Nuclear Thermal Propulsion (NTP): A Proven, Growth Technology for “Fast Transit” Human Missions to Mars

Stanley K. Borowski
Glenn Research Center, Cleveland, Ohio

David R. McCurdy and Thomas W. Packard
Vantage Partners, LLC, Brook Park, Ohio
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National Aeronautics and
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Glenn Research Center
Cleveland, Ohio 44135

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National Aeronautics and Space Administration
Glenn Research Center
Cleveland, Ohio 44135

David R. McCurdy and Thomas W. Packard
Vantage Partners, LLC
Brook Park, Ohio 44142

Abstract
The “fast conjunction” long surface stay mission option was selected for NASA’s recent Mars Design Reference Architecture (DRA) 5.0 study because it provided adequate time at Mars (~540 days) for the crew to explore the planet’s geological diversity while also reducing the “1-way” transit times to and from Mars to ~6 months. Short transit times are desirable in order to reduce the debilitating physiological effects on the human body that can result from prolonged exposure to the zero-gravity (0-gE) and radiation environments of space. Recent measurements from the RAD detector attached to the Curiosity rover indicate that astronauts would receive a radiation dose of ~0.66 Sv (~66 rem)—the limiting value established by NASA—during their 1-year journey in deep space. Proven nuclear thermal rocket (NTR) technology, with its high thrust and high specific impulse ($I_{sp} \approx 900$ s), can cut 1-way transit times by as much as 50 percent by increasing the propellant capacity of the Mars transfer vehicle (MTV). No large technology scale-ups in engine size are required for these short transit missions either since the smallest engine tested during the Rover program—the 25 klbf “Pewee” engine is sufficient when used in a clustered arrangement of three to four engines. The “Copernicus” crewed MTV developed for DRA 5.0 is a 0-gE design consisting of three basic components: (1) the NTP stage (NTPS); (2) the crewed payload element; and (3) an integrated “saddle truss” and LH$_2$ propellant drop tank assembly that connects the two elements. With a propellant capacity of ~190 t, Copernicus can support 1-way transit times ranging from ~150 to 220 days over the 15-year synodic cycle. The paper examines the impact on vehicle design of decreasing transit times for the 2033 mission opportunity. With a fourth “upgraded” SLS/HLV launch, an “in-line” LH$_2$ tank element can be added to Copernicus allowing 1-way transit times of 130 days. To achieve 100 to 120 day transit times, Copernicus’ saddle truss/drop tank assembly is replaced by a “star truss” assembly with paired modular drop tanks to further increase the vehicle’s propellant capacity. The HLV launch count increases (from ~5 to 7) and a fourth engine is needed to reduce total mission burn time and gravity losses. Using a “split mission” approach, the NTPS, in-line tank and the saddle truss/LH$_2$ drop tank elements can be configured as a pre-deployed Earth Return Vehicle/propellant tanker supporting 90-day crewed mission transits. The split mission approach also eliminates the need for on-orbit assembly. Mission scenario descriptions, key features and operational characteristics for five different vehicle configurations are presented.

Nomenclature

<table>
<thead>
<tr>
<th>Symbol</th>
<th>Description</th>
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<tbody>
<tr>
<td>IMLEO</td>
<td>initial mass in low Earth orbit</td>
</tr>
<tr>
<td>K</td>
<td>temperature (degrees Kelvin)</td>
</tr>
<tr>
<td>klbf</td>
<td>thrust (1000’s of pounds force)</td>
</tr>
<tr>
<td>LEO</td>
<td>low Earth orbit (= 407 km circular)</td>
</tr>
<tr>
<td>LOX/LH$_2$</td>
<td>Liquid oxygen/liquid hydrogen propellant</td>
</tr>
</tbody>
</table>
RAD Radiation Assessment Detector carried on Mars Science Laboratory’s Curiosity rover
SLS/HLV Space Launch System/Heavy Lift Vehicle
Sv Sievert (1 Sv = 100 rem)
t metric ton (1 t = 1000 kg)
ΔV velocity change increment (km/s)

1.0 Introduction and Background

In 2007 and 2008, NASA conducted a multi-center, agency-wide study that examined the mission, payload and transportation system options and requirements needed to land humans on Mars in the post-2030 timeframe. The Mars Design Reference Architecture (DRA) 5.0 study (Ref. 1) provided an update to NASA’s earlier DRM 3.0 and 4.0 studies (Refs. 2 to 5) and was conducted in two phases. In Phase I, key architectural approaches, mission drivers and technology options were evaluated to support a down-selection process. Selected architectural approaches and systems concepts were then further refined in Phase II while also evaluating different surface exploration strategies. A three mission set of Mars landing missions over an ~10 year period was assumed with each mission carrying a crew of six. Opposition and fast-conjunction trajectories supporting short and long surface stay landing missions were examined (shown in Fig. 1) during Phase I. Mission characteristics and features associated with each trajectory type were evaluated and compared against the in-space transportation system requirements and the amount of scientific investigation that could be achieved using each mission type. The scientific community overwhelmingly preferred going to different sites and staying as long as possible so the crew could sample the planet’s rich geological diversity and maximize the scientific return on each mission. While opposition trajectories typically have short round trip mission times, the stay times at Mars are also short. This trajectory class also usually has one short and one long transit leg with the latter involving a Venus swing-by (VSB) maneuver that can bring the spacecraft relatively close to the Sun. By contrast, fast conjunction trajectories have long round trip and surface stay times, but transit times to and from Mars can be relatively short and no VSB is required on either mission leg. The long surface stay mission also has a lower total ΔV requirement than the short stay option, and therefore requires less propellant and less mass delivered to LEO. Because of its attractive features—lower total ΔV, a long surface stay, short transit times, and no VSB—the fast conjunction/long stay mission option was selected as the baseline for DRA 5.0.

Figure 1.—Features and characteristics associated with opposition and fast conjunction Mars missions.
Recent measurements of the energetic particle radiation environment inside the Mars Science Laboratory (MSL) spacecraft during its 253-day journey to Mars were made by the RAD attached to the Curiosity rover located inside MSL (Ref. 6). The detector measured an average of \( \sim 1.8 \times 10^{-3} \) Sv per day indicating that astronauts could receive a radiation dose of \( \sim 0.66 \) Sv (\( \sim 66 \) rem)—the limiting value established by NASA—during a fast conjunction Mars mission involving a 1-year journey through deep space. This dose rate could also be higher depending on the time in solar cycle that the mission occurs. When coupled with the potential for additional crew dose received during the lengthy exploration phase on Mars, it is of interest to see what proven NTP can do to reduce transit times further. Can NTP cut transit times in half and if so what are the engine requirements and vehicle configuration impacts?

The NTR was selected as the propulsion system of choice in DRA 5.0 because of its high thrust (10’s of klbf) and high specific impulse (\( I_{sp} \sim 900 \) to 940 s), its increased tolerance to payload mass growth and architecture changes, and its lower IMLEO which is important for reducing the HLV launch count, overall mission cost and risk. With a 100 percent higher \( I_{sp} \) than today’s best chemical rockets, the use of NTP reduced the required launch mass for the baseline long stay mission by over 400 t—the equivalent mass of the International Space Station. For the higher energy, short stay opposition missions examined during Phase I, NTP provided an even greater mass savings—over 530 t compared to chemical propulsion. This is important to note since short stay missions will likely be selected by NASA when it decides to send humans to orbit Mars in the early 2030s as outlined in the United States’ National Space Policy (Ref. 7). Lastly, it should not be overlooked that NTP is a proven technology successfully ground tested at the performance levels required for a human mission to Mars. The smallest and highest performing engine tested during the Rover/NERVA programs (Ref. 8)—the 25 klbf “Pewee” engine is sufficient for a wide variety of human mission to Mars when used in a clustered engine arrangement. This means there are no large scale-ups in size or performance needed for NTP as there are with other proposed non-chemical propulsion options.

DRA 5.0 featured separate cargo and crewed Mars transfer vehicles each using a common “core” NTPS with three 25 klbf Pewee-class engines. Two cargo vehicles were used to pre-deploy surface and orbital assets to Mars ahead of the crew who arrived during the next mission opportunity (~26 months later). The “Copernicus” crewed MTV (Fig. 2) is a 0-g\( _e \) vehicle (Refs. 9 and 10) consisting of three basic components: (1) the NTPS; (2) the crewed payload element; and (3) an integrated “saddle truss” and LH\( _2 \) propellant drop tank assembly that connects the two elements. With a propellant capacity of \( \sim 190 \) t in its two tanks, Copernicus can achieve 1-way transit times to Mars and back ranging from \( \sim 150 \) to 220 days over a 15-year synodic cycle starting in 2033.

Figure 2.—“Copernicus” crewed MTV in LEO prior to departure.
This paper examines the possibilities for achieving short transit times by reconfiguring the Copernicus design to allow it to carry larger quantities of LH$_2$ propellant needed for these more demanding missions. The paper covers the following topic areas. First, the operational principles and performance characteristics of the baseline 25 klbf Pewee-class engine used in this analysis are presented along with a summary of the Rover/NERVA programs’ technical accomplishments. Key NTP development activities currently funded under NASA’s Advanced Exploration Systems (AES) program are also mentioned. Mission and transportation system ground rules and assumptions used in the analysis are then presented along with trajectory data for a fast conjunction 2033 mission that covers a range of transit times. Copernicus’ key components and mission capabilities are reviewed next. To reduce 1-way transit times down to 130 days, then to 100 days, modifications are made to the Copernicus configuration incrementally—first with the addition of an in-line tank, then with the addition of paired sets of drop tanks—to increase the vehicle’s propellant loading. A “split crewed mission” option that does not require orbital assembly and uses Copernicus’ components is also described that can support a crewed mission with 90-day transit times. Mission scenario descriptions, key features and characteristics for five different vehicle configurations are presented along with the engines’ operational requirements. The paper ends with a summary of our findings and some concluding remarks.

2.0 NTR System Description and Performance Characteristics

The NTR uses a compact fission reactor core containing 93 percent “enriched” Uranium (U)-235 fuel to generate 100’s of megawatts of thermal power (MWt) required to heat the LH$_2$ propellant to high exhaust temperatures for rocket thrust. In an “expander cycle” Rover/NERVA-type engine (Fig. 3), high pressure LH$_2$ flowing from either a single or twin turbopump assembly (TPA) is split into two paths with the first cooling the engine’s nozzle, pressure vessel, neutron reflector, and control drums, and the second path cooling the engine’s tie-tube assemblies. The flows are then merged and the heated H$_2$ gas is used to drive the TPAs. The hydrogen turbine exhaust is then routed back into the reactor pressure vessel and through the internal radiation shield and core support structure before entering the coolant channels in the reactor core’s fuel elements. Here it absorbs energy produced from the fission of U-235 atoms, is superheated to high exhaust temperatures ($T_{ex}$ ~2550 to 2950 K depending on fuel type and uranium loading), then expanded out a high area ratio nozzle (~300:1) for thrust generation.

Controlling the NTR during its various operational phases (startup, full thrust and shutdown) is accomplished by matching the TPA-supplied LH$_2$ flow to the reactor power level. Multiple control drums, located in the reflector region surrounding the reactor core, regulate the neutron population and reactor power level over the NTR’s operational lifetime. The internal neutron and gamma radiation shield, located within the engine’s pressure vessel, contains its own interior coolant channels. It is placed between the reactor core and key engine components to prevent excessive radiation heating and material damage.

Figure 3.—Schematic of “expander cycle” NTR engine with dual LH$_2$ turbopumps.
The fuel elements (FE) tested in the Rover/NERVA program (Ref. 8) consisted of a “graphite matrix” material that contained the U-235 fuel in the form of either coated particles of uranium carbide (UC\(_2\)) or as a dispersion of uranium and zirconium carbide (UC-ZrC) referred to as “composite” fuel. Each FE (see Fig. 4) had a hexagonal cross section (~0.75 in. across the flats) and 19 axial coolant channels that were coated with niobium carbide (NbC) initially, then with zirconium carbide (ZrC) using a chemical vapor deposition (CVD) process. This protective coating, applied to the exterior FE surfaces as well, helped reduce hydrogen erosion of the graphite. Individual elements were 1.32 m (52 in.) in length and produced ~1 MW\(_e\).

This basic FE shape and length was introduced in the KIWI-B4E reactor and became the standard used in the 75 klb\(_f\) Phoebus-1B, 250 klb\(_f\) Phoebus-2A, 25 klb\(_f\) Pewee and the 55 klb\(_f\) NERVA NRX series of engines (Ref. 8). The Rover program’s Pewee engine (Ref. 8) was designed and built to evaluate higher temperature, longer life fuel elements and improved coatings. It set several performance records including the highest average fuel element exit gas temperature of ~2550 K, and the highest peak fuel temperature ~2750 K. Other performance records included average and peak power densities in the reactor core of ~2340 MW/m\(^3\) and ~5200 MW/m\(^3\), respectively. A new CVD coating of ZrC was also introduced in Pewee that showed performance superior to the NbC coating used in previous reactor tests.

In addition to FEs, the engine reactor cores also included tie tube (TT) elements of the same hexagonal shape that provided structural support for the FEs. A coaxial coolant tube of Inconel inside each TT supplied a source of heated hydrogen for turbine drive power and a sleeve of zirconium hydride (ZrH) moderator material could also be incorporated in the TTs to help raise neutron reactivity (shown in Fig. 4). In the larger size engines tested during the Rover/NERVA programs, a “sparse” FE—TT arrangement was used with each FE having two adjacent TTs and four adjacent FEs comprising its six surrounding elements (Ref. 11). In this sparse pattern, the FE to TT ratio is ~3 to 1.

In the Small Nuclear Rocket Engine (SNRE) design developed by Los Alamos National Laboratory near the end of Rover/NERVA (Ref. 12), shorter (0.89 m/35 in.) FEs were used so additional TTs were included in the reactor to increase core reactivity. With the “SNRE” FE—TT pattern each FE has three adjacent TTs and three adjacent FEs surrounding it (shown in Fig. 4) and the FE to TT ratio is ~2 to 1. An important feature common to both the sparse and SNRE FE—TT patterns is that each tie tube provides mechanical support for six adjacent fuel elements.

![Figure 4.—Coated particle and composite SNRE fuel element and tie tube arrangement.](image-url)
Recent Monte Carlo N-Particle (MCNP) transport modeling of engine reactor cores by Schnitzler et al., (Refs. 11, 13, and 14) has shown that the SNRE design can be scaled down to even lower thrust levels (~5.3 to 7.4 klbf) or up to the 25 klbf Pewee-class engine used in Mars DRA 5.0. For low thrust engines, short length elements (~0.89 m) and a “dense” FE—TT pattern are used consisting of parallel rows of FEs and TTs. Each FE has four adjacent TTs and two adjacent FEs surrounding it and the FE to TT ratio is ~1 to 1. By using the SNRE FE—TT pattern and increasing the FE length from 0.89 to 1.32 m (the same length used in the Rover program’s Pewee engine), the U-235 fuel loading in the core can be lowered from ~0.60 to 0.25 grams/cm³ allowing the FEs to operate at higher peak fuel temperatures (Tpeak) while still staying safely below the melt temperature for composite fuel of ~3050 K. It also allows higher hydrogen exhaust/higher chamber inlet temperatures thereby increasing the engine’s Isp capability. Higher Isp operation can help stretch the available LH₂ propellant loading to meet mission requirements, or in the case of an emergency to allow a safe return of the crew. Engine and reactor performance characteristics for small and Pewee-class engines using composite fuel are summarized in Table 1.

The reference NTR engine used in DRA 5.0 and this paper is a 25 klbf Pewee-class engine that is based on an “axial-growth” version of the SNRE. It utilizes an expander cycle and has the following nominal performance parameters: Tex ~2790 K, chamber pressure ~1000 psia, nozzle area ratio (NAR) ~300:1, and Isp ~906 s. The LH₂ flow rate is ~12.5 kg/s and the engine thrust-to-weight ratio is ~3.50. The overall engine length is ~7.01 m, which includes an ~2.16 m long, retractable radiation-cooled nozzle skirt extension. The corresponding nozzle exit diameter is ~1.87 m. The higher chamber pressure helps to maintain reasonable nozzle dimensions at the assumed NAR. The engine’s reactor core contains 564 FEs and 241 TTs that are 1.32 m long like those used in Pewee. The core power level and fuel matrix power density are ~560 and ~3.44 MW/liter, respectively. The U-235 fuel loading used in the reactor FEs is ~0.25 grams/cm³ and the inventory of 93 percent enriched U-235 in the core is just under 37 kg. At this fuel loading the engine is able to operate at a Tpeak of ~2860 K providing a temperature margin to fuel melt of ~190 K. During enhanced operation Tpeak is raised to ~3010 K (a margin-to-melt temperature of ~40 K) and Tex to ~2940 K resulting in an ~35 s increase in Isp to ~940 s if needed. Table 1 also shows a “radial-growth” version of a Pewee-class engine using shorter FEs and a sparse FE—TT pattern for comparison.

<table>
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<tr>
<th>Performance Characteristic</th>
<th>7.420-lbf Option Baseline</th>
<th>Axial Growth Option Nominal Enhanced</th>
<th>Radial Growth Option Nominal Enhanced</th>
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<td>U-235 Mass (kg)</td>
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TABLE 1.—PERFORMANCE CHARACTERISTICS FOR SMALL AND FULL SIZE “COMPOSITE” FUEL NTR ENGINES
High temperature UC-ZrC in graphite “composite” fuel with ZrC coating was selected as the primary fuel form in this analysis. Composite FEAs were first tested in the “Nuclear Furnace” element test reactor (Ref. 8) and withstood peak power densities of ~4500 to 5000 MW/m³. They also demonstrated better corrosion resistance than the standard coated particle graphite matrix fuel element used in the previous Rover/NERVA reactor tests. Composite fuel’s improved corrosion resistance is attributed to its higher coefficient of thermal expansion that more closely matches that of the protective ZrC coating, thereby helping to reduce coating cracking. Electrical-heated composite fuel elements were also tested by Westinghouse in hot hydrogen at 2700 K for ~600 min—equivalent to ten 1-hr cycles. At the end of Rover/NERVA program, composite fuel performance projections (Ref. 15) were estimated at ~2 to 6 hr at full power for hydrogen exhaust temperatures of ~2500 to 2800 K.

In addition to these carbide-based fuels, a ceramic-metallic or “cermet” fuel consisting of uranium dioxide (UO₂) in a tungsten (W) metal matrix was developed in the GE-710 and ANL nuclear rocket programs (Refs. 16 and 17) as a backup to the Rover/NERVA fuel. While no integrated reactor/engine tests were conducted, a large number of fuel specimens were produced and exposed to non-nuclear hot H₂ and irradiation testing with promising results. Both fuel options are under development today in NASA’s Advanced Exploration Systems (AES) program and the Nuclear Cryogenic Propulsion Stage (NCPS) project (Ref. 18) which began in FY’12. This 3-year project is a collaboration between NASA and DOE involving five key task activities that include state-of-the-art engine modeling, mission design and requirements definition, fuel element fabrication and non-nuclear testing, and analysis of affordable options for nuclear ground testing and engine development. If successfully completed by the end of FY’14, a NCPS-Phase II effort would begin that could culminate in ground testing a small NTR engine at the Nevada Test Site (NTS) in the early 2020’s, followed by a flight demonstration mission several years later (Ref. 19).

Finally, the motivation for selecting NTP as the propulsion system of choice for Mars is simple—it is a proven technology with a specific impulse that is twice that of today’s best chemical rockets. During the Rover/NERVA programs (1955 to 1972), a technology readiness level (TRL ~5 and 6) was achieved (Ref. 8). Twenty rocket reactors were designed, built and ground tested in integrated reactor/engine tests that demonstrated: (1) a wide range of thrust levels (~25, 50, 75, and 250 klb); (2) high temperature carbide-based nuclear fuels that provided hydrogen exhaust temperatures up to 2550 K \(\text{achieved in Pewee}\); (3) sustained engine operation \(\text{over 62 min for a single burn achieved in the NRX-A6}\); as well as (4) accumulated lifetime at full-power; and (5) restart capability \(\text{>2 hr with 28 startup and shutdown cycles achieved in the NRX-XE experimental engine}\)—all the requirements needed for short transit time Mars missions. Just as important, NTP requires no large scale-ups in size or performance that are required with other advanced propulsion options. A cluster of three to four Pewee-class engines are more than adequate.

### 3.0 ΔV Budgets, Mission and Transportation System Ground Rules and Assumptions

The fast conjunction 2033 opportunity was selected as the baseline mission for this analysis since missions performed in the 2033 to 2035 timeframe have ΔV budgets near their minimum values over the 15-year synodic cycle. As a result, it is expected that the shortest achievable transit times would be found in this opportunity. Table 2 shows the ideal ΔV budgets as a function of decreasing transit time for the outbound Earth to Mars mission leg. It is assumed that the crewed MTV departs from a 407 km circular LEO, then captures into and departs from a 250 by 33,793 km elliptical Mars orbit (EMO) with a 24-hr orbital period. Gravity losses are also added to the ideal ΔVs and depend on the particular mission C₃, vehicle thrust-to-weight (F/W), and propulsion system specific impulse.
Table 3 shows the ideal $\Delta V$ budgets for the same range of transit times for the inbound Mars to Earth mission leg with Mars departure occurring in late 2035. The last column in the table shows the corresponding stay time at Mars assuming equal outbound and inbound transit times. As transit times decrease, it becomes necessary to perform a slowdown burn on the return leg to ensure that arrival $V_{\text{inf}}$ at Earth is less than 6.81 km/s at an altitude of ~125 km.

For 180-day transits to and from Mars and a 575-day stay time at Mars, the total round trip time and $\Delta V$ budget (minus g-loss) for the mission are 935 days and 6.045 km/s, respectively. Cutting the transit times in half to 90-days increases the Mars stay time to ~673 days, but decreases the round trip time to ~853 days. The mission $\Delta V$ budget however, increases by ~250 percent to ~15.175 km/s! While challenging, transit times this short are possible using NTP.

The mission and transportation system ground rules and assumptions used in this paper are summarized in Tables 4 and 5, respectively. Table 4 provides information about the mission profiles and operational scenarios, and the assumed parking orbits at Earth and Mars. In addition to the large $\Delta V$ requirements for the primary propulsion maneuvers—trans-Mars injection (TMI), Mars orbit capture (MOC), and trans-Earth injection (TEI)—smaller $\Delta V$ maneuvers are needed for rendezvous and docking (R&D) of vehicle components during the LEO assembly phase, for spacecraft attitude control during in-space transit, and for “split crewed mission” orbital operations at Mars.
TABLE 4.—MISSION AND PAYLOAD GROUND RULES AND ASSUMPTIONS

| Mission Profiles: | • Split mission; cargo pre-deployed to Mars before crew leaves Earth  
  • Cargo missions use “1-way” minimum energy trajectory  
  • Round trip crewed missions use “fast conjunction” trajectories  
  • Direct return uses Orion capsule for crew recovery at mission end  
  • For “split crewed mission option”, a pre-deployed Earth Return Vehicle (ERV) uses minimum energy trajectory out to Mars and a fast conjunction transfer trajectory back to Earth |
|---|---|
| Missions depart for Mars from low Earth orbit (LEO); capture into and depart from a 24-hr elliptical Mars orbit (EMO) | • LEO: 407 km circular  
  • 24-hr EMO: 250 km x 33,793 km |
| Mars Mission ΔV Budgets: | • Mission dates, trip times and ΔV budgets provided in Tables 1 & 2 |
| Additional ΔV Requirements: Advanced Material Bipropellant Rocket (AMBR) RCS thrusters used to perform non-primary propulsion maneuvers | • LEO R&D between orbital elements: ~15 – 100 m/s  
  • Coast attitude control and mid-course correction: ~15 m/s and ~50 m/s, respectively  
  • Mars orbit maintenance plus R&D: ~100 m/s |
| Crewed Mission Payload Masses: Varies with crew size and mission duration; consumables based on a crew consumption rate of ~2.45 kg/person/day; payload also includes a short sattelite truss (SST), a transfer tunnel with second docking module (TDM), and an exterior contingency consumables container that is jettisoned prior to the trans-Earth injection (TEI) maneuver | • Transit Habitat: 22.7 t – 27.5 t (minus consumables)  
  • SST/TDM/Container: 5.08 t / 1.76 t / 23% of stored food  
  • Crew (4-6): 0.4 t – 0.6 t  
  • Total Consumables: 3.581 – 5.37 t (4 – 6 crew for 1-yr); with extra consumables stored in exterior container  
  • Orion / MPCV: ~10.0 t  
  • Returned Samples: 0.25 t – 0.5 t (Mars) |
| Mission Abort Strategy: | • Outbound: Abort to Mars Surface  
  • At Mars: Abort to the orbiting crew MTV which carries contingency consumables onboard |

TABLE 5.—NTR TRANSPORTATION SYSTEM GROUND RULES AND ASSUMPTIONS

| NTR System Characteristics | • Engine / Fuel Type: SNRE-derived / UC-ZrC “Composite”  
  • Propellant: LH₂  
  • Thrust Level: 25 kNb (“Pewee-class” is baseline) (3 – 4 engine cluster on “Core” NTPS)  
  • Fuel Element Length: 1.32 m  
  • Exhaust Temp: \( T_e \approx 2790 – 2940 \) K  
  • Chamber Pressure: \( P_{ch} \approx 1000 \) psi  
  • Nozzle Area Ratio: 300:1  
  • \( I_{sp} \) Range: 906 s (2790 K) – 941 s (2940 K) |
| Propellant Margins | • Cooldown: 3\% of usable LH2 propellant  
  • Performance reserve: 1\% on ΔV  
  • Tank trapped residuals: 2\% of total tank capacity |
| Reaction Control System (LEO R&D, Setting, Attitude Coast Control, and Mid-course Correction Burns) | • Propulsion Type: AMBR 200 lb thrusters  
  • Propellant: NTO / N₂H₂  
  • Nominal \( I_{sp} \): 335 seconds |
| LH₂ Cryogenic Tanks and Passive Thermal Protection System (TPS) | • Material: Aluminum-Lithium (Al/Li) and Composite; (Al/Li is the baseline)  
  • Tank OD: 10.0 m  
  • Tank L: ~19.7 m (for “core” NTPS) ~22.7 m (for “in-line” and drop tanks)  
  • Geometry: cylindrical with root 2/2 ellipsoidal domes  
  • Insulation: 1” SOFI (~0.78 kg/m³) + 60 layers of MLI (~0.90 kg/m³) |
| Active Cryo-Fluid Management / Zero Boil-Off (ZBO) LH₂ Propellant System | • Reverse turbo-Brayton ZBO cryocooler system powered by PVAs  
  • ZBO system mass and power requirements driven by core stage size; ~930 kg and ~8.87 kW (10.0 m D) |
| Photovoltaic Array (PVA) Primary Power System | • Circular PVA sized for ~7 kW, at 1 A.U., two arrays provide power for ZBO cryocoolers on core NTPS and in-line tanks when used, PVA mass is ~455 kg & array area ~25 m²; to supply 1 kW at Mars requires ~10 m² of array area  
  • “Keep-alive” power supplied by lithium-ion battery system |
| Dry Weight Contingency Factors | • 30\% on NTR system and composite structures (e.g., saddle truss)  
  • 15\% on established propulsion, propellant tanks, spacecraft systems |
| SLS / HLV Launch Requirements:  
  – Lift Capability to LEO  
  – Cylindrical Payload (PL) Envelope | • ~140 t; NTPS with external crew radiation shields  
  • 11 m D x ~33.8 m L (for DRA 5.0 crewed PL element) |
For the crewed mission, the outbound payload mass varies with the crew size and mission duration. For long surface stay Mars missions, the crewed MTV carries contingency consumables equivalent to that found on the habitat lander. This allows the crew MTV to function as an orbital “safe haven” in the event of a major failure of a key surface system. For the nominal surface mission, the contingency consumables are jettisoned prior to the TEI maneuver. Fixed mass payload elements include the Orion Multi-Purpose Crew Vehicle (MPCV). For Mars DRA 5.0 and in this analysis it is assumed that ~250 to 500 kg of samples are returned.

Table 5 lists the key transportation system ground rules and assumptions. The NTR engine and fuel type, operating characteristics, and thrust levels examined are summarized first. All engines use composite fuel with a U-235 loading of 0.25 g/cm³. With Tex ~2790 K and NAR ~300:1, the Isp is ~906 s with higher Isp values achievable by increasing the fuel operating temperature. The total mission LH₂ propellant loading consists of the usable propellant plus performance reserve, post-burn engine cooldown, and tank-trapped residuals. For the smaller auxiliary maneuvers, a storable bipropellant RCS system is used. All transfer vehicle configurations utilize a “split RCS” with 16 of 32 AMBR thrusters and approximately half of the bipropellant mass located on the rear propulsion stage and the forward most saddle truss adaptor ring just behind the TransHab module.

The LH₂ propellant carried by the various vehicles is stored in the same “state-of-the-art” Al/Li LH₂ propellant tank being developed for the SLS/HLV that will support future human exploration missions. For this analysis, tank sizing assumes a 30 psi ullage pressure, 5 gE axial/2.5 gE lateral launch loads, and a safety factor of 1.5. A 3 percent ullage factor is also assumed. All tanks use a combination foam/multilayer insulation (MLI) system for passive thermal protection. A zero boil-off (ZBO) “reverse turbo-Brayton” cryocooler system is used on the core NTPS and “in-line” LH₂ tanks (where required) to eliminate boil-off during LEO assembly and the remainder of the mission which can be as long as 950 days in duration. The propellant tank heat load is largest in LEO and sizes the ZBO cryocooler system. Solar photovoltaic arrays are baselined for supplying the primary electrical power needed for all key transfer vehicle subsystems. Because of the decreased solar radiation at Mars (~486 W/m²), array areas can become quite large (~10 m²/kWₑ) necessitating multiple arrays for human Mars missions.

Table 5 also includes the “dry weight contingency” (DWC) factors, along with the lift and payload envelope requirements for the “upgraded” SLS/HLV. A 30 percent DWC is used on the NTR system and advanced composite structures (e.g., stage adaptors, trusses) and 15 percent on heritage systems (e.g., Al/Li tanks, RCS, etc.). The NTPS and crewed payload element (PL) determine the required lift and PL envelope. The optimum tank diameter and mass for a Mars NTPS are 10 m and ~140 t (Ref. 9). The crewed PL element includes the “packaged” TransHab module with PVA power system, the short saddle truss, consumables container and transfer tunnel with secondary docking module (DM), and the Orion MPCV (shown Figs. 2 and 6). The PL envelope’s diameter is ~11 m (the saddle truss outer dimension) and its length can be up to ~33.8 m (the DRA 5.0 PL length including the Orion MPCV).

### 4.0 Mars DRA 5.0: “7-Launch” NTR Mission Overview

The 7-Launch NTR Mars mission strategy (Ref. 9) for a 2033 human landing mission is illustrated in Figure 5. It assumes a long surface stay, split cargo/piloted mission approach. Two cargo flights pre-deploy a cargo lander to the surface and a habitat lander into Mars orbit where it remains until the arrival of the crew on the next mission opportunity (~26 months later). The cargo flights utilize “1-way” minimum energy, long transit time trajectories. Four HLV flights carried out over 90 days (~30 days between launches), deliver the required components for the two cargo vehicles. The first two launches deliver the NTR propulsion stages each with three 25 klbf NTR engines. The next two launches deliver the cargo and habitat lander payload elements which are enclosed within a large triconic-shaped aeroshell that functions as a payload shroud during launch, then as an aerobrake and thermal protection system during Mars aerocapture (AC) and subsequent entry, descent and landing (EDL) on Mars. Vehicle assembly involves Earth orbit rendezvous and docking (R&D) between the propulsion stages and payload elements with the NTR stages functioning as the active element in the R&D maneuver.
Once the operational functions of the orbiting habitat and surface cargo landers are verified, and the Mars Ascent Vehicle (MAV) is supplied with ISRU-produced ascent propellant, the crewed MTV is readied and departs on the next mission opportunity. The baseline Copernicus design is capable of 1-way transit times ranging from ~150 to 220 days depending on the particular opportunity. For the 2033 mission, 150-day transit times are possible. Like the cargo MTVs, Copernicus is assembled in LEO using Earth orbit R&D. It uses the same “common” NTPS but includes additional external radiation shielding on each engine for crew protection during engine operation. Three HLV launches over 60 days are used to deliver the vehicle’s key elements which include: (1) the NTPS; (2) the integrated “saddle truss” and LH₂ drop tank assembly; and (3) the crewed payload. The payload element includes the TransHab module with its six crew, the Orion MPCV for vehicle-to-vehicle transfer and “end of mission” re-entry, a secondary T-shaped DM, a contingency consumables container and connecting structure. Four 12.5 kWₑ/125 m² rectangular PVAs provide ~50 kWₑ of electrical power at Mars for crew life-support (~30 kWₑ), a ZBO Brayton cryocooler system (~10 kWₑ), and high data-rate communications (~10 kWₑ) with Earth. When assembly is completed, the Mars crew is delivered to LEO on either the SLS or a commercial launch vehicle and docks with Copernicus on its underside using the secondary DM that connects the TransHab crew module and contingency consumables container (shown in Figs. 6 and 7).

Following the TMI maneuver, the drained LH₂ drop tank, attached to the saddle truss, is jettisoned and the crewed MTV coasts to Mars under 0-g conditions with its four PVAs tracking the Sun. Attitude control and mid-course correction maneuvers are provided by Copernicus’ split RCS that uses 200 lbₜ storable bipropellant Advanced Material Bipropellant Rocket (AMBR) thrusters located on the rear NTPS and the short saddle truss forward adaptor ring just behind the TransHab module. After the MOC burn, Copernicus rendezvous with the orbiting Hab lander using engine cool-down thrust and the vehicle’s RCS. The crew then transfers over to the lander using the MPCV. After crew transfer, the MPCV returns and docks to the TransHab autonomously. The crew then initiates EDL near the cargo lander and begins the surface exploration phase of the mission. After ~592 days on the surface, the crew lifts off using the
MAV and returns to *Copernicus* using its secondary DM (shown in Fig. 6). Following the transfer of the crew and samples, the MAV is jettisoned. The crew then begins a weeklong checkout and verification of all MTV systems, jettisons the DM and contingency consumables and performs the TEI burn to begin the journey back to Earth. After a 150-day return trip, the crew enters the MPCV, separates from the MTV and re-enters Earth’s atmosphere while *Copernicus* flies by Earth at a “sufficiently high altitude” and is disposed of into heliocentric space. Although *Copernicus* was operated in an “expendable mission mode” in DRA 5.0 to reduce total IMLEO and number of HLV launches, it can readily be modified to operate in a “reuse mode” by providing the vehicle with additional propellant capacity as discussed elsewhere (Ref. 10).

The *Copernicus* crewed MTV has an overall length of ~93.7 m (Fig. 7) and an IMLEO of ~331.7 t for the 150-day transit 2033 mission. Included are (1) the NTPS (~138.8 t); (2) the saddle truss and LH$_2$ drop tank (~125.6 t); and (3) the crew payload section (~67.3). The NTPS uses a three-engine cluster of 25 klbf NTR engines and also carries additional external radiation shield mass (~6 t) for crew protection. The NTPS uses an Al/Li LH$_2$ tank size which has a diameter (D) and length (L) of 10 m D by 19.7 m L. The LH$_2$ tank has a propellant capacity of ~87.2 t but is slightly off-loaded for this particular mission and carries ~86.8 t. The NTPS also carries avionics, RCS, auxiliary battery and PVA power, docking and Brayton-cycle ZBO refrigeration systems located in the forward cylindrical adaptor section. To remove ~78 W of heat penetrating the 60 layer MLI system in LEO (where the highest tank heat flux occurs), the 2-stage cryocooler system requires ~8.9 kW$_e$ for its operation. Twin circular PVAs on the NTPS provide the electrical power for the ZBO system in LEO until the four primary PVAs on the crewed PL section are deployed prior to TMI.

*Copernicus*’ second major component is its saddle truss and LH$_2$ drop tank assembly. The saddle truss is a rigid, spine-like composite structure that wraps around the upper half of the LH$_2$ drop tank and connects the NTR stage to the forward payload section. It is ~27.7 m long and has a mass of ~9 t. The saddle truss is open underneath allowing the drained LH$_2$ drop tank to be jettisoned after the TMI burn is completed. The ~22.7 m long LH$_2$ drop tank has a mass of ~21 t and a propellant capacity of ~102.4 t but is also off-loaded and only carries ~91.9 t of LH$_2$ propellant.
*Copernicus*’ third and final component is its payload section. In DRA 5.0 it was designed for launch as a single integrated unit and thus determines the overall size of the PL envelope. The integrated payload element is ~33.8 m long and includes the short saddle truss, “T-shaped” DM and transfer tunnel, consumables container, TransHab and crew, and the Orion MPCV. The DM provides “secondary access” to *Copernicus* for the crew delivery MPCV and the MAV (see Figs. 6 and 7). Following the crew’s return from Mars and MAV separation, the DM and attached consumables container are both jettisoned to reduce vehicle mass prior to TEI (see Fig. 5).

The total crewed payload mass at TMI is ~67.3 t distributed as follows: (1) short saddle truss (~5.1 t); (2) DM and transfer tunnel (~2.8 t); (3) contingency consumables and jettisonable container (~10.7 t); (4) TransHab with its primary PVAs (~27.5 t); (5) transit consumables (~4.6 t); (6) crew (~0.6 t); (7) MPCV (~10 t); and (8) forward RCS and propellant (~6 t). *Copernicus*’ total RCS propellant loading is ~11.9 t with the “post-TMI” RCS propellant load split between the NTPS (~7 t) and the short saddle truss forward cylindrical adaptor ring (~4.9 t).

Lastly, for this particular mission, the performance requirements on operating time and restart for *Copernicus*’ three 25.1 klbf NTR engines are quite reasonable. For the round trip mission, there are four primary burns (three restarts) that use ~168.7 t of LH₂ propellant. With ~75.3 klbf of total thrust and a $I_{sp}$ of ~906 s, the total engine burn time for the mission is ~74.6 min (~53.2 min for the “2-perigee burn” TMI maneuver, ~13.6 min for MOC, and ~7.8 min for TEI), well under the ~2 hr accumulated engine burn time and 27 restarts demonstrated by the NERVA eXperimental Engine—the NRX-XE (Ref. 8).
5.0 Copernicus “Growth” Options Allowing Faster Transit Time Missions

As discussed in Section 4.0, Copernicus was sized so it can perform all fast conjunction missions in the 2033 to 2048 timeframe. For the 2033 mission opportunity, with its propellant tanks near their maximum capacity of ~190 t, Copernicus can achieve 150-day transit times. Besides human Mars missions, the “3-element” Copernicus spacecraft (Fig. 8) has significant capability that can be exploited for a variety of other mission applications such as a reusable human mission to the large, high energy near Earth asteroid—Apophis in 2028. “Scaled down” versions of Copernicus can also be used for reusable lunar cargo delivery and crewed lunar landing missions. Mission details and vehicle characteristics for these applications are provided elsewhere (Ref. 10).

Figure 9 shows the round trip mission ∆V versus transit time for the 2033 opportunity using the trajectory data found in Tables 2 and 3. As is evident, the ∆V requirements rise dramatically as the transit times are shortened. Decreasing the transit times by 60 days—from 180 to 120 days—increases the total mission ∆V by ~46 percent. A further 60-day reduction in transit time—down to 90 days each way—increases the total mission ∆V by ~250 percent.

To achieve transit times less than 150 days, it is necessary to modify the Copernicus spacecraft so it can carry additional propellant. This will be done in incremental steps—first with the addition of an in-line tank, then with the addition of paired sets of drop tanks—to show what can be achieved, but also to show what is required in terms of extra hardware, engine operating time and additional HLV launches.

Figure 8.—Growth paths and alternative missions for Copernicus spacecraft using modular components.
5.1 Copernicus-2 Option: Addition of an In-Line Tank to Achieve 130-day Transit Times

To reduce transit times by 20 days—from 150 to 130 days, the Copernicus-2 configuration positions an “in-line” LH$_2$ tank between the NTPS and integrated saddle truss/drop tank assembly (shown in Figs. 8 and 10). This configuration has an IMLEO of ~443.8 t consisting of the “wet” NTPS (~136.4 t), in-line tank (~118.5 t), saddle truss/drop tank assembly (~119 t) and the crew PL section (~69.9 t). The overall vehicle length is ~117.6 m including the Orion MPCV at ~8.9 m. The LH$_2$ tank length in the NTPS is ~19.7 m and it carries ~86.1 t of propellant (close to its maximum capacity of ~87.2 t). The in-line and
drop tank lengths are the same at ~20.5 m and each tank carries ~91.5 t of LH₂ propellant. With these equal tank lengths, the masses of in-line tank element and the saddle truss/drop tank assembly are balanced at ~118.8 t.

For this 130-day transit mission, there are four primary burns (with three restarts) and ~253.8 t of LH₂ propellant is used. A fourth engine is added to the NTPS to reduce the total engine operating time and minimize g-losses (~326 m/s) during the TMI maneuver but this increases NTPS inert mass at the expense of LH₂ and RCS propellant loading. With ~100.4 klbₜ of total thrust and a T thrust of 906 s, the total engine burn time for this mission is ~84.2 min. The first TMI perigee burn is the longest single burn at ~35 min. After this burn, the drop tank is drained and then jettisoned to reduce vehicle mass and propellant consumption during the second perigee burn (~21.8 min). The in-line tank and NTPS supply the propellant for the remaining MOC (~17.5 min) and TEI burns (~9.9 min).

**Copernicus-2** also requires one additional HLV launch as well as the development of the in-line tank element with its own ZBO cryocooler system. For the longer in-line tank, the cryocooler system requires ~9.13 kWₑ of electrical power supplied by PVAs on the NTPS and in-line tank. Even with a fourth engine added, the total burn time requirement on the engines is increased by 13 percent (from 74.6 to 84.2 min). An alternative to adding a fourth NTR engine would be to add oxygen “afterburner” nozzles to **Copernicus’** three NTR engines. In the “LOX-Augmented” NTR (LANTR) option (Refs. 20 and 21), oxygen is injected into the divergent section of the nozzle downstream of the sonic throat. Here it mixes with reactor-heated H₂ and undergoes supersonic combustion adding both mass and chemical energy to the engine’s exhaust. By operating the LANTR engines with a LOX-to-LH₂ mixture ratio (MR) = 1 during the first TMI perigee burn, the engine’s thrust level can be increased by over 62 percent—from 25 to ~40.6 klbₜ. The addition of oxygen lowers the Isp to ~725 s but this is still ~270 s higher than that achieved by LOX/LH₂ engines.

### 5.2 Copernicus-3 Option: Addition of Twin Drop Tanks to Achieve 120-day Transit Times

In the **Copernicus-3** configuration (shown in Figs. 8 and 11), the saddle truss/drop tank assembly is replaced by a 4-sided “star truss” with twin LH₂ drop tanks that supplement the propellant carried by the NTPS and in-line tank. **Copernicus-3** has an IMLEO of ~551.3 t including the NTPS (~139.9 t), the in-line tank (~133.8 t), the “star truss” with drop tanks (~210.3 t) and the crew PL section (~67.3 t). Two drop tanks (~190 t) carrying ~159.4 t of LH₂ (~79.7 t per tank) are delivered to LEO on two HLV launches, then attached to the ~25.4 m long star truss (~20.3 t) during the LEO assembly phase. Each drop tank is ~18.2 m long. The overall vehicle length is ~118.2 m including the 8.9 m Orion MPCV. The LH₂ loading in the off-loaded NTPS, in-line and twin drop tanks is ~83.1, ~98.6, and ~159.4 t, respectively. The lengths of NTPS and in-line LH₂ tanks are ~19.7 and ~21.9 m. During the four primary mission burns ~321 t of LH₂ propellant is used. With 100.4 klbₜ of total thrust and Tₚ = ~906 s, the total engine burn time for this mission is ~106.4 min, close to the ~2 hr demonstrated on the NRX-XE engine. After completing a long first perigee burn (~52 min), both drop tanks are jettisoned. The remaining burn durations are as follows: second perigee burn (~20.7 min), MOC (~22.2 min) and TEI (~11.5 min). Four HLV launches are required to deliver the major transportation systems elements (the NTPS, in-line tank and two drop tanks) to LEO. An upgraded SLS is used to deliver the crewed payload element and the large but lightweight star truss.

### 5.3 Copernicus-3a Option: Addition of Second Set of Drop Tanks to Achieve 100-day Transit Times

In transitioning from 120 to 100-day transit times the total mission ΔV increases by ~42 percent—from ~8.83 to ~12.5 km/s. To satisfy the high-energy requirements of this mission, **Copernicus-3a** (shown in Figs. 8 and 12) adds a second set of drop tanks to the star truss. Adjustments are made to the PL as well. The crew size is reduced to four and the mass of the TransHab and consumables needed for the
mission are reduced accordingly. Composite LH₂ tanks are also used instead of Al/Li. Lastly, the operating temperature of the engines’ (UC-ZrC) graphite composite fuel is increased to achieve a specific impulse of ~940 s (see Table 1). The resulting vehicle has an IMLEO of ~720.7 t, which includes the NTPS (~137.6 t), the in-line tank (~123.7 t), the star truss with four common LH₂ drop tanks (~397.5 t) and the crew PL section (~61.9 t). The PL element includes the Orion MPCV (~15.1 t) which carries additional propellant (~5.1 t) for the final Earth slowdown burn (ΔV ~1.266 km/s) before re-entry while the MTV flies past Earth and is disposed of in heliocentric space.

Figure 11.—Copernicus-3 Option: Supports 120-day transit times for the 2033 mission.

Figure 12.—Copernicus-3a Option: Supports 100-day transit times for the 2033 mission.
The LH₂ tank lengths in the NTPS and in-line element are ~19.7 and ~21.9 m, and they carry ~87.2 and ~95.5 t of propellant, respectively. *Copernicus-3a*’s four drop tanks carry ~327.8 t of propellant (~82 t per tank), and each tank is ~18.6 m long. The drop tanks are jettisoned in pairs after the first and second perigee burns to reduce vehicle mass and propellant consumption during the TMI maneuver. The overall vehicle length is ~117.6 m.

The total usable propellant for the mission is ~480.7 t. With 100.4 klbf of thrust and \( Iₜ \approx 940 \) s, the total engine burn time is ~165.4 min. This value exceeds the ~2 hr demonstrated on the NRX-XE engine by ~45 min. The first perigee burn is the longest single burn at ~78.3 min and also exceeds the single burn duration record of ~62 min set by the NRX-A6 engine. The duration of the remaining burns are as follows: second perigee burn (~37.3 min), MOC (~33.3 min), and TEI (~16.5 min). Seven HLV launches are required to deliver the major transportation systems elements (the NTPS, in-line tank, four drop tanks, and PL element) to LEO. A single SLS with sufficient PL volume will be used to launch the large but lightweight star truss.

### 5.4 Copernicus “Split Crewed Mission” Option for Achieving 100—90-day Transit Times

*Copernicus-3a* is a large complex vehicle that requires eight launches to deliver the various elements to LEO. From the initial launch of the NTPS until the last drop tank is attached to the star truss (via orbital assembly) could take ~7 months assuming 30 days between SLS/HLV launches. Failure to deliver the elements on time and in the prescribed order could delay vehicle assembly and jeopardize mission departure within the specified launch window.

Using the *Copernicus-2* spacecraft plus its three key components—the NTPS, in-line tank and integrated saddle truss/drop tank assembly—now configured as an Earth Return Vehicle (ERV), a “split crewed mission” architecture can be utilized that can reduce vehicle size and complexity, eliminate the need for orbital assembly and an extra engine, and can achieve transit times as low as 90 days. Use of a pre-deployed ERV is not a new idea. In NASA’s DRM 1.0 study in 1993 (Ref. 22), the ERV was one of the key transportation system elements used in the mission architecture.

The split crewed architecture for very fast 1-way transit missions is outlined in Figure 13. Using the *Copernicus-2* spacecraft plus an ERV, a 2033 landing mission would require a round trip time of ~857 days with ~667 days spent in Mars orbit and on the surface. The outbound transit time on *Copernicus-2* is 100 days and the inbound transit time using the ERV is 90 days. In this split mission, the ERV is pre-deployed to Mars orbit in advance of the crew. It departs from LEO in December 2030 (departure \( Cₜ \approx 10.794 \text{ km}^2/\text{s}^2 \), \( ΔV_{\text{TMI}} \approx 3.662 \text{ km/s} \), on a 283-day “minimum-energy” trajectory out to Mars. The ERV then arrives at Mars (arrival \( V_{\text{inf}} \approx 3.480 \text{ km/s} \)) and propulsively captures into a 24-hr elliptical Mars orbit (~250 by 33,793 km, \( ΔV_{\text{MOC}} \approx 1.34 \text{ km/s} \)) in October 2031 where it remains until the crewed *Copernicus-2* MTV arrives during the next opportunity ~26 months later.

*Copernicus-2* departs LEO with its four crew on May 25, 2033 and arrives at Mars 100 days later. It then captures into the same parking orbit as the ERV on September 2, 2033. In the process, *Copernicus-2* uses up ~97 percent of its available propellant. To return to Earth, *Copernicus-2* rendezvous with the ERV and the forward crewed PL element is switched over to the ERV (shown in Fig. 13). No propellant transfer is required just a R&D maneuver. The reconfigured ERV then rendezvous with the orbiting Hab lander, followed by crew transfer, then EDL near the cargo lander at which point the surface exploration phase of the missions begins. After ~653 days on the surface, the crew lifts off using the MAV and returns to the ERV. Following the transfer of the crew and samples, the MAV is jettisoned. Before leaving Mars orbit, the crew finally jettisons the exterior consumables container and connecting tunnel from the PL element to reduce vehicle mass and propellant consumption. The ERV then performs the TEI burn on July 1, 2035 and begins the 90-day transfer back to Earth. On the final approach to Earth, the ERV does a final slowdown burn (\( ΔV \approx 2.257 \text{ km/s} \)) after which the crew separates and re-enters using the Orion MPCV capsule. The ERV flies past Earth at a sufficiently high altitude and is disposed of into heliocentric space.
For the outbound mission leg, the *Copernicus-2* spacecraft (Fig. 14) has an IMLEO of ~402 t consisting of the NTPS (~134.5 t), the in-line tank (~92.4 t), the integrated saddle truss/LH$_2$ drop tank assembly (~122.8 t), and the crewed PL element (~52.3 t). The NTPS, in-line and drop tank elements (again using composite tanks) carry ~87.2, ~71.3, and ~95 t respectively, for a total LH$_2$ propellant loading of ~253.5 t. By using two separate vehicles for the crewed mission, three engines are again sufficient for each. With ~75.3 klbf of total thrust, an $I_{sp}$ of 940 s, and ~239 t of usable propellant, the total engine burn time for this mission is ~109.6 min. The first TMI perigee burn is the longest at ~52.6 min. After this burn, the drop tank is drained and jettisoned to reduce vehicle mass and propellant consumption during the second perigee burn (~28.8 min). The NTPS supplies all the propellant needed for MOC (~28.2 min).

The “round trip” ERV has an IMLEO of ~360.3 t consisting of the NTPS (~131.3 t), the in-line tank (~101.4 t) and the integrated saddle truss/LH$_2$ drop tank assembly (~122.9 t). The composite tanks used in the ERV’s NTPS, in-line and drop tank elements carry ~87.2, ~77, and ~95 t respectively, for a total propellant loading of ~259.2 t. The ERV’s NTPS performs five primary burns (and four restarts) and its three engines provide ~75.3 klbf of total thrust. With ~244.3 t of usable propellant and a $I_{sp}$ of ~940 s, the total engine burn time for this mission is ~112.1 min. The first TMI perigee burn is the longest at ~41.3 min. After this burn, the drop tank is drained and jettisoned. The durations of the remaining burns are as follows: second perigee burn (~13.7 min), MOC (~13 min), TEI (~28.6 min), and final slowdown burn (~15.5 min). The long TEI burn is attributed to the addition of *Copernicus-2*’s ~40.2 t crewed PL mass plus the high TEI $\Delta V$ requirement (~3.075 km/s) for the 90-day transit back to Earth. The total and longest single engine burn times required for the ERV mission while substantial, are still below the capabilities demonstrated by the NRX-XE (~2 hr) and NRX-A6 engines (~62 min).
Lastly, the total mission IMLEO for the two spacecraft is ~762.4 t. Seven HLV launches are required for the split mission option—four for the Copernicus-2 and three for the ERV—compared to eight total launches required for the “all up” Copernicus-3a configuration. No on-orbit assembly (e.g., attachment of multiple drop tanks) is required just autonomous R&D of several vehicle components.

6.0 Summary and Conclusions

In NASA’s recent DRA 5.0 study (NASA/SP—2009-566), the fast conjunction long surface stay mission was identified as the preferred approach for exploring Mars. Plenty of time for scientific investigation and study, coupled with short ~6-month transit times to reduce crew exposure to the 0-gE and radiation environments of space were key reasons for this selection. Similarly, NTP was identified as the preferred propulsion option because of its high performance (100 percent higher Isp than LOX/LH2 chemical propulsion), its demonstrated technology record, plus the strong synergy that exists between it and chemical propulsion systems and stages. Recent measurements of the energetic particle radiation environment inside MSL spacecraft during its journey out to Mars indicate that astronauts could receive a radiation dose of ~0.66 Sv (~66 rem)—the limiting value established by NASA—during a 1-year journey out to Mars and back. With the potential for the crew to receive additional dose during the exploration phase of the mission, several questions immediately arise: Can NTP’s performance capability be used to reduce transit times further and by how much? Also, what are the engine operating requirements and vehicle design impacts?

The answer to the first question is yes, and analysis presented here indicates transit time reductions as much as 50 percent are possible. The key requirements are more propellant and reduced U-235 loading in the engine’s reactor fuel elements to allow higher operating temperatures. For the 2033 mission opportunity baselined in this paper, Copernicus can achieve 1-way transit times of ~150 days when filled to its maximum propellant capacity of ~190 t. To shorten transit times further, the Copernicus spacecraft can be reconfigured in incremental steps that allow it to carry larger quantities of propellant needed for
these more demanding missions. To reduce the 1-way transit times from 150 to 130 days, an in-line tank is inserted between Copernicus’ NTPS and saddle truss/drop tank assembly that increases the vehicle’s propellant capacity by ~42 percent to ~270 t. A fourth engine is also added to the NTPS to reduce the total engine operating time and minimize g-losses.

To achieve transit times in the range of 120 to 100 days, Copernicus’ saddle truss/drop tank assembly is replaced by a 4-sided star truss to which paired sets of LH₂ drop tanks are attached as needed. For 120-day transits to Mars and back, twin drop tanks—each carrying ~80 t of propellant—are used. Further reductions in transit times from 120 to 100 days requires an ~42 percent increase in total mission ΔV—from ~8.83 to ~12.5 km/s. To compensate, the crew size is reduced to four, composite propellant tanks are utilized and the engines are operated at their maximum temperature capability to achieve a specific impulse of ~940 s. Four drop tanks—each carrying ~82 t of propellant—are also used. The “all up” vehicles for 120- and 100-day cases require 5 and 7 HLVs respectively, as well as on-orbit assembly. The engine operational requirements for the 100-day case also exceed those demonstrated in the NERVA program.

A “split crewed mission architecture” is outlined that uses the Copernicus-2 spacecraft configuration plus an ERV. It can reduce vehicle size and complexity, eliminate the need for orbital assembly and an extra engine, and can achieve transit times as low as 90 days. Finally, it is important to note that both the total and longest single engine burn times required for the split mission, while substantial, are still below the capabilities demonstrated by the NRX-XE (~2 hr) and NRX-A6 engines (~62 min).

References


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Borowski, Stanley, K.; McCurdy, David, R.; Packard, Thomas, W.

**7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES)**
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**14. ABSTRACT**
The “fast conjunction” long surface stay mission option was selected for NASA’s recent Mars Design Reference Architecture (DRA) 5.0 study because it provided adequate time at Mars (~540 days) for the crew to explore the planet’s geological diversity while also reducing the “1-way” transit times to and from Mars to ~6 months. Short transit times are desirable in order to reduce the debilitating physiological effects on the human body that can result from prolonged exposure to the zero-gravity (0-gE) and radiation environments of space. Recent measurements from the RAD detector attached to the Curiosity rover indicate that astronauts would receive a radiation dose of ~0.66 Sv (~66 rem) — the limiting value established by NASA during their 1-year journey in deep space. Proven nuclear thermal rocket (NTR) technology, with its high thrust and high specific impulse (Isp ~900 s), can cut 1-way transit times by as much as 50 percent by increasing the propellant capacity of the Mars transfer vehicle (MTV). No large technology scale-ups in engine size are required for these short transit missions either since the smallest engine tested during the Rover program—the 25 klbf “Pewee” engine—is sufficient when used in a clustered arrangement of three to four engines. The “Copernicus” crewed MTV developed for DRA 5.0 is a 0-gE design consisting of three basic components: (1) the NTP stage (NTPS); (2) the crewed payload element; and (3) an integrated “saddle truss” and LH2 propellant drop tank assembly that connects the two elements. With a propellant capacity of ~190 t, Copernicus can support 1-way transit times ranging from ~150 to 220 days over the 15-year synodic cycle. The paper examines the impact on vehicle design of decreasing transit times for the 2033 mission opportunity. With a fourth “upgraded” SLS/HLV launch, an “in-line” LH2 tank element can be added to Copernicus allowing 1-way transit times of 130 days. To achieve 100 to 120 day transit times, Copernicus’ saddle truss/drop tank assembly is replaced by a “star truss” assembly with paired modular drop tanks to further increase the vehicle’s propellant capacity. The HLV launch count increases (from ~5 to 7) and a fourth engine is needed to reduce total mission burn time and gravity losses. Using a “split mission” approach, the NTPS, in-line tank and the saddle truss/LH2 drop tank elements can be configured as a pre-deployed Earth Return Vehicle/propellant tanker supporting 90-day crewed mission transits. The split mission approach also eliminates the need for on-orbit assembly. Mission scenario descriptions, key features and operational characteristics for five different vehicle configurations are presented.

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