ANOMALIES IN COSMIC RAYS: NEW PARTICLES \textit{vs} CHARM?

G.L.\ BALAYAN, A.YU.\KHODJAMIRIAN, A.G.\OGANESSIAN

Yerevan Physics Institute, Markarian st.2
375036, Yerevan, Armenia, USSR

For a long time two anomalies are observed in cosmic rays at energies $E \sim 100$ TeV: 1) the generation of "long-flying" cascades in the hadron calorimeter (the so-called Tien-Shan effect) /1/; 2) the enhancement of direct muon yield as compared with the accelerator energy region /2/. The aim of this paper is to discuss the possibility that both anomalies have common origin arising from production and decays of the same particles. The main conclusions are the following: 1) direct muons cannot be generated by any new particles with mass exceeding $10+20$ GeV; 2) if both effects are originated from the charmed hadrons, then the needed charm hadroproduction cross section is unexpectedly large as compared with the quark-gluon model predictions.

When energy grows from $E \sim 10$ TeV to $E \sim 100$ TeV, an increase of the hadron cascade mean penetration length is observed. This phenomenon is interpreted by the authors of Tien-Shan experiment /1/ as production (in the cosmic shower interactions with the calorimeter nuclei) of unstable particles which carry a substantial part ($\langle x \rangle \approx 0.25$) of the incident energy and simultaneously have small inelasticity (or small cross section of the interactions with the calorimeter substance). When the incident energy becomes $\sim 100$ TeV, the mean decay path of these particles $\ell \sim \langle x \rangle E c c /m$ reaches the scale of the calorimeter vertical size ( $\ell \sim 1$ m) so that the resulting hadron cascade is observed as "long-flying". The mass range of these particles is estimated /1/ as $2 \text{ GeV} \leq m \leq 450 \text{ GeV}$, their lifetime being respectively in the interval $10^{-12} \text{ sec} \leq \tau \leq 10^{-10} \text{ sec}$. The hadroproduction cross section of "long-flying" particles is $\sigma \sim 3+5 \text{ mb/nucleon}$ at $E \sim 100$ TeV.
The direct muons produced in the primary pN-collisions with definite energy $E_{\mu}$ are characterized by the ratio of their yield to the number of pions produced with the same energy. We have obtained for this ratio $R$ integrated over the primary spectrum with power $\chi + 1 = 2.65$ the following estimate

$$R(E_{\mu}) \approx 11.5 \frac{n_{\mu}}{\langle x_{\mu} \rangle} \int_0^1 \frac{dN}{dx_{\mu}} \frac{\chi_{\mu}^\chi}{d\chi_{\mu}}$$

Here the familiar parametrization of pion inclusive cross section from /3/ was used and scaling for direct muons was assumed ($\langle n_{\mu} \rangle = \frac{\phi(\mu^+\mu^-)}{\phi_{\text{tot}}}$ is the direct muon mean multiplicity, $dN/dx_{\mu}$ is the normalized spectrum over $\chi_{\mu} = E_{\mu}/E$; $\langle x_{\mu} \rangle$ is the value at which the integrand in eq.(1) reaches its maximum). Due to the $\chi_{\mu}$ factor the ratio $R$ becomes a good analyzer of muons carrying a large part of the incident energy. Data from a few experiments /2/, although with large uncertainties, indicate that in the region $E_{\mu} = 1\text{--}100$ TeV $R$ may reach an order of $10^{-3}$ value.

Suppose that starting from $10\text{--}100$ TeV some new X-particles (of hadronic origin) are produced, which simultaneously are "long-flying" and have muonic decays. It turns out that already available SPS collider ($\sqrt{s} = 540$ GeV, i.e. $E \sim 150$ TeV) data sample allows us to reject too heavy X-particles generating large $p_{\perp}$ muons. To demonstrate this, we have calculated the ratio

$$\Delta(M_x) = \frac{\int d\phi_{x_{\mu}} \int d\phi_{x_{\mu}} \frac{d\phi_{x_{\mu}}}{dx_{\mu} d\phi_{x_{\mu}}}}{\phi(\rho \tilde{p} \rightarrow x^+ \ldots)}$$

which determines that part of direct muons from X-particles to which the collider detectors are sensitive. In fact, $\Delta(M_x) > 10\%$ at $M_x > 10$ GeV (20 GeV), if the muon creates in $X \rightarrow \mu^+\nu, \mu^-\mu^+ (X \rightarrow \mu^+\nu + H, \mu^-\mu^+ + H)$decays, where $M_H$ may be $\lesssim M_X/2$. The X-particle mean $p_{\perp}$ was chosen $\sim 1$ GeV/c (i.e. too low for real heavy particles) to decrease artificially the share of large $p_{\perp}$ muons. Nevertheless due to the large X-mass the substantial part of muons are produced with $p_{\perp,\mu} > 5$ GeV. At
the same time, the value of $\Delta$ shows practically no change when the longitudinal X-spectrum is varied in wide limits (from $d\rho/dx \sim (1-x)^5/x$ to $d\rho/dx \sim x(1-x)$). From the UA1-data sample on the $W \rightarrow \mu\nu$ search we conclude that $\rho(p\bar{\nu}(\mu_\mu > 5 \text{GeV}) + ...) \sim 0.2$ at $E \sim 150$ TeV. By means of coefficient (2) we immediately obtain the upper limit for the cross section of hypothetical X-particles:

$$\rho(pN \rightarrow X + ...) B(X \rightarrow \mu) \leq 2 \mu \delta$$

which holds for all variants of X-production and decays mentioned above (we suppose that $\rho(pN \rightarrow X + ...) \simeq \rho(p\bar{\nu} \rightarrow X + ...)$ which is justified at such energies). It is easy to verify that the upper bound given by eq. (3) allows too few direct muons from X-particles ($R \ll 10^{-4}$). In addition, we think that: 1) more detailed scenarios of X-decays (involving a second muon and/or neutrino) will decrease the upper limit of the allowed X-masses up to the b-flavored hadrons; 2) the analogous consideration of jets from pure hadronic X-decays will allow also to reject too heavy X as "long-flying" cascade source. At the same time it is evident that the collider data obtained in the large $p_\perp$ region are not too sensitive to the muons from charm decays produced predominantly with $p_\perp \mu \leq 1$ GeV.

As it was noted earlier, the charmed hadrons ($\Lambda_c$ and $\Xi$) are really good candidates for "long-flying" particles. Particularly important is that their expected inelasticity is small as compared with light hadrons due to the specific effect of c-quark leading inside $\Lambda_c$ or $\Xi$ (6). It is known also that the mean part of incident energy $\langle x \rangle$ carried by charmed hadrons may be really large. In particular, $\Lambda_c$-baryons were detected at ISR ($E \sim 2$ TeV) (7) with spectrum

$$d\rho/dx \simeq 3/2 \delta \sqrt{1-x}$$

where $\delta \leq 0.3$ mb, $\langle x \rangle \sim 0.4$. The whole effect of "long-flying" cascades may be understood in terms of charm if it achieves $\sim 3$ mb at $E \sim 100$ TeV. Therefore, it is necess-
ary to postulate a rapid growth of charm yield in the
fragmentation region \( \chi > 0.1 \). We have chosen it as fol-
lows:

\[
\mathcal{O} \approx 0.3 \text{mb} \left( 1 + \ln \left( \frac{\sqrt{s} (\text{GeV})}{60} \right) \right)^2
\]  

(5)

The muon spectrum from semileptonic decays of charm
\( B(c \rightarrow \mu) = 10\% \) produced according to eqs. (4), (5) was cal-
culated. This spectrum behaves like \( (1 - \chi_{\mu})^6 / \chi_{\mu} \) at \( \chi_{\mu} > 0.05 \) and gives \( R \sim 1.3 \times 10^{-3} \) at \( E_{\mu} \sim 10 \text{ TeV} \).

Therefore, both effects: the "long-flying" cascades and
direct muons are simultaneously explained if we suppose
abundant charm production in the fragmentation region. Note
that this phenomenon is rather unexpected in quark-gluon
models of charm hadroproduction. Although the whole picture
of heavy quark hadroproduction is yet far from understand-
ing, it is clear that in the fragmentation region some non-
perturbative mechanisms are essential. The models describ-
ing this region predict /8-10/ too slow rise \( \sim \ln(\frac{E}{100 \text{ TeV}}) \) of
the cross section fragmentation part so that at \( E \sim 100 \text{ TeV} \)
in eq. (4) \( \mathcal{O} \leq 0.5 \pm 0.8 \text{ mb} \).

We conclude that more detailed information about inclu-
sive cross section of charmed hadroproduction at energies
\( \sim 100 \text{ TeV} \) would be of great interest from the viewpoint of
the quark-gluon physics. New data on the cosmic ray anom-
aries at superhigh energies are therefore needed. Note that
the proposed ANI installation /11/ will have a unique pos-
sibility to detect both "long-flying" cascades and direct
muons.

REFERENCES

/3/ Hillas A.M. Proc.Paris Workshop on Cascade Simulations,
1981, 39.
/10/ Khodjamirian A.Yu., Oganessian A.G. Proc.18th ICRC,