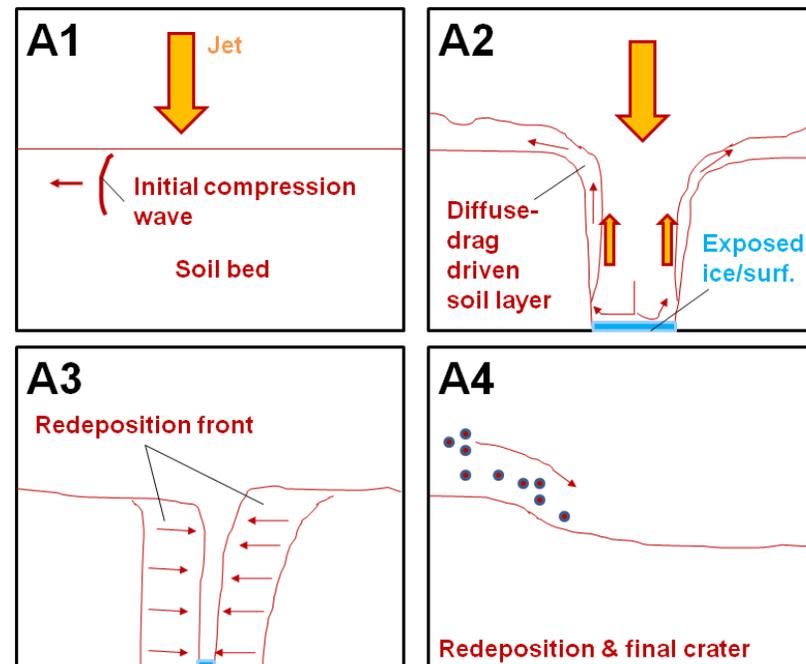


Quantification of plume-soil interaction and excavation due to the sky crane descent stage

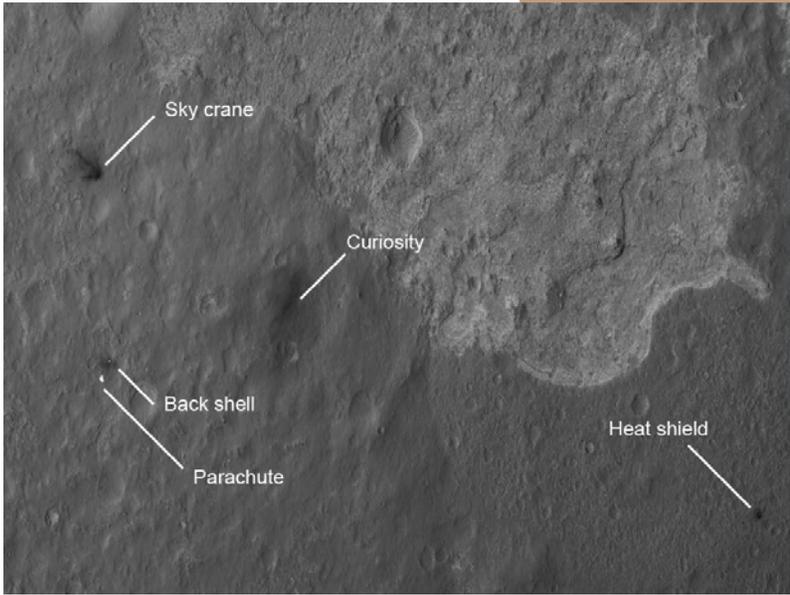
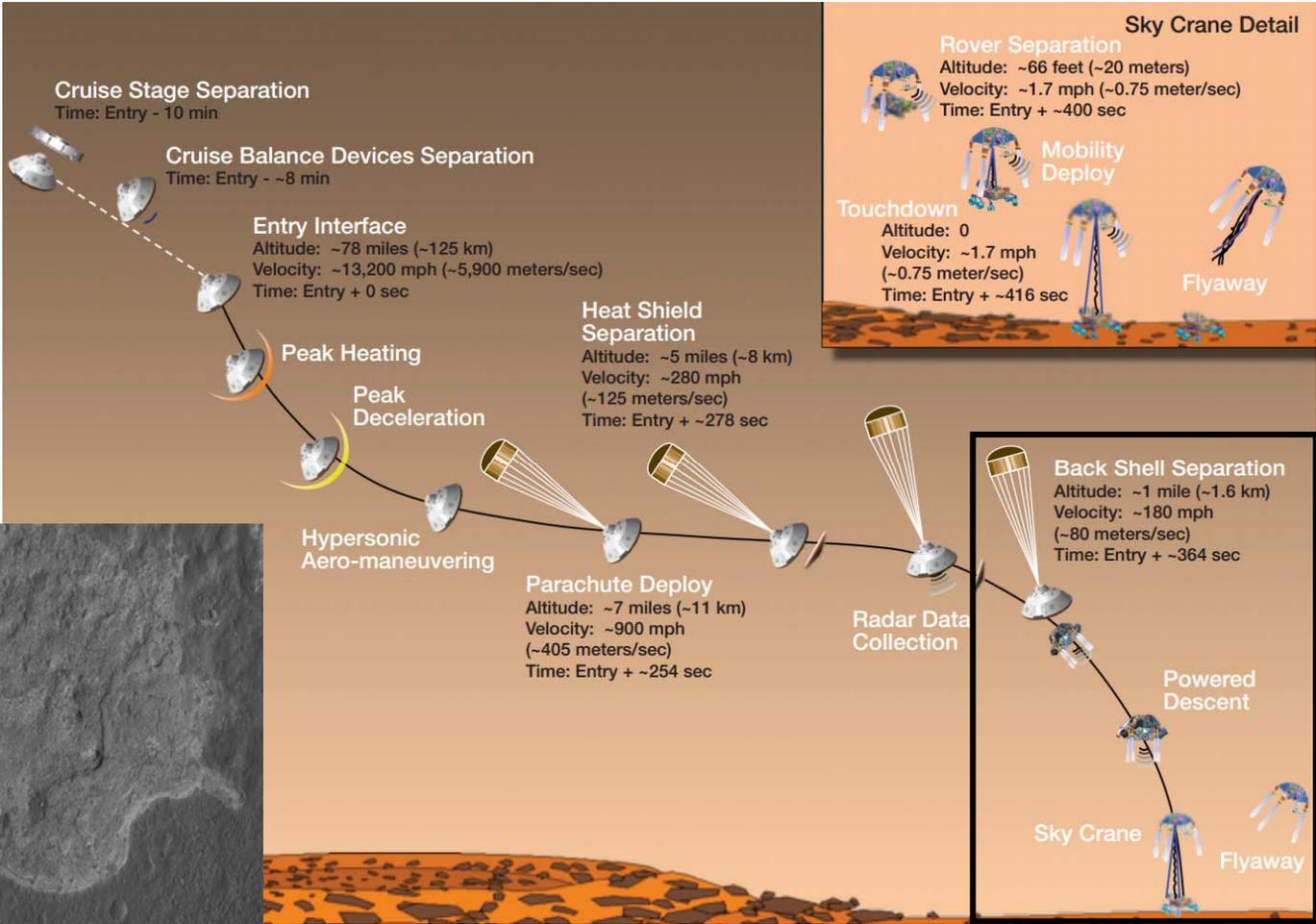
Jeff Vizcaino
Manish Mehta

Plume Induced Erosion

- Viscous Shear Erosion (VSE)
 - Dusting or saltation caused by shear forces of fluid friction parallel to the surface.
- Bearing Capacity Failure (BCF)
 - The formation of transient craters with steep walls occurs by a combination of two distinct processes termed bearing capacity failure (BCF) diffusion-driven flow (DDF) when the jet's ground pressure exceeds the shear strength of the soil.



Descent Profile



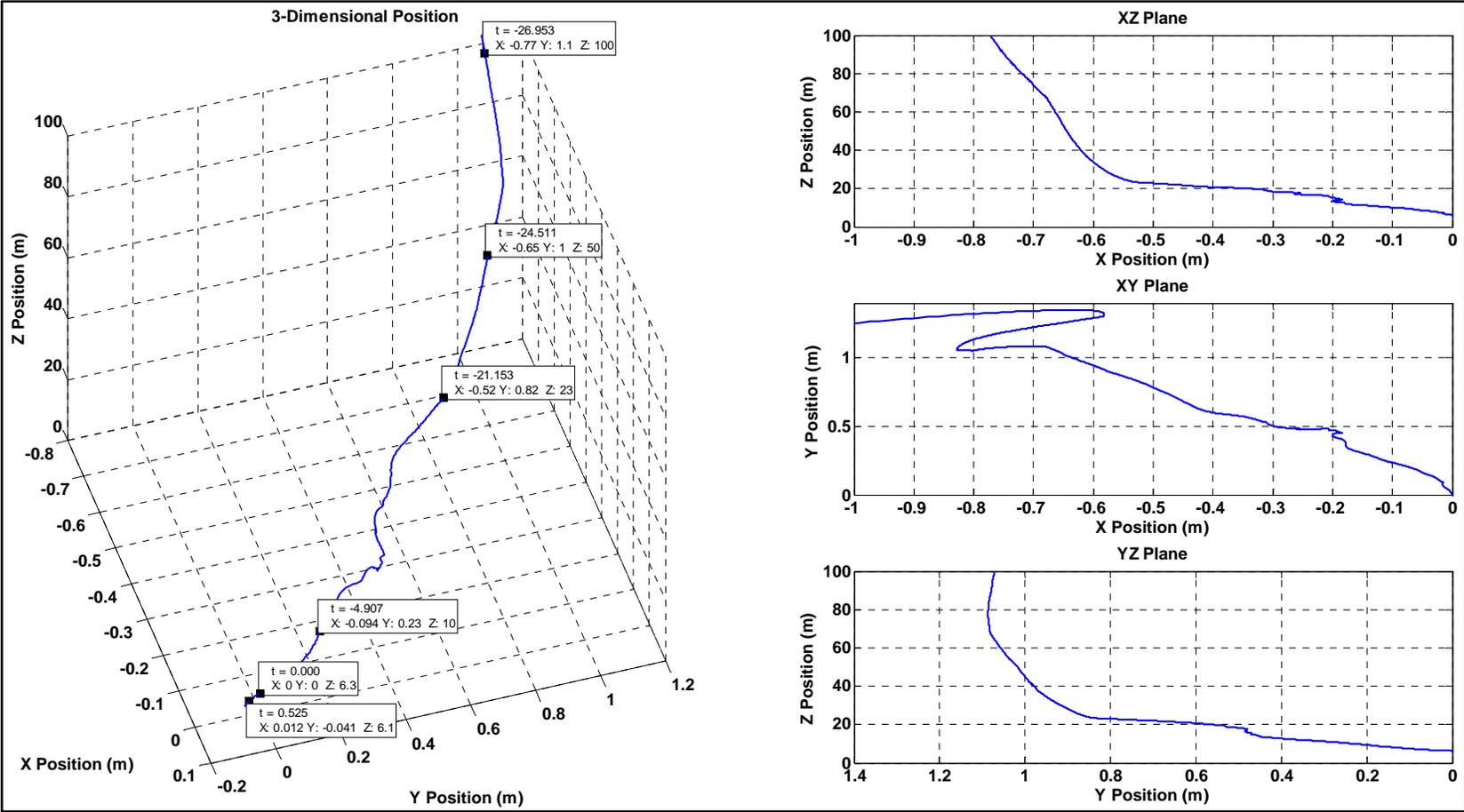
Mars Descent Imager (MARDI)

- Erosion Onset (VSE):
Approximately 63 meters
- Severe Erosion (BCF):
~18 meters



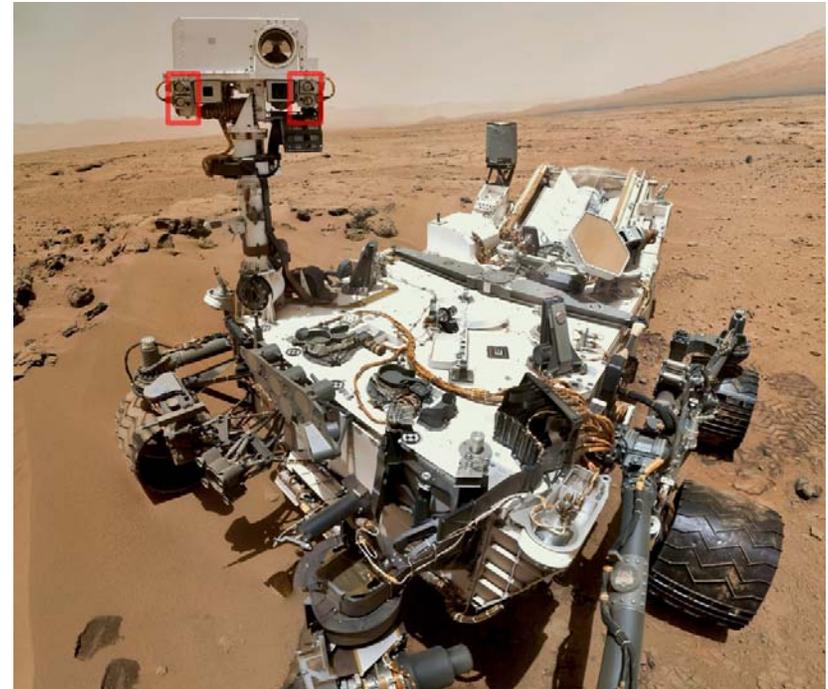
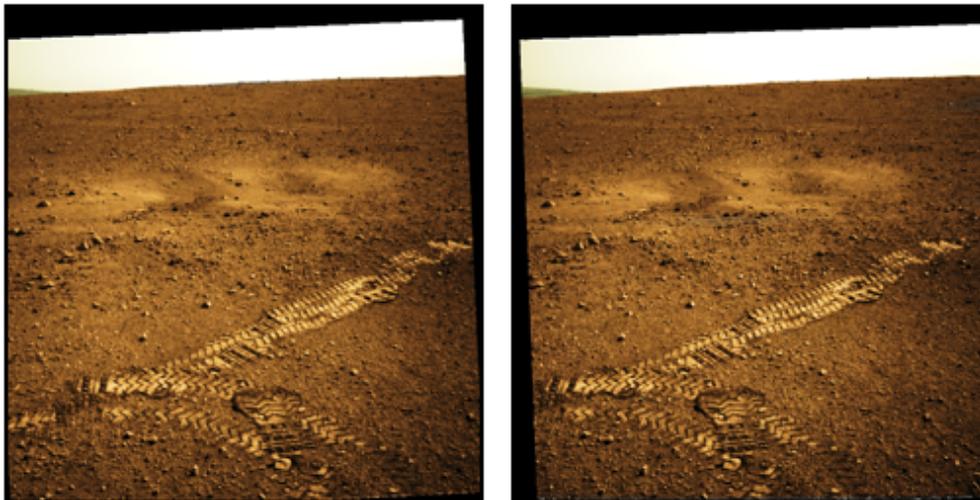
MARDI images taken during descent phase. Correlated MARDI images with trajectory data shows rocket thruster – soil interaction occurs at roughly 63 meters above the surface at 25 seconds before touchdown. Significant erosion occurs within the last 15 seconds.

Descent Ground Track



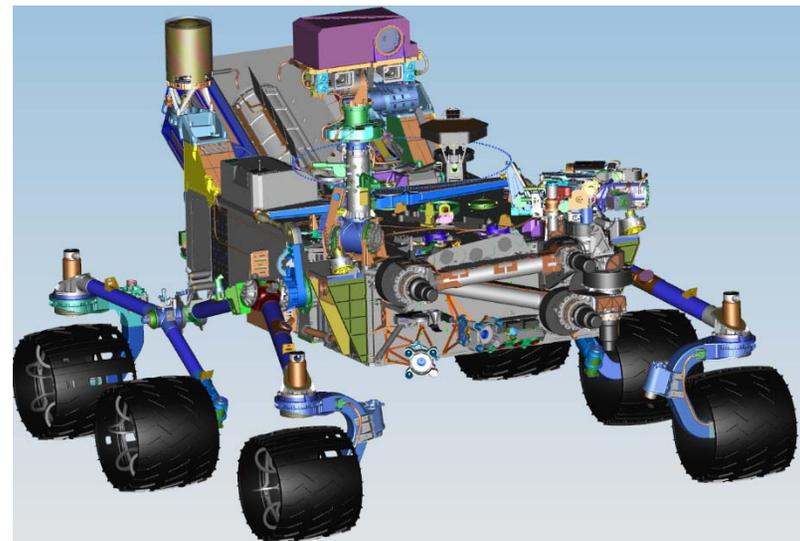
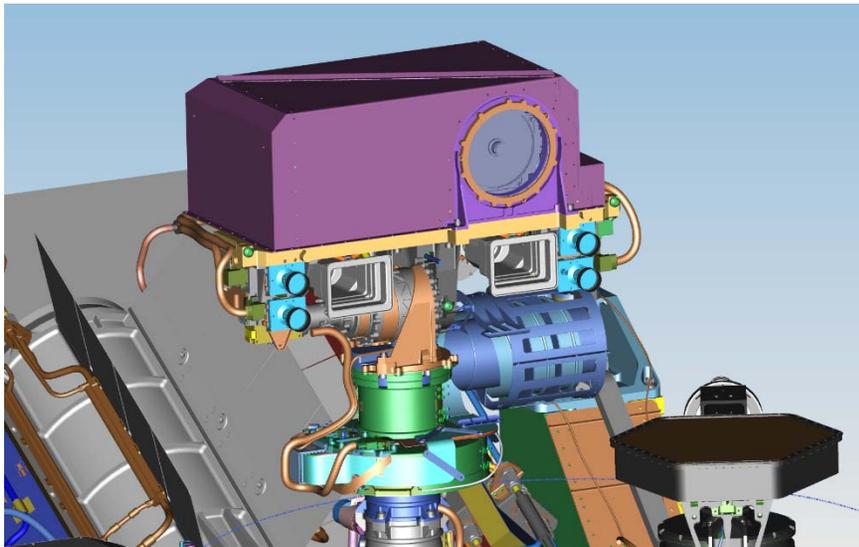
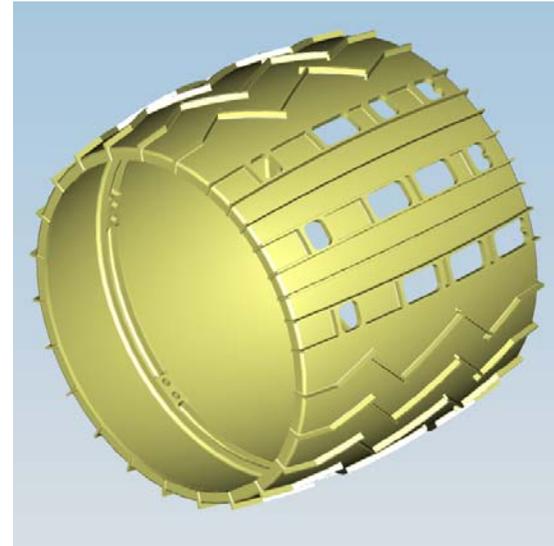
Digital Terrain Mapping

- Image Stereoscopy using MSL NAVCAMs
 - 1024 x 1024 (1 MP resolution)
 - 42 cm spacing
- Composition with Cardinal Systems VR Mapping software
 - Calibration
 - DTM generation



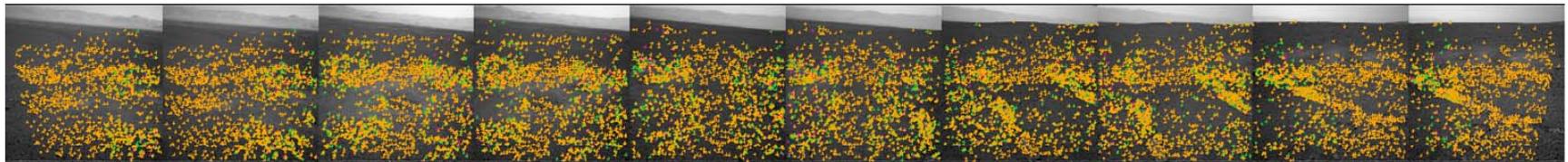
Mapping Calibration: CAD Models

- Detailed engineering CAD model of the Curiosity rover used to provide scaling information for the 3D terrain model.
 - Tire Width
 - Center Tread Length , Tread Spacing, and Wheel Spacing
 - “JPL” Morse Code Spacing and width

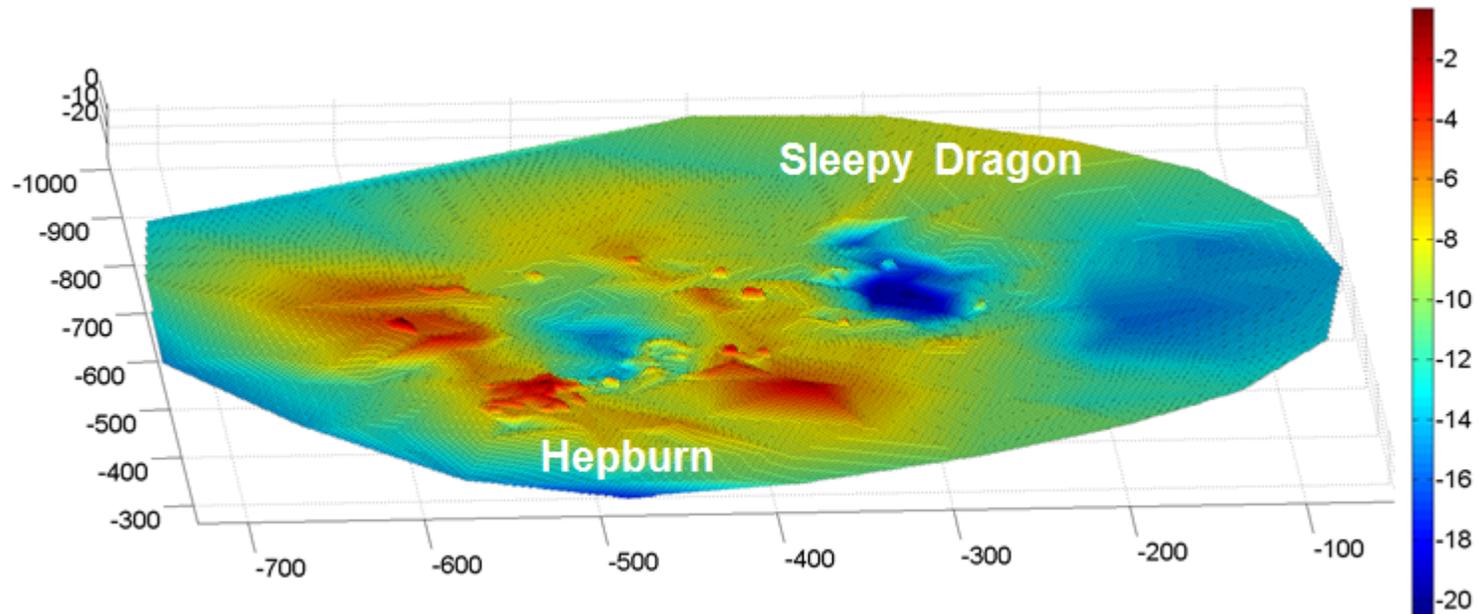


Mapping Calibration: Camera Calibration

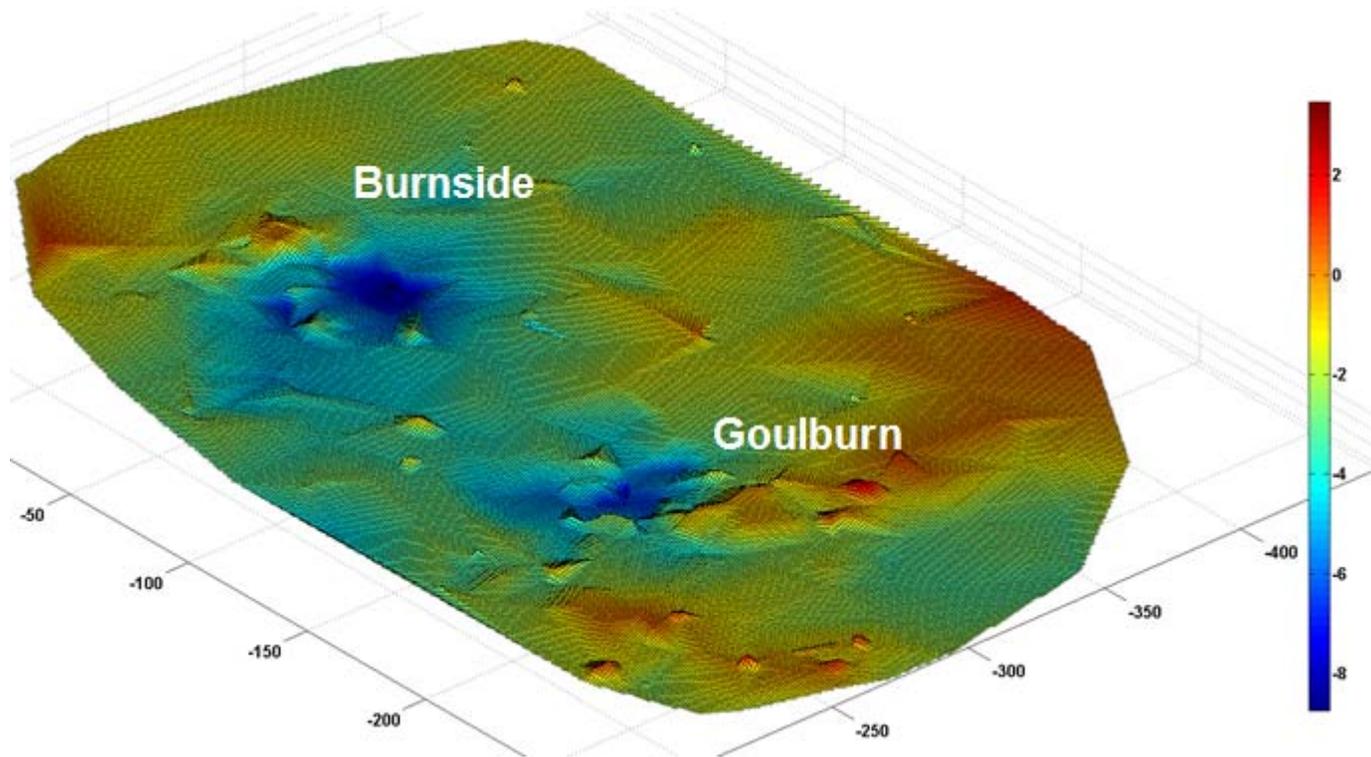
- Worked closely with the Cardinal Systems to develop a process to increase the accuracy of the model. “Autotie” function correlates image measurements and scales across multiple images.



Sleepy Dragon & Hepburn



Burnside & Goulburn



Erosion Rates & Volumes

Eroded Volume / Rate

Test	Test #	Soil Type	Volume (m ³)	Time (s)	Volumetric Erosion Rate (m ³ /s)
MSL	8	Mars (Fine Sand)	3.29E-04	0.94	3.51E-04
MSL	7	Mars (Fine Sand)	4.91E-04	0.96	5.12E-04
MSL - Flight	Goulburn	N/A	1.09E-02	15	7.27E-04
MSL	2	Mars (Fine Sand)	5.86E-04	1.1	5.35E-04
MSL	23	Mars (Coarse)	9.51E-04	1	9.51E-04
MSL	13	Mars (Fine Sand)	1.07E-03	1	1.07E-03
MSL	21	Mars (Coarse)	1.11E-03	0.95	1.17E-03
MSL	6	Mars (Fine Sand)	1.10E-03	0.94	1.17E-03
MSL	16	Mars (Coarse)	1.44E-03	0.95	1.52E-03
MSL	5	Mars (Fine Sand)	1.65E-03	0.98	1.69E-03
MSL	12	Mars (Fine Sand)	1.97E-03	0.98	2.01E-03
MSL	18	Mars (Coarse)	1.64E-03	0.69	2.39E-03
MSL	20	Mars (Coarse)	2.92E-03	1	2.93E-03
MSL	4	Mars (Fine Sand)	2.35E-03	0.78	3.01E-03
MSL	15	Mars (Coarse)	3.85E-03	1	3.85E-03
MSL - Flight	Burnside	N/A	5.38E-02	15	3.59E-03
MSL - Flight	Hepburn	N/A	6.36E-02	15	4.24E-03
MSL - Flight	Sleepy Dragon	N/A	8.47E-02	15	5.65E-03
MSL	22	Mars (Coarse)	5.01E-03	0.89	5.65E-03
MSL	17	Mars (Coarse)	6.69E-03	0.96	6.98E-03
MSL	19	Mars (Coarse)	6.88E-03	0.97	7.09E-03
MSL	14	Mars (Coarse)	1.67E-02	0.97	1.72E-02

Estimated Eroded Mass

Test	Test #	Soil Type	Eroded Mass (m ³)
MSL - Flight	Goulburn	Mars (Coarse)	4.25
MSL - Flight	Burnside	Mars (Coarse)	21.0
MSL - Flight	Hepburn	Mars (Coarse)	24.8
MSL - Flight	Sleepy Dragon	Mars (Coarse)	33.0

Conclusions

- The MARDI camera allowed us to determine, for the first time, when plume-soil interaction began to occur which was found to be approximately 63 meters above the ground level.
 - Soil erosion continuously increased and visibility decreased as the Sky Crane descends to its final altitude.
- The data extraction methods employed to the MSL data were a value added benefit and was performed without a need to alter the vehicle or collaborate with mission planners during design phase.
 - Scientists analyzing future human and robotic missions utilizing dual stereo camera systems will also benefit from this method as well.
- The effect of subsoil bedrock had a significant effect, as predicted, reducing overall crater diameter and depth (Goulburn and Burnside) when compared to a similar region with loosely packed soil (Hepburn and Sleepy Dragon).
 - The increased thrust loads associated with any possible human Mars mission will naturally intensify the erosion problem. If future vehicles are to use retro-propulsive landing system, it will be important to either choose landing sites with solid foundations or construct a landing site ahead of time.
- Volumetric erosion rates agreed well with experimental tests conducted in similar conditions using soil simulants.
 - This agreement validates vacuum chamber testing methodologies for analyzing plume-soil erosion and will allow for better prediction of erosion rates for similar and derived vehicles in the future.

Future Work & Implications

- This data is integrated into an erosion database compiled from surveys of simulated lunar and Mars plume impingement based erosion for development of empirical models to predict erosion severity.
- Validation future vacuum chamber tests, empirical codes, and CFD codes for plume based erosion.
- Potential for use of plume based erosion as a method of excavation as an alternative for traditional methods

Acknowledgements

- Cardinal Systems for their excellence in support and expertise
- Orrin Thomas of Cardinal Systems for his assistance in creating the geometrical model calibration that was the basis for this analysis.
- Anita Sengupta of NASA's JPL for her support and provisioning of the MSL data.