Flame Radiation, Structure, and Scalar Properties in Microgravity Laminar Fires

Douglas Feikema
Glenn Research Center, Cleveland, Ohio

Jongmook Lim and Yudaya Sivathanu
En’Urga, Inc., West Lafayette, Indiana
Since its founding, NASA has been dedicated to the advancement of aeronautics and space science. The NASA Scientific and Technical Information (STI) program plays a key part in helping NASA maintain this important role.

The NASA STI Program operates under the auspices of the Agency Chief Information Officer. It collects, organizes, provides for archiving, and disseminates NASA's STI. The NASA STI program provides access to the NASA Aeronautics and Space Database and its public interface, the NASA Technical Reports Server, thus providing one of the largest collections of aeronautical and space science STI in the world. Results are published in both non-NASA channels and by NASA in the NASA STI Report Series, which includes the following report types:

- **TECHNICAL PUBLICATION.** Reports of completed research or a major significant phase of research that present the results of NASA programs and include extensive data or theoretical analysis. Includes compilations of significant scientific and technical data and information deemed to be of continuing reference value. NASA counterpart of peer-reviewed formal professional papers but has less stringent limitations on manuscript length and extent of graphic presentations.

- **TECHNICAL MEMORANDUM.** Scientific and technical findings that are preliminary or of specialized interest, e.g., quick release reports, working papers, and bibliographies that contain minimal annotation. Does not contain extensive analysis.

- **CONTRACTOR REPORT.** Scientific and technical findings by NASA-sponsored contractors and grantees.

- **CONFERENCE PUBLICATION.** Collected papers from scientific and technical conferences, symposia, seminars, or other meetings sponsored or cosponsored by NASA.

- **SPECIAL PUBLICATION.** Scientific, technical, or historical information from NASA programs, projects, and missions, often concerned with subjects having substantial public interest.

- **TECHNICAL TRANSLATION.** English-language translations of foreign scientific and technical material pertinent to NASA’s mission.

Specialized services also include creating custom thesauri, building customized databases, organizing and publishing research results.

For more information about the NASA STI program, see the following:

- Access the NASA STI program home page at [http://www.sti.nasa.gov](http://www.sti.nasa.gov)

- E-mail your question via the Internet to help@sti.nasa.gov

- Fax your question to the NASA STI Help Desk at 301–621–0134

- Telephone the NASA STI Help Desk at 301–621–0390

- Write to:
  NASA Center for AeroSpace Information (CASI)
  7115 Standard Drive
  Hanover, MD 21076–1320
Flame Radiation, Structure, and Scalar Properties in Microgravity Laminar Fires

Douglas Feikema
Glenn Research Center, Cleveland, Ohio

Jongmook Lim and Yudaya Sivathanu
En’Urga, Inc., West Lafayette, Indiana

Prepared for the
45th Aerospace Sciences Meeting and Exhibit
sponsored by the American Institute of Aeronautics and Astronautics
Reno, Nevada, January 8–11, 2007

National Aeronautics and Space Administration

Glenn Research Center
Cleveland, Ohio 44135

June 2007
Acknowledgments

This work is supported by NASA under contract no. NAS3–01085 by the Microgravity Combustion Program, from 2000 to 2004. The authors also acknowledge the support of Jason Taylor and Brian Clark who completed the drop tests and data reduction. The drop tower support personnel are also gratefully acknowledged.

Available from
NASA Center for Aerospace Information
7115 Standard Drive
Hanover, MD 21076–1320

National Technical Information Service
5285 Port Royal Road
Springfield, VA 22161

Available electronically at http://gltrs.grc.nasa.gov
Flame Radiation, Structure, and Scalar Properties in Microgravity Laminar Fires

Douglas Feikema
National Aeronautics and Space Administration
Glenn Research Center
Cleveland, Ohio 44135

Jongmook Lim and Yudaya Sivathanu
En’Urga, Inc.
West Lafayette, Indiana 47906

Results from microgravity combustion experiments conducted in the Zero Gravity Research Facility (ZGF) 5.18 second drop facility are reported. The results quantify flame radiation, structure, and scalar properties during the early phase of a microgravity fire. Emission mid-infrared spectroscopy measurements have been completed to quantitatively determine the flame temperature, water and carbon dioxide vapor concentrations, radiative emissive power, and soot concentrations in microgravity laminar methane/air, ethylene/nitrogen/air and ethylene/air jet flames. The measured peak mole fractions for water vapor and carbon dioxide are found to be in agreement with state relationship predictions for hydrocarbon/air combustion. The ethylene/air laminar flame conditions are similar to previously reported results including those from the flight project, Laminar Soot Processes (LSP). Soot concentrations and gas temperatures are in reasonable agreement with similar results available in the literature. However, soot concentrations and flame structure dramatically change in long-duration microgravity laminar diffusion flames as demonstrated in this report.

1. Introduction

The present study of microgravity laminar diffusion flames was motivated by a requirement for improved understanding of fires in manned spacecraft. NASA’s fire safety approach for manned spacecraft has been based primarily upon controlling the flammability of materials onboard and eliminating ignition sources (refs. 1 to 3). To develop effective fire protection systems, the behavior of fires in various environments encountered in space exploration must be properly understood. Fire behavior and suppression processes in the space exploration missions are strongly influenced by low-gravity environments in flight and on the planetary surfaces. Thus, fire safety technology must be tailored to respond to the unusual fire characteristics in low-gravity environments (ref. 1). The nonbuoyant microgravity environment causes substantial changes in flame structure and consequently changes the properties of the produced smoke and/or soot. In particular, the increased residence time in the high temperature zone where the smoke and/or soot is formed, increases the probability that the particulate will become larger and agglomerated. Among the many technical issues to be answered, this report addresses the issue of flame structure in a simple laminar fire including chemical composition and soot production. Thermal radiation from soot contributes to the growth and spread of unwanted fires on Earth, and soot-containing plumes emitted from these flames inhibit fire-fighting efforts on Earth. No less problematic are the carbon monoxide and unburned hydrocarbon emissions that intrinsically are associated with soot; carbon monoxide emissions are the main cause of fatalities in unwanted fires on Earth. In quiescent microgravity (µg) environments, fires or diffusion flames are relatively weak because of the lack of buoyantly driven convective flow and the slower resulting diffusional transport. The flame on a thick sheet or rod tends to self-extinguish under quiescent conditions (refs. 4 to 6). However, there is normally a constant air circulation, typically 20 cm/s, inside the crew cabin which makes the microgravity flame stronger since oxygen is transported to the fire by forced convection. The flame spread rate in microgravity generally increases when the oxidizer flow velocity increases.
Laminar hydrocarbon diffusion flames under microgravity conditions have shown distinct characteristics relative to normal-gravity flames. Flames are up to two times longer and up to four times larger in diameter, and are often more sooting in microgravity. These characteristics arise from the result of a significant reduction in buoyant convective forces which makes diffusion a dominant mechanism of transport. As a result, increased residence time, enhanced soot formation, radiative cooling due to larger flame size, and the possible onset of a chemical kinetics limitation on heat release processes are observed (ref. 7).

2. Specific Objectives

The objectives of this study are to examine scalar, radiative, and chemical properties of a well-characterized laminar fire (ref. 9) in short duration microgravity. The detailed results help quantify the structure of a nonbuoyant laminar fire which is significant since fire safety technology must be tailored to respond to the unusual fire characteristics in low-gravity environments (ref. 8). In addition, methane/air and ethylene/nitrogen/air laminar jet flame case studies are reported for comparison and validation.

3. Experimental Methods and Instrumentation

The experiment consisted of a general purpose drop rig which was configured with a gas feed system and a laminar jet flame burning inside a chamber similar to the LSP flame (ref. 9). The experiment package was dropped in the Zero Gravity Facility (ZGF) 5.18 second drop facility. The flame conditions are presented in table I for a jet exit diameter of 1.65 mm. Seven test cases are discussed and presented in this report for nearly identical flame conditions. A scanning mid-IR spectrometer, Spectraline Model No. ES100 (En’Urga Inc., West Lafayette, IN) (refs. 10 and 11), was used to obtain spectrally resolved radiation measurements which enabled quantitative determination of gas temperature, CO₂ and H₂O concentrations, and soot volume fraction. Other instrumentation included a radiometer (i.e., thermopile

| TABLE I.—SUMMARY OF TEST CONDITIONS FOR MICROGRAVITY LAMINAR JET FLAMES |
| Test number in Zero Gravity Research Facility |
| Fuel type | CH₄ | CH₄ | 50% C₂H₄ | 50% N₂ | 50% C₂H₄ | 50% N₂ | C₂H₄ |
| Test pressure, kPa | 100 | 100 | 100 | 100 | 100 | 100 | 100 |
| Test temperature, K | 300 | 300 | 300 | 300 | 300 | 300 | 300 |
| Average burner exit velocity, mm/s | 1075 | 1075 | 935 | 935 | 935 | 935 | 935 |
| Initial chamber oxygen concentration, vol% | 21 | 21 | 21 | 21 | 21 | 21 | 20.6 |
| Luminous flame length, mm | 45 | 45 | 32 | 32 | 56 | 55 | 48 |
| Maximum luminous flame width, mm | 32 | 32 | 14 | 14 | 25 | 25 | 26 |
| Fuel flow rate, mg/s | 1.49 | 1.49 | 2.3 | 2.3 | 2.3 | 2.3 | 2.3 |
| Test duration, s | 5.08 | 5.08 | 5.08 | 5.08 | 5.08 | 5.08 | 5.08 |
| Characteristic flame residence time, ms | 84 | 84 | 68 | 68 | 120 | 120 | 120 |
| Fuel exit Reynolds number | 86 | 86 | 117 | 117 | 170 | 170 | 170 |
| Purpose of test | Spectrometer scan at x/d = 13.6 | Spectrometer scan at x/d = 17.6 | Spectrometer scan at x/d = 17.6 | Spectrometer scan at x/d = 12.0 | Spectrometer scan at x/d = 13.6 | Spectrometer scan at x/d = 17.6 | Thermocouple at x/d = 12.0 |

*aHere x is the axial distance measured from the fuel tube exit and d is the fuel tube exit diameter, 1.65 mm.
detector), R-type thermocouple rake, top and side views from the video camera sampling at 30 Hz, and flow and pressure transducers. The flames are ignited in microgravity with a retractable hot-wire igniter approximately 0.1 s after the package is released. By igniting in microgravity, buoyant flow is eliminated enabling the flame to evolve in a quiescent environment.

Figure 1 shows a color side view and black and white top view of the methane/air flame after 5 s into the drop. The flame is “blue” without a glowing “yellow” soot continuum radiation region. The flame becomes less visible as microgravity duration increases, and after 2 s achieves near constant flame length and shape. The flame has an open tip end as seen in figure 1(b) and is nearly symmetric with a maximum visible flame length of 45 mm. Figure 2 shows a color side view and black and white top view of the diluted 50 percent ethylene/50 percent nitrogen/air flame after 5 s into the drop. The flame is “blue” with glowing “yellow” soot continuum radiation region. The flame has a closed tip end as seen in figure 2(b) and is nearly symmetric. Figure 3 shows a color side view and black and white top view of an ethylene/air flame after 5 s into the drop. The flame is “blue” with glowing “yellow” soot continuum radiation region. The flame has an open tip end as seen in figure 3(b) and is nearly symmetric.

The flame length measurements shown in figure 4(a) have been taken from the side view color video camera for the diluted ethylene/nitrogen/air flame shown in figure 2(a). Initially, the flame is short at ignition and grows to attain a flame length of 32 mm on centerline after 5.0 s. In normal gravity under buoyant conditions the same flame is 38 mm in length (ref. 9). The flame length measurements shown in figure 4(b) have been taken from the side view color video camera for the ethylene/air flame shown in figure 3(a). Initially, the flame is short at ignition and grows to attain a flame length of 55 mm on centerline after 5.0 s, which is very similar to results obtained on the KC–135 parabolic aircraft. In normal gravity under buoyant conditions the same flame is 38 mm in length (ref. 9). Under longer duration microgravity conditions (i.e., 170 s) the same flame increases in length to approximately 62 mm in length and reduces in width from 25 mm to 14 mm (ref. 9). Even after 5.0 s the flame is still not at steady state.

Figure 1.—(a) Side view color video camera still after 5.0 seconds for drop number X–12–6, methane/air flame. (b) Top view black/white video still after 5.0 seconds for same flame in figure 1(a) showing open tip.
Figure 2.—(a) Side view color video camera still after 5.0 seconds for drop number X–12–19, 50% Ethylene and 50% nitrogen/air flame. (b) Top view black/white video still after 5.0 seconds for same flame in figure 2(a) showing closed tip.

Figure 3.—(a) Side view color video camera still after 5.0 seconds for drop number X–12–17, ethylene/air flame. (b) Top view black/white video still after 5.0 seconds for same flame in figure 3(a) showing open tip.
A spectrally broadband radiometer (thermopile) detector (ultraviolet to far infrared) was positioned inside the chamber approximately 20 cm above the nozzle and to the side of the flame. The radiometer signal, $I_s$, is recorded during the drop and is proportional to the total flame radiation, which is roughly the surface area of the flame

$$I_s \approx \frac{\cos^2(\theta/2)}{r^2}$$

where $r$ is the distance between the radiation source and the detector (20 cm), and $\theta$ is the angle between the line of sight and the central axis of the detector. For the cases considered in this study, $\theta$ is small and $\cos^2(\theta/2) \sim 1$. The results are shown in figure 5 for a laminar methane/air jet flame in drop X–12–6. The initial peak is the result of ignition and excess fuel gas being burned off; however, after 0.75 s this transient is complete, and the flame begins a steady increase in both flame length and radiative output. After 2 s the flame radiative output has achieved a near constant value but appears to be increasing slowly. The flame length, shape, and size also remain nearly constant after 2 s. Some technical difficulties prevented a full radiometer signal capture on drop X–12–6 after 4.5 s. Figure 6(a) shows flame radiation data for drops X–12–18 and X–12–19 for a laminar jet diffusion flame diluted 50 percent ethylene and 50 percent nitrogen in air. The initial peak is the result of ignition and excess fuel gas being burned off. In less than 0.5 s this transient is complete, and the flame begins a steady increase in both flame length and radiative output. Both the flame length and the radiative output of the flame are increasing at the completion of the 5.18 second drop. Very similar results were obtained for an ethylene/air laminar jet diffusion flame shown in figure 6(b) for drops X–12–16 and X–12–17. Again an ignition transient is observed which disappears in less than 0.5 s, and then the radiative output and the flame length increase throughout the duration of the drop. For the flames considered during the initial phase of the fire, the flame radiation output correlates with the flame size and shape. Throughout longer microgravity duration, other factors begin to affect flame radiation in addition to flame size and surface area or shape.
Figure 5.—Radiometer output proportional to total flame broadband spectral and continuum radiation emitted for the flame shown in figure 1(a) and 1(b), drop X–12–6, methane/air flame.

Figure 6.—(a) Radiometer output proportional to total flame broadband spectral and continuum radiation emitted for the flame shown in figure 2(a) and 2(b), drop X–12–18 and X–12–19, 50% ethylene/nitrogen/air flame. (b) Radiometer output proportional to total flame broadband spectral and continuum radiation emitted for the flame shown in figure 2(a) and 2(b), drop X–12–16 and X–12–17, ethylene/air flame.
A scanning mid-IR spectrometer (refs. 10 and 11) was positioned on the drop rig to measure radial profiles of path-integrated intensities at 160 wavelengths from 1.3 to 4.8 \( \mu m \) horizontally through the flame in figures 1 to 3 at two heights above the burner exit, at 22.5 and 29 mm or \( x/d = 13.6 \) and 17.6, respectively. The spectrometer was fixed outside of the chamber, and the flame was viewed through a sapphire window. The spectrometer is calibrated against a black body radiation source, thus quantitative intensity measurements can be obtained. Radiation in the 1.3 to 3.3 \( \mu m \) band has contributions from both H\(_2\)O and CO\(_2\). Radiation in the 4.3 to 4.9 \( \mu m \) band has contributions only from CO\(_2\). Continuum radiation is also present from the radiating soot and is most evident in the 1.3 to 3.3 \( \mu m \) band.

4. Results and Discussion

As shown in table 1, three different fuels were used in order to observe global differences in flame structure such as flame shape, luminous flame length, flame color(s), and total flame emissive power. In addition, local properties have been measured and estimated from path integrated intensity measurements using a mid-IR spectrometer. Local flame properties have been determined using a method know as fan beam emission tomography (FBET) (ref. 10) from the path integrated intensities including soot volume fraction, radiative emissive power, carbon dioxide and water vapor concentrations, and gas temperature. The characteristic residence times shown in table I are calculated based upon the luminous flame length and the average streamwise velocity or the average fuel exit velocity divided by 2.

\[
\tau_{char} = \frac{2L}{u_0}
\]  

The characteristic residence time for these flames range from 68 ms to 120 ms and are significantly larger than in practical, normal gravity fires where \( \tau_{char} \) is typically less than 10 ms.

4.1 Methane/Air Laminar Diffusion Flames in Microgravity

In normal gravity a laminar methane/air jet diffusion flame appears mostly “blue” to the human eye with a slight amount of “yellow” continuum radiation from soot particles. In microgravity the flame appears only blue after 2 s and is dimmer in the visible wavelengths of light than in normal gravity with no visible soot radiation as shown in figure 1.

Figure 7(a) is the “raw” spectrometer data and shows a contour plot of the scanned measured path integrated intensities as a function of wavelength and radial position at \( x/d = 17.6 \). The path integrated intensities shown in figure 7(a) have been averaged over 2.2 s from 2.8 s to 5.0 s during the drop. As seen in figure 5 the flame radiation is nearly steady during this time period of the drop. Figure 7(b) is one slice or one path integrated intensity spectrally resolved from figure 7(a) at a radial position of \( r = 11.2 \) mm. The peak at 4.3 \( \mu m \) arises from CO\(_2\) emission from vibrational and rotational modes of the molecule. The peak at 2.7 \( \mu m \) appears from H\(_2\)O and CO\(_2\) emissions also from both rotational and vibrational modes of the molecule. It is noteworthy that only discrete gas band radiation is observed in figure 7(b) without the presence of continuum soot radiation.

A deconvolution algorithm and computer code (refs. 10 and 11) were developed which transforms the path integrated intensities into spatially resolved radial profiles of gas temperature, carbon dioxide and water vapor profiles, and soot volume fraction in an axisymmetric, nonhomogeneous flame like those shown in figures 1 to 3. The FBET method makes use of the RADCAL flame radiation database and mathematical model (ref. 12).
Figure 8 shows the radial profiles of flame temperature at $x/d = 13.6$ and 17.6 for the methane/air flame shown in figure 1. The peak flame temperatures decrease from 1675 K at $x/d = 13.6$ to 1500 K at $x/d = 17.6$. For comparison to methane/air laminar diffusion flames in normal gravity, the peak flame temperatures have been measured with FBET and a radiation corrected thermocouple to be ~2100 K, which is significantly hotter than in microgravity but consistent with other measurements and predictions (ref. 10). The microgravity flames are cooler than in normal gravity due to significantly higher fraction of the heat release being lost as radiative heat loss (ref. 9). The principal reason for this phenomenon can be explained by noting that microgravity flames considered here have a long residence time. Fluid particles of fuel exiting the fuel tube reside within the hot flame zone for a relatively long period of time and thereby lose more heat from radiation, cooling the local gas and/or flame temperature than do short residence time flames. Shown in figure 9, at $x/d = 13.6$ and 17.6, respectively, are the radial profiles of gas concentration obtained for the methane/air laminar jet diffusion flame in microgravity. At $x/d = 13.6$ in figure 9(a), the peak values of CO$_2$ and H$_2$O are located at a radial position of approximately 5 mm, which is theoretically the stoichiometric contour and the location of maximum heat release. Here the peak mole fractions are measured to be $X_{H_2O} = 0.08$ and $X_{CO_2} = 0.165$. Some comparable data exists in the literature to provide an evaluation. In a normal gravity methane/air diffusion flame, Norton et al. (ref. 13)
measured and computed peak values of CO₂ and H₂O mole fractions to be 0.08 and 0.165 with an experimental uncertainty of ±20 percent. Sivathanu and Faeth (ref. 14) obtained peak values of CO₂ and H₂O to be 0.075 and 0.172, respectively, for methane/air diffusion flames in normal gravity, based on generalized state relationships for hydrocarbon flames. Very good agreement is observed, and an important observation can be made that the state relationship for major species is valid during the early combustion phase for methane/air combustion. The adiabatic flame temperature for methane/air combustion is 2140 K, assuming the composition is at the state relationship.

4.2 Ethylene/Nitrogen/Air Laminar Diffusion Flames in Microgravity

In normal and microgravity, a laminar 50 percent ethylene and 50 percent nitrogen in air jet diffusion flame appears mostly “yellow” at visible wavelengths due to continuum radiation from soot particles with a slight amount of “blue” radiation outside the soot layer. The stoichiometric contour for this flame lies approximately at the boundary where the luminous yellow radiation ends and the blue radiation appears. The flame is closed-tipped indicating sufficient oxidizer entrainment is present to consume the fuel in an overventilated configuration. In comparison with pure ethylene fuel, the nitrogen dilution has a dramatic effect on reducing the flame length and size.

Figure 10(a) shows the spectrometer data and a contour plot of the scanned measured path integrated intensities as a function of wavelength and radial position at \( x/d = 17.6 \). The path integrated intensities have been averaged over 2.2 s from 2.8 s to 5.0 s during the drop. Figure 4(a) shows the flame radiation is continually increasing during the drop. Figure 10(b) is one line of sight or one path integrated intensity spectrally resolved from figure 10(a) at a radial position of \( r = 11.2 \) mm. The peak at 4.3 \( \mu \)m arises from CO₂ emission, and the peak at 2.7 \( \mu \)m appears from H₂O emission. It is noteworthy that in contrast to the methane/air case shown in figure 7(b), a broadband presence of continuum soot radiation appears in addition to the discrete gas band radiation.

A radial profile of flame temperature at \( x/d = 17.6 \) for the ethylene/nitrogen/air flame is shown in figure 11(a). The peak flame temperature is 1675 K at \( x/d = 17.6 \) near the centerline of the flame at \( r = 0.5 \) mm. The adiabatic flame temperature is 2390 K for this case indicating significant flame enthalpy losses.
TABLE II.—SUMMARY OF STATE RELATIONSHIP MOLE FRACTIONS FOR 50 PERCENT ETHYLENE AND 50 PERCENT NITROGEN/AIR FLAME INITIALLY AT STANDARD TEMPERATURE AND PRESSURE AND AT STOICHIOMETRIC CONDITIONS, $f_5 = 0.05507$

<table>
<thead>
<tr>
<th>$\phi$</th>
<th>$X_{N_2}$</th>
<th>$X_{O_2}$</th>
<th>$X_{CO_2}$</th>
<th>$X_{CO}$</th>
<th>$X_{H_2O}$</th>
<th>$X_{H_2}$</th>
</tr>
</thead>
<tbody>
<tr>
<td>1.0</td>
<td>0.73557</td>
<td>0.01990</td>
<td>0.09707</td>
<td>0.02294</td>
<td>0.11529</td>
<td>0.00922</td>
</tr>
</tbody>
</table>

The radial profiles for the mole fractions of water vapor and carbon dioxide at $x/d = 17.6$ are given in figure 11(b). The peak values for $X_{H_2O}$ and $X_{CO_2}$ are 0.12 and 0.13, respectively. The state relationship predictions for this flame are given in table II and compared for $X_{H_2O}$ and $X_{CO_2}$ are 0.113 and 0.0955, respectively, which is within reasonable agreement. Differences in the measured and predicted values can be explained by the fact that the measurements were made near the end of the closed-tip flame where
interactions from the closed tip are encountered. The peak values of mole fraction for water vapor and 
carbon dioxide are located at approximately the same radial location at \( r = 8 \) mm where the stoichiometric 
mixture fraction is located. Also, it should be noted that the luminous flame zone extends to \( r = 7 \) mm as 
indicated in table I. The blue glow outside the luminous flame boundary at \( r = 8 \) mm can be seen in 
figure 2(a), which is the location of the peak carbon dioxide and water vapor production in the flame. 
Figure 12 presents the soot volume fraction radial profile for drop X–12–18 obtained in an 
ethylene/nitrogen/air flame at \( x/d = 17.6 \). The stoichiometric composition is positioned at a radius 
of 8 mm where the soot production ends. The soot measured from \( r = 0 \) to 6 mm arises since the 
measurement is made near the closed tip of the flame (i.e., \( x = 29 \) mm) and is being transported from 
upstream locations and produced as well since figure 2(b) clearly shows combustion occurring at the end 
of the flame tip. The luminous flame ends at \( x = 32 \) mm so this measurement was made within the closed-
tip end of the flame.

4.3 Ethylene/Air Laminar Diffusion Flames in Microgravity

In normal and microgravity, a laminar 50 percent ethylene and 50 percent nitrogen in air jet diffusion 
flame appears mostly “yellow” at visible wavelengths due to continuum radiation from soot particles 
with a slight amount of “blue” radiation outside the soot layer. The flame is open tipped indicating insufficient 
oxidizer is present to consume the fuel in an underventilated configuration.

Figure 13(a) shows a contour plot of the scanned measured path integrated intensities as a function of 
wavelength and radial position at \( x/d = 13.6 \). These path integrated intensities shown in figure 10(a) have 
been averaged over 2.2 s from 2.8 s to 5.0 s during the drop. As seen in figure 6(b), the total flame 
radiation is nearly steady during this time period of the drop. Figure 13(b) is one line of sight or one path 
integrated intensity spectrally resolved from figure 10(a) at a radial position of \( r = 11.2 \) mm. The peak at 
4.3 \( \mu \)m arises from CO\(_2\) emission, and the peak at 2.7 \( \mu \)m appears from H\(_2\)O and CO\(_2\) emission. In 
contrast to the methane/air case shown in figure 7(b) and similar to figure 10(b), a broadband presence of 
continuum soot radiation appears in addition to the discrete gas band radiation.

A deconvolution algorithm has been developed using an iterative approach to transform the measured 
path integrated intensities into radial profiles of gas temperature, soot concentration, and both H\(_2\)O and 
CO\(_2\) concentration profiles (refs. 10 and 11). The domain is divided into 128 concentric gas layers, and 
the flame is understood to be symmetric, which is a reasonable assumption, as seen in figures 1 to 3. The 
path integrated intensities have been averaged over 2.2 s from 2.8 to 5 s. After approximately 8 iterations 
the gas temperature is obtained and is shown in figure 14(a) at \( x/d = 13.6 \) and in figure 14(b) at \( x/d = 17.6 \).
The peak temperature is approximately 1725 K at a radial position of 12.5 mm and $x/d = 13.6$ and 1675 K at $x/d = 17.6$ and agrees with the results from the LSP experiment (ref. 9). They report a peak temperature of approximately 1850 K at $x = 20$ mm and 1725 K at $x = 30$ mm. The luminous flame boundary at $x = 22.5$ mm is taken from the color video (fig. 3(a)) to lie at approximately 12.5 mm. For comparison, a normal gravity laminar ethylene/air coflow flame was measured with FBET and thermocouples and is shown in figure 15. The peak flame temperature is about 2050 K at $x = 15$ mm, and the reaction zone is much narrower than in figure 14. The theoretical adiabatic flame temperature for ethylene/air is 2475 K.

A three-channel thermocouple (TC) rake, type R platinum/rhodium alloys, with a wire diameter of 75 μm and a bead diameter of approximately 110 μm was implemented. The thermocouple rake was positioned at a height of 20 mm above the burner and at radial locations of $r = 5, 10$ and 15 mm. Figure 16 shows the uncorrected temperature readings. The flame is disturbed somewhat by the invasive probes and has a luminous flame length of 51 mm rather than 55 to 56 mm without the probes. The results show that the flame does broaden as microgravity duration increases and the readings at $r = 5$ and 15 mm are nearly equal after 5 s in agreement with figure 6(a). In addition the thermocouple reading at $r = 10$ mm is higher than at 5 or 15 mm indicating that peak temperature lies between 5 and 15 mm.
Figure 17 shows radial profiles of both H$_2$O and CO$_2$ mole fraction at $x/d = 13.6$ and 17.6, respectively. The peak value of mole fraction for H$_2$O is approximately 0.10, and for CO$_2$ the peak value is approximately 0.09. Both H$_2$O and CO$_2$ concentrations at $x = 15$ mm were measured in the same normal gravity ethylene/air flame as described in figure 15, and the result is shown in figure 18. Again, the peak value of mole fraction for H$_2$O is approximately 0.10, and for CO$_2$ the peak value is approximately 0.09—very similar to the microgravity flames. It is also observed that the flame reaction zone is narrower in normal gravity than in nonbuoyant conditions. Generalized state relationship predictions of major gas species concentrations (ref. 14) in an ethylene/air diffusion flame are shown in table III. The measurements are within 20 percent of this value.
TABLE III.—SUMMARY OF STATE RELATIONSHIP MOLE FRACTIONS FOR ETHYLENE/AIR FLAME INITIALLY AT STANDARD TEMPERATURE AND PRESSURE AND AT STOICHIOMETRIC CONDITIONS, $f_r = 0.06366$

<table>
<thead>
<tr>
<th>$\phi$</th>
<th>$X_{N_2}$</th>
<th>$X_{O_2}$</th>
<th>$X_{CO_2}$</th>
<th>$X_{CO}$</th>
<th>$X_{H_2O}$</th>
<th>$X_{H_2}$</th>
</tr>
</thead>
<tbody>
<tr>
<td>1.0</td>
<td>0.7187</td>
<td>0.0212</td>
<td>0.1033</td>
<td>0.0244</td>
<td>0.1227</td>
<td>0.0098</td>
</tr>
</tbody>
</table>

The radial profiles of soot volume fraction have been determined from the continuum radiation present in the path integrated intensities and are shown in figure 19 at $x/d = 13.6$ and 17.6. The peak value is about 0.8 ppm at 8 to 10 mm radial position. The soot peak lies on the fuel-rich side of the flame and in the preheat region. The soot is being produced on the fuel rich side of the flame between a fuel equivalence ratio of 3.0 and 1.0 with a peak value at $1.6\pm0.3$ according to Gore and Faeth (1986) (ref. 15). This value of 0.8 ppm is much lower than measurements reported previously in microgravity flames. For validation and comparison, the soot volume fraction was measured in a normal gravity sooting ethylene/air flame in a Santoro-type coflow laminar burner (ref. 16) with a fuel flow rate of 3.85 cm$^3$/s.
As shown in figure 20, at \( x = 15 \) mm the peak soot volume fraction was measured to be 3 ppm at a 4 mm radius. The soot layer is much narrower (i.e., 2 mm) as compared with the 5 s drop test which is 5 mm thick. Also, soot measurements were made in a long duration microgravity ethylene/air flame or LSP experiment (ref. 9). The peak soot concentrations in the LSP experiment were approximately 5 ppm at \( x = 20 \) mm and 8 to 12 ppm at \( x = 30 \) mm with microgravity duration times of 90 s and 170 s. The width of the soot layer in the LSP experiment was about 2 mm at \( x = 20 \) and 30 mm.

5. Conclusions

The following conclusions can be obtained from the study

1. A fan beam emission tomography (FBET) system to obtain spectral radiation intensities at several wavelengths and at several view angles from laminar flames was successfully implemented in microgravity.
2. The laminar microgravity ethylene/air jet flame has a broader reaction zone, has a longer luminous flame length, and is cooler than a similar normal gravity flame. This is in agreement with Bahadori et al. (ref. 7).

3. The soot layer in microgravity after 5 s duration is 5 mm which is thicker than in normal gravity (i.e., 2 mm). For longer microgravity duration (i.e., 170 s) the flame shape changes to be longer and narrower, and the soot layer reduces to about 2 mm thickness.

4. The peak mole fractions for H₂O and for CO₂ estimated using FBET are in agreement with the values derived from generalized state relationships for methane/air, ethylene/nitrogen/air, and ethylene/air diffusion flames.

References

Flame Radiation, Structure, and Scalar Properties in Microgravity Laminar Fires

Feikema, Douglas; Lim, Jongmook; Sivathanu, Yudaya

National Aeronautics and Space Administration
John H. Glenn Research Center at Lewis Field
Cleveland, Ohio 44135-3191

Mid-IR spectroscopy; Flames; Combustion; Concentration and temperature measurements; Optical instrumentation

Results from microgravity combustion experiments conducted in the Zero Gravity Research Facility (ZGF) 5.18 second drop facility are reported. The results quantify flame radiation, structure, and scalar properties during the early phase of a microgravity fire. Emission mid-infrared spectroscopy measurements have been completed to quantitatively determine the flame temperature, water and carbon dioxide vapor concentrations, radiative emissive power, and soot concentrations in microgravity laminar methane/air, ethylene/nitrogen/air and ethylene/air jet flames. The measured peak mole fractions for water vapor and carbon dioxide are found to be in agreement with state relationship predictions for hydrocarbon/air combustion. The ethylene/air laminar flame conditions are similar to previously reported results including those from the flight project, Laminar Soot Processes (LSP). Soot concentrations and gas temperatures are in reasonable agreement with similar results available in the literature. However, soot concentrations and flame structure dramatically change in long-duration microgravity laminar diffusion flames as demonstrated in this report.

12. DISTRIBUTION/AVAILABILITY STATEMENT
Unclassified-Unlimited
Subject Categories: 31, 34, 25, 74, and 35
Available electronically at http://gltrs.grc.nasa.gov
This publication is available from the NASA Center for AeroSpace Information, 301-621-0390