

ERRATUM: IRAS OBSERVATIONS OF IRREGULAR GALAXIES

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In "IRAS Observations of Irregular Galaxies" by D. Hunter *et al.* and "Measuring Star Formation Rates in Blue Galaxies" by J. Gallagher and D. Hunter in these proceedings, as well as in D. Hunter *et al.* (1986, *Ap. J.*, 303, 171), galactic blue luminosities are based on standard optical definitions. Thus we derive  $L_B$  from the blue absolute magnitude  $M_B$  using

$$L_B = 7.81 \times 10^{34} \text{ dex}(-0.4 M_B) \text{ erg s}^{-1}, \tag{1}$$

or from the in band flux derived via

$$f_B = \int B_\lambda S_\lambda(B) d\lambda = \text{dex}(-0.4 B_T - 5.19) \text{ erg s}^{-1} \text{ cm}^{-2}, \tag{2}$$

where  $S_\lambda(B)$  is the Johnson B response function.

However, the  $L_B$  system adopted by de Jong *et al.* (1984, *Ap. J.* (Letters), 278, L67) for spiral galaxies was based on quasi-bolometric (rather than in band) fluxes given by

$$f_B^* = \lambda_B f_\lambda(4400 \text{ \AA}) \text{ erg s}^{-1} \text{ cm}^{-2}. \tag{3}$$

The  $L_B^*$  on this system are a factor of 4.5 times larger than the  $L_B$  from eq. (1). Thus our statements that the  $L(\text{IR})/L(\text{B})$  ratios for irregular galaxies are systematically higher than those of spirals are an incorrect result of comparing data on two different  $L_B$  systems. In fact, the irregulars cover roughly the same range in  $L(\text{IR})/L(\text{B})$  as the spirals when a consistent  $L_B$  system is used. A corrected  $L(\text{IR})/L(\text{B})$  versus  $S(100)/S(60)$  plot is given below.

