RADIATION MONITORING EQUIPMENT DOSIMETER EXPERIMENT

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INTRODUCTION

Spacecraft crews risk exposure to relatively high levels of ionizing radiation. This radiation may come from charged particles trapped in the Earth’s magnetic field, charged particles released by solar flare activity, galactic cosmic radiation, energetic photons and neutrons generated by interaction of these primary radiations with the spacecraft and crew, and man-made sources (e.g., nuclear power generators). As mission are directed to higher radiation level orbits, viz., higher altitudes and inclinations, longer durations, and increased flight frequency, radiation exposure could well become a major factor for crew stay times and career lengths. To more accurately define the radiological exposure and risk to the crew, active, real-time radiation monitoring instrumentation must be flown capable of identifying and measuring the various radiation components. This presentation describes a radiation dosimeter instrument which has been successfully flown on the Space Shuttle, the RME-III.
The RME-III is a self-contained, portable, active (real-time) dosimeter system developed by E.G.&G., Inc. for the U.S. Air Force (USAF) and is adapted for use in measuring the ionizing radiation environment on the Space Shuttle. RME stands for "radiation monitoring equipment", the title given to 2 prototype radiation instruments sucessfully flown by the USAF on several pre-Challenger missions (1). These prototype instruments had relatively short battery life and very limited memory capacity; thereby, requiring an inordinate amount of crew interaction time. The RME-III was developed to incorporate the capabilities of both radiation instruments into a single unit and to minimize crew interaction times with longer battery life and expanded memory capacity. A description of the RME-III instrument and the results from flight of several Space Shuttle missions is presented in this report.

INSTRUMENT DESCRIPTION

The RME-III is a portable, self-contained, active dosimeter system. The system features a 3-channel tissue equivalent proportional counter (TEPC) which measures particle fluence and computes dose and dose equivalent at operator selected time intervals. The total accumulated absorbed dose and dose equivalent are displayed real-time on a liquid crystal display while the data and time of the interval dose readings are stored in memory modules for future analysis. Analysis of the time-resolved data permits correlation of the radiation exposure with geographic position, altitude and instrument location in the space craft. The instrument consists of 2 components: (1) the
Main Module, and (2) the Memory Module. The memory module is
designed to plug into the Main Module and is exchanged with a new
module when the memory capacity or primary system battery life is
reached. The overall dimensions of the system are 28.6 cm x 10.2
cm x 5.1 cm. The combined total weight is 1550 g (3.3 lbs).

Main Module Description

The main module contains the analog and digital operating
circuitry of the instrument and the TEPC. At the top of the
module are 2 locking toggle switches. One is the system power
"ON-OFF" switch. The other is a keyboard locking switch, which
prevents accidental entries via the keyboard during system
operation. The display is a 6.4 cm x 3.8 cm liquid crystal
display (LCD), displaying 8 lines of data with 20 characters per
line. The system operating parameters are entered via a 12 key
keyboard. Primary operating power is provided by 5 AA alkaline
batteries in the memory modules. Zinc-air batteries are used in
the Main Module to maintain the instrument operating parameters
and internal operating clocks during Memory Module replacement.

Memory Module Description

The Memory Module contains a 0.5 Megabyte Random Access
Memory (RAM) into which the RME-III data is stored for future
analysis. The RAM is also backed up by zinc-air batteries. The
primary operating power for the entire system is provided by 5 AA
batteries in the Memory Module, providing approximately 42 hours
of continuous operation. The dimensions of the Memory Modules
are 10.2 cm x 10.7 cm x 4.4 cm. The stand-alone weight of the
Memory Module is 595 grams. Several Memory Modules are flown
with the Main Module on a Shuttle mission, sufficient enough to cover the entire mission.

**TEPC Description**

The heart of the system is the cylindrical tissue equivalent proportional counter developed by E.G.&G. The TEPC is a cylindrical (1.27 cm ID, 5.05 cm long) ionization chamber, lined with tissue equivalent plastic (type A-150), and filled with methane gas at reduced pressure (80 mmHg). The reduced gas pressure allows the chamber to simulate a tissue volume of about 2 micrometers in diameter. The TEPC is designed to measure the rate of energy loss of individual particles passing through the detector. The RME-III TEPC permits measurement of the radiation energy loss in a simulated volume on the order of the size of a typical human cell. Thus, the RME-III is a microdosimetry based instrument.

**RME System Operation**

The system block diagram of the RME is shown in the next figure. Pulses from the TEPC are amplified and shaped by the pre-amplifier and linear amplifier, and are fed into comparators that separate the spectrum into 3 pulse height or linear energy transfer (LET) bins: 0.35 to 6.5 keV/um, 6.5 to 30 keV/um, and 30 to 120 keV/um. Pulses greater than 120 keV/um are placed in the highest bin. The comparators in effect serve as a 3-channel pulse height analyzer. The actual counts in each pulse height bin are stored in the memory and can be used after-the-fact to more accurately estimate exposure, if required.

The instrument uses an algorithm to convert counts in each
pulse height bin into absorbed dose exposure values. For rads
tissue a multiplicative constant is obtained for each of the 3
bins by exposing the instrument of Cs-137 and Cf-252 sources. The
Cs-137 gamma source is used to determine the conversion constant
in the low pulse height energy bin. The Cf-252 spontaneous
fission source permits determination of the constants for the two
higher energy pulse height bins. The constants determined by the
calibration procedure are stored in the Main Module processor.
The instrument calibration is checked before and after each
mission.

The dose equivalent is computed by multiplying the measured
absorbed dose per channel by a present quality factor. The
quality factors used in the Shuttle missions presented in this
report were 1 for the lowest pulse height channel, 9 for the
intermediate channel, and 15 for the upper channel.

RME-III Setup Menu

The RME-III is menu driven. When the RME is initially turned
on the set-up menu appears as shown in DISPLAY 1. The operator
can the select the following operational parameters:

a. Time interval for dose integration. Ten seconds is the
value that is used on the Shuttle.

b. Alarm levels for total dose and dose rate. These are
normally set to OFF on the Shuttle flights.

c. Calendar Date

d. Coordinated universal Time (UTC)

e. Mission Elapsed Time (MET)

Normally all of the operating parameters are preset by the
ground crew at L-5 with the exception of the MET. The MET is entered by the Shuttle crewmember when the unit is turned on after launch. After presetting by the ground crew, the system is turned off and stowed in a mid-deck locker on the Shuttle. After lift-off the RME-III is removed from the locker by the crewmember, the power turned on, the keyboard unlocked and the MET entered. The crewmember then presses the START button, locks the keyboard and places the RME in the appropriate operating location. DISPLAY 3 is the display the operator normally sees during operation. The Shuttle crewmembers are trained in the operation of the RME and can be called upon to change or correct the operating parameters during flight, if required.

DATA

Data may be read out directly from any memory module by using the Main Module display and control keys. This function is a lengthy process since there are usually many sets of data (30,000 to 50,000) for a typical Shuttle mission. The normal procedure is to download the data from the Memory Modules to a personal computer (PC) after the mission. This task is accomplished by attaching the Memory Module to a special readout module which transfers the data from the Memory Module to a file in the PC via an RS-232 protocol. An 80C31 processor with data conversion algorithms is used to handle the transfer and conversion of the Memory Module data files. These files contain the raw data, the timing information (date, UTC, MET, and interval length), the absorbed dose in rads and the dose equivalent in rem for each time interval, and the integral dose in rads and rem for the
whole time the Memory Module recorded data.

The next figure shows representative data files obtained during Shuttle missions. The top file shows the data obtained at the start of one data set from a memory module. The counts in channel #1 are the low LET counts. The counts in Channels #2 and #3 are the intermediate and the high LET counts. The data in this file is representative of the count rate data obtained while traversing the equatorial latitudes, where the radiation background is due primarily to high energy cosmic rays. The middle file is an example of count rate data obtained at higher latitudes, where weakening of the Earth's magnetic field permits more cosmic ray particles to reach the Shuttle. The bottom file is an example of data collected during a pass through the South Atlantic Anomaly (SAA) on STS-31. The bulk of the dose is due to the high energy, low LET component of the inner Van Allen Belt protons, but the increased counts in the intermediate and high-LET channels indicate the presence of either high-LET, low energy protons, and/or the presence of high LET secondaries, which have a marked effect on the resulting dose equivalent.

The data for the first six Shuttle missions flown by the RME-III are summarized in Table I. Shown for comparison are the average Passive Radiation Dosimeters (PRD) doses measured on the Shuttle with thermoluminescent dosimeters by NASA/JSC. Six PRDs are routinely flown at different locations in the Shuttle. The highest doses/dose rates are measured in the high altitude missions which penetrate the SAA. In these orbits the dose contribution is due primarily to the Van Allen Belt protons. In the lower altitude orbits, the Galactic Cosmic Ray component predominates.
The data files are in ASCII or binary format and can readily be manipulated to produce plots of counts or dose vs. time or position. The next figures represent data obtained from various Space Shuttle missions displayed in this format. Figure 8 illustrates dose rate at 10s intervals vs MET for a low altitude, high inclination mission in which there is little interaction with the SAA and the exposures are dominated by the cosmic ray background. Notice the cyclic nature of the cosmic ray background with the peaks occurring at the high latitudes and the valleys at the equator caused by the geomagnetic shielding effect of the Earth. Figure 9 illustrates dose rate vs. MET for a Shuttle mission in which the altitude was sufficiently high to penetrate deeply into the SAA. Notice the peaks due to the penetration of the SAA. Dose rates on the order of several hundred millirad per hour were measured on this mission during some of the SAA passes.

The RME-III data readily lends itself to other forms of analysis. Figure 10 is a plot of dose rate as a function of orbiter geographic position (longitude and latitude) which were made from 38241 individual measurements taken during 113 hours of operation on the STS-28 Space Shuttle mission. The next figure (fig. 11) illustrates a plot of the South Atlantic Anomaly in two dimensions obtained from Space Shuttle RME-III data.

CONCLUSION AND DISCUSSION

Flight data from the Space Shuttle missions has demonstrated that a microdosimetric based radiation instrument such as the RME-III with the capability of registering counts and exposures vs. MET and UTC can be used to accurately assess dose form various sources of exposure, such as that encountered in the complex radiation environment of space. The RME-III is presently manifested to fly on STS-53 next month, and on STS-56 and -51 in 1993. Our goal is to fly
the RME-III on at least two or three more flights in 1994, the point of solar
minimum. The RME-III flew its first mission on STS-27 in Dec 88 just before
solar maximum. The RME-III is the only real-time radiation instrument to fly
on the Shuttle from Solar Maximum to Solar Minimum. The data from all of these
mission are being analyzed. The data is being used to enhance spacecraft shield-
ing and space radiation models; and to investigate phenomena such as the
drift of the SAA with time, the variation of the GCR background in LEO with
the solar cycle, and diurnal (day vs. night) variations in the space radiation
quality factors.

"Smart" dosimetric instrumentation such as the RME-III have definite
applicability to the Aerospace operational environment. Newer instruments
with larger memory capacities and greater spectral resolution are currently
being developed which will ultimately be flown on the Space Station Freedom
and future long duration space missions. Modified versions of the RME-III
will soon be used by the Air Force to investigate the cosmic ray and solar
radiation exposure of air crews at high altitudes.

REFERENCES
1. Cash, S. E., Madonna, R. G., McClellan, M. R., Steskal, M. J., and Fields,
M. E. Results for Radiation Monitoring Equipment Experiments of STS-41-C, 41D,
RME-III

RADIATION MONITORING EQUIPMENT

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USAF Armstrong Laboratory
The Space Radiation Environment

SOLAR PARTICLE EVENT
(Protons to Iron Nuclei)

INNERTH RADIATION BELT
(Protons)

DEBRIS CLOUD
(Photons, Electrons, Ionized Weapon Debris)

OUTER RADIATION BELT
(Electrons)

EXOATMOSPHERIC NUCLEAR DETONATION
(Prompt Neutron and Photons)

TRAPPED ELECTRONS
(Electrons)

SOUTH ATLANTIC ANOMALY
(Protons)

GALACTIC COSMIC RADIATION (GCR)
(Protons to Iron Nuclei)
RME III  BLOCK DIAGRAM
Figure 3. RHE-III operating displays.
### Figure 4a. Count rate data at equatorial latitudes (STS-27)

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<th>GMT</th>
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<th>CNT#2</th>
<th>CNT#3</th>
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### Figure 4c. Count rate data in South Atlantic Anomaly (SAA) measured on (STS-31)

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<td>Shuttle Mission</td>
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<td>RME run Time (hrs.)</td>
<td>RME Total Dose (mrad)</td>
<td>RME Total Dose Equiv. (mrem)</td>
<td>Average Q.F.</td>
<td>Av. PDR Dose (mrad)</td>
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</tbody>
</table>
STS-31 RME-III DATA
MET 0/08:11:36:0 TO 5/03:59:43:0

TOTAL DOSE RATE (uGy/min)

MET (DAYS)

26.5°: 612.93 km x 575.47 km
: 616.27 km x 609.33 km
Middeck Starboard Bulkhead
STS-28
RME-III Measured Dose Rate

NOTES
08 AUG 89 -- 13 AUG 89
SPACECRAFT ALTITUDE = 305 KM
ORBIT INCLINATION = 67.0 DEGREES
RME DATA PLOTTED FOR THE TIMES
MET 0/ 6:58:13:0 TO 4/17:17:52:0
Fig. 1. Isodose contours derived from the STS-28 RME-III data. The contours were derived from 33,000 data points which were smoothed over a 249 x 249 element latitude-longitude array.

STANDARD
RME-III Measured Dose Rate (uGy/min)

NOTES
MISSION DURATION: 09 AUG 89 -- 13 AUG 89
ORBITAL INCLINATION: 67.9 DEGREES / AVERAGE ALTITUDE: 302 KM
PLOTTED DATA SET DURATION: MET 9/8 8:58: 3 TO 4/17 17:52
RECORDS IN DATA SET: 38340
MAP PROJECTION: CYLINDRICAL EQUIDISTANT
DATA SMOOTHING GRID (LATxLOH): 249 x 249
FLUX CONTOURS: uGy/min AT x10 INTERVALS
MAGNETIC FIELD MODEL: IGRF 1985
MAGNETIC FIELD EPOCH: 1986.7

DATA SUMMARY
MAXIMUM DOSE RATE: 2.40 uGy/min
MAXIMUM SMOOTHED DOSE RATE: 1.84 uGy/min
MAXIMUM DOSE EQUIVALENT RATE: 4.17 uGy/min
MAXIMUM L-LET COUNT RATE: 282 cts/10 sec
MAXIMUM M-LET COUNT RATE: 7 cts/10 sec
MAXIMUM N-LET COUNT RATE: 5 cts/10 sec
RANGE OF G VALUES (gauss): 0.2067 -- 0.5768
RANGE OF L VALUES (Earth Radii): 0.85 -- 10.76

M.J. GOLIGHTLY/NASA JSC SN31(713) 483-8190
Outline

Purpose: To Understand the optical emissions that arise from the interactions between spacecraft in low earth orbit and the ambient atmosphere.

Procedures: Instrument Description
Experiment Description

Results: Shuttle Plume Emission
Shuttle Glow
Airglow

Conclusions: Lessons Learned
Future Experiments

8 November 1992
Purpose:

Satellite Assessment:

Optical Signatures for Targeting and tracking

Understanding the Physics that give rise to the spectral, spatial, and temporal variations in the optical emissions around space craft.

Characterizing the spectral, spatial, and temporal variations in the background airglow emissions

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Flight Experiments Technical Interchange Meeting

Instrument Components:

- **Lenses (2)**: 50 mm, f-1.2. (NASA Provided Hardware)
- **Slit**: 25 by 0.03 mm, 1.1 nm resolution
- **Grating**: 600 lines/mm, blazed at ~500 nm
- **Intensifier**: S-20, 25 mm photocathode
- **Camera**: Nikon 35 mm (NASA Provided Hardware)
- **Film**: 1600 ASA, Kodak Ektapress, color negative

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Experiment Description:

Thruster Plumes
  Spectral Assessment
  Near vs far field observations
  Ram angle Dependance
  Comparison with other space observations and with ground

Shuttle Glow
  Spectral Assessment
  Thruster Effects
  Altitude Effects

Airglow Observations
  Spectral Assessment
  Identification of Emissions and emission altitudes

* Future Experiment or Analysis

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Procedures:

Spectrograph mounted in aft-flight-deck.
Aimed towards tail of orbiter
Slit parallel to centerline of plume
Series of 4 sec exposures

Condition:

<p>| | |</p>
<table>
<thead>
<tr>
<th></th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>Orbiter Altitude</td>
<td>240 km (night)</td>
</tr>
<tr>
<td>Orbiter Attitude</td>
<td>Bay to Earth</td>
</tr>
<tr>
<td></td>
<td>Right wing to ram</td>
</tr>
<tr>
<td>Thruster</td>
<td>From OMS pod into ram</td>
</tr>
<tr>
<td></td>
<td>820 lb Thruster</td>
</tr>
<tr>
<td></td>
<td>Fired for 4 seconds</td>
</tr>
<tr>
<td></td>
<td>MMH and H₂O₄ film cooled</td>
</tr>
</tbody>
</table>

8 November 1992
Synthetic Spectrum of HNO (000)-(000)

APE Data of Shuttle Thruster Plume

Laboratory Spectrum of HNO created by the mechanism H + NO
Results:

Thruster Plumes

Spectral Features Identified:  
- CN (B→X) $\Delta v=0,-1,-2$
- CH (B→X) $\Delta v=0$
- C$_2$ (A→X) Swan Bands
- NO$_2$ continuum
- HNO$^-$
- Na
- O$^1$D and O$^1$S$^-$

Spatial Features:

- CN, CH, C$_2$, Na NO$_2$ and HNO O$^1$D and O$^1$S  
  - Decrease Rapidly from nozzle, Down to zero within 1 meter
  - Rise and peak at 1 meter, then fall to 20% at 3 meters
  - Seem to get stronger with increasing distance. O$^1$S more evident in ram burns than in wake burns.

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Shuttle Glow

Confirmation of NO₂ as the primary emitting species in the visible
Proper calculation of window transmission resulting in red-shift in the shuttle glow.

![Graph with wavelengths and intensity]

8 November 1992
Airglow

Identification of NO₂ in the atmosphere.
Observation of a Double airglow layer.

INTENSITY (ARBITRARY UNITS)

HEIGHT (km)

NO₂
OH (6-1)
OH (7-2)
OH (8-3)
OH (6-0)
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Conclusions:

Lessons Learned

Small mid-deck locker experiments provide a short turn around and a very high return for the amount of effort.

Shuttle is a difficult platform from which to do science.

Safety requirements

Large attitude dead-band and uncertainty

Windows with poor optical quality for observations

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Future Experiments

Several experiments planned for the next 12 months.

Further map the spatial and temporal distribution of plumes

Extend the wavelength observations into the UV and IR (APE-C)

Speed up the data rate by using Video instead of Film.

Collaborate the observations with other experiments