Water bodies are challenging terrain hazards for terrestrial unmanned ground vehicles (UGVs) for several reasons. Traversing through deep water bodies could cause costly damage to the electronics of UGVs. Additionally, a UGV that is either broken down due to water damage or becomes stuck in a water body during an autonomous operation will require rescue, potentially drawing critical resources away from the primary operation and increasing the operation cost. Thus, robust water detection is a critical perception requirement for UGV autonomous navigation.

One of the properties useful for detecting still water bodies is that their surface acts as a horizontal mirror at high incidence angles. Still water bodies in wide-open areas can be detected by geometrically locating the exact pixels in the sky that are reflecting on candidate water pixels on the ground, predicting if ground pixels are water based on color similarity to the sky and local terrain features. But in cluttered areas where reflections of objects in the background dominate the appearance of the surface of still water bodies, detection based on sky reflections is of marginal value. Specifically, this software attempts to solve the problem of detecting still water bodies on cross-country terrain in cluttered areas at low cost.

Still water bodies are indirectly detected in cluttered areas of cross-country terrain by detecting reflections of objects in the water bodies using imagery acquired from a stereo pair of color cameras, which are mounted to the front of a terrestrial UGV. Object reflections can be from naturally occurring (e.g. vegetation, trees, hills, mountains, clouds) or man-made entities (e.g. signs, poles, vehicles, buildings, bridges). Color cameras provide a lower-cost solution than specialized imaging sensors (such as a polarization camera) and laser scanners. In addition, object reflections can be detected in water bodies with stereo vision at further ranges than with lidar scanners.

A hole in Stereo Range Data may be a water body still too small to be detected in image space. In this example, the hole was labeled a potential hazard in the world map in frame N. In the next frame, where there was previously a hole, there was range data that was detected as an object reflection, providing confirmation of a water body.
SATPLOT for Analysis of SECCHI Heliospheric Imager Data
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Determining trajectories of solar transients such as coronal mass ejections is not always easy. White light images from SECCHI’s (Sun Earth Connection Coronagraph and Heliospheric Investigation) heliospheric imagers are difficult to interpret because they represent a line-of-sight projection of optically thin solar wind structures. A structure’s image by itself gives no information about its angle of propagation relative to the Sun-spacecraft line, and an image may show a superposition of several structures, all propagating at different angles. Analyzing SECCHI heliospheric imager data using plots of elongation (angle from the Sun) versus time at fixed position angle (aka “jplots”) has proved extremely useful in understanding the observed solar wind structures. This technique has been used to study CME (coronal mass ejection) propagation, CIRs (corotating interaction regions), and blobs.

SATPLOT software was developed to create and analyze such elongation versus time plots. The tool uses a library of cylindrical maps of the data for each spacecraft’s panoramic field-of-view. Each map includes data from three SECCHI white-light telescopes (the COR2 coronagraph and both heliospheric imagers) at one time for one spacecraft. The maps are created using a Plate Carree projection, optimized for creating the elongation versus time plots. The tool can be used to analyze the observed tracks of features seen in the maps, and the tracks are then used to extract information, for example, on the angle of propagation of the feature.

The role of PLUTO (Plug-in Port Utilization Officer) and the growth of the International Space Station (ISS) have exceeded the capabilities of the current tool PiP (Plug-in Plan). Its users (crew and flight controllers) have expressed an interest in a new, easy-to-use tool with a higher level of interactivity and functionality that is not bound by the limitations of Excel.

The PiP Tool assists crewmembers and ground controllers in making real-time decisions concerning the safety and compatibility of hardware plugged into the UOPs (Utility Outlet Panels) onboard the ISS. The PiP Tool also provides a reference to the current configuration of the hardware plugged in to the UOPs, and enables the PLUTO and crew to test Plug-in locations for constraint violations (such as cable connector mismatches or amp limit violations), to see the amps and volts for an end item, to see whether or not the end item uses 1553 data, and the cable length between the outlet and the end item. As new equipment is flown or returned, the database can be updated appropriately as needed. The current tool is a macro-heavy Excel spreadsheet with its own database and reporting functionality.

The new tool captures the capabilities of the original tool, ports them to new software, defines a new dataset, and compensates for ever-growing unique constraints associated with the Plug-in Plan. New constraints were designed into the tool, and updates to existing constraints were added to provide more flexibility and customizability. In addition, there is an option to associate a “Flag” with each device that will let the user know there is a unique constraint associated with it when they use it. This helps improve the safety and efficiency of real-time calls by limiting the amount of “corporate knowledge” overhead that has to be trained and learned through use.

The tool helps save time by automating previous manual processes, such as calculating connector types and deciding which cables are required and in what order.