Using the Bootstrap Method for a Statistical Significance Test of Differences Between Summary Histograms

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Abstract

A new method is proposed to compare statistical differences between summary histograms, which are the histograms summed over a large ensemble of individual histograms. It consists of choosing a distance statistic for measuring the difference between summary histograms and using a bootstrap procedure to calculate the statistical significance level. Bootstrapping is an approach to statistical inference that makes few assumptions about the underlying probability distribution that describes the data. Three distance statistics are compared in this study. They are the Euclidean distance, the Jeffries-Matusita distance and the Kuiper distance.

The data used in testing the bootstrap method are satellite measurements of cloud systems called “cloud objects.” Each cloud object is defined as a contiguous region/patch composed of individual footprints or fields of view. A histogram of measured values over footprints is generated for each parameter of each cloud object and then summary histograms are accumulated over all individual histograms in a given cloud-object size category. The results of statistical hypothesis tests using all three distances as test statistics are generally similar, indicating the validity of the proposed method. The Euclidean distance is determined to be most suitable after comparing the statistical tests of several parameters with distinct probability distributions among three cloud-object size categories. Impacts on the statistical significance levels resulting from differences in the total lengths of satellite footprint data between two size categories are also discussed.
1. Introduction

In atmospheric and oceanic sciences, the amount of data created from observational and modeling studies is enormous. Satellite measurements can produce a few gigabytes of data a day. High-resolution global climate modeling can easily produce a similar amount of data for a monthly simulation. There are a few possible methods to analyze the large amount of data. One is to average the individual satellite footprint (or field of view) data over a large region and over an extended period so that climate signals are adequately sampled. A similar averaging is also often taken in the climate modeling community. This allows a direct comparison between observations and models so that possible deficiencies of model simulations can be identified. However, this approach does not provide direct clues for improving model simulations because of cancellation of errors in the averaging. Another method is to present observational and modeling data in histograms for each characteristic or parameter to be measured and modeled. This method drastically reduce the data volume because the data volume is only determined by the bin size of the histogram regardless the size of the original data. A significant advantage of this method is that physical clues can be found by comparing model simulations with observations if the data are adequately stratified, for example, according to cloud system types (Xu et al. 2005).

There are many other examples of modeling and observational comparisons that use the histogram representation. For example, Luo et al. (2003) compared the statistical properties of cirrus clouds from cloud-resolving model simulations with ground-based cloud radar measurements. Cloud-resolving models explicitly resolve the cloud-scale dynamics with parameterized microphysics, turbulence and radiation. The cirrus properties were represented in Luo et al. (2003) by histograms of cloud thickness, cloud base and top heights and cloud temperature, etc. Due to large differences between modeled and observed cirrus cloud properties, no vigorous statistical testing was used to gauge the magnitudes of the differences. This has been the case in many similar observational and modeling studies (e.g., Dong and Mace 2003; Sengupta et al. 2004). When the differences between two histograms are not visually large, it is necessary to design a vigorous statistical testing, for example, to address how closely two histograms resemble to each other.
Many statistical methods have been used to analyze observational data and modeling output and to compare the similarities and differences between observational and modeling results in the atmospheric and oceanic sciences (Wilks 1995). To quantify the similarities and differences, statistical significance tests are commonly used to judge, for example, whether or not the means of two time series (or data samples) are statistically different, whether or not two time series are correlated, and whether or not a regression function fits the data adequately. A parametric distribution of the sample is often assumed so that a simple mathematical expression can be used for the statistical significance test, for example, the student \( t \)-test for sample mean and chi-square test for sample standard deviation if the parametric distribution is normal. If the distribution of the population is not known, if it is different from the normal distribution, or if the sample size of the data is small, computationally intensive methods have to be used to perform the statistical significance test, such as the permutation procedure and the bootstrap method (Efron 1982; Livezey and Chen 1983; Zwiers 1990; Efron and Tibshirani 1993; Gong and Richman 1995; Briggs and Wilks 1996; Wilks 1996, 1997; Chu and Wang 1997; Bove et al. 1998; Beersma and Adri Buishand 1999; Hamill 1999; Chu 2002; Sardeshmukh et al. 2000; McFarquhar 2004).

This particular type of statistical significance test is better known as “statistical hypothesis test,” a procedure for evaluating, to a certain required level of significance, the null hypothesis that two sets of data samples are drawn from the same population distribution function. But this procedure cannot determine that two sets of data samples come from a single distribution. A null hypothesis constitutes a specific logical frame of reference between a hypothesized value and an observed value of the test statistic. The test statistic can be any characteristic of the sample. A statistical hypothesis test consists of calculating the probability of obtaining a statistic that is as different or more different from the null hypothesis than the statistic obtained in a sample for which the null hypothesis is correct. If this probability value (\( p \) value) is sufficiently low, the difference between the parameter and the statistic is said to be statistically significant; that is, the null hypothesis is rejected. The level of significance, \( \alpha \), is customarily chosen to be 0.05 or 0.01 in the literature. The \( \alpha \) value is the probability that one will reject the null hypothesis when it is true,
which is called “Type I error.” If one fails to reject the null hypothesis when the null hypothesis is
in fact not true, which is called “Type II error.” The probability (called the $\beta$ value in the litera-
ture) for making the Type II error diminishes as the sample size increases, but it is very difficult to
determine it.

Statistical hypothesis tests have been widely used to explore physical insights of observa-
tions and to evaluate model performance and improvement in many subfields of the atmospheric
and oceanic sciences (e.g., Chu and Wang 1997; Wilks 1997; Bove et al. 1998; Hamill 1999). As
mentioned earlier, information from cloud observation and modeling data can be presented as a
histogram for each characteristic or parameter of a particular cloud system from measurements
made in grids, the points of which are referred to as footprints (or fields of view) in satellite
remote sensing and grid boxes in cloud-resolving modeling. Combining information from several
observed/simulated cloud systems requires combining these histograms into a summary histo-
gram. Summary histograms, which represent signals of combined data samples of observed/simu-
lated cloud systems, can then be compared to each other to infer the statistical behavior of a large
ensemble of cloud systems in different places at different times and in different climatological
conditions or about the performance of a particular cloud-resolving model. Statistical comparison
of histograms can be a robust approach to evaluating and improving cloud-resolving models using
observational data because it is much more difficult for cloud-resolving models to be tuned to
expect improvement in their performance when large numbers of cases are used, compared to that
obtained using a limited number of cases. A detailed discussion of this topic can be found in Xu et
al. (2005) and Eitzen and Xu (2005).

As mentioned previously, it is necessary to decide whether any apparent difference
between histograms is statistically significant. The usual methods of deciding statistical signifi-
cance when comparing histograms do not apply in this case because they assume that the data are
independent. For example, satellite footprint observations or cloud-resolving model output data
are spatially correlated within a cloud system/object. A cloud object is defined as a contiguous
region/patch composed of individual footprints that satisfy a set of physically-based selection cri-
teria. It is possible to design a method to remove this spatial correlation within a cloud object or system (e.g. Zwiers 1990; Wilks 1996; Ebisuzaki 1997). By doing so, the goal of combining large numbers of cloud objects/systems to diagnose climatic signals with small amplitudes (Xu et al. 2005) cannot be achieved. Therefore, a new method is necessary to perform statistical analyses of large numbers of cloud objects/systems. The objective of this study is to present a new method for comparing the statistical differences between summary histograms using statistical hypothesis testing. The proposed method consists of choosing a distance statistic for measuring the difference between summary histograms and using a bootstrap procedure (Efron and Tibshirani 1993) to calculate the statistical significance level. Three distance statistics will be extensively compared in this study using satellite cloud-object data.

Section 2 introduces the basics of the bootstrap procedure. Section 3 discusses the detailed method of comparing histograms. Data used in this study and results are shown in Section 4. Section 5 gives a summary and discussion.

2. Basics of the bootstrap procedure

The bootstrap is an approach to statistical inference that makes few assumptions about the underlying probability distribution that describes the data (Efron 1982; Efron and Tibshirani 1993). This approach assumes that the empirical cumulative distribution function is a reasonable estimate of the unknown cumulative distribution function of the population. That is, the empirical density function approximates the population density function. Using the data as an approximation to the population density function, data are resampled randomly with replacement from the observed sample, i.e. generating uniform random numbers to draw data from the observed sample, to create an empirical sampling distribution for the test statistic under consideration. If one knew the population density function, e.g. the normal probability distribution function (PDF), one could just sample from this PDF, as in a Monte Carlo simulation, and then generate the sampling distribution to within any degree of precision. Due to the lack of this information, one has to assume that the data are a good estimate of the population density function.
Specifically, the empirical cumulative distribution function of the data is a step function, taking jumps of height $1/n$ at each of the sample points. As $n$ increases, the function becomes smooth, and looks more like the population cumulative distribution function, and it converges to the population cumulative distribution function as $n$ approaches infinity. In the bootstrap procedure, the data sample is used as a proxy for the population. That is, one resamples with replacement to produce another data set whose length is equal to the length of the original data sample. This resampling scheme is known as bootstrap resampling, and the resampled data sets are known as bootstrap samples. In this procedure, bootstrap samples may contain duplicate copies of the original data points.

For statistical hypothesis testing, two sets of data are usually used. A test statistic is chosen based upon the null hypothesis to be tested. For example, the null hypothesis is that these two sets of data are from a distribution with the same mean value, and the alternative hypothesis is that they are not. Under the null hypothesis, the two samples are from the same distribution, so merging the samples yields an approximation of the underlying distribution. The bootstrap resampling procedure is then performed to calculate the probability of obtaining a statistic that is as different or more different from the null hypothesis than the statistic obtained from the two samples for which the null hypothesis is correct. If this probability value is sufficiently low, the null hypothesis is rejected. A level of significance of 5% is most often used. It is important to note that it is possible for bootstrapping to yield incorrect results, specifically, the Type I and Type II errors mentioned in Section 1. However, this fact is true of any statistical hypothesis test procedure.

An alternative statistical testing procedure is the permutation procedure, in which each sample is drawn randomly only once without replacement from the merged data. Similar to the bootstrap procedure, the permutation procedure does not need to make any assumptions of the sample distribution. Each sample is randomly chosen from the merged data set by a uniform random number generator, but each sample in the merged data set is labeled and can only be drawn once. Two randomized data sets are then formed. This process is repeated many times to calculate the probability of obtaining a statistic that is as different or more different from the null hypothe-
sis than the statistic obtained from the two samples for which the null hypothesis is correct. In many applications, this procedure yields similar results as the bootstrap procedure (e.g., Livezey and Chen 1983; Livezey et al. 1997). This will also be verified by the results shown in Section 4.

3. Statistical significance test of the difference between summary histograms

Statistical hypothesis tests can be used to detect statistically significant differences between summary histograms, which are the histograms summed over a large ensemble of individual histograms. These tests seem to be more complicated than some simple examples given in the literature because the data samples discussed subsequently are composed of many individual histograms (Fig. 1) with different lengths of data. However, the key aspect in the procedure being outlined below is to choose the test statistic for measuring the difference/dissimilarity of summary histograms, which is also the case for cluster analysis as extensively discussed in Gong and Richmond (1995). Xu et al. (2005) and Eitzen and Xu (2005) briefly discussed a procedure similar to that discussed in detail below. In particular, three test statistics are chosen and compared in detail.

First, a distance statistic is chosen to represent the difference between summary histograms. Three distance statistics are tested in this study. They are the Euclidean distance, the Jeffries-Matusita (JM) distance, and the Kuiper (Kp) distance. The Euclidean distance measures the root-mean-square difference between histograms, which is also called the L2 distance. This distance statistic is defined as

\[ L_2 = \left\{ \sum_{i=1}^{N} \left[ f(x_i)\Delta x - g(x_i)\Delta x \right]^2 \right\}^{1/2} \]  

where \( f \) and \( g \) are two histograms, with a total of \( N \) bins where the \( i \)th bin is located at \( x_i \). The bin width is assumed to be uniform and denoted by \( \Delta x \). The frequency of occurrence is normalized by the bin width. That is, \( f \) and \( g \) satisfy \( \sum_{i=1}^{N} f(x_i)\Delta x = \sum_{i=1}^{N} g(x_i)\Delta x = 1 \). The maximum possible value of \( L_2 \) is \( \sqrt{2} \), and occurs for two single-point histograms that are not collocated.
The Jeffries-Matusita (JM) distance is also commonly used and is designed to find small differences in the bins with small probability densities (Matusita 1955). It is successfully used in pattern classification and image retrieval (e.g. Duda and Hart 1973; Cha and Srihari 2002) and is defined as

\[
JM = \left\{ \sum_{i=1}^{N} \left[ \sqrt{f(x_i) \Delta x} - \sqrt{g(x_i) \Delta x} \right]^2 \right\}^{1/2}
\]

(2).

The values of JM are usually higher than those of L2, as shown in Section 4c, because it gives more weight to the differences at the tails of histograms with small densities than those obtained from the L2 distance. However, the maximum possible value of JM is also \(\sqrt{2}\), as in L2.

The Kolmogorov-Smirnov (KS) distance and its closely related Kuiper distance have often been used in the atmospheric and oceanic sciences for many different applications (e.g. Gong and Richman 1995; Anderson and Stern 1996; Sardeshmukh et al. 2000; Gille and Llewellyn Smith 2000; Gille 2004). The KS distance is determined by the maximum value of the absolute difference between two cumulative density functions. It can be expressed by

\[
KS = \text{Max}_i \left[ \sum_{1}^{i} |f(x_i) - g(x_i)| \Delta x \right] \quad \text{where } i = 1, 2, \ldots, N
\]

(3).

As seen from (3), the KS distance statistic does not integrate over the entire range of histogram bins. This is fundamentally different from the L2 or JM distance statistics. The Kuiper distance statistic (Kuiper 1960) overcomes this deficiency slightly by including the contribution from the tails of the probability distributions, which is defined as

\[
Kp = \text{Max}_i \left[ \sum_{1}^{i} \{f(x_i) - g(x_i)\} \right] + \text{Max}_i \left[ \sum_{1}^{i} \{g(x_i) - f(x_i)\} \right], \quad \text{where } i = 1, 2, \ldots, N
\]

(4).

This distance statistic measures the sum of the maximum distances of accumulative \(f\) above and \(below\) accumulative \(g\). It does not integrate over the entire range of histogram bins, either. The difference between the KS and Kuiper distances is very small from the calculations performed by
the author (not shown). Therefore, only the testing results of the Kuiper distance statistic will be discussed later. The maximum possible values of KS and Kp are 1.

Second, the bootstrap procedure (Efron and Tibshirani 1993) is used to determine whether the difference between summary histograms of a particular parameter is statistically significant. A statistically significant difference between summary histograms means the individual histograms forming the summary histograms came from two different populations. Each cloud object is composed of many adjacent footprints in a contiguous region/patch (see Section 4a for further details). There are several parameters measured at each footprint. Each histogram is calculated from the measurements of the particular parameter of all footprints in a cloud object. Individual histograms of the top-of-the atmosphere (TOA) albedo and cloud top heights shown in Fig. 1 exhibit large variabilities, compared to summary histograms shown in Figs. 2d and e. Furthermore, the number of footprints for calculating the histogram varies from one cloud object to another.

It is assumed that cloud objects, but not their individual footprints, are independent from each other. That is, the histogram of measured values of a particular parameter is not related to that of another cloud object. This assumption can be justified because cloud objects occur in different locations and at different times and their histograms of measured parameters are very different from each other (Fig. 1; Xu et al. 2005). An argument could, however, be made that the cloud objects are dependent based on similar dynamics over a large geographic region where many cloud objects are developed and maintained. Any dependency involved would be very difficult to quantify. On the other hand, the individual satellite footprints within a cloud object, as mentioned earlier, are highly correlated with each other. This is because each footprint occupies an area of a cloud object that is smaller than a relatively homogenous patch within a cloud object, but not small enough to reveal finer-scale details within the relatively homogenous patch.

The null hypothesis for the bootstrap procedure or the permutation procedure is that all cloud objects came from the same population for a particular parameter of the measurements, which allows the merging of two cloud-object populations for bootstrap or permutation resampling. That is, the choice of which cloud objects were from population A and which were from
population B was equivalent to a random choice from the merged data set. Therefore, the distance between the histograms for the “true” ordering is essentially a random number picked from the sampling distribution of the bootstrapped distances.

Specifically, the two data sets of \( m \) and \( n \) cloud objects are first combined into one population for a particular parameter at a time in this study. For example, one of the data sets is composed of cloud objects occurred in the eastern Pacific while the other is composed of cloud objects occurred in the western Pacific. Then, two bootstrap sets of \( m \) and \( n \) cloud objects are resampled randomly from the merged cloud-object population, and the values of the distance statistics between the summary histograms of two bootstrap sets of cloud objects are calculated. Note that different cloud objects have different lengths of satellite footprint data. The summary histograms for a particular parameter are, therefore, not simply the arithmetic means of all individual histograms. Any cloud object in the merged population can be sampled once, more than once, or not all at any given time. This resampling procedure is repeated \( B \) (\( B \) is chosen to be 9999, because the sample sizes are relatively large) times to generate a statistical distribution of the test statistic (L2, JM or Kp). The bootstrap distance value is compared to the value from the true arrangement of cloud objects, i.e. two separate sets or categories of cloud objects. If the bootstrap value of the test statistic of the particular parameter is greater than the observed value of the test statistic in less than 5% of a total calculation of \( B \) times, the two sets of cloud objects are deemed to be statistically different for this particular parameter. That is, the null hypothesis is rejected at the 5% statistical significance level. Because any of the test statistic distances cannot be negative, the 5% level of significance is referred to as the level of significance of one-tailed hypothesis testing. This is true for both the bootstrap and permutation procedures.

4. Results

a. Satellite data for bootstrap analysis

The cloud-object histogram data used in this study are generated from the Clouds and the Earth’s Radiant Energy System (CERES; Wielicki et al. 1996) data from the Tropical Rainfall
Measuring Mission satellite using the cloud object analysis method (Xu et al. 2005). A cloud object is defined as a contiguous region composed of individual cloud footprints that satisfy a set of physically-based cloud-system selection criteria. A "region-growing" strategy based on imager-derived cloud properties is used to identify the cloud objects within a single satellite swath (Wielicki and Welch 1986). Deep convective cloud objects including cumulonimbus and its associated thick upper tropospheric anvils are analyzed, using the following selection criteria: the footprint being overcast, cloud optical depth greater than 10 and cloud top height greater than 10 km. The cloud object data product includes many physical parameters [see Xu et al. (2005) for details]. Only six of them are examined in this study: cloud optical depth, outgoing longwave radiation (OLR) flux, ice water path (IWP), top-of-the-atmosphere (TOA) albedo, cloud top height and sea surface temperature (SST).

Deep convective cloud objects were observed over the tropical Pacific between 25 °S and 25 °N during March 1998. They are combined to some size categories according to the range of their equivalent diameters. Three size categories are considered for the statistical significance testing presented subsequently. They are defined by the diameter ranges of 100 - 150 km (small size), 150 - 300 km (medium size) and greater than 300 km (large size). For convenience, they are termed the S, M and L size categories, respectively. The number of cloud objects for the S, M and L size categories is respectively 126, 136 and 68. The total number of satellite footprints is 14562, 45382, and 109977 for the S, M and L size categories, respectively. The number of satellite footprints also varies from one cloud object to another within each size category. Table 1 shows the statistics of the number of footprints of cloud objects for all three size categories. The differences of the total numbers of footprints among the size categories will have some impacts on the bootstrap results, which will be discussed in Section 4d.

b. The observed histograms

There are large differences in the summary histograms among the six parameters chosen for this study and among the size categories (Fig. 2) although they are much smaller than those
among the individual histograms (Fig. 1). Qualitatively speaking, cloud optical depth exhibits an exponential distribution (Fig. 2a) and a lognormal distribution can be seen in IWP (Fig. 2c). The other four parameters exhibit distributions that deviate from the normal distribution to various degrees. This set of diverse types of histograms will be helpful to understand the bootstrap results presented in Section 4c.

The differences among the three size categories are the target of this investigation. To what degree do the histograms between two size categories differ? Can they be captured by the bootstrap method outlined in Section 3? Without a detailed analysis, it can be concluded that the histograms of OLR and cloud top height are different between the size categories (Figs. 2b, e) and the histograms of other parameters are perhaps more similar between the size categories. The histograms are more different between the S and L size categories than either between the S and M size categories or between the M and L size categories. Will the bootstrap testing support these qualitative conclusions? To address these and other questions, three sets of comparisons are performed, between the S and M size categories, and between the M and L size categories, and between the S and L size categories.

c. The bootstrap results

Tables 2-4 show the observed distances and the associated probability values (p values) for the comparisons between the S and M size categories, between the M and L size categories, and between the S and L size categories, respectively. The empirical sampling distributions (ESDs) of bootstrap distances are shown in Fig. 3 (the L2 statistic), Fig. 4 (the JM statistic) and Fig. 5 (the Kp statistic) for the comparisons among the size categories. A detailed discussion of the results is given below.

Figs. 3-5 show that the ESD of a given bootstrap distance statistic is somewhat similar in terms of the overall shapes among the six parameters in spite of large differences in the histograms of the observed data between two size categories, as discussed in Section 4b. For the L2 distance statistic, in particular, all six ESDs have narrow spreads and high peak frequencies at
small bootstrap distances, except for that of SST (Fig. 3). All ESDs are somewhat symmetric around their modes, except for long tails at large bootstrap distances. The distribution of SST has a much larger spread than that of other five parameters. This is likely related to the existence of point-like PDFs of SST in many of the individual cloud objects. The JM and Kp distance statistics also exhibit the similar ESDs (Figs. 4 and 5), but with some significant differences to be discussed below.

The ESDs of the three different bootstrap distances differ mainly in both the modes and the spreads. The ESDs of the bootstrap JM distance have the widest spreads and the largest modes for all six parameters (Fig. 4). This is consistent with the large values of the observed JM distances shown in Tables 2-4. The observed distances are also represented by vertical bars shown in Figs. 3-5. The relative positions between the vertical bars and the right tails (with the accumulated frequency of 5% starting from the end of the right tail) of their respective ESDs indicate the p values. The closer they are, the higher the p values are, and thus the histograms are less statistically different between a pair of size categories. For example, the vertical bars for the comparisons between the S and M size categories in Figs. 3a, 4a and 5a are located within the interiors of the respective ESDs. All three p values are greater than 0.05 (Table 2). Those vertical bars for the comparison between the S and L size categories are located on the extreme right tails of the respective ESDs. Those three p values are less than 0.01 (Table 4).

The large spreads in the ESDs of the bootstrap distances are related to both the definitions of the distance statistics, discussed in Section 4a, and, to lesser degree, the differences in the histograms of the observed data between two size categories. For example, the large differences at the tails of the histograms of observed OLR, cloud top height and SST corresponds to the large spread in the ESDs of the JM bootstrap distance (Figs. 4b, e and f), which is most sensitive to the low probability densities at the tails. The definition of the L2 distance represents a minimization of the distance between two histograms, compared to the more linear behavior of the Kp distance, which can be seen in both Figs. 3 and 5 because the former is mostly smaller than the latter. This may explain the differences between the distributions of the L2 and Kp bootstrap distances for all
parameters except for the SST. The similarity in the SST ESDs between the L2 and Kp statistics is probably related to the discrete nature of individual cloud-object histograms of SST because an individual cloud object likely occurs over an area with SST variations that are smaller than the bin size of 0.5 K of the histograms.

The statistical hypothesis tests can validate the qualitative comparison between two histograms given in Section 4b, subject to the Type I and II errors discussed in Section 1. These errors can exist for any statistical hypothesis test procedures. The p values are negligible or much smaller than the 5% level for both OLR and cloud top height for any test statistic among the three size categories (Tables 2, 3 and 4). However, the p value for the JM statistic between the M and L size categories is very close to 5%. This difference can be attributed to the difference in the total numbers of footprints between two size categories. This will be discussed in Section 4d.

Statistical hypothesis test results of most of the test statistics are not sensitive to the chosen distance statistic (Tables 2, 3 and 4). This result validates the method proposed in this study. For example, the SSTs are different between the S and M size categories, but similar between the M and L size categories for all three distance statistics. The cloud optical depths are similar between the S and M size categories, but different between the M and L size categories. This result can be attributed to the large difference between the categories that occurs at low cloud optical depths. The albedos are more similar between the M and L size categories than between the S and M size categories, but IWPs are more similar between the S and M size categories than between the M and L size categories.

Statistical significance test results can differ slightly among the distance statistics. This result suggests that some of the distance statistics are more suitable for some specific applications. For example, the p values for the L2 and Kp distance statistics of albedos exceed the 5% threshold between the S and M size categories, but the p value for the JM distance statistic does not. The opposite is true for IWPs between the M and L size categories. These results are not due to the difference in the total numbers of footprints between two size categories shown in Section 4d. There-
fore, it may be suggested that the JM distance statistic is less suitable than the other two statistics for the application presented in this study.

It is difficult to distinguish the statistical significance test results between the L2 and Kp distance statistics because both the observed distances and p values are rather similar (Tables 2 and 3). The narrower spreads in the L2 ESDs in five of the six parameters examined (except for SST) may suggest that the L2 distance statistic is more suitable. A slightly stronger reason for choosing the L2 distance statistic may be its definition, i.e. integrating over the entire bin range in the L2 distance but only over parts of histograms in the Kp distance. In the Kp definition, the differences in the bins between the maximum distances of accumulative $f$ above and below accumulative $g$ are neglected. However, it can be argued that the bins with the largest differences between histograms are more heavily weighted in the L2 distance than those in the Kp distance.

4d. Impact of different footprint numbers of size categories

As mentioned in Section 4a, the total footprint number of cloud object size categories can differ by an order of magnitude, e.g. between the S and L size categories, and so are the average footprint numbers of individual cloud objects (Table 1). This may exaggerate the bootstrap distances if one bootstrap set picks up many more large size cloud objects than the other set. This is because large-size cloud objects contribute greatly to summary histograms. To eliminate such a potential deficiency, we introduce a normalization procedure, as briefly discussed by Eitzen and Xu (2005), to reduce the relative contribution of all cloud objects from the set of cloud objects with a larger total number of footprints. For example, this multiplication ratio is 0.321 (14562/45382) when comparing the S and M size categories, 0.413 (45382/109977) when comparing the M and L size categories. This ratio is multiplied to the histogram of every cloud object within the set of cloud objects with larger total number of footprints. This normalization procedure does not change the observed distances. So, the bootstrap distances obtained from this normalization procedure can be meaningfully compared with the observed distances.
The normalization procedure generally reduces the p values by up to 10% (Tables 4 and 6) for almost all parameters. It is, however, important to note that no p values are changed from over 5% to under 5%, or vice versa. Because a smaller p value means that a result is less likely due to chance, the statistical differences are better captured with the normalization procedure. As mentioned in Section 4c, the p values for the JM statistic between the M and L size categories were very close to 5% for OLR and cloud height and higher than those of the L2 and Kp statistics (Table 3). After the normalization, the p values of the three statistics are 0.9, 0.7 and 0.4 for OLR (1.5, 1.6 and 0.8 for cloud height), respectively. It should be pointed out that a few of the p values (cloud optical depth and IWP) are increased between the S and M size categories by the normalization procedure, but the original p values were very high. So, the normalization does not impact the conclusion at all.

When the p values are very small, for example, less than 1.0%, the normalization procedure gives nearly identical results, as seen from the p values of OLR and cloud top height shown in Table 4. This is what happened in Eitzen and Xu (2005), in which the modeled and observed histograms were compared. The additional insight given in this study should be helpful for improving the statistical significance testing procedure.

5. Summary and discussion

This study has presented a new method to compare statistical differences between summary histograms. The new method consists of choosing a distance statistic for measuring the difference between summary histograms and using a bootstrap procedure to calculate the statistical significance level. Three distance statistics have been compared in this study using satellite data. They are the Euclidean distance, the Jeffries-Matusita distance and the Kuiper distance.

The results of statistical significance tests for all three distance statistics are generally similar, indicating the validity of the proposed method. After comparing the statistical significance tests of several parameters with distinct distributions of satellite footprint data for different cloud-object size categories, the Euclidean distance is determined to be most suitable as the test statistic.
for comparing the statistical difference between histograms. The difference in their definitions is the crucial factor in choosing the L2 distance over the Kuiper distance. The latter does not integrate over the entire range of histogram bins. The bootstrap results based upon the Jefferies-Matusita distance statistic are sometimes inconsistent with those obtained from using the other two distance statistics because it is designed to find small differences at the tails of histograms.

Impacts on the statistical significance levels resulting from differences in the total numbers of satellite footprints between two size categories have also been discussed. In general, the proposed normalization procedure using the ratio of the total numbers of satellite footprints between two populations improves the p values somewhat but does not change the p values from over 5% to under 5%, or vice versa. Because a smaller p value means that a result is less likely due to chance, statistical differences are slightly better captured in applications with large differences in total number of data between two sets of samples by adopting the normalization procedure.

Finally, this new method will be helpful for analyzing large volumes of satellite data and evaluating model performance against the statistics of observational data. Statistical differences between histograms can be quantitatively discussed in the literature of atmospheric and oceanic sciences if this method is adopted.

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References


Ebisuzaki, W., 1997: A method to estimate the statistical significance of a correlation when the data are serially correlated. *J. Climate*, **10**, 2147-2153.


Sengupta, M., E. E. Clothiaux, and T. P. Ackerman, 2004: Climatology of warm boundary-layer clouds at the ARM SGP site and their comparison to models. J. Climate, 17, 4760-4782.


Table captions

Table 1: Statistics of the number of footprints for cloud objects in the S, M and L size categories.

Table 2: The L2, JM and Kp distances and the corresponding probability values (in percent) for the comparison between the S and M size categories.

Table 3: Same as Table 2 except between the M and L size categories.

Table 4: Same as Table 2 except between the S and L size categories.

Table 5: Same as Table 2 except the total numbers of footprints of the two sets of cloud objects have been normalized, where $\Delta$ represents the difference from the p value shown in Table 2.

Table 6: Same as Table 3 except the total numbers of footprints of two sets of cloud objects have been normalized, where $\Delta$ represents the difference from the p value shown in Table 3.

Table 7: Same as Table 3 except the total numbers of footprints of two sets of cloud objects have been normalized, where $\Delta$ represents the difference from the p value shown in Table 3.
Figure captions

Figure 1: Histograms of (a) top-of-the-atmosphere (TOA) albedo and (b) cloud top height for eight selected tropical convective cloud objects observed during the March 1998 period.

Figure 2: Observed histograms for (a) cloud optical depth, (b) outgoing longwave radiation flux, (c) ice water path, (d) top-of-the-atmosphere albedo, (e) cloud top height and (d) sea surface temperature for S (equivalent diameters of 100 - 150 km), M (equivalent diameters of 150 - 300 km) and L (equivalent diameters greater than 300 km) size categories.

Figure 3: Empirical sampling distributions of bootstrap distances for the six parameters shown in Fig. 2 between the S (equivalent diameters of 100-150 km) and M (equivalent diameters of 150-300 km) size categories (solid lines), between the M and L (equivalent diameters greater than 300 km) size categories (short dashed line) and between the S and L size categories (long dashed line) using the L2 test statistic. The bin interval is 0.004 in all panels, except for (e), which uses a bin interval of 0.008. The vertical bars in each panel indicate the observed L2, JM and Kp distances. Note that the axes in (f) are different from those in the rest of the panels.

Figure 4: Same as Fig. 3 except for using the JM test statistic.

Figure 5: Same as Fig. 3 except for using the Kuiper test statistic.
Table 1: Statistics of the number of footprints for cloud objects in the S, M and L size categories.

<table>
<thead>
<tr>
<th>Size category</th>
<th>Mean</th>
<th>Median</th>
<th>Standard deviation</th>
<th>Minimum</th>
<th>Maximum</th>
</tr>
</thead>
<tbody>
<tr>
<td>100 - 150 km (S)</td>
<td>116</td>
<td>112</td>
<td>27</td>
<td>77</td>
<td>170</td>
</tr>
<tr>
<td>150 - 300 km (M)</td>
<td>334</td>
<td>285</td>
<td>140</td>
<td>175</td>
<td>674</td>
</tr>
<tr>
<td>&gt; 300 km (L)</td>
<td>1617</td>
<td>1222</td>
<td>1103</td>
<td>690</td>
<td>7554</td>
</tr>
</tbody>
</table>
Table 2: The L2, JM and Kp distances and the corresponding probability values (in percent) for the comparison between the S and M size categories.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>L2</th>
<th>JM</th>
<th>Kp</th>
<th>p value, L2</th>
<th>p value, JM</th>
<th>p value, Kp</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cloud optical depth</td>
<td>0.030</td>
<td>0.063</td>
<td>0.044</td>
<td>14.1</td>
<td>12.3</td>
<td>18.6</td>
</tr>
<tr>
<td>OLR</td>
<td>0.066</td>
<td>0.148</td>
<td>0.131</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
</tr>
<tr>
<td>IWP</td>
<td>0.024</td>
<td>0.071</td>
<td>0.047</td>
<td>23.6</td>
<td>16.7</td>
<td>18.2</td>
</tr>
<tr>
<td>Albedo</td>
<td>0.050</td>
<td>0.087</td>
<td>0.063</td>
<td>7.5</td>
<td>3.9</td>
<td>6.5</td>
</tr>
<tr>
<td>Cloud height</td>
<td>0.086</td>
<td>0.139</td>
<td>0.122</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
</tr>
<tr>
<td>SST</td>
<td>0.170</td>
<td>0.289</td>
<td>0.168</td>
<td>3.3</td>
<td>1.9</td>
<td>4.4</td>
</tr>
</tbody>
</table>
Table 3: Same as Table 2 except between the M and L size categories.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>L2</th>
<th>JM</th>
<th>Kp</th>
<th>p value, L2</th>
<th>p value, JM</th>
<th>p value, Kp</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cloud optical depth</td>
<td>0.049</td>
<td>0.078</td>
<td>0.066</td>
<td>0.2</td>
<td>1.6</td>
<td>1.3</td>
</tr>
<tr>
<td>OLR</td>
<td>0.044</td>
<td>0.137</td>
<td>0.098</td>
<td>2.3</td>
<td>2.5</td>
<td>1.5</td>
</tr>
<tr>
<td>IWP</td>
<td>0.035</td>
<td>0.074</td>
<td>0.063</td>
<td>2.2</td>
<td>6.6</td>
<td>3.2</td>
</tr>
<tr>
<td>Albedo</td>
<td>0.036</td>
<td>0.051</td>
<td>0.045</td>
<td>26.7</td>
<td>41.1</td>
<td>23.4</td>
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<tr>
<td>Cloud height</td>
<td>0.057</td>
<td>0.121</td>
<td>0.096</td>
<td>3.5</td>
<td>4.7</td>
<td>1.9</td>
</tr>
<tr>
<td>SST</td>
<td>0.090</td>
<td>0.182</td>
<td>0.103</td>
<td>63.0</td>
<td>56.7</td>
<td>50.5</td>
</tr>
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</table>
Table 4: Same as Table 2 except between the S and L size categories.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>L2</th>
<th>JM</th>
<th>Kp</th>
<th>p value, L2</th>
<th>p value, JM</th>
<th>p value, Kp</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cloud optical depth</td>
<td>0.075</td>
<td>0.123</td>
<td>0.100</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
<td>0.7</td>
</tr>
<tr>
<td>OLR</td>
<td>0.093</td>
<td>0.243</td>
<td>0.207</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
</tr>
<tr>
<td>IWP</td>
<td>0.056</td>
<td>0.116</td>
<td>0.100</td>
<td>0.1</td>
<td>0.5</td>
<td>0.1</td>
</tr>
<tr>
<td>Albedo</td>
<td>0.079</td>
<td>0.128</td>
<td>0.091</td>
<td>0.9</td>
<td>0.6</td>
<td>1.1</td>
</tr>
<tr>
<td>Cloud height</td>
<td>0.125</td>
<td>0.230</td>
<td>0.190</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
<td>&lt; 0.1</td>
</tr>
<tr>
<td>SST</td>
<td>0.195</td>
<td>0.333</td>
<td>0.244</td>
<td>10.3</td>
<td>2.3</td>
<td>1.7</td>
</tr>
</tbody>
</table>
Table 5: Same as Table 2 except the total numbers of footprints of the two sets of cloud objects have been normalized, where $\Delta$ represents the difference from the p value shown in Table 2.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>p value, L2</th>
<th>$\Delta$ p value, L2</th>
<th>p value, JM</th>
<th>$\Delta$ p value, JM</th>
<th>p value, Kp</th>
<th>$\Delta$ p value, Kp</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cloud optical depth</td>
<td>16.8</td>
<td>2.7</td>
<td>14.0</td>
<td>1.7</td>
<td>17.7</td>
<td>-0.9</td>
</tr>
<tr>
<td>OLR</td>
<td>&lt; 0.1</td>
<td>0.0</td>
<td>&lt; 0.1</td>
<td>0.0</td>
<td>&lt; 0.1</td>
<td>0.0</td>
</tr>
<tr>
<td>IWP</td>
<td>28.4</td>
<td>4.8</td>
<td>24.2</td>
<td>7.5</td>
<td>20.4</td>
<td>2.2</td>
</tr>
<tr>
<td>Albedo</td>
<td>6.1</td>
<td>-1.4</td>
<td>3.9</td>
<td>0.0</td>
<td>5.7</td>
<td>-0.8</td>
</tr>
<tr>
<td>Cloud height</td>
<td>&lt; 0.1</td>
<td>0.0</td>
<td>&lt; 0.1</td>
<td>0.0</td>
<td>&lt; 0.01</td>
<td>0.0</td>
</tr>
<tr>
<td>SST</td>
<td>1.1</td>
<td>-2.2</td>
<td>3.2</td>
<td>1.3</td>
<td>2.7</td>
<td>-1.7</td>
</tr>
</tbody>
</table>
Table 6: Same as Table 3 except the total numbers of footprints of two sets of cloud objects have been normalized, where Δ represents the difference from the p value shown in Table 3.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>p value, L2</th>
<th>Δ p value, L2</th>
<th>p value, JM</th>
<th>Δ p value, JM</th>
<th>p value, Kp</th>
<th>Δ p value, Kp</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cloud optical depth</td>
<td>0.3</td>
<td>0.1</td>
<td>1.4</td>
<td>-0.2</td>
<td>1.2</td>
<td>-0.1</td>
</tr>
<tr>
<td>OLR</td>
<td>0.9</td>
<td>-1.4</td>
<td>0.7</td>
<td>-1.8</td>
<td>0.5</td>
<td>-1.0</td>
</tr>
<tr>
<td>IWP</td>
<td>1.4</td>
<td>-0.8</td>
<td>5.7</td>
<td>-0.9</td>
<td>2.6</td>
<td>-0.6</td>
</tr>
<tr>
<td>Albedo</td>
<td>20.7</td>
<td>-6.0</td>
<td>32.2</td>
<td>-8.9</td>
<td>18.5</td>
<td>-4.9</td>
</tr>
<tr>
<td>Cloud height</td>
<td>1.5</td>
<td>-2.0</td>
<td>1.6</td>
<td>-3.1</td>
<td>0.8</td>
<td>-1.1</td>
</tr>
<tr>
<td>SST</td>
<td>54.5</td>
<td>-8.5</td>
<td>52.3</td>
<td>-4.4</td>
<td>40.5</td>
<td>-10.0</td>
</tr>
</tbody>
</table>
**Fig. 1:** Histograms of (a) top-of-the-atmosphere (TOA) albedo and (b) cloud top height for eight selected tropical convective cloud objects observed during the March 1998 period.
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