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ULTRAVIOLET SPECTROMETER AND POLARMETER
(UVSP) SOFTWARE DEVELOPMENT AND HARDWARE
TESTS FOR THE SOLAR MAXIMUM MISSION

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ULTRAVIOLET SPECTROMETER AND POLARMETER
(UVSP) SOFTWARE DEVELOPMENT AND HARDWARE
TESTS FOR THE SOLAR MAXIMUM MISSION

FINAL REPORT

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Contact NAS5-24119

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UVSP PROGRAM

FINAL REPORT

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UVSP PROGRAM
CONTRACT NAS5-24119
FINAL REPORT
PART I
HARDWARE PROGRAM

The Lockheed Role in UVSP

The Ultraviolet Spectrometer / Polarimeter Instrument (UVSP) for the Solar Maximum Mission was based on re-use of the engineering model of the high resolution ultraviolet spectrometer developed at the University of Colorado for the OSO-8 mission. Lockheed became involved in the UVSP program when Dr. Bruner, who had been the principal investigator on the OSO-8 program, joined the Space Astronomy Group at the Lockheed Palo Alto Research Laboratory. Lockheed assumed four distinct responsibilities in the UVSP program; technical evaluation of the OSO-8 engineering model, technical consulting on the electronic, optical and mechanical modifications to the OSO-8 engineering model hardware, design and development of the UVSP software system, and scientific participation in the operations and analysis phase of the mission. Lockheed also provided technical consulting and assistance with instrument hardware performance anomalies encountered during post launch operation of the SMM observatory. Appendix 1 to this report contains an index to the quarterly reports delivered under the contract, and serves as a useful capsule history of the program activity.

Initial Evaluation of the OSO-8 Hardware

The initial evaluation of the OSO-8 engineering instrument was carried out at the Lockheed Palo Alto Research Laboratory prior to delivery of the instrument to General Electric, the prime contractor in preparing the UVSP hardware. This initial evaluation established a performance baseline for the spectrometer, and revealed some problems with the existing electronic hardware. These tests focused on the performance of the wave-

length drive, particularly the computer controlled slew mode, which had given problems in operation of the OSO-8 instrument in orbit.

We also performed resolution tests on the spectrometer, using a mercury 198 lamp as a narrow line source. The resonance line at 2537 Angstroms was observed in first order, and found to have a width of about 0.025 A, a value consistent with its original performance when assembled at the University of Colorado.

Once the instrument had been delivered to GE, we supported the disassembly and inspection of the spectrometer with the specific objective of discovering the source of stray light that had affected the OSO-8 performance. This study revealed a design problem in the baffle system which, when coupled with certain misalignment conditions, would allow extreme off-axis rays from the entrance slit to be reflected by the Ebert mirror directly into the exit slit. The misalignment cannot be discovered when the instrument is fully assembled, as it is automatically compensated by adjusting the grating shaft angle during wavelength calibration. Laboratory calibration sources are too weak to reveal the stray light, and the condition only became known after the OSO-8 spectrometer had been launched. Our recommendations for correcting the condition were followed by the GE design team and were successful in controlling stray light in the UVSP instrument.

Optical System

In order to facilitate an understanding of the stray light problem and to serve as background for the discussion of the hardware modifications, we will briefly discuss the OSO-8 instrument hardware and optical system. A diagram of the OSO-8 instrument is shown in Figure 1. The OSO-8 spectrometer was an Ebert-Fastie system with a 1 meter focal length. It had an aperture ratio of $f/19$ in the plane of dispersion, and $f/15$ in the orthogonal direction. It was fed with a cassegrainian telescope of 12 cm aperture and 1.8 meter focal length. Dispersion was produced by a diffraction grating with a ruling frequency of 3600 grooves per millimeter operating in the second order for the range 1200

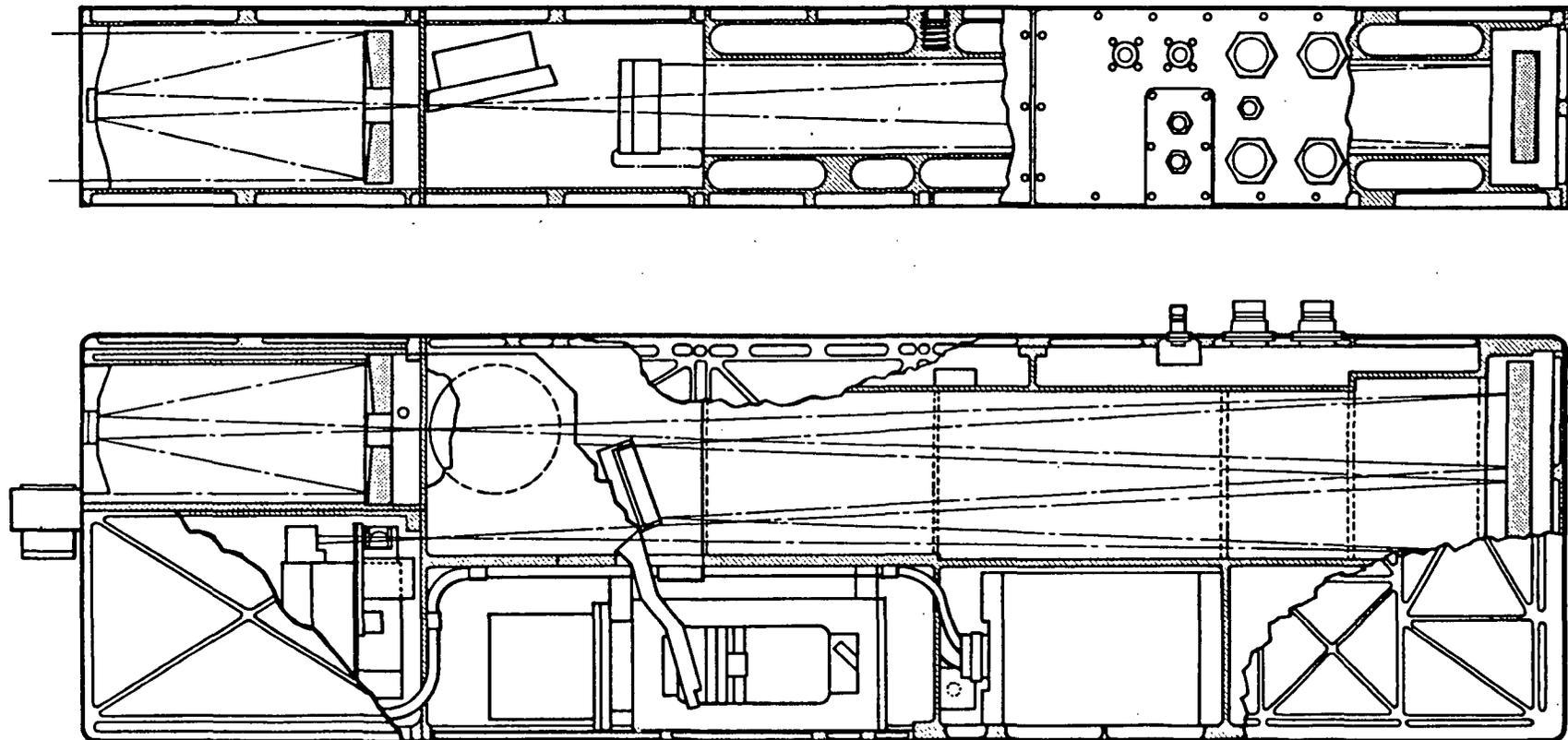


Figure 1. OSO-8 Instrument Diagram

to 1800 Angstroms. The spectrometer had fixed, straight entrance and exit slits of 8 micron width, corresponding to a wavelength interval of 0.01 Angstroms, again in second order. A movable slit mask, or dekker, was provided to allow the effective length of the slit to vary between about 40 microns and 8 millimeters, corresponding to an angular range of about 5 arc sec to 15 arc minutes on the sun. The wavelength passed by the spectrometer could be varied by rotating the grating about a shaft parallel to the grating grooves. Radiation emerging from the exit slit of the spectrometer was detected by a sealed photomultiplier tube operating in the pulse counting mode.

The control system for the OSO-8 spectrometer was based on a small, dedicated, general purpose computer which performed all of the primitive instrument functions under control of a flight software package. OSO-8 was the first such instrument to be flown in the NASA space program, and paved the way for the wide scale use of microprocessor control that characterizes contemporary instruments.

Mechanical System

The OSO-8 optics were supported by a modular structural system consisting of six major assemblies; the telescope, the spectrometer case, the master metering bracket, the wavelength drive, the detector assembly, and the Ebert mirror cell. The heart of the system was the master metering bracket. The grating assembly, which included the grating, the grating cell and its precision bearings and the grating arm, was built into the master metering bracket. The master metering bracket was kinematically mounted to the central wall of the instrument case. The grating drive was supported from three posts on the master metering bracket that protruded through holes in the central wall between the spectrometer compartment and the wavelength drive compartment. The telescope was also cantilever mounted to the master metering bracket via a set of three posts that passed through holes between the telescope cavity and the spectrometer compartment. The slit assembly was fastened to the front of the master metering bracket, just behind the telescope.

In this way, all major optical components except for the Ebert mirror were maintained in strict alignment through a single compact and extremely stiff structural element that was not subject to externally induced distortions of the instrument case. Moreover, this master bracket and the modules it carried could be assembled outside of the instrument case so that critical alignment could be done on a surface plate with standard mechanical metrology techniques.

Wavelength Drive

The wavelength drive was based on a screw and follower nut of a type that is manufactured by the Moore Special Tool Co. for use in their line of ultra-high precision machine tools and measuring engines. The screw was supported by multiple ball bearing races in a titanium housing whose thermal expansion coefficient closely matched that of the nitrided steel screw. The screw was coupled by a flexible metal bellows universal joint to the output of a precision spur gear reducer which, in turn, was driven by a 48 step brushless four phase DC stepping motor. A precision flat carried by the follower nut assembly contacted a steel ball mounted on the end of the grating arm, causing the latter to rotate the grating when the flat was moved by rotating the screw. Each step of the motor moved the flat approximately 15 microinches, altering the spectrometer wavelength setting by 5 milliangstroms. Our tests at Palo Alto demonstrated that the reproducibility of the drive for multiple settings of the screw was of the order of 5 microinches (1 sigma) or about 2 milliangstroms. The geometry of the arm, ball and flat was arranged so that they functioned as a sine-bar, forcing a linear relationship between wavelength setting and the screw position.

Since the screw had to be oil lubricated in order to function, the entire screw drive system was enclosed in a hermetically sealed housing. A stainless steel bellows allowed the nut assembly and flat to travel longitudinally and also prevented the nut assembly from rotating about the screw axis. The enclosure was fitted with redundant pressure relief valves to bleed the air out of the interior of the drive when the instrument entered the

vacuum of space. These pressure relief valves were spring loaded so as to close after the initial venting was complete.

Stray Light Problem

The optical condition that led to the stray light problem may be understood in terms of the diagram in Figure 2, which shows the effect of rotating the Ebert mirror through a small angle E about the vertex. This rotation causes the center of curvature to move up in the figure by an amount $2 \times E \times F$ where F is the focal length. This, in effect, redefines the axis of symmetry, since the spherical Ebert mirror has no unique axis. The new axis of symmetry, which passes through the center of curvature and the midpoint between the slits, will be displaced from the old one by an angle $2E$. If the grating is rotated by this same amount, then the image will again fall on the exit slit. Thus, an initial angular alignment error of the Ebert mirror cannot be discovered in the assembled instrument since the grating shaft angles are initially determined by scanning the spectrum and identifying lines. The error $2E$ will be absorbed in the calibration constants.

Notice that the new axis of symmetry of the system no longer intersects the Ebert mirror in its physical center as it would have in the case of nominal alignment. This means that a ray from the entrance slit to the displaced vertex will be reflected through the exit slit without ever striking the grating. These direct rays are normally blocked in an Ebert-Fastie spectrometer by a series of 3 stops, S1, S2, and S3 as shown. Due to a design error in the OSO-8 system, the stop S1 was too far from the chief ray, allowing radiation from the entrance slit to strike the Ebert mirror below the original axis of symmetry. Although the stop S2 would normally have blocked the undesirable central ray, the misalignment condition discussed above made S2 ineffective. Existence of this misalignment condition and of the improper location of the stop S1 were confirmed during the instrument disassembly at GE. Once the condition was understood, the corrective measures were clear. S1 was placed in its proper location, the Ebert mirror was carefully aligned to center the axis

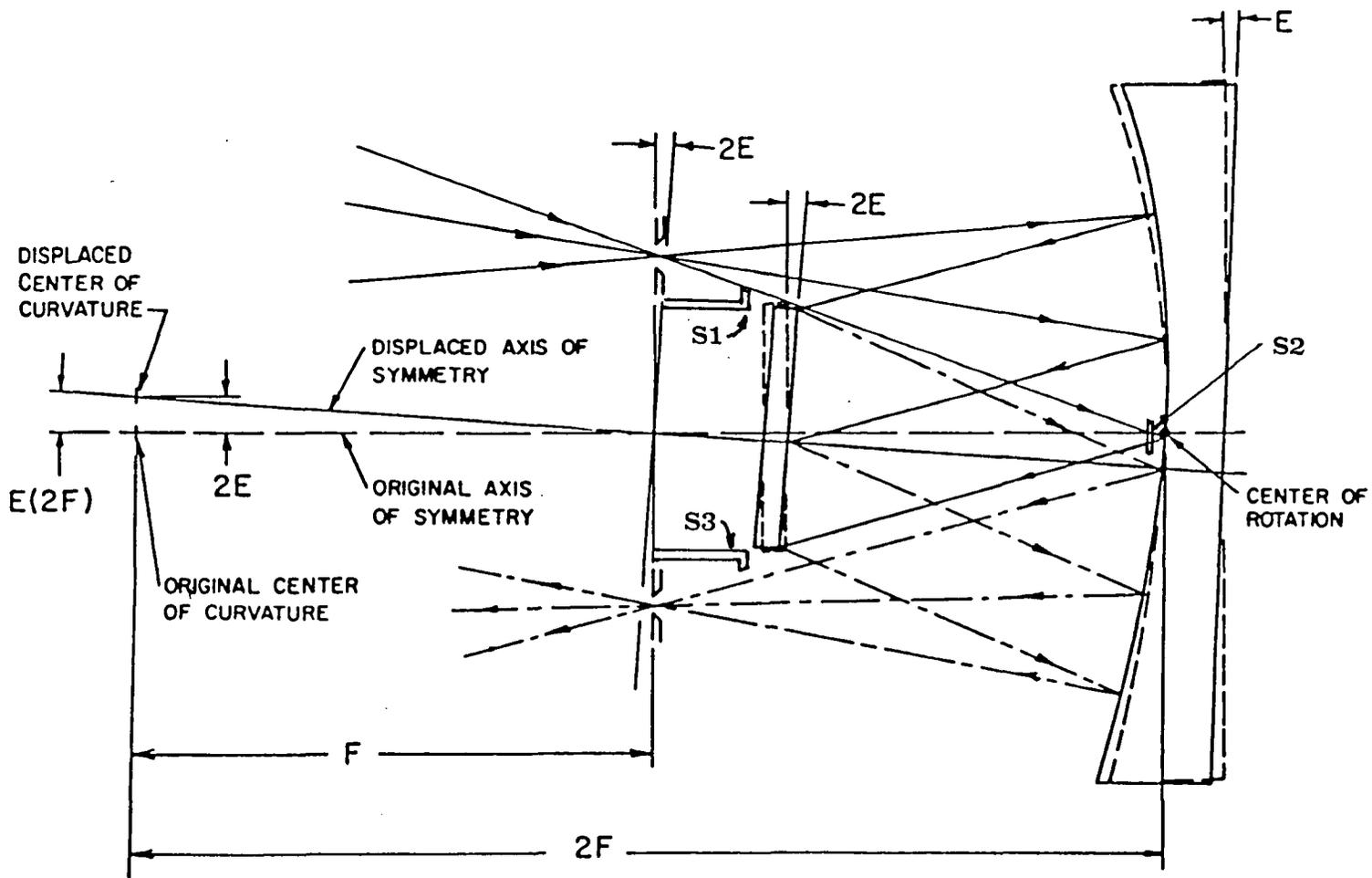


Figure 2 Effect of Ebert Mirror Misalignment

of symmetry, and stop S2 was redesigned to provide superior blocking of the central ray.

Modifications Required for the UVSP Mission

The UVSP instrument differed from its OSO-8 incarnation in four important ways: it used a gregorian telescope with an articulated secondary mirror instead of the original fixed cassegrainian telescope, it had a polarimeter capability, it had five detector channels instead of the two of OSO-8 and finally, it had a mechanism to interchange slits. It also featured a second generation operating system in the instrument computer. The polarimeter has been described by *Calvert. et al., 1979, Opt. Engin., 18, 287*). Addition of the polarimeter represented a major new capability with respect to OSO-8. The articulated telescope secondary was required both because the SMM spacecraft lacked a raster capability, and because rastering the spacecraft would be incompatible with several of the other SMM instruments. The multiple detector array and the interchangeable slits allowed us to define polychromatic positions for which two or more lines could be observed simultaneously. A diagram of the UVSP configuration is given in Figure 3.

Control of Sensitivity Loss

Many of the changes introduced into the UVSP instrument were designed to control a severe sensitivity loss experienced in orbit by the OSO-8 instrument. OSO-8 lost nearly two decades of sensitivity during the first week after launch, and the sensitivity at H-lyman alpha continued to drop by a factor of two every two weeks for the next few months. The cause of this severe and continuing loss was postulated to be the polymerization of outgassing contaminants onto the surfaces of the optics, especially in the telescope. Tests conducted at GSFC had showed that the degradation rate on test mirrors subjected to UV radiation under vacuum was controlled both by the concentration of outgassing effluent in the vicinity of a surface, and by the level of UV irradiance on that surface. The contamination rate in a cassegrainian instrument will be most severe on the secondary

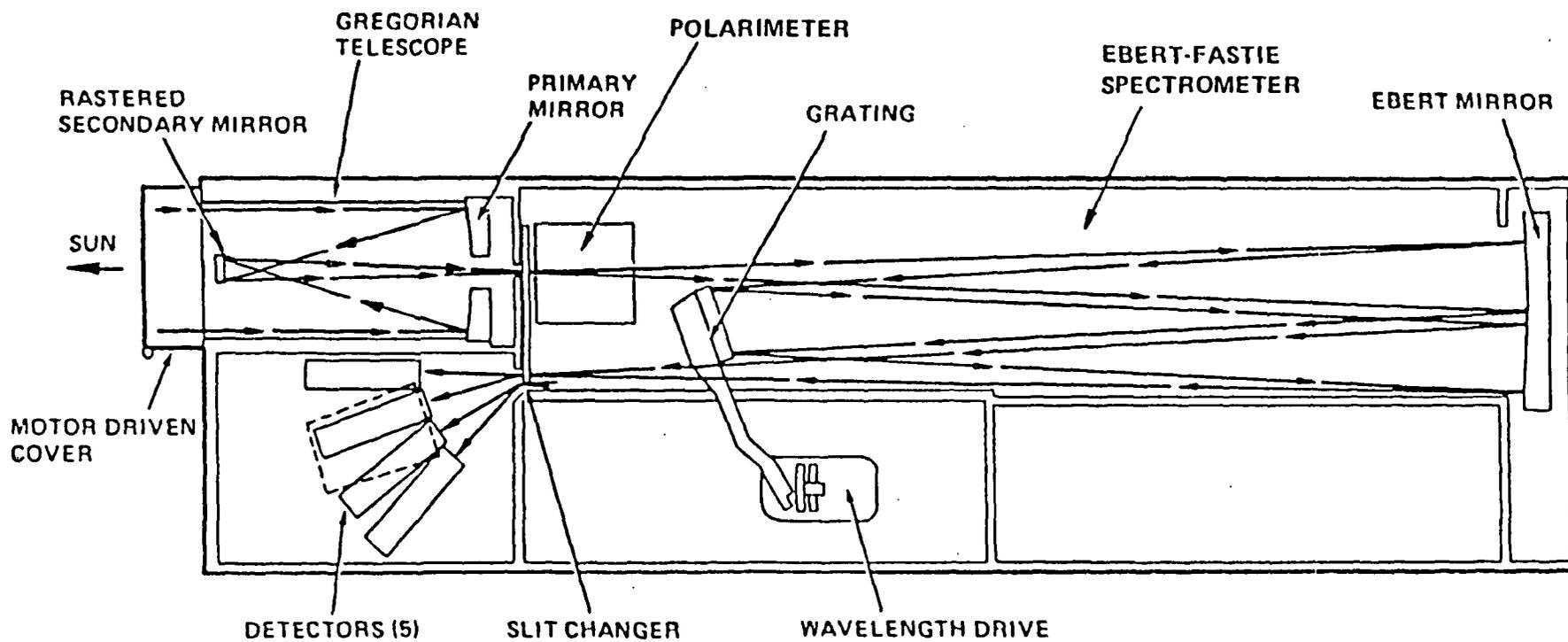


Figure 3. Layout of Ultraviolet Spectrometer and Polarimeter

mirror because of its proximity to the prime focus. Here, the irradiance is concentrated by the ratio of the unocculted area of the primary to the area of the illuminated portion of the secondary mirror. In OSO-8, this concentration factor was nearly twenty to one.

Use of gregorian, rather than cassegrainian optics allowed us to put a field stop at the prime focus of the UVSP telescope. This stop served to intercept most of the flux collected by the primary, passing only an 8×8 arc minute field to the secondary. In this way, most of the heat load from the incoming solar beam was captured by the stop and conducted away from the instrument structure by a system of copper bars and heat pipes. A more important effect of the field stop was to reduce the flux load on the secondary mirror so as to reduce the rate of sensitivity loss due to polymerization of outgassing contaminants. By carefully sizing the field stop, it was possible to make it large enough to intercept the solar disk for all pointing positions on the sun, yet small enough to fall completely within the shadow cone formed by the secondary mirror and its mount. The field stop reduced the flux load on the secondary mirror to less than 1.5 solar constants, and allowed a field of 256×256 arc seconds to be scanned by the secondary without vignetting.

Other measures taken to reduce the sensitivity loss due to contaminants were based on the philosophy of treating the instrument as an ultra-high vacuum system that needed to be baked out prior to being placed into service. An aperture door was added to the front of the instrument to keep solar radiation from entering until after the bakeout was complete. The existing structure heaters were replaced with heaters of higher capacity, and heaters were added to the backs of the two telescope mirrors. Personnel assembling and handling the instrument wore cotton or nylon gloves, rather than the vinyl gloves that had been used on OSO-8, as the vinyl had been shown to contaminate surfaces with plasticizers.

An in-orbit bakeout procedure was defined in three stages. In the first, the instrument heaters would be operated for several days with the aperture door closed during the

daylight portion of each orbit. The aperture door would be opened during each eclipse in order to assist in venting the payload cavity. This phase was designed to remove the bulk of the water vapor and other condensables from the main structure. In the second phase, the structure heaters would be turned off, but the mirror heaters left on so as to drive away residual contaminants left on the mirror surfaces. In the third phase, the structure heaters would be turned on as needed to maintain the instrument at its operating temperature, which was to be several degrees below the bakeout temperature. After the system was stabilized at the operating temperature, the telescope door would be opened and scientific observations started.

This philosophy was generally successful in lowering the sensitivity loss rate substantially below that of its OSO-8 predecessor. Some loss was noted, however, a few weeks after launch, and has been tentatively traced to a failure to carryout the bakeout procedure fully through the cool-down phase. The case heaters were left on to continue the outgassing in parallel with early operations of the instrument. When the satellite's orbit precessed into an orientation that provided a longer exposure to the sun during each orbit, the case temperature rose above the bakeout limit. We have postulated that this caused the residual internal pressure in the wavelength drive enclosure to rise, forcing the pressure relief valves to open, venting oil laden air into the instrument case. The case temperature was lowered as soon as the effect was noted, but it was, of course, too late to stop the degradation of reflectivity that had already occurred. Armed with the wisdom of hindsight, we now recognize that it would have been a good idea to provide for overboard dumping of the air vented from the wavelength drive mechanism. It is also clear that the bakeout procedure was fundamentally sound, and should have been followed more strictly.

Other Modifications

Another modification made to the OSO-8 baseline instrument was in the Ebert mirror cell. The OSO-8 version of the cell was provided with a focussing mechanism based on the

proving ring principle used in many Ebert-Fastie slit mechanisms. Although conceptually sound, the OSO-8 mechanism was found to have an undesirable flexure mode that reduced the position stability of the Ebert mirror. We modified the cell to remove the focus capability in favor of a much stiffer structure. Final focus and alignment were set in the laboratory by adjusting the thicknesses of spacers between the Ebert mirror cell and the spectrometer case. The only in-orbit focus capability lay in our ability to set the operating temperature of the instrument by controlling the case heaters.

The final modification that should be mentioned was the addition of a co-alignment system (the CAS) that supported the UVSP instrument in the SMM instrument support plate. This system was added at the suggestion of the SMM program manager, as it allowed him to delete what would have otherwise been a very expensive environmental test to assure that the UVSP would remain co-aligned with the other SMM instruments during and after launch. The CAS was a two axis gimbal system that fastened to the narrow edge of the UVSP case in the vicinity of the master metering bracket. Structural analysis of the case and CAS attachment was performed by GE and showed that the scheme would not degrade the mechanical integrity or stability of the case.

Optical Performance Evaluation

Lockheed performed an advisory and assistance role during the optical alignment, performance evaluation and testing of the UVSP. Our work on the alignment of the telescope is discussed in our progress report for the first quarter of 1978, and will not be treated at length here. The performance evaluation of the completed instrument began in the first quarter of 1979 and was carried out at the Goddard Space Flight Center. Here, we were concerned with evaluating the focus and resolving power of the telescope and of the spectrometer. Telescope focus was assessed with a Foucault test, using an incandescent lamp to backlight the entrance slit. The knife edge was placed at the focus of an auxiliary telescope which had previously been set to the infinity focus by autocollimation. Visual inspection

of the aperture illumination pattern suggested the presence of residual aberrations (principally coma), but at a level that was within the UVSP performance requirements. The focus was judged to be acceptable.

Focus of the spectrometer was assessed with a modified form of the Hartmann test, discovered accidentally during the course of performance evaluation. In this test, the telescope was illuminated with a "Pen Ray" mercury lamp, oriented such that its long dimension was parallel to the rulings on the diffraction grating. The lamp was uncollimated, and was placed a few centimeters in front of the telescope aperture. The lamp was mounted on a rack and pinion mechanism so that it could be translated in a direction perpendicular to the rulings. The optics of the telescope, together with the entrance slit, acted to admit to the spectrometer, a single fan of rays that was parallel to the grating rulings; i.e. a sagittal fan. Motion of the lamp with the rack and pinion mechanism swept this fan across the grating. If the spectrometer was in focus, then the position of the spectrum in the focal plane would be independent of where the fan struck the grating. If, however, the focus was incorrect, as proved to be the case, then the image of the spectrum would appear to move when the lamp position was changed. By measuring the position of a spectrum line on the wavelength drive as a function of the position of the lamp, we were able to infer both the amount and the direction of the focus error in the spectrometer. The focus error was corrected on the first attempt by re-shimming the Ebert mirror cell.

In the calibration of the completed UVSP spectrometer, Lockheed played primarily a supporting role, assisting in the operation of the instrument in a calibration system designed and prepared by GSFC personnel. Lockheed personnel were in residence at GSFC during the calibration activity, helping both with the installation and checkout of instrument control software and with the collection of the primary calibration data set. Responsibility for the reduction and analysis of the calibration data lay with the GSFC project scientist, and will not be discussed here.

During the post launch checkout of the SMM instruments, Lockheed personnel were in residence at GSFC. Mr. R. Rehse, who developed the flight software package for the instrument, carried the responsibility of verifying the operation of the instrument under software control, identifying and correcting logic problems that came to light as a result of in-orbit experience. Dr.'s Bruner and Schoolman each spent periods of time in the SMM Experiment Operations Facility (EOF). taking part in the observatory operation, developing quick-look data inspection procedures as required by the newly acquired data. These procedures were typically written, checked out, and installed in the PDP-11/34 computer in the EOF when complete. Schoolman also supported the command generation software package, making modifications as needed to improve its operation. Subsequent analysis codes, developed during the course of scientific investigations carried out by the Lockheed experimenters, have been included in the relevant quarterly reports when they were felt to be of general utility. The software system will be discussed in Part II of this report, and the program of scientific investigations is treated in Part III.

PART II

UVSP SYSTEM SOFTWARE

Software Systems Approach

The UVSP software system was based on experience with the UV spectrometer experiment on OSO-8, and many of the elements of that system were carried over directly to the UVSP system. The software system has two major components; ground test software, and mission software. The ground test software included a test interpreter language which operated the instrument in its primitive modes, and a set of hardware test procedures written in the test interpreter language that were used to test individual instrument components. The test interpreter was also used to operate the instrument during performance evaluation and calibration.

The mission software system includes a Command Generation System, a Flight Software Package, a Data Acquisition Package, a Data Reformatting Package, and three data analysis languages. The heart of the UVSP mission software system is the flight computer software package, which contains all of the control logic required for making solar observations. The computer executes observing sequences defined by an observing list contained within its memory. This list is loaded from the ground on a daily basis, according to the needs of the overall SMM observing strategy. Contents of the observing list are prepared with another software package called the Command Generation System, which translates the observing requirements from human readable form into the bit packed format required by the flight software package. When an experiment is executed by the computer, the data stream is tagged with identifying information so that it can later be automatically sorted into logical data files.

The data stream from the entire SMM observatory is transmitted to the experiment operations facility via high speed data lines either in real time during ground contacts with the spacecraft, or as tape recorder playback data that has been recorded at the

ground tracking stations and then re-transmitted at later times. The incoming telemetry stream is scanned by the Data Acquisition Package, which captures the relevant portion and stores it in a large disk memory. Once the data is resident on the disk, it is processed by the Reformatting Package which uses the identifying information to block the data into logical experiment sequences, arrange it into formats convenient for analysis, append record headers containing the identifying information and other pertinent spacecraft data, and store the results in disk files. The data analysis languages allow an experimenter to readily access experiment data files to inspect the results in tabular or graphic form, or as images where appropriate. The languages are also general purpose computational tools that can be used to mathematically manipulate the data in order to extract its physical information content.

UVSP Test Interpreter Language

This software package was prepared and delivered under a previous contract, but is discussed briefly here for completeness. It was originally written for the OSO-8 program at the University of Colorado. The purpose of the package was to provide an easy-to-use system that could be used to operate the mechanisms of the spectrometer during instrument development and testing, and to inspect the resulting telemetry stream. The software is written in Macro, the PDP-11 assembly language, and operated under the DOS disk operating system. It permitted one to send commands via any of the command interfaces, accepting its input in the form of mnemonics that were abbreviations of the respective command functions. Commands that required the transmission of a numerical value would have the value appended to the mnemonic. Commands could be transmitted directly from the keyboard, or could be grouped together into a procedure and transmitted as part of an automated sequence. When a procedure was being run, the package would accept telemetry data from the instrument, loading it into a buffer that could be accessed by instructions in the test language. It was also possible to capture and store data on

disk files for more extensive later analysis. The language was provided with rudimentary programming instructions, including loop and branch capability, and simple arithmetic operations. Subroutine jumps were not implemented as such, though we found that they could be made through an indirect jump to a numerical label. Lockheed work on the software package consisted of implementing the new set of mnemonic instructions required by the revised electronic package, accomodating the new telemetry format and rate, and interfacing it to the new spacecraft simulator. Ideally, the package would have been re-written so that it would operate under the RSX-11M operating system rather than DOS. We elected to use it in the DOS environment both to save cost and to assure that the schedule would be met, although this decision left us with a somewhat awkward situation in which we had to translate the data files and change operating systems in order to evaluate the data with one of the more powerful data analysis languages.

Flight Software Package

The flight computer for the OSO-8 instrument has, for historical reasons, been known as "Junior" or Jr for short. We followed this notation throughout the UVSP program, and will use it in this document. The architecture of the Jr mission software was studied early in the development effort to determine its optimum form. Parts of the code, including wavelength drive and double precision arithmetic subroutines and the monitor section, were taken directly from the OSO-8 code. We added an extensive conditional response facility called the command mode, which allows the scientist to specify a flexible experiment whose actual execution will depend on the conditions detected from the sun. A sequence to locate the brightest point in a field and use it as a target for subsequent observations is an example of the use of the command mode. This type of sequence was heavily used during the mission.

The parameters that specify the sequence of device motions and other operations needed to define an observing mode were packed into a nine word parameter block con-

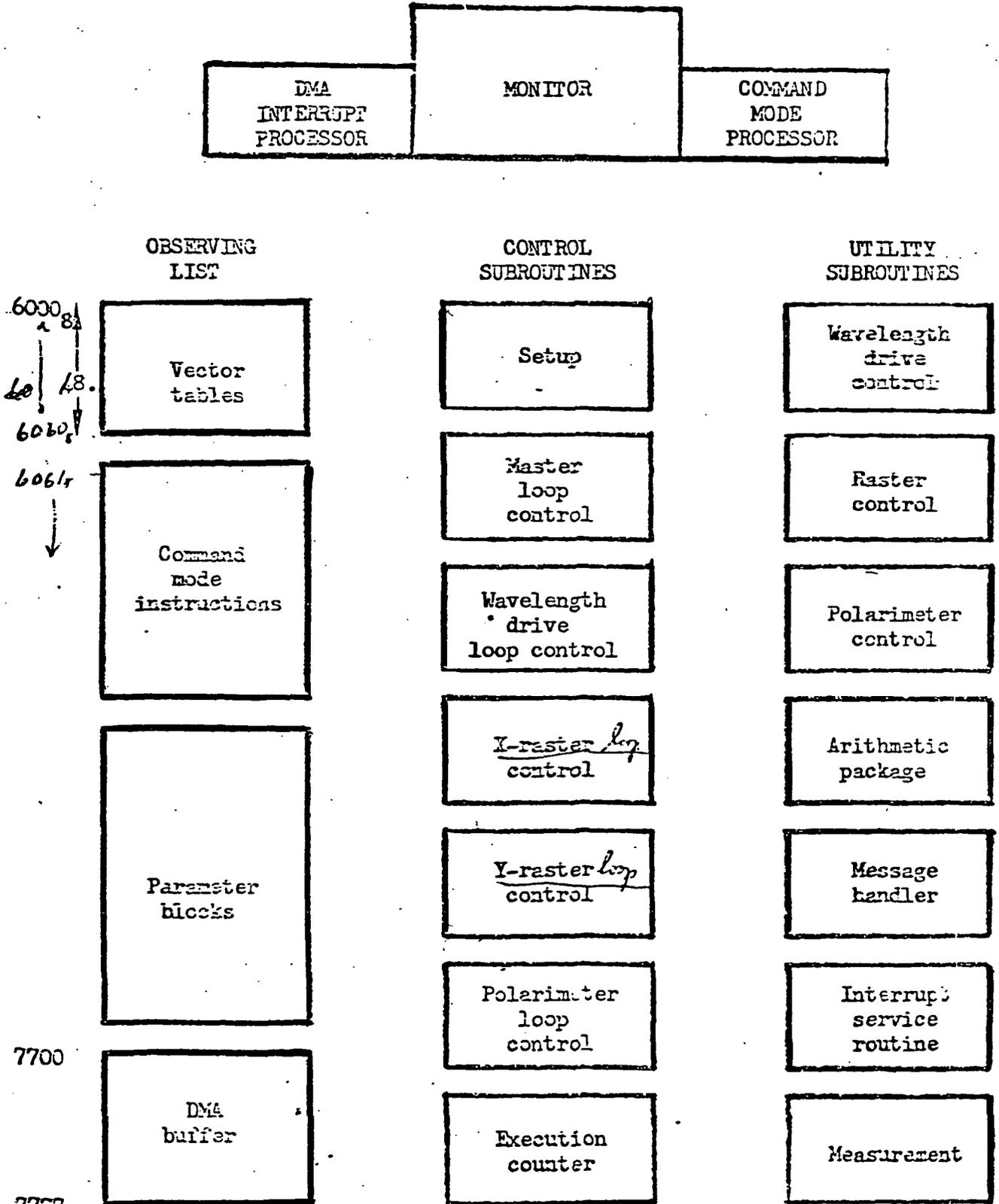
taining twenty different parameters. New control software was written to unpack the parameter block, configure the electronics and mechanisms, and load the appropriate loop counters. Goals defined for the flight software package were to accommodate the new hardware, to implement the new observing modes that it made possible, and to provide at least the same level of control as allowed by the OSO-8 software.

The following discussion of the UVSP flight software package is taken from the paper by *Woodgate et al.* (*Solar Phys.* 1980, 65, 73) discussing the instrument. The software system for UVSP is similar to that of the OSO-8 instrument from which it was derived. The OSO-8 operating system was described by *Hansen and Bruner* (*Space Science Instrumentation*, 1979, 5, 3). The overall organization of the flight software package is shown in Figure 4. Control of the program is directed by the monitor which processes normal instrument and timing interrupts, initiates and terminates observing sequences, and controls auxiliary functions such as operation of instrument heaters, etc. Important sub-functions of the monitor are the DMA interrupt processor, which accepts and interprets the one-word messages from the spacecraft through which the instrument operation is externally directed; and the command mode processor, which decodes the pseudo-instructions of the command mode language and calls up the appropriate subroutines implied by each command code.

The remainder of the code consists of three parts: the control subroutines, a set of utility subroutines, and the observing list. The control subroutines contain the control loops for each of the mechanisms. The ordering of the control loops as well as the extent and increment size for each is set by the nine word parameter block discussed above. The utility subroutine section contains the subroutines for the operation of all of the instrument mechanisms and for the control of the data flow to the telemetry system. Approximately 75 percent of the 4096 word memory is devoted to the monitor and the various subroutines.

Most of the remaining quarter of the memory is set aside as an observing list for the storage of the parameter blocks corresponding to the observing modes needed for a day's

Figure 4. JR FLIGHT SOFTWARE



observations. Contents of the observing list is divided into three major sectors, called the A, B, and C lists respectively. Each list is further divided into 16 sub-sectors, each with its own entry point. The A and B lists were each intended to hold the set observing modes that would typically be required to carry out a single day's operation, with the B list being loaded while the A list was running, and vice-versa. The C list was intended to hold a set of resident operating modes that could be loaded and held in readiness for use in observing rare events for which a quick response would be needed. The last 64 words of memory are a software status buffer which is read out synchronously into the telemetry system. This status buffer provided much the additional information needed to identify the operating mode of the instrument and its current state. The structure of the observing list is shown in Figure 5.

The command mode, which provides for data dependent control of the instrument, consists of a set of pseudo-instructions that can be entered in the observing list along with the experiment parameter blocks. These pseudo-instructions are one word coded subroutine calls to the master program. Through them, the experimenter has access to a block of 32 words of storage which is set aside as a user memory. User memory locations are allocated for the raster mechanism position, a wavelength drive reference position, flare coordinates from the Hard X-ray Imaging System instrument (HXIS), spacecraft flare status, and a number of critical parameters derived from the observations. Eight of the words are assigned as general use registers. Pseudo instructions are defined to move data from one place in user memory to another, to initiate an observing mode (parameter block), to transfer control to another pseudo-instruction or group of instructions, and to insert messages into the telemetry stream. A timekeeping function is also available. A list of pseudo-instructions and their corresponding bit patterns is given in Figure 6. Contents of the parameter block are given in Figure 7.

During the execution of an observing sequence, the monitor continually scans the

UVSP EXPERIMENT CONTROL

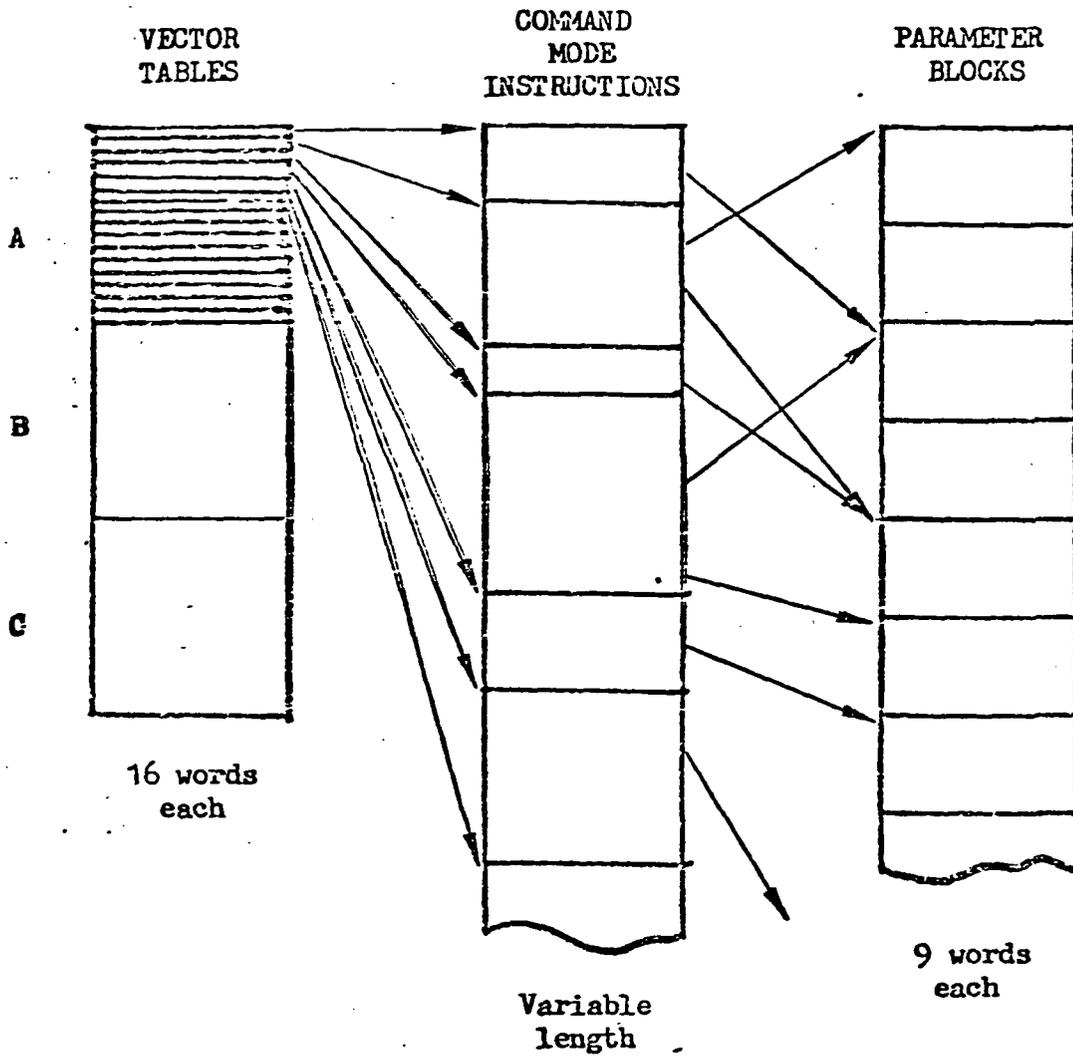


Figure 5.
OBSERVING LIST

COMMAND MODE INSTRUCTIONS

REVISED 31 OCT 78 RJL

- 0 EXECUTE**
- | | | |
|-----------|----------------------|---------|
| WAIT / GO | PARAMETER LIST INDEX | 0 0 0 0 |
|-----------|----------------------|---------|
- Use parameters in specified list to do an experiment $0 < N < 1617_6$
- 1 START**
- | | | | |
|-----------|---|------------|---------|
| WAIT / GO | R | LIST INDEX | 0 0 0 1 |
|-----------|---|------------|---------|
- R=0 Transfer control to the start of a segment in the current list
R=1 Transfer control to the start of a segment in the 'C' list.
- 2 GOTO**
- | | | |
|-----------|-----------------|---------|
| WAIT / GO | RELATIVE OFFSET | 0 0 1 0 |
|-----------|-----------------|---------|
- Transfer control to a command within the current segment $PC \leftarrow PC + 1 + N$
N is a 2's complement number and is non-zero.
- 3 COMPARE**
- | | | | | |
|-----------|---|---|---|---------|
| WAIT / GO | C | A | B | 0 0 1 1 |
|-----------|---|---|---|---------|
- | | | | |
|---|---|---|---|
| C | A | B | d |
|---|---|---|---|
- 1 0 dst Immediate value compared to RW or RD location as $(B) \leq d$ then skip
0 src dst RW or RD value compared to RW or RD location as $(B) \leq (A)$ then skip
- 4 ADD**
- | | | | | |
|-----------|---|---|---|---------|
| WAIT / GO | C | A | B | 0 1 0 0 |
|-----------|---|---|---|---------|
- | | | | |
|---|---|---|---|
| C | A | B | d |
|---|---|---|---|
- 1 0 dst Add immediate value to RW location $(B) \leftarrow (B) + d$
0 src dst Add RW or RD value to RW location
- 5 SUBTRACT**
- | | | | | |
|-----------|---|---|---|---------|
| WAIT / GO | C | A | B | 0 1 0 1 |
|-----------|---|---|---|---------|
- | | | | |
|---|---|---|---|
| C | A | B | d |
|---|---|---|---|
- 1 0 dst Subtract immediate value from RW location $(B) \leftarrow (B) - d$
0 src dst Subtract RW or RD value from RW location $(B) \leftarrow (B) - (A)$
- 6 MOVE**
- | | | | | |
|-----------|---|---|---|---------|
| WAIT / GO | C | A | B | 0 1 1 0 |
|-----------|---|---|---|---------|
- | | | | |
|---|---|---|---|
| C | A | B | d |
|---|---|---|---|
- 1 0 dst Move immediate value to RW or RD location $(B) \leftarrow d$
0 src dst Move RW or RD value to RW or RD location $(B) \leftarrow (A)$
- 7 COMPARE REVERSE**
- | | | | | |
|-----------|---|---|---|---------|
| WAIT / GO | C | A | B | 0 1 1 1 |
|-----------|---|---|---|---------|
- | | | | |
|---|---|---|---|
| C | A | B | d |
|---|---|---|---|
- 1 0 dst Compare immediate value with RW or RD location as $(B) \geq d$ then skip
0 src dst Compare RW or RD value with RW or RD location $(B) \geq (A)$ then skip
- 10 MSG**
- | | | |
|-----------|-------------------|---------|
| WAIT / GO | USER MESSAGE DATA | 1 0 0 0 |
|-----------|-------------------|---------|
- Pass user message data into TM data stream.

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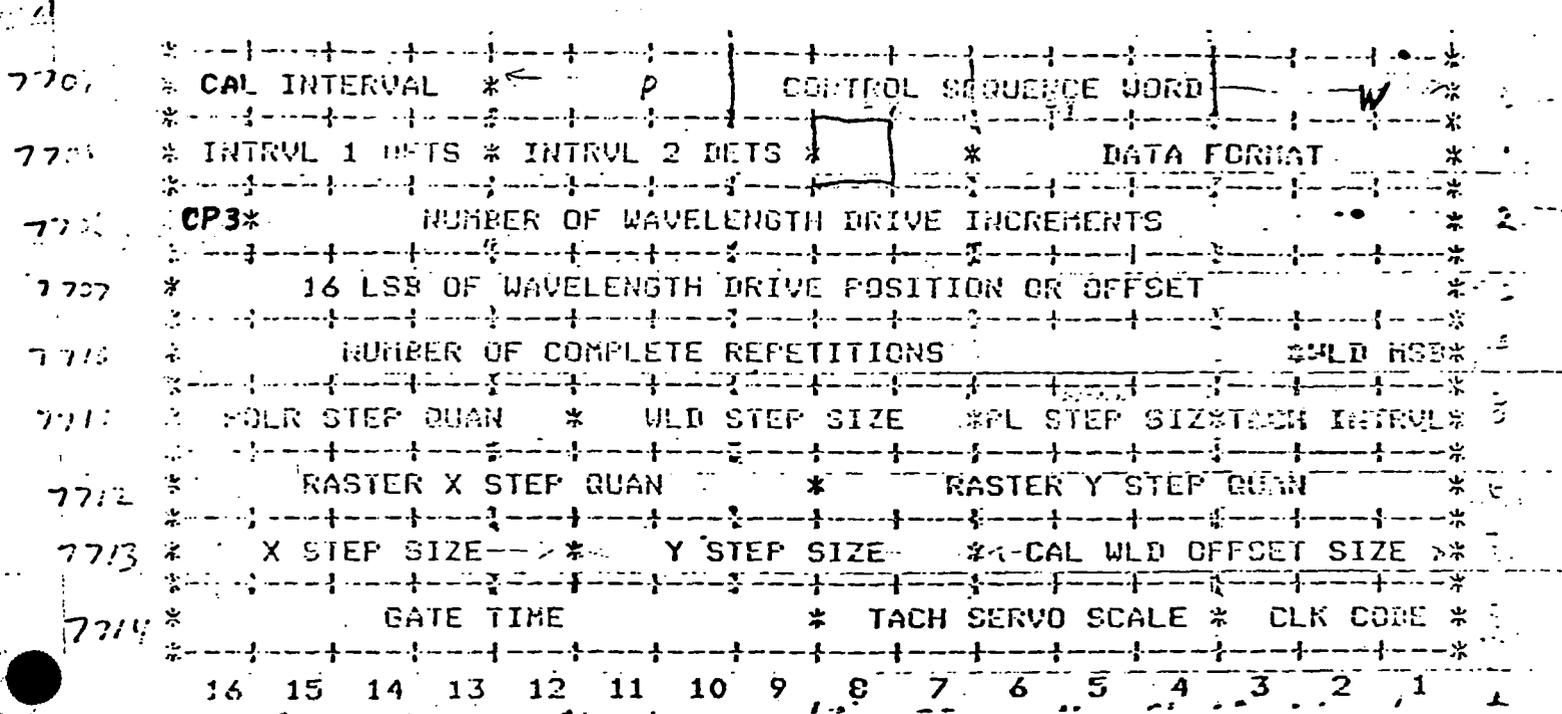


Figure 7a. SMM OBSERVING LIST PARAMETER BLOCK STRUCTURE

PARAMETER BLOCK
WORD 1



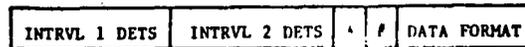
Cal Interval

- 0 disables calibration offset stepping
- > defines $2^n - 1$ as the calibration interval

Control Sequence Word

- 12 bit field is subdivided into four 3 bit fields
- Each 3 bit field uses MS bit as calibration internal count flag
- Only one 3 bit field may have the calibration flag bit set.
- No calibration flag bits is legal
- Low 2 bits of each 3 bit field specifies the device control loop sequence
- Bits 1-3 are assigned to WLD
- Bits 4-6 are assigned to Raster Y (outer gimbal)
- Bits 7-9 are assigned to Raster X (inner gimbal)
- Bits 10-12 are assigned to Polarimeter Rotation
- Sequence codes in the least significant 2 bits of each 3 bit field
- 0 Inner nested loop
- 1 Next to inner nested loop
- 2 Next to outer nested loop
- 3 Outer nested loop

PARAMETER BLOCK
WORD 2



Interval 1/2 Detectors contain bit code for detector routing and power

- 0 none
- 1 Det. 1
- 2 Det. 2
- 3 Det. 3
- 4 Det. 4
- 5 Det. 5
- 6 none
- 7 clock
- 10 Det. 1/2
- 11 Det. 2/1
- 12 Det. 1/4
- 13 Det. 4/1
- 14 Det. 2/3
- 15 Det. 3/2
- 16 Det. 3/4
- 17 Det. 4/3

For detector pair ordering, the first detector should be on blue side of line, second detector on red side of line.

PARAMETER BLOCK, WORD 2 (CONTINUED)

- * bit controls experiment number incrementing for multiline scans
- 0 suppresses incrementing experiment number
- 1 experiment number increments - normal condition

bit is unused

Data Format is a 6 bit code related to experiment type. Not used for control purposes.

PARAMETER BLOCK
WORD 3



OFS

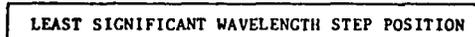
OFS is a bit specifying an offset from a previous wavelength position

- 0 Wavelength is direct from parameter block words 4 and 5
- 1 Wavelength is an offset from a local or global maximum result
- Offset is 16 bits in word 4
- Local/global selection bit is in word 5

Number of Wavelength Drive Increments

- 0 disables operation of WLD control loop - constant λ used
- >0 n+1 observations are made in control loop

WORD 4



Wavelength Step Position/Offset

- For OFS bit (offset) = 0 this is 16 bits of WLD position
- For OFS bit = 1 this is 16 bit WLD offset (2's complement).
- Offset base value may be selected - see word 5

PARAMETER BLOCK
WORD 5



Number of complete repetitions provides 14 bit value of experiment repetitions

WLD* is dual purpose field.

- For OFS = 0 (no offset) this is the most significant bit of WL position
- For OFS = 1 (offset) this selects whether a local or global WL value is selected
- 0 selects the global wavelength base for offset experiments
- 1 selects the local wavelength base for offset experiments

Figure 7b.

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PARAMETER BLOCK
WORD 6

POLR STEP QUAN.	WLD STEP SIZE	PL STEP SIZ.	TACH INTRVL
16	12 11	7 6	4 3 1

Polarimeter Step Quantity

0 disables polarimeter control loop
n>0 n cycles of polarimeter control loop are made

WLD Step Size

Simple count of WLD steps for each pass of WLD control loop

PL Step Size

n + 1 steps are made for each pass of enabled polarimeter loop

Tach Interval

0 disables tachogram servo loop
>0 causes 2ⁿ control loop passes to occur between WLD servo corrections.

WORD 7

RASTER X STEP QUAN	RASTER Y STEP QUAN
16	8

Raster X Step Quan.

0 disables X raster control loop
0 causes n raster positions before next control loop level

Raster Y Step Quan. same as X Step Quan.

WORD 8

X STEP SIZE	Y STEP SIZE	CAL. WLD OFFSET
16	12 11	6

X Step Size determines change of raster position for each control loop call when enabled.

Y Step Size determines change of Y raster position for each control loop call when executed.

Cal. WLD Offset determines quantity of WLD steps for calibration offset if calibration is enabled.

PARAMETER BLOCK
WORD 9

GATE TIME COUNT	DET. BALANCE	CLOCK
16	9	4 3 1

Gate Time Count

Value used to set gate time count down register

Det. Balance

Signed 4 bit value applied to balance detector output for tachogram servo control and user mode velocity values.
The equation for 1st detector is:

$$((30 + \text{Det. Bal.}) * \text{Blue Counts})/30$$

The equation for 2nd detector is:

$$((30 - \text{Det. Bal.}) * \text{Red Counts})/30$$

This has the effect of differential corrections of up to $\pm 3 X$ with steps of about .07 X.

Clock determines period of pulses which decrement the gate time count register

- 0 62.5 μ sec
- 1 500 μ sec
- 2 8 msec
- 3 32 msec
- 4 128 msec
- 5-7 not used

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Figure 7c.

data stream and maintains a record in the 32 word user memory of critical data elements, including the locations and intensities of the brightest and faintest elements of each raster, the most red-shifted and most blue-shifted elements of each raster, and the most intense wavelength of each spectral scan. After completion of a raster or spectral scan, command mode instructions may be used to test the critical data elements against pre-determined thresholds and alter the observing program accordingly, to adjust the instrument pointing so as to view one of the identified spatial elements, or to select a spectral line for subsequent observations.

Suitable parameter blocks together with appropriate groups of command mode instructions define the identified observing modes of the UVSP. The basic modes are the spectroheliogram, the dopplergram, the spectrogram, the polargram, and the magnetogram. Command mode instructions are not required by these modes, except to initiate their execution by activating the appropriate parameter block. In these basic modes, either the wavelength drive or the raster mechanism is scanned, but not both. Combined operation of the two mechanisms is provided in two modes, the profile matrix, in which an entire line profile is measured at each point in the raster, and the raster-over-line mode (RL) in which a complete raster is made at each of several wavelengths in a line profile. The basic observing modes are supplemented with command mode instructions to form another class of combined modes which includes a bright point finder, a faint point finder (useful for locating sunspots in the continuum), a flare finder (bright point finder plus threshold test), an upflow (blue shift) finder, a downflow (red shift) finder, and a spectrum line finder. These modes are useful for identifying targets and initiation times for subsequent observations. Within each mode, the sampling frequency and ranges of spatial and spatial scanning, the integration time, and the number of observations are all parameters that may be adjusted in order to optimize the observing program.

Experiment modes may be linked in memory by short programs of command mode

instructions to form more complex observing sequences. Each such sequence occupies one of the 48 observing list sub-sectors. Since the sector boundaries are defined by a vector table, the sector sizes may be adjusted from day to day to accommodate changing observing requirements. Initiation of an observing sequence is under control of the spacecraft and requires only one command, designating which sector is to be used. This command is identified by the DMA interrupt processor, which passes the appropriate vector label to the monitor. As shown in Figure 5, the vector points to the starting address of the memory sector containing the command mode instructions for that sequence. The instruction sector will, in turn, contain one or more command mode instructions that call for execution of observations under control of a parameter block, and will contain a pointer to that block. Note that parameter blocks do not need to be in any particular order, that a given sector may call up two or more parameter blocks, and that a parameter block may be called from more than one sector.

Once a parameter block has been identified, its contents are unpacked by the monitor and used to load the various loop counters and to set all required internal parameters and switches appropriate to the observations to be made. The nesting order of the control subroutine loops is also set at this time, and then execution begins. When the observing sequence specified by the parameter block is complete, control returns to the monitor, which fetches and processes the next command mode instruction. The last command mode instruction in each sector contains a flag signalling the end of the entire experiment sequence, whereupon control passes back to the monitor and the system waits for a new DMA interrupt command from the spacecraft.

An important aspect of the flight software package is that it makes the telemetry stream self-identifying. This is done in two steps. The first is to place the parameter block being executed into the software status buffer so that it is present in every major frame of telemetry. A unique serial number is assigned to each sequence by the master

program, permitting the telemetry stream to be divided into logical experiment sequences on the ground. These logical sequences become separate files in the data base after they are processed by the ground based reformatting program.

In the second step, the progress of execution of an experiment sequence is reported by having the computer inject messages into the spectrometer data stream at the conclusion of each pass through each of the control subroutine loops. Mode initiation is flagged by a unique message word, followed by the nine word parameter block defining the mode. This feature also permits the identification of experiments shorter than one major frame.

The most significant bit of each pair of spectrometer data words is used as a flag to permit the ground software to discriminate between messages and intensity data. Fill data is distinguished from intensity information by a hardware feature that resets the pulse counters to unity rather than zero. Fill data enters the telemetry stream as a string of zeros, while a zero intensity count enters the stream as a one. In two's complement arithmetic, messages will be negative numbers, fill will be zero, and valid intensity measurements will be non-negative; allowing the different data types to be rapidly and efficiently sorted during the reformatting operation.

The combination of the injected messages in the data stream and the information in the software status buffer, permit the ground software to completely identify each bit of data, including the experiment sequence that produced it, the implied data format, the dimensions of all matrices in the format, the location of each datum within its matrix, and the file name that will be assigned to the sequence in the final data archive. Furthermore, this information can be completely developed from a segment of telemetry as short as one major frame (about 8 seconds in the SMM system), making it very easy to evaluate data received during short real time passes or from partial orbits.

Additional discussion of the UVSP software system is given by *Rehse, et al.* (Journal of Spacecraft and Rockets, 1982, 19, 186). A complete listing of a recent version of the Jr

code is included in Appendix 3 of this report.

Command Generation System

The command generation system provides the software interface through which the scientist can design and execute an experiment in readily understandable terms. It frees one from the requirement to know the internal details of the instrument, a general knowledge of the basic instrument modes being sufficient for most purposes. The command generation system is composed of a two programs, an experiment generator (or compiler), and an experiment assembler. The experiment generator portion is known as Phase 1. This program allows the scientist or daily planner to design an experiment sequence and create readable text files that serve as input to the assembler, Phase 2, which prepares the actual memory load for the flight computer. Phase 1 was designed to provide the user with maximum convenience and flexibility. In the intermediate text file produced by Phase 1, only those parameters that are relevant to the type of experiment being created will appear in the readable parameter block text. Phase 1 also provides facilities for correcting inputs, and inspecting results before output.

The experiment assembler, Phase 2 takes the various experiments requested for the day in the form of intermediate output files from Phase 1 and creates a new memory image for the instrument control computer. Input to Phase 2 may include several Phase 1 files, making it easy to combine observing requirements of several different investigations into one computer load. The output from Phase 2 consists of three files and a listing. The first file contains the complete instrument computer memory image and is retained in the ground computer's storage. The second file contains the data required to create the instrument computer load. The third file associates vectored entry points with experiment descriptor filenames and is used to annotate the daily timeline print. The listing includes octal values for all memory locations loaded by the current command generation, as well as resolved listings of all command mode statements and breakdowns of each parameter

block into its component bit fields with a verbal description.

A complete discussion of the Command Generation System is given in the UVSP command generation handbook, which is included as Appendix 2 of this report.

Data Acquisition Package

The data acquisition package was developed by Dr. R. Shine of the GSFC staff in collaboration with Lockheed personnel. Although the preparation of this code was not a Lockheed responsibility, a brief overview of its operation is included here for the sake of completeness. The purpose of this code is to capture the incoming data stream as it arrives in the Experiment Operations Facility [EOF] at GSFC. The data flow from the SMM satellite into the EOF is illustrated schematically in Figure 8. In the early part of the mission, our primary contact with SMM was via the Satellite Tracking and Data Network (STDN) and the NASCOM communications system. Later in the mission when TDRSS became available, this system took over part of the STDN workload. In either case, data transmitted to GSFC over NASCOM arrived both in the EOF and at the Information Processing Division (IPD). At IPD, the data were recorded for later processing and error correction, and eventually resulted in the production of final data tapes. The data arriving at the EOF entered a PDP-11/34 computer through an electronic interface called the EOF Interface Unit (EOFIU). The EOFIU was developed at Lockheed for the SMM mission, and served both the UVSP and the XRP instruments. Additional information on the EOFIU is contained in Appendix 4 of this report. The data from the EOFIU entered the PDP-11 memory via a Direct Memory Access (DMA) channel, where it was captured and processed by the Data Acquisition System.

The operating philosophy of the Data Acquisition System was to allocate a very large block of storage in a disk memory system, map this block so that each word in the block corresponded to a particular word in the anticipated telemetry stream, and then to load each received datum into its predetermined storage location when it arrived. The disk

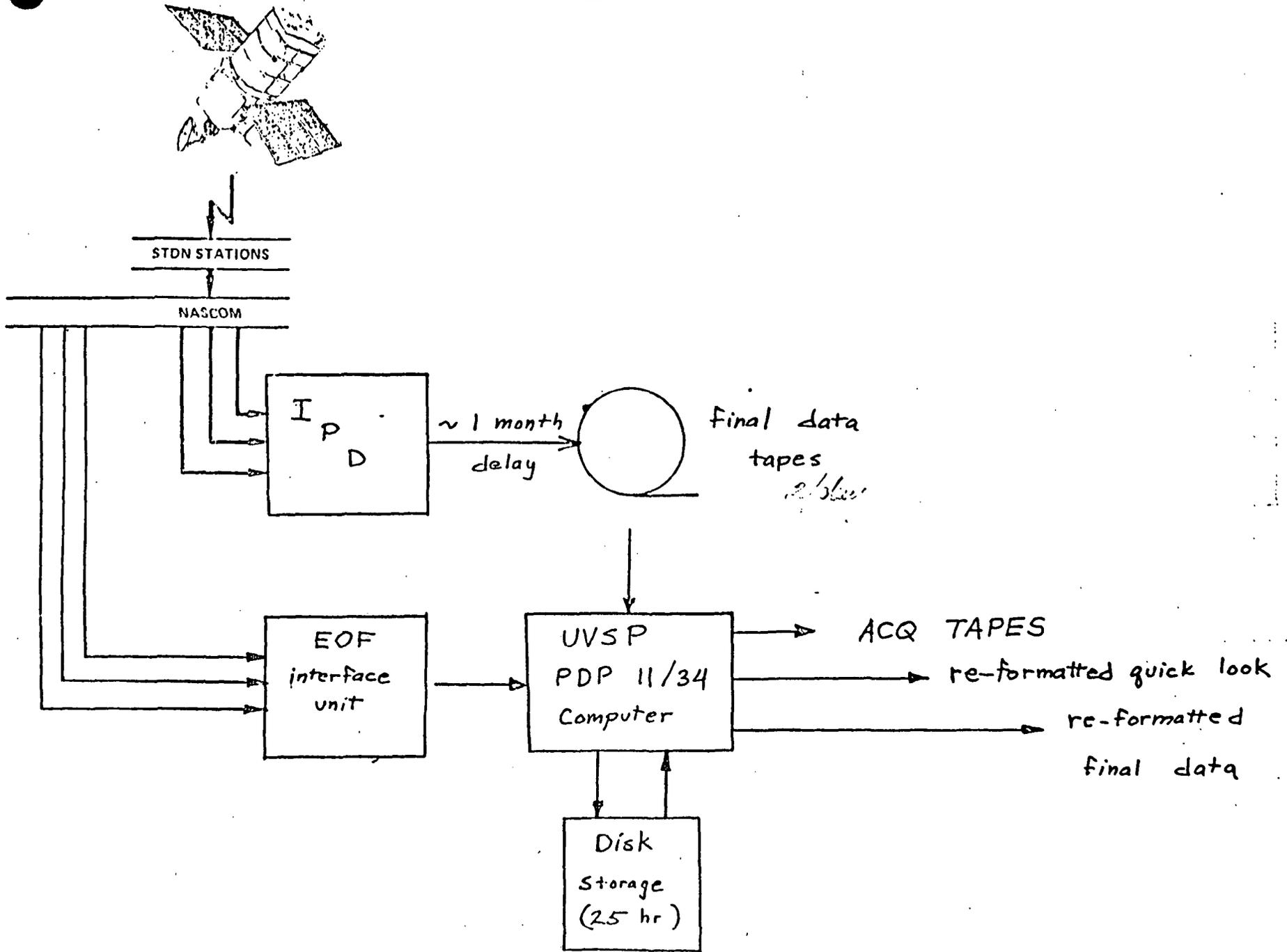


Figure 8. DATA ACQUISITION

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memory system selected for the task was large enough to accommodate roughly one day's data from the two instruments after extracting that portion of the entire telemetry stream that was pertinent. On a daily basis, the contents of the disk would be transferred to tape for storage in a quick-look archive, and the memory re-mapped for the next period of observing. Since the mapping for each period was predetermined, the order in which the telemetry from the different orbits arrived in the EOF was not critical. It was possible for the system to handle three data sources simultaneously by using multiple buffering to interleave operations on the data from different sources. This allowed us, for example, to simultaneously receive playback data from the spacecraft's tape recorder and real time data from the telemetry transmitter. Data in the disk were, of course, on line and available to the computer for processing by the data reformatting task.

As shown in Figure 9, the data entered the PDP-11 memory via a DEC DR-11B direct memory access interface. Data were initially captured in one of two "burst buffers" of 257 word capacity each. Each burst held one minor frame of data together with some overhead and ancillary data. Since not all of each minor frame was relevant to either the UVSP or the XRP instrument, an initial sorting was done at this stage to discard all unwanted words. This was done by extracting the desired words information from each burst buffer using a table-lookup algorithm. Two burst buffers were used so that one could be processed while the other was being loaded. The retained fraction of each minor frame was placed in the proper place in one of six major frame buffers in memory. These buffers were also arranged in pairs, so that loading and processing were asynchronous. Two major frame buffer pairs were allocated to playback data, and a third pair to real time data, so that a total of three simultaneous data sources could be captured. Spacecraft clock data, contained in each minor frame, determined the location of each minor frame in the major frame buffer and later, the location of each major frame in the "Today's Data" buffer in the disk memory. Major frames whose spacecraft clock data fell outside the boundaries of the "Today's Data" buffer were stored in an "Oddball file" to be handled separately. Once

the data were resident in the disk, they could be accessed by the reformatting program for conversion to the science file format for inspection and analysis. Additional information on the Data acquisition program is given in Appendix 5.

Data Reformatting Package

The reformatter software converts nearly raw data from the instrument and SMM spacecraft into a data format compatible with the data analysis language SOL. The functions of the reformatter are to strip out the fill data, block the data stream into logical experiments, identify the experiment in progress and determine the appropriate file format, intercept computer messages that identify the proper location of each datum in the format, load the intensity data into the format, create file header information, and write the results as a logical file on the disk. Ground reception of the data is sometimes noisy or occasionally drops out, so the reformatter allows for gaps in the data. The reformatter can also reconstruct the experiment parameters if the initial parameter information is missing. The initial version of the reformat code was an adaptation of the one developed at the University of Colorado for the OSO-8 instrument. It was prepared at Lockheed and delivered in a single detector version as discussed in the quarterly report for the period 1 January to 31 March 1980. Work on the extended version of the reformatter which could handle multiple detector experiments and accommodate a variety of data anomalies was suspended at the request of GSFC so that additional effort could be devoted to refinements in the flight software package and the consequential modifications required in the command generation system. The final version of the reformatter code was prepared by Dr. R. A. Shine of GSFC. A discussion of the final data file formats is given in Appendix 6.

Data Analysis Languages

The format of the files in the data archive followed the convention established for the OSO-8 spectrometer in order to make the data immediately and easily accessible to the

SOL data analysis language developed during the OSO-8 program. SOL, which stands for Spectrum Oriented Language, was written by D. M. Stern of the University of Colorado. It was a general purpose language that had several features that made it particularly convenient for use in the analysis of spectroscopic data. Procedures for opening, closing, reading, and writing data files were embedded within the language, and a graphics package operating a Tektronics 4010 terminal was included. The language handled vectors and arrays automatically in ordinary arithmetic operations so that loops over array indices did not have to be explicitly written. As part of this contract, Lockheed modified and delivered a version of SOL for use by the UVSP team. The modifications affected primarily the internal workings of the program and removed several unused sections that were relevant only to the original OSO-8 hardware configuration. The program remains functionally the same as the original, and is fully documented in the SOL language manual written at the University of Colorado.

Two other languages were also available for UVSP data analysis in the EOF. The first of these was IDL (Interactive Data Language), which was written by D.M. Stern after leaving the University of Colorado to form Research Systems, Inc. IDL used many of the ideas embodied in SOL, but added many extensions. Automatic handling of arrays was retained, and generalized to handle arrays with more than two dimensions. IDL also featured a greatly enhanced string handling ability, and the graphics package was improved. There are a number of detailed differences between the two languages, such as the range of array indices, which run from 1 to N in SOL but from 0 to N-1 in IDL. IDL did not have the built-in OSO-8 file reading procedures of SOL, although Stern provided a rudimentary read procedure for these files for our use. A disadvantage of Stern's file read procedure was that it gave no access to logical record header information. Lockheed wrote and delivered an improved procedure that retained the logical record header as part of this contract. We also provided a number of other utility procedures that were developed during the course of our scientific study of SMM data. These procedures are discussed in the appropriate

quarterly reports, and will not be treated here.

The final language that was prepared for UVSP is called ANA, and was developed by Dr. R. M. Shine of GSFC. It was designed to make the manipulation of UVSP image arrays particularly convenient, and features some powerful array manipulation commands. This language, though available to us, was not extensively used in the Lockheed data analysis program and will not be treated here. It is fully discussed in a manual prepared by Dr. Shine.

PART III

THE LOCKHEED SCIENTIFIC PROGRAM

Discovery and study of C IV Post Flare Loops

The discovery of post flare loops seen in C IV was one of the early results to which LMSC has made substantial contributions. The so-called 'Logo Raster' observation, carried out at the west limb on March 27, 1980, was planned by M. Bruner during an early period of residence at GSFC. The image, made during the rising part of the soft X-ray time history, shows a system of loops rising above the limb, clearly guided by the influence of the magnetic field. The observation was made in the dopplergram mode and shows that the northern legs of the loops are redshifted, while the southern ones were blueshifted.

The loops appear to have originated in NOAA region 2339, which was on the west limb at the time of the observation. Magnetograms are available from the Kitt Peak National Observatory for March 23rd, 25th, and 28th. There were two groups of spots seen in the Mt Wilson drawings. The leader spots (showing black polarity on the magnetograms) were approximately 10 degrees west and 5 degrees north of the trailer spots. On the basis of this, it appears that the most likely orientation of the loop system we observed was with the northern footpoints further from the Earth than the southern ones. If this interpretation is correct, then the observed doppler shifts correspond to downflowing material at transition zone temperatures in both legs of the loop system, rather than a syphon flow. The loops were transient in nature, as shown by a time series of smaller rasters made immediately after completion of the 'Logo Raster' observation. The lifetimes of individual loops in the system (as defined by their visibility in C IV) was of the order of a few minutes. This set of observations has been the subject of a detailed study by a team including M. Bruner, G. Poletto, R. Kopp, and G. Noci. A paper on the results of the study has been accepted for publication in Solar Physics.

Transition Zone Signature of Ephemeral Regions

An investigation arising from our participation in the FBS activity was a study of the growth of ephemeral regions and their signature in the transition zone lines. This study was coordinated by F. Tang, and concentrated on observations made on 11 Sept. 1980. The UVSP observations were made in the dopplergram mode in C IV. Magnetic field observations were made at Kitt Peak national observatory, Big Bear Solar Observatory, and the Mount Wilson Observatory. Ephemeral regions were identified in the magnetograms, which were then compared to the UVSP dopplergrams to search for cospatial signatures in the transition zone. Of the 31 bipolar ephemeral regions that were observed in the magnetograms, three were in the field of view covered by the UVSP. Study of the UVSP images showed two regions, co-spatial with the ephemeral regions, that both brightened and expanded in area during the period of observations. The results of the study were presented at the COSPAR meeting in Canada in the summer of 1982, and are published in the proceedings.

Density Enhancements of Flare Footpoints

An early investigation of flare observations on the disk concerned the 1980 April 8 flare. The observations of this flare were made in the UVSP density diagnostic line set, consisting of the Si IV, O IV, and S IV lines. Measurements were made in the RL mode, in which a series of rasters is made with the wavelength drive being advanced between rasters. Each raster represents a different position in the line profile, with the entire profile being covered in a series of five rasters. Data taken in this mode may be analyzed to determine the line intensities, widths, and positions (with respect to some global average) for each of the lines and for each pixel in the raster pattern. Electron densities may be estimated from the ratio of the Si IV and O IV lines. At the time of the impulsive phase, the 8 April flare showed a sudden brightening at the flare footpoint, accompanied by an increase in derived electron density. A preliminary presentation of the observations were made by Bruner et al. at the AAS Solar Division meeting at the University of Maryland. A more definitive

paper by Cheng, et al. appeared in the *Astrophysical Journal*.

The April 8 Flare - a Critical Review of the Experimental Results

The 1980 April 8 flare became the object of an extended investigation during the SMM workshop; one of five selected for study by the energetics team, of which M. Bruner was a member. Density diagnostics for this flare were available both from the UVSP results, and from concurrent P78-1 measurements, allowing us to derive the total thermal energy content of the flaring plasma and its evolution with time. This was the only data set available to us for which this was possible. By the time of the workshop, a considerable body of analysis of this event was in existence. M. Bruner prepared a critical review of the results, that was subsequently incorporated into the energetics chapter of the forthcoming monograph on the workshop. The complete text of the review was included in the quarterly report on this program for the period 1 April to 30 June 1984.

Energy Flux Transportable by Sound Waves

Another early investigation involved the study of N V dopplergram sequences in an attempt to estimate the energy flux transported across the transition zone by acoustic waves. This study was done in collaboration with Dr. G. Poletto of the Arcetri Astrophysical Observatory in Florence, Italy. The observations were made in the dopplergram mode in a series of 21×21 arc sec rasters. The results were generally consistent with earlier studies conducted by Bruner who had analyzed C IV and Si IV observations made with the UV spectrometer on OSO 8; finding that the inferred flux of energy that could be carried by the waves was inadequate by two or three orders of magnitude to explain the heating of the corona. A short contribution discussing the N V work has appeared in *Memoria della Societa Astronomica Italiana*. A more extended paper including a new theoretical treatment of wave propagation was prepared and submitted to *Solar Physics*. This paper met difficulties with a referee who raised several objections to the theoretical treatment, and is now awaiting revision.

Radiating Properties of Solar Plasmas

A more recent investigation that was partially inspired by the SMM workshop activities was a study of the radiating properties of solar plasmas. In this study, which was initiated by and carried out in collaboration with Dr. R.W.P. McWhirter of the Rutherford-Appleton Laboratory, we compared the total power radiated by an atmosphere with the power in a single spectral line. The calculations were based on a carefully selected set of atomic data and were carried out for a series of empirical emission measure distributions taken from the literature. The object of the study was to discover to what extent the intensity of a single line could be used as a diagnostic to estimate the total radiated power from an unknown atmosphere. Such an implied relationship is not unreasonable, since the general shapes of emission measure distributions tend to be very similar.

In a preliminary test, McWhirter found that for the several distributions tested, the total radiated power was directly proportional to the intensity of the C IV resonance lines at 1548 and 1550 Å, with an uncertainty of about 20 percent. We extended this study to incorporate a larger set spectral lines that are commonly observed by SMM, and added several more emission measure distributions to the empirical data base. The final data base included sample distributions for both quiet and active regions as well as for flares. The results of the extended study confirmed the existence of an apparent systematic relationship between the two quantities, but with a larger uncertainty. We confirmed the approximately linear relationship between total radiated power and the intensity of the C IV line, but found that a power law with an exponent of 1.1 (e.g. a linear relationship in the logarithms of the quantities) gave a slightly better fit to the data. The power law relation held for the C IV, N V, and O V lines observed by UVSP, though with different exponents. For the O VII and Ne IX lines observable by the XRP experiment, we found that the data were well represented by a quadratic relationship between the logarithms of the two quantities. These relationships are illustrated in Figure 10.

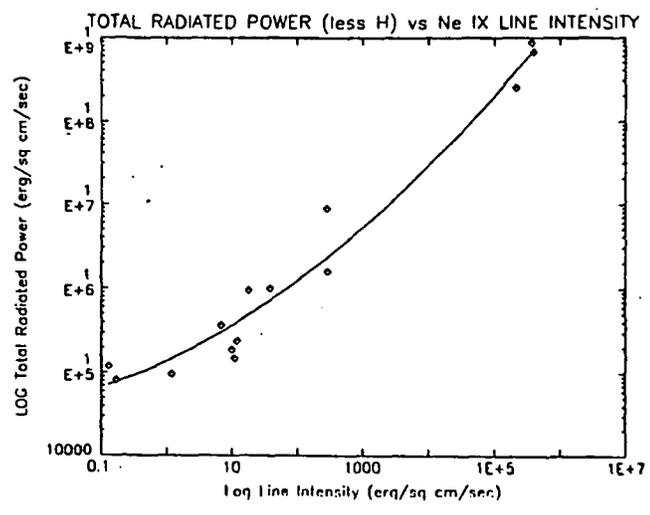
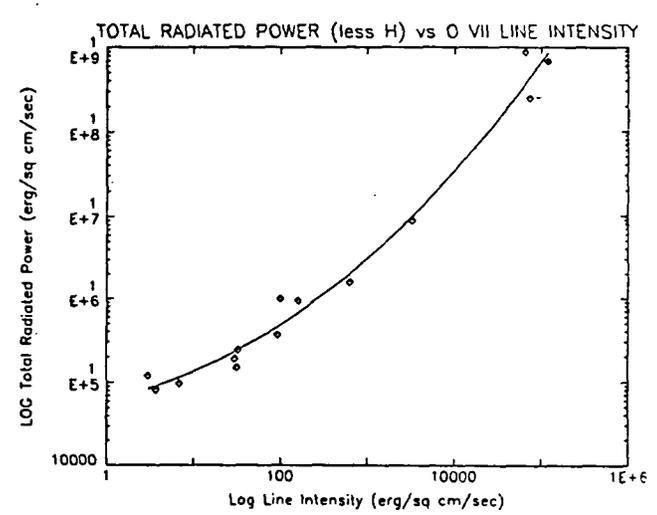
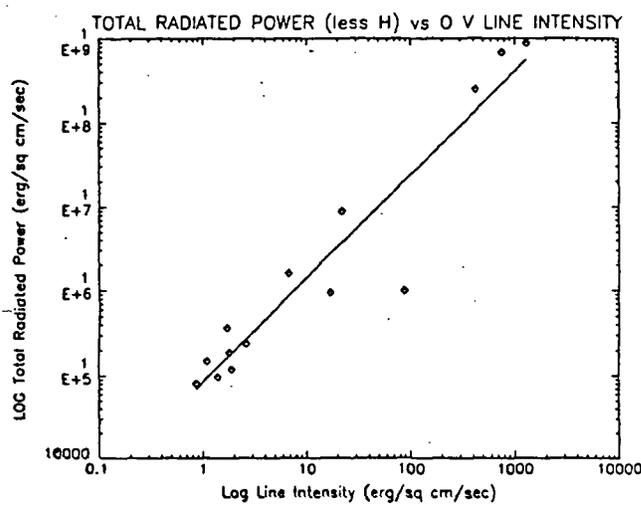
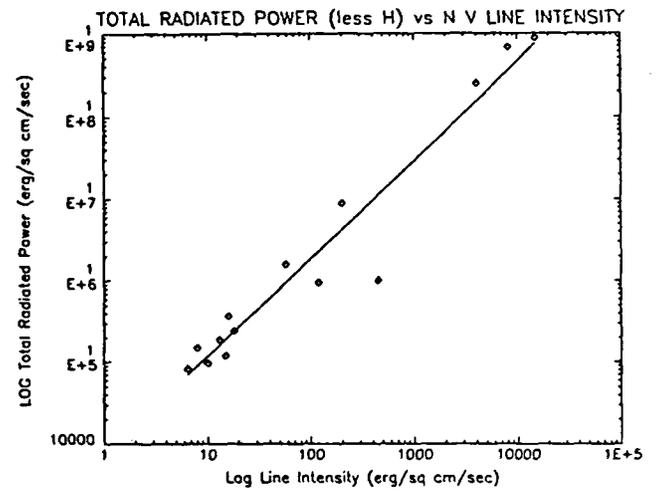
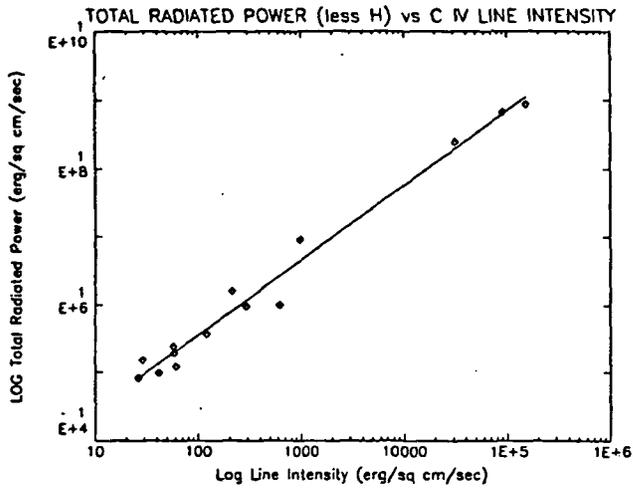


Figure 10

Another aspect of the radiated power study was the computation of effective values of the so-called G(T) functions for each spectral line considered. To illustrate the concept, we consider the conventional expression for the intensity of a spectral line in the effectively thin case. The intensity is given by

$$I = \frac{1}{4\pi} \int_0^{\infty} n_e^2 \frac{n(H)}{n_e} \frac{n(z)}{n(H)} \frac{n(g,z)}{n(z)} \chi(g,T) \frac{dh}{dT} dT$$

where Ne is the electron density, n(H)/Ne represents the ionization balance of hydrogen, n(z)/n(H) is the abundance of the element z with respect to hydrogen, n(g,z) is the fraction of the element z that is in ionization state g, T is the absolute temperature, and h is a unit of distance along the line of sight. This expression may be written as

$$I = \frac{1}{4\pi} \int_0^{\infty} n_e^2 G(T) \frac{dh}{dT} dT$$

where the abundance and the temperature dependent terms depending on the physics of the particular ion have been combined in the function G(T). We now define an average, or effective value of G through the expression:

$$I \equiv \frac{1}{4\pi} \frac{n(z)}{n(H)} \bar{G}(T_m) \int_{\frac{T_m}{\sqrt{2}}}^{\sqrt{2}T_m} n_e^2 \frac{dh}{dT} dT$$

where Tm is the median temperature below which exactly half of the intensity of the line arises. Note that this integral is carried out over a finite range in T, amounting to a factor of two between the lower limit and the upper limit, and with Tm being the geometric mean of the two limiting values. This is the convention used, for example, by Jordan (ref.). Combining the first and third expressions, we may compute G(Tm) as

$$\bar{G}(T_m) \equiv \frac{\frac{1}{4\pi} \frac{n(z)}{n(H)} \int_0^{\infty} n_e^2 \frac{n(H)}{n_e} \frac{n(g,z)}{n(z)} \chi(g,T) \frac{dh}{dT} dt}{\frac{1}{4\pi} \frac{n(z)}{n(H)} \int_{\frac{T_m}{\sqrt{2}}}^{\sqrt{2}T_m} n_e^2 \frac{dh}{dT} dT}$$

We see that $G(T_m)$ is a special kind of weighted average, where the emission measure is used as the weighting function. It is not the usual weighted average, because the normalization and averaging integrals are carried out over different ranges in T .

We computed values of $G(T_m)$ for each spectrum line and for each of the emission measure models considered for the radiated power study. For the transition zone lines, we found the $G(T_m)$ and T_m values to vary only slightly from one model to the next, suggesting that the mean values could be used to compute a very good first approximation to the emission measure distribution, given a set of line intensity measurements. In the case of the O VII and Ne IX lines, the values varied considerably between flaring and non-flaring models, being influenced by the slope of the high temperature part of the emission measure distribution. A summary of these results is given in Table 1. The entries marked in the tables with asterisks represent cases where the high temperature end of the emission measure model did not completely cover the range of formation of the ion in question.

The utility of the $G(T_m)$ averages is that they permit us to quickly estimate values for the emission measure at temperatures in the vicinity of T_m , with the assurance that the derived values will represent something better than a zeroth order approximation. A possible extension of this utility will be discussed in the next section. An oral paper covering some the results of the radiated power study was presented at the 1985 summer meeting sponsored by NSO at the Sacramento Peak Observatory. A definitive paper on the results is in preparation.

Absolute Wavelengths of Solar Lines

Another research topic that was recently addressed is the question of the absolute wavelengths of solar lines that have been observed with the UVSP. This observing program had as its objective, the measurement of the wavelengths of several chromospheric lines with respect to the geocoronal absorption line in O I, which is taken as a reference wavelength. The significance of the program is as follows: In the study of velocity fields

G(T) - Model Weighted Average

Model	C IV	N V	O V	O VII	Ne IX
1	1.04(-8)	7.9(-9)	1.22(-10)	1.8(-12) *	0.2(-12) *
2	1.08(-8)	8.5(-9)	1.16(-10)	1.9(-12) *	0.3(-12) *
3	0.97(-8)	7.3(-9)	1.14(-10)	2.6(-12)	2.8(-12)
4	1.07(-8)	7.5(-9)	1.14(-10)	2.7(-12)	2.8(-12)
5	1.16(-8)	5.3(-9)	1.04(-10)	2.9(-12)	2.3(-12)
6	1.05(-8)	7.5(-9)	1.21(-10)	1.5(-12) *	0.1(-12) *
7	1.00(-8)	7.6(-9)	1.22(-10)	1.6(-12) *	0.1(-12) *
8	1.10(-8)	9.1(-9)	1.08(-10)	2.7(-12)	0.2(-12) *
9	0.97(-8)	7.7(-9)	1.15(-10)	2.6(-12)	2.9(-12)
10	1.02(-8)	7.6(-9)	1.15(-10)	2.5(-12)	1.0(-12) *
11	1.16(-8)	7.2(-9)	1.14(-10)	1.6(-12)	1.2(-12) *
14	0.87(-8)	6.7(-9)	1.20(-10)	1.6(-12)	2.6(-12)
15	0.94(-8)	7.2(-9)	1.23(-10)	1.4(-12)	2.6(-12)
16	0.94(-8)	6.4(-9)	1.17(-10)	0.5(-12)	1.5(-12)
**	1.24(-8)	8.0(-9)	1.24(-10)	3.5(-12)	4.1(-12)

** Model independent values

on the sun, all past experiments have suffered from the fact that none of the available instruments have been equipped with on-board wavelength reference sources. Since velocity measurements are based on the measurement of doppler shifts, this has meant that there was effectively no rest frame to which velocity measurements could be referred. In order to study problems such as mass balance in the transition zone and corona, investigators have had to assume that some observable quantity such as the wavelength position averaged over a large field represented a reproducible working standard of wavelength, and that this wavelength represented material that was at rest with respect to the center of the sun. Although these assumptions are plausible, they lacked experimental confirmation. A systematic red or blueshift of the reference wavelengths would have been undetectable.

By measuring a set of chromospheric wavelengths with respect to a non-solar reference, the question of possible systematic motions or wavelength shifts originating in the solar atmosphere is avoided. The geocoronal absorption lines in the O I triplet near 1302 Å provide such a reference. The O I line profile is very similar to that of the much broader H-Lyman alpha line, and shows two quite distinct regions of line reversal in the vicinity of the core. The broad, shallow core is caused by non-LTE radiative transfer effects in the solar chromosphere, which is optically thick at these wavelengths. The narrow central part of the core, however, is an absorption line formed in the Earth's upper atmosphere, which is at a much lower temperature. The geocorona is substantially at rest with respect to the center of mass of the Earth, affected at most by the effects of diffusion related to the gradual escape of atoms in the high energy tail of the outer layers of the oxygen geocorona. The physics of the escape process in the geocorona is well understood so that this effect can be evaluated with confidence. Similarly, the radial velocity of the Earth with respect to the sun is a function of orbital mechanics, and can be accurately computed. Thus, the O I line can serve very well as a standard of absolute wavelength for solar UV observations.

In applying the method, the UVSP instrument was used to carefully measure the

positions of several UV emission lines formed in the solar chromosphere with respect to absorption cores in the resonance emission triplet of atomic oxygen. The selected lines were close in wavelength to the O I triplet in order to minimize the required motion of the wavelength drive and consequently, the uncertainty introduced by any non-linearities in the drive performance. Steps in the analysis included the determination of the observed line positions in step numbers on the wavelength drive, conversion these position numbers to apparent wavelengths, correction of the apparent wavelengths for systematic effects (principally the orbital motion of the spacecraft) and finally, computing the corrected wavelengths of the solar lines from the observed offsets from the geocoronal O I absorption lines.

Computation of the line-of-sight component of the spacecraft velocity vector was based on a complete solution of the spherical triangle defined by the position vectors from the center of the Earth to the sun, the spacecraft velocity vector, and the geocentric pole. Input data to the computation were the time of the observation; the times of spacecraft sunset, sunrise, and ascending node passage taken from the orbit predictions on the SMM planning charts; and the right ascension and declination of the Sun from the American Ephemeris and Nautical Almanac.

The analysis showed very good internal consistency among the several measured positions of the O I lines at 1302.169 and 1304.858, based on the pre-launch values for the polynomial coefficients in the wavelength drive position prediction formula. The results of the wavelength measurements of the solar lines were very surprising. Both the 1300.91 line of S I and the 1318.998 line attributed to N I were found to be blue shifted with respect to their rest positions. The observed blueshifts corresponded to upflow velocities of about 3 km/sec, and the shift exceeded 3σ . In our first observing run, we had also observed the 1318.998 line which was classified as arising from N I in the NRL atlas of L. Cohen (NASA publication 1069, 1981). This line showed a considerable departure from

its expected position based on the pre-launch calibration of the UVSP wavelength drive. If this departure is attributed to doppler shift due to motion in the sun's atmosphere, an upward velocity of about 8 km/sec is implied. In the second run, the C I line at 1311.404 was observed in lieu of the 1318.998 line since it is closer to the nearest reference line. To our surprise, we found this line to the red of its rest position by about 8 km/sec. The line profiles are well developed in all cases, and display good signal to noise ratios, so that the displacements cannot be attributed to statistical errors. There is a possibility that the identification of the 1318.998 is in error, since the 1319.67 line which arises from the same multiplet is not observed in any of the UVSP spectra. The 1319.67 line is expected to be nearly twice as bright as the 1318.998 line (Kelly and Palumbo - NRL report 7599).

We have considered a number of possibilities apart from a systematic velocity in the chromosphere that could be advanced to explain the observations. The effect of solar rotation, which can be as high as 1.9 km/sec, is not a problem for these observations, since they were carried out at sun center. The radial velocity of the Earth was computed from ephemeris data for the day of the measurement to be about 0.19 km/sec, which is a decade too low to explain the observations. The effect of the Earth's motion about the Earth-Moon barycenter is even smaller; about 12 m/sec. There is a possibility that one or both of the lines have been mis-identified. The line at 1300.91 angstroms is not listed in the Kelly and Palumbo table, but has been classified by *Tondello* (1972, Ap. J. 172, 771) as arising from S I. It was identified in the solar spectrum by *Chipman and Bruner* (1975, Ap. J. 200, 765), who also reported most of the other nearby S I transitions. The other S I transitions are also seen in the UVSP spectrum. Thus this identification seems fairly secure. The 1318.917 line is classified in the NRL ATLAS (L. Cohen, 1981, NASA Publication 1089) as arising from N I. Kelly and Palumbo list a N I doublet whose fainter component lies close to our observed wavelength. The other component, however, has not been observed either in the Chipman and Bruner spectrum or in the UVSP spectrum. Thus this identification is suspicious and may be wrong.

These results were presented during the 1985 annual meeting of the Solar Physics Division of the American Astronomical Society. An abstract of the paper has been published in *Bull. Am. Astron. Soc.* Vol 17, 630, (1985).

Comparison of Photospheric Electric Currents and Ultraviolet and X-ray Emission in a Solar Active Region

Recently it has become possible to infer the presence of electric currents in the solar photosphere using vector magnetograph measurements. An important question that can now be addressed is whether heating of the upper solar atmosphere takes place via electric current dissipation. This can be studied empirically by comparing regions of inferred J_z (vertical component of the photospheric electric current density) with areas of enhanced emission in the chromospheric, transition region and coronal structure. Recently *deLoach et al.* (1984) used MSFC vector magnetograms and UVSP raster maps in Lyman alpha and N V to investigate spatial correlations of J_z and enhanced emission within an active region. A marginal correlation was found.

As summarized in a paper to appear in the *Astrophysical Journal* (1 January 1986; "A Comparison of Photospheric Electric Current and Ultraviolet and X-ray Emission in a Solar Active Region" by Haisch, Bruner, Hagyard and Bonnet) we have completed a more comprehensive intercomparisons of vector magnetograph, UVSP, XRP and high-resolution UV rocket images and filtergrams to search for evidence of heating by current dissipation. Specifically, we used UVSP spectroheliograms in C IV, Si IV and O IV. Empirical correlations between J_z and bright emission regions in Lyman-alpha and in the 1600 A UV continuum (rocket data) were found. There appeared to be a lesser degree of correlation between J_z and the UVSP transition region emission. However none of these correlations were consistent with expected scaling relations between simple ohmic heating and radiative losses. The present status of this approach for empirically investigating the nature of the heating mechanism of the structures in the upper solar atmosphere is that

there are suggestive correlations involving electric currents, but further correlative studies are necessary.

Directions for Future Investigations

In this section, we discuss some research topics that have been identified as logical extensions to the investigations performed under the present contract. Some of these topics are logical extensions of work that we have already done or that is in progress. Others have been identified in the past, but postponed in favor of the work discussed above, while still others are new. We anticipate that additional topics will present themselves as the study of the existing data base continues.

Flare Filling Factors

This project has been treated in several of the progress reports on this contract. It is an outgrowth of studies done for the SMM Flare Workshop, specifically with respect to the April 8 flare. A striking result of the compilation of observations of this flare was the comparison of estimates of the flaring volume as functions of time using different methods. In one method, based on atomic physics computations, line ratios are used to estimate the electron density. These densities are combined with values of the volume emission measure determined from line intensities to determine the effective emitting volume. A second method uses an analysis of the HXRBS data combined with radio observations to determine an effective area for the optically thick radio emitting region, which, in turn is used to estimate the volume. A third method rests on the apparent area observed with one of the imaging instruments such as UVSP, XRP, HXIS, or P78-1. This area is again used to infer a volume.

In the case of the April 8 flare, we found agreement between the two volume estimates based on area measurements, but a large discrepancy between these values and the effective volume estimated from the density / emission measure analysis. This result is in accordance

with previous findings by others, who attributed the differences to the incomplete filling of the emitting volume with plasma. The new result from the April 8 study was that as the flare developed, the volumes based on atomic physics estimates approached those estimated from areas until they were in substantial agreement at the end of the gradual phase. The result is illustrated in Fig. 11, which was prepared for Chapter 5 of the SMM Flare Workshop monograph. It is seen that the volume estimates based on areas rise during the impulsive phase of the flare, and then gradually decrease with time. The volumes derived from the spectroscopic diagnostics, however, show a large (though uncertain) initial decrease, followed by a gradual rise. This seems to imply a time evolution of the filling factor, which would be an important result, if confirmed. This idea could be followed by examining both the SMM and the P78-1 data bases for other flares where this type of comparison can be made. Our preliminary checks have revealed a number of candidate events that could be examined as an extension to the present study.

Radiated Power Study

The basis of this study was discussed at some length earlier in this report. There are two directions in which the study could be extended. The first of these is to broaden the empirical data base by identifying and adding more examples of emission measure distributions derived from observations, and to incorporate more of the ions for which we have good atomic data. The emphasis in this extension should be to add more examples of flaring plasmas to the set of emission measure distributions, and to include more of the lines that are typically used by the SMM instruments, particularly UVSP and XRP.

The second extension emphasizes the effective values for the $G(T)$ functions, and their utility in computing emission measure distributions. As discussed previously, the quantity $G(T_m)$ may be used to derive a good estimate of the emission measure at temperatures in the vicinity of T_m . If we were to do this for several lines spanning the desired temperature range, the result would be a first order emission measure distribution. The method is

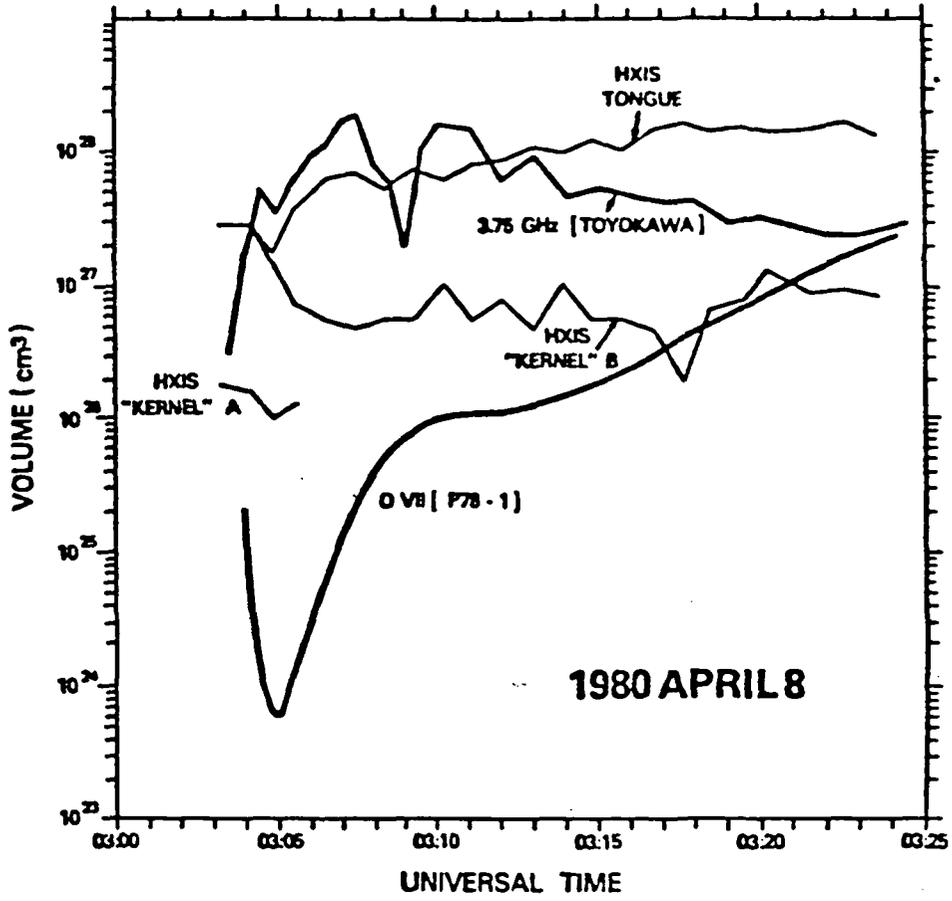


Figure 11

similar to the original method of *Pottasch* (Space Science Rev. 3, 816, 1964; Bull Astron. Inst. Neth., 19, 113, 1967), but with atomic data that are more realistically weighted. If these derived emission measure values are now connected by some reasonable technique such as cubic spline interpolation, we may use the methods developed in this study to recompute the $G(T_m)$ functions for this particular emission measure distribution. The new $G(T_m)$ functions would, in turn, be used to compute a second approximation to the emission measure distribution, and the iterative process continued until convergence is obtained. Since, as we have already shown, the $G(T_m)$ values are insensitive to the shape of the emission measure distribution, we may expect convergence to come very quickly, probably within one or two iterations.

Preflare Oscillation Study

This project is based on a suggestion by E. Antonucci that it might be possible to observe oscillatory behavior in the transition zone lines during the last few minutes before onset of the impulsive phase of a flare. We have found some observations that are suggestive of oscillations in the TRANSVEL and TRANSMAP observations that were made in the N V line during the early part of the mission in 1980. These data sets should be studied more carefully, subjecting them to power spectrum analysis to discover the extent to which they display quasi-periodic behavior. There appears to be an adequate data base in the existing UVSP archives, so that additional observations will probably not be needed.

Chromospheric Depression Study

The process of chromospheric evaporation or ablation is, by now, a widely accepted idea. Observations made with the XRP instrument, particularly the bent crystal spectrometer, have revealed the blue shifted material that would be expected on the basis of the model. The question to be addressed here is the fate of the region from which the material is ablated during and immediately after the impulsive phase. Since chromospheric

material is removed from a relatively restricted area. we may logically expect to find a depression, or region of low density, in the vicinity of the footpoints of the flare. Presumably, the higher temperature in this region would provide the pressure necessary to prevent the depression from being filled by material flowing in laterally from the surroundings for as long as the strong chromospheric heating persists. Such a depression is expected whether the heating mechanism is thermal conduction, as suggested by Hyder during the SMM workshops, or by non thermal electrons, as discussed by Woodgate during the 1985 NSO summer workshop at the Sacramento Peak Observatory.

It may be possible to find evidence for chromospheric depressions by examining the maximum transition zone densities seen at flare footpoints as a function of the position of the flare on the disk. What we are seeking is a simple geometric effect. If the footpoint is near disk center, then we expect to see all the way to the bottom of the depression (assuming that the hot ejecta are transparent to the transition zone radiation), while for a footpoint near the limb, the bottom may be obscured by the intervening wall. Since the density is expected to be highest at the bottom of a depression, we may expect to find that flares observed near the disk center show systematically higher maximum densities than those observed near the limb. This idea could be tested by surveying the UVSP data base for flares and sub flares observed with the O IV - Si IV density diagnostic line pair. A correlation plot of maximum observed density as a function of distance from disk center should reveal the effect if it is present, provided that a sufficiently large set of samples can be found.

Prominence and Filament Studies

The object of this investigation would be to study the formation of and evolution filaments by examining them in as many temperature regions as possible. The question that would be addressed is their mechanism of formation. Some schools of thought contend that prominences are formed by cooling and condensation (recombination) of hot coronal

material, while others postulate a direct formation from cooler material coming from the chromosphere. Observations of the higher temperature regions should allow this question to be resolved in a straightforward way. Much of the effort would be focussed on ground based observations. A systematic survey of the existing data set could be made to search for examples of UVSP observations that are cospatial with filaments that have been observed from the ground. Of particular interest is the period of time from May, 1984 through September, 1984 when the wavelength drive was inoperative. Subsequent work by Bruner and later by Henze showed that the spectrometer was tuned to the C II lines during this period. C II is interesting for the study of filaments and prominences, as it is formed at a temperature of about 30000 deg K; only slightly higher than the 10000 deg typical of prominences. It will also be interesting to conduct a similar search for signatures of prominences in the C IV lines.

LMSC/F067851

APPENDIX 1
INDEX OF QUARTERLY REPORTS
Prepared Under Contract NAS5-24119

INDEX OF QUARTERLY REPORTS
Prepared Under Contract NAS5-24119

The purpose of this index is to identify the topics discussed in the various quarterly reports prepared during the course of this contract. It is intended to assist the interested reader in locating additional information pertinent to topics discussed in the final report. It also serves as a convenient short-form history of the work performed under the contract.

1977 Quarter 1

- o Completion of first phase of work under contract NAS5-23691
- o Installation and modification of S/C simulator software in PDP-11/34
- o Hardware interface definition for PDP-11 to SCI interface drawer

1977 Quarter 2

- o Integration of S/C simulator interface drawer and PDP-11
- o Wavelength drive performance test completed
- o JR Cross-assembler written and tested
- o Instrument test procedures for use in JR defined

1977 Quarter 3

- o JR software architecture defined, including new command mode
- o Command generator compiler input specifications defined
- o Science meetings at Culham Lab and UCL attended
- o Design review of Electronic system at SCI systems
- o OSO-8 hardware failures found

1977 Quarter 4

- o Mission flight software work (Revision A) completed
- o Parameter blocks defined
- o Field support of test software
- o Phase 2 command generator coding
- o Investigator's Working Group (IWG) meeting in Sunnyvale
- o JOS Working Group formed
- o Data acquisition codes defined
- o R. A. Shine detailed to LMSC from GSFC
- o PDP-11/34 sysgen (V 3)
- o PDP-11/34 Tape drive specifications defined
- o EOF computer load analysis

1978 Quarter 1

- o Huntsville test support
- o Electronics breadboard / SC simulator software development tool defined
- o Mission software installed and tested
- o Phase 2 command generator completed
- o Phase 1 command generator conceptual design complete, and coding started
- o Preliminary version of data acquisition code completed by R. A. Shine
- o UVSP telescope alignment at GE.

1978 Quarter 2

- o JR flight software package installed and tested at G.E.
- o Phase 1 command generator coding
- o IWG meeting in Huntsville, Ala.

1978 Quarter 3

- o Analysis of "missed interrupt" problem, development of redundant timer operation as work-around
- o Completion of Phase 1 Parameter Block Generator code
- o IWG meeting at Culham Laboratory
- o SSO.007 Coronal Bright Point program defined

1978 Quarter 4

- o Continued JR software checkout at GSFC
- o Instrument calibration at GSFC
- o Phase 1 command mode section complete and integrated with Phase 1 parameter block generator
- o IWG meeting at GSFC
- o Development of coronal heating SSO

1979 Quarter 1

- o Enhanced baseline JR software package completed and tested on software development tool
- o Diagnostic work on JR hardware
- o Performance Evaluation of completed instrument
- o Definition and execution of modified Hartmann test for setting Ebert mirror focus
- o Telescope focus and resolution tests
- o IWG meeting in Boulder, Colo. (High Altitude Observatory)

1979 Quarter 2

- o Reformatter defined and work started
- o Flight software modified to add three level priority interrupt
- o Baseline revisions to SOL defined
- o JWG, IWG meeting in Durham, New Hampshire

1979 Quarter 3

- o Reformatter development continued
- o New flight software package delivered with two level flare priority interrupt response
- o Flare test series package delivered
- o Baseline SOL conversion completed and tested
- o Command generation Phase 1, Phase 2 package completed
- o User's manual for command generation in preparation
- o JWG meeting at GSFC
- o FBS meeting in Montreal, Ca.

1979 Quarter 4

- o Updated flight software package delivered and installed

- o Further flight software package enhancements defined, coded, and tested on software test tool
- o Reformatter work deferred in favor of work on JR software at the direction of GSFC program scientist
- o Command generator package modified to reflect JR software changes
- o Final performance evaluation of completed UVSP at GSFC
- o JWG meeting at GSFC
- o IWG meeting at Huntsville, ALa.

1980 Quarter 1

- o Launch Support
- o Post launch checkout of UVSP
- o Delivery of 1 detector version of reformat program
- o Jr. flight software package in-orbit checkout
- o Command generator updated to reflect new performance information
- o Baseline version of SOL delivered and installed
- o Review of in-orbit testing of spectrometer
- o Measures implemented for preventing photometric sensitivity loss are detailed in this report
- o Sensitivity decrease diagnosed and control procedures defined. Loss mechanism is discussed
- o Definition of "workhorse mode" experiment concept and initial examples
- o Preparation of utility codes for quick-look analysis; delivery of documented codes
- o Observation of 27 March post-flare loops ("Logo Raster")

1980 Quarter 2,3

- o Post launch mission operations support
- o Major upgrade of flight software package
- o Evaluation of flare data in C IV, N V dopplergram mode
- o Presentation of 8 April flare density measurement at AAS meeting
- o Development of conversion program to allow image display on LMSC HP-1000 / Ramtek system
- o Big Bear Solar Observatory meeting

1980 Quarter 4

- o In-residence work in EOF by Bruner, Schoolman
- o Initiation of N V sound wave study with Poletto at Arcetri Obs.
- o Paper on SMM control system presented to AIAA
- o Development and installation of Command Generator enhancements
- o Preparation and execution of spicule observing program
- o Contribution to HXIS study of Hard X-ray imaging of post flare radio burst

1981 Quarter 1

- o Analysis of UVSP data supporting NASA sounding rocket 27.036
- o Continuation of N V sound wave study
- o Flare Buildup Study (FBS) meeting at GSFC
- o Bright point study initiated under FBS; one region identified

1981 Quarter 2

- o Bright point study continues - codes developed to mask images and develop light curves
- o Development of blinking color table for identification of image elements
- o Continued analysis of rocket support data, velocity computation, normalization for absolute intensities
- o N V preflare study begun with E. Antonucci
- o Continuation of N V sound flux project - modification of analysis codes to correct problems

1981 Quarter 3

- o Rocket support data analysis continued by L.W. Acton
- o G. Poletto visit to Palo Alto, discussion of theoretical results
- o Beginning of Noci, Antiochos loop model project

1981 Quarter 4

- o Calibration of C IV dopplergrams for Poletto / Noci loop study
- o Analysis of March 27 post flare loop system
- o Modifications to N V sound flux codes completed
- o Preparation for Jan, 1982 AAS paper
- o Development of velocity transfer function analysis

1982 Quarter 1

- o Presentation of N V work at AAS meeting in Boulder, Co
- o Bright Point study continues with determination of background levels
- o Bright Point project with M. Kundu defined.
- o Test of Hyder Vortex model of flares
- o Development of 48 level pseudo-grey scale for Ramtek

1982 Quarter 2

- o Comprehensive review of loop models completed by B. Haisch
- o Analysis of 27 March loop observation continues, R. Kopp joins analysis team
- o A. Walker and students begin a new loop study
- o Fe XXI limb scan survey initiated

1982 Quarter 3

- o Walker et al. study continues
- o Kopp and Poletto visit re: 27 March loop analysis
- o Loop lifetimes determined to be 15 - 30 minutes in C IV
- o Antiochos suggests formation is due to cutoff of heating to a pre-existing loop so that C IV loop is result of a cooling process

1982 Quarter 4

- o Problem discovered with velocity computation algorithm in N V program. Results are re-computed
- o N V paper in final preparation
- o Walker et al. work continues, finding a number of Fe XXI loops
- o Paper on "Transport and containment of plasma, particles and energy"

within flares" presented in Japan and accepted for publication in workshop proceedings

1983 Quarter 1

- o Wavelength drive reference method developed for analysis of 27 March flare loop system. Time development of velocity fields determined
- o Shell model of post flare loop system developed and applied to 13 July, 1982 flare
- o Paper on 13 July flare presented to AAS Solar Phys. Div. meeting in Pasadena, Ca.
- o SMM workshop begins, M. Bruner joins energetics group

1983 Quarter 2

- o Sept, 1980 active region study continues
- o Initial study of limb flares showing ejecta begins at Culham Lab
- o Initiation of radiated power study (RADPWR) with RWP McWhirter
- o SMM workshop at GSFC. M. Bruner accepts responsibility to prepare complete presentation of April 8 flare
- o UVSP data for Team E (Energetics) analyzed and presented to team members
- o Codes to analyze limb flares prepared and checked out on Rutherford "Starlink" computer (IDL procedures)
- o SOL version of radiated power code written and checked

1983 Quarter 3

- o RADPWR project continued in Palo Alto - effective collision rate concept developed for O V line at 1371.2 Angstroms, also applied to Fe XXI line
- o Effective G(T) values computed for major UVSP lines and used to derive emission measure conversion constants for UVSP observations
- o April 8 study continues with collection of available observations and published results

1983 Quarter 4

- o N V sound wave study continues - Paper returned by critical referee
- o N V flux computation procedures reviewed; small discrepancies corrected, and results re-computed - no substantial change in results
- o 23 Sept Active Region Study continues. Current density maps received from MSFC to be compared with UVSP data, rocket filtergraph data

1984 Quarter 1

- o April 8 critical review of all observations completed and presented to Team E at SMM workshop
- o Critical discussion of UVSP data from Team E flares completed and submitted to Team E leader. Complete text is included in this quarterly report
- o Magnetic field plotting capability developed to display MSFC magnetic field models on Lockheed HP-1000 system

1984 Quarter 2

- o Critical discussion of April 8 data set completed and submitted to Team E leader. Full text is included in this quarterly report

1984 Quarter 3

- o UVSP wavelength drive problem diagnosed and corrected - a discussion of the hardware, its problem, and the analysis of the problem is given in this quarterly report
- o Post-recovery data analyzed to show that the UVSP had been observing the C II lines at 1334.5 and 1335.7 Angstroms during the time when the wavelength drive was inoperative
- o Radiated power study continues with expansion of the atomic physics data base
- o 23 Sept active region study continues, concentrating on comparison of inferred electric current and images in H-Lyman alpha, 1600 A continuum, and C IV. Results do not support a current heating hypothesis

1984 Quarter 4

- o No work performed in October due to a gap in funding
- o WZERO program to determine absolute wavelength reference for UVSP defined by M. Bruner and run at GSFC
- o Bruner and Crannel initiate project resulting from 8 April study.

1985 Quarter 1

- o 23 Sept active region study completed. Paper submitted to Ap J for publication. Preprint of paper contained in this quarterly report
- o Data analysis methods developed for WZERO data. Wavelength drive system shown to be remarkably accurate
- o Bruner / Crannel study continues with identification of Feb 26 event for which both SMM and P78-1 data are available
- o IDL utility procedures developed for analysis of WZERO experiment will be widely applicable to UVSP data analysis. Procedures and documentation submitted to NASA in this quarterly report

1985 Quarter 2

- o Second observing run of WZERO experiment is analyzed
- o WZERO paper presented to AAS meeting in Tucson, Ariz.
- o Major wavelength drive anomaly analyzed - test procedures defined and tested. WLD problem shown to be apparently due to lubrication failure between WLD screw and follower nut. Recovery procedures defined
- o Corrected IDL procedures for computing line of sight velocity of S/C from planning sheets completed and included in this quarterly report
- o RADPWR work continues - FORTRAN version of the code is prepared

1985 Quarter 3

- o Results of RADPWR study presented to 1985 National Solar Observatory conference at the Sacramento Peak Observatory. Methodology and results are given in this quarterly report.

- o Wavelength drive tests show that the WLD motor is now free to run, but the WLD does not move. Failure determined to be most probably in the flexible coupling between the gear box and the WLD screw. In-orbit recovery from this failure is not possible, and the instrument will need to be returned to the laboratory for repair.

1985 Quarter 4

- o SMM observing program defined to support launch of NASA sounding rocket 27.090. Successful flight develops new data base for active region studies
- o Work initiated on contract final report

APPENDIX 2
UVSP COMMAND GENERATION
Updated 21 JAN 80

UVSP COMMAND GENERATION

Updated 21-JAN-80

Address inquiries to Steve Schoolman at Lockheed Palo Alto
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NOTICE

By popular demand, the formats for specifying the motions of the four hardware mechanisms of UVSP which are controlled within an experiment (X and Y rasters, polarimeter, wavelength drive) have been changed. Instead of indicating the number of STEPS which the mechanism will take, the user now specifies the number of POSITIONS it will occupy. Thus, a 3x5 raster is now created with the numbers 3 and 5 instead of 2 and 4, as was the case with the version of Phase-1 delivered in October, 1979. Note that previously created Experiment Definition Files will not be accepted by the new version of Phase-2.

CHAPTER 1

INTRODUCTION

Command Generation is the process by which the daily observing program is loaded into the UVSP's onboard computer, named JR. Command Generation has been divided into two parts, called Phase-1 and Phase-2. In Phase-1, the user is led through the procedures required to create Experiment Definition files. These are text files which contain all of the instructions which permit the UVSP to carry out scientific observing programs. In Phase-2, a number of Experiment Definition files are compiled into a JR memory load to be uplinked to the spacecraft. This memory load will control the operations of the UVSP instrument during a day's observations.

Each of these daily memory loads is called an "observing list". The control area in JR's memory is divided logically into three areas, called A-list, B-list, and C-list. The basic operations philosophy is that A-list and B-list will be used on alternate days, so that each of them can be re-loaded on the day during which the other is active. C-list will contain experiments which will remain resident for an extended period of time, either because they are used repeatedly or because they are held in reserve for special occasions like super flares. Because of a quirk in the software, a new C-list JR load can only be uplinked on a day when A-list is active.

Although Phase-1 and Phase-2 are both parts of Command Generation, they are obviously very different processes. Phase-1 gives the user the opportunity to use a considerable amount of imagination and flexibility in creating experiments. Phase-2, on the other hand, creates an actual memory load for JR, so it must do extensive error checking and will reject any input which is not perfect. Phase-1 and Phase-2 of Command Generation will typically be done at different times, and perhaps by different people. Any knowledgeable user can use Phase-1 to create Experiment Definition files at any time he/she finds convenient. These files are simply stored on a disk for inclusion in some future JR load. Phase-2, on the other hand, will generally be once per day in the late afternoon, following the daily planning meeting, to prepare the JR load which is to uplinked before the beginning of the next observing day.

Chapter 2 of this manual describes in detail the uses of the two types of text which go into the Experiment Definition files, namely Command Mode text and Experiment Parameter Blocks. Chapter 3 describes how the Phase-1 processor is used to create the Experiment Definition files. Chapter 4 describes how the Phase-2 processor is used to compile a number of Experiment Definition files into a single JR load.

CHAPTER 2

EXPERIMENT DEFINITION FILES

An Experiment Definition File is the output of Phase-1 Command Generation and the input to Phase-2. The file is a fully readable ASCII file which can be printed on a terminal or printer and can be modified with any of the RSX editors.

Each Experiment Definition File has two sections. The first section contains the Command Mode text, while the second contains the Experiment Parameter Blocks. Command Mode is a simple language with which the flow of scientific operations is controlled. Experiment Parameter Blocks contain the parameters which control the actual data-taking operations of the UVSP. These two types of text are described in detail in the following sections.

There are two basic rules governing any Experiment Definition File which is input into the Phase-2 processor. The first is that all Command Mode text must precede all Experiment Definition Blocks; the sections are separated by a line containing the symbol ".PBLK". The second is that the corresponding Experiment Parameter Blocks must exist within the file for all experiments referenced by the EXECUTE command, even if the experiment has been declared global. Any file created by the Phase-1 processor will of course meet these requirements. However, since the files are ordinary ASCII text files, the user cannot be prevented from generating them with an editor, or altering those created by Phase-1. Such a procedure may at times be quite useful, but these restrictions as well as the syntax rules of the two sections must be kept in mind if this is to be done successfully.

There is no requirement that an Experiment Definition File contain any Command Mode text. While it makes no sense to input a file having only Parameter Blocks into Phase-2, there is a good reason for creating such files with Phase-1. When the dialog through which a Parameter Block is created begins, the user is first asked whether this will be a new experiment. If the answer is NO, he/she is then asked for a file name. The program will search the named file to find a Parameter Block having the same label (symbolic name) as the one about to be created. If such a Block is found, it is simply copied by Phase-1 and the need for the dialog is eliminated. Thus, if there are experiments which will be run from many Command Mode sequences, the user may wish

PERIMENT DEFINITION FILES

to create an appropriately named file containing that Parameter Block and simply reference the file whenever the Block is needed thereafter.

COMMAND MODE

"Command Mode" is a pseudo assembly language which allows the user to control the flow of an orbit's operation, do simple arithmetic, test results, and make real-time decisions on how to use the UVSP instrument based on the results of the previous experiment and the state of the Sun.

The Command Mode instructions may reference a 32 word "user buffer" which contains status information as well as scratch memory. Some of these words are "read-only"; the user can read the contents of the word but cannot modify it. Others are "read-write" and can be altered as desired. Each word in the buffer has a symbolic name by which it is referenced. The buffer is defined as follows:

Read-Write Memory

XRASTR - X-raster coordinate

The X-raster position within the UVSP's field of view to be used as the center for the next experiment. Range 0-255.

YRASTR - Y-raster coordinate

The Y-raster position within the UVSP's field of view to be used as the center for the next experiment. Range 0-255.

ITHRSR - Intensity threshold for Dopplergram servo correction

The Dopplergram experiment has an option which allows a drift correction to be applied to the wavelength. The points used in calculating the correction must exceed this threshold to prevent statistical noise and roundoff errors at low intensity levels from unduly affecting the result.

FLAG - Flare Flag

When the HXIS flare flag is issued, the SMM spacecraft computer (OBC) sets the top bit (bit 15) of this word to 1, thereby making the word negative. If bit 14 was previously set to 1, the experiment in progress is terminated; otherwise, it runs to completion. Thus, the user can cho-

ose to respond immediately to the flare flag or to finish his current observation first. If HXIS reports a "super flare", the OBC will set both bits 13 and 15, and the experiment in progress will automatically terminate.

GLMAXH - Global Lambda-max (high)

The high order 2 bits of the wavelength drive position as determined by the last Global Lambda-max experiment.

GLMAXL - Global Lambda-max (low)

The low order 16 bits of the wavelength drive position as determined by the last Global Lambda-max experiment. The user should not normally write into these two locations. However, they are defined as Read-Write because they are loaded by some internally generated Command Mode code and must therefore be legal destinations for the MOVE instruction.

R1, R2, R3, R4, R5, R6, R7, R8 - User scratch registers

The user may use these words as he wishes.

Read-Only Memory

LLMAXH - Local Lambda-max (high)

The high order 2 bits of the wavelength drive position as determined by the last Local Lambda-max experiment.

LLMAXL - Local Lambda-max (low)

The low order 16 bits of the wavelength drive position as determined by the last Local Lambda-max experiment.

FLAREX - X-coordinate of flare

When the flare flag is issued, its X-position as determined by HXIS is loaded into this word. If the user wishes to look at the HXIS location, he simply moves this word to XRASTR.

FLAREY - Y-coordinate of flare

The Y coordinate of the HXIS flare location.

IMIN - The intensity measured at the darkest point during the previous raster.

IMINX - X-coordinate of darkest point measured during the previous raster.

IMINY - Y-coordinate of darkest point measured during the previous raster.

IMAX - The intensity measured at the brightest point during the previous raster.

IMAXX - X-coordinate of brightest point measured during the previous raster.

IMAXY - Y-coordinate of brightest point measured during the previous raster.

BMAX - The wavelength shift measured at the most blue shifted point during the previous raster.

BMAXI - The intensity measured at the most blue shifted point during the previous raster.

BMAXX - X-coordinate of most blue shifted point measured during the previous raster.

BMAXY - Y-coordinate of most blue shifted point measured during the previous raster.

RMAX - The wavelength shift measured at the most red shifted point during the previous raster.

RMAXI - The intensity measured at the most red shifted point during the previous raster.

RMAXX - X-coordinate of most red shifted point measured during the previous raster.

RMAXY - Y-coordinate of most red shifted point measured during the previous raster.

Command Mode

The set of Command Mode instructions contained in a file is called a "Command Mode sequence". A special label, called an "entry point", is used to indicate places where the execution of a sequence can be initiated. Entry point labels are distinguished from other labels by the fact that the first character in an entry point label must be a dollar sign (\$). Each sequence must contain at least one entry point.

Each line of Command Mode text may contain up to five fields. Except that they must be in the proper order, there are no rules as to where the fields must be located on the line. Tabs and spaces are ignored, except that they serve as terminators for opcodes and operands and may not be imbedded within fields.

The first field, which is optional, is the label. A label consists of one to six alphanumeric characters, the first of which must be a letter, and it is terminated with a colon (:). If the label is preceded by a dollar sign (\$), it becomes an entry point. (The \$ is not actually part of the label. Thus, \$ENTRY: is an entry point label, but references to it are written as ENTRY, not \$ENTRY. For example, use GOTO ENTRY to branch to its line.) If an entry point label is terminated with two colons (e.g., \$ENTRY::), it becomes globally defined and can be referenced from other Command Mode sequences. That is, when several Experiment Definition files are combined during Phase-2 to create a single JR load, a START command in one file can cause a transfer to a globally defined entry point in a different file.

Examples:

\$EINAR:	An entry point label
\$GRANT::	A global entry point label
ELMO:	An ordinary label, usable for GOTO ELMO
BRUCE::	Illegal. Only entry points can be global.
JACQUES:	Legal, but the 7th letter will be ignored.

The second field (which may be the first on the line) is the opcode field. The opcodes represent the set of legal instructions which the Command Mode language is capable of executing. Only the first three characters of the opcode are checked for validity, but the user may type the whole word if he desires. Thus, the "start" command may be shown as START or STA, etc. The opcode field is terminated with a space or tab (or semicolon or RETURN, if no operand is required). The legal opcodes are described in the next section.

The third and fourth fields contain the operands, the parameters which the opcode requires in order to function. If two op-

codes are required, they must be separated by a comma. Operands may be either symbols (statement labels, parameter block names, or user buffer locations) or numbers. A number may optionally be preceded by a number sign (#). A number will be interpreted as decimal unless it is preceded by a double quote mark ("), in which case it is treated as octal. (The # must precede the " if both are present.) A trailing decimal point is NOT permitted. The last operand can be terminated with a space, tab, semicolon, or RETURN. The last possible field is the comment field. The comment field is initiated with a semicolon (;). Anything after a semicolon is assumed to be a comment and is ignored. The semicolon is only required when the comment is the only text on the line. If the comment follows Command Mode text, processing of the line ends when the fields required by the opcode have been verified, so the use of the semicolon becomes optional.

To prevent accidental transfer to an undefined location, the last statement in any command mode sequence should be a START, GOTO, or STOP, or the last required field should be terminated with an exclamation point (!) which forces a stop. If this is not done, the PHASE2 compiler will insert a stop bit in the last instruction.

Opcodes

STOP (STO)

Operands: None

Terminates execution of the command mode sequence.

START (STA)

Operand: Entry point name

Causes a jump to an entry point. The entry point name does NOT include the dollar sign (\$). If the entry point is not found within this Command Mode sequence, it must be globally defined in another sequence included in the PHASE2 command generation. Note that, if the entry point IS found in this sequence, there is no effective difference between the START and the GOTO commands.

There is a special form for starting C-list sequences from either A-list or B-list. Instead of using an entry point name as the operand for START, use a backslash (\) followed immediately by a number between 1 and 16. This will transfer control to the n-th C-list entry point. Note that there is no way within Command Mode to return to the

original list once the transfer to C-list has occurred. It requires a command from the OBC to accomplish that.

GOTO (GOT)

Operand: Any label found in this file, including entry points.

GOTO is the "branch" instruction and works in the same way as the Fortran GOTO.

EXECUTE (EXE)

Operand: Experiment parameter block name

This command causes the UVSP instrument to actually take data in the manner specified in the experiment parameter block referenced by the command. When the experiment is completed, processing of Command Mode statements resumes on the following line.

MOVE (MOV)

First operand: Any user buffer location or a number.

Second operand: Any read-write location in the user buffer.

The MOVE command copies requested data from one place to another. It can only write into a word for which the user has write access.

ADD

First operand: Any user buffer location or a number.

Second operand: Any read-write location in the user buffer.

The ADD command performs 16-bit signed integer addition, adding the first operand to the second and storing the sum in the second operand location.

SUBTRACT (SUB)

First operand: Any user buffer location or a number.
Second operand: Any read-write location in the user buffer.

The SUBTRACT command performs 16-bit signed integer subtraction, subtracting the first operand from the second and storing the difference in the second operand location.

COMPARE (COM or CMP)

Operands: Any user buffer locations or a user buffer location and a number.

The COMPARE command compares the two operands, treating them as 16-bit signed integers. If it finds that the first operand is greater than or equal to the second operand, the next Command Mode line is skipped; otherwise, it is executed. Note that the order of the operands is important. COMPARE A,B should be thought of as

IF(A.GE.B) SKIP

COMPARE is the only opcode which can accept a number as its second operand.

AND

First operand: Any user buffer location or a number.
Second operand: Any read-write location in the user buffer.

The AND command performs a 16-bit Boolean "and" of the two operands and leaves the result in the second operand location.

SLIT (SLI)

Operand: A number between 1 and 22 or letter between "A" and "V"

The SLIT command causes the UVSP spectrograph slit to change. There is a dual designation system in which each slit can be identified either by a letter or a number; the SLIT command will accept either type of identifier.

MESSAGE (MES or MSG)

Operand: An unsigned number not exceeding 4095.

The MESSAGE command inserts the designated number into the telemetry stream, encoded in such a way that the receiving software on the ground will recognize it as a message rather than UVSP data. A list of standard messages will be developed at some future date.

TIME (TIM)

Operand: Any read-write location in the user buffer (but should be one of the scratch registers R1 through R8).

JR keeps a count of the number of spacecraft telemetry minor frames which have occurred since sunrise. Since a minor frame takes .064 seconds, this counter can be used as an elapsed time clock. The TIME command copies the minor frame counter into the designated user word.

EXPERIMENT PARAMETER BLOCKS

The Experiment Parameter Blocks are placed at the end of the input file. All of the command mode statements must precede the parameter blocks. The parameter block section is introduced with the line:

.PBLK

This is the only occurrence of this symbol within the file. The first line of each parameter block must begin with the symbolic name of the experiment. It must contain 1 to 6 alphanumeric characters, beginning with a letter, and it must be terminated with one or two colons, depending on whether it is to be a local or globally-defined name. The parameter block consists of a subset of the following lines:

```
EXPER TYPE =  
SLIT =  
LOOP CONTROL =  
INTVL-1 =  
INTVL-2 =  
WAV POSITION =  
WAV OFFSET =  
# OF WAVLENS =  
WAV STEP SIZ =  
POL POSN NUM =  
POL STEP SIZ =  
X POSN NUM =  
X STEP SIZ =  
Y POSN NUM =  
Y STEP SIZ =  
OBSERVATIONS =  
DISABLE INCR  
SERVO INTRVL =  
CALIB INTRVL =  
CALIB AFTER  
CALIB STPSIZ =  
GATE TIME =
```

Some of the lines are mandatory. Others are optional depending on the type of experiment being defined. However, the lines which do appear must occur in the indicated order.

EXPER TYPE =

EXPER TYPE = must be followed by a number between 1 and 21, corresponding to one of the 21 defined types of experiments. They are:

1. SPECTROHELIOGRAM
2. DOPPLERGRAM
3. POLARGRAM
4. MAGNETOGRAM
5. I-MAX
6. I-MIN
7. FLASHWATCH
8. RED-MAX
9. BLUE-MAX
10. SPECTROGRAM
11. LAMBDA-MAX (GLOBAL)
12. LAMBDA-MAX (LOCAL)
13. LAMBDA-MIN (GLOBAL) (Not implemented)
14. LAMBDA-MIN (LOCAL) (Not implemented)
15. SPECIAL
16. PROFILE MATRIX
17. MULTI-LINE PROFILE MATRIX
18. RASTERS THRU THE LINE
19. POLARIZED PROFILE MATRIX
20. POLARIZED MULTI-LINE PROFILE MATRIX
21. POLARIZED RASTERS THRU THE LINE

SLIT =

The slit is designated by a letter between A and V. This line is advisory only, since the experiment control block in JR contains no reference to the slit. However, since the wavelength drive setting for any given wavelength is determined by the slit in use, the Phase-2 processor requires the information. Note that the experiment may not work properly if the wrong slit is in the beam when the experiment is run.

LOOP CONTROL =

JR operates the UVSP through a set of nested DO loops. The order of the nesting and the number of repetitions per loop determine the function of an experiment. The user has control of 4 loops: the X and Y rasters, the wavelength drive, and the polarimeter. The control order is always specified from inner loop to outer loop. Thus, for example, the loop control XYPW would cause a line to be scanned in the X-direction, then the Y raster would be stepped and another X line would be scanned, etc., until the entire raster has been built up. Then the polarimeter wheel would be rotated

and another full raster made. Finally, when an entire set of polarized spectroheliograms had been taken, the wavelength drive would be stepped and the whole process repeated. Of course, the repeat count on some of the loops could be set to 1, effectively removing the operation from the experiment.

Once the experiment type has been chosen, only certain of the operations are relevant. For example, a spectroheliogram requires an X and Y raster only, with neither the wavelength drive nor the polarimeter participating (except for the initial wavelength setting). Only the relevant loop identifiers (X, Y, W, and P) can appear on the line, and all of the relevant ones MUST appear.

INTVL-1 =

The UVSP instrument contains 5 detectors (numbered 1 through 5) and two pulse counters. Because there are only two counters, only two of the detectors can be taking data at any one time. Since it will often be desirable to use four detectors in an experiment (4 lines, or the two wings of 2 lines), each position within an experiment can be divided into two data gathering intervals, with different detectors connected to the counters during each interval. For each interval, one or two detectors may be specified (or, for Interval-2, none). There are two rules governing how detectors can be combined. The first is that, if detector-5 is specified, it must be used alone during that interval. The other is that, if two detectors are specified for an interval, they must be an even- and an odd-numbered detector. Thus, 1 and 2 can be combined, or 1 and 4, but not 1 and 3.

INTVL-2 =

Same as for INTVL-1, except that the Interval-2 detectors may be set to OFF, which means that Interval-2 is not used. When Interval-2 is OFF, all data taking time is used by Interval-1; when both intervals are used, each has a 50 percent duty cycle.

WAV POSITION =

This line indicates an absolute wavelength setting, in Angstroms, at which the experiment is to be started.

WAV OFFSET =

If this line appears, the wavelength drive will be moved to the specified distance (in Angstroms) from the position identified by a previously run Lambda-Max experiment. The offset is followed by the field (LCL) if it is to be interpreted as a local offset (that is, the offset is to be calculated from the position stored in the Local Lambda-Max position contained in the words LLMAXH and LLMAXL in the user buffer) or by (GBL) if it is to be interpreted as a global offset (i.e., using GLMAXH and GLMAXL). Note that the WAV POSITION and WAV OFFSET lines are mutually exclusive; one but not both must appear.

OF WAVLENS =

If the wavelength drive is to move during this experiment, i.e., it is to be a spectral scan of some type, this line contains the number of different wavelengths to be sampled. If the wavelength drive does not move during the experiment, this and the following line do not appear.

WAV STEP SIZ =

This line contains the number of wavelength drive increments which the grating is to be moved for each spectral step. If the grating is being used in 2nd order, which will be the case for most of the slits, one drive increment corresponds to about 50mÅ.

POL POSN NUM =

If the polarimeter is to move, this is the number of positions at which polarimetry measurements will be taken. If the polarimeter does not move, this and the following line do not appear.

POL STEP SIZ =

The polarimeter wheel moves in steps of 22.5 degrees (1/16 rotation). This line shows the number of these 22.5 degree steps the waveplate is to be moved between each measurement. Since the retardation of the waveplate is highly wavelength dependent, one cannot automatically associate a given rotation with a corresponding retardation without knowing the wavelength.

X POSN NUM =

This is the number of points which the X-raster mechanism will take along each X-line. If the X-raster mechanism does not move, this and the next line do not appear.

X STEP SIZ =

This is the size of each raster step in the X-direction. The UVSP rastering mechanism has been designed so that each step is equivalent to one arcsec on the Sun.

Y POSN NUM =

This is the number of points which the Y-raster mechanism will take along each Y-line. If the Y-raster mechanism does not move, this and the next line do not appear.

Y STEP SIZ =

This is the size of each raster step in the Y-direction. The UVSP rastering mechanism has been designed so that each step is equivalent to one arcsec on the Sun.

OBSERVATIONS =

This is the number of times the complete operation is repeated in order to constitute the experiment. For example, a 30-frame movie consists of a spectroheliogram repeated 30 times.

DISABLE INCR

This line commands JR not to increment the experiment sequence number on the second and subsequent times the EXECUTE of this experiment is performed from Command Mode (the sequence number is always incremented on the first EXECUTE). This will allow multiple executions of the experiment to be formatted into a single data file for ground analysis. The DISABLE will remain in effect until one of several conditions, usually the START command or its OBC equivalent, is encountered.

SERVO INTRVL =

After N repetitions of a velocity-type experiment, the grating drive can automatically be moved to center the slits on the mean line position found during those measurements. This line shows the number of complete repetitions of the experiment which must occur before this "servo" balancing is done. The number must be such that $N = (2 * I - 1)$, where I is an integer. If this line does not appear, servo balancing will not be done.

CALIB INTRVL =

Calibration involves offsetting the grating drive by some amount and repeating the measurement cycle. This would generally involve either moving to the nearby continuum to provide a null signal or shifting a spectral by some amount to provide a known signal level. This line shows the number of complete repetitions of the experiment which must be completed before making a calibration run. The number must be of the form $N = (2 * I - 1)$. If no calibration is to be done, either because this is not a velocity-type experiment or because it was not requested, this and the next two lines will not appear.

CALIB AFTER

Calibration is performed after the completion of a specified loop in the loop control. The loop letter can either be one of those shown on the LOOP CONTROL line above, or it can be an S for Servo loop, which is always the outermost loop.

CALIB STPSIZ =

This is the number of wavelength drive steps by which the grating is to be offset to make the calibration number. Steps correspond to 50mA for wavelengths between 1000A and 1850A, and to 100mA for longer wavelengths.

GATE TIME =

This is the integration time per measurement, in seconds.

EXAMPLES

```
; FILE DEMO1.DEF

$ENTRY: EXE LMAX FIRST DO A GLOBAL LAMBDA-MAX
        MOV LLMAXH, GLMAXH
        MOV LLMAXL, GLMAXL

        EXE MOVIE THEN A SPECTROHELIOGRAM MOVIE
        STOP

.PBLK

LMAX:   EXPER TYPE = 11 LAMBDA-MAX (GLOBAL)
        SLIT = M
        LOOP CONTROL = W
        INTVL-1 = 1
        INTVL-2 OFF
        WAV POSITION = 1234.567
        # OF WAVLENS = 128
        WAV STEP SIZ = 3
        OBSERVATIONS = 1
        GATE TIME = 1

MOVIE:: EXPER TYPE = 1 SPECTROHELIOGRAM
        SLIT = M
        LOOP CONTROL = XY
        INTVL-1 = 1-2
        INTVL-2 = 3-4
        WAV OFFSET = -2.004 (LCL)
        X POSN NUM = 16
        X STEP SIZ = 3
        Y POSN NUM = 16
        Y STEP SIZ = 3
        OBSERVATIONS = 30
        GATE TIME = 0.064
```

; FILE DEMO2.DEF

\$START: : EXE FLASH
COM IMAX, 8000
GOTO START
MOV IMAXX, XRASTR
MOV IMAXY, YRASTR
COM 20000, IMAX
START \4
EXE MTRX!

THIS IS A GLOBAL ENTRY POINT

SUPER FLASH?
IF SO, USE EXP4 IN C-LIST
IF NOT, PROFILE MATRIX, THEN STOP

.PBLK

FLASH: EXPER TYPE = 7
SLIT = B
LOOP CONTROL = XY
INTVL-1 = 1
INTVL-2 OFF
WAV OFFSET = 0. (LCL)
X POSN NUM = 8
X STEP SIZ = 3
Y POSN NUM = 8
Y STEP SIZ = 3
OBSERVATIONS = 1
DISABLE INCR
GATE TIME = 0.128

FLASH WATCH

MTRX: EXPER TYPE = 16
SLIT = B
LOOP CONTROL = WXY
INTVL-1 = 1-2
INTVL-2 = 3-4
WAV OFFSET = 0. (LCL)
OF WAVLENS = 11
WAV STEP SIZ = 3
X POSN NUM = 8
X STEP SIZ = 3
Y POSN NUM = 8
Y STEP SIZ = 3
OBSERVATIONS = 1
GATE TIME = 1

PROFILE MATRIX

CHAPTER 3

PHASE-1

PHASE1.TSK is an RSX-11M task which is initiated with the usual RUN PHASE1 command to MCR (or simply PH1 if the task has been installed). After Phase-1 identifies itself, it asks what type of terminal it is being run from. There are three legal answers: answer T if the terminal is a Tektronix 4000-series terminal; answer L if the terminal is a Lear-Siegler ADM-3A; answer D if the terminal is a Decwriter or other printing terminal. (A simple RETURN will default to a T.) The answer allows the program to provide the proper control characters to erase the video terminals and to provide a suitable number of lines per page on the terminal screen. After the terminal question has been answered, the program will prompt with PH1> and wait for a Phase-1 command. The commands are described in the next section.

Phase-1 is program designed to facilitate the creation of Experiment Definition Files. It contains two basic sections, corresponding to the two types of text contained in the Experiment Definition Files: Command Mode text and Experiment Parameter Blocks. The former is handled by a very basic editor capable of inserting, deleting, and listing lines. It also does some simple syntax checking. However, it is by no means idiot-proof. The user can easily create Command Mode text which will be rejected by the Phase-2 processor, which demands perfection. The text will generally be syntactically legal as long as the lines are entered sequentially. However, if lines are deleted or are inserted in the middle of existing text, Phase-1 bears no responsibility for the results, and the user must depend on his/her own proper understanding of the rules for Command Mode instructions. On the other hand, the dialog which creates Experiment Parameter Blocks IS idiot-proof (we hope) and will always produce a legal block.

Some of the experiment types are not completely defined by their Parameter Blocks, but rather require accompanying Command Mode instructions to implement their action. For example, a FLASH WATCH experiment is actually an I-MAX (Intensity Maximum) experiment. The I-Max value is compared to the threshold with Command Mode instructions to determine whether a "flash" has occurred and requires special action. Phase-1 will automatically insert these needed lines of Command Mode text, but to do so it must know the experiment type. Therefore, whenever the user in-

serts a line containing the EXECUTE instruction followed by a Parameter Block name which has not been previously defined. Phase-1 will immediately jump into the Parameter Block dialog. This tends to be annoying, so the user is advised to create all of the Parameter Blocks needed for the file before beginning to insert Command Mode text. This is, however, only a suggestion and not a requirement.

The Experiment Definition Files created as output from Phase-1 can be used directly as input to Phase-2. In particular, Phase-1 automatically provides the .PBLK statement which must separate Command Mode text from Parameter Blocks. The file name is also added as a comment line at the beginning of the text, so that listings can be placed in a documentation file (notebook) without additional identification.

Any Experiment Definition File created by Phase-1 can contain up to 60 lines of Command Mode text and up to 32 Parameter Blocks. Each text line can hold up to 72 characters. (Only 64 columns are printed with the list commands, but all characters are written into the output file. Note that a TAB is a single character but may account for up to 8 columns.) Blank lines are permitted to improve readability, but they count as part of the 60 line limit.

There are no defaults for the names of output files from Phase-1. However, Phase-2 accepts .DEF as the default file type for inputs to it, so the user may find it convenient to use that type unless other naming conventions are developed.

PHASE-1 COMMANDS

When the Phase-1 processor prompts with PH1>, the user must enter a command. Each command consists of a single letter which may or may not be followed by a number. All commands are terminated with a RETURN. Only one command can be entered in response to a prompt. There are nine defined commands. They are:

- A - Abort and restart.
- D - Delete.
- E - Write output file, then exit.
- I - Insert.
- L - List.
- P - Create Parameter Block.
- R - Review Parameter Block.
- T - List top of buffer.
- W - Write output file.
- Z - Exit.

A - Abort and restart.

The A command cancels all of the input received to that point, both Command Mode text and Parameter Block definitions, and allows the user to begin again.

Dn - Delete line-n.

The Delete command requires that a line number be included as part of the command. The command deletes the specified line from the Command Mode buffer. All following lines are immediately re-numbered to reflect their new position in the text. Note that, if you wish to delete a number of successive lines, you must either do it from the bottom up or you must specify the SAME line number for each Delete command, since the following lines get re-numbered each time. For example, to delete lines 4, 5, and 6, use either

PH1>D6
PH1>D5
PH1>D4

or

PH1>D4
PH1>D4
PH1>D4

E - Write output file, then Exit

The E command provides a convenient means of terminating a Phase-1 command generation session. Phase-1 will first ask for the output filename (see the "W" command below for details), then exit after creating the file.

I or In - Insert

The I command allows the user to insert lines of Command Mode text. If the command is used by itself, the text is placed at the end of buffer, following all previously entered lines. If a number is associated with the command, the text will be inserted ahead of the line which currently bears that line number. All lines are terminated with the RETURN key. Command Mode input will continue until the user types the ALTMODE or ESCAPE key. (The exception occurs when a text line includes the EXECUTE opcode for an experiment which has not been previously defined. The program will automatically terminate insert mode and transfer the user to the Parameter Block definition dialogue.)

When Phase-1 is in insert mode, it will automatically place the start of each line 8 spaces from the terminal's left hand margin. This is done to allow room for the line numbers provided by the listing commands (L and T) and then to align new input with the listed text. This spacing is NOT part of the inserted line, and the user will normally want to start the line with a TAB unless it contains a label.

L, Ln, and T - List

The listing commands cause up to 30 lines of Command Mode text (20 lines on a Lear-Siegler terminal) to be displayed. The L command lists the last 30 lines. The T (top) command lists the first 30 lines. The Ln command (where n is a number) lists 30 lines beginning at line-n. If the total Command Mode text does not exceed 30 lines, the L and T com-

mands produce identical results.

P - Create Parameter Block

The command initiates the dialogue required to define an experiment parameter block.

R - Review Parameter Block.

After entering the R command, you will be asked for a Parameter Block name. If the name you specify is that of a Block which has been defined, Phase-1 will list the block on the terminal.

W - Write output file

Once an experiment has been completely defined, it must be written into a disk file. Phase-1 asks for a filename, which the user must fully enter; there are no defaults for either name or type. Once the file is written out, the Phase-1 buffers are cleared, allowing a new experiment to be defined.

Z - Exit

The Z command causes the Phase-1 processor to exit. No output is created at that time, although files previously written out are of course preserved.

PARAMETER BLOCK DIALOG

The Experiment Parameter Block section of Phase-1 is constructed as an interactive dialog which leads the user through the steps required to create a Parameter Block. The hardware controls which are needed in any given Parameter Block depend on the experiment being defined. For example, if you have specified a spectroheliogram as your experiment type, you will NOT be asked for a wavelength step size, since a spectroheliogram is by definition a single-frequency experiment. All inputs to the program are terminated with a RETURN.

For every question you are asked, there will be a default answer. The default will usually be shown between square brackets, i.e. []. You can accept the default by simply typing a RETURN. (Note that you can't use a RETURN to enter a zero unless the default happens to be 0; you must type an explicit 0.) The program contains an internal set of defaults for each experiment type. Whenever you specify a new type, the program resets the defaults accordingly. However, if you are creating an experiment of the same type as the previous one, your values from last time in general become the current defaults.

A facility has been built into the dialog to allow the user to back up any time he decides he has made a mistake. To back up, type CTRL-P when the next question is asked. The program will echo ^P on the terminal and then repeat the previous question. Note that your previous answer has become the new default. You may back up as many steps as you like.

This section describes the prompts and responses needed to create an Experiment Parameter Block. When the dialogue is completed, the entire Parameter Block is printed on the terminal, and the user is asked whether it is OK. If the response is positive, the Block is stored for inclusion in the next output file created. If the user responds with an N, the Block is not saved. However, the user could rapidly step through the dialog to create a slightly different Block because his answers have become the defaults unless a different experiment type is specified. It is also still possible to back up from the OK question using CTRL-P.

PARAMETER BLOCK SYMBOLIC NAME

This question is only asked if you have arrived here by using the Phase-1 "P" command. If you entered the dialog by inserting an EXECUTE line, the Parameter Block name was specified as the operand, so this question is skipped. The default is EXPn, where n is a number which increments automatically if you accept the default. If you specify a name of your own, it must consist of one to six alphanumeric char-

acters (letters and numbers) beginning with a letter.

NEW EXPERIMENT?

A new experiment is one which has to be defined by means of the dialog. However, you may wish to pick up an experiment which was previously defined in a different Experiment Definition File. If your answer is YES, the dialog continues. If your answer is NO you will be asked for a file name. Phase-1 scans the file for a Parameter Block having the correct name. If it finds the block, it copies it in and skips the dialog entirely. The user is shown the contents of the Block by Phase-1.

DECLARE BLOCK NAME GLOBAL?

If identical Parameter Blocks with the same global symbolic names exist in two or more files input to PHASE2 during the creation of a JR load, the Phase-2 processor will only create one copy of the corresponding Experiment Parameter Block in the JR load, thereby saving JR memory. Global symbols will appear in the Parameter Block followed by two colons, while local symbols are followed by a single colon. (Note: you do NOT specify the colon(s) as part of the symbol. The program adds them automatically.)

EXPERIMENT TYPE

The program next prints a numbered list of the possible experiment types with an arrow pointing to the default, and asks for your type selection. It then erases the screen and proceeds with the questions which determine the physical control of the instrument.

SLIT

The answer must be a letter between A and V or a number between 1 and 22. There are two naming systems in use for designating slits, one using letters and the other using numbers. Phase-1 will accept either system. The slit you select determines some of the defaults for other parameters. In particular, the slit width becomes the default value for X step size and for wavelength step size, while the slit length becomes the Y step size. The slit must also be known so that the requested wavelength can be converted to wavelength drive position, which is highly slit dependent. Note, however, that this will NOT cause the selected slit to be moved into

Parameter Block Dialog

the optical path in the UVSP when the experiment is run. The slit mechanism can only be changed with the Command Mode SLIT command or by a command from the OBC. If the wrong slit is in place, the experiment may not produce usable data.

LOOP CONTROL

The experiment control program in JR operates as a set of nested DO loops whose order can be specified. There are five loops to be considered: X-raster, Y-raster, polarimeter step, wavelength step, and Doppler servo. Servo is always the outer loop, but the other four can be put in any order. The first function specified will be the inner loop, the next will be the 2nd loop, etc. All of the loops are not relevant to all types of experiments, and you are only allowed to specify the required ones.

INTVL-1 DETECTOR(S)

The UVSP instrument contains 5 detectors (numbered 1 through 5), and two pulse counters. Because there are only two pulse counters, only two of the detectors can be taking data at any one time. Since it will often be desirable to use four detectors in an experiment (4 lines, or the two wings of 2 lines), each position within an experiment can be divided into two data gathering intervals, with different detectors connected to the counters during each interval. For each interval, one or two detectors may be specified (or, for Interval-2, none). There are two rules governing how detectors can be combined. The first is that, if detector-5 is specified, it must be used alone during that interval. The other is that, if two detectors are specified for an interval, they must be an even- and an odd-numbered detector. Thus, 1 and 2 can be combined, or 1 and 4, but not 1 and 3. To specify two detectors, the user can either type the two numbers consecutively or can separate them with a dash (-). That is, detectors 1 and 4 can be entered either as 14 or as 1-4.

INTVL-2 DETECTOR(S)

Same as for INTVL-1, except that the Interval-2 detectors may be set to OFF, which means that no data is taken during Interval-2. Enter either 0 or OFF to specify the OFF condition.

WAVELENGTH

The user must respond to this query with a floating point number which gives the wavelength in Angstroms. The number will be interpreted in one of two ways, depending on its value. If the number is at least 1000., it will be interpreted as an absolute wavelength. However, if it lies between -1000. and 1000. (exclusive), it will be used as a wavelength offset. In either case, only three places to the right of the decimal point are significant, and trailing zeros may be omitted. The decimal point is optional if a whole number is being entered.

LOCAL OR GLOBAL OFFSET?

If an absolute wavelength was specified in the preceding question, this one will not be asked. If an offset was selected, you must specify whether the offset is to be calculated with respect to the wavelength found by the most recent Local or Global Lambda-Max experiment.

NUMBER OF WAVELENGTHS

This is the number of different equally spaced wavelengths at which data will be taken. The acceptable range of answers is 1-32767. This and the next question are only asked for experiment types which require motion of the grating, not including the initial wavelength setting.

WAVELENGTH STEP SIZE

Your response to this question must be the number of mechanical steps of the grating drive mechanism which you desire. If the grating is being used in 2nd order, each step corresponds to 50mÅ. The range of acceptable responses is 1-31.

NUMBER OF POLARIMETER POSITIONS

The acceptable range is 1-32. This will be the number of measurements you wish to take in different polarization states at a given point.

C-2

POLARIMETER STEP SIZE

Each polarimeter step is a 22.5 degree rotation, or 1/16 of a full circle. The acceptable range of responses is 1-8.

X-POSITION NUMBER

This is the number of points in a line along the X-direction. Answers in the range 1-256 are acceptable. The initial default will be the maximum range (255) divided by the width of the selected slit.

X-STEP SIZE

This question controls the size of the raster step in the X-direction. The default value is the width of the selected slit, unless a velocity type experiment (Dopplergram) is being done. In that case, the default X-step is 1 arcsec, so that measurements can be averaged to suppress spurious velocity signals due to intensity inhomogeneities across the slit (the Beckers effect). The legal range is 1-255. However, the product of the number of steps and the step size cannot exceed 255.

Y-POSITION NUMBER

This is the number of points in a line along the Y-direction. The acceptable range is 1-256. The default is set such that a Y-step size equal to the slit length will produce a square raster.

Y-STEP SIZE

The default value is the length which will produce a square raster. Answers in the range 1-255 are legal, but step size times step number cannot exceed 255.

NUMBER OF OBSERVATIONS

This is the total number of repetitions of the experiment, including the first, but excluding servo and calibration cycles. The answer must be in the range 1-16383. Certain types of experiments, those which are looking for a minimum or maximum value within the scan, are by their nature restricted to a single execution. For these experiment types,

this question is skipped.

DISABLE SEQUENCE INCREMENT?

In normal operation, the "sequence" number which is internally generated by JR is incremented at the beginning of each experiment. When the data stream is processed on the ground, the reformatter program will use this number to determine when a new data file should be started. If the sequence number does not change, more than one experiment will be placed in a single file. Some types of experiments (flash-watch, multi-line profile matrix) are intrinsically designed to have multiple observations in a single file, so the sequence increment will be automatically disabled. Other types (Lambda-max, I-max, etc.) are by their nature single measurements, so the number always increments. For the remaining types, the user is offered the option of disabling the automatic incrementation. Note that this feature only affects the re-using of an experiment after first returning to Command Mode; multiple observations as specified in the answer to the OBSERVATIONS question are always placed in a single data file.

SERVO INTERVAL

After N repetitions of a velocity-type experiment, the grating drive can be automatically moved to center the slits on the mean line position found during those measurements. The desired number of complete experiment repetitions (including the first) which must occur before this balancing is done is called the "servo interval". Legal responses range from 0 (to suppress the operation) to the number of observations previously specified, but must satisfy the equation $N = (2 * I - 1)$.

CALIBRATION INTERVAL

Calibration involves offsetting the grating drive by some distance and repeating the previous measurement cycle. It generally involves moving the spectrograph to a nearby continuum position to provide a null signal or shifting the spectral line by a set amount to inject a known signal level. The legal range is the same as for the servo interval.

CALIBRATE AFTER

Calibration is performed after the completion of a specified loop in the loop control. The legal responses (given with a single letter) are the loops used in this type experiment or "S", which refers to the servo loop and is always the outermost loop.

CALIBRATION STEP SIZE

Respond with the number of grating steps by which the spectrometer must be offset to do the calibration measurement. The grating will automatically be returned to its previous position after completion of the calibration cycle.

GATE TIME

The gate time is specified in seconds, and is a floating point number (although the decimal point is optional for a whole number). Any value greater than zero is legal, although anything less than .064 sec will merely waste photons, since that is the telemetry period between data values.

CHAPTER 4

PHASE-2

OPERATION

The Phase-2 Command Generation processor is the program which compiles the desired Experiment Definition Files into a JR load. Phase-2 will normally be run once per day by the daily planner or his/her appointee. Before Phase-2 can be run, the entire load must have been planned. Because Phase-2 is generating an actual memory load for JR, it is extremely intolerant of errors. If you specify an Experiment Definition File which it can't find, it will notify you and let you try again. ALL OTHER ERRORS ARE CONSIDERED FATAL, and Phase-2 exits after issuing an error message.

From the user's point of view, the operation of Phase-2 is extremely simple. The program is initiated with RUN PHASE2 (or simply PH2 if the task has been installed). Phase-2 first informs the user of the load which is to be replaced by the one about to be created. For example:

```
---> SUPERCEDING A-LIST LOAD CREATED AT 11:29AM ON 18-OCT-79
```

The operator should verify that this was indeed the previous load uplinked to JR. If it was not, Phase-2 will have incorrect knowledge of what part of JR memory is available to it, and unpredictable results may occur. (If this message does not match reality, the operator should exit from Phase-2 and, using PIP, find the version of the file JRMAP which was created at the time and date of the previous load creation, then copy that file using the /NV switch to make it the latest version of JRMAP.)

Phase-2 then types LIST (A, B, OR C): and the operator responds with the observing list to be created. (RETURN with no letter causes a clean exit.) If the answer is C, the program types

```
*** WARNING - CURRENT B-LIST WILL BE DESTROYED ***  
DO YOU WANT TO PROCEED? [Y/N]
```

Any answer other than Y will cause an exit. Due to the way in which Phase-2 manages JR's memory, a C-list load can only be

uplinked on a day on which B-list is not active (i. e., an A-day). The B-list load which will normally be sent up on the same day must be created after the C-list load.

Finally, Phase-2 will ask for a FILENAME, and the operator responds with the name of an Experiment Definition File. Phase-2 accepts .DEF as its default file type, but there is no filename default. After Phase-2 has processed that file, it will prompt for another one, and will continue the process until the operator responds to a FILENAME prompt with a simple RETURN. Phase-2 then completes the creation of the JR load and exits.

OUTPUT FILES

The Phase-2 processor creates a number of files. In the discussions which follow, we will assume that an A-list load has been created. If it had been a B-list or C-list load, those filenames which are shown beginning with an "A" would begin with a "B" or "C" instead.

ALOAD. JRO

This is the actual binary load file. It or a derivative of it must be passed to the SMM Command Management System for uplink to the spacecraft. The file is in a format compatible with the JR Test Interpreter designed by Roger Rehse.

ALIST

This is the listing file which will be used during the observing day to monitor the action of the UVSP and should be retained as part of the archival record of mission operations.

The listing contains two columns. The left hand side shows all of the Command Mode text contained in the load, along with the absolute address and octal contents of each word. Entry points are flagged with their appropriate letter and number, followed by a right angle bracket (e.g., A4>). References to entry points via START instructions have the corresponding entry point flag shown between angle brackets (e.g., START FLARE <A12>). Parameter Blocks are similarly referenced, except that A-list Parameter Blocks are flagged with an X, B-list Blocks with a Y, and C-list Blocks with a Z, and they are enclosed in square brackets (e.g., EXECUTE BLOCK [X3]).

The right hand column shows the symbolic and flag names of each Parameter Block, along the address and contents of each word in JR memory. The meanings of the sub-fields (bit patterns) within each word are verbally described beside the word. This display is better suited to showing what JR will do with the UVSP mechanisms than what the scientific intent of the experiment is, and it should therefore be a useful tool for trouble-shooting if necessary.

ALIST.VEC

This is the "Vector Association" file. It is a readable text file which lists, for each entry point, the symbolic label of that entry point and the name of the Experiment Definition File in which the entry point is defined.

JRMAP

JRMAP is a file which maintains a record of the 923 words of JR memory to which Phase-2 has access. The file contains three records. The first contains the list, date, and time of the load creation. The second record contains two 923-byte arrays. Each byte in the first array contains one of the characters A, B, C, X, Y, or Z, or a zero. A, B, and C refer to A-list, B-list, and C-list Command Mode words respectively, while X, Y, and Z refer to A-list, B-list, and C-list Parameter Blocks. A byte containing a zero is not assigned to any of the observing lists. For each byte which contains a letter, the corresponding byte in the other array contains the number of the entry point or Parameter Block within that list. Thus, if the bytes corresponding to a given JR word contain "B" and 4, the word is part of the Command Mode code following entry point B4, while Z12 would belong to the 12th Parameter Block in C-list. (The first 49 bytes in each array, which correspond to the observing list vectors, are not filled in.) The third record is a 923 word array containing the JR memory image, that is, the actual contents of JR's memory after this load has been uplinked.

The records can be read and the heading typed with the following Fortran code:

```
DIMENSION JRIMAG(923),JRMAP(923)
BYTE MAP(2,923),LIST,AMPM,TIM(8),DAT(9)
EQUIVALENCE (MAP,JRMAP)

CALL ASSIGN(1,'JRMAP')
READ(1) LIST,TIM,AP,DAT
READ(1) JRMAP
READ(1) JRIMAG
TYPE 201, LIST,(TIM(I),I=1,5),AMPM,DAT
201 FORMAT(' 'A1,'-LIST LOAD CREATED AT '5A1,A1,'M ON '9A1)
```

APPENDIX 3
FLIGHT SOFTWARE PACKAGE LISTING

```

1      ;      *** JR SOFTWARE: POST LAUNCH; PATCHED ***
2      ;      ** WLD 'FLYBACK' DISABLED
3      ;      ** WLD STEPPED AT UNIFORM 100-HZ RATE
4      ;
5      ; MONITOR: EXPERIMENT EXECUTION CONTROL
6      ;      ENTRY POINTS:  MONITR = START NEW EXPERIMENT
7      ;                      MONITI = ACT ON NEW S/C DATA FROM SCP
8      ;                      POWRUP = POWER UP AND INITIALIZE
9      ;                      PWROFF = POWER DOWN ENTIRE INSTRUMENT
10     ;                      HALT = STOP EXPERIMENT IN PROGRESS
11     ;      . = 0
12     0 140260      JMP      POWRUP ; POWER UP AND INITIALIZE INSTRUMENT
13     ;
14     ; INTERRUPT VECTORS
15     1 140007      JMP      LEVEL1 ; POWER FAIL INTERRUPT
16     2 133000      HLT      ; DMA INTERRUPT
17     3 170102      JST      LEVEL3 ; HOUSEKEEPING INTERRUPTS
18     ;
19     ; WORD 4 IS A RESTART LOCATION AFTER LOADS OR HALTS
20     4 133400      CIL      ; CLEAR INTR 3
21     5 160316      JST      TSTD   ; TEST DAY FOR RESTART
22     6 140265      JMP      FWUP2  ; RESTART
23     ;
24     7 137002      LEVEL1: LSA 2
25     10 095777     DM5%     PFFLAG ; DECREMENT POWER FAIL FLAG
26     11 140327     JMP      ALLOFF ; TURN OFF MOTORS
27     ;
28     12 140012     HALT: JMP    ; HALT CURRENT EXPERIMENT
29     13 161612     JST      DTIM   ; HALT GATES, JR TM CTRL, DET OFF
30     14 133400     CIL      ;
31     15 061727     LIO      MSG3   ; OUTPUT 'END OF EXPT' MSG
32     16 171353     JST      USMSG  ;
33     17 140012     JMP      HALT   ; RETURN
34     ;
35     20 137003     MONITI: LSA 3 ; MONITOR ENTRY POINT FROM SCI CMD
36     21 125775     SIN%     TEMPI  ; SAVE INDEX TEMPORARILY
37     22 045700     LAC%     SCI     ; SEE IF ITS INLINE PROCESS
38     23 021732     ORR      SMASK  ; IE: BITS 11-10 ARE SET
39     24 132400     NOT      ;
40     25 130000     TAZ      ;
41     26 140071     JMP      INLINE ; YES - DO OPERATION
42     27 055777     LAN%     BSYGN  ; SEE IF IS ALL RIGHT TO PROCESS COMMAND
43     30 130400     TAN      ; NOT IF BUSY = NEG (POWERUP OR WLD FAIL)
44     31 151265     JMP      SAVSA  ; RESTORE REGISTERS AND RETURN
45     32 045700     LAC%     SCI     ; BUSY IS POS; GO AHEAD
46     33 021730     ORR      SGNBIT ; SET BIT 16 = 1 (CMD PROCESSED)
47     34 115700     SAC%     SCI     ; REPLACE SCI
48     35 134012     ASR      10     ; NOT INLINE, NEW EXPERIMENT REQUESTED
49     36 011715     AND      THREE  ; BITS 11-12 DEFINE ACTION
50     37 160044     JST      INJMP  ; JUMP TO APPROPRIATE ROUTINE
51     ; *****
52     40 140125     JMP      NEW     ; 00 = NEW EXPERIMENT VECTOR
53     41 155663     JMP      DIVISE  ; 01 = DOOR/SLIT/POLR COMMAND TO JR
54     42 140124     JMP      XFLARE  ; 10 = FLARE EXPERIMENT FLAG
55     43 140051     JMP      OTHER  ; 11 = OTHER OPERATION
56     ; *****
57     44 142000     INJMP: JMPI 0

```

ORIGINAL PAGE IS
OF POOR QUALITY

```

1   45 115774      SAC$  TEMP   ; SAVES WORDS
2   46 045700      LAC$  SCI    ; PASS (SCI) IN AC REG
3   47 075774      LIN$  TEMP   ; CREATE INDEXED JUMP
4   50 140044      JMP    INJMP
5
6   51 134004      OTHER: ASR 4   ; SPECIAL OPERATION AC=(SCI)
7   52 011715      AND   THREE  ; CMD TYPE IS BITS 10-5
8   53 160044      JST   INJMP   ; INDEXED JUMP TO TYPE
9
10  54 140514      JMP    JRLOAD  ; JR LOAD UPCOMING, WAIT
11  55 141753      JMP    ENDWT   ; WAIT
12  56 140060      JMP    CPUDMP  ; MEMORY DUMP
13  57 140057      JMP    .+0
14
15
16  50 160012      CPUDMP: JST HALT
17  61 137003      LSA   3
18  62 045700      LAC$  SCI    ; RENEW INSTRUCTION
19  63 011715      AND   THREE  ; GET SEGMENT DATA
20  64 134412      ASL   10.   ; MOVE TO POSITION
21  65 137002      LSA   2
22  66 115701      SAC$  AB    ; SET ADDRESS REG.
23  67 171151      JST   MEMDMP  ; PERFORM DUMP
24  70 140070      JMP    .+0    ; WHAT NOW?
25
26  71 045700      INLINE: LAC$ SCI  ; INLINE OPERATION - DON'T RESET PHANTOM
27  72 021730      ORR   SGNBIT  ; SET COMMAND PROCESSED FLAG
28  73 115700      SAC$  SCI
29  74 134404      ASL   4
30  75 130400      TAN
31  76 151501      JMP    INLHTR  ; DO IN LINE HEATER SWITCHING
32  77 134405      ASL   5
33  100 134015     ASR   13.   ; BITS 9-5 INDICATE OPERATION
34  101 160044     JST   INJMP  ; RESTORE INDEX, AC, SA
35
36  102 140111     JMP    FLFLAG  ; SET FLARE FLAG 700X X=0,2
37  103 140116     JMP    ALIST   ; SELECT LIST A
38  104 140120     JMP    BLIST   ; SELECT LIST B
39  105 145427     JMP    FIXWL   ; CHANGE WLD POSN COUNTER
40  106 061747     LIO   0377   ; DEBUG AID TO DISABLE INTRPTS
41  107 136520     EXI   MASK
42  110 140122     JMP    INRTN   ; RETURN AFTER MASKING
43
44  111 134414     FLFLAG: ASL 12. ; INLINE FLARE FLAG AC=(SCI)
45  112 025733     ORR$  FLAG
46  113 021730     ORR   SGNBIT
47  114 115733     SAC$  FLAG
48  115 140122     JMP    INRTN
49
50  116 041735     ALIST: LAC D16 ; SET POINTER OFFSET TO NEW LIST
51  117 140121     JMP    BLIST+1
52
53  120 041736     BLIST: LAC D32 ; SET ADDRESS INCREMENT TO 40 OCTAL
54  121 114060     SAC$  LIST
55  122 075775     INRTN: LIN$ TEMPI ; RESTORE INDEX REGISTER
56  123 151265     JMP    SAVSA  ; RESTORE REST OF REGISTERS AND RETURN
57

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1 124 041744 XFLARE: LAC 017 ;DEFAULT TO SEQUENCE 15.
2 125 011744 NEW: AND 017 ;SCI CMD ENTRY TO GONEW AC=(SCI)
3 126 004060 NEWFL: ADD$ LIST ;ADDING THE VALUE FOR A/B LIST
4 127 115703 GONEW: SAC$ OLS ;NEW INDEX TO VECTOR TABLE SET
5 130 040071 LAC INLINE ;=045700
6 131 115702 SAC$ LPC ;LIST PROG CNTR IS FLAGGED
7 132 041713 LAC ONE ;RESTORE EXPT NO. CHANGE
8 133 115677 SAC$ EXPADV
9 ;DEVPON ;BRING DEVICE POWER UP FOR EXPT SEQ.
10 134 133400 CIL
11 135 136755 EXI ECTR ;CLEAR EVENT COUNT
12 136 061740 LIO 0177
13 137 136610 EXI RXLD ;START RASTER AT ELECT NULL
14 140 136614 EXI RYLD
15 141 041716 LAC FOUR ;POWER RASTER UP
16 142 164232 JST RWENAB ;OUTER GIMBAL
17 143 137002 LSA 2
18 144 045631 LAC$ 05153 ;128 MS WAIT
19 145 170360 JST UWTX
20 146 041714 LAC TNO
21 147 164232 JST RWENAB ;INNER GIMBAL
22 150 137002 LSA 2
23 151 045634 LAC$ 03777
24 152 170360 JST UWTX ;98 MS WAIT
25
26 153 136761 EXI QBRN ;'A' PLATE :: REVERSAL FOUND 11/79
27 154 045623 LAC$ D1137
28 155 170360 JST UWTX
29 156 045735 LAC$ PRREFC
30 157 130000 TAZ
31 160 140162 JMP MONITR
32 161 170706 JST GETPRF ;GET PLATE REF POSN
33
34 MONITR: ;ENTRY POINT AFTER SCI 'NEW' EXPMT OR AFTER
35 ;:::::: 'A 'START', ALSO AT END OF ALL EXPERIMENTS.
36 162 137003 LSA 3
37 163 045702 LAC$ LPC ;TEST WAIT/GO BIT IN LIST PROGRAM CNTR
38 164 130400 TAN
39 165 140170 JMP 1$ ;NOT SET
40 166 161632 JST DEVOFF ;UNPOWER RASTER AND ROTATING PLATE
41 167 140307 JMP WAIT ;DO WAIT
42
43 170 134401 1$: ASL 1
44 171 130400 TAN ;TST IF LPC FLAGGED BY SCI/START
45 172 140205 JMP NXTCMI ;CONTINUATION OF O.L. SEGMENT
46 173 045703 LAC$ OLS ;SCI OR START WAS LAST CALL
47 174 130400 TAN ;TEST IF OBSERVING LIST SEGMENT INDEX IS VALID
48 175 140177 JMP .+2 ;NOT NEGATIVE, OK
49 176 140253 JMP BADMOD ;ILLEGAL
50 177 001733 ADD M48D ;NEEDS TO BE .LT. 48
51 200 130400 TAN
52 201 140253 JMP BADMOD ;ILLEGAL
53 202 075703 LIN$ OLS ;LOAD OBSERVING LIST SEGMENT=VECTOR TABLE INDEX
54 203 046000 LACIS VECTOR ;LOAD NEW LIST PROGRAM COUNTER
55 204 115702 SAC$ LPC ;LPC IS LIST PROGRAM COUNTER
56
57 205 045702 NXTCNI: LAC$ LPC ;NEXT POINTER TO COMMAND MODE INSTR

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```

1  206 001734      ADD    M1616  ;TEST FOR UPPER LIMIT
2  207 130400      TAN      ;MUST BE .LT. END OF PARAM TABLE
3  210 140253      JMP     BADMOD  ;ILLEGAL ADDRESS
4  211 045702      LAC%   LPC
5  212 021726      ORR    0150K
6
7  213 171403      ; ***
8                      ; ***
9  214 137003      LSA    3
10 215 075702      LIN%   LPC      ;GET INDEX TO OBS. LIST
11 216 066061      LIO%I  OESLST  ;UPDATE INSTRC. REG FROM OBS. LIST
12 217 105701      SIO%   LIR
13 220 171353      JST    USMSG   ;SEND THE CMD INSTRC
14 221 137003      LSA    3
15 222 045701      LAC%   LIR
16 223 011742      AND    01TZ17 ;SAVE SIGN AND LOW 4 BITS
17 224 115776      SAC%   TEMPJ  ;MASKED INSTRUCTION
18 225 075776      LIN%   TEMPJ  ;USE AS INDEX
19 226 011730      AND    SGNBIT ;SAVE SIGN BIT FOR WAIT FLAG
20 227 025702      ORR%   LPC
21 230 001713      ADD    ONE    ;SET WAIT/GO AND INCREMENT LIST PC
22 231 115702      SAC%   LPC
23
24 232 142233      JMPFI  .+1
25
26                      ; JUMP TABLE 16 POSSIBLE ENTRYS  NOP COMMAND INSTRUCTIONS ARE ERRORS
27
28 233 150000      JMP    EXECUT  ; 10 BIT PARAM LIST INDEX
29 234 140466      JMP    START   ; LIKE SCI CMD. C BIT / 4 LIST INDX / 0001
30 235 140476      JMP    GOTOO   ; 5 BIT REL OFFSET / 0010
31 236 141664      JMP    COMPAA  ; SKIP IF A/IMED .GE. B
32 237 141674      JMP    ADDD    ; IMED BIT / 2 5-BIT FIELDS / 0100
33 240 150041      JMP    SUBB    ; IMED BIT / 2 5-BIT FIELDS / 0101
34 241 140511      JMP    MOVEE   ; IMED BIT / 2 5-BIT FIELDS / 0110
35 242 141703      JMP    CMPR    ; SKIP IF A/IMED .LE. B
36 243 140162      JMP    MONITR  ;UNUSED BITS ARE A MESSAGE
37 244 150340      JMP    MITIM   ; MINOR FR CLK SINCE JR ON OR LOADED
38 245 151512      JMP    SIZE
39 246 150044      JMP    IAND    ;BIT AND FUNCTION
40 247 134000      NOP      ; FORMERLY "ADD OVFL BIT TO 'B' "
41 250 134000      NOP
42 251 134000      NOP
43 252 134000      NOP
44
45                      ; END OF TABLE
46
47 253 171425      BADMOD: JST UERROR ; ILLEGAL CMD MODE OR LIST ADDRESS
48 254 140307      JMP     WAIT
49 255 171425      BADLST: JST UERROR ; ILLEGAL PARAMETER WORD
50 256 137003      LSA    3
51 257 140205      JMP     NXTCMD ; TRY NEXT CMD
52
53 260 136561      POWRUP: EXI SMN ; POWER UP STAT MON.--MAY BE OFF
54                      ; ***
55 261 165073      JST    GTBUSY ;SET BUSY=NEG SO SCI CMDS WON'T DISRUPT
56                      ; ***
57 262 170365      JST    UWTB

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1 263 136713      EXI   TM2N   ; POWER BUS 2
2 264 136717      EXI   TM3N
3 265 161612      FWUP2: JST DTIM ; SET CLOCKS, TM
4 266 161632      JST   DEVOFF ; SET RASTER PWR
5 267 136575      EXI   ORDN   ; SET DAY/NIT OVERRIDE LATCH
6 270 170365      JST   UWT8   ; 8 MSEC
7
8 271 136645      EXI   WAVN   ; ENABLE WLD
9 272 170365      JST   UWT8
10 273 165103      JST   SMWLD  ; SET COUNTER = WLD POSN
11 274 132000      CAC
12 275 137002      LSA 2
13 276 115637      SAC$  CNTOBC ; RESET OBC 'OK' INDICATOR
14 277 137003      LSA 3 ; RESTART AFTER JR LOAD
15 300 115771      SAC$  ACONFG ; SET ACONFG IN DMA BUFFER
16 301 115773      SAC$  MICKL  ; RESET MINOR FRAME CLOCK
17 302 061711      LIO   UNMSK  ; UNMASK MANY INTERRUPTS
18 303 136520      EXI   MASK   ; FLAR, MI, MI/4, G3, G2
19 304 041730      LAC   SGNBIT
20 305 115702      SAC$  LPC    ; SET WAIT/GO TO 'WAIT'
21 306 161641      JST   BSYCLR ; SET BUSY TO + TO ALLOW SCI PROCESSING
22 307 160012      WAIT: JST HALT ; CLEAR INTERRUPT VECTORS
23 310 055677      LAN$  HEATWD
24 311 130400      TAN
25 312 171463      JST   HEATR  ; SERVICE HEATER
26 313 140313      JMP   .+0
27
28 ; DAY/NIGHT TRANSITION: NOP IF TO DAY, POWER OFF IF TO NIGHT
29
30 ; ***
31 314 160316      D2N:  JST   TSTD  ; TEST DAY/NIGHT LEVEL
32 315 140260      JMP   POWRUP
33 ; ***
34 316 140316      TSTD: JMP
35 317 061737      LIO   D&2   ; GET DAY NIGHT STATUS
36 320 170051      JST   USMR   ; INTERROGATE STATUS MONITR
37 321 134407      ASL  7 ; BIT 9 IS LEVEL
38 322 130400      TAN ; DAY OR NIGHT?
39 323 151175      JMP   TNLP   ; TST WLD POWER
40 324 140316      JMP   TSTD
41
42 325 160012      EXOFF: JST   HALT  ; TURN OFF EVERYTHING AT NIGHT
43 326 170371      JST   UNTONE ; ***2ND 2-SEC. WAIT REMOVED
44 327 161612      ALLOFF: JST DTIM ; TURN OFF EVERYTHING
45 330 161632      JST   DEVOFF
46 331 133400      CIL
47 332 140310      JMP   WAIT+1
48 ; ***
49
50 ; BOOTSTRAP LOADER: ALTERNATE LOAD/DUMP ROUTINE
51 ; ENTRY POINTS: BOOT = FROM L3 POLLING LIST AFTER S/C COMMANDS
52 ; RT, MPR
53 ; S4, BOOT
54 ; ARGUMENTS: BOOT IS OPERATED BY REPEATEDLY SENDING THE FOLLOWING
55 ; S/C COMMAND SEQUENCE:
56 ;
57 ; RT, S3, DRDL+X

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1          ;          RT, S3, WAVI+Y
2          ;          RT, JRR
3          ;
4          ;          WHERE X = 10 = SET MS 8 BITS OF AB = Y
5          ;
6          ;          = 20 = SET LS 8 BITS OF AB = Y
7          ;
8          ;          = 40 = SET SM 8 BITS OF DATA = Y
9          ;
10         ;          = 100 = SET LS 8 BITS OF DATA = Y AND
11         ;          STORE DATA IN ADDRESS SPECIFIED
12         ;          IN AB, THEN INCREMENT AB
13         ;
14         ;          = 200 = BEGIN DUMP AT ADDRESS IN AB.
15         ;          DUMP WILL END AT LOCATION 7777.
16         ;
17         ; NOTE: MP MUST BE RESET TO USE BOOT
18         ;
19 333 061747 BOOTX: LIO 0377
20 334 136520      EXI   MASK   ;DISABLE INTERRUPTS
21 335 136713      EXI   TM2N   ; ENABLE BUS 2 AND WAV JUST IN CASE
22 336 136717      EXI   TM3N
23 337 136645      EXI   WAVN
24 340 133400      CIL
25         ; ***
26 341 060352 BOOT: LIO CLEAR+7 ; SET PATH TO 241
27         ; ***
28 342 136721      EXI   321
29 343 170401 CLEAR: JST GET8 ; SET LS 8 BITS OF WLD TO ZERO
30 344 136642      EXI   SLWR
31 345 111710      SAC   TEMPB
32 346 061710      LIO   TEMPB
33 347 136654      EXI   WAVI
34 350 041711      LAC   UNMSK ; WAIT 8 MS FOR WAVI
35 351 170360      JST   UWTX ; SA RETURNED = 2
36 352 136641      EXI   SLWF ; RESET FWD FOR WAVI
37 353 133000      HLT   ; WAIT FOR JRR == INTERRUPTS NEED TO BE MANAGED.
38
39 354 060432      LIO   D47   ; GET PATH FROM SM
40 355 170051      JST   USMR
41 356 134410      ASL   8.    ; LOOK AT ONLY BITS 8-5
42 357 134014      ASR   12.
43
44 360 130000      TAZ
45 361 140401      JMP   ABHI ; PATH = 10 ?
46 362 001722      ADD   M1   ; MS 8 BITS OF AB
47 363 130000      TAZ
48 364 140406      JMP   ABLO ; PATH = 20 ?
49 365 001722      ADD   M1   ; LS 8 BITS OF AB
50 366 130000      TAZ
51 367 140412      JMP   DATAHI ; PATH = 40 ?
52 370 001723      ADD   M2   ; MS 8 BITS OF DATA
53 371 130000      TAZ
54 372 140416      JMP   DATALO ; PATH = 100 ?
55 373 001724      ADD   M4   ; LS 8 BITS OF DATA / STORE
56 374 130000      TAZ
57 375 140425      JMP   DUMP ; PATH = 200 ?

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1
2 376 061747      LIO  0377  ;PATH UNDEFINED
3 377 136721      EXI  321  ;SET PATH = 377
4 400 140343      JMP  CLEAR  ; CLEAR WLD AND WAIT
5
6 401 170401  ABHI: JST GET8 ;GET MS 4 BITS OF AB
7 402 134414      ASL  12.  ; MASK AND MOVE TO MS HALF
8 403 134004      ASR  4
9 404 115701      SAC$ AB  ; STORE
10 405 140341     JMP  BOOT  ; NEXT OPERATION
11
12 406 170401  ABLO: JST GET8 ; GET LS 8 BITS OF AB
13 407 025701     ORR$ AB  ; INSERT INTO AB
14 410 115701     SAC$ AB
15 411 140341     JMP  BOOT
16
17 412 170401  DATAI: JST GET8 ; GET MS 8 BITS OF DATA
18 413 134410     ASL  8.  ; MOVE TO MS HALF
19 414 111707     SAC  DATA ; STORE
20 415 140341     JMP  BOOT
21
22 416 170401  DATAO: JST GET8 ; GET LS 8 BITS OF DATA
23 417 021707     ORR  DATA ; INSERT INTO DATA
24 420 111707     SAC  DATA
25 421 060433     LIO  SACDLR ; LOAD SAC$ INSTRUCTION
26 422 170410     JST  ASSMBL ; ASSEMBLE AND EXECUTE INSTRUCTION
27 423 170425     JST  INCAB  ; STORE DATA, THEN INCREMENT AB
28 424 140341     JMP  BOOT
29
30 425 133400  DUMP: CIL      ; ENABLE L3 INTERRUPTS
31 426 061711     LIO  UNMSK
32 427 136520     EXI  MASK  ; SET MASK REGISTER
33 430 171151     JST  MEMDMP ; USE STANDARD DUMP FORMAT
34 431 140333     JMP  BOOTX
35
36 432 000057  D47: 47.
37 433 114000  SACDLR: SAC$ 0
38
39 ; GET ARGUMENTS FROM COMMAND INSTRUCTION FIELDS A, B
40 ; A DATA IS RETURNED IN IO REG, 'TEMPJ'
41 ; B INDEX IS RETURNED IN 'IN' OR ACC REGS
42
43 434 075702  FIELD1: LINS  LPC  ;GET IMMED INDEX
44 435 066061  LIO$I  OBSLST ;LD IMMED DATA
45 436 137401  BIN      ;INCREMENT PC VALUE
46 437 125702  SINS  LPC
47
48 440 045701  FIELD2: LAC$  LIR  ;GET INDEX OF 'B' DATA
49 441 134004  ASR  4
50 442 011745  AND  037
51 443 001731  ADD  RWBUF  ;ADD SAME BASE FOR READABLE MEMORY
52 444 115776  SAC$  TEMPJ ;SAVE 'B' INDEX VALUE FOR LATER USE
53 445 075776  LINS  TEMPJ ;INDEX SET FOR 'B' DATA
54 446 105776  SIO$  TEMPJ ;PASS 'A' INDEX
55 447 140447  FIELDS: JMP  +0 ;FIND INDEXES FROM INSTRUCTION INDEX FIELDS
56 450 045701  LAC$  LIR  ;FIRST TEST C FIELD IN INSTRUCTION
57 451 134401  ASL  1

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1 452 071712      LIN      ZERO
2 453 130400      TAN          ;SKIP IF 'C' IS SET
3 454 071713      LIN      ONE
4 455 134012      ASR      10.    ;ALIGN BITS FOR 'A' INDEX FIELD
5 456 011745      AND      037    ;MASK REMAINING BITS
6 457 142460      JMPI     .+1    ;ACTUAL TEST OF 'C' BIT
7 460 140434      JMP      FIEL1  ;'C' WAS SET; USE IMMED. DATA
8 461 001731      ADD      RWBUF  ;ADD BASE OF RD/WRT MEMORY
9 462 115776      SAC$    TEMPJ  ;SAVE FOR INDEX
10 463 075776     LINS     TEMPJ
11 464 066000     LIOSI   6000    ;'A' DATA FROM SEG 3
12 465 140440     JMP      FIEL2
13
14 466 045701     ; START: LAC$ LIR ;REFRESH THE INSTRUCTION
15 467 071713     LIN      ONE
16 470 134407     ASL      7      ;SHIFT 'C' FLAG BIT TO SIGN POSN
17 471 130400     TAN
18 472 071712     LIN      ZERO    ;INDEX 0 IF NOT C LIST EXPMT
19 473 134013     ASR      11.   ;SHIFT LIST INDEX TO POSITION
20 474 011744     AND      017    ;MASK 4 BIT LIST INDEX
21 475 142126     JMPI     NEWFL  ;NORMAL I=0; C I=1
22
23
24 476 045701     ; GOTOD: LAC$ LIR ;REFRESH INSTRUCTION IN ACCUM
25 477 134405     ASL      5.    ;TEST IF JUMP IS BACKWARD
26 500 130400     TAN
27 501 140505     JMP      1$
28 502 134011     ASR      9.    ;EXTEND SIGN BIT OF 2'S COMPL OFFSET
29 503 021741     ORR      017600
30 504 140506     JMP      GON      ;PROCEED TO UPDATE POINTER
31 505 134011     1$: ASR      9.    ;FORWARD JUMP FOUND
32 506 005702     GON: ADD$ LPC  ;SECONDARY ENTRY
33 507 115702     SAC$    LPC
34 510 140162     JMP      MONITR  ;POSSIBLE RETURN POINT
35
36 511 160447     ; MOVEE: JST FIELDS
37 512 106000     SIOSI   6000
38 513 140162     JMP      MONITR
39
40 514 161612     ; JRLOAD: JST DTIM
41 515 061747     LIO     0377
42 516 136520     EXI     MASK  ;MASK OUT CLOCK INTERRUPTS
43 517 133000     HLT
44
45 520 040434     ; SETUP: LAC FIEL1 ;MIN MUST BE 175702
46 521 115736     SAC$    MINIC  ;PRESET MIN COUNT
47 522 045733     LAC$    FLAG
48 523 011725     AND      047777 ;RESET SUPER & FLARE BIT
49 524 115733     SAC$    FLAG
50 525 132000     CAC
51 526 115743     SAC$    MAXIC  ;PRESET MAX COUNT
52 527 115746     SAC$    MAXBV  ;BLUE VELOC.
53 530 115752     SAC$    MAXRV  ;RED VELOC.
54 531 115756     SAC$    OFFTTL ;SERVO OFFSET TOTAL
55
56 532 065704     LIOS   0L      ;GET SEQUENCE CONTROL WORD
57 533 137002     LSA     2

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```

1
2 534 115647 SAC% DLYP ;CLEAR DEVICE DELAYS
3 535 115650 SAC% SPC1
4
5 ;SETUP COUNTERS AND INDEXES:
6 536 115774 SAC% COUNT ;ASCENDING SEQUENCE INDEX
7 537 115766 SAC% SCOUNT ;ASCENDING CONTRL SLOT COUNTER
8
9 540 115765 CTRLST: SAC% DIRECT ;ASCENDING CONTROL WRD INDEX
10 541 105775 SIO% CHAN ;RESET SCRATCH SEQ. WORD
11 542 045775 LAC% CHAN ;TEST IF CONTROL BLOCK INDEX...
12
13 543 011715 CTRLCHK: AND THREE ;MATCHES SEQUENCE INDEX
14 544 115763 SAC% OLD ;COMPARISON BITS
15 545 055763 LAN% OLD
16 546 005774 ADD% COUNT
17 547 130000 TAZ ;SKIP WHEN NOT MATCHED
18 550 140573 JMP CTLFND ;INDEX MATCHED...TRANSFER JST INSTR.
19 551 055765 LAN% DIRECT
20 552 001714 ADD TWO
21 553 130400 TAN ;TEST IF CHECK IS COMPLETE ( )2)
22 554 140557 JMP CTLNXT ;NOT YET SO MOVE POINTERS
23 555 171425 JST UERROR ;SEND ERROR MESSAGE TO TM
24 556 140162 JMP MONITR ;HOLD TEMPORARY INSTRUCTION
25
26 557 045765 CTLNXT: LAC% DIRECT ;INCREMENT CONTROL WORD TABLE INDEX
27 560 001713 ADD ONE
28 561 115765 SAC% DIRECT
29 562 045775 LAC% CHAN ;SHIFT SCRATCH SEQUENCE WORD
30 563 134003 ASR 3 ;RIGHT BY 3
31 564 115775 SAC% CHAN
32 565 140543 JMP CTLCHK
33
34 566 045774 CTLMOR: LAC% COUNT ;INCREMENT SEQUENCE INDEX
35 567 001713 ADD ONE
36 570 115774 SAC% COUNT
37 571 132000 CAC ;CLEAR TABLE INDEX
38 572 140540 JMP CTRLST ;RESET CTRL WD INDX * SCRATCH SEQ WD
39
40 573 075765 CTLFND: LIN% DIRECT
41 574 042634 LACI CTLINS ;LOAD INDEXED CONTROL JUMP
42 575 075766 LIN% SCOUNT
43 576 117640 SAC%I CTLSLT ;SET PROPER CONTROL SLOT
44 577 137401 BIN ;INCREMENT SLOT COUNTER
45 600 125766 SIN% SCOUNT
46 601 045775 LAC% CHAN ;TEST IF CALIBRATION RUNS
47 602 135003 RSR 3 ;...AFTER LOOP JUST SET
48 603 130400 TAN
49 604 140611 JMP NOTCAL ;CAL. BIT NOT SET
50 605 040640 LAC CTLCAL ;CAL. FLAG SET
51 606 117640 SAC%I CTLSLT
52 607 137401 BIN ;INCREMENT SLOT COUNTER
53 610 125766 SIN% SCOUNT ;UPDATE LOCATION
54
55 611 055774 NOTCAL: LAN% COUNT ;* TEST IF END OF PROGRAMMABLE SLOTS
56 612 001714 ADD TWO
57 613 130400 TAN ; WAS IT ) 2 ?

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1  614 140566      JMP      CTLMOR      ;NO
2  615 040641      LAC      CTLTAC      ;YES... SET TACH SERVO CALL
3  616 117640      SACI%    CTLSLT
4  617 137401      BIN
5  620 125766      SIN%     SCOUNT
6  621 055766      LAN%     SCOUNT      ;TEST IF ALL SLOTS ARE FILLED (05)
7  622 001743      ADD      DS
8  623 130400      TAN
9  624 140630      JMP      DFALT      ;NOT DONE FILLING CONTROL SLOTS
10 625 040642      FILSLT: LAC CTLOTR      ;*SET OUTER LOOP CONTROL CALL
11 626 117640      SAC%I    CTLSLT
12 627 140643      JMP      SETOTS     ;DONE SETTING CONTROL SLOTS
13
14 630 040640      DFALT: LAC CTLCAL      ;DEFAULT CALIBRATION
15 631 117640      SAC%I    CTLSLT
16 632 137401      BIN
17 633 140625      JMP      FILSLT
18
19 634 161434      CTLINS: JST DWSTP      ;WAVELENGTH CONTROL LOOP CALL
20 635 161351      JST      RYCTRL      ;Y RASTER CONTROL LOOP CALL
21 636 161400      JST      RXCTRL      ;X RASTER CONTROL LOOP CALL
22 637 161575      JST      PRCTRL      ;POLARIMETER CONTROL LOOP CALL
23
24 640 161476      CTLCAL: JST CCALIB
25 641 161160      CTLTAC: JST TSCTRL
26 642 141304      CTLOTR: JMP COUNTER
27
28 643 137003      SETOTS: LSA 3      ; OUTER CONTROL LOOP SETUP
29 644 132000      CAC
30 645 115770      SAC%     REPEAT      ; INIT REPEAT COUNT
31 646 045710      LAC%     OL+4
32 647 134002      ASR      2
33 650 137002      LSA      2
34 651 130000      TAZ
35 652 001713      ADD      ONE      ; ALLOW 0 AS REPEAT PARAM.
36 653 115703      SAC%     PASCNT
37
38 654 137003      CALIB: LSA 3      ;CALIB. INTERVAL COUNTER SETUP
39 655 045704      LAC%     OL
40 656 134014      ASR      12.
41 657 137002      LSA      2
42 660 170250      JST      EXPON2
43 661 001722      ADD      M1
44 662 115754      SAC%     CCYCSZ      ; 01 DISABLES CAL
45 663 115756      SAC%     CALCYC
46
47 664 137003      LSA      3
48 665 045705      LAC%     OL+1      ;TEST IF NEW EXPMT NO. NEEDED
49 666 134007      ASR      7
50 667 011713      AND      ONE      ;IF BIT SET, INCR EXPMT.
51 670 025677      ORR%     EXPADV
52 671 115677      SAC%     EXPADV      ;THIS IS ADDED TO EXP. NO.
53
54
55 672 137002      SETUPW: LSA 2      ;WAVELENGTH DRIVE SETUP
56 673 045755      LAC%     CALLAM      ;TEST IF ANY CALIB OFFSET REMAINING
57 674 130000      TAZ

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1  675 140677      JMP      .+2      ;NONE REMAINING
2  676 161476      JST      CCALIB   ;JUST GET RID OF IT AND PROCEED
3
4  677 137003      LSA      3
5  700 132000      CAC
6  701 115763      SAC%    WLSCAN   ;INIT SCAN NO.
7  702 045706      LAC%    OL+2     ;LD OFFSET FLG + WLD.INCR
8  703 065707      LIO%    OL+3     ;LOAD OFFSET OR LSB
9  704 130400      TAN      ;NEG MEANS OFFSET FROM LAST MAX
10 705 140741      JMP      3%      ;POSITION COUNT SPECIFIED
11      SETUP  OFFSET FROM LAST LAMBDA MAX VALUE
12 706 132000      CAC
13 707 137002      LSA      2
14 710 115771      SAC%    MT3
15 711 105772      SIO%    MT4
16 712 137003      LSA      3
17 713 045710      LAC%    OL+4     ;LOAD AND TEST BIT 0 OF OL+4 (WLD MSB)
18 714 135001      RSR      1
19 715 130400      TAN
20 716 140722      JMP      1%      ;GLOBAL WHEN = 0
21 717 045741      LAC%    MAXIWH   ;USE LOCAL VALUE WHEN = 1
22 720 065742      LIO%    MAXIWL
23 721 140724      JMP      2%
24 722 045721      1%:    LAC%    MAXHI
25 723 065722      LIO%    MAXLO   ;USE LAMDA OFFSET BASE VALUE
26 724 137002      2%:    LSA      2
27 725 115767      SAC%    MT1
28 726 105770      SIO%    MT2
29 727 045772      LAC%    MT4     ;LOAD LIST OFFSET LSB
30 730 130400      TAN
31 731 140734      JMP      .+3     ;POSITIVE OFFSET SPECIFIED
32 732 051713      LAN      ONE     ;NEGATIVE OFFSET IS SPECIFIED
33 733 115771      SAC%    MT3     ;SET HIGH ORDER NEGATIVE
34 734 165542      JST      MDPA
35 735 137002      LSA      2
36 736 065772      LIO%    MT4     ;LOAD RESULTANT WLD STEP
37 737 045771      LAC%    MT3
38 740 140744      JMP      .+4
39
40 741 045710      3%:    LAC%    OL+4     ;DIRECT POSITION SPECIFIED
41 742 137002      LSA      2
42 743 011715      AND      THREE
43 744 105744      SIO%    DWLDLO
44 745 115743      SAC%    DWLDHI
45 746 137003      LSA      3
46 747 045711      LAC%    OL+5     ;SET WL INCREM
47 750 134006      ASR      6
48 751 011745      AND      037
49 752 137002      LSA      2
50 753 115740      SAC%    DELTAL   ;SET DELTA LAMBDA
51 754 164274      JST      DNPOS   ;SET WAVELENGTH DRIVE TO NEW POSN
52
53 755 137003      LSA      3
54 756 045706      LAC%    OL+2     ;NUMBER OF INCREMENTS
55 757 065711      LIO%    OL+5     ;POLR STEPS / WLD STEP SIZE / PL STEP / TACH
56 760 137002      LSA      2
57 761 015636      AND%    077777

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1	762	130000		TAZ	
2	763	041722		LAC	M1 ;DISABLE WLD LOOP
3	764	115741		SAC\$	WLCYSZ ;SET # OF INCREMENTS
4	765	115742		SAC\$	NLCYC
5	766	105774		SIO\$	COUNT
6					
7	767	045774		LAC\$	COUNT ;SET TACH SERVO INTERVAL
8	770	011720		AND	SEVEN
9	771	130000		TAZ	
10	772	140774		JMP	4\$;ZERO SPECIFIED TURNS OFF SERVO
11	773	170250		JST	EXPON2 ;RETURNS VALUE OF NONZERO EXPON. OF 2
12					;CHOICES--128, 64, 32, 16, 8, 4, 2
13	774	115736	4\$:	SAC\$	TCYCSZ
14	775	115737		SAC\$	TCYC
15					
16	776	045774		LAC\$	COUNT ;SET POLR INCREM SIZE
17	777	134003		ASR	3
18	1000	011720		AND	SEVEN
19	1001	001713		ADD	ONE ;ADD ONE TO PRODUCE RANGE OF 1 - 8
20	1002	115734		SAC\$	QSTPSZ ;STEP SIZE IS ALWAYS SET .GT. 0
21	1003	045774		LAC\$	COUNT ;SET POLR STEP QUAN
22	1004	134013		ASR	11.
23	1005	130000		TAZ	;IF ZERO SPECIFIED, SET TO NEG. WHICH
24	1006	001722		ADD	M1 ;DISABLES LOOP CONTROL FUNCTION
25	1007	115732		SAC\$	QCYSZ ;STEP QUAN. IS ALSO LOOP SIZE
26	1010	115733		SAC\$	PRCYCL
27					
28	1011	137003		LSA	3 ;SETUP OF MOSTLY RASTER ITEMS
29	1012	132000		CAC	
30	1013	115766		SAC\$	CXCYCL ;SET X CYCLE COUNTER
31	1014	115767		SAC\$	CYCYCL ;SET Y CYCLE COUNTER
32	1015	115765		SAC\$	CPRCYC ;SET POLR CYCLE COUNTER
33					
34	1016	045712		LAC\$	OL+6 ; RASTER X STEP QUAN / RASTER Y STEP QUAN
35	1017	065713		LIO\$	OL+7 ; XSTEP SIZE / Y STEP SIZE / CAL WLD OFFSET SIZE
36	1020	137002		LSA	2
37	1021	115774		SAC\$	COUNT ;SET RAS Y STEP QUAN
38	1022	011747		AND	0377
39	1023	130000		TAZ	;TEST IF DISABLED
40	1024	001722		ADD	M1 ;SET DISABLED FLAG
41	1025	115731		SAC\$	RYCYCL
42	1026	115730		SAC\$	YCYCSZ
43					
44	1027	045774		LAC\$	COUNT ;SET RAS X STEP QUAN
45	1030	134010		ASR	8.
46	1031	011747		AND	0377
47	1032	130000		TAZ	;TEST IF DISABLED
48	1033	001722		ADD	M1 ;SET DISABLED FLAG
49	1034	115727		SAC\$	RXCYCL
50	1035	115726		SAC\$	XCYSZ
51					
52	1036	105724		SIO\$	DELTX ;SET CALIBRATION STEP SIZE
53	1037	045724		LAC\$	DELTX ; FROM OL+7
54	1040	011746		AND	077
55	1041	115753		SAC\$	CALSIZ
56					
57	1042	137003		LSA	3

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1 1043 065717      LIO$   RAYCEN
2 1044 137002      LSA   2
3 1045 105765      SIO$   DIRECT ; CENTER OF RASTER
4
5 1046 045730      LAC$   YCYCSZ
6 1047 130400      TAN           ; TEST IF DISABLED
7 1050 141052      JMP    5$
8 1051 141061      JMP    6$ ; SETS DISABLED POSITION
9
10 1052 001722 5$:   ADD    M1
11 1053 115773      SAC$   MTS ; NOT DISABLED
12 1054 045724      LAC$   DELTX ; SET RAS Y STEP SIZE
13 1055 134006      ASR    6 ; AND INITIAL DIRECTION
14 1056 011745      AND    037 ; PASS Y STEP SIZE
15
16 1057 071713      LIN    ONE ; INDEX FOR Y RASTER
17 1060 171054      JST    RASLMC
18
19 1061 105717 6$:   SIO$   POSNY
20 1062 136614      EXI    RYLD
21 1063 045634      LAC$   03777 ; FOR 98 MS WAIT
22 1064 170360      JST    UWTX
23
24 1065 137003  RAXPOS: LSA   3
25 1066 065716      LIO$   RAXCEN ; POINTING POSN
26 1067 137002      LSA   2
27 1070 105765      SIO$   DIRECT ; USED LAYER IN CALC
28
29 1071 045726      LAC$   XCYCSZ ; SETUP FOR CALCULATING X LIMITS
30 1072 130400      TAN           ; ... IF NEEDED
31 1073 141075      JMP    1$
32 1074 141103      JMP    2$ ; UNNEEDED, DISABLED
33
34 1075 001722 1$:   ADD    M1
35 1076 115773      SAC$   MTS ; PASS STEP QUAN. FOR LIM CHECKING
36
37 1077 045724      LAC$   DELTX ; UNPACK X STEP QUAN
38 1100 134013      ASR    11. ; PASS STEP SIZE IN AC
39
40 1101 071712      LIN    ZERO ; X VARIABLE INDEX
41 1102 171054      JST    RASLMC ; CALCULATE LIMITS
42 1103 105716 2$:   SIO$   POSNX
43 1104 136610      EXI    RXLD
44 1105 045634      LAC$   03777
45 1106 170360      JST    UWTX ; 98 MS WAIT
46
47 1107 170216      JST    DETLIM ; FIND MAX ALLOWED CTR VAL.
48
49 ; DETECTOR SETUP FOLLOWS
50 1110 137003      LSA   3
51 1111 045705      LAC$   0L+1 ; PICK FOR HI BYTE
52 1112 137002      LSA   2
53 1113 134014      ASR    12. ; GET HI 4 BITS OF O.L. WORD
54 1114 115774      SAC$   MT6 ; TRANSFER AS INDEX ARG.
55 1115 134403      ASL    3 ; SET BITS FOR LO HALF OF ACONF6
56 1116 021713      ORR    ONE ; SET INTERVAL = 1
57 1117 170204      JST    DETPW ; GET DETS POWERED UP

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1
2 1120 045705 LAC$ OL+1
3 1121 137002 LSA 2
4 1122 134404 ASL 4 ; PICK LO 4 BITS OF MS 8
5 1123 134014 ASR 12.
6 1124 115774 SAC$ MT6
7 1125 130000 TAZ ;TEST FOR NON-ZERO INTERVAL 2
8 1126 141132 JMP 4$ ;NO INTVL 2
9 1127 134403 ASL 3 ;SET INTERVAL 2 PART OF ACONFG
10 1130 021714 ORR TWO ;SET INTERVAL = 2
11 1131 134410 ASL 8. ;POSITION IN MS HALF
12 1132 137003 4$: LSA 3
13 1133 025771 ORR$ ACONFG
14 1134 170204 JST DETPW
15 ;BMSG$: SEND MESSAGES AND RUN EXPMT.
16 1135 061751 LIO MSG1 ;SEND 'BEGIN EXPERIMENT' MSG
17 1136 171353 JST USMSG
18
19 1137 137003 LSA 3
20 1140 045677 LAC$ EXPADV ; TEST IF ADVANCING EXPT NO
21 1141 130000 TAZ
22 1142 141324 JMP REPM$G ;NO EDB OUTPUT FOR REPEATS
23 1143 005715 ADD$ EXPNUM ;INCREMENT EXPERIMENT SEQ NO.
24 1144 115715 SAC$ EXPNUM
25
26 1145 045705 LAC$ OL+1 ;CONDITIONAL STOP EXPT. INCREM
27 1146 134007 ASR 7 ;SHORTCUT ASSUMES EXPADV IS 0 OR 1
28 1147 015677 AND$ EXPADV ;BIT 1 (LSB) AFFECTED
29 1150 115677 SAC$ EXPADV ;CLEAR SERIAL NO. INCREMENT
30
31 1151 041750 LAC EDBLO ;EXPERIMENT DEFINITION BLOCK LOW ADDR
32 1152 137002 LSA 2
33 1153 115701 SAC$ AB
34 1154 001735 ADD D16 ;SIZE OF EDB DUMP...16.0 WORDS
35 1155 115700 SAC$ LAST
36 1156 171130 JST DBKDMP ;PERFORM EXP. DEF. BLK. DUMP
37 1157 141324 JMP REPM$G
38
39
40 ;
41 ; **** TACHOGRAM SERVO ROUTINE ****
42 ; (LAST MODIFIED ON JULY 2, 1980)
43 1160 141160 TSCTRL: JMP ;TACH SERVO CONTROL
44 1161 137002 LSA 2 ;USES COMPENSATED COUNT DATA
45 1162 045737 LAC$ TCYC
46 1163 130000 TAZ ;INACTIVE LOOP TEST
47 1164 141160 JMP TSCTRL ;NO SERVO SO RETURN
48
49 1165 001722 ADD M1
50 1166 130000 TAZ ;TEST IF 1-->0 TRANSITION
51 1167 141172 JMP 1$ ;YES
52 1170 115737 SAC$ TCYC ;UPDATE COUNTER
53 1171 141160 JMP TSCTRL ;RETURN
54
55 1172 045736 1$: LAC$ TCYCSZ ;SET INTERVAL COUNT BACK
56 1173 115737 SAC$ TCYC
57 1174 045675 LAC$ NTACH
58 1175 130000 TAZ

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1 1176 141160      JMP    TSCTRL  ;PREVENT DIVISIONS BY ZERO !
2 1177 115773      SAC$   MT5     ;SET DIVISOR FOR LATER NORMALIZATIONS
3
4 1200 045714      LAC$   TOTRDH  ;TRANSFER TOTALS OF RED COUNTS
5 1201 065715      LIO$   TOTRDL
6 1202 115771      SAC$   MT3
7 1203 105772      SIO$   MT4
8
9 1204 045712      LAC$   TOTBLH  ;LOAD BLUE ACCUMULATION
10 1205 065713      LIO$   TOTBLL
11 1206 115767      SAC$   MT1
12 1207 105770      SIO$   MT2
13
14 1210 165542      JST    MDPA    ;SUM THE BLUE AND RED INTENSITIES(IN MT3-4)
15 1211 165665      JST    DIVIDE  ;NORMALIZE AS THE ARITH MEAN
16 1212 045772      LAC$   MT4     ;PREVENTS OVERFLOWS IN LATER CALCULATIONS
17 1213 115765      SAC$   DIRECT  ;SAVE NORMALIZED INTENSITY-- (R+B)/N
18
19 1214 045714      LAC$   TOTRDH
20 1215 065715      LIO$   TOTRDL
21 1216 115771      SAC$   MT3
22 1217 105772      SIO$   MT4
23
24 1220 165572      JST    MDPS    ;FIND DIFFERENCE FOR NUMERATOR(IN MT3-4)
25 1221 045675      LAC$   NTACH   ;FORM ARITHMETIC MEAN OF RED-BLUE
26 1222 115773      SAC$   MT5
27 1223 165665      JST    DIVIDE  ; (B-R)/N
28
29 1224 137003      LSA    3       ;SERVO GAIN (TO BE SET BY GRND COMMAND)
30 1225 045777      LAC$   BSYGN   ;LOAD TACH SERVO GAIN VARIABLE
31 1226 011721      AND    01777   ;MASK OFF BUSY BIT IN SIGN
32 1227 137002      LSA    2
33 1230 115773      SAC$   MT5
34
35 1231 165634      JST    MLTPLY  ;MULTIPLY BY GAIN VALUE-- ((B-R)/N)*GAIN
36                ; CORRECTION: ((B-R)/N)*GAIN / ((R+B)/N)
37 1232 045765      LAC$   DIRECT
38 1233 115773      SAC$   MT5
39 1234 165665      JST    DIVIDE  ;FIND CORRECTION--INTENSITY NORMALIZED
40 1235 061722      LIO    M1
41 1236 045772      LAC$   MT4
42 1237 130400      TAN
43 1240 141247      JMP    45
44 1241 001717      ADD    SIX     ;TEST IF LT -6
45 1242 130400      TAN
46 1243 141256      JMP    55
47 1244 051717      LAN    SIX
48 1245 115772      SAC$   MT4
49 1246 141256      JMP    55
50
51 1247 061712      LIO    ZERO    ;TEST IF GT 6
52 1250 055772      LAN$   MT4
53 1251 001717      ADD    SIX
54 1252 130400      TAN
55 1253 141256      JMP    55
56 1254 041717      LAC    SIX
57 1255 115772      SAC$   MT4

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1 1256 105771 51: SIO% MT3 ;SET HIGH BITS
2 1257 045772 LAC% MT4
3 1260 137003 LSA 3
4 1261 005756 ADD% OFFTTL ;UPDATE ACCUMULATION OF SERVO CORRECTION
5 1262 115756 SAC% OFFTTL ;STORE IN DMA LOCN
6
7 1263 061752 LIO MSG12 ;SEND CORRECTION MESSAGE
8 1264 171353 JST USMSG
9 1265 065772 LIO% MT4 ;NO SEGMENT CORRECTION REQ'D
10 1266 171353 JST USMSG
11
12 1267 132000 CAC
13 1270 115714 SAC% TOTRDH ; CLEAR SUM OF COUNTER DATA
14 1271 115715 SAC% TOTRDL
15 1272 115712 SAC% TOTBLH
16 1273 115713 SAC% TOTELL
17 1274 115675 SAC% NTACH ;RESET N OF SUMMATION
18
19 1275 045743 LAC% DWLDHI ;ADJUST LAMBDA SCAN RESET VALUE
20 1276 115767 SAC% MT1
21 1277 045744 LAC% DWLDLO
22 1300 115770 SAC% MTR
23 1301 165542 JST MDPA
24
25 1302 161467 ; *** JST MDWPS ;IMMEDIATE WLD ADJUSTMENT
26 1303 141160 JMP TSCTRL ;RETURN TO NEXT CONTROL SLOT
27
28 ; **** OUTER CONTROL LOOP ****
29 ;
30 1304 137002 COUNTER: LSA 2
31 1305 035703 DM5% PASCNT
32 1306 141316 JMP CTR1
33 1307 061727 COUT: LIO MSG3
34 1310 171353 JST USMSG ;SEND END OF EXPMT MSG
35 1311 045650 LAC% SPC1 ; EXPMT COMPLETED
36 1312 130000 TAZ
37 1313 140162 JMP MONITR
38 1314 170433 JST DUALP ; DIGEST LAST DATA PAIR
39 1315 140162 JMP MONITR
40
41 1316 137003 CTR1: LSA 3
42 1317 132000 CAC
43 1320 115766 SAC% CXCYCL
44 1321 115767 SAC% CYCYCL ;CLEAR OUTER LOOP COUNTERS
45 1322 115763 SAC% WLSCAN
46 1323 115765 SAC% CPRCYC
47
48 1324 137003 REPM5G: LSA 3
49 1325 045770 LAC% REPEAT ;PRODUCE LOOP MESSAGE
50 1326 001713 ADD ONE
51 1327 115770 SAC% REPEAT
52 1330 021336 ORR MSG160 ;FORMAT IS 160000+ N WHERE
53 1331 171403 JST SNDMSG ; @ < N < [OL+4] AND OL+4 .LE. 2E13 -1
54 1332 055677 LAN% HEATWD
55 1333 130400 TAN
56 1334 171463 JST HEATR
57 1335 141756 JMP MEASUR ;INITIATE DATA INPUT

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1 1336 160000 MSG160: 160000
2
3
4
5 1337 045730 RYCFIN: LAC$ YCYCSZ
6 1340 115731 SAC$ RYCYCL
7 1341 171025 JST RAYUND :RESET TO TOP
8
9 1342 041713 LAC ONE
10 1343 137003 LSA 3 :SEND END OF Y SCAN MSG.
11 1344 005767 ADD$ CYCYCL
12 1345 115767 SAC$ CYCYCL
13 1346 011721 AND 01777 :MASK FOR MSG INTEGRITY
14 1347 021365 ORR MSG16Y
15 1350 171403 JST SNDMSG
16
17 1351 141351 RYCTRL: JMP
18 1352 137002 LSA 2
19 1353 045731 LAC$ RYCYCL : TEST IF LOOP IS ACTIVE
20 1354 130400 TAN
21 1355 141357 JMP 1$ :ACTIVE
22 1356 141351 JMP RYCTRL : DISABLED ... RETURN
23
24 1357 001722 1$: ADD M1
25 1360 130000 TAZ : :TEST IF LOOP CYCLE IS COMPLETE
26 1361 141337 JMP RYCFIN :COMPLETED
27
28 1362 115731 SAC$ RYCYCL
29 1363 171025 JST RAYUND
30 1364 141756 JMP MEASUR
31
32 1365 116000 MSG16Y: 116000
33
34
35
36 1366 045726 RXCFIN: LAC$ XCYCSZ :RXCTRL-11 THIS IS FINAL PART OF CONTROL PROG
37 1367 115727 SAC$ RXCYCL :RESET LOOP COUNTER
38 1370 171003 JST RAXUND :RESET X AXIS
39 1371 041713 LAC ONE
40 1372 137003 LSA 3
41 1373 005766 ADD$ CXCYCL
42 1374 115766 SAC$ CXCYCL :INCREMENT COUNTER
43 1375 011721 AND 01777
44 1376 021414 ORR MSG12X :PRODUCE MESSAGE
45 1377 171403 JST SNDMSG
46
47 1400 141400 RXCTRL: JMP
48 1401 137002 LSA 2
49 1402 045727 LAC$ RXCYCL
50 1403 130400 TAN :TEST IF CONTROL LOOP DISABLED
51 1404 141406 JMP 1$
52 1405 141400 JMP RXCTRL :DISABLED, SO RETURN
53
54 1406 001722 1$: ADD M1
55 1407 130000 TAZ
56 1410 141366 JMP RXCFIN
57

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1 1411 115727      SAC$   RXCYCL
2 1412 171003      JST    RAXUND  ; UNIDIRECTIONAL RASTER DRIVER
3
4 1413 141756      JMP    MEASUR
5
6 1414 112000      MSG12X: 112000
7
8                ;      **** WAVELENGTH CONTROL ****
9
10 1415 045741      DWSFIN: LAC$ WLCYSZ  ; DISPLACED ENTRY BELOW
11 1416 115742      SAC$   WLCYC  ; SET DEPLETED CYCLE COUNTER
12 1417 041713      LAC    ONE
13 1420 137003      LSA    3
14 1421 005763      ADD$   WLSCAN
15 1422 115763      SAC$   WLSCAN  ; INCREMENT SCAN COUNTER
16 1423 021466      ORR    MSG2    ; 120000 + N      N ( 2(13) - 1
17 1424 171403      JST    SNDRSG
18 1425 171047      JST    DLYTST  ; WAIT FOR POWER LIMIT DELAYS
19                ; ** RESET WLD TO START POINT (FLY BACK) ONLY IF THE WLD SCAN IS REPEATED
20 1425 137002      LSA    2
21 1427 055611      LAN$   ONE2
22 1430 005703      ADD$   PASCNT  ; LAST REPEAT OF SCANS?
23 1431 130000      TAZ
24 1432 141434      JMP    DWSTP   ; YES: SKIP RESET OF WLD
25 1433 164274      JST    DWPOS   ; NO: RESET WLD TO START POSN.
26                ; **
27
28                ; ENTRY POINT
29 1434 141434      DWSTP: JMP .
30 1435 137002      LSA    2
31 1436 045742      LAC$   WLCYC
32 1437 130400      TAN
33 1440 141442      JMP    15
34 1441 141434      JMP    DWSTP
35
36 1442 130000      15:   TAZ      ; TEST FOR END OF LINE
37 1443 141415      JMP    DWSFIN  ; REALLY WAS EOL
38
39 1444 001722      ADD    M1      ; DECREMENT CYCLE COUNT
40 1445 115742      SAC$   WLCYC
41
42 1446 171047      JST    DLYTST  ; WAIT FOR POWER DELAYS
43 1447 137003      LSA    3
44 1450 045761      LAC$   AWLDHI
45 1451 065762      LIO$   AWLDLO
46 1452 137002      LSA    2
47 1453 115767      SAC$   MT1
48 1454 105770      SIO$   MT2
49 1455 132000      CAC
50 1456 065740      LIO$   DELTAL
51 1457 115771      SAC$   MT3
52 1460 105772      SIO$   MT4
53 1461 165772      JST    MDPS    ; SUBTRACT DOUBLE PRECISION FOR REVERSING
54 1462 045771      LAC$   MT3
55 1463 065772      LIO$   MT4
56 1464 164216      JST    MOVLOW  ; MOVE DRIVE AND REDUCE POWER
57

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1 1465 141756 JMP MEASUR
2
3 1466 120000 MSG2: 120000
4 ; ***
5 1467 141467 MDWPS: JMP
6 1470 045771 LAC% MT3
7 1471 065772 LIO% MT4
8 1472 115743 SAC% DWLDHI
9 1473 105744 SIO% DWLDLO
10 1474 164274 JST DWPOS
11 1475 141467 JMP MDWPS
12 ; ***
13 ; ***** CALIBRATION CONTROL *****
14 ;
15 1476 141476 CCALIB: JMP
16 1477 137002 LSA 2
17 1500 045755 LAC% CALLAM :TEST OFFSET PRESENTLY SET
18 1501 130000 TAZ
19 1502 141523 JMP CLINT :FOUND NONE =0
20 1503 065755 LIO% CALLAM :FOUND SOME ... REMOVE IT
21 1504 132000 CAC
22 1505 115755 SAC% CALLAM :CLEAR OFFSET PRESENTLY SET
23 ;
24 1506 115771 SAC% MT3
25 1507 105772 SIO% MT4
26 1510 137003 LSA 3
27 1511 045761 LAC% AWLDHI
28 1512 065762 LIO% AWLDLO
29 1513 137002 LSA 2
30 1514 115767 SAC% MT1
31 1515 105770 SIO% MT2
32 1516 165572 JST MDPS :REMOVE BY SUBTRACTING
33 ; ***
34 1517 161467 JST MDWPS :MOVE WLD AND REDUCE STEP PWR
35 ; ***
36 1520 061557 LIO MSG7 :GET END OF CALIBRATION MESSAGE
37 1521 171353 JST USMSG
38 1522 141476 JMP CCALIB
39 ;
40 1523 045756 CLINT: LAC% CALCYC :GET CALIBRATION INTERVAL
41 1524 001722 ADD M1
42 1525 130400 TAN :TEST FOR ZERO OR NEG
43 1526 141530 JMP 1$
44 1527 141476 JMP CCALIB :NO CALIBRATION SETUP-- RETURN
45 ;
46 1530 130000 1$: TAZ :TEST INTERVAL COUNTER
47 1531 141534 JMP 2$ :CALIB. TIME IS NOW
48 1532 115756 SAC% CALCYC :INTERVAL NOT UP--DECREMENT
49 1533 141476 JMP CCALIB :RETURN
50 ;SET INTERVAL COUNTER TO REPEAT
51 1534 045754 2$: LAC% CCYCSZ :FOLLOWING A CALIBRATION CYCLE
52 1535 115756 SAC% CALCYC
53 1536 065753 LIO% CALSIZ
54 ;
55 1537 105755 SIO% CALLAM :SET PRESENT CALIBRATION OFFSET
56 1540 132000 CAC
57 1541 115771 SAC% MT3

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1 1542 105772      SIO$   MT4
2 1543 137003      LSA     3
3 1544 045761      LAC$   ANLDHI
4 1545 065762      LIO$   ANLDLO
5 1546 137002      LSA     2
6 1547 115767      SAC$   MT1
7 1550 105770      SIO$   MT2
8 1551 165342      JST    MDPA      ;ADD CALIBRATION OFFSET
9
10 1552 161467     ; ***      JST    MDWPS      ;MOVE WLD AND REDUCE STEP PWR
11
12 1553 061556     ; ***      LIO     MSG6
13 1554 171353      JST    USMSG
14 1555 141756      JMP    MEASUR    ;MAKE A CALIBRATION PASS
15
16 1556 155006     MSG6: 155006
17 1557 155007     MSG7: 155007
18
19                ;          **** POLARIMETER CONTROL ****
20
21 1560 045732     PRCFNS: LAC$ 0CYCSZ
22 1561 115733      SAC$   PRCYCL    ;ENTRY POINT IS DISPLACED BELOW
23 1562 137003      LSA     3
24 1563 041713      LAC     ONE
25 1564 005765      ADD$   CPRCYC
26 1565 115765      SAC$   CPRCYC
27 1566 137002      LSA     2
28 1567 115773      SAC$   MTS
29
30 1570 170706      JST    GETPRF    ;RESET TO REFERENCE
31 1571 045773      LAC$   MTS
32 1572 011721      AND    01777
33 1573 021611      ORR    MSG10P    ;SEND CYCLE MESSAGE
34 1574 171403      JST    SNDMSG    ;FORMAT 110000 + N
35
36 1575 141575     PRCTRL: JMP    ;POLARIMETER CONTROL ROUTINE
37 1576 137002      LSA     2
38 1577 045733      LAC$   PRCYCL
39 1600 130400      TAN    ;TEST FOR NULL CONTROL LOOP
40 1601 141603      JMP    1$
41 1602 141575      JMP    PRCTRL    ;RETURN FROM NULL CONTROL LOOP
42
43 1603 130000     1$:  TAZ    ;TEST FOR END OF LOOP
44 1604 141560      JMP    PRCFNS    ;END WAS FOUND
45
46 1605 001722      ADD    M1      ;DECREMENT CYCLE COUNT
47 1606 115733      SAC$   PRCYCL    ;UPDATE COUNTER LOCATION
48 1607 170734      JST    MOVPR    ;CALL FOR DEVICE STEP
49
50 1610 141756      JMP    MEASUR    ;RETURN TO DATA INPUT PROG.
51
52
53 1611 110000     MSG10P: 110000
54
55 1612 141612     DTIM: JMP    ;HALT GATES
56 1613 137003      LSA     3
57 1614 061713      LIO     ONE

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1 1615 136711      EXI    C2LD
2 1616 136715      EXI    C3LD
3 1617 045703      LAC$   OLS      ;CLEAR SETUP BIT
4 1620 011721      AND    01777    ;MASK OFF SIGN BIT
5 1621 115703      SAC$   OLS
6 1622 136432      EXI    JRC2
7 1623 136422      EXI    JRC3
8 1624 136564      EXI    D1F
9 1625 136566      EXI    D2F
10 1626 136570      EXI    D3F
11 1627 136572      EXI    D4F
12 1630 136576      EXI    D5F
13 1631 141612      JMP    DTIM
14
15 1632 141632      DEVOFF: JMP
16 1633 136762      EXI    ORF      ; PLATE OFF
17 1634 137003      LSA    3
18 1635 061713      LIO    ONE
19 1636 105772      SIO$   RPWD    ;RASTER; WLD OFF
20 1637 136600      EXI    RPLD
21 1640 141632      JMP    DEVOFF
22
23 1641 141641      BSYCLR: JMP      ; SET BUSY TO + SO LEVEL 2 INTRPTS
24 1642 137003      LSA    3        ; CAN BE SERVICED
25 1643 045777      LAC$   BSYGN   ; BUSY BIT AND T5 GAIN ARE COMBINED
26 1644 011721      AND    01777   ; LIMITS SIZE OF GAIN
27 1645 115777      SAC$   BSYGN
28 1646 141641      JMP    BSYCLR
29
30
31 1647 136660      ; OBC 'OK' HAS TRIPPED...SHUT DOWN NOW
32 1650 136461      CLOSE: EXI    H1F ;HTR 1 OFF, & CLOSE DOOR
33 1651 041747      EXI    DORN    ; DOOR FWR ON
34 1652 170360      LAC    0377    ; (12 MSEC WAIT)
35 1653 136463      JST    UWTX
36 1654 061663      EXI    DORC   ;CLOSE DOOR COMMAND
37 1655 170051      LIO    013
38 1656 134407      JST    USMR   ;WAIT TILL DOOR IS CLOSED
39 1657 130400      ASL    7
40 1660 141654      TAN
41 1661 136460      JMP    4$
42 1662 133000      EXI    DORF   ; THEN DOOR POWER OFF,
43 1663 000013      HLT    ; AND HALT JR.
44
45 1664 160447      Q13: 13
46 1665 056000      ; ***
47 1666 005776      COMPAA: JST   FIELDS ;PRELIM INSTRUCTION PARSE
48 1667 130400      LAN$I  6000    ;LOAD NEGATIVE 'B' DATA
49 1670 141672      ADD$   TEMPJ   ;ADD 'A' DATA; TEST 2-A
50 1671 140162      COMT: TAN
51 1672 041713      JMP    .+2    ;POSITIVE OR 0, SO B .GE. A
52 1673 140506      JMP    MONTR  ;NEGATIVE, SO B .LT. A
53
54 1674 160447      LAC    ONE    ; ***
55 1675 045776      JMP    GON    ;USE SECONDARY ENTRY POINT
56 1676 131000      ;
57 1677 141700      ADDD: JST   FIELDS ;PRELIMINARY INSTRUCTION PARSE
58 1678 045776      LAC$   TEMPJ   ;LOAD 'A' DATA
59 1679 131000      TOF    ;CLEAR OVFL LATCH
60 1680 141700      JMP    .+1

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1 1756 164022 MEASUR: JST CLLCT
2 1757 170631 JST CKINTN
3 1760 137002 LSA 2
4 1761 045654 LACS BOTH ;TEST FOR DUAL COUNTER DATA
5 1762 130400 TAN ;SKIP IF DUAL
6 1763 141774 JMP TANDM
7 1764 045761 LACS TM2WD
8 1765 115650 SACS SPC1
9 1766 045762 LACS TM3WD
10 1767 115651 SACS SPC2
11 1770 045716 LACS POSNX
12 1771 115652 SACS SPX
13 1772 045717 LACS POSNY
14 1773 115653 SACS SPY
15 1774 137003 TANDM: LSA 3
16 1775 045771 LACS ACONFG ;TEST IF 2 INTERVALS
17 1776 135406 RSL 6 ;INTRVL 1 BIT:2 TO SIGN POSN
18 1777 130400 TAN ;SKIP IF USING 2 INTRVLS
19 2000 144010 JMP FLRTST ;ONLY ONE
20 2001 135402 RSL 2 ;COMPLETE CFG. BYTE SWAP
21 2002 115771 SACS ACONFG ;SET STATUS WORD
22 2003 164022 JST CLLCT ;TAKE OTHER DATA AND OUTPUT
23 2004 137003 LSA 3
24 2005 045771 LACS ACONFG ;REPLACE SWAPPED BYTES OF CFG.
25 2006 135010 RSR B.
26 2007 115771 SACS ACONFG
27 2010 045733 FLRTST: LACS FLAG
28 2011 130400 TAN
29 2012 151640 JMP CTLSLT ;EXIT TO CONTROL SLOT PGM.
30 2013 134401 ASL 1
31 2014 134016 ASR 14. ;MASK FOR TEST
32 2015 130000 TAZ ;SKIP IF SUPER OR IMMED ARE SET
33 2016 151640 JMP CTLSLT
34 2017 141307 JMP COUT ;FLARE BITS SET, END EXPMT.
35
36 2020 060241 COLFIN: LIO ZERO1
37 2021 136721 EXI 321 ;DET ROUTE CLEARED
38 2022 144022 CLLCT: JMP
39 2023 137003 LSA 3
40 2024 045714 LACS OL+8.
41 2025 065771 LIOS ACONFG ;SET CLOCK CODE / GATE TIME AND ROUTING
42 2026 137002 LSA 2
43 2027 105765 SIOS DIRECT ;SAVE FOR TEMPR. USE
44 2030 135010 RSR B. ;MOVE TIME TO LOW HALF
45 2031 115775 SACS MT7
46 2032 065775 LIOS MT7
47 2033 136710 EXI T2LD ;SET GATE 2 COUNT
48 2034 136714 EXI T3LD ;SET GATE 3 SAME
49
50 2035 134405 ASL 5
51 2036 134015 ASR 13.
52 2037 115774 SACS MT6 ;GET CLOCK INDEX
53 2040 075774 LINS MT6 ;USE 3 BIT INDEX TO CLOCK TABLE
54 2041 062173 LIOI CLKTB L ;LOAD CLOCK SELECT WORD
55 2042 136711 EXI C2LD
56 2043 136715 EXI C3LD
57

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1 2044 132000      CAC    ;CLEAR INTRPT FLAGS
2 2045 115761      SAC$    TM2WD
3 2046 115762      SAC$    TM3WD
4
5 2047 045765      LAC$    DIRECT ;SET ROUTING OF DATA
6 2050 134411      ASL    9    ;MOVE DUAL FLAG BIT TO SIGN
7 2051 115654      SAC$    BOTH    ;SAVE SINGLE/DUAL DET. FLAG
8 2052 134014      ASR 12.
9 2053 115774      SAC$    MT6    ;USE 4 BIT INDEX TO ROUTING TABLE
10 2054 075774      LINS    MT6
11 2055 066153      LIOIS$ DETTBL ;LD 8 BIT ROUTING DATA
12 2056 136721      EXI    321    ;SET PATH OF DETECTOR ROUTING
13
14 2057 134414      ASL 12.
15 2060 130400      TAN            ;SKIP IF 2 DETS USED
16 2061 144063      JMP    1$      ;ONLY ONE SPECIFIED
17 2062 136716      EXI    G3I    ;INITIATE GATE 3
18 2063 136712    1$: EXI    G2I    ; GATE 2
19 2064 045650      LAC$    SPC1    ;TEST IF LAST DATA CHOMPED
20 2065 130000      TAZ
21 2066 144071      JMP    2$
22 2067 170433      JST    DUALP    ;NO
23 2070 137002      LSA 2
24 2071 045761    2$: LAC$    TM2WD    ;TEST FLAG CLEARED AT START
25 2072 130400      TAN            ;FLAG SET NEG BY INTRPT
26 2073 144071      JMP    2$
27 2074 136442      EXI    RTM2
28 2075 105761      SIO$    TM2WD
29 2076 045761      LAC$    TM2WD
30 2077 015636      AND$    077777
31 2100 115772      SAC$    MT4
32 2101 060252      LIO    076
33 2102 170051      JST    USMR    ;GET OVF2 STATUS
34 2103 135001      RSR    1      ;OVF2 TO MSB
35 2104 115773      SAC$    MTS
36 2105 130400      TAN
37 2106 144120      JMP    3$
38 2107 060200      LIO    PCTOV    ;SEND OVERFLOWED MESSAGE
39 2110 171275      JST    U2MSG    ;USE SAME DATA CHANNEL
40 2111 136713      EXI    TM2N    ;CLEAR OVF2 LATCH
41 2112 065772      LIO$    MT4
42 2113 171275      JST    U2MSG    ;SEND DATA ON BUS
43 2114 045761      LAC$    TM2WD    ;TEST TM2WD MSB
44 2115 130400      TAN
45 2116 164201      JST    DETFF    ;LARGE OVERFLOW, DET. OFF
46 2117 144122      JMP    4$
47
48 2120 065772    3$: LIO$    MT4
49 2121 171275      JST    U2MSG
50 2122 045761    4$: LAC$    TM2WD
51 2123 134001      ASR    1
52 2124 115761      SAC$    TM2WD    ;SCALE TO IGNORE SIGN BIT
53 2125 055676      LAN$    LTDCTS
54 2126 005761      ADD$    TM2WD
55 2127 130400      TAN
56 2130 164201      JST    DETFF    ;DETECTORS OFF
57 2131 045654      LAC$    BOTH    ;DUAL DETECTOR FLAG TEST

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1 2132 130400      TAN
2 2133 144020      JMP   COLFIN ;NOT EXPECTED SO RETURN
3
4 2134 045762 5%:   LAC%  TM3WD  ;TEST FLAG CLEARED AT START
5 2135 130400      TAN           ;FLAG SET NEG BY INTRPT
6 2136 144134      JMP           5%
7 2137 136443      EXI   RTM3
8 2140 105762      SIO%  TM3WD
9 2141 045762      LAC%  TM3WD  ;LEAVE COLLECTED DATA INTACT
10 2142 015636     AND%  077777
11 2143 115772     SAC%  MT4
12 2144 045773     LAC%  MT5   ;SAVED OFL STAT
13 2145 135001     RSR   1     ;OVFL 3 TO MSB
14 2146 130400      TAN
15 2147 144161     JMP   6%
16 2150 060200     LIO   PCTOV
17 2151 171323     JST   U3MSG ;SEND MSG ON DATA CHANNEL
18 2152 136717     EXI   TM3N  ;CLEAR OVFL3 LATCH
19 2153 065772     LIO%  MT4
20 2154 171323     JST   U3MSG
21 2155 045762     LAC%  TM3WD
22 2156 130400      TAN
23 2157 164201     JST   DETFF ;LARGE OVFL; DET. OFF
24 2158 144163     JMP   7%
25
26 2161 065772 6%:   LIO%  MT4
27 2162 171323     JST   U3MSG
28 2163 045762 7%:   LAC%  TM3WD
29 2164 134001     ASR   1
30 2165 115762     SAC%  TM3WD
31 2166 055676     LAN%  LTDCTS
32 2167 005762     ADD%  TM3WD
33 2170 130400      TAN
34 2171 164201     JST   DETFF ;COUNT EXCEEDED LIMIT
35 2172 144020      JMP   COLFIN ;RETURN
36
37
38 2173 000001     ONE1:  CLKTBL: 1.;   16 KHZ 62.5 MICROSEC
39 2174 000002     2.;   2 KHZ  500 MICROSEC
40 2175 000004     FOUR1: 4.;   125 HZ  8 MILLISEC
41 2176 000010     10.;  31.25 HZ  32 MILLISEC
42 2177 000020     20.;  7.8125 HZ 128 MILLISEC
43
44 2200 155550     PCTOV: 155550
45
46 2201 144201     DETFF: JMP   ;DETECTOR POWER OFF
47 2202 161612     JST   DTIM  ;USED IF COUNTS TOO HIGH
48 2203 045733     LAC%  FLAG  ;SET STATUS BIT
49 2204 020254     ORR   010K1
50 2205 115733     SAC%  FLAG
51 2206 137002     LSA   2
52 2207 144201     JMP   DETFF ;CONTINUE WITH EXPMT.
53
54 2210 144210     RDSIZ: JMP
55 2211 060243     LIO   D10
56 2212 170051     JST   USMR  ;SIZE NUMBER
57 2213 134412     ASL   10.

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1 2214 134013      ASR      11.
2 2215 144210      JMP      RDSIZ
3
4 2216 144216      MOVLOW: JMP      ; ARGS PASSED IN AC: IO REGS
5 2217 115763      SAC$    WANTHI  ; STORE WLD HI
6 2220 105764      SIG$    WANTLO
7 2221 164363      JST     MOVWLD  ; MOVE THE DRIVE
8
9 2222 137003      LSA     3
10 2223 040176     LAC     CLKTBL+3 ; =10 OCTAL
11 2224 132400     NOT
12 2225 015772     AND$    RPWD
13 2226 115772     SAC$    RPWD
14 2227 065772     LIO$    RPWD
15 2230 136600     EXI     RPLD   ; STEP POWER OFF
16 2231 144216     JMP     MOVLOW
17
18
19 2232 144232     RWENAB: JMP
20 2233 137003     LSA     3
21 2234 025772     ORR$    RPWD
22 2235 115772     SAC$    RPWD
23 2236 065772     LIO$    RPWD
24 2237 136600     EXI     RPLD
25 2240 144232     JMP     RWENAB ; USAGE REQUIRES SA=3 ON EXIT
26
27
28 2241 000000     ZERO1:  0
29 2242 000003     THREE1: 3
30 2243 000012     D10:    10.
31 2244 000026     D22:    22.
32 2245 000030     D30:    30
33 2246 000041     D41:    41
34 2247 000050     D40:    40.
35 2250 000060     D48:    48.
36 2251 000070     D56:    56.
37 2252 000076     D76:    76
38 2253 000321     D207:   207.
39 2254 010000     D10K1: 10000
40
41 2255 064000     LIODLR: LIO$  0
42
43      . =2274
44
45      ; WAVELENGTH DRIVE - MOVE DRIVE OR CHANGE POSITION COUNTER
46      ; ENTRY POINTS: DWPOS - MOVE DRIVE TO SPECIFIED POSN
47      ; DWSTP - MOVE DRIVE REV SPECIFIED NO OF STEPS
48      ; CONTROL LOOP CALL FROM CONTROL SEQUENCE
49      ; SMWLD - SET SM COUNTER = JR POSN COUNTER
50      ; ARGUMENTS: DWLD = DESIRED POSITION (2 WORDS)
51      ; DELTA = DESIRED CHANGE
52
53 2274 144274     DWPOS: JMP
54 2275 137002     LSA     2
55 2276 045740     LAC$    DELTA   ; IF DELTA = 0 AND AWLD=DWLD, DONT MOVE
56 2277 130000     TAZ
57 2300 144302     JMP     1%     ; YES - NOW SEE IF POSN OK

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1 2301 144315      JMP MOVIT      ; DELTAL NOT ZERO - MUST MOVE DRIVE
2 2302 055744 15:  LANS DWLDLO  ; CHECK LO 16 BITS FIRST
3 2303 137003      LSA 3
4 2304 005762      ADDS AWLDLO
5 2305 130000      TAZ           ; LO 16 BITS THE SAME ?
6 2306 144310      JMP ZS        ; YES - NOW CHECK HI 2 BITS
7 2307 144315      JMP MOVIT     ; POSN NOT RIGHT, MUST MOVE IT
8 2310 055761 25:  LANS AWLDHI
9 2311 137002      LSA 2
10 2312 005743     ADDS DWLDHI
11 2313 130000     TAZ           ; EQUAL ?
12 2314 144274     JMP DWPOS     ; DELTAL =0 AND POSN OK, RETURN
13                ; MOVE WLD TO DWLD, WITH FINAL MOTION
14 2315 051322 MOVIT: LANS SLACK  ; REVERSE AT LEAST 'SLACK' STEPS. IF
15 2316 137002     LSA 2
16 2317 005740     ADDS DELTAL  ; DELTAL > SLACK, MOVE DRIVE TO 'DELTAL'
17 2320 130400     TAN           ; STEPS FORWARD OF DWLD, THEN MOVE
18 2321 144341     JMP DELOK    ; REVERSE 'DELTAL' STEPS TO DWLD.
19 2322 045740     LACS DELTAL  ; IF DELTAL < SLACK, MOVE TO 'DELTAL +
20 2323 001322     ADD SLACK    ; SLACK' STEPS FORWARD OF DWLD, THEN
21 2324 115772     SACS MT4    ; REVERSE 'SLACK' STEPS, THEN FINALLY
22 2325 132000     CAC           ; REVERSE 'DELTAL' STEPS TO DWLD
23 2326 115771     SACS MT3
24 2327 045743     LACS DWLDHI  ; FIRST TEST DELTAL, IF .LT. SLACK,
25 2330 065744     LIO% DWLDLO  ; SET WANT = DWLD + DELTAL + SLACK
26 2331 115767     SACS MT1
27 2332 105770     SIO% MT2
28 2333 165542     JST MDPA
29 2334 045771     LACS MT3
30 2335 065772     LIO% MT4
31 2336 115763     SACS WANTHI
32 2337 105764     SIO% WANTLO
33 2340 164363     JST MOVWLD
34
35 2341 045743 DELOK: LACS DWLDHI  ; NOW SET WANT = DWLD + DELTAL
36 2342 065744     LIO% DWLDLO
37 2343 115767     SACS MT1
38 2344 105770     SIO% MT2
39 2345 132000     CAC
40 2346 065740     LIO% DELTAL
41 2347 115771     SACS MT3
42 2350 105772     SIO% MT4
43 2351 165542     JST MDPA
44 2352 045771     LACS MT3
45 2353 065772     LIO% MT4
46 2354 115763     SACS WANTHI  ; STORE IN WANT
47 2355 105764     SIO% WANTLO
48 2356 164363     JST MOVWLD  ; MOVE DRIVE TO 'WANT'
49
50 2357 045743 FINAL: LACS DWLDHI
51 2360 065744     LIO% DWLDLO  ; SET WANT = DWLD
52 2361 164216     JST MOVLOW  ; MOVE WLD, REDUCE POWER
53 2362 144274     JMP DWPOS    ; RETURN
54
55
56 2363 144363 MOVWLD: JMP      ; MOVE DRIVE FROM AWLD TO WANT
57 2364 137003     LSA 3      ; CALCULATE DOWN = AWLD - WANT

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1 2365 045761      LAC$ AWLDHI      ; OR WANT - AWLD, SO THAT DOWN IS
2 2366 065762      LIO$ AWLDLO      ; POSITIVE
3 2367 137002      LSA 2
4 2370 115767      SAC$ MT1
5 2371 105770      SIO$ MT2
6 2372 045763      LAC$ WANTHI
7 2373 065764      LIO$ WANTLO
8 2374 115771      SAC$ MT3
9 2375 105772      SIO$ MT4
10 2376 165572     JST MDPS
11 2377 050173     LAN ONE1        ; DIRECTION WAS ASSUMED REVERSE
12 2400 115765     SAC$ DIRECT
13 2401 045772     LAC$ MT4        ; SEE IF DOWN IS ZERO
14 2402 025771     ORR$ MT3        ; INCLUDE HI PART
15 2403 130000     TAZ
16 2404 144566     JMP NOMOV       ; YES - EXIT MOVWLD
17
18 2405 137003     RUNWLD: LSA 3
19 2406 045772     LAC$ RPWD
20 2407 010176     AND CLKTB+3    ; = 010
21 2410 134002     ASR 2
22 2411 115776     SAC$ TEMPJ     ; CONDITIONAL JUMP INDEX
23 2412 075776     LINS TEMPJ
24 2413 040176     LAC CLKTB+3    ; = 010
25 2414 164232     JST RWENAB     ; STEP PWR ON
26
27 2415 146416     JMPI 1$
28 2416 041475     1$: LAC UND420    ; CONDITIONAL EXECUTION
29 2417 170360     JST UWTX       ; WAIT 20 MS
30 2420 137002     LSA 2
31 2421 045771     LAC$ MT3       ; NOW SEE IF DOWN IS +
32 2422 130400     TAN
33 2423 144427     JMP 2$        ; YES - DIRECTION IS REV
34 2424 040173     LAC ONE1      ; FWD, NEGATE DOWN
35 2425 115765     SAC$ DIRECT
36 2426 165625     JST MDFNOT
37 2427 045771     2$: LAC$ MT3    ; STORE DOWN
38 2430 065772     LIO$ MT4
39 2431 115711     SAC$ DOWNHI
40 2432 105776     SIO$ DOWNLO
41 2433 045765     LAC$ DIRECT    ; LOAD STEP COMMANDS ACCORDING TO DIREC
42 2434 130400     TAN
43 2435 144441     JMP 3$        ; FWD
44 2436 136642     EXI SLWR      ; SET REVERSE
45 2437 051370     LAN STEPR
46 2440 144443     JMP 4$
47 2441 136641     3$: EXI SLWF   ; SET FORWARD
48 2442 051367     LAN STEPF
49 2443 164607     4$: JST LDCMD
50 2444 137003     LSA 3        ; MOVE AWLD TO MT3/4
51 2445 045761     LAC$ AWLDHI
52 2446 065762     LIO$ AWLDLO
53 2447 137002     LSA 2
54 2450 115771     SAC$ MT3
55 2451 105772     SIO$ MT4
56 2452 065765     LIO$ DIRECT    ; MOVE DIRECT TO MT1/2 FOR UPDATING
57 2453 045765     LAC$ DIRECT    ; AWLD AFTER EACH STEP

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1 2454 130400      TAN          ; IF LO BYTE .LT. 0, HI BYTE MUST BE
2 2455 132000      CAC          ; -1, AND ZERO IF LO BYTE .GT. 0
3 2456 115767      SAC$ MT1
4 2457 105770      SIO$ MT2
5 2460 070241      LIN ZERO1    ; INDEX INITIAL FREQUENCY
6 2461 136403      EXI FANS     ; DISABLE L3 INTERRUPTS
7 2462 063323      LIOI RGTG
8 2463 136710      EXI T2LD
9 2464 136714      EXI T3LD
10 2465 061364      LIO GTGWAY   ; .0625 MS AUTO STEP/SYNC 101
11 2466 136711      EXI C2LD
12 2467 060246      LIO 041     ; SET SAME FOR GT3, NO STEP
13 2470 136715      EXI C3LD
14 2471 043336      LACI RCOUNT ; GET NUMBER OF STEPS FOR THIS SPEED
15 2472 115766      SAC$ SCOUNT
16
17 2473 164666      STEP1: JST POLL ; POLL INTERRUPTS FOR MEASURED INTERVAL
18 2474 045752      LAC$ DSOLD   ; THIS STEP USES OLD NULL DATA
19 2475 115751      SAC$ DSLAT
20 2476 164724      JST YESSTP   ; DO STEP ACCOUNTING
21 2477 035776      DMS$ DOWNLO  ; DEC DOWN COUNT...TST IF DONE
22 2500 144521      JMP LOOP2B   ; NOT DONE, CONTIN
23 2501 045711      LAC$ DOWNHI  ; COULD BE DONE...TST HI COUNT
24 2502 130000      TAZ
25 2503 144556      JMP THERE    ; DONE NOW
26 2504 144516      JMP LOOP2A   ; NOT DONE, CONTINUE WITH RAMP
27
28 2505 115766      LOOP1: SAC$ SCOUNT
29 2506 164572      LOOP2: JST DNCHK ; ANY STEPS AT THIS SPEED?
30 2507 063323      LIOI RGTG   ; YES - LOAD TIME REGISTER
31 2510 136710      EXI T2LD
32 2511 136714      EXI T3LD
33 2512 164666      JST POLL
34 2513 164724      JST YESSTP
35 2514 035776      DMS$ DOWNLO  ; DECREMENT DOWN
36 2515 144521      JMP LOOP2B
37 2516 035711      LOOP2A: DMS$ DOWNHI ; DECREMENT OF HIGH WORD DONE 1 EARLY
38 2517 144507      JMP LOOP2+1  ; FOR EASE OF PROGRAMMING
39 2520 144507      JMP LOOP2+1
40 2521 025766      LOOP2B: DMS$ SCOUNT
41 2522 144506      JMP LOOP2    ; SEE IF SHOULD RAMP DOWN YET
42 2523 137401      BIN          ; INCREMENT INDEX TO NEXT SPEED
43 2524 043336      LACI RCOUNT ; SEE IF NEXT SPEED IS TOP SPEED
44 2525 130400      TAN
45 2526 144505      JMP LOOP1    ; NO - STILL ACCELERATING
46
47 2527 164572      LOOP3: JST DNCHK ; TOP SPEED REACHED - GO AT THIS RATE
48 2530 063323      LIOI RGTG   ; UNTIL TIME TO START SLOWING DOWN
49 2531 136710      EXI T2LD
50 2532 136714      EXI T3LD
51 2533 164666      JST POLL
52 2534 164724      JST YESSTP
53 2535 035776      DMS$ DOWNLO  ; DECREMENT DOWN AS BEFORE
54 2536 144527      JMP LOOP3    ; DO MORE STEPS
55 2537 035711      DMS$ DOWNHI
56 2540 144530      JMP LOOP3+1
57 2541 144530      JMP LOOP3+1

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1
2 2542 137404 LOOP4: TIN ; RAMP DOWN
3 2543 144545 JMP 13 ; IF INDEX IS 0 OR 1, FINISH REST OF
4 2544 144547 JMP 25 ; STEPS AT 100 HZ
5 2545 137402 1%: DIN
6 2546 144530 JMP LOOP3+1 ; STILL AT INTERMEDIATE SPEED
7 2547 061323 2%: LIO RGTG ; LOAD LOWEST SPEED
8 2550 136710 EXI T2LD
9 2551 136714 EXI T3LD
10 2552 164666 JST POLL
11 2553 164724 JST YESSTP
12 2554 035776 DMS% DOWNLO ; STEP UNTIL DOWN GOES TO ZERO
13 2555 144547 JMP 25
14
15 2556 060173 THERE: LIO ONE1 ; DOWN = 0, STOP GTG AFTER REGULAR
16 2557 136711 EXI C2LD ; STEP INTERVAL, NO RESTART
17 2560 136715 EXI C3LD
18 2561 165073 JST GTBUSY ; RETURN TO INTRPT SERVICING
19 2562 040254 LAC 010K1 ; DO 200 MS WAIT
20 2563 170360 JST UWTX
21 2564 136440 EXI RDS ; GET NULL FROM DIR STAT
22 2565 105752 SIO% DSOLD ; SAVE NULL FOR NEXT MOVE
23 2566 165103 NOMOV: JST SMWLD ; SET HARDWARE POSN COUNT TO AWLD
24 2567 161641 JST ESYCLR ; ON WITH L2 INTRPTS
25 2570 137002 LSA 2
26 2571 144363 JMP MOVWLD ; RETURN
27
28 2572 144572 DNCHEK: JMP ; IF DOWN .LT. RAMP(I), GO TO LOOP4
29 2573 045711 LAC% DOWNHI ; ELSE, RETURN
30 2574 130000 TAZ ; IF DOWN > 65,535, IS ALSO > RAMP
31 2575 144577 JMP 13
32 2576 144572 JMP DNCHEK
33 2577 055776 1%: LANS DOWNLO ; IF DOWN > 32,767, IS ALSO > RAMP
34 2600 130400 TAN ; SO RETURN TO ADDR + 2
35 2601 144572 JMP DNCHEK
36 2602 053351 LANI RAMP ; CALCULATE DOWN - RAMP
37 2603 005776 ADD% DOWNLO
38 2604 130400 TAN ; IF POSITIVE, DOWN > RAMP AND RETURN
39 2605 144572 JMP DNCHEK ; TO ADDR + 2
40 2606 144542 JMP LOOP4
41
42 2607 144607 LDCMD: JMP ; SET UP STEP COMMAND
43 2610 005250 ADD% CMD ; SEE IF SAME AS LAST DIRECTION
44 2611 130000 TAZ ;
45 2612 144607 JMP LDCMD ; YES - RETURN
46 2613 045250 LAC% CMD ; STORE CHANGE OF DIRECTION
47 2614 065252 LIO% XCMD ; MOVE CMD TO XCMD AND XCMD TO CMD
48 2615 105250 SIO% CMD
49 2616 115252 SAC% XCMD
50 2617 045745 LAC% PRIOR ; SEE IF LAST STEP WAS NULL
51 2620 130400 TAN
52 2621 144650 JMP LDNUL ; YES
53 2622 055750 LANS MOTOR ; SET MOTOR AND NULL SO THAT DRIVE
54 2623 000251 ADD DS6 ; APPEARS TO HAVE COME FROM NEW DIRECTIO
55 2624 115750 SAC% MOTOR
56 2625 045746 RENU: LAC% NULLHI ; CALCULATE POSN OF VIRTUAL NULL
57 2626 065747 LIO% NULLLO ; IE, ADD 48 TO NULL IF NEW DIRECTION IS

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1 2627 115767 SAC$ MT1 ; REVERSE, AND SUBTRACT 48 IF NEW IS FWD
2 2630 105770 SIO$ MT2
3 2631 132000 CAC
4 2632 060250 LIO D48
5 2633 115771 SAC$ MT3
6 2634 105772 SIO$ MT4
7 2635 045765 LAC$ DIRECT ; TEST DIRECTION
8 2636 130400 TAN
9 2637 144642 JMP 1$
10 2640 165542 JST MDPA ; REVERSE, SO ADD 43
11 2641 144643 JMP 2$
12 2642 165572 1$: JST MDPS ; FORWARD, SO SUBTRACT
13 2643 045771 2$: LAC$ MT3
14 2644 065772 LIO$ MT4
15 2645 115746 SAC$ NULLHI ; REPLACE LAST NULL WITH VIRTUAL NULL
16 2646 105747 SIO$ NULLLO
17 2647 144607 JMP LDCMD
18 2650 045752 LDNUL: LAC$ DSOLD ; LAST STEP WAS NULL, SEE ABOUT THIS
19 2651 135001 RSR 1 ; GET NULL FROM DIRECT STATUS
20 2652 130400 TAN ; SET ?
21 2653 144660 JMP 1$ ; NO
22 2654 055750 LAN$ MOTOR ; YES - MULTIPLE NULL, STORED NULL IS OK
23 2655 000244 ADD D22 ; RESET NULL COUNTER
24 2656 115750 SAC$ MOTOR
25 2657 144607 JMP LDCMD ; RETURN
26 2660 055750 1$: LAN$ MOTOR ; TURNING JUST BEYOND NULL
27 2661 005621 ADD$ TWOD15 ; FOOL MOTOR INTO THINKING IT IS COMING
28 2662 115750 SAC$ MOTOR ; FROM CURRENT DIRECTION
29 2663 050173 LAN ONE1 ; BACK TO NULL FROM NOT-NULLS
30 2664 115745 SAC$ PRIOR
31 2665 144625 JMP RENU1L ; CALCULATE POSN OF VIRTUAL NULL
32
33 2666 144666 POLL: JMP . ; UPDATE AWLD AT EOGT AND CHECK NULLS
34 2667 050715 LAN ACCTP ; CODE TIME - 2 = 16
35 2670 003323 ADDI RGTG ; FIND FREE TIME TO NEXT STEP
36 2671 115774 SETFR: SAC$ MT6
37 2672 136140 EXR ISFC ; POLL 16 MS CLK
38 2673 144717 JMP DDMI
39 2674 136141 EXR IMI ; POLL 64 MS CLK
40 2675 144721 JMP DMIF
41
42 2676 050716 LAN LUPTIC ; MINUS CLOCK TICKS PER PASS
43 2677 005774 ADD$ MT6 ; TEST IF TIME REMAINS
44 2700 130400 TAN
45 2701 144671 JMP SETFR
46
47 2702 132000 SAMPLT: CAC
48 2703 136440 1$: EXI RDS ; DIRECT STATUS FOR NULL
49 2704 105774 SIO$ MT6
50 2705 025774 ORR$ MT6 ; COMBINE NULL SIGNAL
51 2706 136322 EXR IPC2 ; TEST IF STEP MADE
52 2707 144713 JMP 2$
53 2710 136323 EXR IPC3 ; BACKUP CLOCK TEST
54 2711 144713 JMP 2$
55 2712 144703 JMP 1$ ; NO STEP YET
56 2713 115751 2$: SAC$ DSLAT ; SAVE NULL BIT
57 2714 144666 JMP POLL ; NO LIMITS

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ORIGINAL PAGE IS
OF POOR QUALITY

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1
2 2715 000037 ACCTP: 31 ;CLOCK PERIODS USED BETWEEN STEPS - 2
3 2716 000007 LUPTIC: 7 ;CLOCK PERIODS PER LOOP PASS (APROX)
4 2717 170317 DQMI: JST QMIFRM
5 2720 144702 JMP SAMPLT
6
7 2721 170301 DMIF: JST MIFRM
8 2722 137002 LSA 2
9 2723 144702 JMP SAMPLT
10
11 2724 144724 YESSTP: JMP . ; THIS USED IN PLACE OF STEP ROUTINE
12 2725 136322 EXR IPC2 ; CLEAR OTHER GT LATCH
13 2726 144727 JMP .+1
14 2727 136323 EXR IPC3 ; DO IT 196 MICROSEC AFTER
15 2730 144731 JMP .+1 ; RECOGNIZED GATE END
16 2731 045751 LAC$ DSLAT ; SEE IF NULL WAS SET
17 2732 010245 AND 030 ; SEE IF A LIMIT SET
18 2733 130000 TAZ
19 2734 144736 JMP 1$
20 2735 151206 JMP LIMCHK ; FOUND LIMIT
21
22 2736 045751 1$: LAC$ DSLAT
23 2737 010173 AND ONE1 ; SEE IF NULL WAS SET
24 2740 130000 TAZ ; SKIP IF NULL SET
25 2741 144762 JMP NONULL
26
27 2742 005745 ADD$ PRIOR ; PRIOR: 0 = LAST STEP WAS NULL
28 2743 130000 TAZ ; -1 = LAST STEP WAS NOT-NULL
29 2744 144751 JMP STEPOK ; HAVE NULL AFTER NOT-NULL - VERY GOOD
30 2745 035750 DMS$ MOTOR ; NULL AFTER NULL - DECREMENT COUNT
31 2746 144775 JMP ENDSTP ; STILL OK
32 2747 165073 JST GTBUSY ; SET BUSY SO NEW SCI'S WILL WAIT
33 2750 145021 JMP HELP4 ; DRIVE STUCK ON NULL
34
35 2751 045771 STEPOK: LAC$ MT3 ; NULL AFTER NOT-NULL, DRIVE OK
36 2752 065772 LIO$ MT4 ; COPY WORKING STORE OF AWLD INTO
37 2753 115746 SAC$ NULLHI ; AWLD AND STORE POSN OF THIS NULL
38 2754 105747 SIO$ NULLLO
39 2755 132000 CAC ; SET PRIOR = 0 = LAST WAS NULL
40 2756 115745 SAC$ PRIOR
41 2757 040243 LAC D10 ; ALLOW 10 NULLS IN A ROW
42 2760 115750 SAC$ MOTOR
43 2761 144775 JMP ENDSTP
44
45 2762 005745 NONULL: ADD$ PRIOR ; THIS STEP NOT-NULL
46 2763 130000 TAZ ; CHECK PREVIOUS STEP
47 2764 144771 JMP STEPA1 ; HAVE NOT-NULL AFTER NULL - GOOD
48 2765 035750 DMS$ MOTOR ; NOTNULL AFTER NOTNULL - DECREMENT COU
49 2766 144775 JMP ENDSTP ; STILL OK
50 2767 165073 JST GTBUSY ; SET BUSY SO NEW SCI'S WILL WAIT
51 2770 145027 JMP HELPS ; MOTOR STUCK ON NOT-NULL
52
53 2771 045750 STEPA1: LAC$ MOTOR ; NOT-NULL AFTER NULL, GOOD
54 2772 000247 ADD D40 ; ALLOW 50 STEPS BETWEEN NULLS
55 2773 115750 SAC$ MOTOR
56 2774 035745 DMS$ PRIOR ; SET PRIOR = -1 = NOT-NULL
57 2775 045770 ENDSTP: LAC$ MT2 ; ADD DIRECT TO AWLD AND REPLACE AWLD

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1 2776 005772      ADD$ MT4      ; MT2 IS DIRECT, MT4 IS LS 16 BITS OF A
2 2777 115772      SAC$ MT4      ; AWLD. ADD THEM FIRST, THEN TEST RESULT
3 3000 130000      TAZ          ; IF ZERO, MT3=MT3+MT1+1
4 3001 145007      JMP 1$
5 3002 000173      ADD ONE1     ; IF MINUS 1, MT3=MT3+MT1
6 3003 130000      TAZ
7 3004 145010      JMP 2$
8 3005 045771      LAC$ MT3     ; OTHERWISE, MT3=MT3
9 3006 145013      JMP 3$
10 3007 000173 1$:  ADD ONE1
11 3010 005771 2$:  ADD$ MT3
12 3011 005767      ADD$ MT1
13 3012 115771      SAC$ MT3
14 3013 065772 3$:  LIO$ MT4      ; STORE MT3/4 INTO AWLD
15 3014 137003      LSA 3
16 3015 115761      SAC$ AWLDHI
17 3016 105762      SIO$ AWLDLO
18 3017 137002      LSA 2
19 3020 144724      JMP YESSTP   ; STEP DATA WAS GOOD--RETURN
20
21 3021 060173  HELP4: LIO ONE1     ; DRIVE STUCK ON NULL
22 3022 136711      EXI C2LD    ; STOP SENDING COMMANDS
23 3023 136715      EXI C3LD    ; STOP BACKUP GTG
24 3024 061365      LIO WMSG4   ; OUTPUT MSG = 133334
25 3025 171353      JST USMSG
26 3026 145052      JMP SETNUL   ; SET AWLD = LAST NULL POSN AND RETRY
27
28 3027 060173  HELP5: LIO ONE1     ; DRIVE STUCK OFF NULL
29 3030 136711      EXI C2LD    ; STOP SENDING STEP COMMANDS
30 3031 136715      EXI C3LD
31 3032 061366      LIO WMSG5   ; OUTPUT MSG = 133335
32 3033 171353      JST USMSG
33 3034 040175      LAC FOUR1   ; SET STEP COUNTER
34 3035 115750      SAC$ MOTOR
35 3036 170371      JST UNTONE  ; WAIT 1 SEC BEFORE CHANGING DIRECTION
36 3037 035750  LOOPF: DMS$ MOTOR  ; STEP IN OTHER DIRECTION AT 50 HZ
37 3040 151252      JMP XCMD    ; UNTIL NULL, KEEP TRACK OF CLOCKS
38 3041 151252      JMP XCMD    ; DECREMENT STEP COUNT FOR EACH STEP
39 3042 041475  RECOVR: LAC UND420 ; WAIT 20 MS
40 3043 170360      JST UWTX
41 3044 136440      EXI RDS    ; CHECK NULL
42 3045 105752      SIO$ DSOLD
43 3046 045752      LAC$ DSOLD
44 3047 010173      AND ONE1
45 3050 130000      TAZ        ; SET ?
46 3051 145375      JMP LIMIT1 ; NO - AND MAKE SURE NULL WITHIN 300 STEPS
47
48 3052 055750  SETNUL: LAN$ MOTOR  ; IF MOTOR .GT. ZERO, DIDNT HAVE TO STEP
49 3053 130400      TAN        ; FAR BACK TO NULL - ASSUME NULL MISSED
50 3054 145057      JMP 1$     ; RATHER THAN DRIVE STUCK
51 3055 055252      LAN$ XCMD  ; SO SET LAST NULL 48 STEPS NEARER
52 3056 164607      JST LDCMD
53 3057 040243 1$:  LAC D10
54 3060 115750      SAC$ MOTOR
55 3061 132000      CAC
56 3062 115745      SAC$ PRIOR
57 3063 045746      LAC$ NULLHI ; NULL - SET AWLD = LAST NULL

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1 3064 065747      LIO$ NULLLO
2 3065 137003      LSA 3
3 3066 115761      SAC$ AWLDHI
4 3067 105762      SIO$ AWLDLO
5 3070 161641      JST BSYCLR      ; RECOVERY COMPLETE, CLEAR BUSY
6 3071 170371      HELPER: JST UNTONE ; WAIT
7 3072 144364      JMP MOVWLD+1    ; TRY AGAIN TO GET TO 'WANT'
8
9 3073 145073      GTBUSY: JMP .    ; SET BUSY TO NEG SO NO NEW SCI COMMAND
10 3074 137003      LSA 3          ; WILL BE PROCESSED UNTIL ERROR IS
11 3075 045777      LAC$ BSYGN     ; RECOVERED
12 3076 021350      ORR RAMP-1    ; SIGN BIT USED AS FLAG
13 3077 115777      SAC$ BSYGN
14 3100 137002      LSA 2
15 3101 133400      CIL
16 3102 145073      JMP GTBUSY    ; RETURN
17
18 3103 145103      SMWLD: JMP .
19 3104 137003      RECHK: LSA 3   ; SET HARDWARE COUNTER = AWLD
20 3105 045761      LAC$ AWLDHI   ; STORE AWLD IN MT1/2
21 3106 065762      LIO$ AWLDLO
22 3107 137002      LSA 2
23 3110 115767      SAC$ MT1
24 3111 105770      SIO$ MT2
25 3112 165176      JST SMPOSN    ; LOAD COUNTER INTO MT3/4
26 3113 165572      JST MDPS      ; CALCULATE AWLD - COUNTER
27 3114 045771      LAC$ MT3      ; SEE IF EQUAL
28 3115 005772      ADD$ MT4
29 3116 130000      TAZ          ; EQUAL ?
30 3117 145103      JMP SMWLD     ; YES - RETURN
31 3120 060173      LIO ONE1     ; ASSUME CHANGE FWD
32 3121 105765      SIO$ DIRECT
33 3122 045771      LAC$ MT3
34 3123 130400      TAN          ; IS IT FORWARD ?
35 3124 145130      JMP 1$       ; YES
36 3125 165625      JST MDPNOT    ; NO - NEGATE CHANGE AND REVERSE DIRECTI
37 3126 055765      LAN$ DIRECT
38 3127 115765      SAC$ DIRECT
39
40 3130 075771      1$: LIN$ MT3 ; SEE HOW BIG CHANGE IS
41 3131 147132      JMPI .+1
42 3132 145150      JMP 6$       ; CLOSE AND IN DIRECTION SET
43 3133 145136      JMP 2$       ; FAR IN DIRECTION SET
44 3134 145141      JMP 3$       ; FAR IN OPPOSITE DIRECTION
45 3135 145144      JMP 4$       ; NEAR IN OPPOSITE DIRECTION
46
47 3136 050173      2$: LAN ONE1 ; FAR IN DIRECTION SET
48 3137 115772      SAC$ MT4     ; SET MT4 = MAX
49 3140 145150      JMP 6$
50
51 3141 050173      3$: LAN ONE1 ; FAR IN OPPOSITE DIRECTION
52 3142 115772      SAC$ MT4     ; SET MT4 = MAX AND CHANGE DIRECTION
53 3143 145146      JMP 5$
54
55 3144 055772      4$: LAN$ MT4 ; NEAR IN OPPOSITE DIRECTION
56 3145 115772      SAC$ MT4     ; NEGATE CHANGE SIZE
57 3146 055765      5$: LAN$ DIRECT ; CHANGE DIRECTION

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1 3147 115765      SAC$ DIRECT
2
3 3150 136641 6$:  EXI SLWF      ; SET DIRECTION FWD
4 3151 055765      LAN$ DIRECT
5 3152 130400      TAN          ; CHANGE TO REV IF DIRECT = -1
6 3153 136642      EXI SLWR
7 3154 170401      JST GET8     ; STORE LO 8 BITS OF START POSN
8 3155 115763      SAC$ OLD
9 3156 045772      LAC$ MT4     ; MAGNITUDE OF CHANGE IN MT4
10 3157 134010     ASR 8        ; PULL OUT HI 8 BITS
11 3160 115771      SAC$ MT3
12 3161 065771      LIO$ MT3
13 3162 136643      EXI LMSH     ; LOAD THEM IN MSH
14 3163 065772      LIO$ MT4     ; LOAD LO 8 BITS IN LSH
15 3164 136654      EXI WAVI     ; AND INITIATE CHANGE
16
17 3165 170401     LOOPW: JST GET8 ; POSN CHANGES EVERY .0625 MICRO SEC
18 3166 115774      SAC$ MT6
19 3167 055763      LAN$ OLD     ; FINISHED WHEN NO MORE CHANGE
20 3170 005774      ADD$ MT6
21 3171 130000      TAZ          ; CHANGED ?
22 3172 145104      JMP RECHK    ; NO - NOW CHECK POSITION
23 3173 045774      LAC$ MT6     ; STILL CHANGING - UPDATE OLD
24 3174 115763      SAC$ OLD
25 3175 145165      JMP LOOPW    ; CHECK AGAIN
26
27 3176 145176     SMPOSN: JMP .      ; READ SM COUNTER INTO MT3/4
28 3177 170401      JST GET8     ; GET LS 8 BITS FROM DS
29 3200 115772      SAC$ MT4     ; STORE IN MT4
30 3201 065617      LIO$ NINE2    ; BITS 18 - 9 ARE IN SM CHAN 9
31 3202 170051      JST USMR
32 3203 115771      SAC$ MT3     ; STORE IN MT3
33 3204 134410      ASL 8        ; MOVE BITS 16-9 TO POSNS 16-9
34 3205 025772      ORR$ MT4     ; INSERT INTO MT4
35 3206 115772      SAC$ MT4     ; STORE IN MT4
36 3207 045771      LAC$ MT3     ; GET MT3
37 3210 134010     ASR 8        ; MOVE TO POSNS 2-1
38 3211 010242      AND THREE1    ; MASK THEM
39 3212 115771      SAC$ MT3     ; STORE
40 3213 145176     JMP SMPOSN    ; RETURN
41
42 3214 165073     LIMITA: JST GTBUSY ; SET BUSY SO NEW SCI'S WILL WAIT
43 3215 171425      JST UERROR   ; RECORD ERROR
44 3216 041371      LAC AHI     ; SET AWLD TO AHI/ALO
45 3217 061372      LIO ALO
46 3220 137003      LSA 3
47 3221 115761      SAC$ AWLDHI
48 3222 105762      SIO$ AWLDLO
49 3223 051370      LAN STEPR   ; STEP REVERSE FROM LIMIT
50 3224 145235      JMP AWAY
51
52 3225 165073     LIMITB: JST GTBUSY ; SET BUSY SO NEW SCI'S WILL WAIT
53 3226 171425      JST UERROR   ; RECORD ERROR
54 3227 041373      LAC BHI     ; SET AWLD TO BHI/BLO
55 3230 061374      LIO BLO
56 3231 137003      LSA 3
57 3232 115761      SAC$ AWLDHI

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1 3233 105762      SIO% AWLDLO
2 3234 051367      LAN STEPF      ; STEP FORWARD FROM LIMIT
3
4 3235 137002      AWAY:  LSA 2
5 3236 164607      JST LDCMD      ; SET UP DIRECTION
6 3237 041245      LAC LSRET1     ; SET UP RETURN FROM 'CMD'
7 3240 115251      SAC% CMD+1
8 3241 040175      LAC FOUR1     ; REPHASE MOTOR BY STEPPING 4 AT 1 HZ
9 3242 115773      SAC% MTS
10 3243 170371     JST UWTONE    ; INITIAL WAIT OF 1 SEC
11 3244 151250     JMP CMD       ; SEND STEP COMMAND
12 3245 145246     LSRET1: JMP .+1 ; RETURN TO HERE
13 3246 170371     JST UWTONE    ; WAIT 1 SEC
14 3247 035773     DMS% MTS     ; DONE 4 ?
15 3250 151250     JMP CMD       ; NO - DO ANOTHER
16 3251 041253     LAC LSRET2    ; NOW MOVE AT 100 HZ AWAY UNTIL
17 3252 115251     SAC% CMD+1    ; LIMIT SWITCH GOES OFF
18 3253 145254     LSRET2: JMP .+1
19 3254 040253     LAC D209     ; WAIT 10 MS
20 3255 170360     JST UWTX
21 3256 136440     EXI RDS      ; READ LIMIT SWITCHES
22 3257 105773     SIO% MTS
23 3260 045773     LAC% MTS
24 3261 134003     ASR 3        ; THEY ARE BITS 4-5
25 3262 010242     AND THREE1
26 3263 130000     TAZ          ; STILL SET ?
27 3264 145266     JMP .+2      ; NO
28 3265 145404     JMP LIMIT2   ; YES - MAKE SURE .LT. 5000 STEPS OUT OF LIMIT
29 3266 041270     LAC LSRET3   ; NOW FIND NEXT NULL
30 3267 115251     SAC% CMD+1
31
32 3270 145271     LSRET3: JMP .+1 ; STEP AT 1 HZ TO NULL
33 3271 170371     JST UWTONE
34 3272 136440     EXI RDS
35 3273 105752     SIO% DSOLD   ; SEE IF NULL SET
36 3274 045752     LAC% DSOLD
37 3275 135001     RSR 1
38 3276 130400     TAN
39 3277 145416     JMP LIMIT3   ; NO - BE SURE NULL WITHIN 300 STEPS
40
41 3300 137003      OUT:  LSA 3
42 3301 045761      LAC% AWLDHI   ; SET LAST NULL = AWLD
43 3302 065762      LIO% AWLDLO
44 3303 137002      LSA 2
45 3304 115746      SAC% NULLHI
46 3305 105747      SIO% NULLLLO
47 3306 040243      LAC D10      ; ALLOW 10 NULLS
48 3307 115750      SAC% MOTOR
49 3310 132000      CAC          ; SET PRIOR = NULL
50 3311 115745      SAC% PRIOR
51 3312 165103      JST SMWLD    ; SET WLD POSN COUNTER TO AWLD
52 3313 136250      EXR ILTA    ; CLEAR LIMIT SWITCH INTERRUPTS
53 3314 134000      NOP
54 3315 136251      EXR ILTB
55 3316 134000      NOP
56 3317 161641      JST BSYCLR   ; RECOVERY COMPLETE, CLEAR BUSY
57 3320 160012      JST HALT     ; HALT EXPERIMENT OR SET UP

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1 3321 140260      JMP POWRUP      ; WAIT FOR NEXT LEVEL2
2
3
4 3322 000005  SLACK: 5          ; STEPS TO TAKE UP SLACK
5 3323 000236  RGTG: 158. ; 100 HZ LIST OF TIME VALUES FOR RAMP UP
6 3324 000236      158. ; ** UNIFORM 100-HZ STEPPING RATE... **
7                ;** 90. ; 174 HZ REVERSE ORDER IS RAMP DOWN
8 3325 000101      65. ; 239 HZ LIST CAN BE ANY LENGTH
9 3326 000065      53. ; 291 HZ
10 3327 000055     45. ; 340 HZ
11 3330 000050     40. ; 381 HZ
12 3331 000044     36. ; 421 HZ
13 3332 000042     34. ; 444 HZ
14 3333 000040     32. ; 471 HZ
15 3334 000037     31. ; 485 HZ
16 3335 000036     30. ; 500 HZ
17 3336 000310  RCOUNT: 200. ; NO OF STEPS TO DO AT EACH FREQUENCY
18 3337 100000     100000 ; ** MEANS 1ST SPEED(100HZ) IS TOP SPEED **
19                ;** 348. ; LISTED IN RGTG
20 3340 000736      478.
21 3341 001106      582.
22 3342 001250      680.
23 3343 001372      762.
24 3344 001512      842.
25 3345 001570      888.
26 3346 001656      942.
27 3347 001712      970.
28 3350 100000     100000 ; NEG COUNT MEANS TOP SPEED REACHED
29                ; SIGN BIT USED AS A FLAG THIS SEGMENT
30 3351 000620  RAMP: 400. ; RAMP LENGTH AT EACH FREQUENCY
31 3352 000310      200.
32 3353 001044      548.
33 3354 002002     1026.
34 3355 003110     1608.
35 3356 004360     2288.
36 3357 005752     3050.
37 3360 007464     3892.
38 3361 011254     4780.
39 3362 013132     5722.
40 3363 015042     6690.
41 3364 000101  GTGWAV: 101 ; 16 KHZ CTR DUMP SYNC/STEP
42
43                ; NLD ERROR MESSAGES
44
45 3365 133334  WMSG4: 133334 ; DRIVE STUCK ON NULL
46 3366 133335  WMSG5: 133335 ; DRIVE STUCK OFF NULL
47
48 3367 136655  STEPF: EXI STPF
49 3370 136656  STEPR: EXI STPR
50 3371 000003  AHI: 3 ; POSN OFF LIMIT A = 219699.
51 3372 045234  ALO: 045234
52 3373 000000  BHI: 0 ; POSN OFF LIMIT B = 2539.
53 3374 004753  BLO: 004753
54
55 3375 045750  LIMIT1: LAC$ MOTOR ; MOTOR SHOULD BE .GT. -300
56 3376 001425  ADD STPMAX
57 3377 130400  TAN

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1 3400 145037      JMP LOOPF      ; STILL OK
2 3401 145073      JST GTBUSY    ; SET BUSY TO INHIBIT SCI'S
3 3402 171425 1$:  JST UERROR    ; WLD IS TOTALLY STUCK (OR NULL GONE)
4 3403 145402      JMP 1$        ; KEEP FLAGGING TM
5
6 3404 035745      LIMIT2: DMS$ PRIOR ; MAKE SURE IT TAKES NO MORE THAN 5000
7 3405 134000      NOP          ; STEPS TO GET OUT OF THE LIMIT
8 3406 132000      CAC          ; CLEAR MOTOR FOR NEXT PHASE
9 3407 115750      SAC$ MOTOR
10 3410 045745      LAC$ PRIOR    ; MAKE SURE PRIOR STILL .GT. -5000
11 3411 001426      ADD STPMXL
12 3412 130400      TAN
13 3413 151250      JMP CMD      ; OK- KEEP STEPPING
14 3414 171425 1$:  JST UERROR    ; WLD SEEMS STUCK INSIDE LIMIT
15 3415 145414      JMP 1$        ; KEEP FLAGGING TM
16
17 3416 035750      LIMIT3: DMS$ MOTOR ; MAKE SURE NULL FOUND WITHIN 300 STEPS
18 3417 045750      LAC$ MOTOR
19 3420 001425      ADD STPMAX
20 3421 130400      TAN
21 3422 151250      JMP CMD      ; STILL OK - DO MORE STEPS
22 3423 171425 1$:  JST UERROR    ; WLD SEEMS STUCK
23 3424 145423      JMP 1$        ; KEEP FLAGGING TM
24
25 3425 000454      STPMAX: 300.
26 3426 015530      STPMXL: 7000.
27
28 ; CHANGE WLD POSITION COUNTER BY THE CONTENTS OF FIX
29 ;   ADDING FIX SHIFTS LINES TO THE BLUE
30 ;   SUBTRACTING FIX SHIFTS LINE TO THE RED
31
32 3427 010173      FIXWL: AND ONE1 ; ADD TO OR SUB FROM AWLD AC=(SCI)
33 3430 130000      TAZ          ; BIT 1 = 1 = ADD
34 3431 145476      JMP RED      ; RED MOVES LINE TO RED
35
36 ; BLUE:
36 3432 045762      LAC$ AWLDLO   ; BLUE MOVES LINE TO BLUE
37 3433 115776      SAC$ TEMPJ   ; ADD VALUE IN 7755 TO AWLD AND NULL
38 3434 005757      ADD$ FIX
39 3435 115762      SAC$ AWLDLO
40 3436 130400      TAN          ; SEE IF HAVE TO CARRY TO MS 2 BITS
41 3437 145441      JMP 1$
42 3440 145447      JMP 2$
43 3441 045776 1$:  LAC$ TEMPJ
44 3442 130400      TAN
45 3443 145447      JMP 2$
46 3444 045761      LAC$ AWLDHI
47 3445 000173      ADD ONE1
48 3446 115761      SAC$ AWLDHI
49 3447 137002 2$:  LSA 2
50 3450 045747      LAC$ NULLLO
51 3451 137003      LSA 3
52 3452 115776      SAC$ TEMPJ   ; ADD VAL TO NULL TOO
53 3453 005757      ADD$ FIX
54 3454 137002      LSA 2
55 3455 115747      SAC$ NULLLO
56 3456 130400      TAN
57 3457 145461      JMP 3$        ; SEE IF HAVE TO CARRY TO MS HALF

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1 3460 145471          JMP 4$
2 3461 137003 3$:     LSA 3
3 3462 045776          LAC$ TEMPJ
4 3463 130400          TAN
5 3464 145471          JMP 4$
6 3465 137002          LSA 2
7 3466 045746          LAC$ NULLHI
8 3467 000173          ADD ONE1
9 3470 115746          SAC$ NULLHI
10 3471 137003 4$:     LSA 3
11 3472 045757          LAC$ FIX          ; STORE FIX VALUE IN DMA
12 3473 115760  DONE: SAC$ LASTFX
13 3474 140122          JMP INRTN          ; RETURN
14
15 3475 000644  UND420: 420.
16
17 3476 045762  RED: LAC$ AWLDLO          ; SUBTRACT VALUE IN FIX FROM AWLD AND NULL
18 3477 115776          SAC$ TEMPJ
19 3500 055757          LAN$ FIX
20 3501 005762          ADD$ AWLDLO
21 3502 115762          SAC$ AWLDLO
22 3503 130400          TAN          ; SEE IF BORROW WAS DONE
23 3504 145514          JMP 2$
24 3505 045776          LAC$ TEMPJ
25 3506 130400          TAN
26 3507 145511          JMP 1$
27 3510 145514          JMP 2$
28
29 3511 050173 1$:     LAN ONE1
30 3512 005761          ADD$ AWLDHI
31 3513 115761          SAC$ AWLDHI
32 3514 137002 2$:     LSA 2
33 3515 045747          LAC$ NULLLO
34 3516 137003          LSA 3
35 3517 115776          SAC$ TEMPJ
36 3520 055757          LAN$ FIX
37 3521 137002          LSA 2
38 3522 005747          ADD$ NULLLO
39 3523 115747          SAC$ NULLLO
40 3524 130400          TAN
41 3525 145537          JMP 4$          ; SEE IF BORROW
42 3526 137003          LSA 3
43 3527 045776          LAC$ TEMPJ
44 3530 130400          TAN
45 3531 145533          JMP 3$
46 3532 145537          JMP 4$
47 3533 137002 3$:     LSA 2
48 3534 050173          LAN ONE1
49 3535 005746          ADD$ NULLHI
50 3536 115746          SAC$ NULLHI
51 3537 137003 4$:     LSA 3
52 3540 055757          LAN$ FIX          ; STORE FIX NEGATIVE FOR SUBTRACT
53 3541 145473          JMP DONE
54
55          ; MATH
56          ; INCLUDES DP ADD, SUBTRACT, NEGATE, MULTIPLY, DIVIDE, AND SHIFT
57

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1      ; DOUBLE PRECISION ADDITION, SUBTRACTION, AND COMPLEMENTATION
2      ;   ENTRY POINTS: MDPA, MDPS, MDPNOT
3      ;   ARGUMENTS: FOR OPERATION OF THE FORM B=A+B, B=A-B, OR B=-B
4      ;               MT1*MT2 = 32 BITS OF A
5      ;               MT3*MT4 = 32 BITS OF B
6
7 3542 145542 MDPA: JMP . ; ENTRY - D P ADDITION
8 3543 137002      LSA 2
9 3544 045772      LAC$ MT4 ; TEST LS HALF OF B
10 3545 130400     TAN      ; NEGATIVE ?
11 3546 145557     JMP 2$   ; NO
12 3547 005770     ADD$ MT2 ; YES - ADD LSS HALF OF A
13 3550 115772     SAC$ MT4 ; STORE IN LS HALF OF B
14 3551 130400     TAN      ; NEW RESULT NEGATIVE ?
15 3552 145565     JMP 4$   ; NO - CARRY ONE TO MS HALF
16 3553 045770 1$: LAC$ MT2 ; YES - TEST LS HALF OF A
17 3554 130400     TAN      ; NEGATIVE
18 3555 145563     JMP 3$   ; NO - DONT CARRY
19 3556 145565     JMP 4$   ; YES - CARRY ONE
20 3557 005770 2$: ADD$ MT2 ; ADD LS HALF OF A TO B
21 3560 115772     SAC$ MT4 ; STORE IN LS HALF OF B
22 3561 130400     TAN      ; NEGATIVE ?
23 3562 145553     JMP 1$   ; NO - TEST LS HALF OF A
24 3563 132000 3$: CAC      ; SET CARRY = 0
25 3564 145566     JMP 5$
26 3565 040173 4$: LAC ONE1 ; SET CARRY = 1
27 3566 005771 5$: ADD$ MT3 ; ADD MS HALF OF B
28 3567 005767     ADD$ MT1 ; ADD MS HALF OF A
29 3570 115771     SAC$ MT3 ; STORE IN MS HALF OF B
30 3571 145542     JMP MDPA ; DONE
31
32 3572 145572 MDPS: JMP . ; ENTRY D P SUBTRACTION
33 3573 137002      LSA 2
34 3574 045772      LAC$ MT4 ; TEST LS HALF OF B
35 3575 130400     TAN      ; NEGATIVE ?
36 3576 145600     JMP 1$   ; NO
37 3577 145611     JMP 3$   ; YES
38 3600 055772 1$: LAN$ MT4 ; SUBTRACT LOW ORDER BYTES FIRST
39 3601 005770     ADD$ MT2
40 3602 115772     SAC$ MT4 ; STORE RESULT IN LS HALF OF B
41 3603 130400     TAN      ; NEW RESULT NEGATIVE ?
42 3604 145621     JMP 5$   ; NO - NO BORROW
43 3605 045770 2$: LAC$ MT2 ; YES - TEST LS HALF OF A
44 3606 130400     TAN      ; NEGATIVE ?
45 3607 145616     JMP 4$   ; NO - BORROW ONE
46 3610 145621     JMP 5$   ; NO BORROW
47 3611 055772 3$: LAN$ MT4 ; DO SUBTRACT ON LS WORDS
48 3612 005770     ADD$ MT2
49 3613 115772     SAC$ MT4 ; STORE IN LS HALF OF B
50 3614 130400     TAN      ; RESULT NEGATIVE
51 3615 145605     JMP 2$   ; NO
52 3616 040173 4$: LAC ONE1 ; BORROW ONE
53 3617 005771     ADD$ MT3 ; ADD BORROW TO MS HALF OF B
54 3620 115771     SAC$ MT3
55 3621 055771 5$: LAN$ MT3 ; DO SUBTRACT ON MS WORDS
56 3622 005767     ADD$ MT1
57 3623 115771     SAC$ MT3 ; STORE RESULT IN B

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1 3624 145572          JMP MDPS          ; DONE
2
3 3625 145625 MDPNOT: JMP .          ; ENTRY DP NEGATE
4 3626 137002          LSA 2            ; SUBTRACT B FROM ZERO
5 3627 132000          CAC
6 3630 115767          SAC% MT1
7 3631 115770          SAC% MT2
8 3632 145572          JST MDPS
9 3633 145625          JMP MDPNOT        ; DONE
10
11                   ; MULTIPLY
12                   ; ENTRY POINTS: MLTPLY
13                   ; ARGUMENTS: MT4 = 16 BIT MULTIPLIER
14                   ;                   MT5 = 16 BIT MULTIPLICAND
15                   ;                   MT3*MT4 = 32 BIT PRODUCT
16 3634 145634 MLTPLY: JMP .          ; ENTRY
17 3635 137002          LSA 2
18 3636 132000          CAC                ; CLEAR MT3 AND MT7
19 3637 115775          SAC% MT7          ; MT7 IS FLAG OF LAST MULTIPLIER BIT
20 3640 115771          SAC% MT3          ; MT3 IS MS HALF OF PRODUCT
21 3641 075622          LINS TWOD16        ; LOAD OPERATION CTR = 16.
22 3642 045772 1%:    LAC% MT4          ; TEST WHETHER CURRENT MULTIPLIER BIT
23 3643 134417          ASL 15.          ; IS THE SAME AS THE LAST MULTIPLIER BIT
24 3644 005775          ADD% MT7
25 3645 130400          TAN                ; IF SAME, NO CHANGE TO PRODUCT
26 3646 145661          JMP 3%          ; SAME
27 3647 005775          ADD% MT7          ; BITS ARE DIFFERENT
28 3650 130400          TAN
29 3651 145655          JMP .+4
30 3652 115775          SAC% MT7          ; UPDATE LAST = CURRENT = 1
31 3653 055773          LANS MT5
32 3654 145657          JMP 2%
33 3655 115775          SAC% MT7          ; UPDATE LAST = CURRENT = 0
34 3656 045773          LAC% MT5
35 3657 005771 2%:    ADD% MT3
36 3660 115771          SAC% MT3
37 3661 165744 3%:    JST MDPRS          ; 32 BIT END-OFF SIGN EXTENDED RIGHT SHIFT
38 3662 137402          DIN                ; DECREMENT OPERATION COUNTER
39 3663 145642          JMP 1%
40 3664 145634          JMP MLTPLY        ; DONE
41
42                   ; DIVIDE AND ROUND TO NEAREST INTEGER
43                   ; ENTRY POINTS: DIVIDE
44                   ; ARGUMENTS: MT3*MT4 = 32 BIT DIVIDEND
45                   ;                   MT5 = 16 BIT DIVISOR
46                   ;                   MT4 = ROUNDED QUOTIENT
47                   ;                   MT3 = MAGNITUDE OF REMAINDER BEFORE ROUNDING
48 3665 145665 DIVIDE: JMP .          ; ENTRY
49 3666 137002          LSA 2
50 3667 045773          LAC% MT5          ; TEST DIVISOR
51 3670 130000          TAZ                ; = 0 ?
52 3671 145665          JMP DIVIDE        ; YES - RETURN
53 3672 015605          AND% HIBIT        ; NO - MASK SIGN BIT
54 3673 115775          SAC% MT7
55 3674 045771          LAC% MT3          ; GET SIGN BIT OF DIVIDEND
56 3675 015605          AND% HIBIT
57 3676 005775          ADD% MT7          ; ADD SIGN BIT OF DIVISOR

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1 3677 115774 SAC$ MT6 ; RESULT IS SIGN OF QUOTIENT, STORE IT
2 3700 055773 LAN$ MT5 ; DIVISOR SHOULD BE NEGATIVE
3 3701 130400 TAN ; IF POSITIVE, NEGATE IT
4 3702 145704 JMP 1$
5 3703 115773 SAC$ MT5
6 3704 045771 1$: LAC$ MT3 ; DIVIDEND SHOULD BE POSITIVE
7 3705 130400 TAN
8 3706 145710 JMP 2$
9 3707 165625 JST MDPNOT ; IS NEGATIVE SO COMPLEMENT IT
10 3710 075622 2$: LIN$ TWOD16 ; LOAD OPERATION COUNTER
11 3711 165764 3$: JST MDPLS ; LEFT SHIFT DIVIDEND
12 3712 045771 LAC$ MT3 ; CALCULATE REMAINDER - DIVISOR
13 3713 005773 ADD$ MT5
14 3714 130400 TAN ; ) 0 OR = 0 ?
15 3715 145737 JMP 7$ ; YES
16 3716 137402 4$: DIN ; NO - DECREMENT OPERATION COUNTER
17 3717 145711 JMP 3$ ; DO NEXT BIT
18 3720 045773 LAC$ MT5 ; CALCULATE 2*REMAINDER - DIVISOR
19 3721 005771 ADD$ MT3
20 3722 005771 ADD$ MT3
21 3723 130400 TAN ; WHICH WAY TO ROUND ?
22 3724 145726 JMP 5$ ; ROUND UP
23 3725 145731 JMP 6$ ; ROUND DOWN
24 3726 040173 5$: LAC ONE1
25 3727 005772 ADD$ MT4
26 3730 115772 SAC$ MT4
27 3731 045774 6$: LAC$ MT6 ; TEST SIGN FLAG
28 3732 130400 TAN ; QUOTIENT SUPPOSED TO BE NEGATIVE ?
29 3733 145665 JMP DIVIDE ; NO - IS OK
30 3734 055772 LAN$ MT4 ; YES - MAKE IT NEGATIVE
31 3735 115772 SAC$ MT4
32 3736 145665 JMP DIVIDE ; DONE
33 ;
34 3737 115771 7$: SAC$ MT3 ; DECREMENT REMAINDER BY DIVISOR
35 3740 045772 LAC$ MT4 ; INCREMENT QUOTIENT BY ONE
36 3741 000173 ADD ONE1
37 3742 115772 SAC$ MT4
38 3743 145716 JMP 4$ ; TEST TO SEE IF FINISHED
39 ;
40 ; SHIFT - 32 BIT END-OFF LEFT SHIFT AND RIGHT SHIFT WITH SIGN EXTENDED
41 ; ENTRY POINTS: MDPRS, MDPLS
42 ; ARGUMENTS: MT3*MT4 = 32 BITS TO BE SHIFTED
43 3744 145744 MDPRS: JMP ; ENTRY
44 3745 045772 LAC$ MT4 ; GET LS WORD
45 3746 134001 ASR 1 ; END OFF SHIFT RIGHT ONE
46 3747 115772 SAC$ MT4
47 3750 045771 LAC$ MT3 ; SET BIT 16 OF PAIR
48 3751 134417 ASL 15 ; (IE, BIT 1 OF MS WORD)
49 3752 025772 ORR$ MT4
50 3753 115772 SAC$ MT4
51 3754 045605 LAC$ HIBIT
52 3755 015771 AND$ MT3 ; GET SIGN BIT
53 3756 115774 SAC$ MT6
54 3757 045771 LAC$ MT3
55 3760 134001 ASR 1 ; END OFF SHIFT MS WORD
56 3761 025774 ORR$ MT6 ; EXTEND SIGN BIT
57 3762 115771 SAC$ MT3

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1 3763 145744      JMP MDPRS
2
3 3764 145764 MDPLS: JMP . ; ENTRY
4 3765 045771      LAC% MT3      ; SHIFT MS WORD 1 LEFT
5 3766 134401      ASL 1
6 3767 115771      SAC% MT3
7 3770 045772      LAC% MT4      ; SHIFT BIT 16 TO BIT POSN 17
8 3771 134017      ASR 15.
9 3772 025771      ORR% MT3
10 3773 115771     SAC% MT3
11 3774 045772     LAC% MT4      ; NOW SHIFT LS WORD
12 3775 134401     ASL 1
13 3776 115772     SAC% MT4
14 3777 145764     JMP MDPLS
```

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1      ;=4000 / DEFINE SEGMENT 2 BOUNDARY
2      ;
3 4000 045701 EXECUT: LACS LIR      ;LOAD INSTRUCTION
4 4001 134401 ASL      1      ;EXTRACT PARAMETER LIST FIELD
5 4002 130400 TAN      ;TEST EVENT BUMP FLAG
6 4003 150005 JMP      2%      ;NOT SET
7 4004 136740 EXI      EVA      ;EVENT A COUNTS RASTER FRAMES
8 4005 134401 2%: ASL 1
9 4006 134006 ASR      6
10 4007 130000 TAZ
11 4010 140253 JMP      BADMOD ;DISALLOW @ PARAM BLK INDX
12 4011 001617 ADD      NINE2
13 4012 111775 SAC      MT7
14 4013 001601 ADD      MM1620 ;TEST RANGE OF PARAM LIST INDEX +9
15 4014 130400 TAN
16 4015 140253 JMP      BADMOD ;TOO LARGE
17      ;
18 4016 041617 LAC      NINE2 ;LOAD LIST LENGTH CONSTANT
19 4017 111774 SAC      MT6 ;SET INDEX TO START W/ END OF LIST
20      ;
21      ; COPY LIST TO DMA BUFFER BACK TO FRONT ORDER
22      ;
23 4020 071775 3%: LIN      MT7 ;GET LOAD INDEX
24 4021 046057 LACIS LIST-1 ;LOAD
25 4022 071774 LIN      MT6 ;GET STORE INDEX
26 4023 117703 SACSI OL-1 ;STORE
27 4024 031775 DMS      MT7 ;DECREMENT BOTH INDEXES
28 4025 031774 DMS      MT6 ;COPYING IS DONE WHEN MT6 IS = 0
29 4026 150020 JMP      3%
30      ;
31 4027 045703 LACS      OLS
32 4030 021605 ORR      HIBIT ;SET EXPMT 'SETUP' BIT
33 4031 115703 SAC      OLS ;SIGN BIT IS THE FLAG
34      ; INITIALIZE FOR TACHOGRAM BEFORE EXP. SETUP
35 4032 132000 CAC
36 4033 111675 SAC      NTACH
37 4034 111714 SAC      TOTRDH
38 4035 111715 SAC      TOTRDL
39 4036 111712 SAC      TOTELH
40 4037 111713 SAC      TOTELL
41 4040 140520 JMP      SETUP ;PERFORM EXPERIMENT SETUP
42      ;
43 4041 160447 SUBB: JST FIELDS
44 4042 055776 LAN3 TEMPJ ;NEGATIVE 'A' DATA LOADED
45 4043 141676 JMP      ADD+2 ;USE COMMON CODE IN OTHER ROUTINE
46      ;
47 4044 160447 IAND: JST FIELDS ;16-BIT .AND.
48 4045 045776 LACS      TEMPJ
49 4046 016000 ANDIS 6000
50 4047 141701 JMP      ADD+5
51
52      ; ***
53      ; USMR
54      ;
55      ; SMREAD - READ A STATUS MONITOR CHANNEL
56      ; ENTRY POINTS: USMR
57      ; ARGUMENTS: IO CONTAINS DESIRED CHANNEL

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1                                     ;          DATA RETURNED IN AC, IO, AND LOCN SMVAL
2 4050 041776 SMR1:  LAC          SMVAL
3 4051 150051 USMR: JMP
4 4052 101775          SIG          CHAN
5 4053 136500          EXI          STAT      ;REQUEST READ
6 4054 132000          CAC          ;SET TIMER
7 4055 111776          SAC          SMVAL
8 4056 136041 1$:     EXR          RSM      ;READ DONE ?
9 4057 150065          JMP          3$      ;YES
10 4050 031776         DMS          SMVAL   ;WAITED 3 SEC FOR RESPONSE ?
11 4061 150056         JMP          1$      ;KEEP CHECKING
12 4062 171425         JST          UERROR  ;NO RESPONSE IN 3 SECONDS
13 4063 061775         2$:     LIO          CHAN
14 4064 150053         JMP          USMR+2 ;TRY AGAIN
15 4065 101776         3$:     SIG          SMVAL   ;STORE DATA
16 4066 041776         LAC          SMVAL   ;COMPARE CHAN RETND
17 4067 134012         ASR          10.    ;SHOULD = REQUESTED
18 4070 132400         NOT
19 4071 001611         ADD          ONE2
20 4072 001775         ADD          CHAN
21 4073 130000         TAZ
22 4074 150050         JMP          SMR1
23 4075 150063         JMP          2$
24                                     ; ***
25 4076 136404         EXI          FANR
26 4077 150102         JMP          LEVEL3
27 4100 041572         LAC          SAVAC
28 4101 133400         RETN3: CIL
29 4102 150102         LEVEL3: JMP
30 4103 136010         EXR          IFAN    ;FIRST TEST LEV 2 INHIBIT
31 4104 150076         JMP          LEVEL3-4 ; ON, RETURN IMMEDIATELY
32
33 4105 111572         SAC          SAVAC
34 4106 136322         EXR          IPC2    ;GT2 INTERVAL
35 4107 151267         JMP          GT2FIN
36 4110 136323         L3A:     EXR          IPC3    ;GT3 INTERVAL
37 4111 151272         JMP          GT3FIN
38 4112 136140         L3B:     EXR          ISPC
39 4113 150351         JMP          QTRMIN
40 4114 136141         EXR          IMI
41 4115 150266         JMP          MINFRM
42                                     ; ***
43 4116 136143         EXR          TMA    ;MAJOR FRAME INTR.
44 4117 150100         JMP          RETN3-1 ;NO MAJOR FRAME INTR. SHOULD HAPPEN!
45                                     ; ***
46 4120 136146         EXR          ID2N   ;DAY/NIGHT TRANSITION
47 4121 140314         JMP          D2N
48 4122 136250         EXR          ILTA
49 4123 151206         JMP          LIMCHK
50 4124 136251         EXR          ILTB
51 4125 151206         JMP          LIMCHK
52 4126 136145         EXR          IPOC   ;POIC-B
53 4127 150375         JMP          VFAULT
54 4130 136147         EXR          IMP
55 4131 151450         JMP          MP
56 4132 136142         EXR          IFLA
57 4133 150263         JMP          FLAINT ; OBC 'OK' SIGNAL

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1
2 4134 136144 ; ***      EXR      ISCM      ;S/C MODE
3 4135 150100      JMP      RETN3-1
4 4136 136056      EXR      ISIM
5 4137 151262      JMP      SIMCLR ;FOR GSE USE ONLY
6 4140 136253      EXR      ISLW
7 4141 150100      JMP      RETN3-1
8 4142 136252      EXR      ISTP
9 4143 150100      JMP      RETN3-1
10 4144 150100     JMP      RETN3-1
11
12 4145 134000     DPWTBL: NOP
13 4146 136565     EXI D1N
14 4147 136567     EXI D2N
15 4150 136571     EXI D3N
16 4151 136573     EXI D4N
17 4152 136577     EXI D5N
18
19                DETTBL: ; DET INTRVL 1 / DET INTRVL 2 / ROUTING WORD /
20                ZERO2:
21 4153 000000     000000 ; OFF
22 4154 010201     010201 ; 1
23 4155 020001     020001 ; 2
24 4156 030202     030202 ; 3
25 4157 040002     040002 ; 4
26 4160 050040     050040 ; 5
27 4161 000000     000000 ; NOT USED
28 4162 000100     000100 ; TEST CLK - 16 KHZ
29 4163 012201     012201 ; 1 / 2
30 4164 012001     012001 ; 2 / 1
31 4165 014204     014204 ; 1 / 4
32 4166 014004     014004 ; 4 / 1
33 4167 023010     023010 ; 2 / 3
34 4170 023210     023210 ; 3 / 2
35 4171 034202     034202 ; 3 / 4
36 4172 034002     034002 ; 4 / 3
37
38 4173 150173     DETON: JMP
39 4174 135403     RSL      3      ;SHIFT 3 BITS TO INDEX POSN
40 4175 111774     SAC      MT6
41 4176 071774     LIN      MT6     ;LOAD 3 BIT INDEX TO DET INSTR TABLE
42 4177 062145     LIQI    DPWTBL ; LOAD INSTRUCTION
43 4200 100201     SIO      15     ;STORE INSTR. IN PLACE
44 4201 134000     15:     NOP      ;EXECUTE DET PWR UP
45 4202 011150     AND     MSG170 ;MASK OUT USED INDEX BITS
46 4203 150173     JMP      DETON
47
48 4204 150204     DETPW: JMP
49 4205 137003     LSA      3
50 4206 115771     SAC3    ACONFG ; STORE DMA ROUTING WORD
51 4207 071774     LIN      MT6     ; TABLE INDEX
52 4210 042153     LACI    DETTBL ; PICK POWER CODE BITS
53 4211 134011     ASR      9
54 4212 134412     ASL     10     ; MASK AND MOVE FOR INDEX
55 4213 170173     JST     DETON ; PASS ARG. IN AC
56 4214 170173     JST     DETON ; TURN UP TO 2 DETS ON
57 4215 150204     JMP     DETPW ; RETURN

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1
2 4216 150216 DETLIM: JMP ;SET MAX ALLOWED DET COUNTS
3 4217 137003 LSA 3 ;FOR SPECIFIED GATE TIME
4 4220 045714 LAC$ OL+8.
5 4221 011616 AND SEVEN2
6 4222 111774 SAC MT6
7 4223 071774 LIN MT6
8 4224 042243 LACI CTLIMIT
9 4225 111772 SAC MT4
10 4226 045714 LAC$ OL+8.
11 4227 134010 ASR 8.
12 4230 111773 SAC MT5
13 4231 165634 JST MLTPLY
14 4232 041771 LAC MT3 ;MS WORD OF RESULT
15 4233 130000 TAZ ;IF NOT 0, LIMIT MUST BE OVFL
16 4234 150237 JMP 1$
17 4235 041610 LAC NEG1
18 4236 150240 JMP 2$
19 4237 041772 1$: LAC MT4
20 4240 134001 2$: ASR 1 ;DIVIDE TO BYPASS SIGN COMPLICATIONS
21 4241 111676 SAC LTDCTS
22 4242 150216 JMP DETLIM
23 4243 000076 CTLIMIT: 62. ;FOR 62.5 MICRO SEC PERIOD
24 4244 000764 500.
25 4245 017500 8000.
26 4246 076400 32000.
27 4247 177777 177777
28
29 4250 150250 EXPON2: JMP ;ARGUMENT IN AC ( 6 BITS
30 4251 130000 TAZ ;TEST IF ZERO
31 4252 150260 JMP 2$ ;ZERO SO SET IT NEG.
32 4253 020262 ORR 3$ ;ARG. ) 0
33 4254 110256 SAC 1$
34 4255 041611 LAC ONE2
35 4256 134400 1$: ASL 0 ;THIS INSTR GETS MODIFIED SHIFT SIZE
36 4257 150250 JMP EXPON2 ;ARG RETURNED IN ACC
37 4260 051611 2$: LAN ONE2
38 4261 150250 JMP EXPON2
39 4262 134400 3$: ASL 0
40
41 ; FLARE COMMAND(= OBC 'OK')-- CLEAR COUNTER
42 4263 132000 FLAINT: CAC
43 4264 111637 SAC CNTOBC
44 4265 150100 JMP RETN3-1 ; EXIT LEVEL 3 INTR. ROUTINE
45
46 4266 041265 MINFRM: LAC SAVSA ;SAVE SEG ADDR
47 4267 131400 TSA
48 4270 111265 SAC SAVSA ;LOC'N EXECUTED ON EXIT
49 4271 041637 LAC CNTOBC ;CHECK HOW LONG SINCE LAST 'OK' SIGNAL
50 4272 001611 ADD ONE2
51 4273 111637 SAC CNTOBC ;COUNT OF FRAMES SINCE LAST 'OK'
52 4274 001607 ADD TOP10H ;SEE HOW LONG SINCE LAST
53 4275 130400 TAN ;'OK' FROM OBC
54 4276 141647 JMP CLOSE
55 4277 170301 JST MIFRM
56 4300 151265 JMP SAVSA
57

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1 4301 150301 MIFRM: JMP
2 4302 137003 LSA 3
3 4303 045773 LAC$ MICKL ; COUNTER MICKL IS RESET BY DAY INTR
4 4304 001611 ADD ONE2 ; INCREMENTED EACH MINOR FRAME
5 4305 115773 SAC$ MICKL ; ROLLOVER PERIOD IS 69 MINUTES
6
7 4306 041702 LAC DIRSEL ; SET DIRSEL AT MINOR FRAME
8 4307 011613 AND THREE2
9 4310 021630 ORR 050000
10 4311 111702 SAC DIRSEL
11 4312 150301 JMP MIFRM
12
13 ; **** QUARTER MINOR FRAME INTERRUPT ****
14 ; CONTROLS OUTPUT ON TM BUSES
15 4313 135402 SFTDIR: RSL 2 ; CORRECT SHIFTING DONE
16 4314 134401 RDIRCT: ASL 1 ; SHIFT FLAG BITS
17 4315 111702 SAC DIRSEL ; FLAG NEXT CHANNEL
18 4316 061573 LIO SAVIO ; RESTORE REG
19
20 4317 150317 @MIFRM: JMP
21 4320 101573 SIO SAVIO
22
23 4321 041702 LAC DIRSEL ; TEST DIRECTION POINTER
24 4322 135002 RSR 2 ; FIRST TEST IF ANY OUTPUT
25 4323 130400 TAN
26 4324 150313 JMP SFTDIR ; NO OUTPUT WAITING
27 4325 135402 RSL 2
28 4326 130400 TAN ; NEG SHOWS 3 NEXT
29 4327 150334 JMP 1$
30
31 4330 061760 LIO B3BUF ; SEND WORD ON TMB3
32 4331 136031 EXR JRO3
33 4332 150314 JMP RDIRCT
34 4333 150331 JMP .-2
35
36 4334 061757 1$: LIO B2BUF ; 1/4 MINOR FRM INTR
37 4335 136031 EXR JRO2 ; TRY DATA TRANSFER
38 4336 150314 JMP RDIRCT ; DONE
39 4337 150314 JMP RDIRCT
40
41 4340 160447 MITIM: JST FIELDS ; GET 'B' INDEX
42 4341 045773 LAC$ MICKL ; GET CLOCK VALUE
43 4342 116000 SAC$I 6000 ; STORE IN 'B'
44 4343 140162 JMP MONITR
45
46 4344 000777 0777: 777
47 4345 150345 @MTIM: JMP ; SET DELAY WORD
48 4346 021647 ORR DLYP ; AC PASSES THE AMT
49 4347 111647 SAC DLYP ; DLYP IS USED FOR CONVIENENCE
50 4350 150345 JMP @MTIM
51
52 4351 170317 @TRMIN: JST @MIFRM
53 4352 041647 LAC DLYP
54 4353 130000 TAZ
55 4354 155651 JMP GETSCI ; NO DELAY, BYPASS
56 4355 134001 ASR 1
57 4356 111647 SAC DLYP ; BITS SHIFT RIGHT

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1 4357 155651      JMP      GETSCI
2
3                ; UWTX
4
5                ; WAIT - 8 MS OR USER SPECIFIED WAIT
6                ; ENTRY POINTS: UWT8,UWTX
7                ; ARGUMENTS: NONE FOR UWT8
8                ; FOR UWTX; THE NUMBER OF 48 MICROSEC IN AC
9
10 4360 150360     UWTX: JMP      ; ENTRY
11 4361 111775     SAC      MT7      ; AC CONTAINS NO OF CYCLES
12 4362 031775     1$:      DMS      MT7      ; COUNT DOWN CYCLES
13 4363 150362     JMP      1$
14 4364 150360     JMP      UWTX ; DONE; RETURN
15
16 4365 150365     UWT8: JMP
17 4366 041577     LAC      T8MILI ; 8 MS WAIT
18 4367 170360     JST      UWTX
19 4370 150365     JMP      UWT8
20
21 4371 150371     UWTONE: JMP
22 4372 041630     LAC      050000
23 4373 170360     JST      UWTX
24 4374 150371     JMP      UWTONE
25
26 4375 041704     VFAULT: LAC      PWRERR
27 4376 001611     ADD      ONE2
28 4377 111704     SAC      PWRERR
29 4400 150100     JMP      RETN3-1
30
31 4401 150401     GET8:  JMP      ; GET LS 8 BITS OF POSN FROM DS
32 4402 136440     EXI      RDS
33 4403 101774     SIO      MT6
34 4404 041774     LAC      MT6      ; POSN IS BITS 14-7
35 4405 134006     ASR      6
36 4406 011576     AND      LOW8      ; MASK THEM
37 4407 150401     JMP      GET8
38
39 4410 150410     ASSMBL: JMP      ; ASSEMBLE AND EXECUTE INSTRUCTION
40 4411 100423     SIO      2$      ; WHOSE OP CODE IS IN IO REG
41 4412 041701     LAC      AB      ; FIRST STORE OP CODE
42 4413 134012     ASR      10.     ; CALCULATE SEGMENT TO USE
43 4414 021600     ORR      LSAWRD ; ASSEMBLE LSA + SECTOR
44 4415 110422     SAC      1$      ; STORE LSA
45 4416 071701     LIN      AB      ; ASSEMBLE REST OF INSTRC FROM
46 4417 120423     SIN      2$      ; LOW 10 BITS OF ADDRESS
47 4420 137000     LSA      0
48 4421 045707     LAC$     DATA ; LOAD AC IN CASE INSTRC IS SAC
49 4422 137000     1$:      LSA      0      ; MODIFIED LOCATION
50 4423 000000     2$:      0      ; INSTRUCTION SLOT
51 4424 150410     JMP      ASSMBL ; RETURN
52
53 4425 150425     INCAB:  JMP      ; INCREMENT AB
54 4426 041701     LAC      AB
55 4427 001611     ADD      ONE2
56 4430 011635     AND      07777 ; AB IS 12 BITS
57 4431 111701     SAC      AB

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```

1 4432 150425      JMP      INCAB      ; RETURN
2
3      ;
4      ;--COMPENSATE DETECTOR DATA BY FACTOR IN PARAMETER BLOCK BY:
5      ;      ADJPC2 = ((SCF + ADJ) * PC2) / SCF
6      ;      ADJPC3 = ((SCF - ADJ) * PC3) / SCF
7 4433 150433      DUALP: JMP
8 4434 137003      LSA 3      ;DET. DATA * .5
9 4435 045714      LAC$      OL+8.      ;COMPENSATION FOR SERVO & VEL
10 4436 134410     ASL      8.
11 4437 021634     ORR      03777      ;SET EXTEND BITS FOR LATER SHIFT
12 4440 130400     TAN
13 4441 150444     JMP      1$      ;NOT NEG SO SHIFT OFF EXTEND BITS
14 4442 135013     RSR      11.      ;ROTATE EXTEND BITS INTO HIGH ORDER
15 4443 150445     JMP      2$
16 4444 134013     1$: ASR      11.      ;REMOVE EXTEND BITS
17 4445 111766     2$: SAC      SCOUNT      ;SAVE ADJUST FACTOR
18 4446 061650     LIO      SPC1      ;PASS PC2 DATA
19
20 4447 170566     JST      ADJPC
21 4450 111776     SAC      DOWNLO      ;ADJPC2 IS FOUND
22
23 4451 051766     LAN      SCOUNT      ;DIFFERENT FOR PC3
24 4452 061651     LIO      SPC2
25
26 4453 170566     JST      ADJPC
27 4454 111766     SAC      SCOUNT      ;ADJPC3 IS FOUND
28 4455 051651     LAN      SPC2      ;TEST IF RED .GT. THRESHOLD
29 4456 137003     LSA 3
30 4457 005720     ADD$     THRSHI
31 4460 130400     TAN
32 4461 150502     JMP      VELC      ;NOT .GT. THRESHOLD
33
34 4462 051650     LAN      SPC1      ;IS .GT. THRESHOLD
35 4463 005720     ADD$     THRSHI      ;TEST IF BLUE .GT. THRESHOLD
36 4464 130400     TAN
37 4465 150502     JMP      VELC      ;NOT BRIGHT ENOUGH
38
39 4466 041737     LAC      TCYC      ;TEST IF TACH SERVO IS ON
40 4467 130000     TAZ
41 4470 150502     JMP      VELC      ;TEST VELOCITIES
42
43 4471 070153     LIN      ZERO2      ;INDEX FOR DUAL USE SUBROUTINE
44 4472 041776     LAC      DOWNLO      ;PROCESS BUS 2 (BLUE) DATA
45 4473 170613     JST      SIGMA
46
47 4474 071612     LIN      TWO2      ;INDEX FOR BUS 3 (RED) DATA
48 4475 041766     LAC      SCOUNT      ;ADJUSTED PC3 COUNTS
49 4476 170613     JST      SIGMA
50 4477 041675     LAC      NTACH      ;INCREMENT NO OF SUMMATIONS
51 4500 001611     ADD      ONE2
52 4501 111675     SAC      NTACH
53
54      ; CALCULATE LINE SHIFT TO RED USING
55      ;      ((I(R) - I(B)) * VELOC) / ((I(R) + I(B)) / 2)
56 4502 132000     VELC: CAC      ;FIRST ADD SUM FOR DENOMINATOR
57 4503 111767     SAC      MT1

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1	4504	111771	SAC	MT3	
2	4505	041776	LAC	DOWNLO	:ADJPC2
3	4506	111770	SAC	MT2	
4	4507	041766	LAC	SCOUNT	
5	4510	111772	SAC	MT4	
6	4511	165542	JST	MDPA	:ADDITION---DOUBLE PRECISION
7					
8	4512	165744	JST	MDPRS	:SCALE DENOMINATOR BY 2
9	4513	041772	LAC	MT4	
10	4514	111765	SAC	DIRECT	
11					
12	4515	051776	LAN	DOWNLO	:SET NUNERATOR
13	4516	001766	ADD	SCOUNT	
14	4517	111772	SAC	MT4	
15	4520	041575	LAC	VELOC	:SET D.P. SIGN AND UPSCALE * VELOC
16	4521	111773	SAC	MT5	:PRODUCES SCALED VELOCITY NUMBERS
17	4522	165634	JST	MLTPLY	
18	4523	041765	LAC	DIRECT	
19	4524	111773	SAC	MT5	
20					
21	4525	165665	JST	DIVIDE	:RESULT IS RED SHIFT
22					
23	4526	137003	LSA	3	
24	4527	041772	LAC	MT4	
25	4530	111765	SAC	DIRECT	
26	4531	130400	TAN		
27	4532	150547	JMP	1\$:TEST FOR RED MAXIMA
28					
29	4533	005746	ADD\$	MAXBV	:BLUE DIRECTION SO TEST BLUE MAXIMA
30	4534	130400	TAN		
31	4535	150563	JMP	TINTN	:NOT A MAXIMA
32	4536	051765	LAN	DIRECT	:YES, SO UPDATE VALUE, INTEN, POSN
33	4537	115746	SAC\$	MAXBV	:UPDATE BLUE VELOCITY
34					
35	4540	041650	LAC	SPC1	:NOW INTENSITY
36	4541	115747	SAC\$	MAXBI	
37					
38	4542	041652	LAC	SPX	:NOW POSITION UPDATE
39	4543	115750	SAC\$	MAXBX	
40	4544	041653	LAC	SPY	
41	4545	115751	SAC\$	MAXBY	
42					
43	4546	150563	JMP	TINTN	:TEST FOR TACH ANALYSIS
44					
45	4547	051765	LAN	DIRECT	
46	4550	005752	ADD\$	MAXRV	:COMPARE WITH MAX RED VELOCITY
47	4551	130400	TAN		
48	4552	150563	JMP	TINTN	
49	4553	041765	LAC	DIRECT	:UPDATE INTENSITY MAX
50	4554	115752	SAC\$	MAXRV	
51					
52	4555	041651	LAC	SPC2	
53	4556	115753	SAC\$	MAXRI	
54					
55	4557	041652	LAC	SPX	:NOW POSITION UPDATE
56	4560	115754	SAC\$	MAXRX	
57	4561	041653	LAC	SPY	

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1 4562 115755          SAC$   MAXRY
2 4563 132000  TINTN: CAC
3 4564 111650          SAC     SPC1   ;CLEAR DATA (FLAG)
4 4565 150433          JMP     DUALP
5
6 4566 150566  ADJPC: JMP     ;PASSED ADJ FACTOR IN AC
7 4567 001574          ADD    SRVDCF ;CONSTANT FOR NUM/DENOM
8 4570 111773          SAC     MT5
9 4571 101772          SIO    MT4   ;SETUP FOR NUMERATOR
10 4572 165634         JST    MLTPLY
11
12 4573 041772          LAC    MT4   ;LOW PART
13 4574 111774          SAC    MT6
14 4575 134005          ASR    5
15 4576 111772          SAC    MT4
16 4577 041771          LAC    MT3   ;HIGH PART
17 4600 134413          ASL    11.
18 4601 021772          ORR    MT4
19 4602 111772          SAC    MT4
20 4603 041771          LAC    MT3
21 4604 134005          ASR    5
22 4605 111771          SAC    MT3
23 4606 041774          LAC    MT6
24 4607 134004          ASR    4
25 4610 011611          AND    ONE2  ;ROUND UP WHEN 1
26 4611 001772          ADD    MT4
27 4612 150566         JMP    ADJPC
28
29 4613 150613  SIGMA: JMP     ;SUM ADJUSTED DET DATA
30 4614 111770          SAC    MT2   ;DATA PASSED - AC
31 4615 132000          CAC
32 4616 111767          SAC    MT1
33
34 4617 043712          LACI   TOTBLH ;BLUE COUNTER HIGH PART
35 4620 111771          SAC    MT3
36 4621 043713          LACI   TOTBLL
37 4622 111772          SAC    MT4   ;LOW PART
38
39 4623 165542          JST    MDPA  ;D. P. TOTAL
40 4624 041771          LAC    MT3
41 4625 113712          SACI   TOTBLH
42 4626 041772          LAC    MT4
43 4627 113713          SACI   TOTBLL
44 4630 150613         JMP    SIGMA
45
46 4631 150631  CKINTN: JMP
47 4632 061761          LIO    TM2WD
48 4633 101765          SIO    DIRECT
49 4634 137003          LSA    3
50 4635 051765          LAN    DIRECT
51 4636 005743          ADD$   MAXIC
52 4637 130400          TAN
53 4640 150663          JMP    3$    ;TEST IF DATA IS I MAX
54 4641 105743          SIO$   MAXIC ;NO, TEST FOR LOW I
55
56 4642 041716          LAC    POSNX ;UPDATE POSITION OF MAX
57 4643 115744          SAC$   MAXIX

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1 4644 041717 LAC POSNY
2 4645 115745 SAC$ MAXIY
3
4 4646 055705 LAN$ OL+1 :TEST IF L.MIN EXPT
5 4647 021607 ORR TOP10H
6 4650 001620 ADD TWOD12
7 4651 130400 TAN
8 4652 150656 JMP 1$ :NOT MIN, UPDATE
9 4653 001612 ADD TW02
10 4654 130400 TAN
11 4655 150631 JMP CKINTN :L.MIN, NO UPDATE AT MAX
12 4656 045761 1$: LAC$ AWLDHI :UPDATE LAMBDA
13 4657 115741 SAC$ MAXIWH :EXPERIMENT I. LAMBDA
14 4660 045762 LAC$ AWLDLO :CAN BE TRANSFERRED TO PERMANENT
15 4661 115742 SAC$ MAXIWL :OFFSET BASE WORDS W/ 2 CONTROL MOVES.
16 4662 150631 JMP CKINTN
17
18 4663 055736 3$: LAN$ MINIC :TEST LO INTENSITY
19 4664 001765 ADD DIRECT
20 4665 130400 TAN
21 4666 150631 JMP CKINTN :NOT MINIMA SO RETURN
22
23 4667 041765 LAC DIRECT
24 4670 115736 SAC$ MINIC
25
26 4671 041716 LAC POSNX
27 4672 115737 SAC$ MINIX :UPDATE POSITIONS OF MINIMA
28 4673 041717 LAC POSNY
29 4674 115740 SAC$ MINIY
30
31 4675 055705 LAN$ OL+1
32 4676 021607 ORR TOP10H
33 4677 001620 ADD TWOD12
34 4700 130400 TAN
35 4701 150631 JMP CKINTN :NOT MIN EXPT
36 4702 001612 ADD TW02
37 4703 130400 TAN
38 4704 150656 JMP 1$ :L.MIN EXPT
39 4705 150631 JMP CKINTN
40
41 4706 150706 GETPRF: JMP
42 4707 041622 LAC TWOD16
43 4710 111774 SAC MT6 :SET STEPPING LIMIT
44
45 4711 136765 1$: EXI 0R5 :TAKE ONE STEP
46 4712 041631 LAC 05153 :128 MS WAIT
47 4713 170360 JST UWTX
48
49 4714 061626 LIG 03792 :CHANNEL TO TEST PLATE REF
50 4715 170051 JST USMR :GET STATUS MON DATA
51 4716 134406 ASL 6
52 4717 130400 TAN :SKIP IF REF IS SET
53 4720 150724 JMP 2$
54 4721 132000 CAC
55 4722 111735 SAC PPREFC :CLEAR REF STEP COUNTER
56 4723 150706 JMP GETPRF
57

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1 4724 031774 2%: DMS MT6 :DECR & TEST STEP LIMIT
2 4725 150711 JMP 1% :TRY ANOTHER STEP
3 4726 171425 JST UERROR :BAD ERROR -NO REF IN 16. STEPS
4 4727 150706 JMP GETPRF :RETURN
5
6 ;
7 ; POLARIMETER ROTATION DRIVER
8 4730 040344 LAC 0777 :LOAD DELAY COUNT SIZE
9 4731 170345 JST QMTIN
10 4732 041625 LAC 0600 :30 MS DELAY
11 4733 170360 JST UWTX
12
13 4734 150734 MOVPR: JMP ;SA IS = 2
14 4735 171047 JST DLYTST :WAIT FOR ANY PENDING DELAY
15
16 4736 041734 LAC QSTPSZ :LOAD THE STEP SIZE
17 4737 111766 SAC SCOUNT :SET SCRATCH COUNT
18 4740 061626 PR1: LIO 03752
19 4741 170051 JST USMR
20 4742 136765 EXI QRS :STEP PLATE
21 4743 134406 ASL 6 :EXAMINE REF BIT FROM STATUS
22 4744 130400 TAN :TEST IF SET
23 4745 150762 JMP NOPRNL :* NOT SET
24 4746 041735 LAC PRREFC :*SET-TEST REF CTR
25 4747 130000 1%: TAZ
26 4750 150757 JMP 2% :EXPECTED REF, OK
27 4751 061001 LIO MSG331 :NOT EXPECTING REF
28 4752 171353 JST USMSG :SEND ERROR MSG
29 4753 041735 LAC PRREFC
30 4754 001610 ADD NEG1
31 4755 111735 SAC PRREFC
32 4756 150747 JMP 1% :SEND ERROR UNTIL PRREFC = 0
33
34 4757 041621 2%: LAC TWOD15
35 4760 111735 SAC PRREFC :RESET REF CTR
36 4761 150773 JMP PRMTST
37
38 4762 041735 NOPRNL: LAC PRREFC :TEST REF CTR
39 4763 130000 TAZ
40 4764 150770 JMP 4% :ERROR
41 4765 001610 ADD NEG1 :NONE EXPECTED, OK
42 4766 111735 SAC PRREFC
43 4767 150773 JMP PRMTST
44
45 4770 061002 4%: LIO MSG332 :EXPECTED REF, GOT NONE
46 4771 171353 JST USMSG :SEND MSG
47 4772 170706 JST GETPRF :RETURN WITH AC = 0 IS REF, # 0 IS NON-REF
48
49 4773 031766 PRMTST: DMS SCOUNT :COUNT DOWN STEP QUAN.
50 4774 150776 JMP 6%
51 4775 150730 JMP MOVPR-4 :DONE
52 4776 041631 6%: LAC 05153 :120 MS COUNT
53 4777 170360 JST UWTX :WAIT FOR INTERVAL
54 5000 150740 JMP PR1
55
56 5001 133331 MSG331: 133331
57 5002 133332 MSG332: 133332

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1
2
3
4 5003 151003 RAXUND: JMP
5 5004 171047 JST DLYTST
6
7 5005 051720 LAN RASHIX : POSN COMPARE
8 5006 001716 ADD POSNX
9 5007 130400 TAN : TEST IF AT END
10 5010 151021 JMP 2% : YES - END ATTAINED
11 5011 041716 LAC POSNX : NO POSNX < HIGH END
12 5012 001724 ADD DELTX : FIND NEXT POSITION
13 5013 111716 SAC POSNX
14 5014 041621 LAC TWOD15 : DELAY VALUE
15 5015 170345 1%: JST QMTIM
16 5016 061716 LIO POSNX
17 5017 136610 EXI RXLD : MOVE TO NEW POSN
18 5020 151003 JMP RAXUND : RETURN
19
20 5021 041722 2%: LAC RASLOX
21 5022 111716 SAC POSNX
22 5023 041633 LAC LS10HI
23 5024 151015 JMP 1%
24
25
26 5025 151025 RAYUND: JMP : UNIDIRECTIONAL Y DRIVER
27 5026 171047 JST DLYTST : WAIT FOR ANY DELAY
28
29 5027 051721 LAN RASHIY : Y END POSITION COMPARISON
30 5030 001717 ADD POSNY : IF NEG RESULT, END NOT REACHED
31 5031 130400 TAN
32 5032 151043 JMP 2% : YES
33 5033 041717 LAC POSNY : NO END YET
34 5034 001725 ADD DELTY
35 5035 111717 SAC POSNY : SET NEXT POSN
36 5036 041626 LAC Q3752 : DELAY FOR STEP TIME
37 5037 170345 1%: JST QMTIM
38 5040 061717 LIO POSNY
39 5041 136614 EXI RYLD
40 5042 151025 JMP RAYUND : RETURN
41
42 5043 041723 2%: LAC RASLOY : GET LOW END OF Y RANGE
43 5044 111717 SAC POSNY : SET NEW Y
44 5045 041635 LAC Q7777 : DELAY FOR RETRACE
45 5046 151037 JMP 1%
46
47 5047 151047 DLYTST: JMP : DELAY SPECIFIED BY HOW MANY
48 5050 041647 1%: LAC DLYP : SHIFTS REQD TO ZERO THIS WD
49 5051 130000 TAZ
50 5052 151047 JMP DLYTST
51 5053 151050 JMP 1%
52 5054 151054 RASLMC: JMP : CALCULATE RASTER END POINTS
53 5055 113724 SACI DELTX
54 5056 111772 SAC MT4
55 5057 121767 SIN MT1
56 5060 165634 JST MLTPLY : FIND RANGE REQD FROM ARGS PASSED
57 5061 071767 LIN MT1 : RESTORE LOST INDEX

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1 5062 041772      LAC      MT4      ;TEST IF LEGAL RANGE
2 5063 134010      ASR      8.
3 5064 130000      TAZ
4 5065 151070      JMP      1$      ;RANGE GOOD
5 5066 171425      JST      UERROR  ;OVERSIZE ERROR
6 5067 140162      JMP      MONITR  ;TRY NEXT EXPERIMENT
7
8 5070 041772  1$:  LAC      MT4      ;GET SIZE
9 5071 134001      ASR      1
10 5072 111773      SAC      MT5      ;BISECT EVEN SIZE DIMENSION
11 5073 051773      LAN      MT5      ;MAKE TRIAL LOW END POINT
12 5074 001765      ADD      DIRECT
13 5075 113722      SACI     RASLOX  ;TEMPORARY SAVE LOW END
14 5076 130400      TAN
15 5077 151103      JMP      HIRAST  ;LOW END IS VALID
16 5100 111771      SAC      MT3      ;SAVE NEG OVERRANGE
17 5101 132000      CAC
18 5102 113722      SACI     RASLOX  ;SET LOW END PT TO 0
19
20 5103 043722  HIRAST: LACI RASLOX ;CALCULATE END PNT
21 5104 001772      ADD      MT4
22 5105 113720      SACI     RASHIX
23 5106 051576      LAN      LOW8
24 5107 003720      ADDI    RASHIX  ;HI .LT. LIM?
25 5110 130400      TAN
26 5111 151113      JMP      2$      ;NO, 0 ALSO ACCEPTABLE
27 5112 151121      JMP      3$      ;YES
28 5113 111771  2$:  SAC      MT3      ;SAVE CORRECTION
29 5114 051771      LAN      MT3
30 5115 003722      ADDI    RASLOX
31 5116 113722      SACI     RASLOX  ;REDUCE LOW END
32 5117 041576      LAC      LOW8
33 5120 113720      SACI     RASHIX
34 5121 051771  3$:  LAN      MT3
35 5122 137003      LSA      3
36 5123 007716      ADDI$   RAXCEN
37 5124 117716      SACI$   RAXCEN
38 5125 137002      LSA      2
39 5126 063722  4$:  LIOI    RASLOX  ;PASS BACK START X/Y
40 5127 151054      JMP      RASLMC  ;RETURN
41
42 ;          ;**** EXPMT DEFN BLOCK DUMP OVER DATA CHANNELS ****
43 ;
44 5130 151130  DEKDMP: JMP
45 5131 041702  1$:  LAC      DIRSEL  ;TEST FLAG WORD
46 5132 134416      ASL      14.     ;LOW TWO BITS SHOW ACTIVITY
47 5133 130000      TAZ
48 5134 151136      JMP      2$
49 5135 151131      JMP      1$
50
51 5136 137001  2$:  LSA      1
52 5137 064255      LIO$     LIODLR
53 5140 170410      JST      ASSMEL
54 5141 170425      JST      INCAB   ;INCREMENT AB (ADDRESS)
55
56 5142 171353      JST      USMSG
57

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1 5143 051700 LAN LAST
2 5144 001701 ADD AB
3 5145 130400 TAN ;NEG MEANS NOT DONE
4 5146 151130 JMP DBKDMP ;RETURN
5 5147 151131 JMP 1$ ;REITERATE
6
7 ;MEMORY DUMP FOR FAST OUTPUT
8 ; OUTPUTS ADDRESS AND MEMORY CONTENTS
9 ; ON BUS 2 AND BUS 3, RESPECTIVELY.
10 ;
11 5150 170000 MSG170: 170000 ;MEMORY DMP MSG. FLAG
12 5151 151151 MEMDMP: JMP
13 5152 051702 1$: LAN DIRSEL ;MIFR INTR SETS DIRSEL = 50000
14 5153 001630 ADD 050000 ;LOW 2 BITS MAY REMAIN SET
15 5154 130000 TAZ ;FIND EMPTY SLOTS AFTER MIFR
16 5155 151157 JMP 2$
17 5156 151152 JMP 1$ ;CONDITION NOT MET, TRY AGAIN
18 ;
19 5157 041701 2$: LAC AB ;SETUP AND SEND..
20 5160 021150 3$: ORR MSG170 ;DUMP MESSAGE
21 5161 111774 SAC COUNT ;ALONG WITH ADDRESS
22 5162 061774 LIO COUNT ;ON TM BUS 2
23 5163 171275 JST U2MSG
24 ;
25 5164 137001 LSA 1
26 5165 064255 LIO% LIODLR ;AN LIO INSTRUCTION
27 5166 170410 JST ASSMBL
28 5167 170425 JST INCAB ;INCREMENT AB POINTER
29 5170 171323 JST U3MSG ;OUTPUT ON TM BUS 3
30 ;
31 5171 041701 LAC AB
32 5172 130000 TAZ ;TEST IF DONE YET
33 5173 151151 JMP MEMDMP ;DUMPED LAST LDCN = 7777 OCTAL
34 5174 151160 JMP 3$ ;REPEAT UNTIL DONE
35 ;
36 5175 171254 TWLP: JST ISV3 ; TEST WLD POWER CONDITION
37 5176 137003 LSA 3
38 5177 045772 LAC% RPWD
39 5200 134414 ASL 12.
40 5201 130400 TAN
41 5202 140325 1$: JMP EXOFF ;NO WLD PWR ON
42 5203 041202 LAC 1$ ;STEP PWR ON
43 5204 111640 SAC CTLSLT ;SET DE-POWER JMP
44 5205 151265 JMP SAVSA ;AT INNER LOOP SLOT
45 ;
46 5206 171254 LIMCHK: JST ISV3 ; THIS VERSION WILL TEST MP LAST
47 5207 161612 JST DTIM
48 5210 161632 JST DEVOFF
49 5211 136250 EXR ILTA
50 5212 134000 NOP
51 5213 136251 EXR ILTB
52 5214 134000 NOP
53 5215 136440 EXI RDS
54 5216 137003 LSA 3
55 5217 105774 SIO% TEMP ; IN DMA BLOCK 7774
56 5220 045774 LAC% TEMP
57 5221 134413 ASL 11.

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1 5222 070153 LIN ZERO2 ;PRESET INDEX
2 5223 130400 TAN ;TEST LIMIT B BIT
3 5224 151237 JMP 25 ;NOT LIMIT B
4 5225 134401 ASL 1
5 5226 130400 TAN
6 5227 151243 JMP 35 ;ONLY LIM B, SO ACT
7 5230 133000 HLT ; FAULT, DSW LIM BOTH ON
8 5231 061621 1$: LIO TWOD15 ;BOTH A & B, TEST MP
9 5232 170051 JST USMR
10 5233 134410 ASL 8.
11 5234 130400 TAN
12 5235 140333 JMP BOOTX
13 5236 151264 JMP IRS3
14 ;
15 5237 134401 2$: ASL 1 ;TEST FOR ONLY LIMIT A
16 5240 130400 TAN
17 5241 151231 JMP 1$: ;NEITHER, TEST MP FOR BOOT
18 5242 071611 LIN ONE2
19 5243 041603 3$: LAC 01052
20 5244 164232 JST RWHNAB
21 5245 153246 JMPI .+1
22 5246 145225 JMP LIMITB
23 5247 145214 JMP LIMITA
24 ;
25 ;
26 5250 136656 CMD: EXI STPR ;THIS COMMAND IS CHANGED BY PROGRAM
27 5251 151251 JMP . ;THIS RETURN ALSO CHANGED
28 5252 136656 XCMD: EXI STPF ;STEP OTHER DIRECTION FROM CMD
29 5253 145042 JMP RECOVR ;USED WHEN NULL IS LOST
30 ;
31 ;SAVE REGISTERS FOR LEVEL 3 INT
32 5254 151254 ISV3: JMP
33 5255 041265 LAC SAVSA
34 5256 131400 TSA ; SA RESTORE, VARIES 0..3
35 5257 111265 SAC SAVSA
36 5260 101573 SIO SAVIO
37 ;
38 5261 151254 ; *** JMP ISV3
39 ;
40 5262 171254 SIMCLR: JST ISV3
41 5263 136057 EXR RSIM ;IRS3 MUST BE CONTIGUOUS
42 ; ***
43 5264 061573 IRS3: LIO SAVIO
44 5265 137000 SAVSA: LSA 0
45 5266 150100 JMP RETN3-1
46 ;
47 5267 041605 GT2FIN: LAC HIBIT ;NEG BIT
48 5270 111761 SAC TM2WD ;FLAG SPECIAL LOC'N
49 5271 150110 JMP L3A
50 ;
51 5272 041605 GT3FIN: LAC HIBIT ;NEG BIT
52 5273 111762 SAC TM3WD ;FLAG SPECIAL LOC'N
53 5274 150112 JMP L3B
54 ; DIRSEL:16 0 BUS 2 OUT NEXT
55 5275 151275 U2MSG: JMP ; 1 BUS 3 OUT NEXT
56 5276 136403 1$: EXI FANS
57 5277 041702 LAC DIRSEL ;TEST IF B2 IS IMMED NEXT

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1 5300 130400      TAN
2 5301 135001      RSR 1 ;USE BIT 2 FOR MARK (YES)
3 5302 135001      RSR 1 ;USE BIT 1 FOR MARK (NO)
4 5303 130400      TAN ;TEST IF B2 IS EMPTY
5 5304 151307      JMP 2$ ;EMPTY
6 5305 171414      JST CHK25
7 5306 151276      JMP 1$
8
9 5307 041702      2$: LAC DIRSEL
10 5310 001611      ADD ONE2
11 5311 130400      TAN ;SKIP TO MARK BIT 1
12 5312 001611      ADD ONE2 ;INTEND TO MARK BIT 2
13 5313 171315      JST B2IT ;USE COMMON OUT ROUTINE
14 5314 151275      JMP U2MSG
15
16 5315 151315      B2IT: JMP
17 5316 101757      SIO B2BUF ;STORE OUTPUT WORD
18 5317 111702      SAC DIRSEL
19 5320 171414      JST CHK25
20 5321 137002      LSA 2 ;FOR COMPATIBILITY
21 5322 151315      JMP B2IT
22
23 5323 151323      U3MSG: JMP
24 5324 136403      1$: EXI FANS
25 5325 041702      LAC DIRSEL ;WHAT IS NEXT
26 5326 130400      TAN ;SKIP IF B3 IS NEXT
27 5327 135401      RSL 1 ;(IT ISN'T)USE BIT 1 FOR MARK
28 5330 135002      RSR 2 ;USE BIT 2 FOR MARK
29 5331 130400      TAN
30 5332 151335      JMP 2$
31 5333 171414      JST CHK25
32 5334 151324      JMP 1$ ;WAIT FOR EMPTY
33 5335 041702      2$: LAC DIRSEL
34 5336 130400      TAN
35 5337 151342      JMP 3$
36 5340 021612      ORR TWO2
37 5341 151343      JMP 4$
38 5342 021611      3$: ORR ONE2
39 5343 171345      4$: JST B3IT
40 5344 151323      JMP U3MSG
41
42 5345 151345      B3IT: JMP
43 5346 101760      SIO B3BUF ;STORE OUT WORD
44 5347 111702      SAC DIRSEL
45 5350 171414      JST CHK25
46 5351 137002      LSA 2
47 5352 151345      JMP B3IT
48
49 5353 151353      U5MSG: JMP
50 5354 136403      1$: EXI FANS ;INHIBIT LEVEL 3 INTS
51 5355 041702      LAC DIRSEL
52 5356 135001      RSR 1 ;TEST IF SECOND SLOT USED
53 5357 130400      TAN
54 5360 151363      JMP 2$ ;BIT 1 NOT SET
55 5361 171414      JST CHK25
56 5362 151354      JMP 1$ ;REPEAT TEST
57

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1 5363 135001 2$: RSR 1 ;BIT 2 SET?
2 5364 130400 TAN ;TEST IF NEXT SLOT USED
3 5365 151376 JMP 5$ ;NO, FILL NEXT
4 5366 135402 RSL 2 ;YES/CORRECT SHIFTS
5 5367 021611 ORR ONE2 ;SET BIT 1
6 5370 130400 TAN ;DETERMINE CROSS FILL
7 5371 151374 JMP 4$
8 5372 171315 3$: JST B2IT
9 5373 151353 JMP USMSG
10 5374 171345 4$: JST B3IT
11 5375 151353 JMP USMSG
12
13 5376 135402 5$: RSL 2
14 5377 021612 ORR TWO2
15 5400 130400 TAN
16 5401 151372 JMP 3$
17 5402 151374 JMP 4$
18
19 5403 151403 SNDMSG: JMP ;COMMONLY USED CODE PASSES MSG IN AC
20 5404 111773 SAC MTS ;MOVE MESSAGE TO IO REG.
21 5405 061773 LIO MTS
22 5406 171353 JST USMSG ;UTILITY MSG ROUTINE
23 5407 151403 JMP SNDMSG ;RETURN
24
25 5410 136404 EXI FANR
26 5411 136140 EXR ISPC
27 5412 151416 JMP A25HZ
28 5413 133400 CIL
29 5414 151414 CHK25: JMP
30 5415 151410 JMP -5
31
32 5416 170317 A25HZ: JST @MIFRM
33 5417 041647 LAC DLYP
34 5420 130000 TAZ
35 5421 151413 JMP CHK25-1
36 5422 134001 ASR 1
37 5423 111647 SAC DLYP
38 5424 151413 JMP CHK25-1
39
40 ; UERROR
41 ;
42 ; ERROR ROUTINE
43 ; ENTRY POINTS: UERROR
44 ; ARGUMENTS: UERROR CONTAINS ADDRESS OF ERROR
45 ;
46 5425 151425 UERROR: JMP
47 5426 061602 LIO ERRMSG ;OUTPUT 'ERROR' MSG
48 5427 171353 JST USMSG
49 5430 137003 LSA 3
50 5431 041425 LAC UERROR
51 5432 134403 ASL 3 ;PUT BITS 13-12 IN POSNS 12-11
52 5433 134004 ASR 4
53 5434 111771 SAC MT3 ;TEMPORARY STORE
54 5435 045764 LAC$ ERRWRD ; THIS ADDS 10000 TO DMA LOCN
55 5436 011606 AND TOPFOR ; TO HELP COUNT ERRORS. CYCLE LENGTH = 16.
56 5437 001604 ADD 010K
57 5440 021771 ORR MT3

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1 5441 115764      SAC$  ERRWRD
2 5442 071425      LIN   UERROR  ;PUT BITS 10-1 IN POSNS 10-1
3 5443 137402      DIN
4 5444 125764      SIN$  ERRWRD
5 5445 065764      LIOS  ERRWRD  ;OUTPUT ERROR WORD
6 5446 171353      JST   USMSG
7 5447 151425      JMP   UERROR
8
9 5450 040102      MP:   LAC   LEVEL3 ;LOG MEM PRO ERROR LOCN
10 5451 011633     AND   LS10HI ;MASK LOW 10 BITS IN
11 5452 111705     SAC   MPERR
12 5453 041613     LAC   THREE2
13 5454 134413     ASL   11.
14 5455 010102     AND   LEVEL3 ;GET SEGMENT
15 5456 134001     ASR   1
16 5457 021705     ORR   MPERR  ;COMBINE
17 5460 001610     ADD   NEG1
18 5461 111705     SAC   MPERR  ;SAVE FOR ANALYSIS
19 5462 150100     JMP   RETN3-1
20
21 5463 151463     HEATR: JMP
22 5464 041677     LAC   HEATWD
23 5465 011576     AND   LOW8   ;CLEAR FLAG BIT
24 5466 111677     SAC   HEATWD
25 5467 134410     ASL   8.
26 5470 130400     TAN
27 5471 151474     JMP   .+3
28 5472 136661     EXI   H1N
29
30 5473 151463     JMP   HEATR
31 5474 134401     ASL   1
32 5475 130400     TAN
33 5476 151463     JMP   HEATR
34 5477 136660     EXI   H1F
35 5500 151463     JMP   HEATR
36
37 5501 045700     INLHTR: LAC$  SCI   ;CORRECTED FROM LAST TIME
38 5502 111677     SAC   HEATWD ;(THIS WAS MISSING)
39 5503 045772     LAC$  RPWD  ;HEATER CHANGE AC=(SCI)
40 5504 134001     ASR   1     ;TEST POWER BITS
41 5505 130000     TAZ
42 5506 171463     JST   HEATR ;DON'T NEED LOW BIT
43 5507 140122     JMP   INRTN ;SKIPS HEATR IF EXPMT IN PROG
44
45 5510 136634     SLSF: 136634
46 5511 136630     SLSR: 136630
47 5512 161632     SIZE: JST   DEVOFF ;TURN ALL HRDWR OFF
48 5513 136624     EXI  SLTN
49 5514 161612     JST   DTIM  ;DETS OFF
50 5515 170371     JST   UHTONE
51 5516 045701     LAC$  LIR
52 5517 134011     ASR   9.
53 5520 011626     AND   03752
54 5521 115776     SAC$  TEMPJ
55 5522 055776     LAN$  TEMPJ
56 5523 001624     ADD   D2052
57 5524 130400     TAN

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1 5525 151527 JMP 1%
2 5526 136622 EXI SLOS
3 5527 001612 1%: ADD TW02
4 5530 130400 TAN
5 5531 151533 JMP 2%
6 5532 151570 JMP SIZOFF
7 5533 164210 2%: JST RDSIZ
8 5534 111770 SAC MT2
9 5535 130000 TAZ
10 5536 151570 JMP SIZOFF
11 ;
12 5537 061510 4%: LIO SLSF
13 5540 055776 LAN$ TEMPJ
14 5541 001770 ADD MT2
15 5542 130400 TAN
16 5543 061511 LIO SLSR
17 5544 101551 SIO 5%
18 5545 130000 TAZ
19 5546 151570 JMP SIZOFF
20 5547 041770 LAC MT2
21 5550 111771 SAC MT3
22 5551 136634 5%: EXI SLSF ; VARIABLE INSTRUCTION
23 5552 045773 LAC$ MICKL
24 5553 111772 SAC MT4 ; START TIME
25 ;
26 5554 164210 6%: JST RDSIZ
27 5555 111770 SAC MT2
28 5556 051771 LAN MT3
29 5557 001770 ADD MT2
30 5560 130000 TAZ
31 5561 151563 JMP 7%
32 5562 151537 JMP 4%
33 ;
34 5563 055773 7%: LAN$ MICKL
35 5564 001772 ADD MT4
36 5565 001576 ADD LOW8
37 5566 130400 TAN
38 5567 151554 JMP 6%
39 5570 136620 SIZOFF: EXI SLTF
40 5571 140162 JMP MONITR
41 ;
42 5572 000000 SAVAC: 0
43 5573 000000 SAVIO: 0
44 ; ***
45 ;
46 ; 16 MS COUNT OF WAIT. ACCURACY -1, +0
47 ; DELPR: 9. ; POL ROTATE )128 MS
48 ; DELSX: 4 ; VALUE FOR X DELAY
49 ; DELSY: 5 ; VALUE FOR Y
50 ; DELRX: 8.
51 ; DELRY: 12. ; VALUE FOR Y RETRACE DELAY
52 ;
53 5574 000040 SRVOCF: 32. ; 30. GIVES 7 PERCENT CHANGES TO 3:1
54 5575 000024 VELOC: 20. ; MULTIPLIER FOR USER VELOCITY DATA
55 ;
56 5576 000377 LOW8: 377
57 5577 000243 TSMILI: 163.

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1 5600 137000 LSAWRD: LSA 0
2 ; ***
3 5601 176204 MM1620: -1574
4 5602 155555 ERRMSG: 155555
5 5603 000010 O10S2: 10
6 5604 010000 O10K: 10000
7 5605 100000 HIBIT: 100000
8 5606 170000 TOPFOR: 170000
9 5607 177700 TOP10H: 177700
10 5610 177777 NEG1: -1
11 5611 000001 ONE2: 1
12 5612 000002 TWO2: 2
13 5613 000003 THREE2: 3
14 5614 000004 FOUR2: 4
15 5615 000005 FIVE2: 5
16 5616 000007 SEVEN2: 7
17 5617 000011 NINE2: 9
18 5620 000014 TWOD12: 12
19 5621 000017 TWOD15: 15
20 5622 000020 TWOD16: 16
21 5623 002161 D1137: 1137
22 5624 000024 D2052: 20
23 5625 001130 D600: 600 ;FOR 30 MS WAIT
24 5626 000037 O3752: 37
25 5627 000077 O77D2: 77
26 5630 050000 O50000: 50000
27 5631 005153 O5153: 5153 ;COUNT FOR 120 MS WAIT
28 5632 057330 O57330: 57330
29 5633 001777 L510HI: 1777
30 5634 003777 O3777: 3777
31 5635 007777 O7777: 7777 ;LOW 12 BIT MASK
32 5636 077777 O77777: 77777
33 ;
34 5637 000000 CNTOBC: 0 ;COUNTER--TIME SINCE LAST OBC 'OK' SIGNAL
35 ;
36 5640 000000 CTLSLT: 0
37 ;=.45 ;CONTROL SLOT AREA
38 ;
39 5646 000000 DELAY: 0 ;ALSO CONTROL SLOT
40 5647 000000 DLYP: 0
41 5650 000000 SPC1: 0
42 5651 000000 SPC2: 0
43 5652 000000 SPX: 0
44 5653 000000 SPY: 0
45 5654 000000 BOTH: 0
46 ;
47 ;=5675
48 5675 000000 NTACH: 0
49 5676 000000 LTDCTS: 0
50 5677 000000 HEATWD: 0
51 5700 000000 LAST: 0
52 5701 000000 AB: 0
53 5702 000000 DIRSEL: 0
54 5703 000000 PASCNT: 0
55 5704 000000 PWRERR: 0
56 5705 000000 MPERR: 0
57 5706 000001 PINDX: 1

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1 5707 000002 XINDX: 2
2 5710 000003 YINDX: 3
3 5711 000000 DOWNHI: 0
4          ;THE FOLLOWING VARIABLE PAIRING IS CRITICAL
5 5712 000000 TOTBLH: 0
6 5713 000000 TOTELL: 0
7 5714 000000 TOTRDH: 0
8 5715 000000 TOTRDL: 0
9 5716 000000 POSNX: 0
10 5717 000000 POSNY: 0
11 5720 000000 RASHIX: 0
12 5721 000000 RASHIY: 0
13 5722 000000 RASLOX: 0
14 5723 000000 RASLOY: 0
15 5724 000000 DELTX: 0
16 5725 000000 DELTY: 0
17          ;***** PAIRS END HERE
18 5726 000000 XCYSZ: 0
19 5727 000000 RXCYCL: 0
20 5730 000000 YCYSZ: 0
21 5731 000000 RYCYCL: 0
22 5732 000000 QCYCSZ: 0
23 5733 000000 PRCYCL: 0
24 5734 000000 QSTPSZ: 0
25 5735 000000 PRREFC: 0
26 5736 000000 TCYCSZ: 0
27 5737 000000 TCYC: 0
28 5740 000000 DELTAL: 0
29 5741 000000 WLCYSZ: 0
30 5742 000000 WLCYC: 0
31 5743 000000 DWLDHI: 0
32 5744 000000 DWLDLO: 0
33 5745 177777 PRIOR: -1
34          NULLHI=5746
35          NULLLO=5747
36          .=5750
37 5750 000061 MOTOR: 49
38 5751 000000 DSLAT: 0
39 5752 000000 DSOLD: 0
40 5753 000000 CALSIZ: 0
41 5754 000000 CCYCSZ: 0
42 5755 000000 CALLAM: 0
43 5756 000000 CALCYC: 0
44 5757 000000 B2BUF: 0
45 5760 000000 B3BUF: 0
46 5761 000000 TM2WD: 0
47 5762 000000 TM3WD: 0
48          WANTHI:
49 5763 000000 OLD: 0
50 5764 000000 WANTLO: 0
51 5765 000000 DIRECT: 0
52 5766 000000 SCOUNT: 0
53 5767 000000 MT1: 0
54 5770 000000 MT2: 0
55 5771 000000 MT3: 0
56 5772 000000 MT4: 0
57 5773 000000 MT5: 0

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1      MTS:
2 5774 000000 COUNT: 0
3      MT7:
4 5775 000000 CHAN: 0
5      DOWNLO:
6 5776 000000 SMVAL: 0
7 5777 000000 PFFLAG: 0
8      ;
9      ;
10 6000 000000 VECTOR: 0
11      ;
12 6036 000400 400      ;VECTOR 0'16'
13      ;
14 6060 000020 LIST: 16 ;LIST A
15 6061 110040 OBSLST: 110040 ;OBSERVING LISTS BEGIN HERE
16      ;
17 6461 110040 110040
18 6462 002013 002013
19 6463 120205 120205
20 6464 100000 100000
21 6465 000000 0
22 6466 000004 4
23 6467 000000 0
24 6470 036074 036074
25 6471 020400 020400
26 6472 077001 077001
27      ;
28      ; ***
29      ;
30      ; LEVEL 3 16 MS INTERRUPT TEST FOR SCI CMD
31 7651 041700 GETSCI: LAC SCI ;SEG 3 TEST FOR SPEED
32 7652 130400 TAN
33 7653 155655 JMP .+2
34 7654 150100 JMP RETN3-1
35 7655 171254 JST ISV3
36 7656 140020 JMP MONITI ; LOOK AT SCI CMD
37      ; 'TYPE 1' SCI: DOOR/SLIT/POLR COMMAND TO JR
38 7657 135460 DEVCD: EXI 060 ;DOOR
39 7660 136620 EXI 220 ;SLIT
40 7661 136760 EXI 360 ;POLRM
41 7662 136660 EXI 260 ;HEATR + INSTR PWR INVERTER
42 7663 137002 DVISE: LSA 2
43 7664 134006 ASR 6 ;SERIES CODE(I.E. DOOR/SLIT/POLR)
44 7665 015613 AND$ THREE3
45 7666 111774 SAC TEMP
46 7667 071774 LIN TEMP
47 7670 041700 LAC SCI
48 7671 015621 AND$ TWOD15 ;SUB-SERIES CODE
49 7672 023657 ORRI DEVCD ;GET FULL COMMAND
50 7673 111675 SAC .+2
51 7674 061704 LID OL ;WANTED SLIT #(PREV. SET IF 'SLLD' COMMAND)
52 7675 134000 NOP ; ** THIS IS THE DOOR/SLIT/POLR COMMAND**
53 7676 140307 JMP WAIT ;DONE
54      ; ***
55      ;
56 7677 000001 EXPADV: 1
57 7700 106020 SCI: 106020

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```
1 7701 000000 LIR: 0
2 7702 000000 LPC: 0
3 7703 000000 OLS: 0
4 7704 000000 OL: 0
5
6 EXPNUM=7715
7 .=7716
8 7716 000177 RAXCEN: 177
9 7717 000177 RAYCEN: 177
10 7720 000005 THRSHI: 5 ; RAW DATA MUST EXCEED 2 * 5
11 7721 000000 MAXHI: 0
12 7722 000000 MAXLO: 0
13
14 .=. +11.
15 FLAG=7733
16
17 7736 000000 MINJC: 0
18 7737 000000 MINIX: 0
19 7740 000000 MINIY: 0
20 7741 000000 MAXIWH: 0
21 7742 000000 MAXIWL: 0
22 7743 000000 MAXIC: 0
23 7744 000000 MAXIX: 0
24 7745 000000 MAXIY: 0
25 7746 000000 MAXBV: 0
26 7747 000000 MAXEI: 0
27 7750 000000 MAXBX: 0
28 7751 000000 MAXBY: 0
29 7752 000000 MAXRV: 0
30 7753 000000 MAXRI: 0
31 7754 000000 MAXRX: 0
32 7755 000000 MAXRY: 0
33
34 7756 000000 OFFTTL: 0
35 7757 000000 FIX: 0
36 7760 000000 LASTFX: 0
37 AWLDHI=7761
38 AWLDLO=7762
39 .=7763
40 7763 000000 WLSCAN: 0
41 7764 000000 ERRWRD: 0
42 7765 000000 CPROCYC: 0
43 7766 000000 CXCYCL: 0
44 7767 000000 CYCYCL: 0
45 7770 000000 REPEAT: 0
46 7771 000000 ACONFG: 0
47 7772 000000 RPWD: 0
48 7773 000000 MICLK: 0
49 7774 000000 TEMP: 0
50 7775 000000 TEMPI: 0
51 7776 000000 TEMPJ: 0
52 7777 000020 BSYGN: 16.
53
54 .END
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DTIM	1612	1-29	5- 3	5-44	8-40	20-55#	21-13	25-47	57-47	61-49
DUALP	4433	16-38	24-22	50- 6#	52- 4					
DUMP	425	6-57	7-30#							
DVISE	7663	1-53	65-42#							
DWLDHI	5743	11-44	16-19	19- 8	27-10	27-24	27-35	27-50	64-31#	
DWLDLO	5744	11-43	16-21	19- 9	27- 2	27-25	27-36	27-51	64-32#	
DWFGS	2274	11-51	18-25	19-10	26-53#	27-12	27-53			
DWSFIN	1415	18-10#	18-37							
DWSTP	1434	10-19	18-24	18-29#	18-34					
EDBLO	1756	14-31	22-46#							
ENDSTP	2775	32-31	32-43	32-49	32-57#					
ENDWT	1753	2-11	22-49#							
ERRMSG	5602	60-47	63- 4#							
ERRWRD	7764	60-54	61- 1	61- 4	61- 5	66-41#				
EXECUT	4000	4-28	44- 3#							
EXOFF	325	5-42#	57-41							
EXPADV	7577	3- 8	10-51	10-52	14-20	14-25	14-29	65-56#		
EXPNUM	7715	14-23	14-24	66- 6						
EXPON2	4250	10-42	12-11	47-29#	47-36	47-38				
FIEL1	434	7-43#	8- 7	8-45						
FIEL2	440	7-48#	8-12							
FIELDS	447	7-55#	8-36	21-45	21-54	22- 5	44-43	44-47	48-41	
FILSLT	625	10-10#	10-17							
FINAL	2957	27-50#								
FIVE2	5615	63-15#								
FIX	7757	38-38	38-53	39-11	39-19	39-36	39-52	66-35#		
FIXWL	3427	2-39	38-32#							
FLAG	7733	2-45	2-47	8-47	8-49	23-27	25-48	25-50	66-15	
FLAINT	4263	45-57	47-42#							
FLFLAG	111	2-36	2-44#							
FLRTST	2010	23-19	23-27#							
FOUR	1716	3-15	22-18#							
FOUR1	2175	25-40#	33-33	36- 8						
FOUR2	5614	63-14#								
GETB	4401	6-29	7- 6	7-12	7-17	7-22	35- 7	35-17	35-28	49-31# 49-37
GETPRF	4706	3-32	20-30	53-41#	53-56	54- 4	54-47			
GETSCI	7651	48-55	49- 1	65-31#						
GON	506	8-30	8-32#	21-52						
GONEW	127	3- 4#								
GOTOO	476	4-30	8-24#							
GT2FIN	5267	45-35	58-47#							
GT3FIN	5272	45-37	58-51#							
GTBUSY	3073	4-55	30-18	32-32	32-50	34- 9#	34-16	35-42	35-52	38- 2
GTGWAV	3364	29-10	37-41#							
HALT	12	1-28#	1-33	2-16	5-22	5-42	36-57			
HEATR	5463	5-25	16-56	61-21#	61-30	61-33	61-35	61-42		
HEATWD	5677	5-23	16-54	61-22	61-24	61-36	63-50#			
HELP4	3021	32-33	33-21#							
HELPS	3027	32-51	33-28#							
HELPER	3071	34- 6#								
HIBIT	5605	41-53	41-56	42-51	44-32	58-47	58-51	63- 7#		
HIRAST	5103	56-15	56-20#							
IAND	4044	4-39	44-47#							
INCAB	4425	7-27	49-53#	50- 1	56-54	57-28				
INJMP	44	1-50	1-57#	2- 4	2- 8	2-34				
INLHTR	5501	2-31	61-37#							
INLINE	71	1-41	2-26#	3- 5						
INRTN	122	2-42	2-48	2-55#	39-13	61-43				
IRS3	5264	58-13	58-43#							

POSNX	5716	13-42	23-11	52-56	53-26	55- 8	55-11	55-13	55-16	55-21	64- 9#				
POSNY	5717	13-19	23-13	53- 1	53-28	55-30	55-33	55-35	55-38	55-43	64-10#				
POWRUP	260	1-12	4-53#	5-32	37- 1										
PR1	4740	54-18#	54-54												
PRCFNS-	1560	20-21#	20-44												
PRCTRL	1575	10-22	20-36#	20-41											
PRCYCL	5733	12-26	20-22	20-38	20-47	64-23#									
PRIOR	5745	30-50	31-30	32-27	32-40	32-45	32-56	33-56	36-50	38- 6	38-10	64-33#			
PRMTST	4773	54-36	54-43	54-49#											
PRREFC	5735	3-29	53-55	54-24	54-29	54-31	54-35	54-38	54-42	64-25#					
PWRERR	5704	49-26	49-28	63-55#											
PWUP2	265	1-22	5- 3#												
QCYC52	5732	12-25	20-21	64-22#											
QMIFRM	4317	32- 4	48-20#	48-52	60-32										
QMTIM	4345	48-47#	48-50	54- 9	55-15	55-37									
QSTPSZ	5734	12-20	54-16	64-24#											
QTRMIN	4351	45-39	48-52#												
RAMP	3351	30-36	34-12	37-30#											
RASHIX	5720	55- 7	56-22	56-24	56-33	64-11#									
RASHIY	5721	55-29	64-12#												
RASLMC	5054	13-17	13-41	55-52#	56-40										
RASLOX	5722	55-20	56-13	56-18	56-20	56-30	56-31	56-39	64-13#						
RASLOY	5723	55-42	64-14#												
RAXCEN	7716	13-25	22-30	56-36	56-37	66- 8#									
RAXPOS	1065	13-24#													
RAXUND	5003	17-38	18- 2	55- 4#	55-18										
RAYCEN	7717	13- 1	66- 9#												
RAYUND	5025	17- 7	17-29	55-26#	55-40										
RCOUNT	3336	29-14	29-43	37-17#											
RDIRCT	4314	48-16#	48-33	48-38	48-39										
RDSIZ	2210	25-54#	26- 2	62- 7	62-26										
RECHEK	3104	34-19#	35-22												
RECOVR	3042	33-39#	58-29												
RED	3476	38-34	39-17#												
REULL	2625	30-56#	31-31												
REPEAT	7770	10-30	16-49	16-51	66-45#										
REFMSG	1324	14-22	14-37	16-48#											
RETN3	4101	45-28#	45-44	46- 3	46- 7	46- 9	46-10	47-44	49-29	56-45	61-19	65-34			
RGTG	3323	29- 7	29-30	29-48	30- 7	31-35	37- 5#								
RFWD	7772	21-19	26-12	26-13	26-14	26-21	26-22	26-23	28-19	57-38	61-39	66-47#			
RUNWLD	2405	28-18#													
RWEUF	1731	7-51	8- 8	22-30#											
RWENAB	2232	3-16	3-21	26-19#	26-25	28-25	58-20								
RXCFIN	1366	17-36#	17-56												
RXCTRL	1400	10-21	17-47#	17-52											
RXCYCL	5727	12-49	17-37	17-49	18- 1	64-19#									
RYCFIN	1337	17- 5#	17-26												
RYCTRL	1351	10-20	17-17#	17-22											
RYCYCL	5731	12-41	17- 6	17-19	17-28	64-21#									
SACDLR	433	7-25	7-37#												
SAMPLT	2702	31-47#	32- 5	32- 9											
SAVAC	5572	45-27	45-33	62-42#											
SAVID	5573	48-18	48-21	58-36	58-43	62-43#									
SAVSA	5265	1-44	2-56	47-46	47-48	47-56	57-44	58-33	58-35	58-44#					
SCI	7700	1-37	1-45	1-47	2- 2	2-10	2-26	2-28	61-37	65-31	65-47	65-57#			
SCOUNT	5766	9- 7	9-42	9-45	9-53	10- 5	10- 6	29-15	29-28	29-40	50-16	50-22	50-26	50-47	51- 4
		51-13	54-17	54-49	64-52#										
SET1UP	520	8-45#	44-41												
SETFR	2671	31-36#	31-45												

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TWO	1714	3-20	9-20	9-56	14-10	22-16#													
TWO2	5612	50-46	53- 9	53-36	59-36	60-14	62- 3	63-12#											
TWOD12	5620	53- 6	53-33	63-18#															
TWOD15	5621	31-27	54-34	55-14	59- 8	63-17#	65-48												
TWOD16	5622	41-21	42-10	53-42	63-20#														
UZMSG	5275	24-39	24-42	24-49	57-23	58-55#	59-14												
UZMSG	5323	25-17	25-20	25-27	57-29	59-23#	59-40												
UERROR	5425	4-47	4-49	9-23	25-43	35-53	38- 3	38-14	38-22	45-12	54- 3	56- 5	60-46#	60-50	61- 2				
		61- 7																	
UND420	3475	28-26	33-39	39-15#															
UNMSK	1711	5-17	6-34	7-31	22-13#														
USMR	4051	5-36	6-40	21-37	24-33	25-56	35-31	45- 3#	45-14	53-50	54-19	58- 9							
USMSG	5353	1-32	4-13	14-17	16- 8	16-10	16-34	19-37	20-13	22-50	33-25	33-32	54-28	54-46	56-56				
		59-49#	60- 9	60-11	60-22	60-48	61- 6												
UWTB	4345	4-57	5- 6	5- 9	49-16#	49-19													
UWTONE	4371	5-43	33-35	34- 6	36-10	36-13	36-33	49-21#	49-24	61-50									
UWTX	4360	3-19	3-24	3-28	6-35	13-22	13-45	21-34	28-29	30-20	33-40	36-20	49-10#	49-14	49-18				
		49-23	53-47	54-11	54-53														
VECTOR	6000	3-54	65-10#																
VELC	4502	50-31	50-36	50-40	50-56#														
VELOC	5575	51-15	62-54#																
VFAULT	4375	45-53	49-28#																
WAIT	307	3-41	4-48	5-22#	5-47	65-53													
WANTHI	5763	26- 5	27-31	27-46	28- 6	64-48#													
WANTLO	5764	26- 6	27-32	27-47	28- 7	64-50#													
WLCYC	5742	12- 4	18-11	18-31	18-40	64-30#													
WLCYSZ	5741	12- 3	18-10	64-29#															
WLSCAN	7763	11- 6	16-45	18-14	18-15	66-40#													
WMSG4	3365	33-24	37-45#																
WMSG5	3366	33-31	37-46#																
XCMD	5252	30-47	30-49	33-37	33-38	33-51	58-28#												
XCYC5Z	5726	12-50	13-29	17-36	64-18#														
XFLARE	124	1-54	3- 1#																
XINDX	5707	64- 1#																	
YCYC5Z	5730	12-42	13- 5	17- 5	64-20#														
YESSTP	2724	29-20	29-34	29-52	30-11	32-11#	33-19												
YINDX	5710	64- 2#																	
ZERO	1712	8- 1	8-18	13-40	15-51	23-14#													
ZERO1	2241	23-36	26-28#	29- 5															
ZERO2	4153	46-20#	50-42	58- 1															

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JR MESSAGES

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<u>TACHOMETER SERVO CORRECTION</u>	114000 + X (5 bits)
<u>END OF WLD SCAN n</u>	120000 + n
<u>END OF POLR. ROTATION CYCLE n</u>	110000 + n
<u>END OF RASTER X CYCLE n</u>	112000 + n
<u>END OF RASTER Y CYCLE n</u>	116000 + n
<u>END OF EXP. REPEAT CYCLE n</u>	160000 + n
<u>END OF EXPERIMENT</u>	155003
<u>START OF EXPERIMENT</u>	155001
(FOLLOWED BY 16-WORD EXP. DEFINITION BLOCK)	_____
	:
<u>COMMAND MODE INSTRUCTION DONE</u>	150000 + PC (REL. TO 6060 _r)
(FOLLOWED BY THE INSTRUCTION)	_____
<u>CALIBRATION IN</u>	155006
<u>" OUT</u>	155007
<u>ERROR DETECTED</u>	155555
(ERROR REGISTRATION)	4-bit COUNT, 12-bit address
<u>OVERFLOW (ON DATA BUS)</u>	155550
<u>WLD STUCK ON NULL</u>	133334
<u>WLD STUCK OFF NULL</u>	133335
<u>POLR. STAYS NULL (RTI.)</u>	133331
<u>" " NO-NULL "</u>	133332
<u>MEMORY DUMP</u>	17 nnnn - address nnnn,
(followed by)	xxxxxx - contents of address

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DMA BLOCK CONTENTS

USER SYMBOLS	SYMBOL	ADDR	TM MINOR	CM OFFSET	COMMENTS
=====	=====	=====	=====	=====	=====
	SCI	7700	2/3		WORD LOADED BY ORC
	LIR	7701	4/5		LIST INSTR. REG.
	LPC	7702	6/7		LIST PROGRAM CNTR.
	OLS	7703	8/9		OBSERVING LIST SEG.
	OL	7704	10/11		OBSERVING LIST WD. 1
	OL+1	7705	12/13		DETS/FORMAT
	OL+2	7706	14/15		WL INCRS
	OL+3	7707	16/17		WL POSN/OFFSET
	OL+4	7710	18/19		EXPMT REFS/WL MSB
	OL+5	7711	20/21		PSQ/WL INC/P INC/T INTVL
	OL+6	7712	22/23		RAS X Q/RAS Y Q
	OL+7	7713	24/25		X INCR/Y INCR/CAL OFS
	OL+8	7714	26/27		GATE TIM CNT/SERVO ADJ/CLK
	EXPNUM	7715	28/29		EXPERIMENT NUMBER
X POSN	RAXCEN	7716	30/31	0	X RASTER CENTER
Y POSN	RAYCEN	7717	32/33	1	Y RASTER CENTER
I THRSN	THRSHI	7720	34/35	2	SERVO COUNT THRSHLD
MAXHI	MAXHI	7721	36/37	3	GLOBAL WLD OFFSET HI BITS
MAXLO	MAXLO	7722	38/39	4	GLOBAL WLD OFFSET LO BITS
RI	'R1'	7723	40/41	5	USER REGISTERS
	'R2'	7724	42/43	6	
	'R3'	7725	44/45	7	
	'R4'	7726	46/47	10	
	'R5'	7727	48/49	11	
	'R6'	7730	50/51	12	
	'R7'	7731	52/53	13	
	'R8'	7732	54/55	14	
	FLAG	7733	56/57	15	FLARE CONTROLLED FLAG WORD
FLAREX	FLARX	7734	58/59	16	X FLARE COORDINATE (HXIS)
FLAREY	FLARY	7735	60/61	17	Y FLARE COORDINATE (HYIS)
IMIN	MINIC	7736	62/63	20	MINIMUM INTEN COUNT / 2
IMINX	MINIX	7737	64/65	21	MINIMUM INTEN X FOSN
IMINY	MINIY	7740	66/67	22	MINIMUM INTEN Y FOSN
WMAXHI	MAXIWH	7741	68/69	23	MAXIMUM WL HIGH BITS
WMAXLO	MAXIWL	7742	70/71	24	MAXIMUM WL LOW BITS
IMAX	MAXIC	7743	72/73	25	MAXIMUM INTEN COUNT / 2
IMAXX	MAXIX	7744	74/75	26	MAXIMUM INTEN X FOSN
IMAXY	MAXIY	7745	76/77	27	MAXIMUM INTEN Y FOSN
BMAX	MAXBV	7746	78/79	30	MAXIMUM BLUE SHIFT
BMAXI	MAXBI	7747	80/81	31	MAXIMUM BLUE COUNT / 2
BMAXX	MAXBX	7750	82/83	32	MAXIMUM BLUE X FOSN
BMAXY	MAXBY	7751	84/85	33	MAXIMUM BLUE Y FOSN
RMAX	MAXRV	7752	86/87	34	MAXIMUM RED SHIFT
RMAXI	MAXRI	7753	88/89	35	MAXIMUM RED COUNT / 2
RMAXX	MAXRX	7754	90/91	36	MAXIMUM RED X FOSN
RMAXY	MAXRY	7755	92/93	37	MAXIMUM RED Y FOSN
	OFFTTL	7756	94/95		SERVO OFFSET TOTAL
	FIX	7757	96/97		WLD FIX SIZE FROM GND CMD
	LASTFX	7760	98/99		LAST WLD FIX VALUE USED
	AWLDHI	7761	100/101		WLD ACTUAL POSITION HIGH BITS
	AWLDLO	7762	102/103		WLD ACTUAL POSITION LOW WORD
	WLSCAN	7763	104/105		WAVELENGTH SCAN COUNTER
	EKRWRD	7764	106/107		ERROR MESSAGE SENT LAST
	CPRCYC	7765	108/109		POLARIMETER PASS COUNTER
	CXCYCL	7766	110/111		X RASTER PASS COUNTER
	CYCYCL	7767	112/113		Y RASTER PASS COUNTER
	REPEAT	7770	114/115		EXPERIMENT PASS COUNTER
	ACONFG	7771	116/117		DETECTOR CONFIGURATION
	RFWD	7772	118/119		RASTER/WLD POWER CONDITION
	MICLN	7773	120/121		MINOR FRAME COUNTER
	TEMP	7774	122/123		MONITOR SAVE LOCATION
	TEMPI	7775	124/125		INDEX SAVE LOCATION
	TEMPJ	7776	126/127		TEMPORARY STORAGE LOCATION
	BSYGN	7777	0/1		SERVO GAIN FACTOR/RUSY (SIGN) BIT

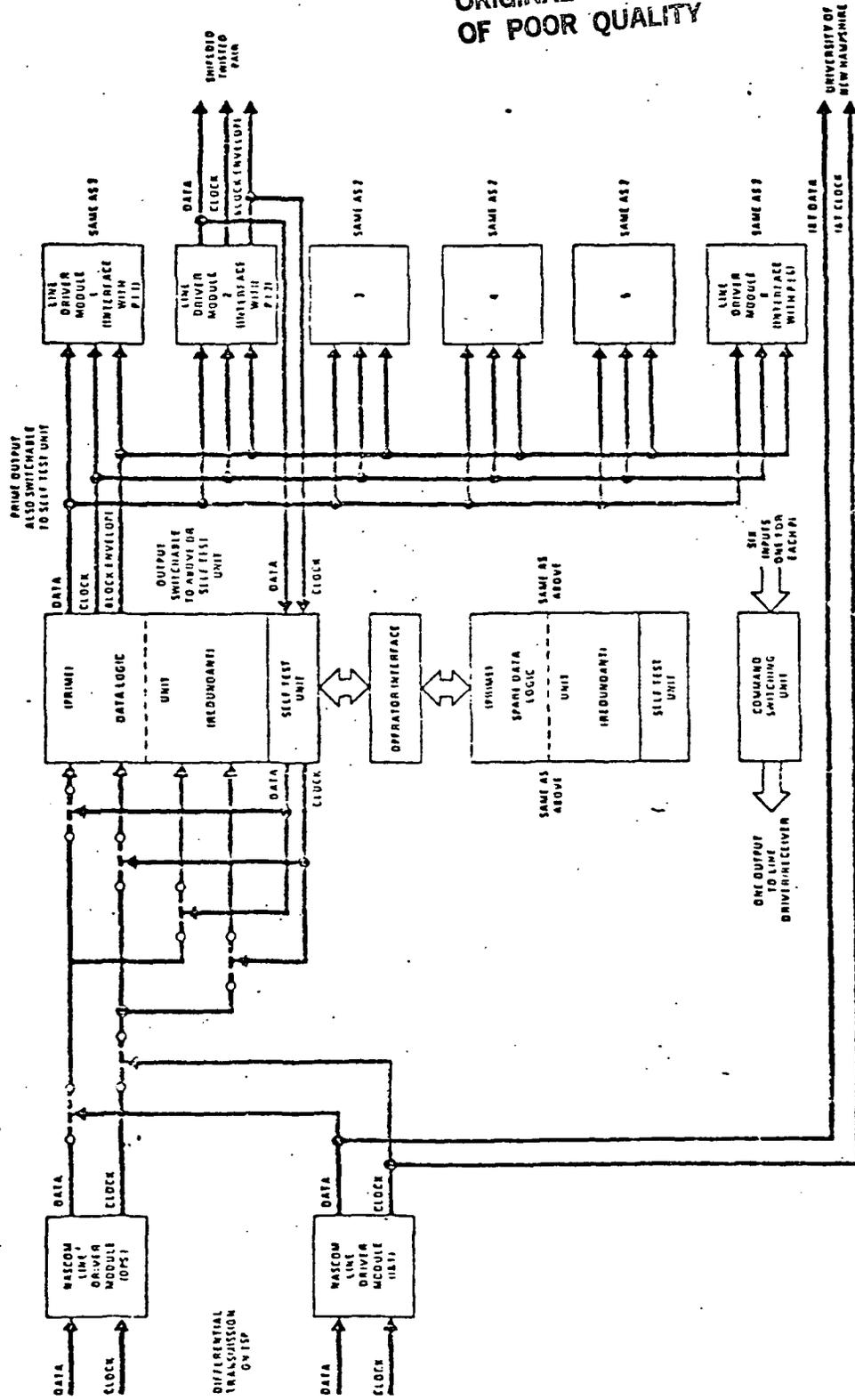
APPENDIX 4
EXPERIMENT OPERATIONS FACILITY
INTERFACE UNIT (EOFIU)

SOLAR MAXIMUM MISSION (SMM)
PRINCIPAL INVESTIGATOR
EXPERIMENT OPERATIONS FACILITY INTERFACE UNIT (EOFIU)

The following is a summary of the format in which data will be provided to the SMM Principal Investigators (experiments) in the Experiment Operations Facility (EOF) at the output of the EOF Interface Unit (EOFIU):

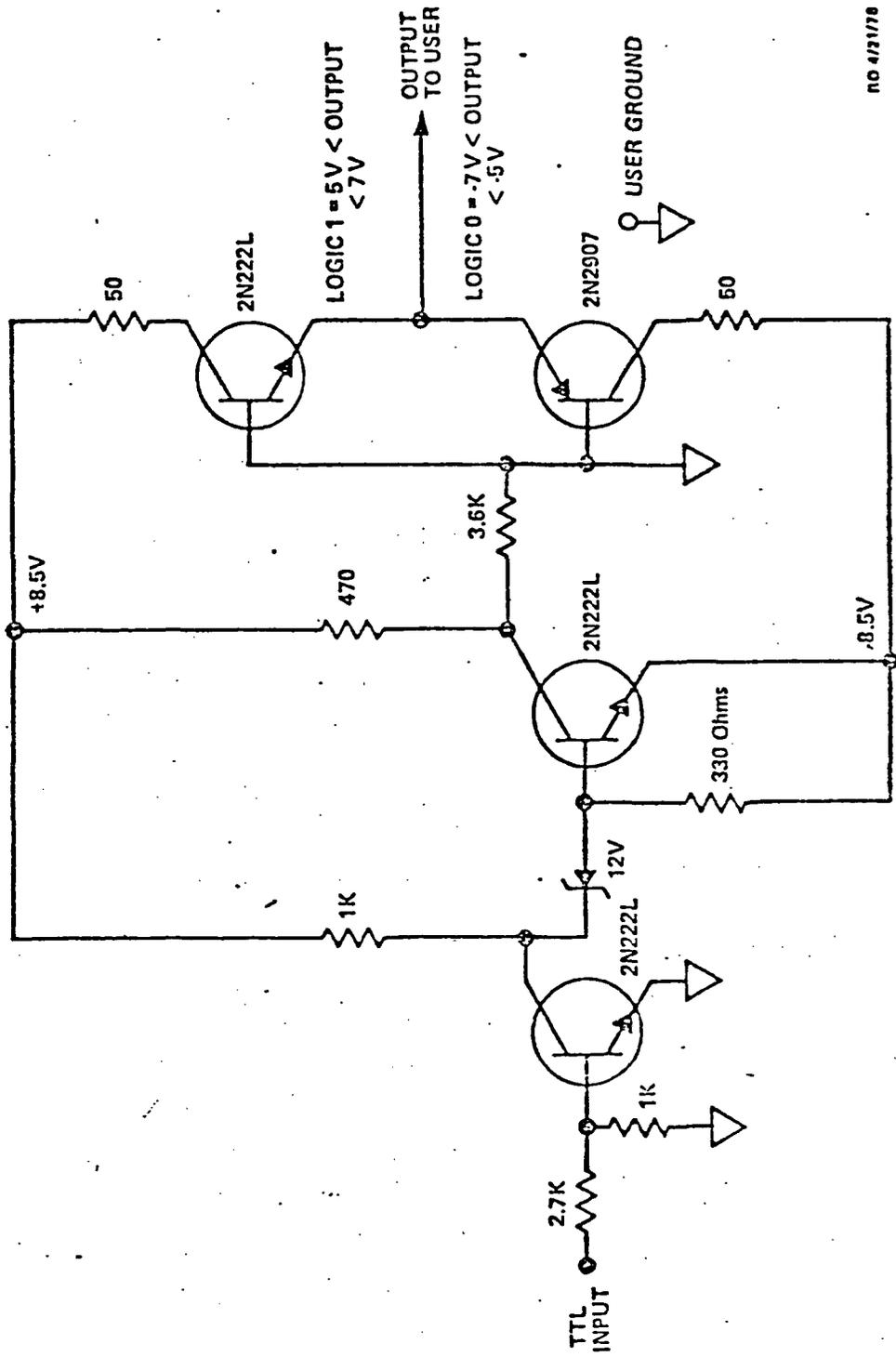
- a. Three lines will be provided to the experimenters (Figure 1). The signals will be output to the user from line drivers. The schematic for these line drivers is shown in Figure 2.
 - (1) clock (continuous, with transitions in the middle of each data bit) (Figure 3)
 - (2) data (bursted)
 - (3) block envelope
- b. Minor frame synchronized SMM data will be bursted to the experimenters four contiguous minor frames at a time at 224 kbps (18.3 ms). The interval between blocks of bursted data will vary from a minimum of 3 ms to a nominal maximum of 238 ms.
- c. The average data rate from the output of the EOFIU to the experimenters will vary from 16 to 191 kbps.
- d. Data within the four minor frame blocks from SMM Integration and Test (I&T) will be contiguous and the same type from block to block. The bursted data within each block will be the same number of words and in the same format as during operations. The average data rate from the output of the EOFIU to the experimenters will be 16 kbps.
- e. During operations, the data within the four minor frame blocks will be contiguous and the same type. Each block, however, could be different and be coming in from a different Space Tracking and Data Network (STDN) site. Two types of data can be received at a time from block to block from any one STDN site, and data from up to three sites can be received from block to block [i.e., one block would be real-time (forward) telemetry data

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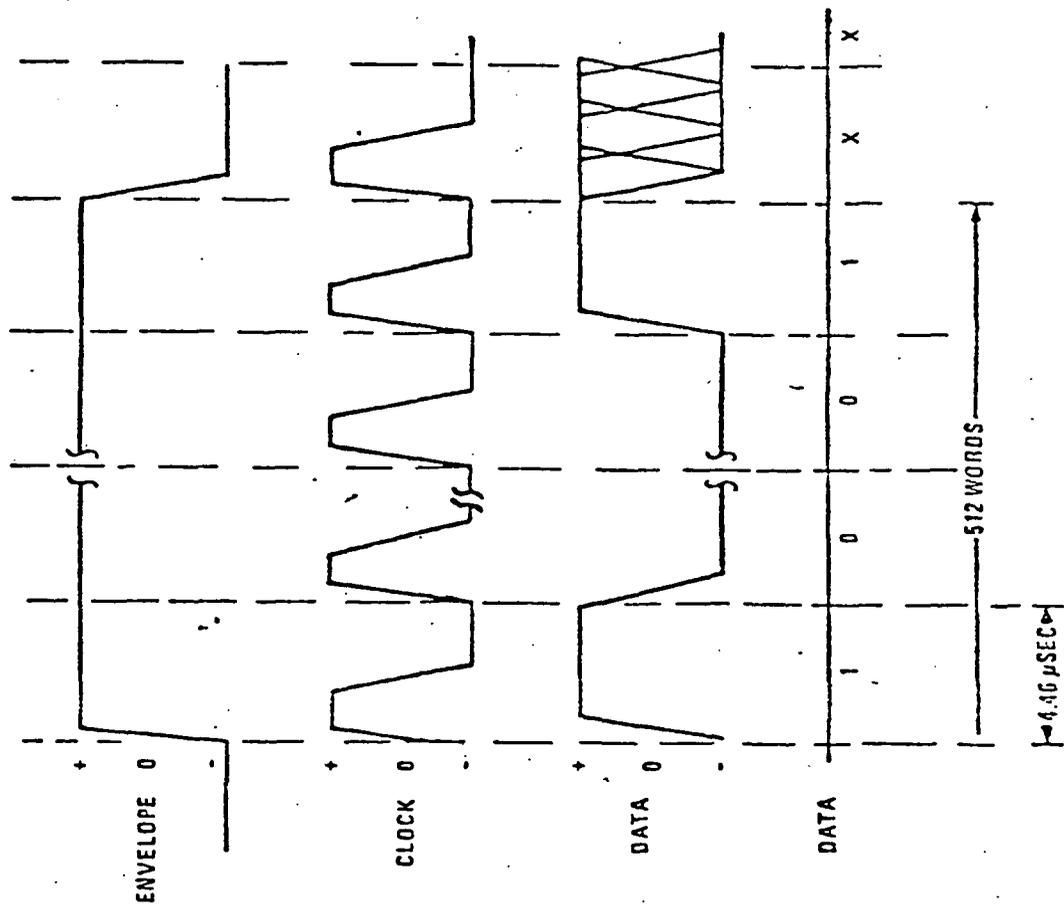
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Figure 1. EOF Interface Unit Elements



NO 4721778

Figure 2. EOFIU Driver Interface Data Out, Envelope Enable Out, and Clock Out



NOTE: VOLTAGE: 5 < V < 7
 REF.: NASCOM 844-71-03

RO 4/21/78

Figure 3. EOFIU Output Signals to Users

from the first site, the next block could be on board computer (OBC) data dump from the first site, the next block could be playback (reverse) telemetry data from a second site, and the fourth block could be playback High Altitude Observatory (HAO) data from a third site]. Any combination of data types (maximum of four) could be received.

- f. In order that the experimenters may distinguish between data types, forward/reverse, and end of data, four words will be used.
- (1) SMM telemetry word 3 bits 3 and 4 each will be 0 during HAO data. The other bits can be 1 or 0.
 - (2) SMM telemetry word 3 bits 3 and 4 will be 01 for engineering telemetry format. The other bits can be 0 or 1.
 - (3) SMM telemetry word 3 bits 3 and 4 will be 10 for science telemetry format. The other bits can be 0 or 1.
 - (4) SMM telemetry word 3 bits 0, 1, 2, 3, and 4 will be 10011 during flexible format telemetry. The other bits can be 1 or 0.
 - (5) SMM byte 3 (8 bit byte) will be 00011000 during OBC dump. The EOFIU will add the ones to the OBC third byte. Bytes 0, 1, and 2 will be SMM sync.
 - (6) SMM telemetry word 3 will be 11111111 for the EOFIU test pattern.
 - (7) SMM telemetry word 67 bit 0 will be a 1 during dwell mode. Bits 1-7 will be dwell identification.
 - (8) SMM telemetry words 8 and 9 will be modified except during HAO data, dwell mode, and OBC dumps.
- g. Word 8 will be source identification (STDN site ID), word 9 will contain flags, words 0, 1, and 2 will be SMM telemetry minor frame sync and words 3, 4, 5, 6, 7, 10-127 will be SMM data.
- h. Word 8 will allow the experimenters to keep track of each data source (possible two data types per source) and by also using word 9, the Principal Investigator (PI) will be able to determine the type of data from that source.

- i. Word 9 will be set up by the EOFIU such that bits 0, 1, 2, 3, 4, 5, and 6 will be flags. Bits 6 and 7 have not been designated at this time. For each block that is transmitted to the experimenters, word 9 bits 0, 1, 2, 3, 4, and 5 will be set by the EOFIU as follows:

<u>Designation</u>	<u>Bit</u>	<u>0</u>	<u>1</u>	<u>2</u>	<u>3</u>	<u>4</u>	<u>5</u>
Not end of data	0	-	-	-	-	-	-
End of data	X	-	-	-	-	-	-
Forward data	-	0	-	-	-	-	-
Reverse data	-	1	-	-	-	-	-
EOFIU self-test mode	-	-	0	-	-	-	-
S/C I&T and operations mode	-	-	1	-	-	-	-
Polynomial check good	-	-	-	0	-	-	-
Polynomial check bad	-	-	-	1	-	-	-
Full NASCOM block	-	-	-	-	0	-	-
Partial NASCOM block	-	-	-	-	1	-	-
Real-time telemetry	-	-	-	-	-	0	-
Playback telemetry	-	-	-	-	-	-	1

- j. Bit 0 in word 9 will be set up to zero in all blocks bursted to the experimenters except for the last block. In the last block of that transmission for that data type from that STDN site, bit 0 or word 9 will be set to a one.
- k. The word order and number of words in each block bursted to each experimenter will be as follows:
- (1) Words 0, 1, 2, 3—127
 - (2) Words 0, 1, 2, 3—127
 - (3) Words 0, 1, 2, 3—127
 - (4) Words 0, 1, 2, 3—127

There will be a total of 512 words in each block bursted to each experimenter. Each word will have 8 bits and be in bit order 0 (MSB), 1, 2—7.

- l. Data can be received by the EOFIU in forward or reverse order (spacecraft realtime or tape recorder playback). In both cases, blocks bursted to the experimenters will be sent word order and bit order in the forward direction as per above. Minor frame order, however, will be different. During forward data, minor frame order will be 0, 1, 2, 3, 4—127. During reverse data minor frame order will be 127, 126, 125, 124, 123—0.
- m. The EOFIU has a self-test mode which will generate two test patterns. These will be two fixed-dummy NASCOM blocks. One will simulate forward data, and the other will simulate reverse data. These blocks will be bursted to the experimenter just like real data at 224 kbps. The details of these patterns are provided in Tables 1 and 2. The experimenter will use these patterns to check out and verify the EOFIU/experimenter interface during equipment installation and checkout and during trouble analysis. During operations, the experimenter may want to reject these patterns or use them for an automatic test whenever these patterns are on the line for trouble analysis.
- n. The NASCOM data blocks from STDN have polynomial error control checkbits within each block. The EOFIU will perform a poly-check on each NASCOM block, compare this check with the STDN error control checkbits and provide the experimenter with the results of this comparison. If this polynomial check is bad, any data within that block can be bad and the experimenter may reject it.
- o. If there is any data dropout at the STDN site, that site will send partial blocks to the users. This means that a block can include one, two, three and a partial fourth minor SMM frame or a partial of any SMM minor frame. During a data dropout, any combination of partial SMM minor frames can be received. The rest of the data within that block can be random bits, old data, or someone elses data. The experimenter may want to reject these data.
- p. Some experimenters may want to process the SMM real-time telemetry in near realtime and the playback telemetry at a later date. Word 9 bit 5 will allow the experimenter to automatically distinguish between these data.

Table 1
EOFIU Test Pattern Normal Forward Output

Subframe 1		Subframe 2		Subframe 3		Subframe 4	
Word	Octal Data	Word	Octal Data	Word	Octal Data	Word	Octal Data
0	372						
1	363						
2	040						
3	377						
4	200						
5	100						
6	300						
7	040						
8	246						
9	030						
10	340						
11	020						
12	220						
13	120						
14	320						
15	060						
16	260						
17	160						
18	360	Same as column 1		Same as column 1		Same as column 1	
19	000						
20	010						
21	004						
22	014						
23	002						
24	012						
25	006						
26	016						
27	001						
28	011						
29	005						
30	015						
31	003						
32	013						
33	007						
34	017						
35-64	252						
65	200	65	100	65	300	65	040
66	377						
67	177						
68	200						
69-127	252						

Note: words 0 1 2 are SFR Subframe sync

372 363 040
1111010 1111011 0010000

Bit 0 is the first bit transmitted to the experimenter

Table 2
 EOFIU Test Pattern Normal Reverse Output

Subframe 1		Subframe 2		Subframe 3		Subframe 4	
Word	Octal Data	Word	Octal Data	Word	Octal Data	Word	Octal Data
0	372						
1	363						
2	040						
3	377						
4	200						
5	100						
6	300						
7	040						
8	246						
9	304						
10	340						
11	020						
12	220						
13	120						
14	320						
15	060						
16	260						
17	160						
18	360	Same as column 1		Same as column 1		Same as column 1	
19	000						
20	010						
21	004						
22	014						
23	002						
24	012						
25	006						
26	016						
27	001						
28	011						
29	005						
30	015						
31	003						
32	013						
33	007						
34	017						
35-64	252						
65	040	65	300	65	100	65	200
66	377						
67	177						
68	200						
69-127	252						

- g. The clock envelope will be activated at the beginning of the transmission of the first bit in that block (bit zero of word 0) and will end with the end of the last data bit in the last word of the fourth minor frame (bit 7 of word 127). This envelope will be the same signal level as a clock bit.

APPENDIX 5

ACQ

C-3

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Running ACQ

Starting up

It is first necessary to install the various tasks that interact with each other in the data acquisition process if this has not already been done. Type the following:

```
LOG 200,204
ASN DB0:=SY:
@INSACQ
```

The command file INSACQ installs the tasks and insures that the disk files MAJORS.RAW and ODDBALL.RAW are unlocked. Now type

```
RUN ACQ
```

ACQ should respond with the following question:

```
ENTER S FOR SCIENCE ONLY, F FOR FLEX ALSO, E FOR ENG.,FLEX, & SCI.
```

The answer determines the type of data that will be accepted for processing. An S response allows only science mode data, an F allows both science and flexible format, and an E allows science, flexible, and engineering. After you respond, the next question is:

```
DO YOU WANT TO BYPASS SOURCE CHECKING [Y OR N]
```

You would normally answer N. A yes response may be necessary to record some types of I & T data since the interface unit does not insert a source code in this mode and the source byte may therefore not be constant. A yes response should never be made if more than one source is expected. The next question:

```
DO YOU WANT STASH (POSITIONAL MODE)?
```

should be answered with a Y to choose the positional recorder. If you answer N, the following question appears:

```
DO YOU WANT THASH (SEQUENTIAL MODE) ?
```

Enter Y for the sequential recorder. Enter N only if you want no recording at all.

The next question to be answered depends on your choice of mode made above. If STASH was chosen you should see:

```
ENTER FIRST LEGAL FRAME # (21 BITS MAX)
?>
```

Your answer determines the smallest major frame # that STASH will consider for recording to disk. The value must not exceed 2097151, which is the maximum possible major frame number. The next question for STASH is:

```
LAST
?>
```

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The answer c determines the last major frame to be considered. In SEARCH the actual record used for a legal (i.e., within the above limits) major frame $\#$ is the major frame modulus 11000 plus 1 or, $REC \# = Mod(mf\#, 11000) + 1$. Hence if limits are set from 100 to 20000, frames 200 and 11200 would be candidates for the same record.

If THASH is used the following question appears:

ENTER OFFSET FOR DISK
?>

Your response determines the first record to use for the first received major frame. Subsequent major frames are recorded in order of reception in the following records.

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Run Time Commands

While ACQ is running, it can accept several commands to allow operator interaction. Each of these options is initiated by a single input character (without a carriage return). Depending on the option, there may then appear some questions to answer. The options are listed below under the initiator character.

L - change limits. This works only when STASH is running. When accepted, the questions concerning the first and last legal major frame # appear. Your answers change them.

T - change time between status reports. The following query appears:

ENTER STATUS INTERVAL IN MINUTES
?)

Your answer must be in integer minutes (fractions are not allowed!).

? - give an immediate status report. The program will also check for expired sources.

A - change acceptance mask. The science, flex, engineering question will reappear, allowing you to enter a new answer.

C - clear a section of the major map. Normally when data is recorded the records are write protected via a map in core. Hence, a re-transmission or a wrap-around would not be recorded. The map is always cleared when ACQ is restarted. The C command allows run time clearing (or unprotecting) of a portion of this map. You must answer the following:

ENTER FIRST MAJOR TO CLEAR
?)
LAST
?)

Modulus 11000 of your answer is used.

P - protect a section of the major map. This is the complement to the C command. Similar questions are asked. This could be used to protect previously recorded data.

ESCAPE KEY - kill current messages. The message buffer is forced empty, stopping any accumulated messages. Subsequent messages will be printed out. This command is also used to recover from the kill all message command below.

CONTROL K - kill all current and future messages. Type a CONTROL K when the clattering of the terminal is driving you crazy. It is also useful if you ran out of paper, or for overnight. Status reports are still printed. To reactivate messages, type ESCAPE KEY.

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Stopping ACQ

ACQ will perform an orderly exit, closing files and aborting tasks, when you type control z. If this doesn't work try typing @DB0:[200,204]ABO on any terminal. If the system can't get this command file started, try to abort the installed tasks individually starting with SNAT as follows:

ABO SNAT
ABO STASH
ABO THASH
ABO ACQ

Then unlock the files:

PIP DB0:[200,204]*.RAW/UN

If all fails, re-boot the computer and start over with @INSACQ when you need ACQ again.

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3.3 PROJECT DATA FORMATS

The project data formats (PDF's) for SMM-A are as follows:

- PDF-A is designed to contain real-time 16 kbps data in a forward direction.
- PDF-B handles the 32 kbps onboard computer data dump, and is sent simultaneously with format A in a forward direction to GSFC.
- PDF-C handles High Altitude Observatory (HAO) real-time data at 256 kbps, as a backup mode of operation in the event that the HAO recorder becomes inoperative. These data will be input to the Digital Data Processing System (DDPS) at 128 kbps and transmitted to GSFC in the forward direction.
- PDF-D contains spacecraft recorder dump data at 512 kbps (these data will be analog recorded and input to the DDPS at 128 kbps, and transmitted to GSFC in reverse order).
- PDF-E contains HAO recorder dump data at 512 kbps (these data will be analog recorded and input to the DDPS at 128 kbps, and transmitted in reverse order).
- PDF-F contains spacecraft recorder dump data, with the same characteristics as PDF-D except the transfer to DDPS is at 256 kbps.
- PDF-G's HAO recorder dump data have the same characteristics as PDF-E except the transfer to DDPS is at 256 kbps.
- PDF-H will contain spacecraft recorder dump data at 512 kbps (data will be analog recorded and played back in reverse order at a 12:1 reduced speed from the recorder). The playback data rate of the analog tape will be 42.666 kbps.
- PDF-I contains HAO recorder dump data at 512 kbps (data will be analog recorded and played back in reverse order at 12:1 reduced speed from the recorder; the playback data rate of the analog type will be 42.666 kbps).
- PDF-J contains spacecraft recorder dump data at 512 kbps (data will be analog recorded and played back in reverse order at 6:1 reduced speed from the recorder). The playback data rate of the analog tape will be 85.333 kbps.
- PDF-K contains HAO recorder dump data at 512 kbps (data will be analog recorded and played back in reverse order at 6:1 reduced speed from the recorder). The playback data rate of the analog tape will be 85.333 kbps.
- PDF-L will be real-time data at 1 kbps. This is an emergency format and will be used to sync the OBC dump in the event it gets out of main frame sync with the real-time 16 kbps data (data will be in a forward direction and should be transmitted off station of 1 block per second).

PDF's C, D, and E can be direct from an analog tape or from digital tapes.

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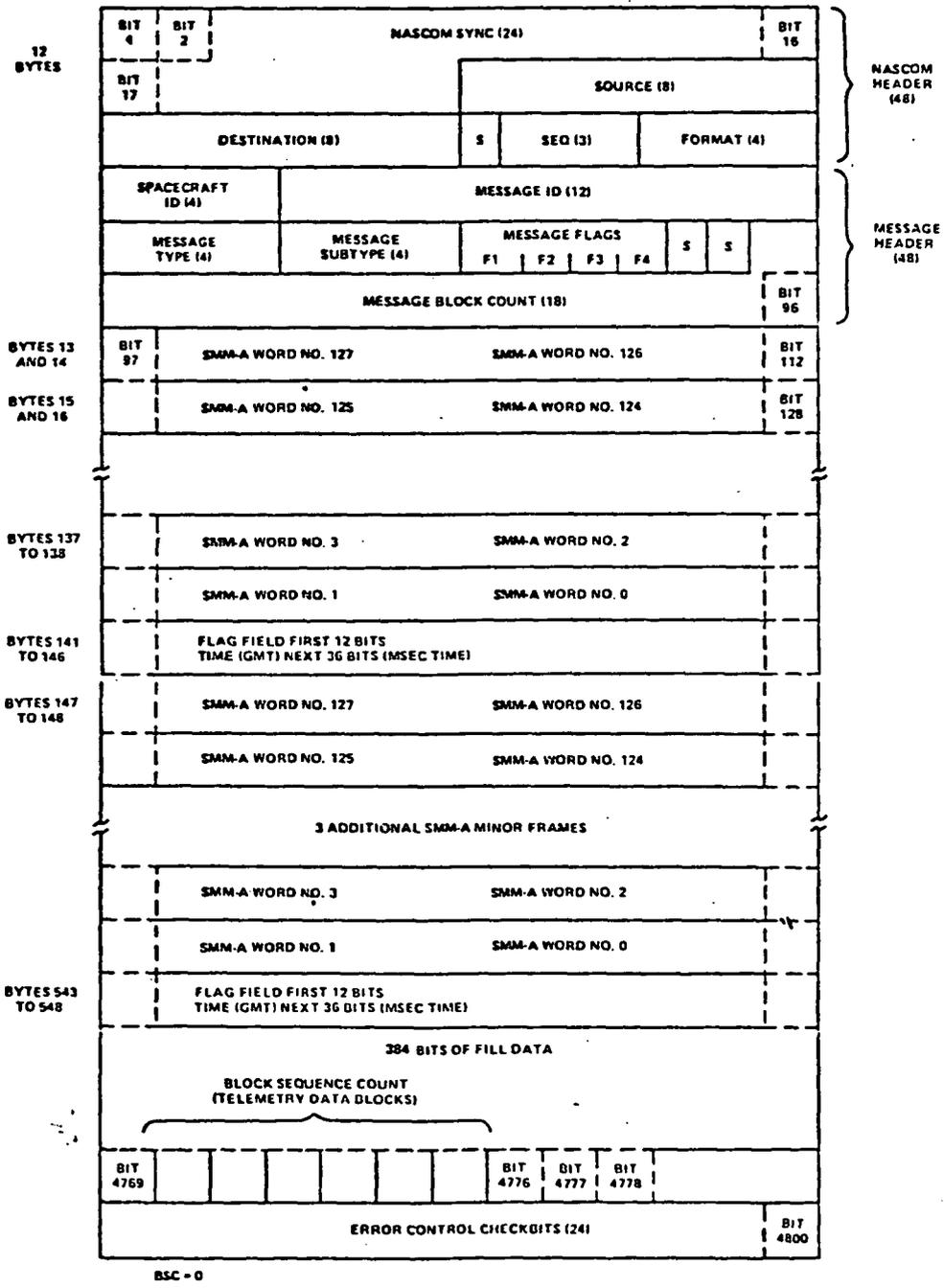


Figure 5-5. SMM-A Project Data Formats D, F, H and J.
Playback Data 512 kbps from Observatory

Table 3-1
SMM-A Minor Frame Telemetry Format Mode 2
(Science Format)

MINOR FRAME WORD NR.	DESCRIPTION	ID
00	FRAME SYNC WORD	CDH 01
01	FRAME SYNC WORD	CDH 01
02	FRAME SYNC WORD	CDH 01
03	TLM FORMAT, RATE & ID	CDH 56
04	SCIENCE DATA	HNIS 04
05	SCIENCE DATA	HNIS 04
06	SCIENCE DATA	HNIBS 22
07	SCIENCE DATA	HNIBS 23
08	UNASSIGNED	SMM
09	UNASSIGNED	SMM
10	SCIENCE DATA	XRP 01
11	SCIENCE DATA	XRP 01
12	DATA SOURCE 1	GRE 26
13	DATA SOURCE 1	GRE 26
14	SCIENCE DATA	HNIS 04
15	SCIENCE DATA	HNIS 04
16	FPSS 1 WORD 1	SACS 01
17	FPSS 1 WORD 2	SACS 02
18	FPSS 1 WORD 3	SACS 03
19	FPSS 1 WORD 4	SACS 04
20	SCIENCE DATA	HNIS 04
21	SCIENCE DATA	HNIS 04
22	PHOTODIODE 1	HNIS 24
23	ANE STATUS 3	HNIS 07
24	SCIENCE DATA	HNIS 04
25	SCIENCE DATA	HNIS 04
26	SCIENCE DATA	XRP 01
27	DMA ADDRESS	UVSP 05
28	DMA DATA	UVSP 06
29	STATUS MONITOR	UVSP 07
30	DATA BUS 2	UVSP 01
31	DATA BUS 2	UVSP 01
32	SUBCOMMUTATOR NR 1	CDH 06
33	SUBCOMMUTATOR NR 2	CDH 07
34	RECEIVER STATUS	CDH 12
35	OBC DATA WORD ID	OBC 01
36	SCIENCE DATA	HNIS 04
37	SCIENCE DATA	HNIS 04
38	SCIENCE DATA	HNIBS 24
39	SCIENCE DATA	HNIBS 25
40	SCIENCE DATA	HNIS 04

Table 3-1 (continued)

MINOR FRAME WORD NR.	DESCRIPTION	ID
41	SCIENCE DATA	HNIS 04
42	SCIENCE DATA	XRP 01
43	SCIENCE DATA	XRP 01
44	DATA SOURCE 1	GRE 26
45	DATA SOURCE 1	GRE 26
46	SCIENCE DATA	HNIS 04
47	SCIENCE DATA	HNIS 04
48	FPSS 2 WORD 1	SACS 05
49	FPSS 2 WORD 2	SACS 06
50	FPSS 2 WORD 3	SACS 07
51	FPSS 2 WORD 4	SACS 08
52	SCIENCE DATA	HNIS 04
53	SCIENCE DATA	HNIS 04
54	SCIENCE DATA	HNIBS 26
55	SCIENCE DATA	HNIBS 27
56	SCIENCE DATA	HNIS 04
57	SCIENCE DATA	HNIS 04
58	SCIENCE DATA	XRP 01
59	SCIENCE DATA	XRP 01
60		
61		
62	DATA BUS 3	UVSP 02
63	DATA BUS 3	UVSP 02
64	S/C CLOCK BITS 7-0	CDH 02
65	FRAME COUNTER	CDH 03
66	CMD COUNTER (SELECTED CU)	CDH 04
67	DWELL MODE & CHANNEL ID	CDH 05
68	SCIENCE DATA	HNIS 04
69	SCIENCE DATA	HNIS 04
70	SCIENCE DATA	HNIBS 28
71	SCIENCE DATA	HNIBS 29
72	SCIENCE DATA	HNIS 04
73	SCIENCE DATA	HNIS 04
74	SCIENCE DATA	XRP 01
75	SCIENCE DATA	XRP 01
76	DATA SOURCE 2	GRE 27
77	DATA SOURCE 3	GRE 28
78	SCIENCE/HSKPG SUBCOM	WLCP 10
79	SCIENCE DATA	HNIS 04
80	OBC DATA WORD 1	OBC 02

Table 3-1 (continued)

MINOR FRAME WORD NR.	DESCRIPTION	ID
81	OBC DATA WORD 2	OBC
82	OBC DATA WORD 3	OBC
83	OBC DATA WORD 4	OBC
84	SCIENCE DATA	HNIS
85	SCIENCE DATA	HNIS
86	S/C CLOCK BITS 15-8	SMM
87	S/C CLOCK BITS 23-16	SMM
88	SCIENCE DATA	HNIS
89	SCIENCE DATA	HNIS
90	SCIENCE DATA	HNIS
91	SCIENCE DATA	HNIS
92	SCIENCE DATA	HNIS
93	SLIRA (COMP SHLD)	HNIBS
94	DATA BUS 2	UVSP
95	DATA BUS 2	UVSP
96	SUBCOMMUTATOR NR 3	CDH
97	SUBCOMMUTATOR NR 4	CDH
98	SUBCOMMUTATOR NR 5	CDH
99	SUBCOMMUTATOR NR 6	CDH
100	SCIENCE DATA	HNIS
101	SCIENCE DATA	HNIS
102	SCIENCE DATA	HNIBS 3
103	SCIENCE DATA	HNIBS 3
104	SCIENCE DATA	HNIS
105	SCIENCE DATA	HNIS
106	SCIENCE DATA	XRP
107	SCIENCE DATA	XRP
108	SCIENCE DATA	ACRIM
109	SUBCOM WORD	XRP
110	SCIENCE DATA	HNIS
111	SCIENCE DATA	HNIS
112	OBC DATA WORD 5	OBC
113	OBC DATA WORD 6	OBC
114	OBC DATA WORD 7	OBC
115	OBC DATA WORD 8	OBC
116	SCIENCE DATA	HNIS
117	SCIENCE DATA	HNIS
118	SCIENCE DATA	HNIBS 3
119	SCIENCE DATA	HNIBS 3
120	SCIENCE DATA	HNIS
121	SCIENCE DATA	HNIS
122	SCIENCE DATA	XRP
123	SCIENCE DATA	XRP
124		
125		
126	DATA BUS 3	UVSP
127	DATA BUS 3	UVSP

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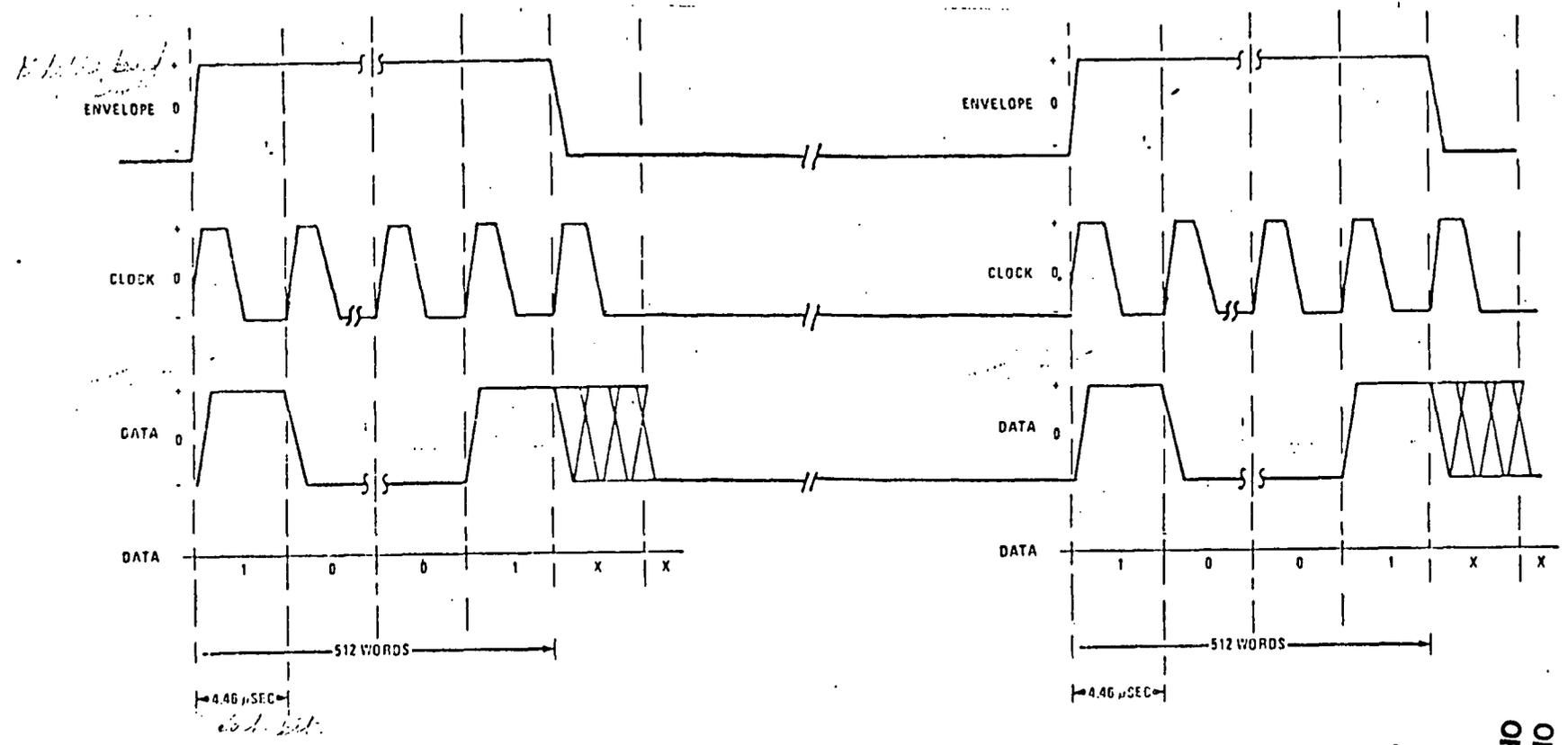
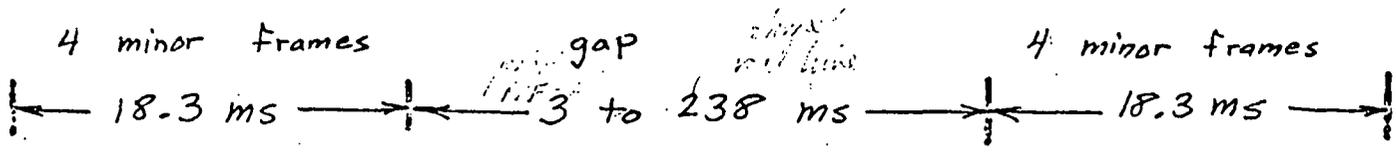
Telemetry Words Saved for UVSP

<u>minor byte #</u>	<u>contents</u>
30,31,62,63	data busses
94,95,126,127	<i>plus unit plus count of monitor</i>
27,28	DMA
29	status monitor
35,80,81,82,83	OBC messages <i>should include word to indicate in flow</i>
112,113,114,115	
32,99	subcom - includes CAS info
33,96,97,98	subcom - includes FPSS
<u>65</u>	<i>File printing can occur</i>
total = 27	<i>Jack - low bit in bit</i>

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DATA FROM EOP INTERFACE UNIT

6)



rate = 224 Kbps

error for some reason looks like it's not the same

NOTE: VOLTAGE: 5 < V
REF.: NASCOM 844-71-03

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Word 3 - Data Format (as seen by UVSP computer)

Format / bit #	7	6	5	4	3	2	1	0
HAO	x	x	x	0	0	x	x	x
Engineering	x	x	x	0	1	x	x	x
Science	x	x	x	1	0	x	x	x
Flexible	1	0	0	1	1	x	x	x
OBC dump	0	0	0	1	1	0	0	0
Self Test	1	1	1	1	1	1	1	1

Word 9 - Flags

flag / bit #	7	6	5	4	3	2	1	0
forward	-	0	-	-	-	-	-	-
backwards	-	1	-	-	-	-	-	-
self test	-	-	0	-	-	-	-	-
operations	-	-	1	<i>normally 1's everything</i>	-	-	-	-
good poly.	-	-	-	0	-	-	-	-
bad poly.	-	-	-	1	-	-	-	-
full block	-	-	-	-	0	-	-	-
partial block	-	-	-	-	1	-	-	-
real time	-	-	-	-	-	0	-	-
playback	-	-	-	-	-	-	1	-

Word 8 - Source ID

Station	ID # (octal)	ID # (decimal)
ACH	6	6
AGO	10	8
BDA	4	4
ETC	30	24
GDS	16	14
GWM	14	12
HAW	15	13
MAD	11	9
III	1	1
ORR	25	21
OUI	5	5
ULA	23	19
VNK	2	2
DEL	17	15

Word 67 - Dwell *add bit*

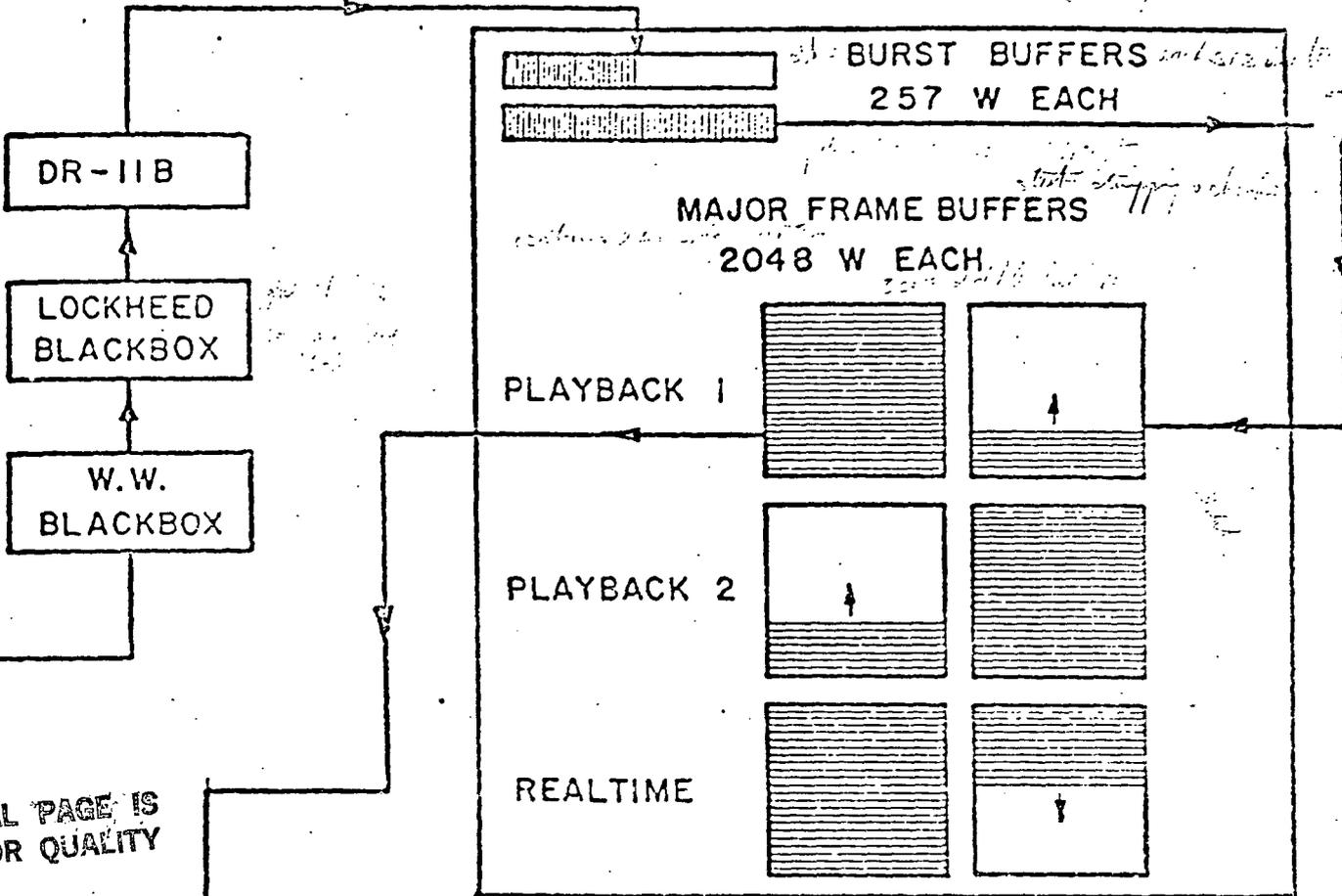
mode / bit #	7	6	5	4	3	2	1	0
dwell	1	dwell identifier						
not dwell	0	-	-	-	-	-	-	-

*an one bit in a row
was not of bits anything in dwell*

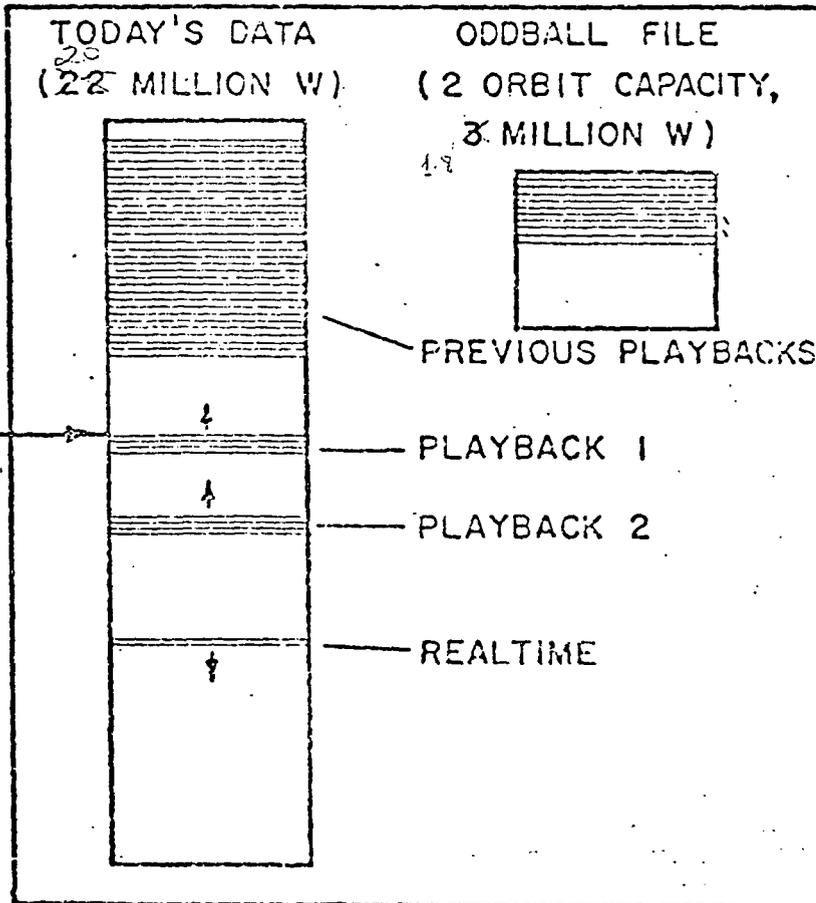
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DATA ACQUISITION

PDP 11/34



DISK



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recorder part of all record... card up on with... instructions...

Stripped Major Frame Format

a)

<u>BYTE #</u>	<u>CONTENTS</u>
0, 1	4761. (synch)
2, 3	clock 2, clock 3
4, 5	16 bit major frame counter
6	source ID (word 8)
7	status flags (word 9)
8 → 39 <small>256 bits</small>	minor frame map - 2 bits/minor ? <small>shows 2 minor</small> 00 ⇒ missing 01 ⇒ bad polynomial check 10 ⇒ bad synch 11 ⇒ good
40 → 55	date & time of reception
56	data format (word 3)
57	# of minor frame bytes stored
58 → 127	not used
128 → 154	minor # 0 bytes 30, 31, 62, 63, 94, 95, 126, 127, 27, 28, 29, 35, 80, 81, 82, 83, 112, 113, 114, 115, 32, 33, 96, 97, 98, 99, 65
155 → 181	minor # 1
⋮	⋮
1766 - 1792	minor # 127
= 26 bits	

128 byte header / minor frame

ACQ MESSAGES

ACQ... 18-OCT-79 16:48:49.7 ** NEW SOURCE, CODE = 000030
 ACQ... 18-OCT-79 16:48:50.3 **PARTIAL BLOCK, SOURCE CODE = 000030
 ACQ... 18-OCT-79 16:48:51.0 ILLEGAL MESSAGE
 ACQ... 18-OCT-79 16:48:51.4 MINOR FRAME OVERWRITE IN MAJOR # 5671.
 ACQ... 18-OCT-79 16:48:52.1 BAD SYNCH 172131
 ACQ... 18-OCT-79 16:48:52.6 DUPLICATED FRAME (STASH) 233
 ACQ... 18-OCT-79 16:48:53.1 OUT OF RANGE FRAME 24501
 ACQ... 18-OCT-79 16:48:53.7 MAJOR FRAME = 000000 000106
 ACQ... 18-OCT-79 16:48:54.3 TOO MANY SOURCES, CODES ARE 000006 000010 000030 000024
 ACQ... 18-OCT-79 16:48:55.1 FRAME RANGE IS 0 TO 10999
 ACQ... 18-OCT-79 16:48:55.6 THASH WRITING REC # 1023, MAJOR # 23
 ACQ... 18-OCT-79 16:48:56.3 STASH--REC # 2048, MAJOR # 48

LIST OF MESSAGES

ACQ... 8-OCT-79 08:55:46.3 MINOR FRAME OVERWRITE IN MAJOR # 8191.
 ACQ... 8-OCT-79 08:57:05.7 MINOR FRAME OVERWRITE IN MAJOR # 8219.
 ACQ... 8-OCT-79 08:57:30.6 MINOR FRAME OVERWRITE IN MAJOR # 8222.
 ACQ... 8-OCT-79 08:57:36.7 MINOR FRAME OVERWRITE IN MAJOR # 8223.
 ACQ... 8-OCT-79 08:57:44.0 BAD SYNCH 171777
 ACQ... 8-OCT-79 08:57:44.5 STASH--REC # 8192, MAJOR # 8192

CURRENT STATUS 8-OCT-79 08:57:47.0

SOURCES = 3.
 CODE [DEC]

STATUS [OCT]

15	000144
24	000144
14	000040

MISSED BURSTS = 1.

DISK 0 TO 10999 LAST = 8192

ACQ... 8-OCT-79 08:57:50.4 MINOR FRAME OVERWRITE IN MAJOR # 8191.
 ACQ... 8-OCT-79 08:58:22.1 MINOR FRAME OVERWRITE IN MAJOR # 1674.
 ACQ... 8-OCT-79 08:59:02.6 MINOR FRAME OVERWRITE IN MAJOR # 1679.
 ACQ... 8-OCT-79 08:59:14.1 STASH--REC # 1690, MAJOR # 1690
 ACQ... 8-OCT-79 08:59:54.8 MINOR FRAME OVERWRITE IN MAJOR # 1685.
 ACQ... 8-OCT-79 09:00:24.3 MINOR FRAME OVERWRITE IN MAJOR # 1639.
 ACQ... 8-OCT-79 09:00:42.7 MINOR FRAME OVERWRITE IN MAJOR # 1691.

CURRENT STATUS 8-OCT-79 09:02:49.1

EXPIRED SOURCE, ID WAS 15.
 EXPIRED SOURCE, ID WAS 24.

SOURCES = 1.

CODE [DEC]

STATUS [OCT]

14	000050
----	--------

MISSED BURSTS = 1.

DISK 0 TO 10999 LAST = 1690

SAMPLE OUTPUT

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APPENDIX 6
DATA TAPE FORMATS

R. Shine
October 27, 1983

Reading UVSP Data Tape

1 OVERVIEW

The data from the Ultraviolet Spectrometer and Polarimeter (UVSP) on board the Solar Maximum Mission satellite are organized as experiments which are archived on magnetic tapes. The format for these experiments was chosen to be consistent with that of the Colorado experiment on OSO-8 in order to allow easy adaptation of software already written for the earlier mission. The processing of the data was done using a PDP 11/34 computer running the RSX-11M operating system which determines some of the physical organization of the tapes. Users of either PDP's running RSX or VAX's running VMS should have easy access to the data tapes using software developed by the UVSP experiment team. Such users may have no need of this document.

Those interested in accessing UVSP without a machine that accepts the already developed software are advised to read both sections 2 and 3. Section 2 describes the physical format of the tapes and how to organize the data into experiments consisting of logical blocks of 512 bytes. Section 3 describes the contents and organization of each logical block in an experiment. Those who are able to easily copy the tape files onto disk may only need information from this section. Section 3 would also be necessary for anyone interested in developing independent software for manipulating UVSP data.

2 PHYSICAL TAPE FILES AND RECORDS

The tapes will generally be labeled with the experiment numbers contained on the tape. About 16,700 experiments have been run but not all are available because of telemetry gaps, etc. The tapes are 9 track 1600 bpi but other formats may be available on request.

The tapes are supposed to be a level 3 implementation of the June 19, 1974 Proposed Revision of the ANSI Standard Magnetic Tape Labels and File Structure for Information Interchange (X3.27-1969). If software is available to handle such structures, you may consider using it. If not, use the following guide to read the tapes and strip out the actual data contents. The tape structure is as follows:

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file #	record #	# bytes	contents
1	1	80	volume label 1 (2 records)
	2	80	
	3	80	label for first file (2 records)
	4	80	
2	1	512	first data file, may contain from 3 to over 800 records depending on amount of data, each record is 512 bytes long
	2	512	
	3	512	
		512	
		512	
3	1	80	end of file label (2 records)
	2	80	
4	1	80	file label for second file
	2	80	
5	1	512	second data file
		512	

more groups of 3
tape files for each
experiment file.

terminated by double end of file

All of the tape records which actually contain the experiment information have a length of 512 bytes. (This is not true of all ANSI tapes but is the standard for UVSP tapes). The 80 byte records just contain various labels. The only information in these labels of interest for UVSP tapes is the file name which is ASCII encoded in bytes 5 through 21 of the first record in each file label (bytes 1 through 4 contain the characters HDR1). This file name would be useful if several files were read from a tape and stored on disk but it is not necessary to identify the experiment.

There are possible exceptions to the structure shown above. Sometimes a tape may appear to have extra file(s) at the beginning, usually because of an error in positioning the start of tape. Anything before the first 80 byte record should be ignored. Sometimes a tape may have errors resulting in records with 1 byte more than they should have. Usually it is safe to just ignore the last byte.

To read the tapes on an arbitrary system, the programmer should have a routine which can read a tape record of length 512 bytes or less and return the actual length as a parameter. The experiments can then be easily identified as files which contain records of 512 bytes each. Any file which contains such records is a data file, any other file is a label file and can be ignored or used as desired. The number of records in an experiment file should always be 3 or more. The largest files can reach 800 or more records. The data is interpreted as 16 bit words (except for a few items in the headers) which may require byte swapping. On these tapes the first byte always represents the least significant bits of the 16 bit word while the second is the most significant bits. Many non-DEC machines have the reverse convention (IBM for example) which implies that you must swap the bytes in each pair. Often there is a flag on the tape read routine which will handle this problem. It may even be possible that some machines interpret the bit order in the bytes backwards although I've never seen this.

Once you have the experiment file records stashed somewhere (on disk or another tape) with the bytes in the proper order, the parameters of the experiment and the counts can be decoded as described below.

3 EXPERIMENT DATA STRUCTURE

Some knowledge of the type of data obtained by the UVSP may be helpful in understanding how the data is stored and how to extract it. Refer to Woodgate, et al, (Solar Physics, 65, p. 73, 1980) for some basic information.

The UVSP experiment files use 512 byte records as a basic building block. This is the size of a physical disk record on many computers and is therefore the basic I/O unit for reading and writing data. Each 512 byte consists of 256 16 bit words. The structure of a file is illustrated below:

block number	contents
1	file header block
2	record header block for logical record 1
3	data for logical record 1
	n blocks (last block may not be entirely filled)
	where n is defined in record header
	n is the same for all logical records
n+3	record header for logical record 2
n+4	data for logical record 2

The first block of each file is an experiment header which contains information about that experiment including a unique experiment number (the experiment numbers are strictly chronological with no known exceptions). Table I shows all the items in this file header. Item 20 is the number of logical records in the file. Note that not all of the 256 words are used.

A logical record (not to be confused with the tape records discussed above) consists of 2 or more blocks. All logical records for a given file have the same length. The first block in the logical record is a record header. The information format of a record header is shown in Table II. Note that the first two items are fixed values which can be used to verify that a given block is a record header. Following this is the actual data. Each data value is a 16 bit number representing the UVSP count. The number of blocks in a logical data record is always an integer. It can be computed from either the file header or the record header. The product of file header items 98 and 99 represents the number of data points in the record. This number rounded up to the next integer multiple of 256 can be used to obtain the number of blocks used for the data. Adding 1 for the record header results in the total blocks per record. The same dimension values are also contained in the record header in items 3 and 4.

Each logical record contains the photons counted by a particular UVSP detector in chronological order. The detector number and the time for the first data point are in the header. If more than one detector is on, their records are consecutive. Often, the experiment data for a given detector is split up into many logical records. If, for example, there are 3 detectors turned on, then the data for a given detector is contained in every third logical record. This detector nesting order is always consistent within a given file and directly corresponds to the nesting during the experiment.

The structure and length of the data records is related to the lengths and nesting orders of the mechanism loops. The following list shows where to find these in the file header:

mechanism	parameter	location in header
x raster	# of values	142
	increment	75
y raster	# of values	143
	increment	76
wavelength	# of values	78
	increment	79
polarimeter	# of values	84
	increment	85

The nesting order is available in items 56 through 59. Once the order is known, the data collected from all the records for a given detector can be considered a 5 dimensional array with the fifth dimension the repeat count (the last repeat may not be complete because of termination by night, etc). The size of the first 4 dimensions is determined by the loop lengths and the nesting. The number of points in each record will be a multiple of some of these loop lengths. This can always be determined by the entries in the file header, but the rule used may be of some use. When generating these files, the program examined the first 2 non-trivial inner loops. If their product was greater than 127, they are used to define each record "array". If not, the next loop length is included until the total product is greater than 127 and this becomes the record size. If the product never reaches 128, the repeat count is used. However, not all the repeats are necessarily used. They may be divided up among several records in order to keep the size 4096 words or less. (This restriction does not apply to cases not using the repeat count). Any non-trivial dimensions not included in the record array will be implicit in the sequence of records for a given detector. The motivation for this scheme was to insure that data blocks are at least half full (to avoid wasted space) and to divide the data into pieces that can usually fit into memory along with the analysis software. When this data is processed on the VAX, the first thing usually done is to clump all the data in the file into one big 6 dimensional array (the sixth dimension is the detectors). In the future the data may also be distributed in this form which will greatly simplify loading it into machines that can memory map the entire file.

VERSION # 7 AND EARLIER?
DESCRIPTION REV. MARCH 12, 1981
RECORD HEADER BLOCK

WORD CONTENTS

1 144444 (OCTAL) SYNCH PATTERN
2 144444
3 DIMENSION 1
4 DIMENSION 2
5 64077 (OCTAL) INTEGER CODE (I.E., DATA IS INTEGER TYPE)
6 # OF DATA POINTS (NOT INCLUDING EMPTY LOOPS)
7 DETECTOR #
13 RECORD #
20-21 32 BIT CLOCK FOR FIRST DATA POINT IN RECORD (1-4)
22 YEAR (E.G. 88)
23 MONTH #
24 DAY OF MONTH
25 HOUR
26 MINUTE
27 SECOND
28 MS
29 00Y
48-79 STATUS MONITOR AT START OF REC (CHANS. 63.8-38)
88 # OF SERVO ADJUSTMENTS IN THIS RECORD
81 TOTAL SERVO SHIFT IN STEPS
128 # OF ACTUAL DATA POINTS FOUND
101-102 MEAN TIME BETWEEN DATA POINTS IN INNER LOOP IN UNITS OF 16 MS (FL)
103, 104 MAXIMUM TIME GAP FOUND (FL PT)
105-106 MINIMUM
107-108 STANDARD DEVIATION (SAMPLE) (FL PT)
109 # OF INNER LOOPS
110 COUNT OF NEXT LOOP
111 COUNT OF NEXT LOOP
112 COUNT OF NEXT LOOP
113 COUNT OF EXECUTIONS IN THIS RECORD (OFTEN 8)
114-115 MEAN TIME BETWEEN INNER LOOPS
116-117 MEAN TIME BETWEEN 2ND LOOPS
118-119 MEAN TIME BETWEEN 3RD LOOPS
120-121 MEAN TIME BETWEEN OUTER LOOPS
122-123 MEAN TIME BETWEEN EXECUTIONS
124-125 MAX TIME BETWEEN INNER LOOPS
126-127 MAX FOR NEXT
128-129 MAX FOR NEXT
130-131 MAX FOR OUTER
132-133 MAX FOR EXECUTIONS
134-135 MIN TIME BETWEEN INNER LOOPS
136-137 MIN FOR NEXT
138-139 MIN FOR NEXT
140-141 MIN FOR OUTER
142-143 MIN FOR EXECUTIONS
144-145 STANDARD DEVIATION FOR INNER LOOP MEAN TIME
146-147 NEXT LOOP
148-149 NEXT LOOP
150-151 OUTER LOOP
152-153 EXECUTIONS
159 MINOR FRAME
168-223 DMA AT START OF RECORD (7788-7777) (MAY INCLUDE
INFO. FROM PREVIOUS EXPERIMENT)
224-255 STATUS MONITOR AT START OF RECORD (CHANS. 31-62)

NOTE - STATUS MONITOR DATA IN TWO SEGMENTS

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VERSION # 7

DESCRIPTION REV. MARCH 12, 1981
11/7/88

EXPERIMENT HEADER BLOCK

WORD CONTENTS

1 =4
2 RE-FORMATTER VERSION # (7)
3 =5
4 # OF CHAINED EXPS. OR # IF NONE
5 # IF COMPLETE, 1 IF BEGINNING MISSING, 2 IF END MISSING
18-13 START TIME(DAY,HR,MIN,SEC)
14-17 STOP TIME
20 NO. OF LOGICAL RECORDS
21 NO. WORDS PER LOG. REC.
22 STATION NO.
23 PITCH (ARC. SEC. *18)
24 YAW (")
25 ROLL (DEG. *188)
27 XCEN- RASTER X CENTER
26 YCEN- RASTER Y CENTER
29-30 TOTAL # OF DATA POINTS (1*4)
31-34 RAW FPSS WORDS
35-36 STARTING MAJOR FRAME (1*4)
37 MINOR
38-39 ENDING MAJOR FRAME (1*4)
40 MINOR
49 YEAR(E.G. 1988)
52 NO. OF DETECTORS IN INTERVAL 1
53 NO. OF DETECTORS IN INTERVAL 2
55 DETECTOR BALANCE FACTOR
56 LOOP NEST CODE FOR POLARIMETER(#=INNER)
57 RASTER X
58 RASTER Y
59 WLD
60 EXPERIMENT TYPE
61 NO. OF DETECTORS USED
62 SCI #
63 DETECTOR WORD(B BITS)
64 POLR. STATUS (#=OUT, 1= A-IN, 2=B-IN)
65 SLIT #
67-68 STARTING WLD STEP #(INTEGER*4)
69-70 ENDING
73-74 GATE TIME(SEC.) (FL.PT.)
75 RASTER DX
76 RASTER DY
78 NO. OF WLD STEPS
79 WLD STEP SIZE(DW)
82 REPEAT COUNT(# OF TIMES TO DO EXP.)
83 RASTER SIZE(NX*NY)
84 NO. OF POLR. POSITIONS
85 POLR. STEP SIZE
86 THRESHOLD LEVEL(DMA 772#)
87 SLIT LETTER CODE
88 CAL. WLD STEP
89 TACHOMETER (SERVO) INTERVAL (=1 FOR NO SERVO)
90 CAL. LOOP CODE (4=SERVO)
97 EXPERIMENT NO.
98 FIRST DIMENSION OF DATA ARRAY
99 SECOND DIMENSION
103-117 (FD FILES) ERROR SUMMARY
105-106 (PB FILES) WLD POSITION AT IMAX
107-121 (PB FILES) ERROR SUMMARY

134 MOST RECENT FLARE X COORD.
135 MOST RECENT FLARE Y COORD.
138 1ST DETECTOR #
139 2ND "
140 3RD "
141 4TH "
142 NO. RASTER X VALUES(NX)
143 NO. RASTER Y VALUES(NY)
144 CALIBRATION INTERVAL(=# IF NO CAL.)

145-158: EXP. RESULTS FROM DMA(INCLUDE LMAX, WORDS 105-6)

145 I MIN/2
146 MIN X
147 MIN Y
148 I MAX/2
149 MAX X
150 MAX Y
151 BLUE VEL.
152 BLUE INT./2
153 BLUE X
154 BLUE Y
155 RED VEL.
156 RED INT./2
157 RED X
158 RED Y

159-174: START-OF-EXPERIMENT DATA

159 SCI WORD
160 LIST INSTRUCTION REG.(JR)
161 LIST PROGRAM CTR.(JR)
162 OBS. LIST SEQUENCE(JR)
163-71 OBSERVING LIST (PARAMETER BLOCK)
172 EXP. NO.
173 XCEN
174 YCEN

175 WLD SCAN TYPE(1=ABS, 2=GBL OFFSET, 3=LCL OFFSET)

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65-35

58-40

58-38

58-32#

57-46

57-36

5254

ISV3

51-B

43-1

42-43#

41-37

3744

MEFRS

52-23 53-43 54-1 65-1A

POLL 2666 29-17 29-33 29-51 30-10 31-33# 31-57

33-48#

23-26

3052

ETNUL

57-344

5-39

5175

TWLP

DEOLD 5752 29-18 30-22 31-19 33-42 33-43 36-35 36-36 64-39H

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