Planetary Crater Detection and Registration Using Marked Point Processes, Multiple Birth and Death Algorithms, and Region-based Analysis

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Introduction

Crater Detection
- Marked Point Process Model
- Energy Function
- Multiple Birth and Death Algorithm
- Region-of-Interest Approach
- Experimental Results

Image Registration
- 2-step Approach
- Experimental Results

Conclusion
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**Introduction**

**Need for automated methods for image registration**

- Launch of several planetary missions
- Design of new and powerful sensors

**Large data sets**
- Multisensor
- Multitemporal
- Both

**Objective**

- **Crater detection** in planetary images
- Development of an image registration method based on the extracted features
Content

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Crater detection based on a marked point process (MPP) model

Marked Point Processes

MPP: Stochastic Process \rightarrow \text{Realizations} \rightarrow \text{Configurations of objects, each described by a marked point}

Mathematical Formulation

A point process \( X \), defined over a bounded subset \( P \) of \( \mathbb{R}^2 \) maps from a probability space to a configuration of points in \( P \).

Realizations of the process \( X \) are random configurations \( x \) of points, \( x = \{x_1, ..., x_n\} \), where \( x_i \) is the location of the \( i \)th point in the image plane (\( x_i \in P \)).

A configuration of an MPP consists of a point process whose points are enriched with additional parameters, called marks and aimed at parameterizing objects linked to the points.

Bayesian approach: Maximum a posteriori (MAP) rule to fit the model to the image is equivalent to minimizing an energy function (computationally challenging)
Marked Point Process for Crater Detection

Craters

Center: \((x_0, y_0)\)
Major Axis: \(a\)
Minor Axis: \(b\)
Orientation: \(\theta\)

Marked Point Process for Crater Detection

Point Process
\[ x = \{x_1, ..., x_7\} \]

Parameters
\[ x_i = (x_{0i}, y_{0i}, a_i, b_i, \theta_i) \]
Energy function of the configuration $X = \{x_i, x_2, ..., x_n\}$ wrt the extracted set $C$ of contour pixels (Canny):

$$U(X|C) = U_P(X) + U_L(C|X)$$

**Prior**

Repulsion coefficient based on the overlapping of the ellipses (overlapping craters are quite unlikely)

$$U_P(X) = \frac{1}{n} \sum_{x_i \land x_j > 0} \frac{x_i \land x_j}{x_i \lor x_j}$$

$x_i \lor x_j$ = area of union of ellipses $x_i$ and $x_j$

$x_i \land x_j$ = area of intersection of $x_i$ and $x_j$

**Likelihood**

Two terms, one based on a correlation measure, the other based on a distance measure (fit between contours and realization of $X$)

$$U_L(C|X) = \sum_{i=1}^{n} \left[ \frac{d_H(x_i^0, C)}{na_i} - \frac{|x_i^0 \cap C|}{|C|} \right]$$

$x_i^0$ = set of pixels corresponding to ellipse $x_i$ in the image plane

$d_H(x_i^0, C)$ = Hausdorff distance between ellipse $x_i$ and the contours:

$$d_H(A, B) = \max \left\{ \sup_{\alpha \in A} \inf_{\beta \in B} d(\alpha, \beta) ; \sup_{\beta \in B} \inf_{\alpha \in A} d(\alpha, \beta) \right\}$$

Classical distance between sets $d(A, B) = 0$
Computation of Death Probability and Death Phase according to Likelihood

Contour Extraction (Canny) and Parameter Initialization

Generation of the Birth Map used to speed up the convergence

Computation of the Birth Probabilities for each pixel

Birth Phase

Energy Computation for all the ellipses

Computation of Death Probability and Death Phase according to Likelihood

Configuration refinement according to Prior

Convergence Test

Markov chain Monte Carlo-type method
Simulated Annealing scheme

Markov chain sampled by a multiple birth and death (MBD) algorithm

Crater Detection – Energy Minimization
**MBD – Birth and Death Steps**

**Birth Step**

For each pixel $s$ in the image, compute the birth probability as $\min\{\delta \cdot B(s), 1\}$, where:

$$B(s) = \frac{b(s)}{\sum_s b(s)}$$

$b(s)$ is the birth map computed from the contour map using generalized Hough transform and Gaussian filtering.

**Death Step**

For each ellipse $x_i$ in the configuration, compute the death probability as $d(x_i)$:

$$d(x_i) = \frac{\delta \cdot a(x_i)}{1 + \delta \cdot a(x_i)}$$

$$a(x_i) = \exp[-\beta(U_L(X\{x_i\}|C) - U_L(X|C))]$$
Crater Detection – Region Based Approach

Region-Based Approach

Why?
- MBD is *computationally heavy*
- Computational burden increases with *image size*

Region Based Flowchart and Example

Initialization

Detection of **regions on interest** based on the birth map

MBD in each **region**

Aggregation
Crater Detection – Data Sets

- 6 THEMIS (Thermal Emission Imaging System) images, TIR, 100m resolution, Mars Odissey mission
- 7 HRSC (High Resolution Stereo Color) images, VIS, ~20m resolution, Mars Express mission
- Image sizes from 1581 × 1827 to 2950 × 5742 pixels

Quantitative Performance Assessment of the crater detection algorithm:
Detection Percentage (D), Branching Factor (B), and Quality Percentage (Q)

<table>
<thead>
<tr>
<th>Data</th>
<th>D = ( \frac{TP}{TP + FN} )</th>
<th>B = ( \frac{FP}{TP} )</th>
<th>Q = ( \frac{TP}{TP + FP + FN} )</th>
</tr>
</thead>
<tbody>
<tr>
<td>Avg on all THEMIS</td>
<td>0.91</td>
<td>0.10</td>
<td>0.83</td>
</tr>
<tr>
<td>Avg on all HRSC</td>
<td>0.89</td>
<td>0.06</td>
<td>0.85</td>
</tr>
<tr>
<td>Avg on all images</td>
<td>0.90</td>
<td>0.09</td>
<td>0.84</td>
</tr>
</tbody>
</table>
Crater Detection – Results

<table>
<thead>
<tr>
<th>Crater</th>
<th>( C = (x_0, y_0) )</th>
<th>Semi-axes ( (a, b) )</th>
<th>Orientation ( \theta )</th>
</tr>
</thead>
<tbody>
<tr>
<td>Crater 1</td>
<td>(139, 393)</td>
<td>(35, 33)</td>
<td>64°</td>
</tr>
<tr>
<td>Crater 2</td>
<td>(258, 756)</td>
<td>(51, 50)</td>
<td>115°</td>
</tr>
<tr>
<td>Crater 3</td>
<td>(343, 23)</td>
<td>(13, 12)</td>
<td>180°</td>
</tr>
<tr>
<td>Crater 4</td>
<td>(591, 215)</td>
<td>(19, 18)</td>
<td>31°</td>
</tr>
<tr>
<td>Crater 5</td>
<td>(919, 157)</td>
<td>(15, 14)</td>
<td>106°</td>
</tr>
</tbody>
</table>
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Why a 2-step Optimization?

**Feature-based registration**
- Min Hausdorff distance ($d_H$) between extracted craters through genetic algorithm
- Fast but sensitive to accuracy of crater maps

**Area-based registration**
- Max Mutual Information ($MI$) through genetic algorithm
- Highly accurate but computationally heavy

**Initialization**

**Fast registration** based on extracted craters $\rightarrow \hat{p}$

**Refinement**: registration based on mutual information in a neighborhood of $\hat{p} \rightarrow p^*$
Image Registration – Region of Interest

Transformation found for an interactively selected region of interest \( \rightarrow p_B^* \)
\[
p_B = (t_x, t_y, \theta, k)
\]

Transformation derived for the entire Image \( \rightarrow p_A^* \)
\[
p_A = (T_x, T_y, \beta, \alpha)
\]

\[
p_A^* = \begin{pmatrix}
-k \cos(\theta) x_0 - k \sin(\theta) y_0 + t_x + x_0 \\
k \sin(\theta) x_0 - k \cos(\theta) y_0 + t_y + y_0 \\
\theta \\
k
\end{pmatrix}
\]

Superposition of Reference and Input
Image Registration – Data Sets

Semi-simulated image pairs

20 pairs composed of one real THEMIS or HRSC image and of an image obtained by applying a synthetic transform and AWGN

Quantitative validation with respect to the true transform (RMSE)

Real multi-temporal image pairs

Real multi-temporal pair of LROC (Lunar Reconnaissance Orbiter Camera) images

100m resolution

Only qualitative visual analysis is available, as no ground truth is available
## Registration Results with Semi-synthetic Data

Below are the RMSE values for different data sets and steps of the registration process.

<table>
<thead>
<tr>
<th>Data set</th>
<th>RMSE [pixel]</th>
<th>Left Image</th>
<th>Right Image</th>
</tr>
</thead>
<tbody>
<tr>
<td>THEMIS (10 data sets)</td>
<td>0.31</td>
<td>$p_{GT}$ (7.05, 35.91, 0.18°, 1.071)</td>
<td>(76.59, 19.96, 2.17°, 1.031)</td>
</tr>
<tr>
<td>HRSC (10 data sets)</td>
<td>0.22</td>
<td>$p^*$ (7.04, 35.92, 0.19°, 1.071)</td>
<td>(76.41, 20.06, 2.18°, 1.031)</td>
</tr>
<tr>
<td>Average (20 data sets)</td>
<td>0.26</td>
<td>RMSE 1\text{st} Step 0.79</td>
<td>0.51</td>
</tr>
<tr>
<td></td>
<td></td>
<td>RMSE 2\text{nd} Step 0.16</td>
<td>0.33</td>
</tr>
</tbody>
</table>
Registration Results with Real Data

Visually accurate matching between reference and registered images in the real multitemporal data set

Checkerboard representation of the registered images (zoom on details)
Registration Results with Real Data

Visually accurate matching between reference and registered images in the real multitemporal data set.
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Conclusions

- **Accurate** crater maps, useful for both image registration and planetary science, were obtained from data from different sensors.

- **Higher accuracy** as compared to previous work on crater detection (not shown for brevity)

- Reduced time for convergence thanks to a **region-based approach**

- **Sub-pixel accuracy and visual precision** in registration: effectiveness of the proposed 2-step registration method

Future Developments

- Test in conjunction with a **parallel** implementation (e.g. computer cluster)

- Validation with **multi-sensor** real images

- Extension to **other applications** requiring the extraction of ellipsoidal or circular features, e.g. optical Earth observation images or medical images
Short Bibliography

For each pixel in the image compute the Birth Probability as \( \min\{\delta \cdot B(s), 1\} \), where:

\[
B(s) = \frac{b(s)}{\sum_s b(s)}
\]

Being \( b(s) \) the Birth Map computed from the Canny Contour Map
For each ellipse $x_i$ in the configuration compute the **Death Probability** as $d(x_i)$, where

$$d(x_i) = \frac{\delta \cdot a(x_i)}{1 + \delta \cdot a(x_i)}$$

and

$$a(x_i) = e^{-\beta (U_L(x|x_i|I_g) - U_L(x|I_g))} = e^{\beta \cdot U_L^I(x_i|I_g)}$$

The complete **Flowchart** of the **Death Step** is as follows:

1. **Computation of the Energy Terms**
2. **Computation of Death Prob. based on Likelihood**
3. **Elimination of ellipses with prob. $d(x_i)$**
4. **Configuration refinement thanks to Prior term**
**Similarity Measures**

**Hausdorff Distance**

\[
\text{Similarity} = \text{mean}_c \left\{ \sum_{i=1}^{N^c} \sum_{t=1}^{P} d_H(x_i^c, x_t) \right\}
\]

- \(c\) = craters in Input Image
- \(N^c = \text{sum}(\text{pixels in crater } c \text{ in Input Image})\)
- \(P = \text{sum}(\text{craters' border pixels in Ref Image})\)
- \(x_i^c = \text{coord of pixel } i \text{ in crater } c \text{ in Input Image}\)
- \(x_t = \text{coord of pixel } t \text{ in Ref Image's craters}\)

**Mutual Information**

\[
MI(X, Y) = \sum_{x\in X} \sum_{y\in Y} p_{X,Y}(x,y) \log \left( \frac{p_{X,Y}(x,y)}{p_X(x) p_Y(y)} \right)
\]

- \(X\): pixel intensity in Reference Image
- \(Y\): pixel intensity in Input Image
- \(p_X(x)\): probability density function (pdf) of \(X\)
- \(p_Y(y)\): probability density function (pdf) of \(Y\)
- \(p_{X,Y}(x,y)\): joint pdf of \(X\) and \(Y\)

\(p_X(x)\), \(p_Y(y)\), \(p_{X,Y}(x,y)\)

estimated through the corresponding image histograms
RST Transformation

Rotation – Scale – Translation Transformation

Transformation vector

\[ p = (t_x, t_y, \theta, k) \]

\( \{t_x, t_y\} \): Translations in \( x \) and \( y \)

\( \theta \): Rotation angle

\( k \): Scaling Factor

Matrix Formulation

\[ T_p(x, y) = \begin{pmatrix} k \cos(\theta) & k \sin(\theta) \\ -k \sin(\theta) & k \cos(\theta) \end{pmatrix} \begin{pmatrix} t_x \\ t_y \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} \]
Region of Interest Approach

$I_A(X, Y), I_B(x, y)$: Two Images

$I_B$: sub-image of $I_A$ such that $I_B(0,0) = I_A(x_0, y_0)$

$p_A = (T_x, T_y, \alpha, \beta)$: RST transformation vector transforming $I_A$ into $I_A^{tr}$

$p_B = (t_x, t_y, \theta, k)$: RST transformation vector transforming $I_B$ into $I_B^{tr}$

$I_B^{tr}(0,0) = I_A^{tr}(x_0, y_0)$

Given:

Transformation: $p_B$

Reference of Region: $(x_0, y_0)$

Find: Transformation: $p_A$

From the image

\[
\begin{align*}
X &= x + x_0 \\
Y &= y + y_0
\end{align*}
\]

Expressing the transformation in Matrix Form

\[
T_{p_A} = \begin{pmatrix}
\alpha \cos(\beta) & \alpha \sin(\beta) \\
-\alpha \sin(\beta) & \alpha \cos(\beta)
\end{pmatrix}
\begin{pmatrix}
T_x \\
T_y
\end{pmatrix}
\]

\[
T_{p_B} = \begin{pmatrix}
k \cos(\theta) & k \sin(\theta) \\
-k \sin(\theta) & k \cos(\theta)
\end{pmatrix}
\begin{pmatrix}
t_x \\
t_y
\end{pmatrix}
\]

This should also hold

$T_{p_A}(x + x_0, y + y_0) = (x' + x_0, y' + y_0)$

Plugging $T_{p_A}$ into this equation and replacing $x'$ and $y'$ according to $T_{p_B}$

\[
\begin{align*}
k \cos(\theta) x + k \sin(\theta) y + t_x + x_0 &= \\
\alpha \cos(\beta)(x + x_0) + \alpha \sin(\beta)(y + y_0) + T_x
\end{align*}
\]

\[
\begin{align*}
-k \sin(\theta) x + k \cos(\theta) y + t_y + y_0 &= \\
\alpha \sin(\beta)(x + x_0) + \alpha \cos(\beta)(y + y_0) + T_y
\end{align*}
\]

Knowing $\alpha = k$ and solving in $P_1 = (0,0)$ and $P_2 = (-x_0, -y_0)$

\[
p_A = \begin{pmatrix}
-k \cos(\theta) x_0 - k \sin(\theta) y_0 + t_x + x_0 \\
k \sin(\theta) x_0 - k \cos(\theta) y_0 + t_y + y_0 \\
\theta \\
k
\end{pmatrix}
\]
RMS Error Computation

Ground Truth Transformation
\[ p_{GT} = (t_{x1}, t_{y1}, \theta_1, k_1) \rightarrow T_{p_{GT}}(x, y) = Q_{p_{GT}} \cdot [x, y, 1]^T \]

Computed Transformation
\[ p = (t_x, t_y, \theta, k) \rightarrow T_p(x, y) = Q_p \cdot [x, y, 1]^T \]

Error Transformation
\[ p_e = (t_{xe}, t_{ye}, \theta_e, k_e) \rightarrow Q_{p_e} = Q_p \cdot Q_{p_{GT}}^{-1} \]
\[ k_e = \frac{k_2}{k_1}, \quad \theta_e = \theta_2 - \theta_1 \]
\[ t_{xe} = t_{x2} - k_e(t_{x1} \cos(\theta_e) + t_{y1} \sin(\theta_e)) \]
\[ t_{ye} = t_{y2} - k_e(t_{y1} \cos(\theta_e) - t_{x1} \sin(\theta_e)) \]

\[(x, y) \in \text{Image}, [x', y', 1]^T = Q_{p_e} \cdot [x, y, 1]^T \]

\[ [x'] = k_e \begin{pmatrix} \cos(\theta_e) & \sin(\theta_e) \\ -\sin(\theta_e) & \cos(\theta_e) \end{pmatrix} [x] + [t_{xe} \\

RMS Error: \[ E(p_e) = \sqrt{\frac{1}{AB} \int_0^A \int_0^B (x' - x)^2 + (y' - y)^2 \, dx \, dy}, \quad \alpha = A^2 + B^2 \]

\[ E^2(p_e) = \frac{1}{AB} \int_0^A \int_0^B (k_e \cos(\theta_e) x + k_e \sin(\theta_e) y + t_{xe} - x)^2 + (-k_e \sin(\theta_e) x + k_e \cos(\theta_e) y + t_{ye} - y)^2 \, dx \, dy \]

\[ E^2(p_e) = \frac{\alpha}{3} (k_e^2 - 2k_e \cos(\theta_e) + 1) + (t_{xe}^2 + t_{ye}^2) - (At_{xe}^2 + Bt_{ye}^2)(1 - k_e \cos(\theta_e)) - k_e (At_{ye} - Bt_{xe}) \sin(\theta_e) \]