THEORETICAL STUDIES OF SOLAR LASERS AND CONVERTERS

By

John H. Heinbockel, Principal Investigator

Progress Report
For the period January 1988 to July 1988

Prepared for the
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ABSTRACT.
The previously constructed one dimensional model for the simulated operation of an iodine laser assumed that the perfluoroalkyl iodide gas $n-C_3F_7I$ was incompressible. The present study removes this simplifying assumption and considers $n-C_3F_7I$ as a compressible fluid.

Introduction.
The iodine laser being modeled uses the gas $n-C_3F_7I$ flowing in a cylindrical tube. As a first approximation to the chemical kinetics and flow it was assumed that the gas was incompressible and operating at a constant temperature. To remove these simplifying assumptions we consider the effects of fluid density variation as a function of both pressure and temperature in the flow. In particular we have added to our previous model:

(i) An equation of state.
We assume the equation of state of an ideal gas and write

$$ P = \eta RT \tag{1} $$

where $P$ is the gas pressure [Pa], $\eta$ is the gas density [Kg/m$^3$], $T$ is the temperature [$^\circ K$], and $R = R^*/296$ [J/Kg$^\circ K$], and $R^*$ is the gas constant, $R^* = 8.317$ [J/mole $^\circ K$]. For $n-C_3F_7I$ we use the value of 296 Kg/mole to convert $R^*$ to $R$.

(ii) Continuity equation.
The continuity equation expressing the conservation of mass flow can be represented

$$ \frac{\partial \eta}{\partial t} + \text{div}(\eta \vec{V}) = 0 \tag{2} $$

where $\eta$ is the fluid density [Kg/m$^3$], and $\vec{V}$ is the fluid velocity [m/s]. For steady state conditions and a flow in the axial direction we let $\vec{V} = \omega \hat{k}$ and reduce the continuity equation to the form

$$ \frac{\partial}{\partial z} (\eta \omega) = 0 \tag{3} $$

An integration of this equation gives

$$ \eta \omega = c_1 \tag{4} $$

where $c_1$ is a constant of integration.
(iii) Momentum equation.

The momentum equation for a control volume having a mass \( \eta d\tau \), where \( d\tau \) is an element of volume, is given by

\[
\dot{M} = \iiint \dot{V} \eta \, d\tau
\]

From Newton's second law we have

\[
\vec{F} = \frac{D\dot{M}}{Dt} = \frac{D}{Dt} \iiint \dot{V} \eta \, d\tau
\]

where \( D/Dt \) is the material derivative. We have

\[
\frac{D\dot{M}}{Dt} = \iiint \dot{V} \left( \eta \dot{V} \cdot d\vec{a} \right) + \frac{\partial}{\partial t} \iiint \dot{V} \eta \, d\tau
\]

where \( d\vec{a} \) is an element of surface area. The surface integral term in (7) represents the efflux of momentum through the control volume and the volume integral term in (7) represents the change in momentum inside the control volume. Using the Gauss–divergence theorem we can change the surface integral to a volume integral and write

\[
\iiint_S \dot{V} \left( \eta \dot{V} \cdot d\vec{a} \right) = \iiint \left[ \nabla \cdot \eta \ddot{V} \right] \, d\tau
\]

Then the momentum equation becomes

\[
\vec{F} = \frac{D\dot{M}}{Dt} = \iiint \text{div} \left( \eta \ddot{V} \right) \, d\tau + \frac{\partial}{\partial t} \iiint (\eta \omega \hat{k}) \, d\tau
\]

where \( \ddot{V} \) is the dyadic \( \omega \hat{k} \hat{k} \) for \( \vec{V} = \omega \hat{k} \), and

\[
\nabla \cdot \eta \ddot{V} = \frac{\partial}{\partial z} \left( \eta \omega^2 \right) \hat{k}
\]

We also make the substitution \( \vec{F} = \iiint \vec{f} \, d\tau \) where \( \vec{f} \) is the average force per unit volume as this allows the momentum equation to be written in the form

\[
\vec{F} = \iiint \vec{f} \, d\tau = \frac{D\dot{M}}{Dt} = \iiint \left[ \frac{\partial}{\partial t} (\eta \omega) \hat{k} + \frac{\partial}{\partial z} (\eta \omega^2) \hat{k} \right] \, d\tau
\]

Neglecting viscosity and using \( \vec{f} = -\nabla P \) as the average force per unit volume which is due to the fluid pressure \( P \), we have

\[
\frac{\partial P}{\partial z} = \frac{\partial}{\partial t} (\eta \omega) + \frac{\partial}{\partial z} (\eta \omega^2)
\]

From equation (4) we use \( \eta \omega = c_1 \) and examine the steady state form of the equation (12) to obtain

\[
\frac{\partial P}{\partial z} = c_1 \frac{\partial \omega}{\partial z}
\]

Integrating this equation gives

\[
P + c_1 \omega = c_2
\]

where \( c_2 \) is a constant of integration.
(iv) Energy Equation.

In terms of the specific enthalpy $h$ per unit mass, the energy equation for the fluid flow is

$$\eta \frac{Dh}{Dt} = \frac{DP}{Dt} + \kappa \nabla^2 T + q$$

where $P$ is the pressure, $T$ is the temperature, $\kappa$ is the thermal conductivity and $q = q(z)$ is the radiation heat flux. In one dimension this can be written

$$\eta \frac{dh}{dt} + \eta \omega \frac{dh}{dz} = \frac{\partial P}{\partial t} + \omega \frac{\partial P}{\partial z} + \kappa \frac{d^2 T}{dz^2} + q$$

The specific enthalpy $h$ can be expressed in terms of the specific heat at constant pressure and

$$h = C_p T$$

Using the relation

$$C_p^* - C_v^* = R^*$$

where $R^* = 8.317 [J/mole - °K]$ is the universal gas constant.

The specific heat at constant pressure $C_p^*$ can be expressed in terms of the specific heat at constant volume $C_v^*$. We use the numerical data

<table>
<thead>
<tr>
<th>Temperature °K</th>
<th>$C_v^*$ for $C_2F_7I [J/mole °K]$</th>
</tr>
</thead>
<tbody>
<tr>
<td>298.15</td>
<td>146</td>
</tr>
<tr>
<td>400.00</td>
<td>169</td>
</tr>
<tr>
<td>500.00</td>
<td>186</td>
</tr>
</tbody>
</table>

and assume that $C_v^*$ can be obtained from the relation

$$C_v^* = C_v^*(T) = \alpha_v \exp(\beta_v (T - 300)), \quad 298.15 \leq T \leq 500$$

By applying a least square's fit of the data with the exponential curve in (20) we obtain the best fit parameter values of

$$\alpha_v = 147.23, \quad \beta_v = 0.0012$$

This gives the steady state energy equation

$$\eta \omega [C_v(T) + R] \frac{dT}{dz} = \omega \frac{dP}{dz} + \kappa \frac{d^2 T}{dz^2} + q(z)$$

where $C_v(T)$ is the correct dimensional form of $C_v^*$.

Having neglected the viscosity terms, we also neglect the effect of thermal conductivity and integrate the equation

$$c_1 [C_v(T) + R]dT + c_1 \omega c_1 dw - q(z)dz = 0$$

to obtain

$$c_1 \int C_v(T)dT + c_1 R(T - 300) + \frac{c_1}{2} \omega^2 - Q(z) = c_4$$

where

$$Q(z) = \int q(z)dz$$

and $c_4$ is a constant of integration.
Implementation of Model.

In order to compare and verify the computer model with laboratory data it is assumed that we are given the boundary conditions as \( z = 0 \) and \( z = L \). We use the subscript 0 to denote values at \( z = 0 \) and subscript \( L \) to denote the values at \( z = L \). By setting \( z = 0 \) we obtain:

\[
\text{from (4)} \quad c_1 = \eta_0 w_0
\]

\[
\text{from (14)} \quad c_2 = P_0 + c_1 \omega_0
\]

from (24) we have

\[
c_4 = c_1 \left[ R(T_0 - 300) + \int C_v(T) \, dT \right] + \frac{c_1 \omega_0^2}{2}
\]

Assuming that \( g = g(z) \) can be represented by a constant, at \( z = L \) we obtain from the equation (24) the relation

\[
g_0 = \frac{1}{z_1} \left\{ c_1 \left[ R(T_L - 300) + \int C_v(T) \, dT \right] + \frac{c_1 \omega_L^2}{2} - c_4 \right\}
\]

Knowing the values of the above constants the temperature, density, pressure and flow rate as a function of distance \( z \) can be calculated as follows:

From the equation of state we can write

\[
P = \eta RT
\]

and multiplication by \( \omega \) gives

\[
P \omega = \eta R T = c_1 RT
\]

Using the relation (14) in the above equation gives

\[
(c_2 - c_1 \omega) \omega = c_1 RT
\]

or

\[
T = \frac{c_2 \omega - c_1 \omega^2}{c_1 R}
\]

We also use the scale factors

\[
SF_1 = \frac{10^3}{296} \text{ mole/Kg} \quad \text{and} \quad SF_2 = \frac{1.01325(10^5)}{760} \text{ (N/m}^2) / \text{Tor}
\]

where applicable to insure all equations balance dimensionally.

For

\[
Q(z) = \begin{cases} 
q_0 z, & 0 < z < z_1 \\
q_0 z_1, & z_1 < z < L
\end{cases}
\]

We may then solve for \( \omega = \omega(z) \) such that the energy equation

\[
F(\omega) = c_1 \int C_v(T) \, dT + c_1 R(T - 300) + \frac{c_1 \omega^2}{2} - Q(z) - c_4 = 0
\]

is satisfied where \( T \) is determined by the relation (25). Since this is a nonlinear equation, the method of interval halving is used. The value of \( \omega \) is changed until \( F(\omega) \) changes sign and then interval halving is used until the desired solution accuracy is obtained. The temperature \( T \) is then determined from (25), the density is obtained from the relation (4) and the pressure is obtained from the relation (1). We then have the flow rate, temperature, pressure and density as functions of distance \( z \).
Reaction Rates.

The reaction rate coefficients

\[ k_1, k_2, k_3, k_4, k_5, k_6, k_7, k_8, \]
\[ c_1, c_2, c_3, c_4, c_5, c_6, \]
\[ Q_1, Q_2, Q_3, Q_4, Q_5 \]

which occur in the chemical kinetics associated with the iodine laser can be assumed to vary with temperature. Various assumptions can be made concerning the effects of temperature on these rate coefficients. Various Arrenius type temperature variation effects can be modeled using one or more of the following expressions:

\[ K = k_0 e^{\theta(T-T_0)} \]
\[ K = k_0 e^{\theta/T} \]
\[ K = k_0 T^{-N} \]
\[ K = k_0 k T^{-N} e^{\theta/T} \]
\[ K = k_0 T^{-N} e^{\theta/T^m} \]

Since the coefficients in these expressions are not known, as a first approximation to temperature on the rate coefficients, we will keep most of the rate coefficients constant.

Incompressible flow model.

The incompressible flow model was presented in the previous progress report (reference 2). Nominal results from that model are illustrated in the figures 1 through 5 for various values of the flow rate \( \omega \). Values of \( \omega \) used in these graphs are:

\[ \omega = 750., 1750., 2750., 5500. \text{ cm/s} \]

while other parameters were held constant. These other parameters were:

\[ P = 23.5 \text{ Torr}, \]
\[ R_1 = .95848 \text{ reflectivity at left end of the tube}, \]
\[ R_2 = .81634 \text{ reflectivity at right end of the tube}, \]
\[ \text{concentration} = 1267 \text{ (peak concentration in solar constants.)} \]

Both the incompressible and compressible flow models are being modified to include temperature effects. Also an analysis by R.C. Costen of the pumping photon density is being studied and incorporated into the model. This analysis is given in the Appendix A along with figures of the results therein.
Figure 1. Log (C₃F₇) vs axial distance.
Figure 2. Log ($I_2$) vs axial distance.
Figure 3. Log (I*) vs axial distance.
Figure 4. Log (I) vs axial distance.
Figure 5. $\pm \log(|1 - \frac{1}{2}I|)$ vs axial distance.
References


Appendix A

Pumping Photon Density within Absorbing Lasant Gas
by R.C. Costen

Assumptions: There is a 2-D cylindrical geometry which is treated as a black body cavity.

A. Consider the relation of 2-D surface brightness $B_2(W/m^2)$ to the 2-D surface flux $F_2(W/m^2)$. Consider a one meter depth of a cylindrical cavity of arbitrary shape as illustrated in the figure. A.1

![Cylindrical geometry](image)

Figure A.1 Cylindrical geometry

We have

$$F_2 = \int_{-\frac{\pi}{4}}^{\frac{\pi}{4}} B_2 \cos \theta \, d\theta = 2B_2$$

W per m of perimeter entering the cavity per m of depth.
B. Determination of the pumping photon density at an interior point resulting from a length $dL$ of perimeter.

![Diagram of density at an interior point]

Figure A.2 Density at an interior point

The power in $d\theta$ due to a length $dL$ is given by

$$W = B_2 \cos \theta \, dL \, d\theta$$

This power $(J/s)$ is due to the photons moving at the velocity of light $c$ $(m/s)$. The energy density in $d\theta$ per m of depth is

$$\left(\frac{J}{m}\right)$$

Using $d\theta = \frac{dL}{r}$ and noting that the photon energy density $d\rho_2$ due to $dL$ is given by

$$d\rho_2 = \frac{B_2}{cr} \cos \theta \, dL \quad (J/m^2)$$

per m depth. For an absorbing gas we have

$$d\rho_2 = \frac{B_2}{cr} \exp\left(-\frac{r}{\delta}\right) \cos \theta \, dL \quad (J/m^2)$$

per m depth and where $\delta$, $(m)$ is the absorption length.
C. 2-D Circular cylinder with absorbing gas and uniform incident flux $F_2$.

At the point $P$ we have due to $dL$ that

$$d\rho = \frac{F_2}{2c} e^{-\frac{r_0}{r}} \cos \theta \ dL \quad (J/m^2)$$

per m of depth. Using

$$\frac{\cos \theta}{r} \ dL = d\psi$$

we can write

$$d\rho = \frac{F_2}{2c} e^{-\frac{r_0}{r}} \ d\psi$$

or

$$\rho = \frac{F_2}{2c} \int_{\psi_0}^{\psi} e^{-\frac{r_0}{r}} \ d\psi = \frac{F_2}{c} \int_{0}^{\psi} e^{-\frac{r_0}{r}} \ d\psi \quad (J/m^2)$$

per m of depth. Using the law of cosines

$$r = -r_1 \cos \psi + \sqrt{r_0^2 - r_1^2 \sin^2 \psi}$$

we have

$$\rho = \frac{F_2}{c} \int_{0}^{\psi} \exp \left( \frac{1}{\delta} (r_1 \cos \psi - \sqrt{r_0^2 - r_1^2 \sin^2 \psi}) \right) \ d\psi$$

Changing $r_1$ to $r$ the pumping photon density is given by:

$$\rho(r|r_0, \delta) = \frac{F_2}{c} \int_{0}^{\psi} \exp \left( \frac{1}{\delta} (r \cos \psi - \sqrt{r_0^2 - r^2 \sin^2 \psi}) \right) \ d\psi$$

in $J/m^2$ per m of depth.
D. Average pumping photon density

The average pumping photon density is given by

\[
< \rho(r_0, \delta) > = \frac{1}{\pi r_0^2} \int_0^{r_0} dr 2\pi r \rho(r|r_0, \delta)
\]

Letting \( s = r/\delta \) where \( s_0 = r_0/\delta \) is the laser tube radius per absorption length, we have

\[
\rho(r|r_0, \delta) = \frac{f_2}{c} \int_0^{\pi} \exp(s \cos \psi - \sqrt{s_0^2 - s^2 \sin^2 \psi}) \, d\psi
\]

and

\[
< \rho(r_0, \delta) > = \frac{\delta^2}{\pi r_0^2} \int_0^{\pi} d\left(\frac{r}{\delta}\right) 2\pi \left(\frac{r}{\delta}\right) \rho(s, s_0)
\]

\[
= \frac{2}{s_0^3} \int_0^{s_0} ds \, s \rho(s, s_0)
\]

<\rho(s_0)>

We define and compute

\[
\rho'(s, s_0) = \frac{1}{\pi} \int_0^{\pi} \exp(s \cos \psi - \sqrt{s_0^2 - s^2 \sin^2 \psi}) \, d\psi
\]

with

\[
< \rho'(s_0) > = \frac{2}{s_0^2} \int_0^{s_0} ds \, s \rho'(s, s_0)
\]

which is the average pumping photon density.

Then

\[
\rho(s, s_0) = \frac{\pi F_2}{c} \rho'(s, s_0)
\]

and

\[
< \rho(s_0) > = \frac{2\pi F_2}{c s_0^3} \int_0^{s_0} ds \, s \rho'(s, s_0)
\]

\[
= \frac{\pi F_2}{c} < \rho'(s_0) >
\]
Note:

\[ \lim_{\delta \to \infty} \rho'(s, s_0) = 1 \quad \lim_{\delta \to \infty} \rho(s, s_0) = \frac{\pi F_2}{c} \]

\[ \lim_{s_0 \to 0} < \rho'(s_0) > = 1 \quad \lim_{s_0 \to 0} < \rho(s_0) > = \frac{\pi F_2}{c} \]

and at the center \( s = 0 \) we have

\[ \rho'(0, s_0) = \frac{1}{\pi} \int_0^{\pi} e^{-s_0} d\psi = e^{-s_0} \]

and on the wall \( s = s_0 \) we have

\[ \rho'(s_0, s_0) = \frac{1}{\pi} \int_0^{\pi} e^0 d\psi = 1 \]
Figure A.5 $\langle \rho' \rangle$ vs $S_0$. 
Figure A.6(a) $\rho'$ vs $S/S_0$. 

$S_0 = 0.1$

$S_0 = 1.0$

$S_0 = 3.0$

$S_0 = 6.0$

$S_0 = 10.0$

$S_0 = 100.0$

$S_0$, NORMALIZED RADIUS

$\rho'$, SCALED PUMPING PHOTON DENSITY

$S/S_0$, NORMALIZED RADIUS

Figure A.6(a) $\rho'$ vs $S/S_0$. 

ORIGINAL PAGE IS OF POOR QUALITY
Figure A.7 $<\rho'>$ vs $S_0$ oil semi-log scale.