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Aerodynamic Database
Development for Mars Smart
Lander Vehicle Configurations

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AERODYNAMIC DATABASE DEVELOPMENT FOR MARS SMART LANDER 
VEHICLE CONFIGURATIONS

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Abstract

An aerodynamic database has been generated for the Mars Smart Lander Shelf-All configuration using computational fluid dynamics (CFD) simulations. Three different CFD codes, USM3D and FELISA, based on unstructured grid technology and LAURA, an established and validated structured CFD code, were used. As part of this database development, the results for the Mars continuum were validated with experimental data and comparisons made where applicable. The validation of USM3D and LAURA with the Unitary experimental data, the use of intermediate LAURA check analyses, as well as the validation of FELISA with the Mach 6 CF₄ experimental data provided a higher confidence in the ability for CFD to provide aerodynamic data in order to determine the static trim characteristics for longitudinal stability. The analyses of the non-continuum regime showed the existence of multiple trim angles of attack that can be unstable or stable trim points. This information is needed to design guidance controller throughout the trajectory.

Nomenclature

\begin{tabular}{ll}
\textbf{A} & Reference area (m²) \\
\textbf{BC} & Ballistic coefficient (m /CD*A) \\
\textbf{CA} & Axial force coefficient \\
\textbf{CD} & Drag coefficient \\
\textbf{CG} & Center of gravity \\
\textbf{CN} & Normal force coefficient \\
\textbf{Cm} & Pitching moment coefficient \\
\textbf{Kn} & Knudsen number \\
\textbf{L/D} & Lift to drag ratio \\
\textbf{m} & Vehicle mass (kg) \\
\textbf{M} & Mach number \\
\textbf{\(\alpha\)} & Angle of attack \\
\textbf{\(\gamma\)} & Ratio of specific heats \\
\textbf{Cm\(_{\alpha}\)} & Static stability derivative \\
\end{tabular}

Subscripts:

\textit{trim} trim condition
Introduction

The exploration of Mars has been driven by the need to understand its planetary evolution and thereby recognize its past, present, or future potential to support life. The discovery of Martian life would require the ability to understand or characterize Mars from its geology, climate, biology, and other natural processes. This requirement ultimately defines a set of overall science objectives/payloads along with a probable set of planetary surface locations. In addition, the accuracy of targeting a specific surface location and achieving a successful precision landing are also essential requirements for fulfilling these science objectives.

The entry, descent, and landing (EDL) phases for a Mars mission can be categorized based on 1st, 2nd, and 3rd generation type systems. These categories describe the development of landing ellipses based on large scale (100s of km), precision (< 10 km), and pinpoint distance accuracies from the desired landing site. The 1st generation EDL system utilizes a ballistic entry. The 2nd generation EDL system (Smart) employs significant improvements in guidance and navigation methods enabling precision entry and utilizes a “Smart” or local surface hazard avoidance and hazard tolerance system in order to provide a safer landing environment. The 2nd generation EDL system also requires the development of a higher fidelity trajectory simulation. The objectives are for the vehicle to land within 10 km of the intended touchdown site and landing safely 99.7% of the time.

The trajectory simulation for the Mars Smart Lander (MSL) configuration series was performed using the 3-degree of freedom (DOF) version of the POST (Program to Optimize Simulated Trajectories). The POST program integrates the equations of motion and includes a set of specific constraints necessary to define the overall planetary, mission, and vehicle parameters. These models include specific science-related information of the planet Mars such as planet definition, gravity, and atmosphere while other models provide specific information related to the vehicle configuration and the vehicle characteristics of motion. Examples of these models include control system emulation, guidance algorithms, navigation system, mass properties, and aerodynamics. The aerodynamics model is a database that provides aerodynamic force and moment data of a specific vehicle configuration for both the entry and descent segments of the trajectory. The entry phase is defined as the region between the atmospheric interface and the supersonic parachute deployment point while the descent phase is defined from that point to landing.

The development of an entry aerodynamic database requires specific knowledge of the geometry and orientation of the vehicle in addition to the freestream quantities along a given trajectory. In accordance with the Mars Smart Lander payload and surface location requirements, a set of preliminary configurations were developed that would have favorable trim characteristics along proposed nominal and dispersed trajectories. These vehicle configurations (Figure 1) are the axisymmetric Baseline, Canted-All, Shelf-All, and a derivative of the Shelf-All tested in the Ames Ballistic range facility. All configurations have the same 70° forebody cone shape with a bi-conic backshell. The Shelf-All model was designed with a blended control tab that is tangent to the forebody surface while the Shelf-2 model has a non-blended control tab while also being tangent to the forebody surface. The Canted-All model has a similar Shelf-All control tab but is canted 10° towards the vehicle nose.

The objectives of this paper are to provide information on the development of an entry aerodynamic database, to provide specific validation of the CFD analysis programs with experimental data, and show the resulting flight trim characteristics for a Mars Smart Lander configuration. The subsequent sections of this paper will describe the analysis codes used, CFD validation, and the creation and results of a flight database.

Code Descriptions

USM3Dns

USM3Dns is a three dimensional, tetrahedral, cell-centered, finite volume Euler and Navier-Stokes flow solver for unstructured meshes. Inviscid flux quantities are computed across each cell face using Roe’s flux-difference splitting (FDS). Spatial discreti-
...zation is accomplished using an analytical reconstruction process for computing solution gradients within tetrahedral cells. The solution is advanced in time to a steady state condition by an implicit backward Euler time-stepping scheme. Flow turbulence is modeled by the Spalart-Allmaras one-equation model, which is optionally coupled with a wall-function formulation to reduce solution stiffness and the number of cells in the sub layer of the boundary layer. USM3Dns runs on massively parallel computers and clusters of personal computers (PC’s). Although a single processor version is available for a variety of computing platforms, the parallel version is the code of choice because it enables rapid turnaround for large problems.

Although USM3Dns is widely in use by aerospace practitioners for several years, the latest Navier-Stokes version was never validated for the Mach number range and shape (especially the blunt fore body) of interest in the present study. In addition, the available version was so far used with perfect gas assumptions, while the aerodynamic database required solutions at Mars atmospheric conditions involving a different gas constant. For these reasons, the flow solver had to be modified before application. This section briefly describes the changes made to the code. The following changes were incorporated to the existing parallel version of USM3Dns:

1. A new flow initialization procedure was introduced whereby for the cases where the free-stream Mach number exceeded 1.2; the solution was initialized to a lower, usually subsonic, value. This strategy helped the flow solution convergence to steady state, especially in the aft, low-Mach number region of the configurations.
2. A special flux-splitting scheme based on the AUSM$^6$ (Advection Upstream Splitting Method) scheme was incorporated to overcome the well-known “Carbuncle phenomenon” usually associated with the application of the Roe scheme to a case with strong shock waves.
3. In order to simulate the Mars atmospheric conditions, the perfect gas assumptions used thus far was modified to allow specification of an effective $\gamma$ for each freestream Mach number, along with coefficients for Sutherland’s for viscosity.

LAURA

The Langley Aerothermodynamic Upwind Relaxation Algorithm (LAURA) is a three-dimensional Na-

vier-Stokes solver developed by Gnoffo for structured grids. LAURA is a finite volume formulation of the integral form of the Navier-Stokes equations. A second order-accurate, symmetric total-variation-diminishing (TVD) scheme is used in conjunction with upwind differencing of the discretized equations. At each cell face, Roe’s averaging defines the flowfield variables based on values from the adjacent cells. The unsteady governing equations are driven to a steady state solution through an implicit time relaxation procedure. During the relaxation process, the grid is periodically adapted in the body-normal direction so that the grid can be tailored to the emerging solution. This elliptic flow solver has the capability of solving chemical non-equilibrium flow through the specification of chemical species and reactions unique to a specific atmosphere as well as the ability to model perfect gas and equilibrium air.

FELISA

The FELISA (Finite Element Langley Imperial College Swansea Ames) software system consists of unstructured surface and volume grid generation with an inviscid flow solver and unstructured post processing utilities. The flow algorithm was developed through the application of the Galerkin finite element method in space to obtain a coupled set of ordinary differential equations in time. The steady state solution of this equation set is achieved by advancing the system using an explicit Runge-Kutta type marching scheme. The hypersonic flow solver has options for perfect gas, equilibrium air, CF$\text{\textsubscript{4}}$, CO$\text{\textsubscript{2}}$, and Mars equilibrium gases. In addition, the hypersonic flow solver has the capability of solving chemical non-equilibrium flow and real gas (chemical and thermal non-equilibrium) flow.

Aerodynamic Database

The aerodynamic database is provided as a FORTRAN routine containing specific vehicle force and moment data as a function of vehicle orientation relative to the flow and flowfield parameters to the POST trajectory simulation. Within this routine, the vehicle trajectory is partitioned into three atmospheric regimes that are defined based on the similarity parameter of Knudsen number. The Knudsen number (Kn) is defined as the mean free path (distance between molecular collisions) divided by the vehicle reference length (typically the aeroshell diameter). These atmospheric regimes are depicted in Figure 2.
The selections and application of these analysis programs were based on several considerations. The first consideration was to develop a methodology that could allow for rapid turnaround of assessing trends for proposed vehicle configurations by taking advantage of grid generation methods and flowfield modeling. Another consideration was the flight regime capability of each program. The final consideration was to ensure confidence and continuity within the solution matrix by using the LAURA code to validate intermediate analysis points with the USM3D and FELISA analysis programs since LAURA has been previously validated with previous Mars missions.18

Table 1. Proposed Flight Analysis Cases for MSL Reference Trajectory.

<table>
<thead>
<tr>
<th>Analysis Program</th>
<th>M</th>
<th>V (m/s)</th>
</tr>
</thead>
<tbody>
<tr>
<td>USM3D</td>
<td>1.6</td>
<td>361.37</td>
</tr>
<tr>
<td></td>
<td>2.3</td>
<td>510.38</td>
</tr>
<tr>
<td></td>
<td>3.5</td>
<td>755.52</td>
</tr>
<tr>
<td></td>
<td>4.5</td>
<td>957.06</td>
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<tr>
<td></td>
<td>6</td>
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<td></td>
<td>8</td>
<td>1677.90</td>
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<td></td>
<td>10</td>
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<td></td>
<td>24.5</td>
<td>4763.74</td>
</tr>
<tr>
<td></td>
<td>28</td>
<td>5295.34</td>
</tr>
</tbody>
</table>

Experimental Data

Two experimental facilities were used to acquire aerodynamic data for CFD validation of the aforementioned analysis programs. The first facility, the Langley Unitary Plan Wind Tunnel, is a closed-circuit, continuous flow, variable density supersonic tunnel that was used to obtain 6-DOF static aerodynamic data for 4 proposed Mars Lander configurations over a Mach number range of 2.3 to 4.5. The complete details of the models, facility calibration, instrumentation, and data acquisition are presented in Reference 17. The second
facility, the 20-inch Mach 6 CF₄ Tunnel, was used to obtain aerodynamic data for the Mars Surveyor configuration. The details of this work are presented in Reference 19.

Results and Discussion

USM3Dns Validation

The USM3D code was required to validate its results with Unitary tunnel experimental data in order to provide confidence in correcting the Unitary Wind tunnel data for the geometric differences between wind tunnel and flight configuration and gas composition of the Mars atmosphere. Along with the previously described changes to USM3D, an unstructured mesh was constructed for the Baseline and Shelf-All tunnel configurations for a series of Mach numbers between 2.3 and 4.5, each for a range of angles-of-attack values. The grid for the full Navier-Stokes solution varied in size typically from 1.4 million tetrahedral cells for the wind-tunnel cases to 2.6 million cells for the flight configurations. All the calculations were run for a Reynolds number of 1.0 million based on the maximum diameter of the configuration (6 inches for the wind tunnel models and 4.05 meters for the flight cases). The CFD cases were run using up to 60 processors on an SGI Origin 2000 computer. For all the computations reported here, a full viscous formulation was used. Some examples of the unstructured surface topology are depicted in Figures 3 and 4.

Figure 3. Unstructured Surface Grid for the Unitary Baseline Model Configuration.

Figure 4. Unstructured Surface Grid for the Unitary Shelf-All Model Configuration.

Figure 5. USM3D/Unitary Data Comparisons for the Baseline Configuration.

However, the USM3D analysis does predict a change in normal force slope near zero degrees angle of attack.
which is a known trend for axisymmetric blunt body configurations at low supersonic speeds.

Figure 6 shows the comparisons for the Shelf-All configuration to the Unitary experimental data at freestream Mach numbers of 2.3, 3.5 and 4.5. As with the previous comparisons shown for the Baseline configuration, the overall agreement for Shelf-All configuration is very good. These comparisons show better agreement in axial and normal force at lower Mach numbers than shown for the Baseline model while the pitching moment and L/D distribution shows excellent agreement throughout the Mach number and angle of attack range.

Figure 7. USM3D/Unitary Data Sideslip Comparisons for the Shelf-All Configuration (β=4°).

Figure 8 shows the comparisons for the Canted-All configuration to the Unitary experimental data at freestream Mach numbers of 2.3 and 3.5. As with the previous comparisons shown for the Baseline configuration, the overall agreement for Canted-All configuration is very good. These comparisons show better agreement throughout the Mach number and angle of attack range for all force, moment and L/D quantities than those shown for the Baseline configuration. However, additional solutions at lower angles of attack would have to be made in order to obtain a more complete assessment of CFD validation.
LAURA Validation

Although the LAURA analysis program has been previously validated with Mars Pathfinder hypersonic flight data, an attempt was made to provide additional validation with the Unitary experimental data for a single Mars Smart Lander configuration at supersonic speeds and to further validate USM3D. Figure 9 shows the LAURA structured grid topology for the Unitary Baseline configuration. This structured grid is comprised of 16 computational blocks with an approximate total of 450,000 cell centers. In addition, this structured grid was adapted in the direction perpendicular to the body surface as a function of the flowfield temperature and grid distribution values.

Figure 9. Structured Grid for the Unitary Baseline Model Configuration

Figure 10 depicts the LAURA Mach number contours for the Baseline Unitary tunnel configuration for freestream Mach number of 4.5 and an angle of attack of 10 degrees with no sideslip angle. This solution imposed perfect gas air ($\gamma=1.4$) at $M=4.5$ tunnel conditions with a laminar, no-slip adiabatic wall boundary condition. Figure 10 shows along the vehicle symmetry and outflow planes the general flowfield characteristics of an axisymmetric blunt body with dominant subsonic regions within the shock layer and vehicle base.

Figure 10. Mach Number Contours for the Unitary Baseline Configuration – (M=4.5, $\alpha=10^\circ$)

Figure 11 depicts the LAURA/Unitary comparisons for the Baseline Unitary tunnel configuration for freestream Mach numbers of 2.3 and 4.5. These comparisons overall show good agreement to the Unitary data for all quantities. The axial force comparisons show better agreement at the higher angles of attack for both Mach numbers while the normal force comparison shows better agreement over the entire angle of attack range at Mach 4.5 while the 10 degree case for Mach 2.3 shows a larger difference as compared to the USM3D/Unitary comparisons in Figure 5.

Figure 11. LAURA/Unitary Data Comparisons for the Baseline Configuration.
FELISA Validation

The FELISA validation work was centered about an earlier Mars 2001 Surveyor configuration that was tested in the 20-inch Mach 6 CF₄ Tunnel. The Mars Surveyor configuration (Figure 12) was similar to Mars Smart Lander configuration, but differed in the control flap geometry and location. The Mars Surveyor flap surface area was at least 50 percent smaller and the flap orientation was 90 degrees perpendicular from the axis of rotation and located at aft of the maximum diameter. The number assigned to each control flap indicated the area ratio of the flap where the larger number has a greater area ratio.

![Figure 12. Mars 2001 Surveyor Wind Tunnel Model](image)

Figures 13 and 14 depict the pitching moment coefficient and L/D ratio for the Mars Surveyor configuration for Flap 1. The FELISA CF₄ results showed very good agreement to the experimental CF₄ data comparisons at 0, 5, 11, and 16 degrees angle of attack. In conjunction with a planned test in the 20-inch Mach 6 facility, a FELISA analysis was performed using a CO₂ test gas as a freestream condition. The FELISA results predicted that the force and moment contributions to the L/D ratio at 0 degrees angle of attack would be insensitive to the CO₂ test gas. However, the need for additional results at different angles of attack would be necessary in order to determine a more generalized trend in CO₂ insensitivity.

![Figure 13. FELISA/CF₄ Pitching Moment Comparisons for the Mars Surveyor Configuration.](image)

Figures 15 and 16 depict the pitching moment coefficient and L/D ratio for the Mars Surveyor configuration with Flap 3. The FELISA CF₄ results shows very good agreement to the experimental CF₄ data comparisons at 0, 5, 11, and 16 degrees angle of attack. The larger Flap 3 configuration does show a slightly larger deviation in pitching moment at 0 and 5 degrees as compared to the Flap1, but overall agreement is still very good. These differences could be attributed to viscous contributions (not being modeled in FELISA) induced by the larger flap or an indication of a required increase in control flap grid resolution.

![Figure 14. FELISA/CF₄ L/D Comparisons for the Mars Surveyor Configuration.](image)
Flight Results – (Shelf-All Continuum Database)

For the continuum regime, an aerodynamic database was created for the Mars Smart Lander “Shelf-All” configuration using the USM3Dns correction analyses for the Unitary experimental data (M=2.3-4.5), the LAURA intermediate analysis check points at Mach 4.5 and 10, and the FELISA analyses at Mach 6, 8, 10, 24, and 28. Figures 17 and 18 show the continuum trim characteristics of the Shelf-All configuration based on the axial CG locations of (x/Diameter) of 0.21 and 0.32. The CG of 0.21 corresponds to an early estimate of the MSL CG location while the 0.32 value corresponds to an L/D value of approximately 0.24 at a velocity of 4700 m/s (Mach 24). In addition, the radial CG offset for this configuration was designed to be zero. For the determination of ballistic coefficient, a mass of 2200 kg. was used.

At the axial CG of 0.21, the trim angle of attack depicts oscillatory behavior of 1 degree within the Mach 2.3 to 4.5 Unitary experimental data range before reaching a maximum value of 15.4 degrees at a velocity of approximately 1300 m/s (Mach 6). Beyond Mach 6, the trim angle of attack at the forward CG decreases nearly linear to a continuum limit value of approximately 12.8 degrees. Likewise, the L/D values depict the same low supersonic behavior as well as the nearly linear behavior beyond the maximum L/D of 0.266 at Mach 4.5.

At the axial CG of 0.32, the trim angle of attack depicts a larger variation in trim angle of attack of about 2.5 degrees within the Unitary data range and also reaching a maximum of 16.9 degrees at a velocity of approximately 550 m/s (Mach 2.7) while reaching another lower peak of 16.1 degrees at Mach 6. Beyond Mach 6, the trim angle of attack at the 0.32 CG decreases nearly linear to a continuum limit value of approximately 13.6 degrees. As with the forward CG results, the L/D shows a corresponding maximum value at the peak trim reaching an overall maximum of 0.298.
location due to only a small (1%) difference in ballistic coefficient over the entire continuum regime.

![Figure 18. Ballistic Coefficient Continuum Trim Characteristics for Shelf-All Configuration.](image)

### Flight Results—(Shelf-All Non-Continuum Database)

For the rarefied regime, an existing set of free molecular data 2 over an angle of attack range of -180 to 180 degrees was used in conjunction with bridging function 20 to provide aerodynamic values within the transition atmospheric regime. Figures 19 and 20 show the trim characteristics for the Shelf-All configuration from the edge of continuum to a value of 20. A Knudsen number (Kn) of 20 was sufficient to cover the range up to the atmospheric interface at approximately 125000 meters above ground level.

Figure 19 shows the oscillatory behavior of the pitching moment within the transition regime. At the continuum limit and up to a Knudsen number of approximately 5, there exists only a single trim point that is also stable (negative Cm_). As the vehicle descends through the transition regime, there are multiple conditions that the vehicle could potentially trim, but could potentially be an unstable or stable trim point. Figure 20 displays Knudsen number and Cm alpha as a function of trim angle of attack (αtrim). For Knudsen numbers above 2, there exist three trim angles of which 2 are unstable and 1 is stable. This information is needed while designing a guidance controller during the entry phase of the trajectory.

![Figure 19. Non-Continuum Pitching Moment for Shelf-All Configuration.](image)

### Concluding Remarks

The CFD validation of USM3Dns and LAURA with the Unitary experimental data, the use of intermediate LAURA check analyses, as well as the validation of FELISA with the Mach 6 CF₄ experimental data provided very good agreement and a higher confidence in each CFD code’s aerodynamic analysis results including the longitudinal static trim characteristics for the Mars continuum atmosphere.
Acknowledgments

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References


