

SPECTRAL COMPARISON AND STABILITY OF RED REGIONS ON JUPITER

A. A. Simon (NASA Goddard Space Flight Center), R. W. Carlson (JPL/Caltech), and A. Sanchez-Lavega (Universidad del País Vasco)

Summary: A study of absolute color on Jupiter from Hubble Space Telescope imaging data shows that the Great Red Spot (GRS) is not the “reddest” region of the planet. Rather, a transient red cyclone visible in 1995 and the North Equatorial Belt both show redder spectra than the GRS (*i.e.*, more absorption at blue and green wavelengths).

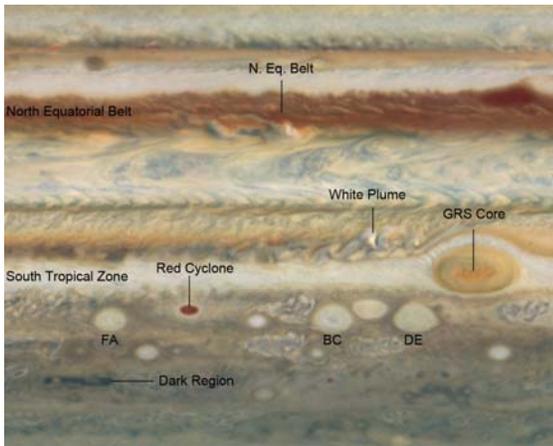


Figure 1. Quasi-true color map of Jupiter from HST, February 1995.

Spectral Variation: Spectral data acquired over multiple epochs and many wavelengths show that most regions on Jupiter are remarkably stable in color; temporal analysis indicates that the darkest regions of the NEB are relative constant in color from 1995 to 2008, while the slope of the GRS core may vary slightly. No region is truly black or white, but all show some absorption at UV and blue wavelengths, see Fig. 2. The most intense coloration is not in the Great Red Spot, but in the North Equatorial Belt, particularly dark cyclonic barges, and in rare transient cyclonic vortices at other latitudes, such as the 1995 Red Cyclone. Regions with quite different cloud structure can show similar color, while regions with similar structure can vary in coloration.

Composition: Multiple factors (variable composition or particle processing history) are likely involved in producing color; the best spectral match seems to require the presence of both NH_4SH and $\text{C}_2\text{H}_2\text{-NH}_3$ interactions, while various levels of irradiation or thermal processing may also explain some of the color variation, see Fig. 3. In addition, there is slight evidence that P_4 or a similar phosphorus-bearing compound is necessary to explain spectral variations at wavelengths beyond 600 nm; further spectral data are required, however.

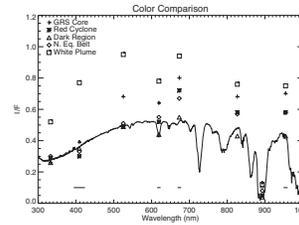


Figure 2. Absolute I/F Spectra

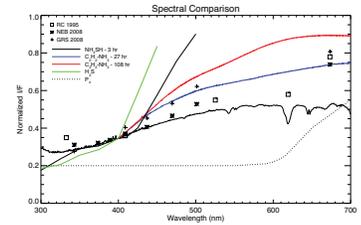


Figure 3. Laboratory Spectra

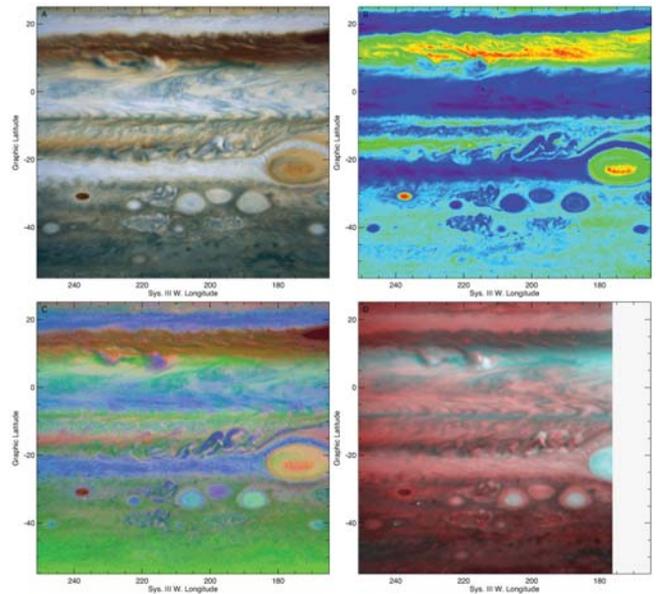


Figure 4. Variations between red regions are evident in A) quasi-true color, B) ratio of 673/410 nm, C) principal components, D) methane band/cloud height

References: [1] Peek, *The Planet Jupiter*, 1958; [2] Rogers, *The Giant Planet Jupiter*, 1995; [3] Morales-Juberias et al., *Icarus* **160**, 325–335, 2002; [4] Brooke et al., *Icarus* **136**, 1-13, 1998; [5] Simon-Miller et al., *Icarus* **145**, 454-461, 2000; [6] Baines et al., *Icarus* **159**, 74-94, 2002; [7] Wong et al., *Plan. and Space Sci.* **52**, 385-395, 2004; [8] Atreya et al., *Plan. and Space Sci.* **53**, 498-507, 2005; [9] Kalogerakis et al., *Icarus* **196**, 202-215, 2008; [10] Sromovsky and Fry, *Icarus* **210**, 211–229, 2010; [11] Sromovsky and Fry, *Icarus* **210**, 230–257, 2010.

Acknowledgments: This work was based on observations made with the NASA/ESA Hubble Space Telescope, obtained from the Data Archive at the Space Telescope Science Institute, which is operated by the Association of Universities for Research in Astronomy, Inc., under NASA contract NAS 5-26555. These observations are associated with programs GO5313, GO5642, GO6009, GO6141, GO6452, GO11096, and GO11498. Sanchez-Lavega was supported by Spanish MICIIN project AYA2012-36666, Grupos Gobierno Vasco IT765-13 and UPV/EHU UFI11/55.