THERMAL AND EVOLVED GAS ANALYSIS OF GEOLOGIC SAMPLES CONTAINING ORGANIC MATERIALS: IMPLICATIONS FOR THE 2007 MARS PHOENIX SCOUT MISSION. H. V. Lauer Jr., D. W. Ming, D. C. Golden, and W. V. Boynton; 1ESCG/Barrios Technology, Houston, TX 77258 (howard.v.lauer1@jsc.nasa.gov); 2ARES NASA/JSC, Houston, TX 77058 (douglas.w.ming@nasa.gov); 3ESCG/Hamilton Sundstrand, Houston, TX 77258, 4University of Arizona, Tucson, AZ 85721

Introduction: The Thermal “and” Evolved Gas Analyzer (TEGA) instrument scheduled to fly onboard the 2007 Mars Phoenix Scout Mission will perform differential scanning calorimetry (DSC) and evolved gas analysis (EGA) of soil samples and ice collected from the surface and subsurface at a northern landing site on Mars. We have been developing a sample characterization data library using a laboratory DSC integrated with a quadrupole mass spectrometer to support the interpretations of TEGA data returned during the mission. The laboratory TEGA test-bed instrument has been modified to operate under conditions similar to TEGA, i.e., reduced pressure (e.g., 100 torr) and reduced carrier gas flow rates [1,2]. We have previously developed a TEGA data library for a variety of volatile-bearing mineral phases, including Fe-oxyhydroxides, phyllosilicates, carbonates, and sulfates [1].

Here we examine the thermal and evolved gas properties of samples that contain organics. One of the primary objectives of the Phoenix Scout Mission is to search for habitable zones by assessing organic or biologically interesting materials in icy soil. Nitrogen is currently the carrier gas that will be used for TEGA. In this study, we examine two possible modes of detecting organics in geologic samples; i.e., pyrolysis using N\textsubscript{2} as the carrier gas and combustion using O\textsubscript{2} as the carrier gas.

Materials and Methods: Green River Shale (SGR-1, U.S. Geological Survey) was used as a geologic standard that contains organics. SGR-1 contains about 28 wt % total C (~3 wt % inorganic C in the form of carbonates and 25 wt % organic C in the form of kerogen). Samples for thermal and evolved gas analyses (TA/EGA) were prepared by diluting the SGR-1 with ultra pure silica powder that had first been heated in air to 1000ºC to combust any organics to formulate samples containing 10 wt % SGR-1 (2.5 wt. % organics).

We have modified our normal TEGA test-bed [1] by removing the DSC and replacing it with a closed system quartz chamber inserted in a vertical tube furnace that will reach 1100ºC (laboratory DSC only obtains temperatures to 725ºC). With this setup we are able to more closely reproduce the flight TEGA geometry and operating conditions with the except of heat flow measurements. All analysis were performed at 100 torr pressure using either O\textsubscript{2} or N\textsubscript{2} as the carrier gas. In preliminary experiments, the quadrupole mass spectrometer was set to scan all mass fragments from 1 to 200 amu to determine organic fragments evolved during either pyrolysis or combustion. Based upon these preliminary experiments, 10 individual mass fragments were selected and monitored during EGA experiments.

Results: As would be expected, there are significant differences in the evolved gas behaviors for samples run under combustion and pyrolysis conditions (Figs. 1 & 2). In combustion experiments, organic C in the form of CO\textsubscript{2} and H\textsubscript{2} evolved between 300-600ºC (Fig. 1); whereas, CO\textsubscript{2} and H\textsubscript{2} evolved between 450-600ºC in pyrolysis experiments (Fig. 2). CO\textsubscript{2} that evolved between 650-750ºC is primarily due to the thermal decomposition of dolomite, although, some of the CO\textsubscript{2} evolved in this range for the pyrolysis experiment may be due to decomposition of the kerogen (see below). Nearly all of the kerogen was “cracked” into CO\textsubscript{2}, H\textsubscript{2}, and H\textsubscript{2}O in the combustion experiments with the exception of a mass 41 fragment that we have assigned to C\textsubscript{3}H\textsubscript{5} and mass 15 that we assigned to methane (Figs. 3 & 4). In addition a small amount of mass 55, 67 and 70 was detected but not included because of space limitations. The methane peak for the combustion experiment has a high sloping background because O\textsubscript{2} is also detected at mass 15 from a small mass 16 drift in our quadrupole mass spectrometer [4].

Figure 1. Evolved CO\textsubscript{2} and H\textsubscript{2} for a combustion experiment (O\textsubscript{2} carrier gas) from a sample containing 10 wt % SGR-1.
THERMAL AND EVOLVED GAS ANALYSIS OF ORGANIC-CONTAINING SAMPLES: H. V. Lauer, Jr. et. al

Figure 2. Evolved CO$_2$ and H$_2$ for a pyrolysis experiment (N$_2$ carrier gas) from a sample containing 10 wt % SGR-1.

Figure 3. Evolved mass 41 molecule during combustion experiments that was assigned as C$_3$H$_5$.

Figure 4. Evolved mass 15 molecule that was assigned as methane in combustion experiments.

Figure 5. Evolved mass 41 molecule during pyrolysis experiments was assigned as C$_3$H$_5$.

Figure 6. Evolved mass 15 molecule that was assigned as methane in pyrolysis experiments.

There were very well-defined releases of C$_3$H$_5$ and methane during the pyrolysis experiments that used N$_2$ as the carrier gas (Figs. 5 & 6). No doubt, H$_2$O and some O$_2$ released during the pyrolysis experiments that will result in some combustion of the kerogen material in SGR-1. However, the N$_2$ carrier gas “preserves” some of the fragments compared to the O$_2$ carrier gas that promotes combustion.

**Implications for 2007 Mars Phoenix Mission.** The N$_2$ carrier gas that is base-lined for TEGA on Phoenix should aid in “preserving” organics (i.e., fragments) that may occur in the icy soil. Additional experiments are underway to look at these systems with trace amounts of organics (e.g., 25 ppm organic C) as well as experiments using mixtures of N$_2$ and trace reducing gases (e.g., CO) that may further “preserve” organics that fragment during pyrolysis.