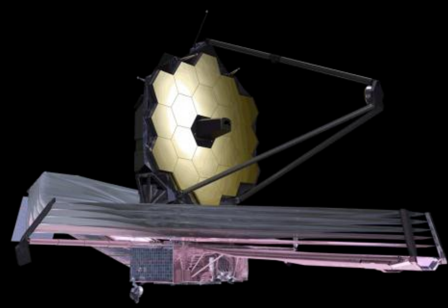
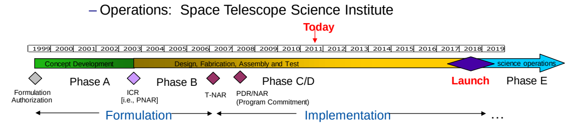
# James Webb Space Telescope (JWST)



The First Light Machine

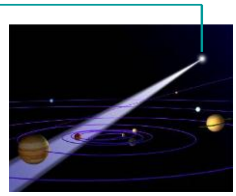
## JWST Summary

- **Mission Objective**
  - Study origin & evolution of galaxies, stars & planetary systems
  - Optimized for near infrared wavelength (0.6 –28 μm)
  - 5 year Mission Life (10 year Goal)
- **Organization**
  - Mission Lead: Goddard Space Flight Center
  - International collaboration with ESA & CSA
  - Prime Contractor: Northrop Grumman Space Technology
  - Instruments:
    - Near Infrared Camera (NIRCam) – Univ. of Arizona
    - Near Infrared Spectrometer (NIRSpec) – ESA
    - Mid-Infrared Instrument (MIRI) – JPL/ESA
    - Fine Guidance Sensor (FGS) – CSA
  - Operations: Space Telescope Science Institute

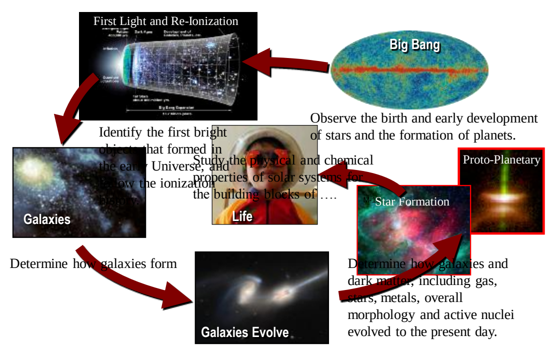



## Origins Theme's ***Two*** Fundamental Questions

- How Did We Get Here?
- Are We Alone?



## JWST Science Themes




## Three Key Facts

There are 3 key facts about JWST that enables it to perform is Science Mission:

It is a Space Telescope

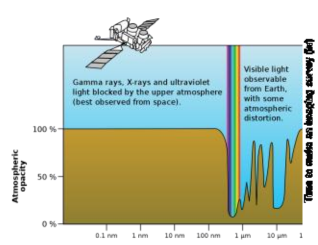
It is an Infrared Telescope

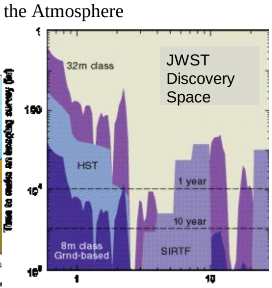
It has a Large Aperture

## Why go to Space

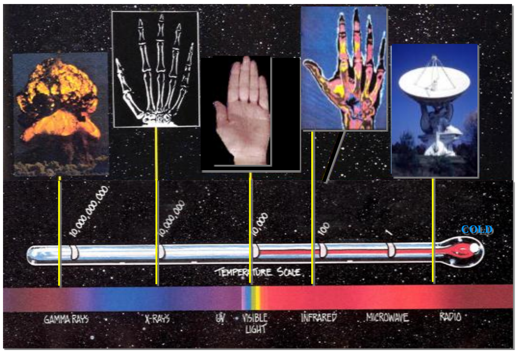
Atmospheric Transmission drives the need to go to space.

Infrared (mid and far/sub-mm) Telescopes (also uv, x-ray, and gamma-ray) cannot see through the Atmosphere



## Infrared Light




## Why Infrared ?

## California Nebula in Center of Picture
Pleiades (Seven Sisters) Cluster is close to bottom.

In the visible this region of space is black and devoid of information.
(from Hubble)

In the sub-mm this region of space is full of the coldest matter of the Universe.
(from Planck, ESA)


## M-83

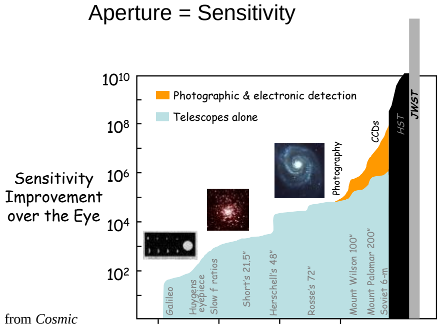
Visible light (taken by the Wide Field Imager on the 2.2-metre MPG/ESO telescope at La Silla in Chile) and infrared (as seen by HAWK-I).

In the infrared, the dust that obscures many stars becomes nearly transparent, making the spiral arms less dramatic, but revealing a whole host of new stars that are otherwise invisible.

Image: ESO/M. Gieles.

(Dr. Emily Baldwin, Astronomy Now, 20 May 2010)

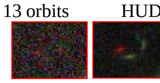
## Why do we need Large Apertures?
Aperture = Sensitivity



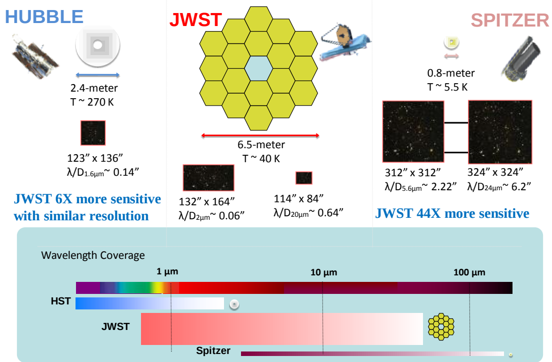

Adapted from *Cosmic Discovery*, M. Harwit

## **Sensitivity Matters**

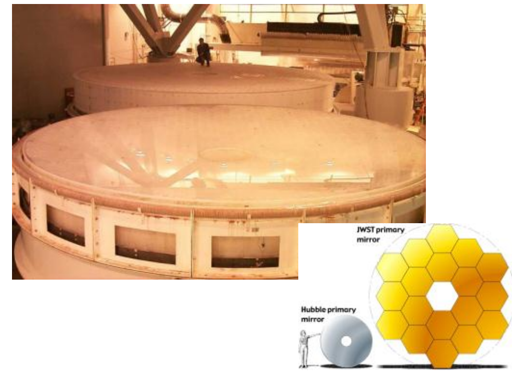
GOODS CDFS – 13 orbits HUDF – 400 orbits



## JWST will be more Sensitive than Hubble or Spitzer




## How big is JWST?




## Full Scale JWST Mockup

21st National Space Symposium, Colorado Springs, The Space Foundation

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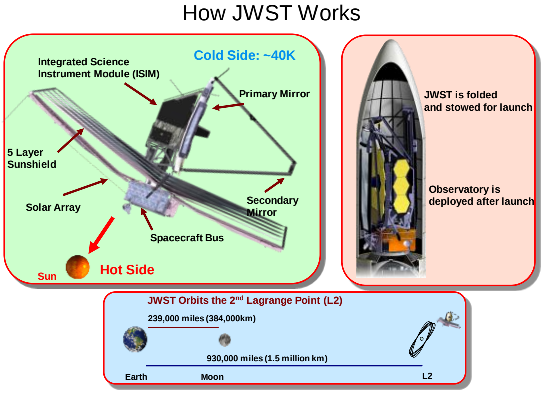
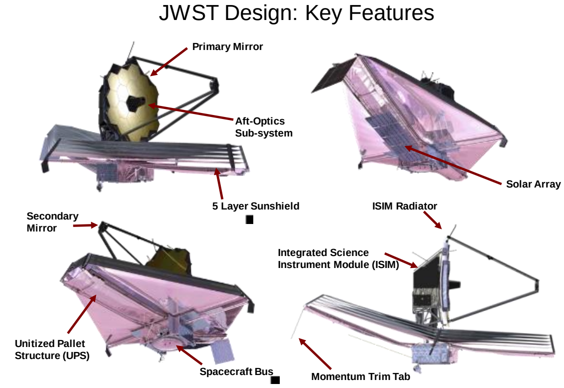
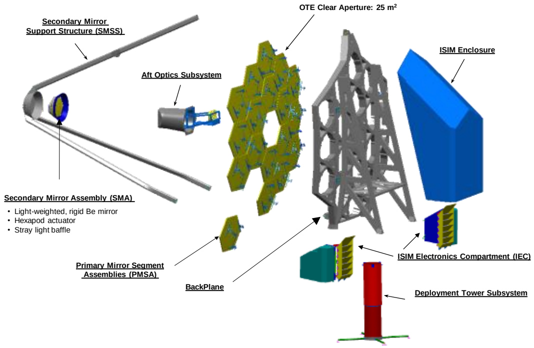
## How JWST Works




## JWST Design: Key Features



## OTE Architecture Concept

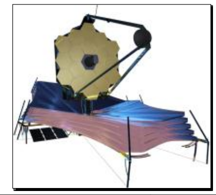



## JWST Requirements

Optical Telescope Element
- 25 sq meter Collecting Area
- 2 micrometer Diffraction Limit
- < 50K (~35K) Operating Temp

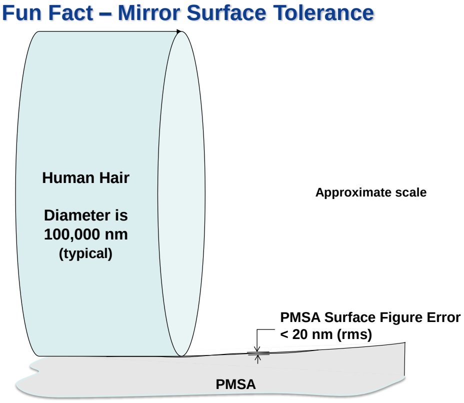
Primary Mirror
- 6.6 meter diameter (tip to tip)
- < 25 kg/m² Areal Density
- < $6 M/m² Areal Cost
- 18 Hex Segments in 2 Rings
- Drop Leaf Wing Deployment

Segments
- 1.315 meter Flat to Flat Diameter
- < 20 nm rms Surface Figure Error

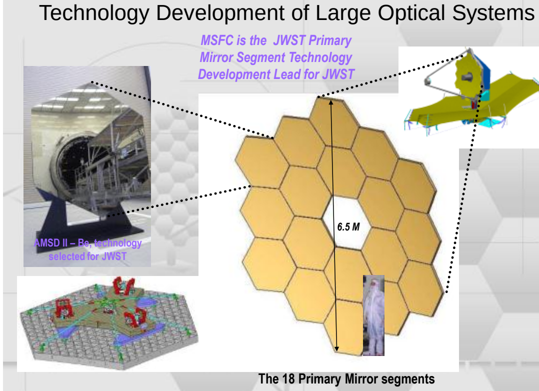


| | |
|---|---|
| Low (0-5 cycles/aper) | 4 nm rms |
| CSF (5-35 cycles/aper) | 18 nm rms |
| Mid (35-65K cycles/aper) | 7 nm rms |
| Micro-roughness | <4 nm rms |

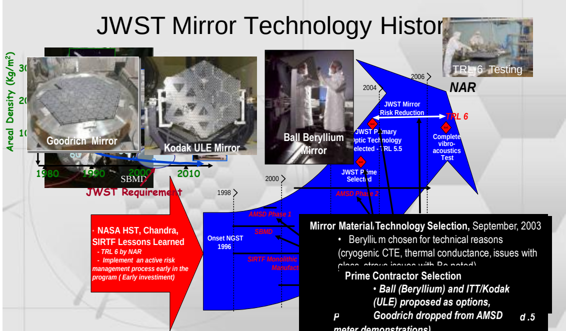
## Fun Fact – Mirror Surface Tolerance




## Technology Development of Large Optical Systems




## JWST Mirror Technology History



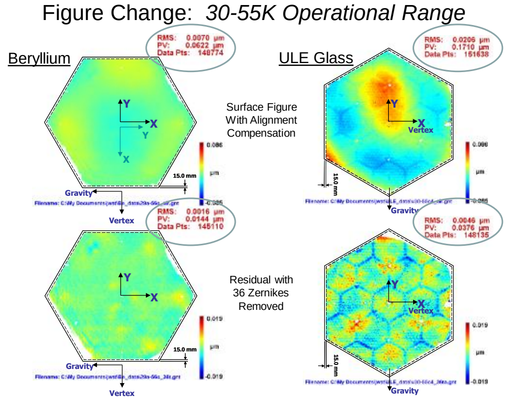

Based on lessons learned, JWST invested early in mirror technology to address lower areal densities and cryogenic operations

## Advantages of Beryllium

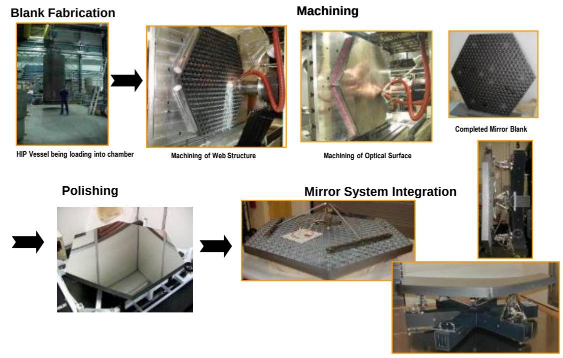
Very High Specific Stiffness – Modulus/Mass Ratio
- Saves Mass – Saves Money

High Conductivity & Below 100K, CTE is virtually zero.
- Thermal Stability

## Figure Change: *30-55K Operational Range*

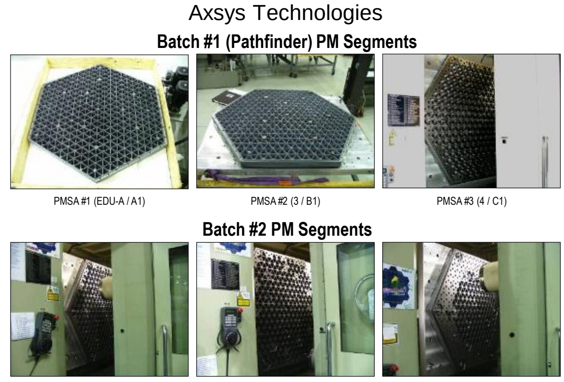



## Mirror Manufacturing Process

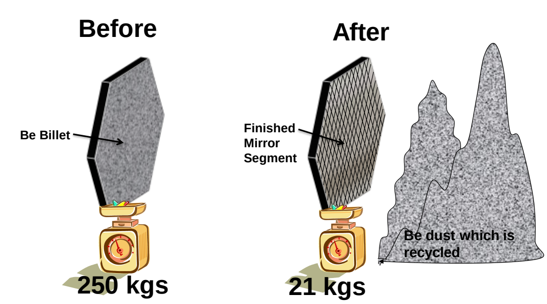



## Brush Wellman

## Axsys Technologies

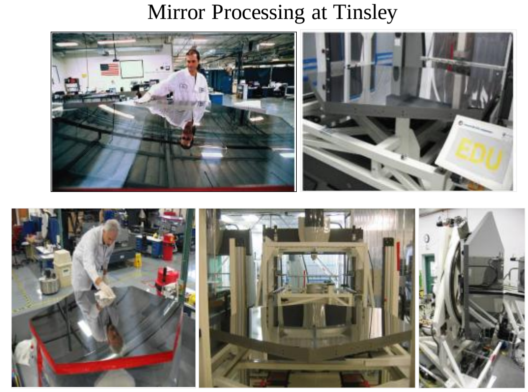



## Fun Facts – Mirror Manufacturing



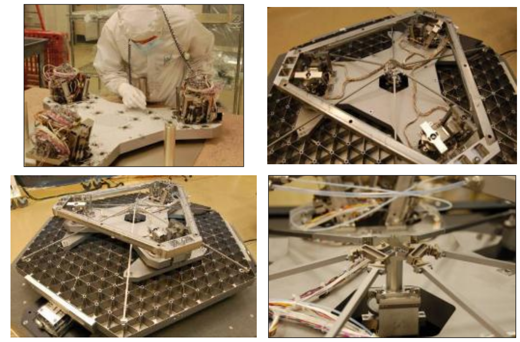

Over 90% of material is removed to make each mirror segment – want a little mirror with your Be dust?

## Mirror Processing at Tinsley

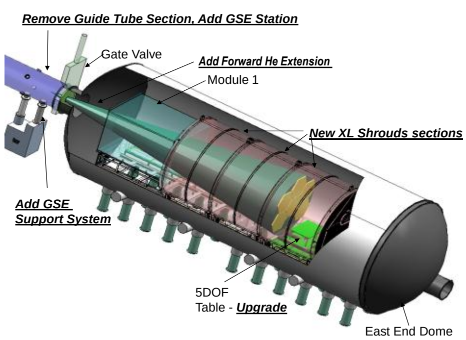


## Tinsley Laboratory – Final Shipment

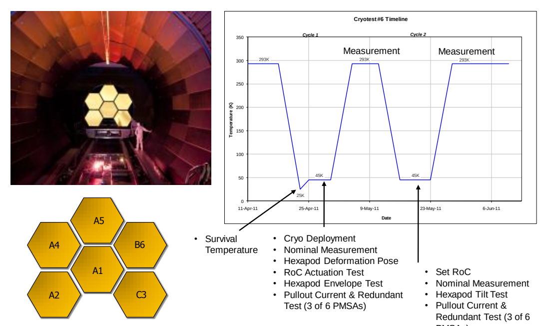

## Primary Mirror Segment Assembly at BATC



## MSFC Cryogenic Test Facility



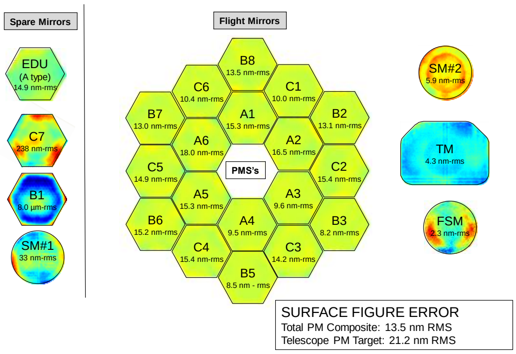

## XRCF Cryo Test



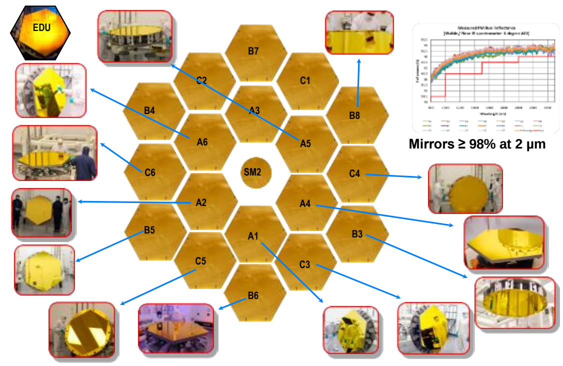

## Flight Mirrors in XRCF

## Primary Mirror Cryogenic Tests

## Mirror Fabrication Status at L-3 SSG-Tinsley – July 11
ALL DONE & DELIVERED



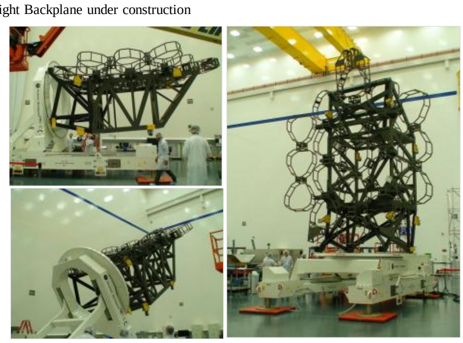

## Gold Coated Mirror Assemblies




## Primary Mirror Backplane

Pathfinder backplane (central section) is complete for test procedure verification at JSC

Flight Backplane under construction



## OTE Integration

**GSFC - JWST Cleanroom**

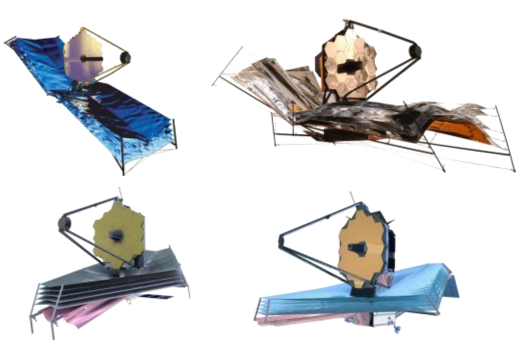
**Watch JWST being built on the web !**

http://jwst.gsfc.nasa.gov/webcam.html

**Backplane Support Structure**

**Robot arm for mirror installation**

## Passive Cooling: Sunshield Evolution



## Sunshield Development

**Evolutionary Pathfinder**

**Bench Test Articles**

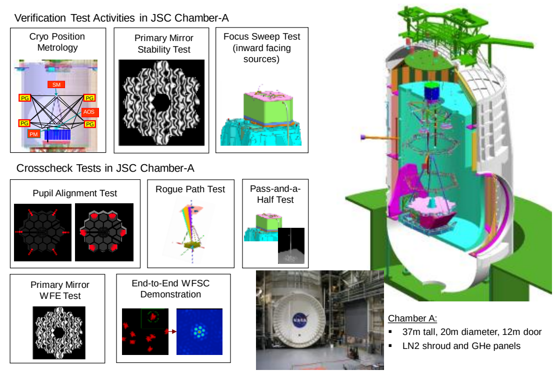
**Deployment tests**

**Clearance check**

## Fight Sunshield at NeXolve in Huntsville

Lee Roop, The Huntsville Times, September 20, 2011

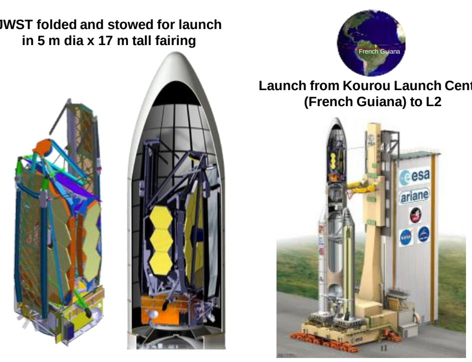
## Observatory level testing occurs at JSC Chamber A

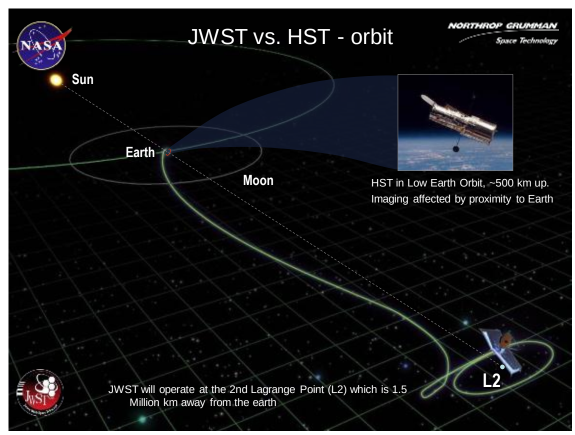



## JWST Launched on Ariane 5 Heavy

**JWST folded and stowed for launch in 5 m dia x 17 m tall fairing**

**Launch from Kourou Launch Center (French Guiana) to L2**

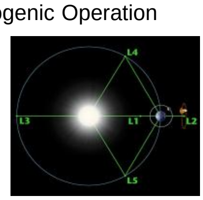


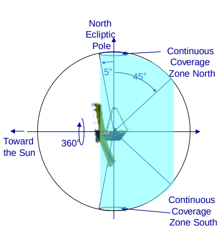



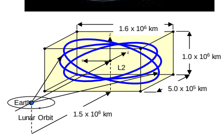

## L2 Orbit Enables Passive Cryogenic Operation

Second Lagrange Point (L2) of Sun-Earth System
- This point follows the Earth around the Sun

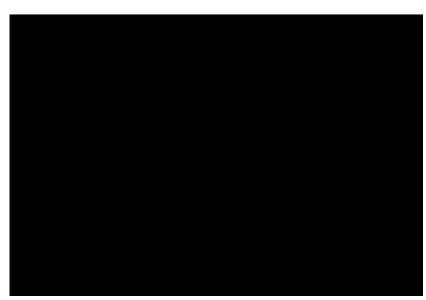
The orbital period about L2 is ~ 6 months
- Station keeping thrusters required to maintain orbit
- Propellant sized for 11 years (delta-v ~ 93



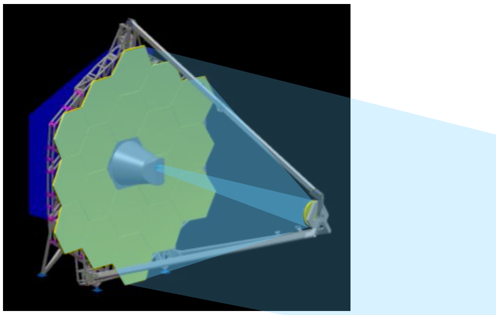



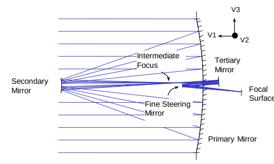

JWST observes whole sky while remaining continuously in shadow of its sunshield
- Field of Regard is annulus covering 35% of the sky
- Whole sky is covered each year



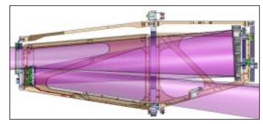
## JWST Optical Path

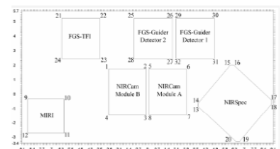


## The JWST telescope is a three mirror anastigmat equipped with a fine steering mirror

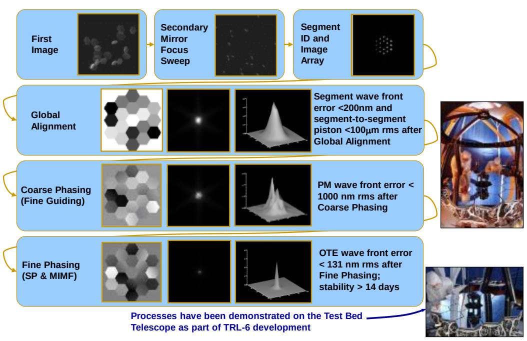



JWST's is a Three Mirror Anistigmat (TMA)
- Optical design: f/20
- Diameter of PM: 6.6 m
- Effective focal length: 131.4 m
- Clear aperture area: 25 m$^2$
- Field of view: 18.2 x 9.1 arcmin
- Elliptical f/1.2 Primary Mirror
- Hyperbolic Secondary Mirror creates f/9 intermediate image
- Elliptical Tertiary Mirror images pupil at Fine Steering Mirror
- Transmitted Wavefront Error is 131 nm rms



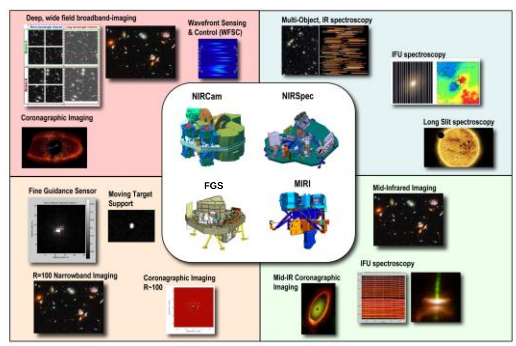


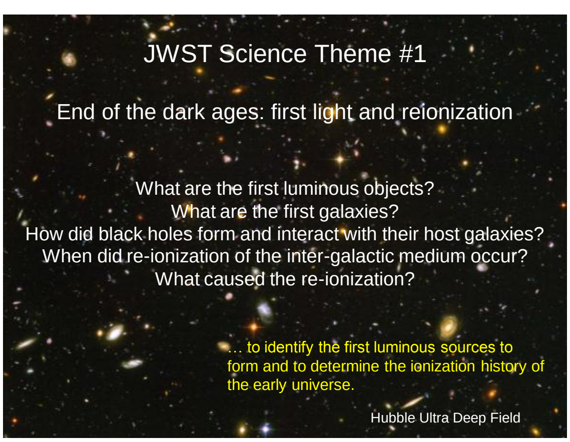

## Deployed Telescope Phasing



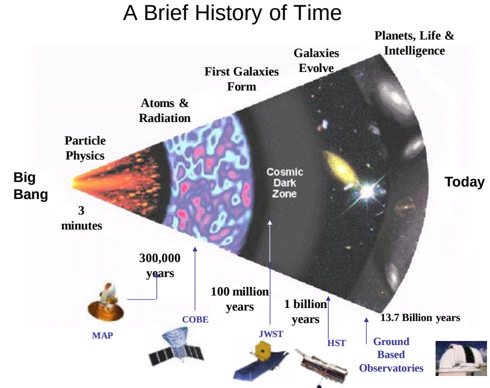

## JWST Science Instruments

enable imagery and spectroscopy over the 0.6 – 29 micron spectrum



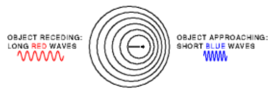



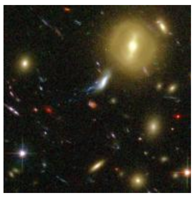

## A Brief History of Time

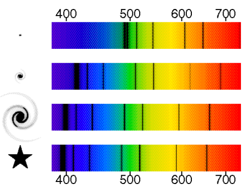


## Redshift

The further away an object is, the more its light is redshifted from the visible into the infrared.





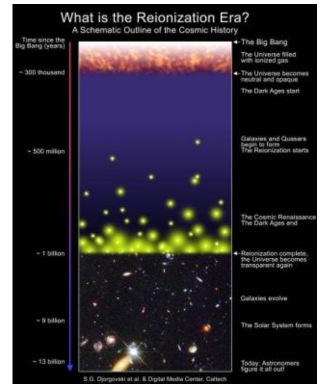

## When and how did reionization occur?

Re-ionization happened at z > 6 or < 1 B yrs after Big Bang.
WMAP says maybe twice?

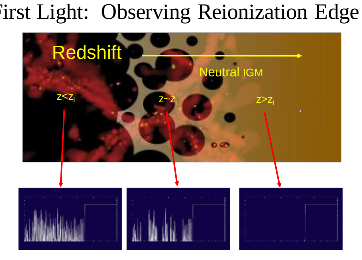
Probably galaxies, maybe quasar contribution

Key Enabling Design Requirments:
- Deep near-infrared imaging survey (1nJy)
- Near-IR multi-object spectroscopy
- Mid-IR photometry and spectroscopy

JWST Observations:
- Spectra of the most distant quasars
- Spectra of faint galaxies




## First Light: Observing Reionization Edge




At 780 M yrs after Big Bang the Universe was up to 50% Neutral. But, by 1 B years after BB is was as we see it today.

Neutral 'fog' was dissolved by very bright 1st Generation Stars (5000X younger & ~100X more massive than our sun).

SPACE.com, 12 October 2011

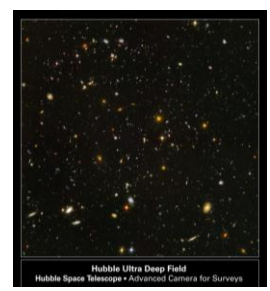
## When and How did the First Galaxies Form

First galaxies are small & faint

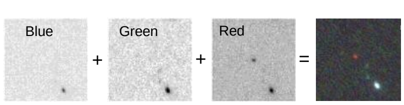
Light is redshifted into infrared.

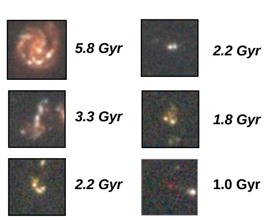
Low-metallicity, massive stars.
SNe! GRBs!

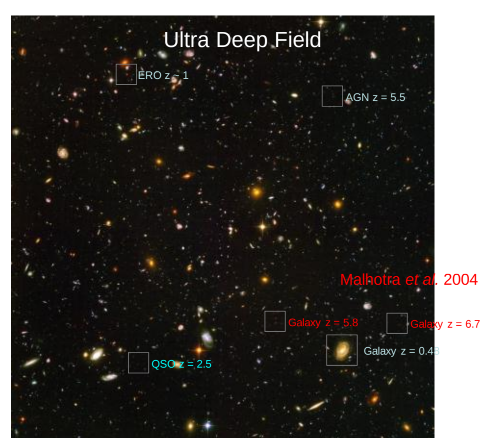
JWST Observations
- Ultra-Deep NIR survey (1.4 nJy), spectroscopic & Mid-IR confirmation.



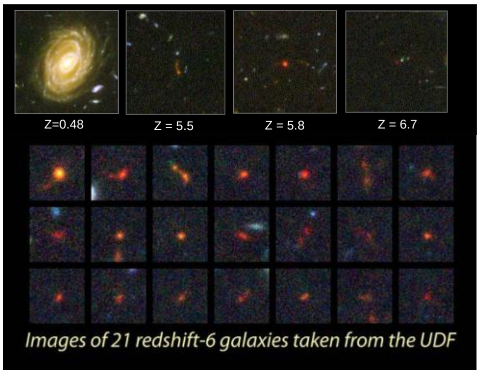

## How do we see first light objects?



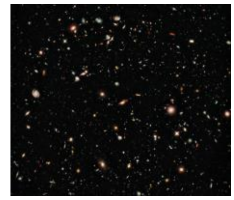



## Results from UDF

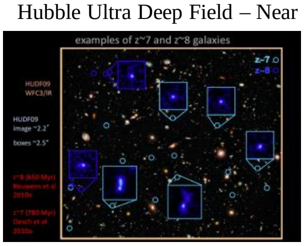



## Hubble Ultra Deep Field – Near Infrared



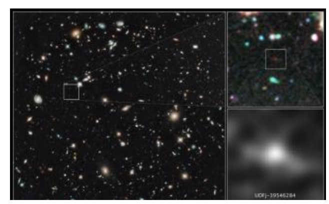
Near-Infrared image taken with new Wide-Field Camera 3 was acquired over 4 days with a 173,000 second exposure.

## Hubble Ultra Deep Field – Near Infrared

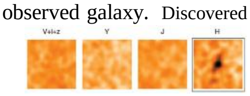



47 Galaxies have been observed at 600 to 650 Myrs after BB.

## Hubble Ultra Deep Field – Near Infrared

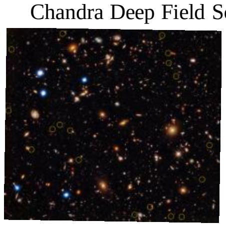


At 480 M yrs after big bang (z ~ 10) this is oldest observed galaxy. Discovered using drop-out technique.




Left image is visible light, and the next three in near-infrared filters. The galaxy suddenly pop up in the H filter, at a wavelength of 1.6 microns (a little over twice the wavelength the eye can detect). (Discover, Bad Astronomy, 26 Jan 2011)

## Hubble Ultra Deep Field – Near Infrared
## Chandra Deep Field South



CREDIT: X-ray: NASA/CXC/U.Hawaii/E.Treister et al; Optical: NASA/STScI/S.Beckwith et al

Keith Cooper, Astronomy Now, 15 June 2011
Taylor Redd, SPACE.com, 15 June 2011

What came first – Galaxies or Black Holes?

Each of these ancient 700 M yrs after BB galaxies has a black hole.

Only the most energetic x-rays are detected, indicating that the black-holes are inside very young galaxies with lots of gas.

## Oldest & Brightest Quasar – 770M yrs after BB

This Quasar is 770 million years after Big Bang, is powered by a black hole 2 billion times the mass of our Sun and emits 60 trillion times as much light as the sun. How a black hole became so massive so soon after the Big Bang is unknown.

Image of ULAS J1120+0641, a very distant quasar powered by a black hole with a mass 2 billion times that of the sun, was created from images taken from surveys made by both the Sloan Digital Sky Survey and the UKIRT Infrared Deep Sky Survey. The quasar appears as a faint red dot close to the centre. CREDIT: ESO/UKIDSS/SDSS

"It is like finding a 6-foot-tall child in kindergarten," says astrophysicist Marta Volonteri, at the University of Michigan in Ann Arbor.

The spectra of the light from this (and other early light objects) indicate that the Universe was still filled with significant amounts of neutral hydrogen even 770 Myrs after big bang.

Nadia Drake , Science News, 29 June 2011
Charles Q. Choi, SPACE.com, 29 June 2011

## Oldest Gamma Ray Burst – 520M yrs after BB

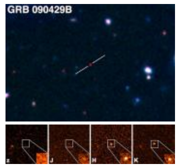
29 Apr 2009, SWIFT detected 5 sec gamma ray burst.
Afterglow in Gemini image has no visible light.
Also, no red-shifted Lyman 'forest' was detected.

Once afterglow faded, nothing was visible

TOO FAR

Estimated Age is 520 million years after big bang, 13.14 billion light-years from Earth (Red Shift 9.4).

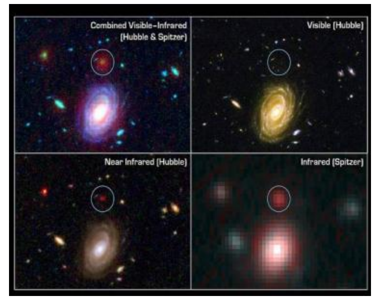
These first light objects are TOO RED SHIFTED for current telescopes. JWST will study them.



Credit: Gemini Observatory

Jonathan Amos Science correspondent, BBC News
Tammy Plotner, Universe Today, May 26, 2011

## Unexpected "Big Babies": 800M yrs after BB




Spitzer and Hubble have identified a dozen very old (almost 13 Billion light years away) very massive (up to 10X larger than our Milky Way) galaxies.

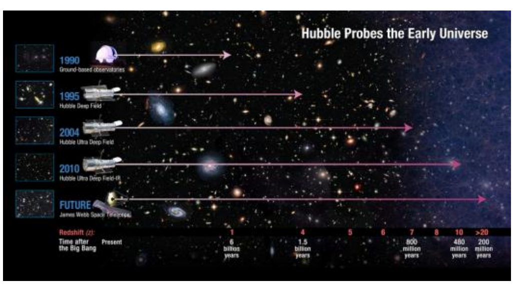
At an epoch when the Universe was only ~15% of its present size, and ~7% of its current age.

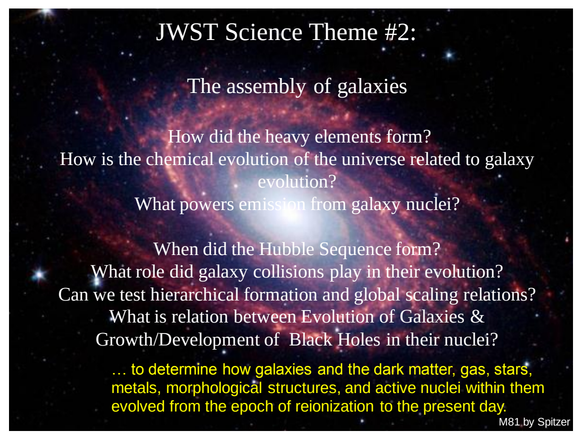
This is a surprising result unexpected in current galaxy formation models.

Michael Werner, "Spitzer Space Telescope", William H. Pickering Lecture, AIAA Space 2007.

## JWST – the First Light Machine

With its 6X larger collecting aperture, JWST will see back in time further than Hubble and explore the Universe's first light.




## Formation of Heavy Elements

Carl Sagan said that we are all 'star dust'.

All of the heavy elements which exist in the universe were formed from Hydrogen inside of stars and distributed via supernova explosions. But observations in the visible couldn't find enough dust.

Dust is cold, therefore, it can only be seen in IR.

Looking in the IR (with Herschel and Spitzer) at Supernova 1987A, 100,000X more dust was seen than in the visible – the total mass of this dust equals about half of our Sun.

Image of Supernova 1987A, taken in the infrared by Herschel and Spitzer, shows some of the warm dust surrounding it.
CREDIT: Pasquale Panuzzo
SPACE.com, Taylor Redd, 7 July 2011

## 2nd Generation Stars – 700M yrs after BB

This star is a 2nd generation star after the big bang because it has trace amounts of heavy elements – meaning that at least one supernova had exploded before it was formed.

But its existence contradicts current theories because it has too much Hydrogen and too much Helium and not enough Carbon and other heavy elements.


Nola Taylor Redd, SPACE.com, 31 August 2011; CREDIT: ESO/Digitized Sky Survey 2

## Chemical make-up of Early Universe

1.8 B yr after BB gamma-ray burst illuminates neighboring galaxies yielding spectra of their chemical makeup.

Metals in the early universe are higher than expected – indicating that star formation in the early universe was much higher than current theory.

GRB 090323 was first detected on 23 March 2009 by NASA's Fermi space telescope and then the Swift satellite, shortly followed by the ground-based GROND system (Gamma-Ray burst Optical and Near-infrared Detector) at the MPG/ESO 2.2-metre telescope in Chile, as well as ESO's Very Large Telescope (VLT). The VLT observations revealed that the gamma-ray burst injected light through its host galaxy and another nearby galaxy, which are both seen at a redshift of 3.57, equivalent to 12 billion years ago.

DR EMILY BALDWIN, ASTRONOMY NOW, 02 November 2011

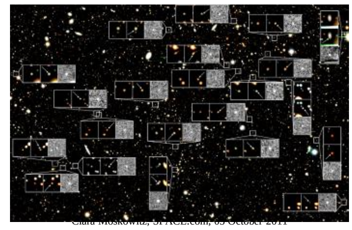
## Subaru Deep Field: Ancient Supernova 3.7B yrs after BB

22 of 150 ancient supernovae in 10% of Subaru Deep Field

12 occurred around 3.7B yrs after big bang.

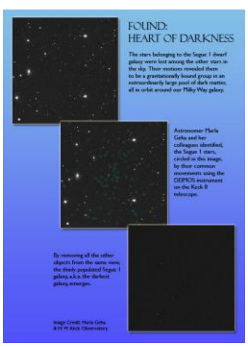
Supernova were 10X more frequent at this time than today.

Supernova helped seed early universe with chemical elements.



Clara Moskowitz, SPACE.com, 05 October 2011

## **Dark Matter**




SPACE.com, 02 August 2011

Dwarf Galaxy Segue 1 has 1000 small, dim, primordial (ancient) low metalicity stars.

But, based on star motion, it has 3400X more mass than can be observed.

Some stars are moving too fast, the only thing keeping the galaxy together is gravity.

Thus, there is Dark Matter.

## Dark Matter Distribution

Current Theory says that to hold galaxies together, Dark Matter should be 'clumped' in a central bulge.

But, observations of two dwarf galaxies, Fornax and Sculptor (which are 99% dark matter), show that the dark matter within them is spread out smoothly.

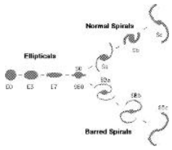
It is possible that dark matter might interact more with ordinary matter than currently thought, allowing the regular matter to stir up the dark matter and spread it out. Alternatively, dark matter might move faster than expected and therefore be less prone to clumping in galactic centers.

Image: ESO/Digitized Sky Survey 2

Adam Mann October 17, 2011

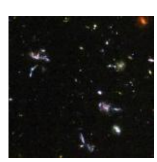
## **The Hubble Sequence**

Hubble classified nearby (present-day) galaxies into Spirals and Ellipticals.




The Hubble Space Telescope has extended this to the distant past.

## Where and when did the Hubble Sequence form? How did the heavy elements form?

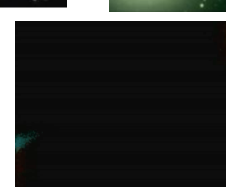


Galaxy assembly is a process of hierarchical merging

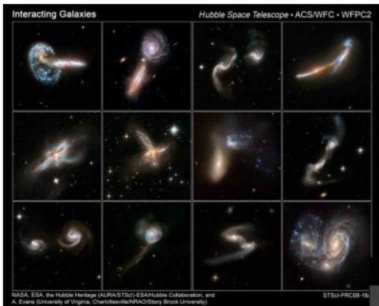
Components of galaxies have variety of ages & compositions

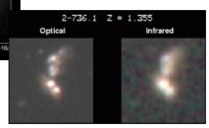
JWST Observations:

- Wide-area near-infrared imaging survey
- Low and medium resolution spectra of 1000s of galaxies at high redshift
- Targeted observations of galactic nuclei



## Distant Galaxies are "Train Wrecks"




## Merging Galaxies = Merging Black Holes

Combined Chandra & Hubble data shows two black holes (one 30M & one 1M solar mass) orbiting each other – separated by 490 light-years. At 160 million light-years, these are the closest super massive black holes to Earth.

Theory says when galaxies collide there should be major disruption and new star formation.

This galaxy has regular spiral shape and the core is mostly old stars.

These two galaxies merged with minor perturbations.

Galaxy NGC3393 includes two active black holes
X-ray: NASA/CXC/SAO/G.Fabbiano et al; Optical: NASA/STScI

Charles Q. Choi, SPACE.com, 31 August 2011

## Galaxy Clusters – 2.6B yrs after BB

Galaxy clusters are the largest structures in the universe. Bound together by gravity, they require billions of years to form.

Galaxy Clusters are thought to have started to form around 3 B-yrs after big bang.

At 2.6 B-yrs old, this is not the oldest observed galaxy cluster. But, spectra indicates that stars in its constituent galaxies are 1 B-yrs old. Thus, may have started forming about 1.5 B-yrs after BB.

X-ray data (similar to image) shows glow from cloud of very hot gas that holds cluster together. Again, it takes many years to trap hot gas.

Hubble NIR Image of CL J1449+0856, the most distant mature cluster of galaxies found. Color added from ESO's VLT and NAOJ's Subaru Telescope. CREDIT: NASA, ESA, R. Gobat (SPACE.com 09 March 2011 )

JKCS 041 at 3.7 B-yr after BB may be one of the Universe's oldest clusters. In Chandra image, X-ray emission is shown in blue. Image: NASA/CXC/INAF/S.Andreon (Astronomy Now, 10 May 2010)

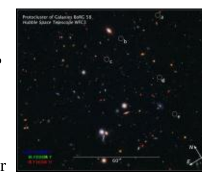
## Galaxy Formation – 0.6 B yrs after BB

The early universe was smooth and without structure. Clumping began small and grew to form large galaxies. But how and when?

At 600 Myrs after big bang, these 5 tiny galaxies (circled) are the youngest galaxy cluster yet observed.

They ranging from 10% to 50% the size of our own Milky Way. But they're about as bright as the Milky Way, because they're feasting on huge amounts of gas via mergers with other galaxies.

Since this time, they may have merged to form a giant galaxy.



Borg 58 galaxy field: composite image taken in visible and near-infrared light, reveals the location of five tiny galaxies clustered together 13.1 billion light-years away. The circles pinpoint the galaxies.

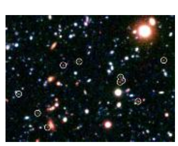
(Space.com 10 Jan 2012)

## Galaxy Formation – 1.1B yrs after BB

Previous oldest cluster is 1.1 B-yrs after BB.

Cluster contains 11 min-galaxies which are all much smaller than the Milky Way. One has a 30 million solar mass black hole.

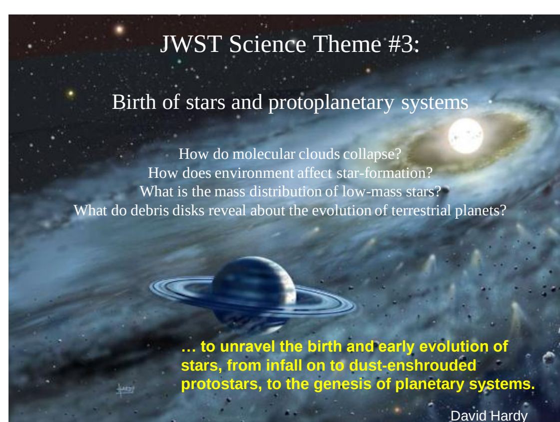
These too may have merged to form a galaxy.



Cluster COSMOS-AzTEC3, located in the Sextans, contains 11 mini-galaxies (circled red dots). Cluster is 1.1 billion yrs after Big Bang.
*Subaru / NASA / JPL-Caltech*

Discovery required observations from: Chandra X-ray, James Clerk Maxwell Sub-MM, Hubble, Subaru, Keck, Spitzer & several Radio Telescopes

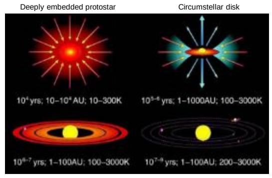
(Sky and Telescope, Robert Naeye, 13 Jan 2011)



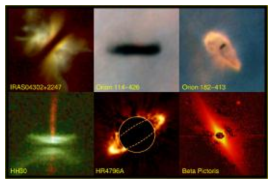
## Birth of Stars and Proto-planetary Systems

- What is the role of molecular clouds, cores and their collapse in the evolution of stars and planetary systems?
- How do protostars form and evolve?
- How do massive stars form and interact with their environment?
- How do massive stars impact their environment by halting or triggering further star formation. How do they impact the evolution of disks?
- What is the initial mass function down to planetary masses?
- How do protoplanetary systems form and evolve?
- How do astrochemical tracers track star formation and the evolution of protoplanetary systems?

Deeply embedded protostar — Circumstellar disk



Agglomeration & planetesimals — Mature planetary system



## How do proto-stellar clouds collapse?

Stars form in small regions collapsing gravitationally within larger molecular clouds.

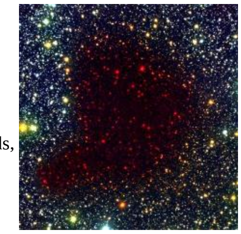
Infrared sees through thick, dusty clouds

Proto-stars begin to shine within the clouds, revealing temperature and density structure.

Key JWST Enabling Requirements:

High angular resolution near- & mid-IR imagery

High angular resolution imaging spectroscopy



Barnard 68 in infrared

## How does environment affect star-formation?

Massive stars produce wind & radiation

Either disrupt star formation, or causes it.

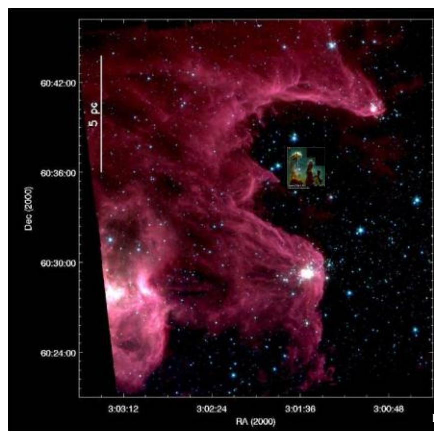
Boundary between smallest brown dwarf stars & planets is unknown

Different processes? Or continuum?

JWST Observations:

Survey dark clouds, "elephant trunks" or "pillars of creation" star-forming regions

The Eagle Nebula
as seen in the infrared



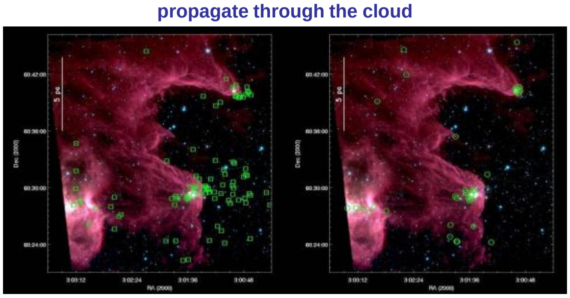
Spitzer has Found "The Mountains Of Creation"

Michael Werner, "Spitzer Space Telescope", William H. Pickering Lecture, AIAA Space 2007.

L. Allen, CfA [GTO]

## The Mountains Tell Their Tale

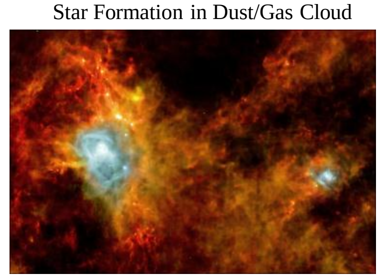
**Interstellar erosion & star formation propagate through the cloud**



**Young (Solar Mass) Stars are Shown in This Panel** — **Really Young Stars are Shown in This Panel**

Michael Werner, "Spitzer Space Telescope", William H. Pickering Lecture, AIAA Space 2007.

## Star Formation in Dust/Gas Cloud



Herschel discovered 700 newly-forming stars condensing along filaments of dust in a never before penetrated dark cloud at the heart of Eagle Nebula.

Two areas glowing brightest in icy blue light are regions where large newborn stars are causing hydrogen gas to shine.

SPACE.com 16 December 2009

## Impossible Stars

100 to 150 solar mass stars should not exist but they do.

When a star gets to 8 to 10 solar mass its wind blows away all gas and dust, creating a bubble and stopping its growth (see Herschel Image).

The bubble shock wave is creating a dense 2000 solar mass region in which an 'impossible' star is forming. It is already 10 solar mass and in a few 100 thousand years will be a massive 100 to 150 solar mass – making it one of the biggest and brightest in the galaxy.

(Space.com, 6 May 2010)

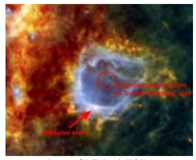
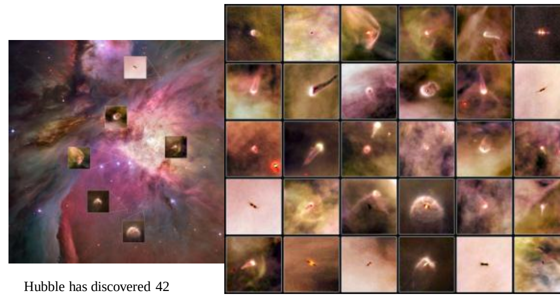


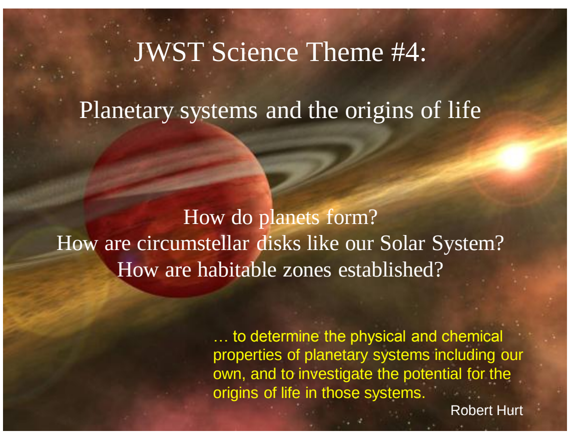
Image of RCW 120 (ESA), Discover.com, Ian O'Neill, 7 May 2010

## Orion Nebula Protoplanetary Discs



Hubble has discovered 42 protoplanetary discs in the Orion Nebula

Credit: NASA/ESA and L. Ricci (ESO)




## Planetary Formation Questions and 2 Models

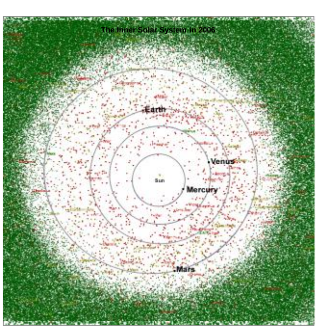
- How do planets and brown dwarfs form?
- How common are giant planets and what is their distribution of orbits?
- How do giant planets affect the formation of terrestrial planets?
- What comparisons, direct or indirect, can be made between our Solar System and circumstellar disks (forming solar systems) and remnant disks?
- What is the source of water and organics for planets in habitable zones?
- How are systems cleared of small bodies?
- What are the planetary evolutionary pathways by which habitability is established or lost?
- Does our solar system harbor evidence for steps on these pathways?


## History of Known (current) NEO Population

2006

Earth Crossing

Outside Earth's Orbit

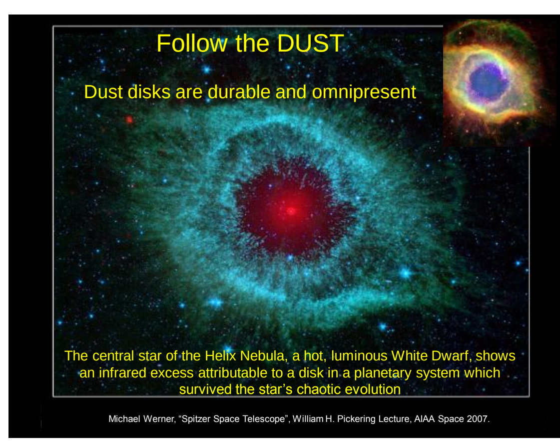



Known
- 340,000 minor planets
- ~4500 NEOs
- ~850 Potentially Hazardous Objects (PHOs)

Scott Manley

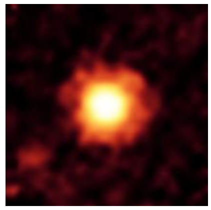
Armagh Observatory

Landis, "Piloted Flight to a Near-Earth Object", AIAA Conference 19 Sep 07



## Planetary System Formation effects Dust

This star has 3 large (10X Jupiter mass) planets (observed by Hubble, Keck & Gemini North) which are causing a huge halo of fine dust particles (indicating lots of colliding objects) around the star. Dust which can be detected by an infrared telescope.

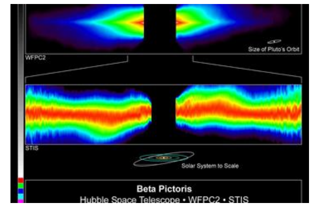


NASA's Spitzer Space Telescope captured this infrared image of a giant halo of very fine dust around the young star HR 8799, located 129 light-years away in the constellation Pegasus. The brightest parts of this dust cloud (yellow-white) likely come from the outer cold disk similar to our own Kuiper belt (beyond Neptune's orbit). The huge extended dust halo is seen as orange-red. Credit: NASA/JPL-Caltech/Univ. of Ariz.

Astrophysical Journal, Nov 2009

## Planetary System Formation effects Dust

'Kinks' in the debris disk around Beta Pictoris was caused by the formation and subsequent migration of a Jupiter-sized planet called Beta Pictoris b.

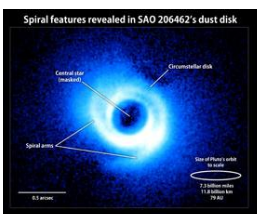



The planet orbiting Beta Pictoris has caused a kink in the debris disk surrounding the star, as seen in this false-color image from the Hubble Space Telescope. CREDIT: Sally Heap (GSFC/NASA)/ Al Schultz (CSC/STScI, and NASA)

Nola Taylor Redd, SPACE.com; 08 December 2011

## Spiral Arms Hint At The Presence Of Planets

Disk of gas and dust around a sun-like star has spiral-arm-like structures. These features may provide clues to the presence of embedded but as-yet-unseen planets.

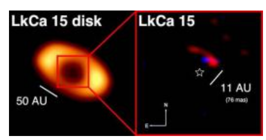



Near Infrared image from Subaru Telescope shows disk surrounding SAO 206462, a star located about 456 light-years away in the constellation Lupus. Astronomers estimate that the system is only about 9 million years old. The gas-rich disk spans some 14 billion miles, which is more than twice the size of Pluto's orbit in our own solar system.

Photonics Online 20 Oct 2011

## Direct Image of an ExoPlanet being Formed

Image shows the youngest exoplanet yet discovered. Its Star (slightly smaller than our Sun) is only 2 million years old. Dust is accreting (falling) into the new planet leaving a gap in the planetary disk. New planet is ~ 6X mass of Jupiter.




Using the Keck Telescope

Left: The dusty disk around the star LkCa 15. All of the light at this wavelength is emitted by cold dust in the disk; the hole in the center indicates an inner gap.

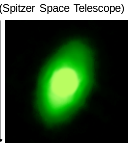
Right: An expanded view of the central part of the cleared region, showing a composite of two reconstructed images (blue: 2.1 microns; red: 3.7 microns) for LkCa 15 b. The location of the central star is also marked.

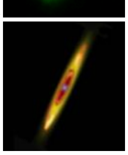
CREDIT: Kraus & Ireland 2011
SPACE.com; 19 October 2011

## Dust in Planetary Systems

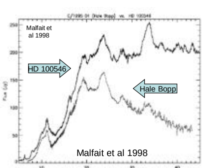
Fomalhaut system at 24 µm
(Spitzer Space Telescope)

72"

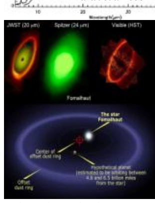




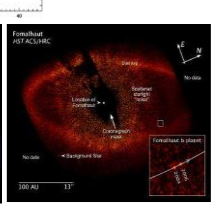
Simulated JWST image Fomalhaut at 24 microns




Malfait et al 1998



Kalas, Graham & Clampin 2005

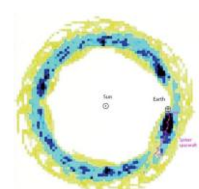


Kalas et al 2008

## Planets and Dust

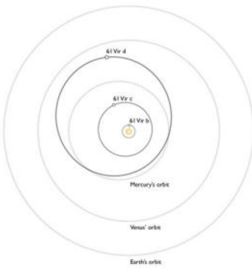
Earth has a 'tail' of dust particles.

10 to 20 micrometer size particles are slowed or captured by Earth's gravity and trail behind Earth. The cloud of particles is about 10 million km wide and 40 million km long.



(Wired.com, Lisa Grossman, 8 July 2010)

## Radial Velocity Method finds planets close to stars




61 Virginis (61 Vir) has 3 planets inside of Venus's orbit.

From their star, the planets have masses of ~5X, 18X & 24X Earth's mass.

They orbit 61 Virginis in 4, 38 & 124 day periods.

Also,direct Spitzer observations indicate a ring of dust at twice the distance of Neptune from the star.

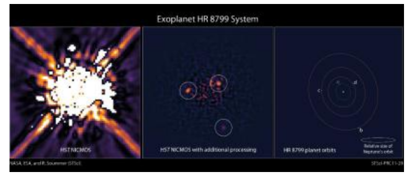
Bad Astronomy
Orbital schematic credit: Chris Tinney

## Direct Imaging detects planets far from their star

HR 8799 has at least 4 planets

3 planets ('c' has Neptune orbit) were first imaged by Hubble in 1998. Image reanalyzed because of a 2007 Keck discovery.

3 outer planets have very long orbits or 100, 200 & 400 years. Multiple detections are required to see this motion.

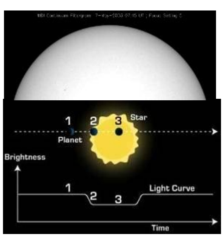



Denise Chow, SPACE.com; 06 October 2011

## Transit Method Finds Planets

Kepler (launched in 2009) is hunting planets by staring continuously at 165,000 stars looking for dips in their light caused when a planet crosses in front of the star.

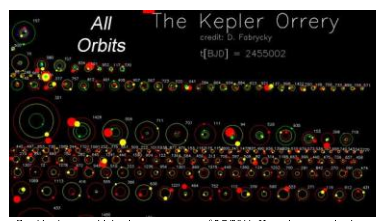
As of Dec 2011, Kepler has found 2326 planets



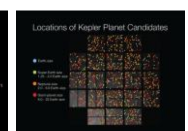

## Kepler Planetary Systems

Of the 2326 planets which Kepler has discovered (Dec 11):

- > 800 in single planet systems,
- > 400 in 170 systems with 2 to 6 transiting planets, and
- 207 potential Earth size; 680 super-Earth size; 48 in Habitable Zone



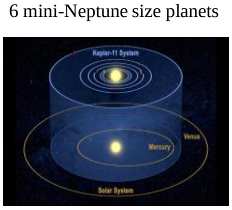

Graphic shows multiple-planet systems as of 2/2/2011. Hot colors to cool colors (red to yellow to green to cyan to blue to gray) indicate big planets to smaller planets. CREDIT: Daniel Fabrycky (SPACE.com, 23 May 2011)




Kepler's planet candidates by size. CREDIT: NASA/Wendy Stenzel (SPACE.com 2 Feb 2011)

## Kepler Mission

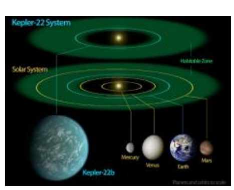
Kepler-11has a star like ours & 6 mini-Neptune size planets




Five of six Kepler-11 exoplanets (all larger than Earth) orbit their star closer than Mercury orbits the sun. One orbits inside Venus.

Credit: NASA/AP (Pete Spotts, Christian Science Monitor.com, 23 May 2011.)

Kepler 22b is the first in the habitable zone.




Kepler-22b is located about 600 light-years away, orbiting a sun-like star. Its is 2.4 times that of Earth, and the two planets have roughly similar temperatures (maybe 22C).

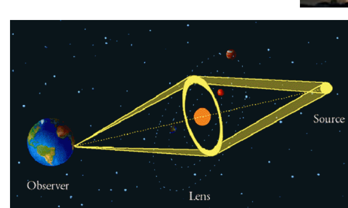
CREDIT: NASA/Ames/JPL-Caltech

## Micro-Lensing Finds Planets

On average every Star has 1.6 planets.

From 2002 to 2007

- 3,247 micro-lensing events were detected
- 440 light curves studied




Gravitational microlensing method requires that you have two stars that lie on a straight line in relation to us here on Earth. Then the light from the background star is amplified by the gravity of the foreground star, which thus acts as a magnifying glass. CREDIT: Nancy Atkinson (January 11, 2012)

## Is There Life Elsewhere in the Galaxy?

Need to multiply these values by $\eta_{Earth}$ x $f_B$ to get the number of potentially life-bearing planets detected by a space telescope.
$\eta_{Earth}$ = fraction of stars with Earth-mass planets in HZ
$f_B$ = fraction of the Earth-mass planets that have detectable biosignatures

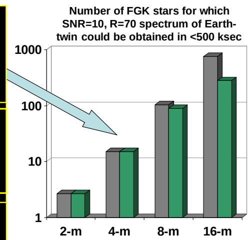
Earth-mass planets within these HZ will be very

If: $\eta_{Earth} \times f_B \sim 1$ then $D_{Tel} \sim 4m$
$\eta_{Earth} \times f_B < 1$ then $D_{Tel} \sim 8m$
$\eta_{Earth} \times f_B << 1$ then $D_{Tel} \sim 16m$

Number of nearby stars capable of hosting

Kepler is finding that $\eta_{Earth}$ maybe 1.5% to 2.5% (SPACE.COM, 21 Mar 2011)

Thus, an 8-m telescope might find 1 to 3 Earth twins and an 16-m telescope might find 10 to 20 Earth twins.




Green bars show the number of FGK stars that could be observed 3x each in a 5-year mission without exceeding 20% of total observing time available to community.

Marc Postman, "ATLAST", Barcelona, 2009; Modified by Stahl, 2011

## How are habitable zones established?

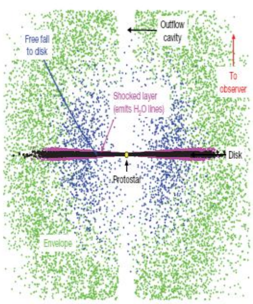
Source of Earth's $H_20$ and organics is not known
Comets? Asteroids?

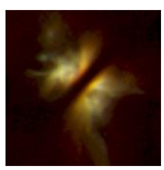
History of clearing the disk of gas and small bodies
Role of giant planets?

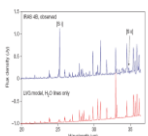
JWST Observations:
Comets, Kuiper Belt Objects
Icy moons in outer solar system



## Spitzer Spectrum Shows Water Vapor Falling onto Protoplanetary Disk





Michael Werner, "Spitzer Space Telescope", William H. Pickering Lecture, AIAA Space 2007.

## Proto-Stars produce Water

In a proto-star 750 light-years away, Herschel detected:

Spectra of Atomic Hydrogen and Oxygen are being pulled into the star, and

Water vapor being spewed at 200,000 km per hour from the poles.

The water vapor freezes and falls back onto the proto-planetary disk.

Discovery is because Herschel's infrared sensors can pierce the dense cloud of gas and dust feeding the star's formation.

(National Geographic, Clay Dillow, 16 June 2011)

A Protostar and its Polar Jets NASA/Caltech

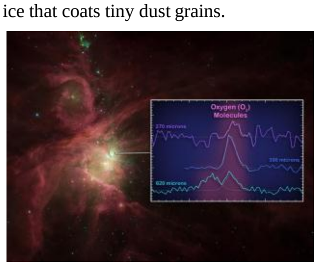
Other Herschel Data finds enough water in the outer reaches of the young star TW Hydrae (175 light-yrs from Earth) to fill Earth's oceans several thousand times over.

Mike Wall, SPACE.com; Date: 20 October 2011

## Molecular Oxygen discovered in space

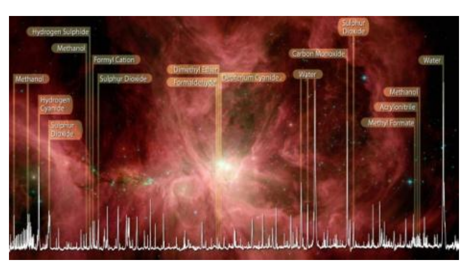
Herschel found molecular oxygen in a dense patch of gas and dust adjacent to star-forming regions in the Orion nebula.

The oxygen maybe water ice that coats tiny dust grains.




SPACE.com, 01 August 2011

## All of Life's Ingredients Found in Orion Nebula

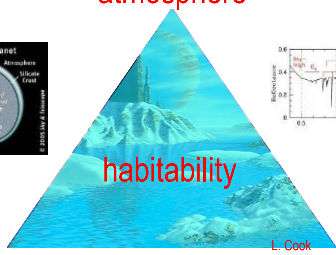


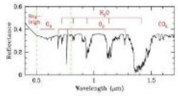

Herschel Telescope has measured spectra for all the ingredients for life as we know them in the Orion Nebula.
(Methanol is a particularly important molecule)

Wired.com Mar 2010

## Search for Habitable Planets

atmosphere


interior

surface

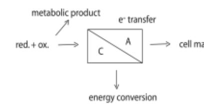
Sara Seager (2006)

## Search for Life

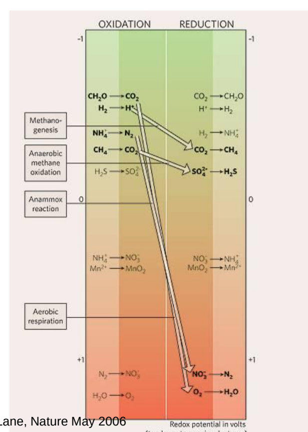
What is life?

What does life do?

Life Metabolizes




Sara Seager (2006)




All Earth life uses chemical energy generated from redox reactions

Life takes advantage of these spontaneous reactions that are kinetically inhibited

Diversity of metabolisms rivals diversity of exoplanets

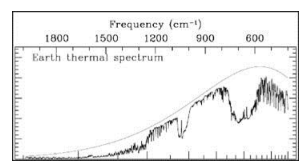
Lane, Nature May 2006

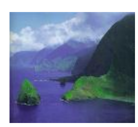
Sara Seager (2006)

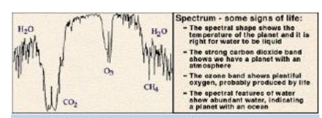
## Bio Markers

Spectroscopic Indicators of Life

Absorption Lines
- CO2
- Ozone
- Water
- “Red” Edge

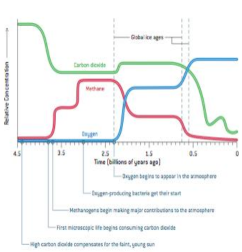


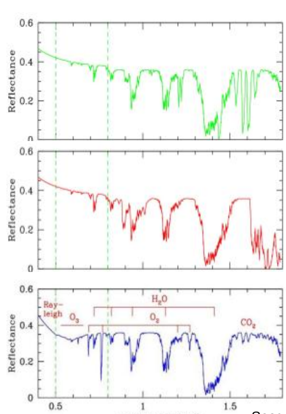






Example signs of life from chemical spectra.
*Credit: NASA JPL*

## Earth Through Time




Kasting Sci. Am. 2004
See Kaltenegger et al. 2006
Earth from the Moon

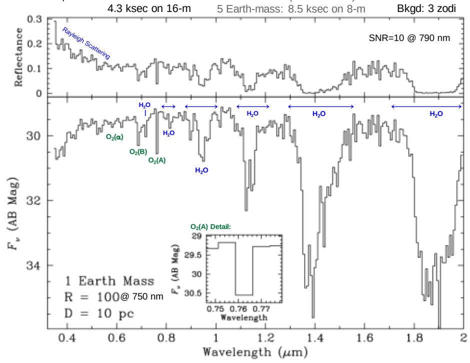
Seager

## Beyond JWST

Heavy Lift Launch Vehicle enables even larger telescopes
- 8-m UV/Optical Telescope or
- 24-m Far-IR Telescope

## R=100 ATLAST Spectrum of 1 Earth-mass Terrestrial Exoplanet at 10 pc

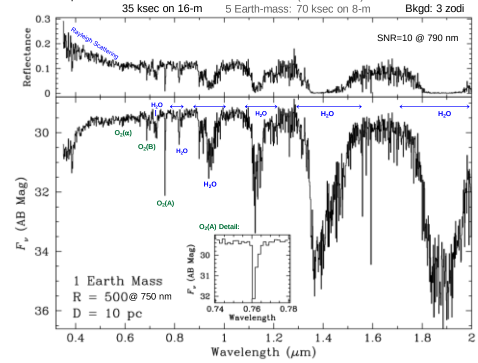
Exposure: 51 ksec on 8-m
4.3 ksec on 16-m
Reflectance ∝ (Planet Mass)$^{2/3}$
5 Earth-mass: 8.5 ksec on 8-m
Bkgd: 3 zodi




Marc Postman, "ATLAST", Barcelona, 2009

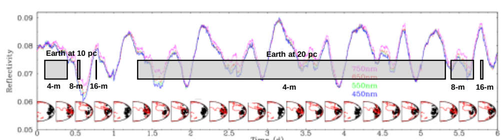
## R=500 ATLAST Spectrum of 1 Earth-mass Terrestrial Exoplanet at 10 pc

Exposure: 502 ksec on 8-m
35 ksec on 16-m
Reflectance ∝ (Planet Mass)$^{2/3}$
5 Earth-mass: 70 ksec on 8-m
Bkgd: 3 zodi




Marc Postman, "ATLAST", Barcelona, 2009

## Detecting Photometric Variability in Exoplanets



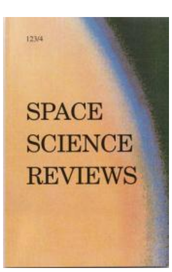

Marc Postman, "ATLAST", Barcelona, 2009

## Countdown to Launch

JWST is

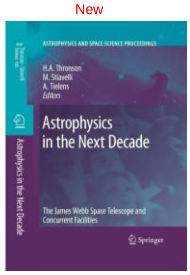
- making excellent technical progress
- will be ready for launch ~2017-2019
- will be the dominant astronomical facility for a decade undertaking a broad range of scientific investigations


## Learn more at: www.jwst.nasa.gov




Download for free at: jwst.gsfc.nasa.gov

New

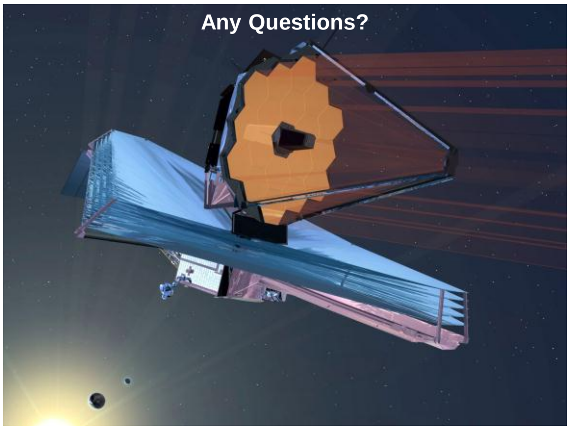



Available Now

New

www.stsci.edu/jwst/science/whitepapers/

## 1000s of Scientists and Engineers in USA and around the world are working to make JWST.


Any Questions?