

PROJECT VESTA - EXPLORATORY MISSION TO THE ASTEROIDS;
SCIENTIFIC OBJECTIVES AND TECHNICAL SPECIFICATIONS

Centre National d'Etudes Spatiales (CNES)

Translation of "Projet VESTA - Interet Scientifique d'Une
Mission d'Exploration vers les Petits Corps", Centre National
d'Etudes Spatiales (CNES), Paris, France, 1985, pp. 1-14

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16. Abstract A description of the VESTA project, a Franco-Soviet collaboration, is presented. The VESTA probe will be launched by the Soviet Union at the end of 1992, and will pass through the asteroid belts before encountering comets. The probe will be an autonomous vehicle under French control. The scientific objectives include obtaining information on asteroids, comets, solar system formation, meteorites and other sources of extra-terrestrial material. The surface morphology, mineral composition and internal structure of asteroids are to be studied.			
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PROJECT VESPA - EXPLORATORY MISSION TO THE ASTEROIDS;
SCIENTIFIC OBJECTIVES AND TECHNICAL SPECIFICATIONS

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- To complete the information gathered on objects which make up the solar system (asteroids, comets).
- To obtain information on the first stages of formation of the solar system: very primitive bodies → processes of differentiation and of collision between planetoids.
- To define the relationship between meteorites and asteroids, micrometeorites and comets: scientific evaluation of the most important and diverse source of extraterrestrial material.

SPECIFIC SCIENTIFIC OBJECTIVES: FLIGHT OVER AN ASTEROID IN /2
THE BELT

Overall characteristics: size, shape, rotation period and axis.
Density → information on internal structure.

Surface morphology and characteristics: existence and thickness of a layer of debris (dust or chunks) → physics of collisions, origin of meteoritic flaws.

Mineral composition of the surface: relationship with the principal groups of meteorites. Existence of geochemical provinces (differentiation).

*Numbers in the margin indicate pagination in the foreign text.

Environment of the asteroids: magnetic field (core)?

Satellites? Weak comet activity → dusts?

Characteristics of the interplanetary environment between 0.7 and 2.5 u.a.

PRINCIPAL SCIENTIFIC RESULTS OBTAINED DURING FLIGHT OVER AN /3
ASTEROID

Target type: Asteroid 50 to 100 km in diameter

Minimum distance: < 1000 km

Relative velocity: 4 to 10 km/sec

- Complete cartography of the light hemisphere with a resolution of ~ 100 m + ~ 10 high-resolution (< 10 m) photographs.
- Cartography of the dark hemisphere resolution ~ 1 km with the high-resolution camera.
- Infrared cartography of the light hemisphere; pixel 3 to 5 km.
- Volume (in qq %), roughness, thermal and dielectric properties.
- Determination of mass → density Δ m/m ~ 3% ("test mass").

SPECIFIC SCIENTIFIC OBJECTIVES: FLIGHT OVER A COMET CORE /4

Overall characteristics: size, shape, rotation period and axis.

Surface morphology: Existence of a silicate crust, of active regions (ice), possible craterization.

Mineral composition of the surface: composition of silicates,
trapped volatiles, ice.

Environment of a low-activity comet: chemical composition of
dusts → direct information of the chemical composition of the
core.

POSSIBLE USEFUL SCIENTIFIC PAYLOAD FOR FLIGHT OVER THE
ASTEROIDS AND COMETS

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	<u>Field</u>	<u>Pixel Size</u>	<u>Other characteristics</u>
Two Cameras			
<u>Wide angle</u>	8°	80 m	navigation
<u>High resol.</u>	0.6°	5 m	"
IR Spectrometer	6.5°	2 km	0.8 to 5 μm
Radar			
<u>Altimeter</u>	1.6°	15 km	±=20 m (alt.)
<u>Radiometer</u>	2.5 to 15°		±=0.5°K
<u>Experiment</u>			
<u>"test mass"</u>			<u>+ 5% of density</u>

Study of the environment and the interplanetary surroundings

Magnetometer ~ 1 nano Tesla of B
PUMA dust detector/analyzer
Cosmic rays, Y jumps, etc.

TOTAL MASS 100 kg

Study of the surface: chemical composition of the soil, of the atmosphere near the soil, of volatile components; geophysical data by passive seismology, electrical readings. Priority is given to the study of areas which might show volcanic activity.

The atmosphere: study of atmospheric circulation through several types of long lifespan balloons, at various altitudes (including a "flying deer" consisting of two balloons linked by a cable); chemical and isotopic composition of gases and condensed phases, possible evidence of products of volcanic activity.

THE VENERA 90/VESTA MISSION: INTERNATIONAL CONTEXT

<u>Mission</u>	<u>Date</u>	<u>Type of Mission</u>
Vega Giotto	1986	<u>High-velocity flight</u> over an <u>active comet</u>
Galileo?	1986	<u>Distant</u> and rapid <u>flight</u> over <u>Amphitrite</u>
Phobos	1988	<u>Very slow</u> and <u>very close</u> (~ 50 m) <u>flight over Phobos</u>
Vesta	1992-1993	<u>Multiple flights over asteroids</u> and <u>comets</u>
Craf	1997-1999	<u>Rendezvous</u> with a <u>short-period comet</u>
Agora	1997-2002?	<u>Multiple rendezvous</u> with the <u>asteroids</u>

STUDY OF VENUS AND THE ASTEROIDS AND COMETS IN THE SOLAR SYSTEM
GENERAL MISSION PLAN

Launch date: End of 1992
Double mission: 2 proton launchers

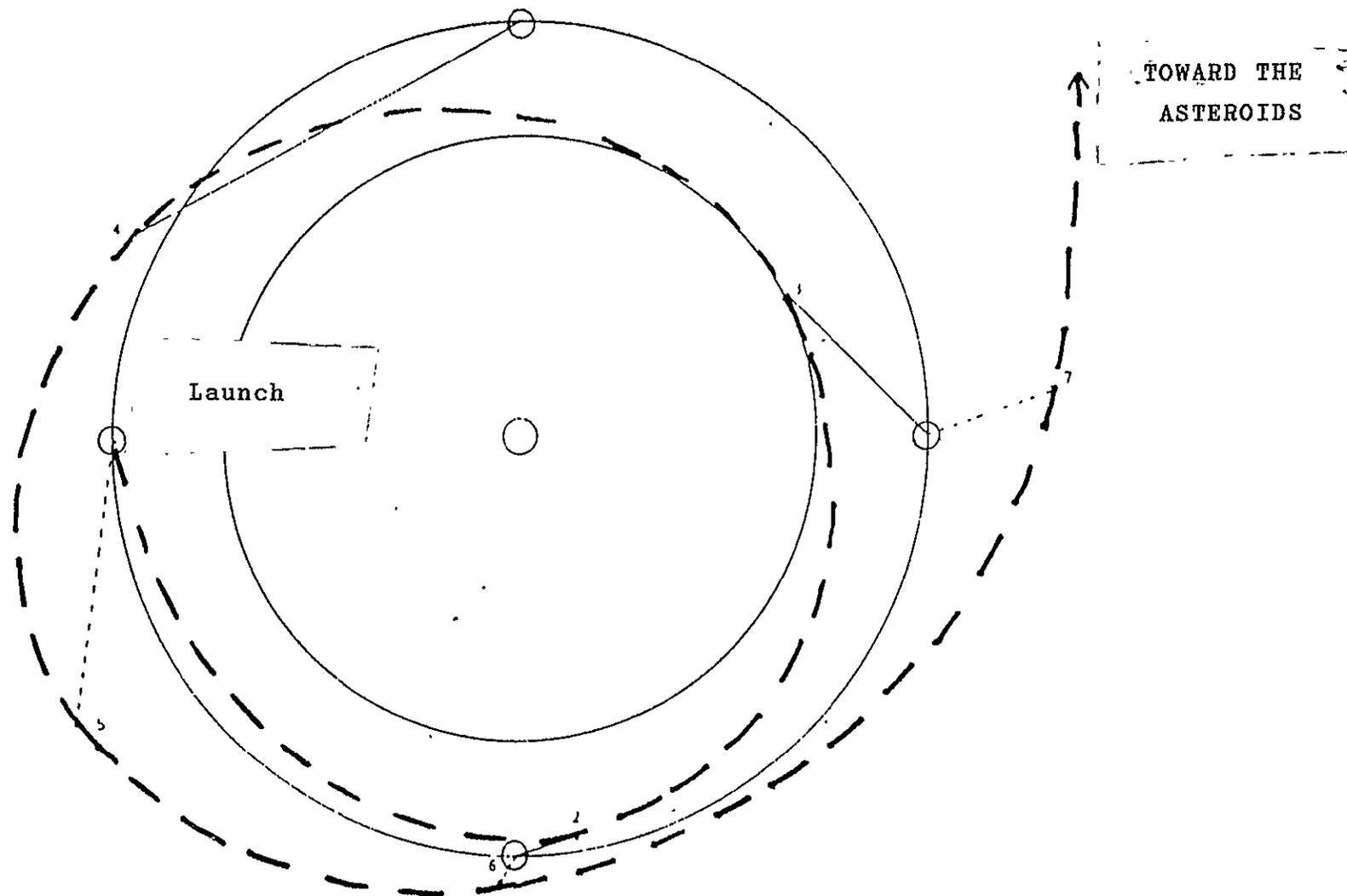
2 different vehicles:

- 1 Soviet vehicle Directed toward Venus
 Launches a descent vehicle there
 Then flies over a small body within
 the orbit of Mars

- 1 independent vehicle Independent from launch
 built under the Uses double gravitational assistance
 responsibility of Venus to Earth
 the CNES Fly over 3 small bodies in the
 principal asteroid belt

8 SMALL BODIES OVERFLOWN

EARTH - VENUS - EARTH TRAJECTORY



AN EXAMPLE OF TRAJECTORY
INCLUDING FLIGHTS OVER COMETS

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EARTH	Depart 12/18/92	V = 3.2 km/sec
VENUS	Arrive 5/13/93	V = 5.05 km/sec
EARTH	Return 2/14/94	V = 8.35 km/sec
33 POLYMNIE	Fly over 7/18/95	Type S asteroid Diameter 62 km V = 11.9 km/sec
EARTH	Return 6/12/96	V = 8.7 km/sec
P/KOPFF	Fly over 7/30/96	Comet (28 days after perihelion) V = 10.5 km/sec
P/GEHRELS	Fly over 7/18/97	Comet (28 days before perihelion) V = 10.1 km/sec

TOTAL DURATION OF MISSION ~ 4.6 years

Δ TOTAL V 400 m/sec

AN EXAMPLE OF TRAJECTORY
INCLUDING FLIGHTS OVER ASTEROIDS

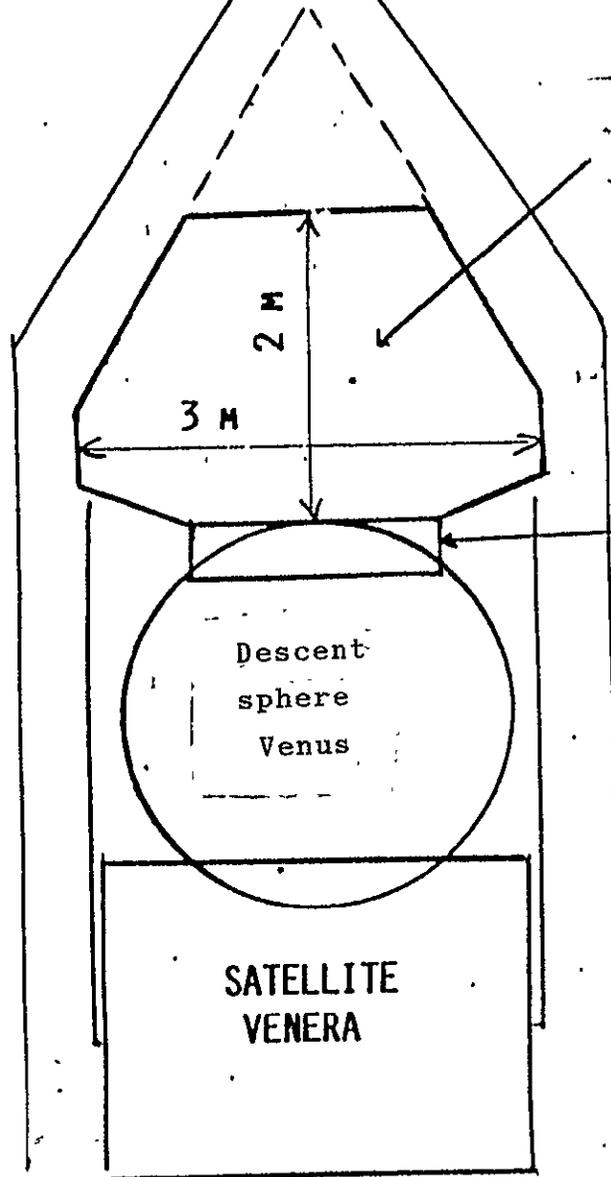
/11

EARTH	Depart 12/27/92	V = 3.1 km/sec
VENUS	Arrive 5/1/93	V = 4.7 km/sec
EARTH	Return 1/22/94	V = 7.9 km/sec
7 IRIS	Fly over 8/7/95	Type S asteroid Diameter 208 km V = 7.8 km/sec
EARTH	Return 1/20/96	V = 8.15 km/sec
4 VESTA	Fly over 8/4/96	Type V asteroid Diameter 576 km V = 12.2 km/sec
286 ICLEA	Fly over 4/27/98	Type CX asteroid Diameter 92 km V = 8 km/sec
TOTAL DURATION	5.3 years	
Δ TOTAL V	230 m/sec	

cnes

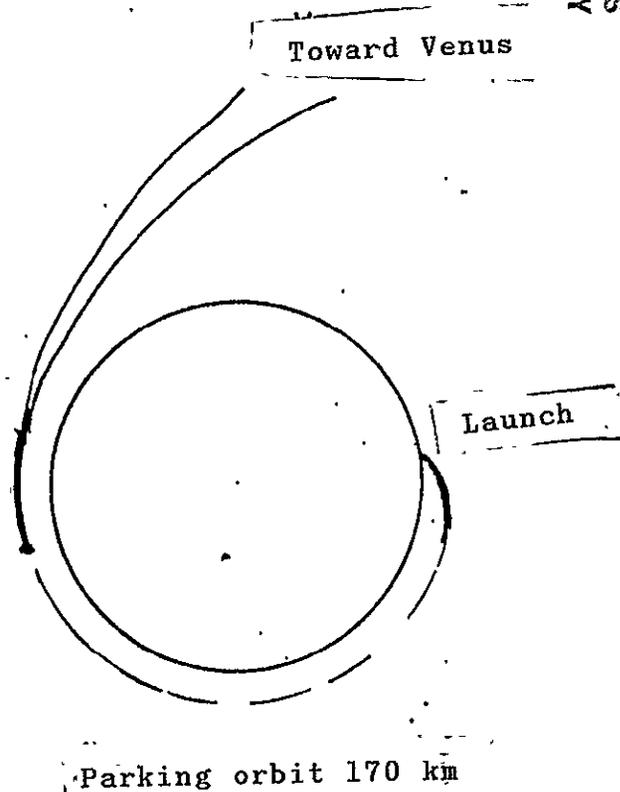
LAUNCH CONFIGURATION

French satellite
Mass \leq 870 kg



Chassis and
separation
(USSR)

Separation
4th stage

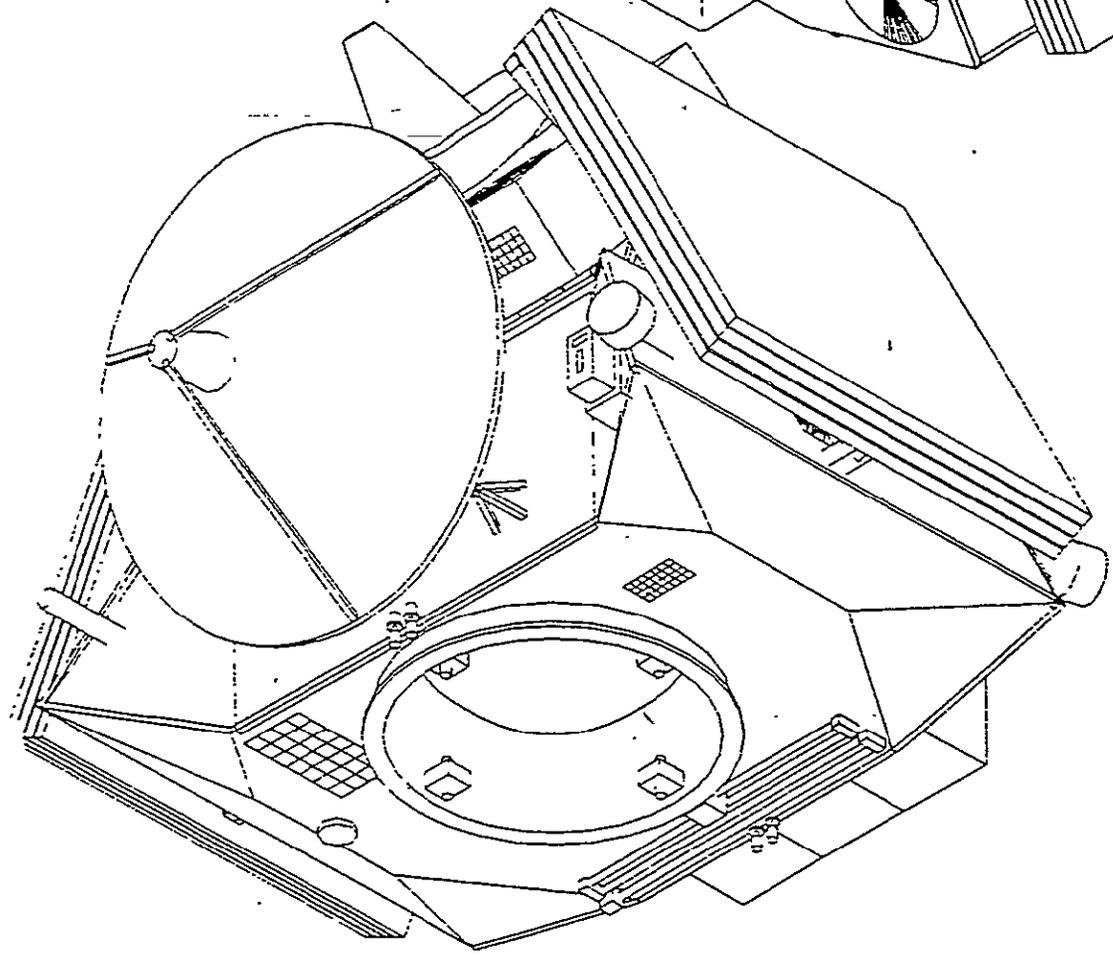
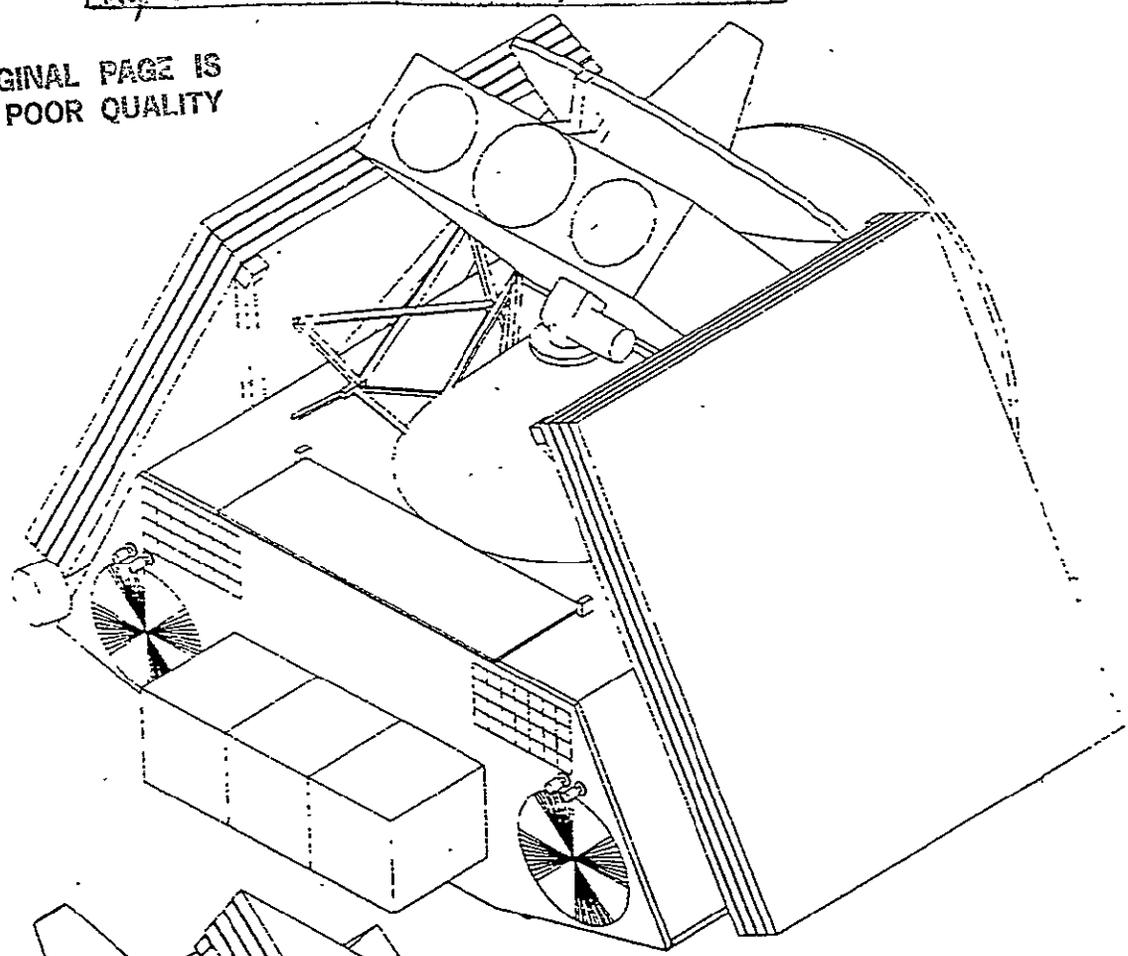


ORIGINAL PAGE IS
OF POOR QUALITY

Parking orbit 170 km

VEHICLE FOR ASTEROIDS AND COMETS

ORIGINAL PAGE IS
OF POOR QUALITY



CONFIGURATION IN OBSERVATION PHASE

