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TURBULENCE AND WAVE PARTICLE INTERACTIONS IN SOLAR-TERRESTRIAL PLASMAS

ANNUAL STATUS REPORT

(1 July 1983 - 30 June 1984)

Grantee: The Regents of the University of Colorado Boulder, Colorado 80309

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Grant Title: Turbulence and Wave Particle Interactions in Solar-Terrestrial Plasmas

Grant Number: NASA Grant NAGW-91 University of Colorado Account 1-5331-48
ANNUAL STATUS REPORT
(1 July 1983 - 30 June 1984)

Turbulence and Wave Particle Interactions
in Solar-Terrestrial Plasmas

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A. Particle and Wave Processes in Solar Flares - led by G. Dulk

We have continued to investigate the amplification of cyclotron maser radiation in solar flares. An RF heating model for the corona surrounding the energy release site was developed (Melrose and Dulk 1984; Dulk and Melrose in SMM Workshop). This model permits a considerably simplified interpretation of several facets of flares: rapid precipitation of fast electrons from magnetic traps, cross-field transport of significant energy at the speed of light, development of macroscopic turbulence which can broaden lines of heavy ions.

Several important aspects of maser emission were investigated: (1) The relative growth and damping rates of various modes and harmonics were determined and it was concluded that the z mode probably grows fastest under many circumstances and coupling between two z mode waves may produce harmonic radiation of the kind observed (Melrose, Hewitt and Dulk 1984). (2) The presence of a background plasma of moderate to high temperature was found to have a significant effect on the growth rate, allowing growth of fundamental, x-mode radiation at densities some four times higher than previously thought possible (Winglee 1984a). (3) The importance of bunching of electrons in particular phases of waves was investigated and found to have probable consequences in triggered VLF emissions in the Earth's magnetosphere (Winglee 1984b).

Other studies included several review papers on radio emission processes by Dulk and collaborators, an investigation of the circumstances when radio emission accompanies shock waves by Gary et al. (1984) and an investigation of the circumstances when plasma emission occurs in the solar wind at the fundamental vs. the harmonic (Dulk, Steinberg and Hoang 1984).
B. Solar Convection Zone Turbulence - led by J. Toomre

Our nonlinear simulations of compressible convection display prominent penetration by plumes into regions of stable stratification at the base of the solar convection zone, leading to the excitation of internal gravity waves there. The inclusion of magnetic fields into these calculations yields regions of very concentrated fields, with such flux sheets or tubes becoming substantially evacuated of plasma.

C. Solar Radio Emission - led by M. Goldman

During this period our work on Langmuir turbulence and radiation associated with Type III solar radio emission and planetary bow-shock emission divides into three main categories: linear saturation of electron beam-driven Langmuir waves by ambient density fluctuations; nonlinear saturation by "strong turbulence" processes; and radiation emission mechanisms. There have been major recognized achievements in all three areas. These are described in 5 published articles, 3 additional submitted articles, and 2 principal invited papers at international conferences.

In the area of linear saturation, our earlier research on the saturation of beam-excited Langmuir waves in the foreshock of the Jovian bow-shock by multiple scatter off (measured) short-scale ion-acoustic turbulence has been published (Russell and Goldman 1983). We have now developed a variation of this idea, and applied it to Type III burst associated Langmuir waves. In a paper submitted to Solar Physics (Muschietti, Goldman and Newman 1984), we have explored the quenching of the beam-plasma instability by diffusion from a 3-D spectrum of large-scale ambient density fluctuations, and come to the
significant conclusion that this effect is so powerful that unless such
density fluctuations are highly anisotropic in space, the Langmuir waves are
often completely suppressed. This indirect evidence for "duct-shaped" ambient
density fluctuations fits nicely with recent theories (Dulk, Steinberg and
Hoang 1984) of ducted radio-wave emission from remote sources.

Our past work on nonlinear saturation by strong turbulence effects has
been summarized in two review articles (Goldman 1983, 1984), the latter
solicited from the editors of Review of Modern Physics after an invited review
talk before the Plasma Physics Division of the American Physical Society in
November, 1983. In recent research still underway, we have studied the
evolution of strong Langmuir turbulence in the vicinity of planetary
bowshocks, where backscatter off thermal ion-acoustic waves may occur, and
calculated the level and angular distribution of 2nd harmonic emission from
the resulting turbulent spectrum (Newman, Muschietti, Moon, and Goldman
1984). In two other recent developments, we have observed driven
self-focusing of Langmuir waves for the first time in a 2-d particle in cell
simulation (Russell, Goldman, and Dubois 1984), and explored the intermittency

Finally, we have completed, and submitted for publication, detailed
calculations on two new mechanisms by which Langmuir turbulence may produce
electromagnetic emission well above the electron plasma frequency. The first
mechanism is Compton conversion of Langmuir waves in the presence of a
relativistic electron beam (Newman 1984), and the second is multiple Raman
scatter of radiation from pre-existing Langmuir waves.
D. Solar Magnetic Fields and Hydromagnetic Waves in Inhomogeneous Media—led by E. Zweibel

Influence of Lower Boundary Condition on MHD Stability on the Solar Corona: There has been some controversy in the literature concerning the correct boundary conditions to use in studies of coronal MHD stability. It is generally agreed that the inertia of the photospheric gas should provide some stabilization by line tying, but the effect of vertical displacements of the corona-photosphere boundary has not been analyzed. I showed (Zweibel 1984) that vertical motions of the boundary are strongly stabilized by the density stratification of the solar atmosphere. This calculation provides a context for choosing boundary conditions and ties together some apparently discrepant results in the literature.

Effect of Magnetic Fields on Stellar Oscillations: Although the field of solar and stellar oscillations is quite active, very few calculations have been done which incorporate the effect of a magnetic field. Calculations presently underway (in collaboration with H. van Horn's group at the University of Rochester) show that the magnetic field can have a strong effect on oscillations which are localized (or trapped) within the surface layers of a neutron star. This is true even though the field may be unimportant over most of the interior of the star (Morrow and Zweibel 1984). We expect a similar effect in magnetic white dwarfs and in solar active regions.

A program focused on magnetic effects of solar oscillations is underway in collaboration with T. Brown (HAO) and C. Morrow (graduate student, CU). This program involves observations of solar oscillations with the Fourier tachometer at Sac Peak, a new instrument developed by Brown, as well as theoretical analysis. We are studying the modification of the oscillation spectrum by magnetic fields which are concentrated into intense flux tubes, as the solar field is observed to be.
REFERENCES


Appendix A

Publications Related To This Grant

"Turbulence and Wave Particle Interactions in Solar Terrestrial Plasmas

1 July 1983 - 30 June 1984\(^\text{1}\)

\(^{1}\text{List compiled October 1984.}\)
PUBLICATIONS RELATED TO THIS GRANT
(Inclusive; status as of October 1984)


Appendix B

Invited Papers

"Turbulence and Wave Particle Interactions in Solar Terrestrial Plasmas"

1 July 1983 – 30 June 1984

List compiled October 1984.
INVITED PAPERS
(inclusive; status as of June 1984)


28. "Plasma Phenomena in Close Binary Stars," George A. Dulk, Department of Theoretical Physics, University of Sydney Colloquium; March 1983.


30. "AM Herculis-Type Binaries," George A. Dulk, CSIRO Colloquium, Sydney, Australia; April 1983.


Appendix C

Contributed Papers Related To This Grant

"Turbulence and Wave Particle Interactions
in Solar Terrestrial Plasmas"

1 July 1983 - 30 June 1984

[List compiled October 1984]
CONTRIBUTED PAPERS RELATED TO THIS GRANT
(Published Abstracts Listed when applicable, status as of June 1984)


Appendix D

Conferences, Seminars, and Other Grant-Related Activities

"Turbulence and Wave Particle Interactions in Solar Terrestrial Plasmas"

1 July 1983 - 30 June 1984
CONFERENCES, SEMINARS AND OTHER GRANT-RELATED ACTIVITIES
(1 July 1983 - 31 June 1984)

George Dulk, Senior Visiting Scientist, at Observatoire de Paris, France, July-December 1983.

George Dulk attended the UK/SMM Workshop, Oxford University, England, September 1983.


George Dulk visited Space Research Laboratory, Utrecht, Netherlands, October 1983.

George Dulk attended meeting of Japan-France Seminar, Paris, October 1983.

George Dulk visited Dept. of Applied Physics, University of Bern, Switzerland, November 1983.


Juri Toomre and Ellen Zweibel attended the Workshop on Solar Physics sponsored by the solar groups in the STTP, Stanford University, March 16-21, 1984.

George Dulk conducted site visit to NRAO Tucson, Az., January 1984.

George Dulk attended SMM Workshop, Goddard SFC, February 1984.

George Dulk attended workshop on cataclysmic variable stars, LSU, Baton Rouge, La., February 1984.

George Dulk conducted observing program on cataclysmic variable stars, VLA, March 1984; on flare stars in Orion and the Pleiades, VLA, March 1984.


George Dulk attended meeting of NRAO Visitors Committee, Charlottesville, VA, April 1984.


George Dulk attended meeting of NRAO Users Committee, Socorro, NM, May 1984.
George Dulk attended General Assembly of COSPAR, Graz, Austria, June 1984.

George Dulk conducted joint EUV rocket and VLA radio experiment, June 1984.

George Dulk conducted invited working visit to Observatoire de Paris, Meudon, June-July 1984.