Mass wasting in planetary environments: Implications for seismicity

RENEE WEBER · NASA MARSHALL SPACE FLIGHT CENTER · AMANDA NAHM · UNIV. OF IDAHO
NICK SCHMERR · UNIV. OF MARYLAND

Overview

On Earth, mass wasting events such as rock falls and landslides are well known consequences of seismic activity. Here we investigate the regional effects of seismicity in planetary environments with the goal of determining whether such surface features on the Moon, Mars, and Mercury could be triggered by fault motion (Fig. 1).

Lobate scarps

Lobate scarps, the typical surface expressions of thrust faults resulting from tectonic compression, are widely observed on the Moon, Mars, and Mercury (Figs. 2&3). Compared to other types of faults, surface-cutting thrust faults require the largest amount of stress to form and/or slip, so they could possibly generate large quakes. While normal faults, graben, and wrinkle ridges may be more abundant on Mars, the Moon, and Mercury respectively, these structures would create smaller theoretical maximum quakes than lobate scarp thrust faults. Thus, we optimize our chances of finding mass wasting associated with faults by studying lobate scarps.

We derive the theoretical maximum quake magnitude a given scarp can produce from basic fault properties. These are either estimated from imagery or derived from laboratory rock experiments or elastic dislocation models, and include the length (L), dip angle (θ), depth of faulting (T), displacement (D), and fault width (W) (Fig. 4).

To determine the dimensions of an area affected by seismic shaking, we model the ground motion resulting from the theoretical maximum quake along a given fault (Fig. 5). We use a numerical code for simulating seismic wave propagation through arbitrary elastic and anelastic media in a 3-D model space (including topography). Peak vertical ground motion typically occurs within a few kilometers of the main shock and drops off rapidly from there. This implies that we should expect most of the mass wasting phenomena to occur in the immediate vicinity of the fault. However, this result may depend on regional effects like surface slope and megaregolith thickness. A thicker megaregolith (as might be expected in the vicinity of craters) would tend to focus shaking in some of the crater basins. Sediments can also enhance seismic shaking; this could be a relevant scenario for Martian craters that may have been lakes at some time in the past.