Longwave Radiative Forcing of Saharan Dust Aerosols Estimated from MODIS, MISR and CERES Observations on Terra

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Abstract

Using observations from the Multi-angle Imaging Spectroradiometer (MISR), the Moderate Resolution Imaging Spectroradiometer (MODIS), and the Clouds and the Earth's Radiant Energy System (CERES) instruments onboard the Terra satellite; we present a new technique for studying longwave (LW) radiative forcing of dust aerosols over the Saharan desert for cloud-free conditions. The monthly-mean LW forcing for September 2000 is 7 Wm\(^{-2}\) and the LW forcing efficiency\(^1\) (LW\(_{\text{eff}}\)) is 15 Wm\(^{-2}\). Using radiative transfer calculations, we also show that the vertical distribution of aerosols and water vapor are critical to the understanding of dust aerosol forcing. Using well calibrated, spatially and temporally collocated data sets, we have combined the strengths of three sensors from the same satellite to quantify the LW radiative forcing, and show that dust aerosols have a "warming" effect over the Saharan desert that will counteract the shortwave "cooling effect" of aerosols.

\[^1\] LW\(_{\text{eff}}\) = Longwave forcing per unit aerosol optical thickness at 0.55 \(\mu\)m
1. Introduction

Desert dust is considered to be one of the major sources of tropospheric aerosol loading, and play an important role in climate forcing studies [Kaufman et al., 2002; Christopher and Zhang, 2002]. Widely prevalent in the tropics [Prospero, 1999], dust aerosols are effective in reflecting solar energy back to space thereby "cooling" the earth's surface [Tegen et al., 1996; Myhre et al., 2003]. Besides their radiative impact in the shortwave portion of the electro-magnetic spectrum, dust aerosols also have an important radiative effect in the longwave (LW) [Sokolik et al., 1998; Liao and Seinfeld, 1998]. Having mean particle sizes on the order of several micrometers [Kaufman et al., 2001], dust aerosols can effectively reduce the earth's LW emission by re-emitting at a colder temperature when compared to the surface and thereby "warming" the earth [Sokolik et al., 1998; Tegen et al., 1996]. However, numerical simulations of dust LW forcing remain a challenging task because of the high spatio-temporal variation of dust properties [Myhre et al., 2003].

Using satellite observations, several studies have attempted to estimate the LW radiative forcing of dust aerosols. For example, using broadband measurements from the Earth Radiation Budget Experiment (ERBE), the changes in LW and shortwave (SW) radiation at the top of atmosphere (TOA) with and without dust aerosols has been reported [e.g. Ackerman and Chung, 1992; Hsu et al., 2000]. Since the ERBE instruments were not designed to detect aerosols, narrowband satellite observations from the Advanced Very High Resolution Radiometer (AVHRR) [Ackerman and Chung, 1992] and the Total Ozone Mapping Spectrometer (TOMS) [Hsu et al., 2000] were used to identify the presence of dust plumes within the large ERBE scanner footprints. Although the AVHRR and the ERBE scanner were on the same satellite (e.g. NOAA-9), accurate detection of dust aerosols over bright targets such as the Saharan desert was not possible due to the limitations of the AVHRR. The TOMS aerosol index which is the difference in aerosol absorptivity between two UV channels has been used to detect dust
aerosols over bright targets such as deserts, but the technique is insensitive to dust layers at low altitudes [Herman et al., 1997] and since the TOMS and ERBE instruments were not on the same satellite, precise temporal collocation was not possible.

With the launch of NASA's Terra satellite, well-calibrated measurements from the same satellite are now available for examining the role of aerosols on climate. For example, using MODIS and CERES data, aerosol radiative forcing over the global oceans has been completed [Christopher and Zhang, 2002]. One of the primary roles of the Moderate Resolution Imaging Spectroradiometer (MODIS) and the Multi-Angle Imaging Spectroradiometer (MISR) is to detect aerosols and retrieve their properties. The MODIS with 36 channels has improved spectral, spatial and radiometric resolutions when compared with previous imagers [King et al., 1992] and has been used to retrieve aerosol properties over the global oceans [Remer et al., 2002] and over land areas with low surface reflectances [Chu et al., 2002]. The MISR with multi-angle measurements and four spectral channels has also been used to identify aerosols and retrieve aerosol properties over the global oceans and land including bright targets such as deserts [Kahn et al., 2001]. In this study, we combine the strengths of three instruments on the Terra satellite to examine cloud-free longwave dust aerosol forcing. We use the MODIS to identify clouds because of its multi-spectral capabilities and higher spatial resolution when compared to MISR and CERES. Although the MODIS provides aerosol properties over the global oceans, aerosol retrievals over highly-reflective surfaces such as deserts are not available. Therefore we use the multi-angle capabilities of the MISR instrument to obtain aerosol optical depth. Finally we use the CERES to obtain broadband TOA longwave fluxes.
2. Data and Methods

Four data sets are used in this study; the MISR Level 2 (MIL2ASAE) 17.6 km- daily aerosol product [Diner et al., 2001]; the 5 km- MODIS Level 2 daily (MOD06) cloud product [Ackerman et al., 1998] and the 30 km (nadir view) pixel level CERES data [Wielicki et al. 1996]. The MISR aerosol product contains aerosol optical thickness and other related properties at 0.558μm. The expected uncertainties in MISR aerosol retrievals are ±0.05 for aerosol optical thickness (AOT) less than 0.5, and ±10% for AOT greater than 0.5 [Martonchik et al., 1998]. Recently, Diner et al., [2001] showed that there was excellent agreement between MISR retrievals and ground based sun-photometer measurements over southern Africa during August and September 2000. The CERES scanner on Terra observes SW and LW radiances that are converted to fluxes using angular distribution models, which take into account the bi-directional characteristics of a reflecting surface [Wielicki et al., 1996]. Spatial collocation is performed using point spread functions of the CERES scanner [Christopher and Zhang, 2002].

The area of study is the Saharan desert (15°- 40°N, 20°W- 40°E). The CERES pixels that are labeled as “clear desert” by the CERES data were first selected. However, due to the large footprint of the CERES scanner, these pixels could still have some cloud contamination [Christopher and Zhang, 2002]. To eliminate these cloud effects, the MOD06 cloud product is collocated with the CERES data and only those CERES pixels that are at least 99.9% cloud free as identified by the MOD06 data were used. For these cloud-free CERES pixels, the best-fit aerosol optical thickness (τ_{0.55}) from the MISR data is obtained. As noted in Christopher and Zhang [2002], stringent cloud rejection criteria such as this could eliminate some aerosol plumes that have high optical thickness.
3. Results

3.1 Satellite Observations

Figure 1a shows the monthly-mean MISR $\tau_{0.55}$ for cloud-free CERES pixels. Data is gridded every $1^\circ \times 1^\circ$ and smoothed for display purposes. The major dust plumes are clustered in three regions; within Mali and Mauritania, Chad and Niger, and Sudan. Areas shown in white are due to missing data, data gaps and also due to cloud-screening criteria\(^2\). Superimposed on Figure 1a are six selected regions with optically thick aerosol plumes. The latitude and longitude boundaries of the six regions are shown in Table 1. For reference purposes, we also compared the spatial distribution of dust aerosols from MISR with those from the TOMS [Herman et al., 1997]. Figure 1b shows the monthly-mean Earth Probe TOMS aerosol index (version 8, level 2) for the same study period. Pixels with reflectivity greater than 20\% in the TOMS 360 nm channel are excluded to avoid possible cloud contamination. The spatial distribution of dust aerosols derived from these two independent instruments are consistent over most of the study regions. However, over region 4, thick dust plumes shown in figure 1a are not seen in figure 1b. Since the TOMS aerosol index is insensitive to dust plumes at low altitudes [Herman et al., 1997], it is possible that these aerosols are not detected by TOMS although we have no ancillary information to verify this.

Figure 1c shows the MISR optical thickness versus the CERES LW flux for the six selected regions. The CERES fluxes are binned for every 0.1 MISR optical thickness intervals. The thick black line in figure 1 shows the mean value for all regions. With the exception of region 4, as MISR AOT increases the CERES LW flux decreases because dust aerosols emit at a colder temperature when compared to the surface. The reason for the weak relationship between MISR $\tau_{0.55}$ and CERES LW flux in region 4 is due to the combined effect of dust aerosol vertical distribution, column water vapor, and surface

\(^2\) Data from September 8, 17, 24, 27, and 30 was not used due to these criteria.
characteristics (see section 3.2). We use this relationship between MISR $\tau_{0.55}$ and CERES LW flux and calculate dust aerosol forcing efficiency for the six selected regions (Table 1). The dust aerosol forcing efficiency ($LW_{eff}$) is defined as the LW flux perturbation due to dust aerosols per unit optical thickness, and is derived from the difference in CERES LW flux for MISR $\tau_{0.55}$ values between 0 and 1. The largest $LW_{eff}$ of 20 Wm$^{-2}$ is found over region 2, and the lowest $LW_{eff}$ of near 0 Wm$^{-2}$ is found over region 4. The other four regions have values ranging from 11 to 19 Wm$^{-2}$. The variability of $LW_{eff}$ is due to differences in surface and atmospheric conditions (see section 3.2). Averaged over all pixels, the regional mean dust aerosol forcing efficiency is 15 Wm$^{-2}$.

To examine the spatial distribution of the aerosol radiative effect, we define the TOA dust LW forcing as $F_{clr} - F_a$ [Christopher and Zhang, 2002], where $F_{clr}$ is the TOA LW flux observed in cloud and aerosol free conditions, and $F_a$ is the TOA LW flux observed over dust regions. The $F_{clr}$ is obtained by averaging the CERES cloud and aerosol free pixels with MISR $\tau_{0.55} < 0.1$. Due to the persistence of dust aerosols, only one $F_{clr}$ value is used for each one of the six selected regions. However for the rest of the study area, $F_{clr}$ is obtained by averaging the LW fluxes for each 3°×3° grid and $F_a$ values are reported for every 1°×1° grid. Figure 1d shows the dust aerosol LW forcing over the Saharan desert. In general, as AOT increases the LW forcing also increases. However, over region 4, 5, and 6, although dust aerosol loading is high, the aerosol induced LW perturbations are low. The averaged dust aerosol LW direct forcing for the six regions range from 0 to 14 Wm$^{-2}$ (Table 1), and the averaged dust aerosol LW direct forcing over the entire study area is 7 Wm$^{-2}$. It is interesting to note that over region 4, where MISR observed optically thick aerosols that was not observed by TOMS, the averaged dust LW direct forcing is near 0. It is possible that the dust plume in this region is at a very low altitude and emits at a similar temperature as the surface.
3.2 Sensitivity studies

Although our analysis clearly shows a longwave "warming" effect, the differences in LW flux as a function of $\tau_{0.55}$ among the six regions are due to both atmospheric and surface properties including surface temperature ($T_s$), and vertical distribution of water vapor and aerosols [Hsu et al., 2000; Ackerman and Chung, 1992]. In this section, we tested the sensitivity of $LW_{eff}$ to changes in $T_s$, column water vapor, and aerosol vertical distribution over cloud-free desert conditions. A four-stream radiative transfer model [Fu and Liou, 1993] was used to compute TOA LW fluxes for various atmospheric and surface conditions with desert aerosol properties as described by Hess et al., [1998]. The vertical distribution of dust aerosol extinction is assumed to be exponential, and the dust extinction scale height is used to represent the vertical distribution. For the atmospheric temperature and water vapor profiles, we use rawinsonde measurements [27.20°N, 2.47°E] from September 1 at 12 UTC (hereafter referred to as profile 1). The $T_s$ is 316K and the column water vapor is 1.6 cm. To test the sensitivity of $T_s$, we used another case that has the same water vapor distribution as profile 1, but with different temperature profiles from the surface to 700 mb. The $T_s$ of the new profile (hereafter referred as profile 2) is 300K. We further varied the water vapor amount from 1.6 to 3.2 cm and aerosol extinction scale height from 1 to 2 km.

Table 2 lists the $LW_{eff}$ as functions of aerosol extinction scale height, $T_s$ and column water vapor. The profiles 1 and 2 are identified by their surface temperatures of 316K and 300K respectively. When the dust height is changed from 1 to 2 km while keeping the surface temperature (316K) and the water vapor loading the same (1.6 cm), the $LW_{eff}$ increases from 11.4 to 18.7 Wm$^{-2}$, because the dust layer now emits at a colder temperature and consequently, water vapor absorption above the dust layer is reduced. Increasing the water vapor loading from 1.6 to 3.2 cm for the same temperature (316K) and aerosol extinction height (1km) reduces the $LW_{eff}$ from 11.4 to 8.0 Wm$^{-2}$ because the absorption due to
water vapor above the dust layer has increased. Increasing the surface temperature from 300 to 316 K for the same water vapor loading (1.6cm) and aerosol scale height (1km) will increase the LW_{eff} from 6.3 to 11.4 Wm^{-2} because the cloud and aerosol free emission temperature increases. We also note that the TOA longwave flux is very sensitive to column water vapor. A one centimeter increase in column water vapor has a similar effect as a unit aerosol optical depth increase in reducing the TOA LW flux [Hsu et al., 2000]. Therefore, the uncertainties in column water vapor could induce uncertainties in the derived F_{clr}. In summary, LW_{eff} is proportional to T_s and aerosol extinction scale height, and inversely proportional to column water vapor. All of these parameters are critical to the understanding of LW dust direct forcing, and precisely collocated observations of the three parameters from other sources are needed for future dust forcing studies that utilize broadband satellite observations.

**4. Summary**

In this paper, we have demonstrated a new technique for estimating the dust aerosol LW direct forcing over the Saharan desert. Through the synergistic use of MODIS, MISR and CERES instruments onboard the same satellite, our study clearly shows that there is a strong dust warming effect of 7 Wm^{-2} over the cloud free Sahara desert regions for September 2000 that can offset the shortwave cooling effect of aerosols. The regional mean dust LW direct forcing efficiency is 15 Wm^{-2} per τ_{0.55}. However, since water vapor dominates the radiative processes in the longwave part of the electromagnetic spectrum, simultaneous measurements of both aerosol height and water vapor are needed to reduce the uncertainties in the longwave radiative forcing calculations of dust aerosols.
Acknowledgements

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References


**Figure Captions**

**Figure 1a.** Spatial distribution of MISR optical thickness at 0.55µm over the area of study for September 2000. Missing data is shown in white and six selected regions (see Table 1) are shown in red boxes.

**Figure 1b.** Spatial distribution of the TOMS aerosol index for September 2000.

**Figure 1c.** MISR $\tau_{0.55\mu m}$ vs. CERES LW fluxes for six selected regions. Solid black line shows the mean values for the whole region.

**Figure 1d.** Spatial distribution of CERES-derived LW direct dust forcing over cloud free desert regions. Positive values denote a "warming" effect.
Table 1. Summary of MISR optical thickness and CERES fluxes for the six selected regions.

<table>
<thead>
<tr>
<th>Region</th>
<th>Latitude degrees</th>
<th>Longitude Degrees</th>
<th>Dust LW forcing Wm$^{-2}$</th>
<th>Dust LW direct forcing efficiency* Wm$^{-2}$ per AOT</th>
<th>AOT (0.55µm)</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>16N - 28N</td>
<td>16W - 4E</td>
<td>14</td>
<td>19</td>
<td>0.69</td>
</tr>
<tr>
<td>2</td>
<td>15N - 20N</td>
<td>5E - 22E</td>
<td>12</td>
<td>20</td>
<td>0.75</td>
</tr>
<tr>
<td>3</td>
<td>15N - 22N</td>
<td>22E - 36E</td>
<td>9</td>
<td>19</td>
<td>0.66</td>
</tr>
<tr>
<td>4</td>
<td>23N - 32N</td>
<td>25E - 35E</td>
<td>0</td>
<td>0</td>
<td>0.52</td>
</tr>
<tr>
<td>5</td>
<td>27N - 33N</td>
<td>15E - 25E</td>
<td>4</td>
<td>11</td>
<td>0.52</td>
</tr>
<tr>
<td>6</td>
<td>23N - 27N</td>
<td>15E - 25E</td>
<td>3</td>
<td>14</td>
<td>0.44</td>
</tr>
<tr>
<td>Whole area</td>
<td>15N - 40N</td>
<td>20W - 40E</td>
<td>7</td>
<td>15</td>
<td>0.55</td>
</tr>
</tbody>
</table>

* Different methods are used in deriving dust LW forcing and LW forcing efficiency (see text). Therefore, the averaged dust LW forcing divided by the averaged AOT may not exactly match the dust LW direct forcing efficiency.

Table 2. Dust LW direct forcing efficiency (LW$_{eff}$)* as functions of surface temperature, aerosol extinction scale height and column water vapor content from radiative transfer calculations.

<table>
<thead>
<tr>
<th>Surface temperature K</th>
<th>Aerosol extinction scale height km.</th>
<th>Column Water vapor cm.</th>
<th>LW$_{eff}$ Wm$^{-2}$ per AOT</th>
</tr>
</thead>
<tbody>
<tr>
<td>316</td>
<td>1</td>
<td>1.6</td>
<td>11.4</td>
</tr>
<tr>
<td>316</td>
<td>1</td>
<td>3.2</td>
<td>8.0</td>
</tr>
<tr>
<td>316</td>
<td>2</td>
<td>1.6</td>
<td>18.7</td>
</tr>
<tr>
<td>316</td>
<td>2</td>
<td>3.2</td>
<td>14.2</td>
</tr>
<tr>
<td>300</td>
<td>1</td>
<td>1.6</td>
<td>6.3</td>
</tr>
<tr>
<td>300</td>
<td>1</td>
<td>3.2</td>
<td>4.2</td>
</tr>
<tr>
<td>300</td>
<td>2</td>
<td>1.6</td>
<td>11.6</td>
</tr>
<tr>
<td>300</td>
<td>2</td>
<td>3.2</td>
<td>9.0</td>
</tr>
</tbody>
</table>

* Defined as the difference in TOA LW fluxes for AOT values between 0 and 1.
Dust radiative forcing from Terra

(a) Map showing MISR AOT (0.55μm) with regions labeled.

(b) TOMS AI map with scale bars.

(c) CERES LW flux (Wm⁻²) vs MISR AOT (0.55 μm) graph with regions 1-6 and all regions.

(d) Dust LW direct Forcing (Wm⁻³) map with scale bars.