Electron–ion recombination rate coefficient measurements in a flowing afterglow plasma

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Abstract

The flowing-afterglow technique in conjunction with computer modeling of the flowing plasma has been used to determine accurate dissociative-recombination rate coefficients $\alpha$ for the ions $O_2^+$, HCO$^+$, CH$^+$, C$_2$H$^+$, H$_3$O$^+$, CO$^+$, HCO$^-$, HN$_2$O$^+$, and N$_2$O$^+$ at 295 K. We find that the simple form of data analysis that was employed in earlier experiments was adequate and we largely confirm earlier results. In the case of HCO$^+$ ions, published coefficients range from $1.1 \times 10^{-7}$ to $2.8 \times 10^{-7}$ cm$^3$/s, while our measurements give a value of $1.9 \times 10^{-7}$ cm$^3$/s.

1. Introduction

Rate coefficients for dissociative recombination (DR) of positive molecular ions with electrons are often needed to construct models of cold, weakly ionized plasmas, e.g. terrestrial and planetary ionospheres [1] and interstellar gas clouds [2,3]. Many ions have been studied subsequent to the work of Biondi et al. who used the stationary afterglow technique [4–10]. More recently other techniques, especially flowing afterglows [11–18], but also merged beams [19,20], and ion storage rings [21] have become important tools for such work. The chemical versatility of flowing afterglows has made this technique particularly productive, but a rigorous analysis of the data obtained by this method is a rather difficult task. In this article we present measurements of the recombination coefficients for several polyatomic ions obtained using a flowing afterglow/Langmuir probe (FALP) apparatus. In part, the present study was motivated by the recent finding [15,22] that the de-ionization of H$_3^+$ and D$_3^+$ afterglow plasmas deviates significantly from the simple behavior that is expected for binary recombination of a plasma containing a single ion species. The commonly used approximate methods of deducing recombination rates from graphs of the reciprocal electron density versus time or position tend to mask deviations from a simple recombination decay, and one can never be sure that the inferred recombination coefficients are truly constant during the afterglow. We have sought to improve this situation by developing better approximations for the data analysis which would reveal subtle deviations from a simple recombination decay.

2. Experimental

Our flowing afterglow technique [22] and the use of Langmuir probes for electron density measurements [23] have been described in detail elsewhere and will only be summarized here. A microwave
discharge in helium is used to generate helium metastable atoms, He\(^+\), which further downstream ionize argon by Penning ionization. In this work, several of the desired ion species were formed by charge transfer from Ar\(^+\).

\[
Ar^+ + XY \rightarrow XY^+ + Ar, \quad (1)
\]

where XY is O\(_2\), CO\(_2\), or N\(_2\)O. To make C\(_2\)H\(_4\)\(^+\) ions, CH\(_4\) was added to the afterglow leading to the reactions

\[
Ar^+ + CH_4 \rightarrow CH_3^+ + H + Ar \quad (2)
\]

\[
CH_3^+ + CH_4 \rightarrow C_2H_5^+ + H_2. \quad (3)
\]

Protonated molecules were produced by first making H\(_3^+\) and subsequently transferring the proton to the molecule under study, i.e.,

\[
Ar^+ + H_2 \rightarrow ArH^+ + H \quad (4)
\]

\[
ArH^+ + H_2 \rightarrow H^+ + Ar \quad (5)
\]

\[
H_3^+ + XY \rightarrow HXY^+ + H_2 \quad (6)
\]

where XY is CH\(_4\), CO\(_2\), N\(_2\)O, CO, and H\(_2\)O. The quadrupole mass spectrometer at the downstream end of the tube was used to monitor the ion composition of the plasma and to ascertain that no significant impurity ions were present.

The movable Langmuir probe measured the electron density as a function of distance \(z\). Under typical experimental conditions the electron densities in the upstream region of the flow tube were of the order of \(5 \times 10^{10}\) cm\(^{-3}\), the flow tube pressure was 0.9 Torr, and the central velocity was \(\sim 5500\) cm/s.

2.1. Data analysis

The usual practice of deducing recombination coefficients from the slopes of the inverse electron density versus distance \(z\) plots is justified only if ambipolar diffusion in radial and axial directions can be safely ignored. The on-axis electron density then follows approximately the equation

\[
\frac{1}{n_e(z)} = \frac{1}{n_e(z = 0)} + a \frac{z}{v} \quad (7)
\]

where \(v\) is an 'effective flow velocity' whose precise relation to the central flow velocity does not seem to be well defined. While it is true that the initial slope of a graph of Eq. (7) using experimental values for \(n_e\) often yields good values for the recombination coefficient, the problem is that further downstream the transition from a recombination-controlled to a diffusion-controlled afterglow leads to a gradual increase in the slope. In the absence of a better method of analysis, it is difficult to distinguish this effect from an actual change of the apparent recombination coefficient which may be present. Such changes may result from changes in the ion composition, quenching of vibrationally excited ions, or contributions of electron-stabilized recombination.

We have developed an accurate numerical model of the afterglow processes [22] that includes the ion chemistry and recombination, mixing of reagent gases, diffusion and transport of all species in both axial and radial directions in a flow with a parabolic velocity profile. In practice, however, the calculation time was found to be extremely long for routine deduction of recombination coefficients, and the model was used mainly to validate a simpler, approximate method of analysis.

In the approximate solution of the electron continuity equation the gas flow velocity is taken to be radially independent and the radial variation of the electron density is approximated by the fundamental diffusion mode in a cylinder. With these simplifications the continuity equation for the on axis electron density is of the form

\[
\frac{\partial n_e(r = 0)}{\partial z} = \frac{1}{v_o} \left[-\alpha n_e^2 + D_o \left(\frac{\partial^2 n_e}{\partial z^2} - \frac{n_e}{\Lambda^2}\right)\right]. \quad (8)
\]

where \(\Lambda = R/2.405\) is the fundamental diffusion length for a cylinder of diameter \(R\), \(D_o\) is the ambipolar diffusion coefficient and \(v_o\) is the central flow velocity. A fourth-order polynomial least-squares fit to the measured \(z\)-dependence of \(1/n_e\) is inserted into Eq. (8), the needed derivatives are obtained analytically, and the equation is then solved for \(\alpha(z)\). If the resulting \(\alpha(z)\) is independent of \(z\), then it can be concluded that a single, constant recombination coefficient prevails during the afterglow. This approximate method of data analysis was tested by creating computer-simulated data using the accurate numerical model and then analyzing them using the approximate method. The results confirmed
that the approximate method is entirely satisfactory for analyzing experimental data.

Ambipolar diffusion coefficients for the different species are needed as input parameters for Eq. (8). For most ions they were calculated from the known mobilities [24–26] in argon and helium using the Einstein formula relating the diffusion coefficients to the mobilities. In cases where the mobilities in argon were not known the Langevin formula [27] was used to obtain an estimate. The central velocity \( v_0 \) was calculated from the known gas flow rate and the total pressure.

3. Results and discussion

The recombination coefficients were measured for the ion species listed in Table 1. Before data were recorded, the mass spectrum was taken to verify that the desired ion species indeed dominated the plasma. As a test case for the technique and the data analysis the recombination coefficient of \( \text{O}_2^+ \) ions was measured. A value of \( (1.9 \pm 0.2) \times 10^{-7} \text{ cm}^3/\text{s} \) was obtained, in good agreement with the generally accepted value of \( (2.0 \pm 0.2) \times 10^{-7} \text{ cm}^3/\text{s} \). Fig. 1 shows an example of the electron-density decay curve.

<table>
<thead>
<tr>
<th>Species</th>
<th>Present work ( \alpha ) (( 10^{-7} \text{ cm}^3/\text{s} ))</th>
<th>Previously reported values</th>
<th>Reference</th>
<th>Method</th>
</tr>
</thead>
<tbody>
<tr>
<td>( \text{O}_2^+ )</td>
<td>1.9 ± 0.2</td>
<td>1.95 Mehr [4]</td>
<td>SA</td>
<td></td>
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<tr>
<td></td>
<td>2.20 Kasner [5]</td>
<td>SA</td>
<td></td>
<td></td>
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<tr>
<td></td>
<td>2.30 Mul [19]</td>
<td>MB</td>
<td></td>
<td></td>
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<tr>
<td></td>
<td>1.90 Wallis [26]</td>
<td>IT</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>2.00 Spaniel [13]</td>
<td>FALP</td>
<td></td>
<td></td>
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<tr>
<td>( \text{HCO}^- )</td>
<td>1.9 ± 0.3</td>
<td>2.0 Leu [6]</td>
<td>SA</td>
<td></td>
</tr>
<tr>
<td></td>
<td>1.1 Adams [17]</td>
<td>FALP</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>2.4 Ganguli [10]</td>
<td>SA</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>2.8 Amano [29]</td>
<td>IR</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>2.2 Rowe [14]</td>
<td>FALP</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>1.5 Smith [15]</td>
<td>FALP</td>
<td></td>
<td></td>
</tr>
<tr>
<td>( \text{CO}_2^- )</td>
<td>3.5 ± 0.5</td>
<td>3.8 Welter [6]</td>
<td>SA</td>
<td></td>
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<td></td>
<td>4.0 Gutcheck [30]</td>
<td>SA</td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>3.1 Geoghegan [16]</td>
<td>FALP</td>
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<tr>
<td>( \text{N}_2\text{O}^- )</td>
<td>3.6 ± 0.6</td>
<td></td>
<td></td>
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</tr>
<tr>
<td>( \text{HCO}_2^- )</td>
<td>4.6 ± 0.7</td>
<td>3.4 Adams [12]</td>
<td>FALP</td>
<td></td>
</tr>
<tr>
<td>( \text{HN}_2\text{O}^- )</td>
<td>4.5 ± 0.7</td>
<td>4.2 Adams [12]</td>
<td>FALP</td>
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<tr>
<td>( \text{H}_2\text{O}^- )</td>
<td>8.0 ± 1.5</td>
<td>10.0 Herd [18]</td>
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<tr>
<td></td>
<td>3.4 Mul [20]</td>
<td>MB</td>
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<td></td>
<td>13.0 Leu [9]</td>
<td>SA</td>
<td></td>
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<tr>
<td></td>
<td>20.0 Heppner [31]</td>
<td>IT</td>
<td></td>
<td></td>
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<tr>
<td>( \text{CH}_4^+ )</td>
<td>9.0 ± 2.0</td>
<td>11.0 Adams [17]</td>
<td>FALP</td>
<td></td>
</tr>
<tr>
<td></td>
<td>14.0 Smith [15]</td>
<td>FALP</td>
<td></td>
<td></td>
</tr>
<tr>
<td>( \text{C}_2\text{H}_4^+ )</td>
<td>9.0 ± 2.0</td>
<td>7.4 Adams [12]</td>
<td>FALP</td>
<td></td>
</tr>
</tbody>
</table>

Fig. 1. Values of $\alpha(z)$ (filled squares, right-hand ordinate) from measured $n_e(z)$ data (open circles, left-hand ordinate) for an O$_2^+$ plasma. The solid line through the $n_e(z)$ data represents a fourth-order polynomial fit. The solid line through the $\alpha(z)$ values represents the average, while the dashed lines indicate the one standard deviation interval.

Fig. 2. Values of $\alpha(z)$ (filled squares, right-hand ordinate) from measured $n_e(z)$ data (open circles, left-hand ordinate) for an CO$_2^+$ plasma. The solid line through the $n_e(z)$ data represents a fourth-order polynomial fit. The solid line through the $\alpha(z)$ values represents the average, while the dashed lines indicate the one standard deviation interval.

Fig. 3. Values of $\alpha(z)$ (filled squares, right-hand side axis) from measured $n_e(z)$ data (open circles, left-hand ordinate) for an HCO$_2^+$ plasma. The solid line through the $n_e(z)$ data represents a fourth-order polynomial fit. The solid line through the $\alpha(z)$ values represents the average, while the dashed lines indicate the one standard deviation interval.

and the deduced $\alpha(z)$ for O$_2^+$ ions. As was expected, $\alpha(z)$ remains constant over a factor of ten in electron density, supporting the conclusion that deionization occurs by binary recombination of a single ion species. Data of equal quality (see examples in Fig. 2 and Fig. 3) were obtained for most of the other ions in Table 1. As a rule, the $\alpha(z)$ values for all ions were constant within 10 to 15% from the mean. None of the ions exhibited the strongly anomalous decay that has been observed for H$_3^+$ ions [15,22].

For most ion species our recombination coefficients agree well with earlier data where such are available. The agreement among various determinations of $\alpha$(HCO$^+$) is poorer than for other ions. The lowest value was reported by Adams et al. [17] at $1.1 \times 10^{-7}$ cm$^3$/s and was revised later by Smith et al. [15] to $1.5 \times 10^{-7}$ cm$^3$/s, while the highest value ($2.8 \times 10^{-7}$ cm$^3$/s) was reported by Amano [29]. With the exception of these two cases the rest of the experiments yielded values near $2.2 \times 10^{-7}$ cm$^3$/s and our result $1.9 \times 10^{-7}$ cm$^3$/s is in good agreement with them.

The state of internal excitation of the various ions is generally not well known. Our data contain no hint that the recombining ion species consisted of mixtures of different vibrational states and that these recombined with significantly different recombination coefficients. If this were the case, the finding of a constant recombination coefficient would be rather
surprising. Since vibrationally excited ions are often quenched to lower states in collisions with molecular gases, the presence of excited ions might also lead to an apparent dependence of the recombination coefficient on reagent pressure. However, measurements at different reagent concentrations resulted in no more than 10 to 15% variation of the recombination coefficient indicating that either the ions were in their ground state or that there is little dependence of \( \alpha \) on vibrational state.

While our data agree well with many other determinations, it should be noted that most other determinations were made in plasmas of similar pressures and electron density. There is still lack of accurate absolute recombination coefficients (or cross sections) measured under essentially single-collision conditions, such as in merged-beam or ion storage rings.

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