Transitional Disks Associated With Intermediate-mass Stars: Results of the SEEDS YSO survey

Talk Roadmap

1. IR SEDs for Intermediate-mass stars
2. The SEEDS Survey –
3. Transitional Disks Imagery and Interferometry
4. The path forward
The SEEDS YSO Survey

- 1-10 Myr old objects, both single and binary/multiple stars

- 210 objects which are to be searched for both disks and the presence of planets in/near the disks.


- Focus in this talk on disks associated with stars of spectral types B-G.

- IR SEDs for these stars come in 2 flavors from 1-200 μm: power law only (Meeus et al. 2001, Group II) and powerlaw + blackbody (Meeus Group I)
2. IR SEDS– Meeus Group II

SED credit: M. Sitko

Kusakabe et al. 12

H band PI, Kusakabe et al. 2012
IR SEDS- Meeus Group I

Acke & van den Ancker 04

HD 135344
Group I

Gap in Sub-mm Data

Muto et al. 2012 H PI, Garufi et al. 2013 Ks PI
Stars in the Survey

- HD 179218 (ND), HD 100546 (NaCo), HD 97048 (NaCo), AB Aur, Oph IRS 48, HD 34282
- MWC 758, HD 169142 (HiCIAO and NaCo), SAO 206462 (HiCIAO and NaCo), CQ Tau (ND)
- HD 142527 (HST NICMOS, NIRI, NaCo, HiCIAO), LkHa 330, SR 21
Meeus I disks are diverse
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- AB Aur
  - Hashimoto+ 10

- Oph IRS 48
  - Follette+ 14

- HD 34282
  - Maruta + 14
  - Grady + 13

- MWC 758
  - Fukagawa+14

- HD 169142
  - Momose+13

- SAO 206462
  - Muto+12

- HD 142527
  - Fukagawa+14

- SR 21
  - Follette + 13

- LkHa 330
  - Bonnefoy+ 14
Are Meeus I Disks the missing Transitional Disks?

* 66% of the HiCIAO sample have inner cavities seen in sub-millimeter interferometry

* 55% have gaps seen in the NIR either with HiCIAO or VLT/NaCo (smaller IWA)

* Gap non-detections include the more distant systems – NIR detection rate is clearly sensitive to IWA and proximity of the disk.

* Conclude that the majority of the well-studied Herbig stars with Meeus I SEDS are transitional disks – same conclusion reached by Maaskant et al. (2013).
Statistics for Meeus I disks

- 9 disks in HiCIAO sample, 11 if include VLT/NACO imagery in the literature; scattered light detection rate ~85% detection rate
- Featureless disks: up to 3 – 25% of sample – may reflect Strehl ratio of data... or distance to star
- Gapped disks visible in NIR – 44% of HiCIAO sample, rises with smaller IWA
- Broken powerlaw systems: 0 – these would be hard to find, given other structure
- 45% have spiral arms/features at some wavelengths
- 35% of the HiCIAO sample have eccentric gaps
Statistics (con’t)

- Disk detection rate in scattered light is high compared with Meeus II disks (hit or miss with HST and NIR ground-based surveys)

- Diverse grain properties: Some disks can be fit with MCRT codes using compact grains (HD 142527), others require porous grain aggregates (Oph IRS 48), and some can be modeled using either.

- 2 flavors of spiral – some dominated by ISM-like grains (AB Aur) and extend beyond mm disk – envelope material, others seen throughout NIR and clearly associated with disks.
Why and When are we seeing structured disks?

- Andrews et al. (2011 - same model for T Tauri and Hae transitional disks - Hwall/R_wall

- Expect loosely wrapped arms for dynamically warm disks – easier to image

- Some of our non-detections (SR 21) have large H/R and were obtained at low Strehl ratio.

- Some of the non-detections are for systems viewed at 50-55° inclination – projection effect?
Gap Properties for Meeus I disks

- All disks are pre-transitional
- Gap size is uncorrelated with FUV radiation field.
- Uncorrelated with system age
- Dynamical origins are suggested by the fact that one of these systems hosts 2 exoplanet candidates.
Potential of Extreme AO

- SCExAO - smaller IWA – 4x better than HiCIAO, 2x better than NACO - image in to 7AU at d=140pc
- More stable PSF – may be able to get I images to go with PI data from HiCIAO – probe of grain properties
- Classification of disks can change with IWA – case in point is SAO 206462 – would classify as a filled disk with spiral arms from HiCIAO, but NaCo data clearly demonstrate a cavity which is smaller than seen in the sub-mm.

HST NICMOS

Muto et al. 2012  Garufi et al. 2013
Synergy between scattered light imagery and ALMA

- Band 9 ALMA continuum for SAO206462 has a gapped ring, similar to SMA, but places the gap and the flux maximum at different locations in the disk.
- SE “wall” coincides with SE arm.
- Gapped ring morphology appears typical for the transitional disks in the sub-mm continuum.
- Can expect ALMA data to evolve as interferometer comes fully on-line.

Perez et al. 2012
Conclusions

* Structure in the form of spiral arms, gaps, etc. extremely common in Intermediate-mass transitional disks. Easier to see in dynamically warm disks, and can be washed out if the image quality is low.

* Gap size and visibility in the scattered light imagery does not correlate with stellar properties (mass, luminosity). Excludes photoevaporation, and suggests an independent process, such as exoplanet formation.

* Expect disks around lower-mass stars to show structure as the IWA, Strehl ratio, and contrast of the imaging data improve.

* Next phase of study will be searches for indirect signatures of exoplanets in these disks such as rotation of planet-induced spiral arms, and direct imaging searches for the exoplanets and/or their accretion luminosity.

* Please See Posters: Baille, Follette, Garufi, Hanawa, Kanagawa, Matter, Takahashi, S., Tsukamoto, Wisniewski, and Yasui