

The Extreme Spin of the Black Hole Cygnus X-1

Lijun Gou,^{1*} Jeffrey E. McClintock,¹ Mark J. Reid,¹ Jerome A. Orosz,²
James F. Steiner,¹ Ramesh Narayan,¹
Jingen Xiang,¹, Ronald A. Remillard,³
Keith A. Arnaud,^{4,5} Shane W. Davis⁶

¹Harvard-Smithsonian Center for Astrophysics,
60 Garden Street, Cambridge, MA, 02138, USA

²Department of Astronomy, San Diego State University,
5500 Campanile Drive, San Diego, CA 92182, USA

³Kavli Institute for Astrophysics and Space Research,
MIT, 70 Vassar Street, Cambridge, MA 02139, USA

⁴CRESST, NASA Goddard Space Flight Center,
8800 Greenbelt Road, Greenbelt, MD 20771, USA

⁵Astronomy Department, University of Maryland,
College Park, MD 20742, USA

⁶Canadian Institute for Theoretical Astrophysics,
University of Toronto, Toronto, ON M5S 3H8, Canada

*To whom correspondence should be addressed; E-mail: lgou@cfa.harvard.edu.

Remarkably, an astronomical black hole is completely described by the two numbers that specify its mass and its spin. Knowledge of spin is crucial for understanding how, for example, black holes produce relativistic jets. Recently, it has become possible to measure the spins of black holes by focusing on the very inner region of an accreting disk of hot gas orbiting the black hole. According to General Relativity (GR), this disk is truncated at an inner radius

that depends only on the mass and spin of the black hole. We measure the radius of the inner edge of this disk by fitting its continuum X-ray spectrum to a fully relativistic model. Using our measurement of this radius, we deduce that the spin of Cygnus X-1 exceeds 97% of the maximum value allowed by GR.