6
$K_{\text{bind}} = 34.6$ keV

$P_K / (P_K + P_L)$

$\frac{A}{(A+B)}$

$\times^{131}\text{Cs in CsI(Na)}$

$\bigcirc^{131}\text{Cs in NaI(Tl)}$

FIG. 3-8
FIG. 3-9

(a) Counting Rate (arb. units)

Energy (keV)

(b) Counting Rate (arb. units)

Surface/Volume (cm⁻¹)

69%
FIG. 3-13
FIG. 3-15

HALF LIFE

FREQUENCY

1 sec 1 min 1 hour 1 day 1 month 1 year 10 years
FIG. 3-16
PLASTIC COUNTER

GAS INLET

GUARD TUBE

PLASTIC SCINTILLATOR

LIGHT GUIDE

EVAPORATED CATHODE

PROPORTIONAL COUNTER

O'RING SEAL

GAS OUTLETS

PHOTOMULTIPLIERS

FIG. 3-17
FIG. 3-18

30p
K PEAK
1.838 keV.
FIG. 3-2C
FIG. 3-21
FIG. 3-23
22Na

THEORY
MacMAHON
VATAI
WILLIAMS
LEUTZ

Z OF THE PARENT NUCLEUS

FIG. 3-24
Maximum $\beta^+$ kinetic energy (MeV)
FIG. 4-1
FIG. 4-2
1.0

EXACT RESULT
LOW k EXPANSION
HIGH k APPROXIMATION

Z=4
Z=24
Z=68

0.2 0.4 0.6 0.8 1.0 1.2 1.4 1.6 1.8
k (UNITS OF mc /h)

FIG. 4-4
FIG. 4-5

<table>
<thead>
<tr>
<th>k (IN UNITS OF mc²)</th>
<th>( \frac{d\omega_{1S}}{dk}/w_K ) (IN UNITS OF ( 10^{-5}$/mc²)</th>
</tr>
</thead>
<tbody>
<tr>
<td>0.1</td>
<td>Approximately 0</td>
</tr>
<tr>
<td>0.2</td>
<td>Approximately 10</td>
</tr>
<tr>
<td>0.3</td>
<td>Approximately 15</td>
</tr>
<tr>
<td>0.4</td>
<td>Approximately 20</td>
</tr>
</tbody>
</table>

Curves labeled MS, EM, and MG.
FIG. 4-6

$3 \times 10^4 \text{ } W_E$

$2S$, $3P$, $2P$

$k$, keV

50 100 150 200 250 300
FIG. 4-8
FIG. 4-10
FIG. 4-11

(dW_1s/dk)/w_k (IN UNITS OF 10^{-5}/mc^2)

k (IN UNITS OF mc^2)

MS-A
MG
MS-F
ZR
FIG. 4-12

COUNTS PER MINUTE x (1/128)

ENERGY / keV

80 keV
163 keV

CHANNEL WIDTH
13.3 keV

131Cs

A
B
C
D
E
F
G
H

FIG. 4-12
Counting rates per channel [sec^{-1}]

FIG. 4-13

Energy E (keV)

- $^{208}$Tl 510.8 keV (BGR)
- $^{208}$Tl 583.1 keV (BGR)
- $^{214}$Bi 609.4 keV (BGR)
- Coincidence sum peak 640.2 keV
- IB - Endpoint 750.8 keV
$^{51}$Cr

Corrected total counting rate $N'$ (sec$^{-1}$)

FIG. 4-14
SOURCE

3X3" NaI

1/32" Cd

1/32" Cu

Pb

3X3" NaI

IB

FIG. 4-16
FIG. 4-17
FIG. 4-19

$^{7}\text{Be}$

$E_\gamma = 477.4$ keV

ENDPOINT
$^{131}\text{Cs}$

**A: TOTAL SPECTRUM**

**B: 1S CAPTURE SPECTRUM**

---

**FIG. 4-20**
Counts per channel

channel number

Photon energy

$^{193}\text{Pt}$

FIG. 4-21
FIG. 4-23

(a) $^{54}\text{Mn}$

(b) $^{85}\text{Sr}$
FIG. 4-24

54 Mn

\[ \chi^2 \]

506 \pm 25 keV

END-POINT ENERGY (keV)

ENERGY (keV)

INTENSITY
FIG. 4-25

\[ \sqrt{\frac{N}{E \cdot R_{15}}} \]

\[ E = 314.5 \]

\[ 165_{\text{Er}} \]

CHANNEL NUMBER

ENERGY (keV)

200 300
FIG. 4-26

ENERGY in keV

RELATIVE INTENSITY

POINT AT 50.2

49V

1S

2P

2S

FIG. 4-26
FIG. 4-27

$\frac{N_{\text{corr}}}{k(k_0-k)^2}$

$\text{Fe}^{55}$

Energy

$K$ (channel number)

keV

533
Experimental-to-theoretical bremsstrahlung yield $\rho_{E,T}$

Atomic number $Z$

FIG. 4-28
FIG. 4-29
FIG. 4-30

Experimental-to-theoretical bremsstrahlung yield $\rho_{E1}$

Atomic number $Z$

INTEMANN
Figure 4.31

$\frac{W_{n\alpha}}{k(k^2+k^2)}$

"distant" geometry

"close" geometry

Glauber Martin

$Q_{\beta^-}(^{46}\text{Ca})$

$Q_{\beta^-}(^{46}\text{Ca})$

$\theta_{EC}$

 ENERGY [keV]

$Zon$

$k^2+q^2$

$Zon$

ENERGY [keV]

ARBITRARY UNITS

$41\text{Ca}$
FIG. 4-32
\[ \delta_{br} \left[ 10^{-5} \right] \]

\[ X \text{ [mm]} \]

**FIG. 4-34**

Diagram showing energy levels and branching ratios for various states: 51Cr, 7/2-, 51V, 5/2-, and spectral lines with branching ratios given as percentages.
FIG. 4-35
$^{119}\text{Sb}$

![Graph showing $A(K)$ vs. $K$ (keV) with error bars and a fitted curve.](image-url)
FIG. 5-1

\[ \frac{d\omega}{dP} \text{ (units of } 10^{-4} \text{/mc)} \]

\[ P \text{ (units of mc)} \]
FIG. 5-2

- \( \frac{dW_1}{dP} \) (units of \( 10^{-5}/mc \))
- \( P \) (units of \( mc \))
FIG. 5-3

EJECTED-ELECTRON ENERGY (keV)